



# Latest results from JUNO

Zeyuan Yu (IHEP)

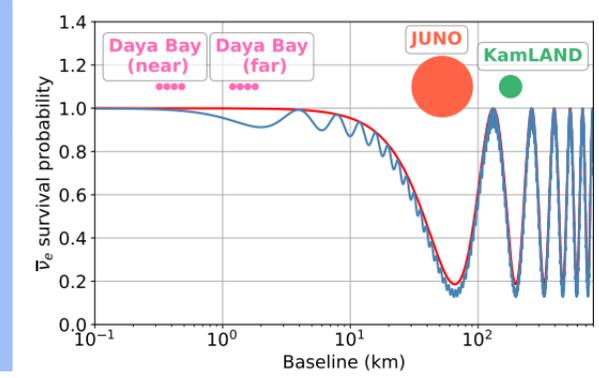
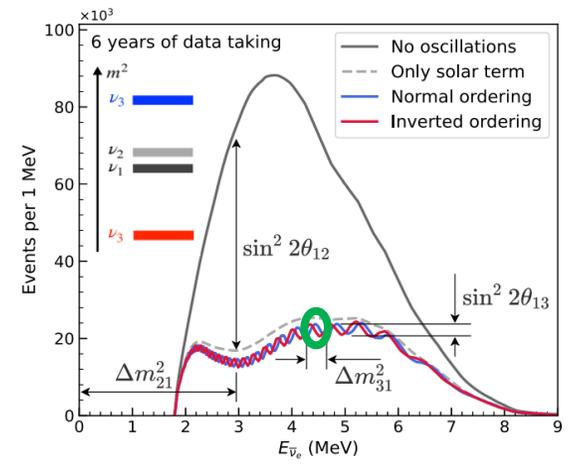
On behalf of the JUNO collaboration

March 17, 2026

# Jiangmen Underground Neutrino Observatory

JUNO studies how reactor  $\bar{\nu}_e$ 's oscillate (disappear) into other flavors

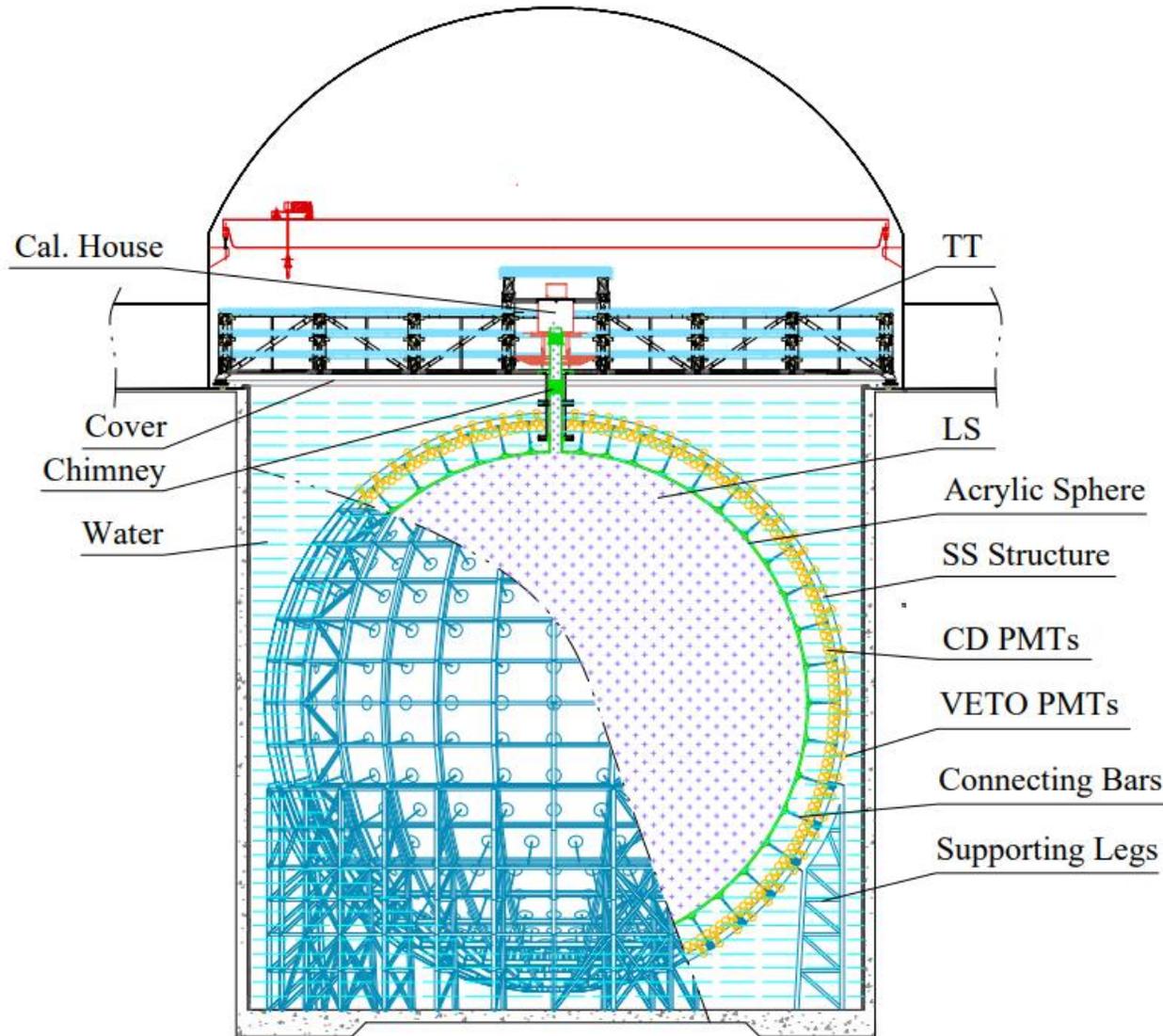
$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}(L, E) = 1 - \sin^2 2\theta_{12} \cos^4 \theta_{13} \sin^2 \frac{\Delta m_{21}^2 L}{4E} - \sin^2 2\theta_{13} \left( \cos^2 \theta_{12} \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \theta_{12} \sin^2 \frac{\Delta m_{32}^2 L}{4E} \right)$$



# JUNO collaboration

= 70 institutes, > 700 participants

| Country | Institute | Country | Institute  | Country      | Institute       |
|---------|-----------|---------|------------|--------------|-----------------|
| Armenia | YerPhI    | China   | SDU        | Italy        | INFN-CT         |
| Belgium | ULB       | China   | SJTU       | Italy        | INFN-FE         |
| Chile   | SAPHIR    | China   | SUSTech    | Italy        | INFN-LNF        |
| Chile   | UNAB      | China   | SYSU       | Italy        | INFN-MI         |
| China   | BISEE     | China   | THU        | Italy        | INFN-MIB        |
| China   | CDUT      | China   | UCAS       | Italy        | INFN-PD         |
| China   | CNPRI     | China   | WHU        | Italy        | INFN-PG         |
| China   | CUGB      | China   | WuYiU      | Italy        | INFN-Roma3      |
| China   | DGUT      | China   | XJTU       | Pakistan     | PINSTECH (PAEC) |
| China   | ECUT      | China   | XMU        | Russia       | JINR            |
| China   | GXU       | China   | ZZU        | Slovakia     | FMFI-UK         |
| China   | HIT       | Czech   | Charles U. | Taiwan-China | NCTU            |
| China   | HunanU    | Finland | JyU        | Taiwan-China | NKNU            |
| China   | IGGCAS    | France  | CPPM       | Taiwan-China | NTU             |
| China   | IHEG-CAGS | France  | IJCLab     | Taiwan-China | NTUT            |
| China   | IHEP      | France  | IPHC       | Thailand     | CHULA           |
| China   | IMP       | France  | LP2iB      | Thailand     | NARIT           |
| China   | JNU       | France  | Subatech   | Thailand     | SUT             |
| China   | JXNU      | Germany | EKUT       | U.K.         | U. Liverpool    |
| China   | KNRC      | Germany | GSI        | U.K.         | U. Warwick      |
| China   | Nankai U. | Germany | U. Hamburg | USA          | UCI             |
| China   | NCEPU     | Germany | JGUMainz   | USA          | UMD-G           |
| China   | NJU       | Germany | RWTH-AC    |              |                 |
| China   | RNCG      | Germany | TUM        |              |                 |



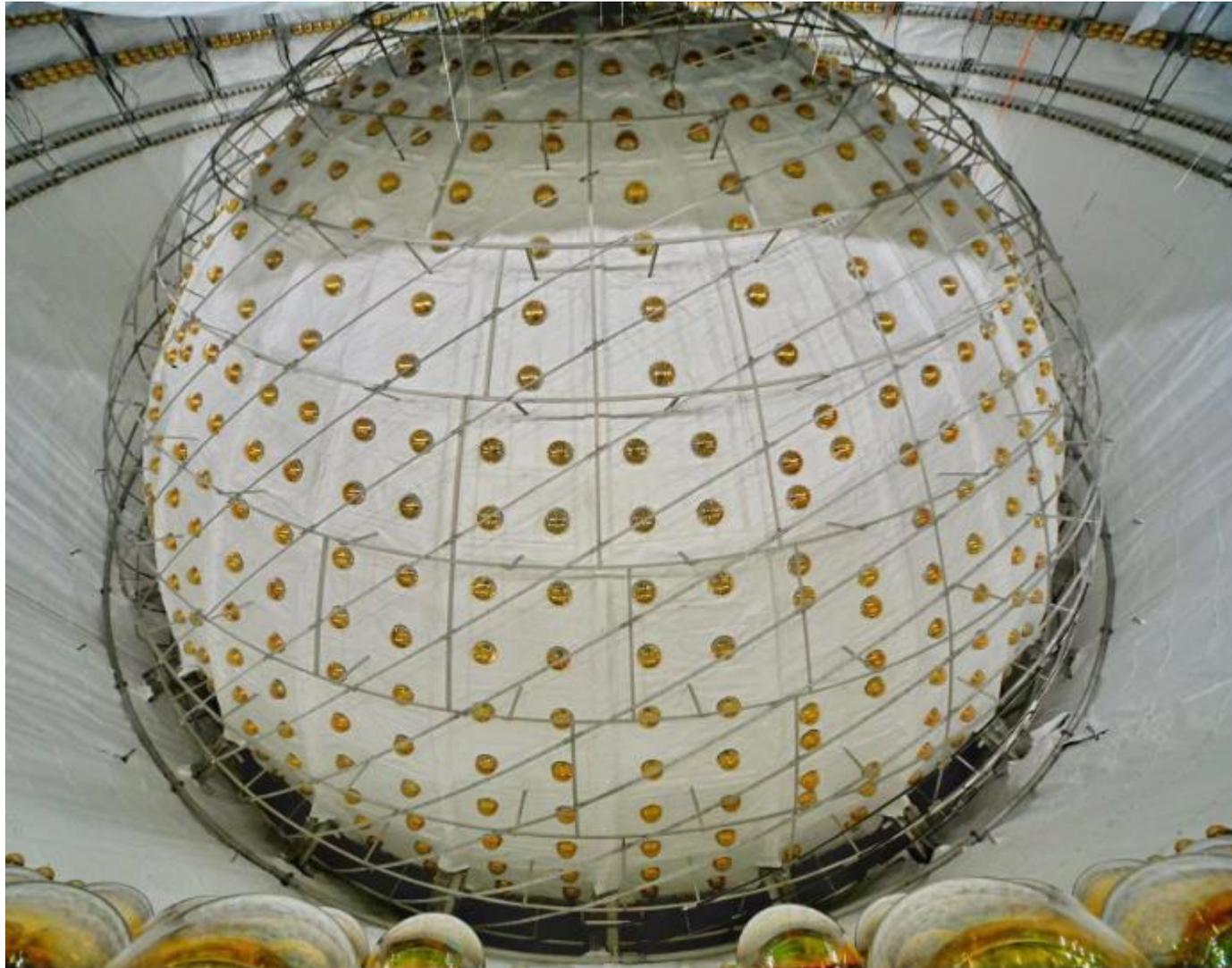
- **Stainless Steel structure**
  - Finished in June 2022
- **Acrylic sphere with 35.4m diameter**
  - Done in Oct. 2024
- **20,012 20" PMTs + 25,600 3" PMTs**
  - Installation finished in Dec. 2024
  - PMTs + electronics working well
- **20kt Liquid scintillator**
  - Filling finished on Aug. 22, 2025

# Group photo on Dec. 4, 2024



# More detector photos

**2024.12: detector assembled, filling start soon**



**2025.6 Top tracker installation finished**



# Liquid filling

Four purification plants to achieve target radio-purity  $<10^{-15}$  g/g U/Th and 20 m attenuation length



5000 m<sup>3</sup> LAB tank



1) Al<sub>2</sub>O<sub>3</sub> for optical

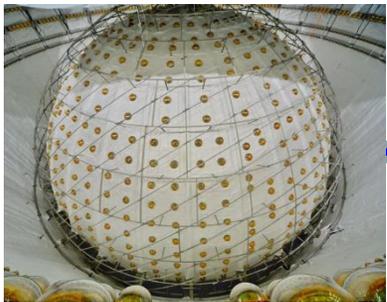


2) Distillation for radiopurity



Mixing PPO and bis-MSB

SS pipes to underground



JUNO detector



CD FOC system



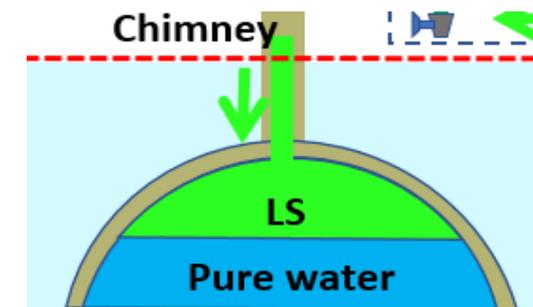
4) Gas stripping to remove Rn and O<sub>2</sub>



3) Water extraction to remove radioactive impurities

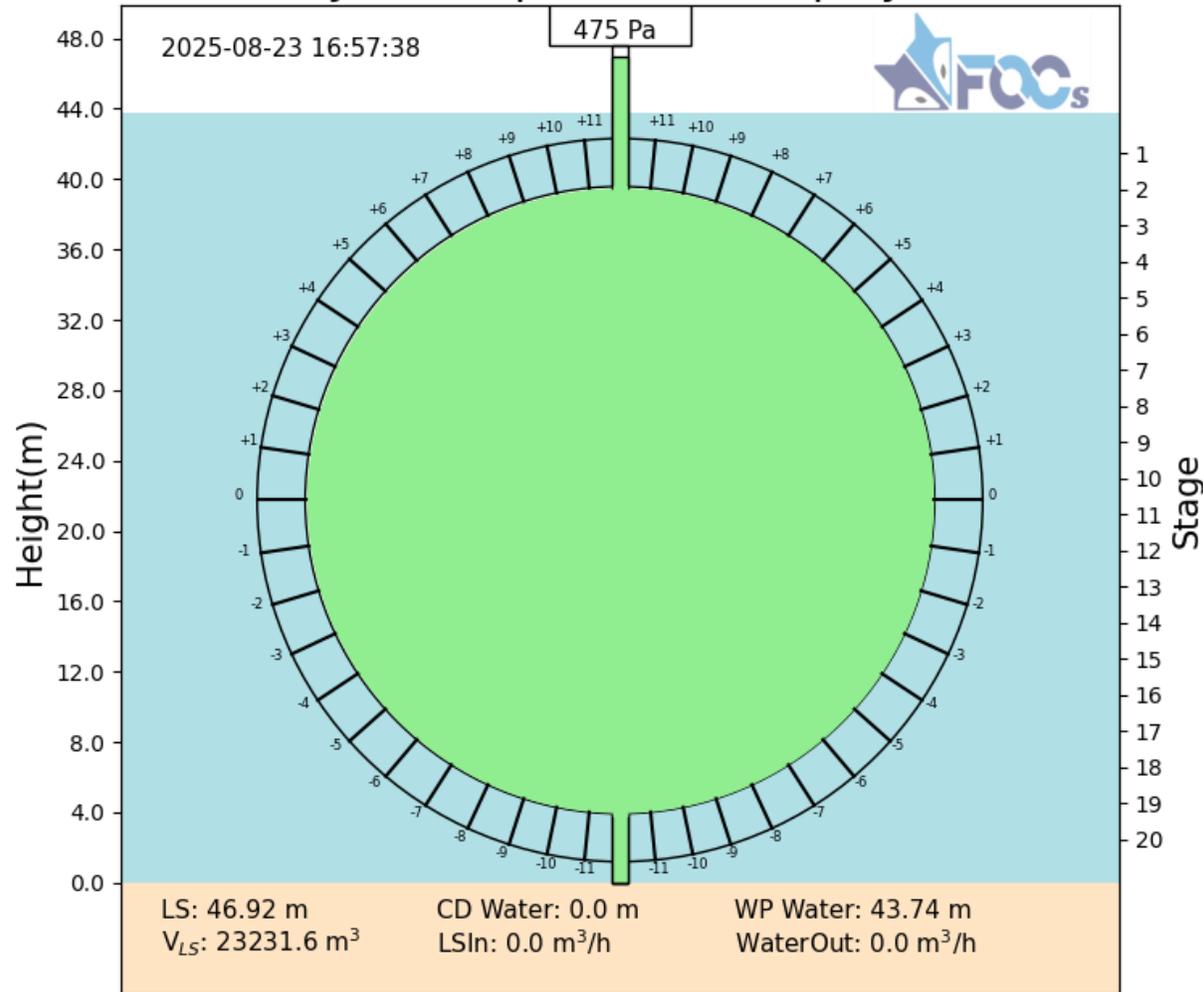
Filling strategies:

- ⇒ Leakage (single component  $< 10^{-6}$  mbar·L/s)
- ⇒ Cleaning vessel before filling
- ⇒ Clean environment
- ⇒ **Water/LS filling**



# Physics data taking started

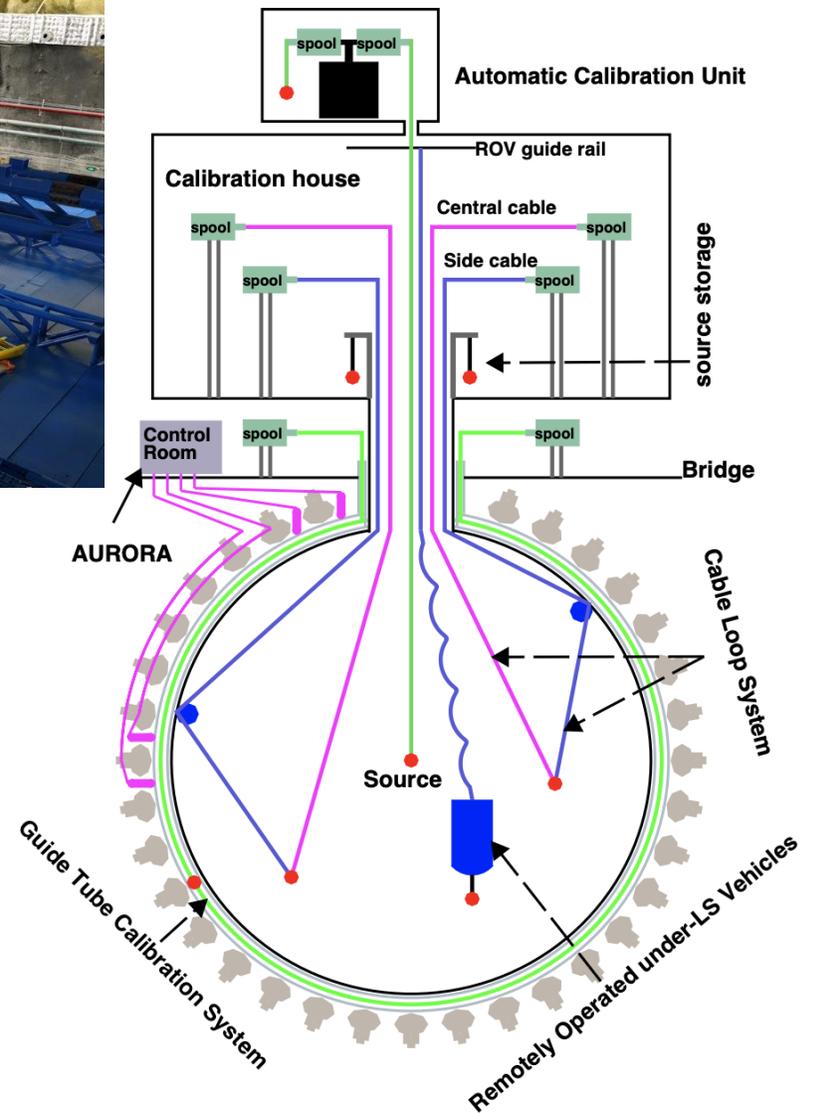
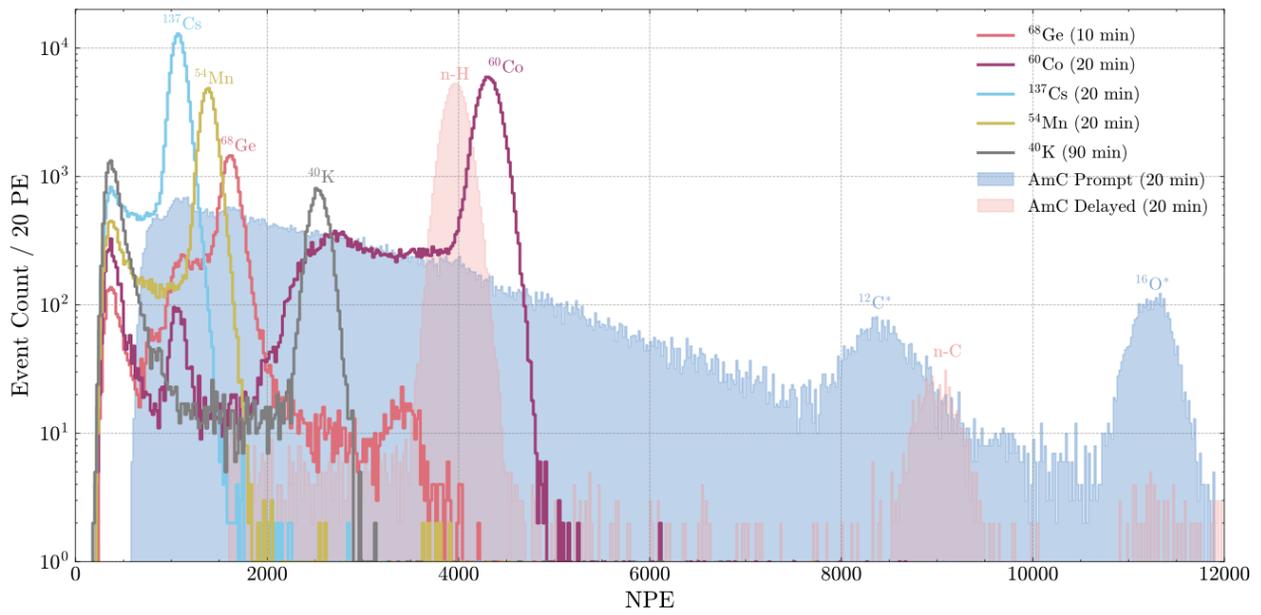
JUNO Liquid Level Display



- Aug. 22, 2025, at ~22:12, water in Acrylic tank was fully replaced by LS
- ~90 t lower quality LS at the bottom were further replaced
- Calibration started at 14:00 of Aug. 23
- Data taking for physics started on Aug. 26 at 5:30 am (Beijing time)
- **Total LS in CD: 23231.6  $\text{m}^3$**
- **Total water in pool: 41225.1  $\text{m}^3$**

# Calibration system

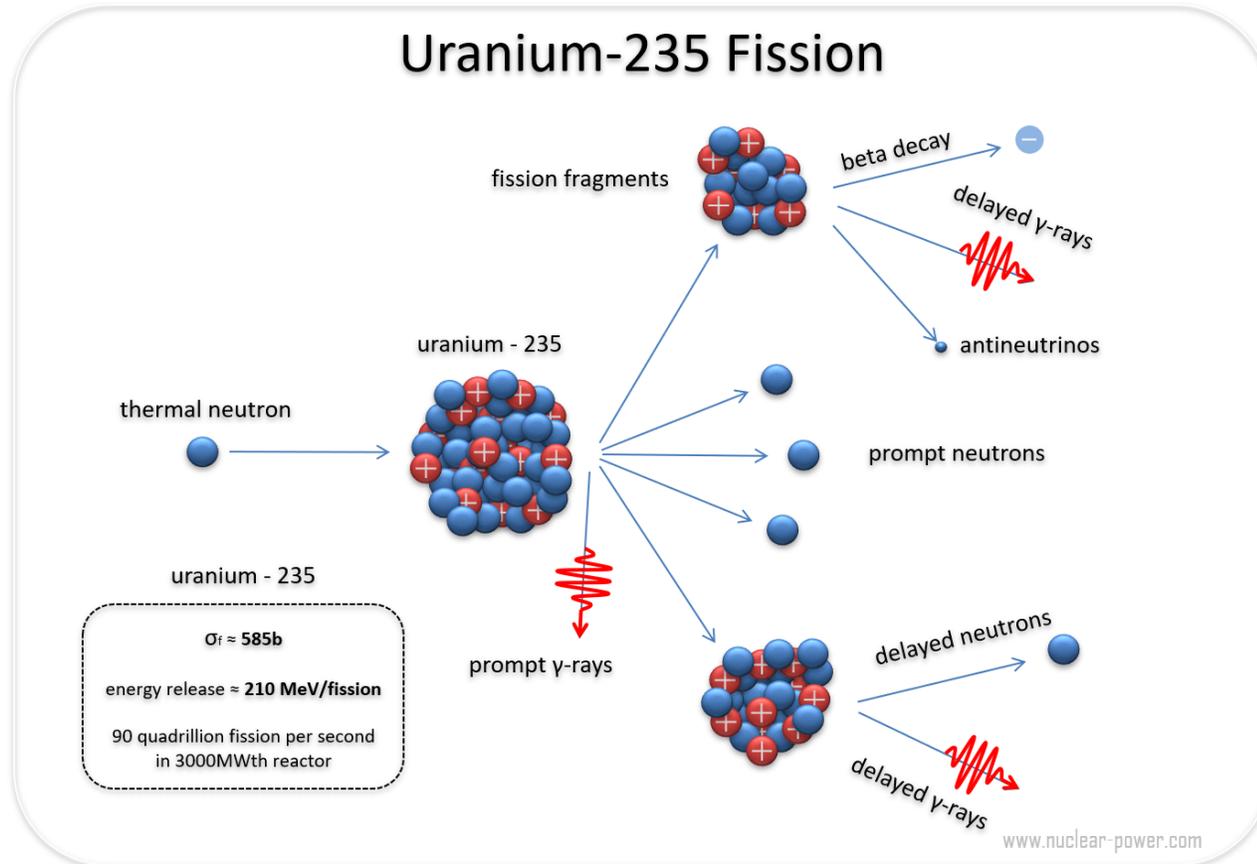
- Precise calibration to understand detector response, using multiple sources, from 1D to 3D
- No Radon leakage during routine and special calibrations



# Reactor neutrinos

## The strongest artificial neutrino source on the Earth

$2 \times 10^{20}$   $\nu$ 's per second per GW thermal power, >99.7% from  $^{235,238}\text{U}$  and  $^{239,241}\text{Pu}$

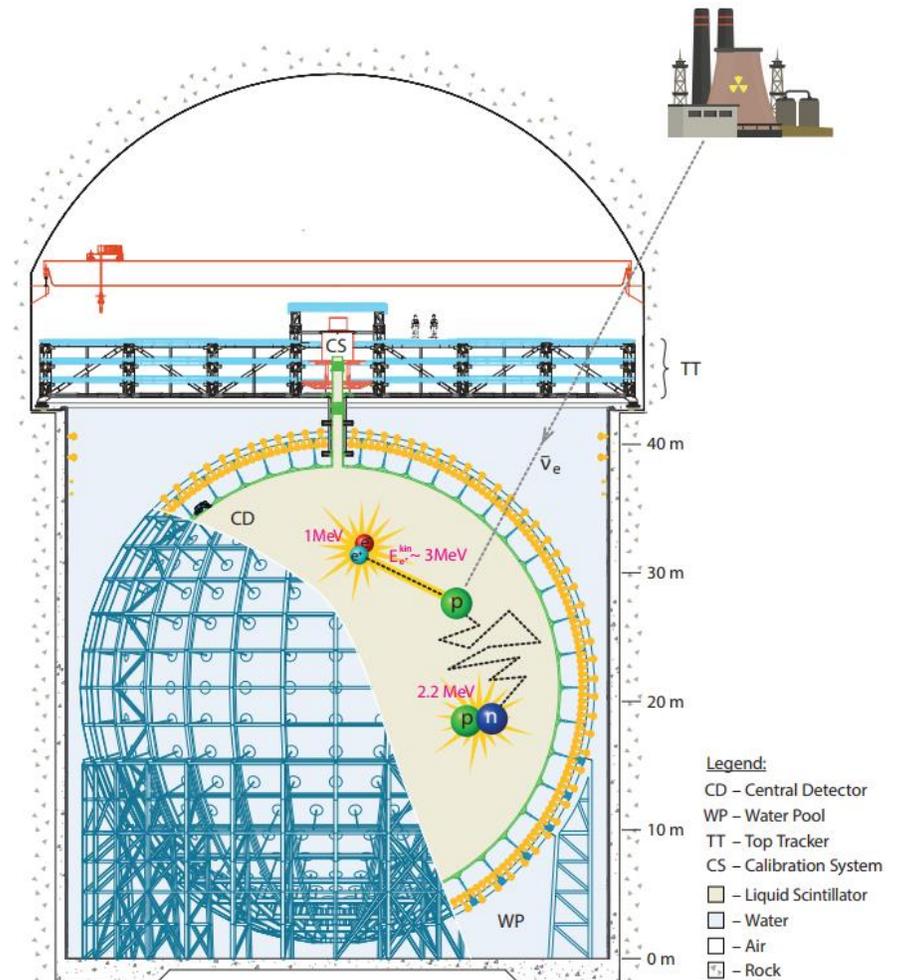
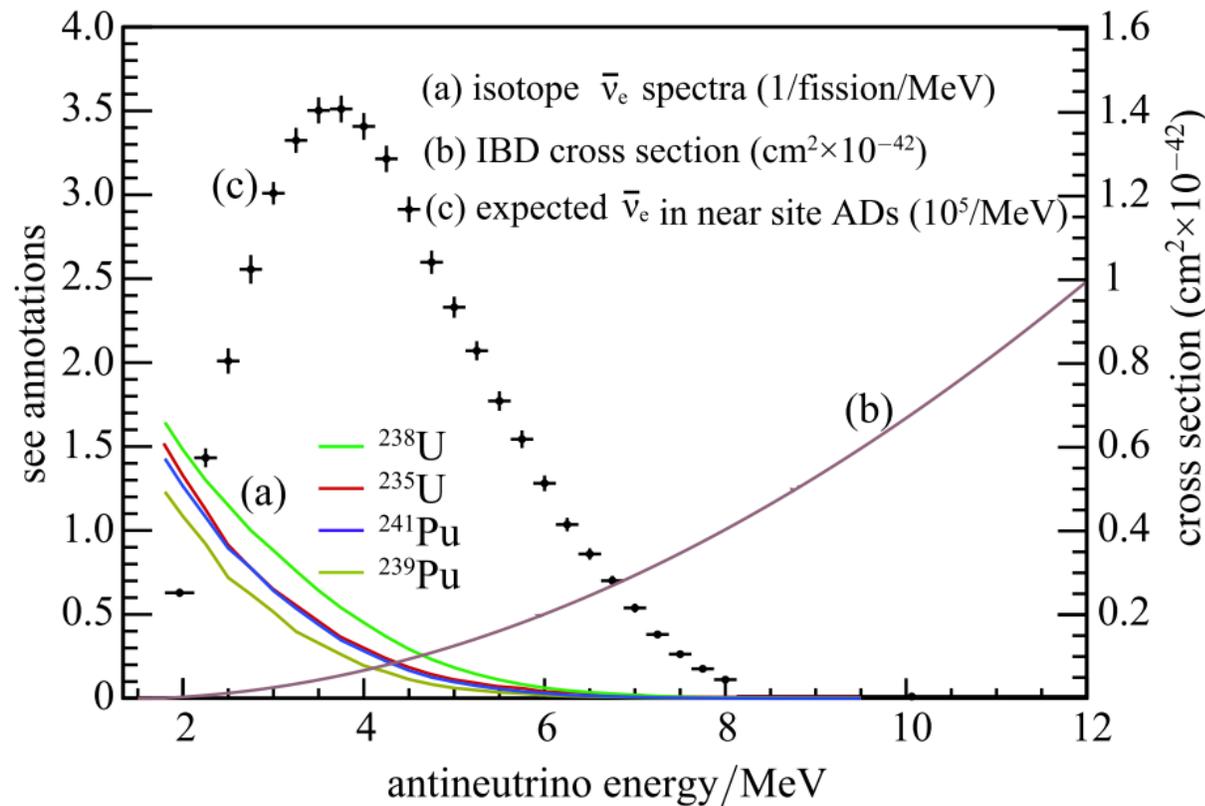


# Reactor neutrinos

The strongest artificial neutrino source on the Earth

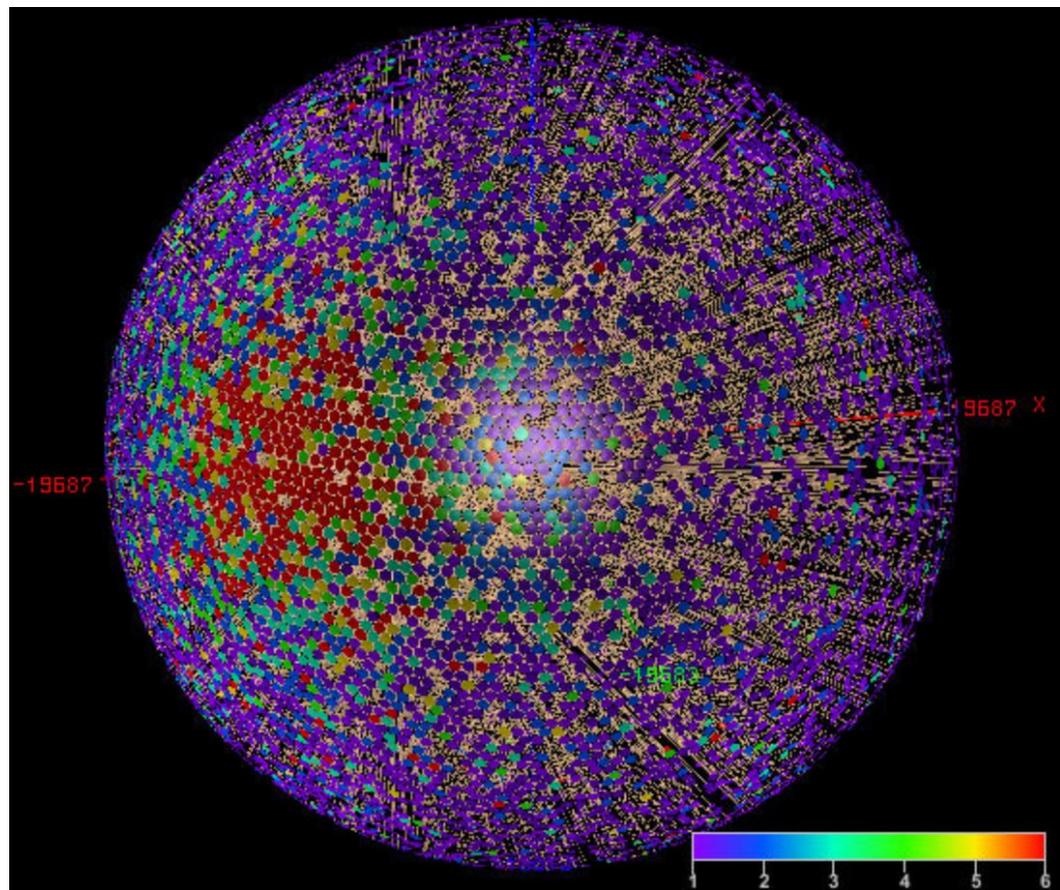
$2 \times 10^{20}$   $\bar{\nu}$ 's per second per GW thermal power, >99.7% from  $^{235,238}\text{U}$  and  $^{239,241}\text{Pu}$

Easy to detect via the Inverse Beta Decays (IBD)



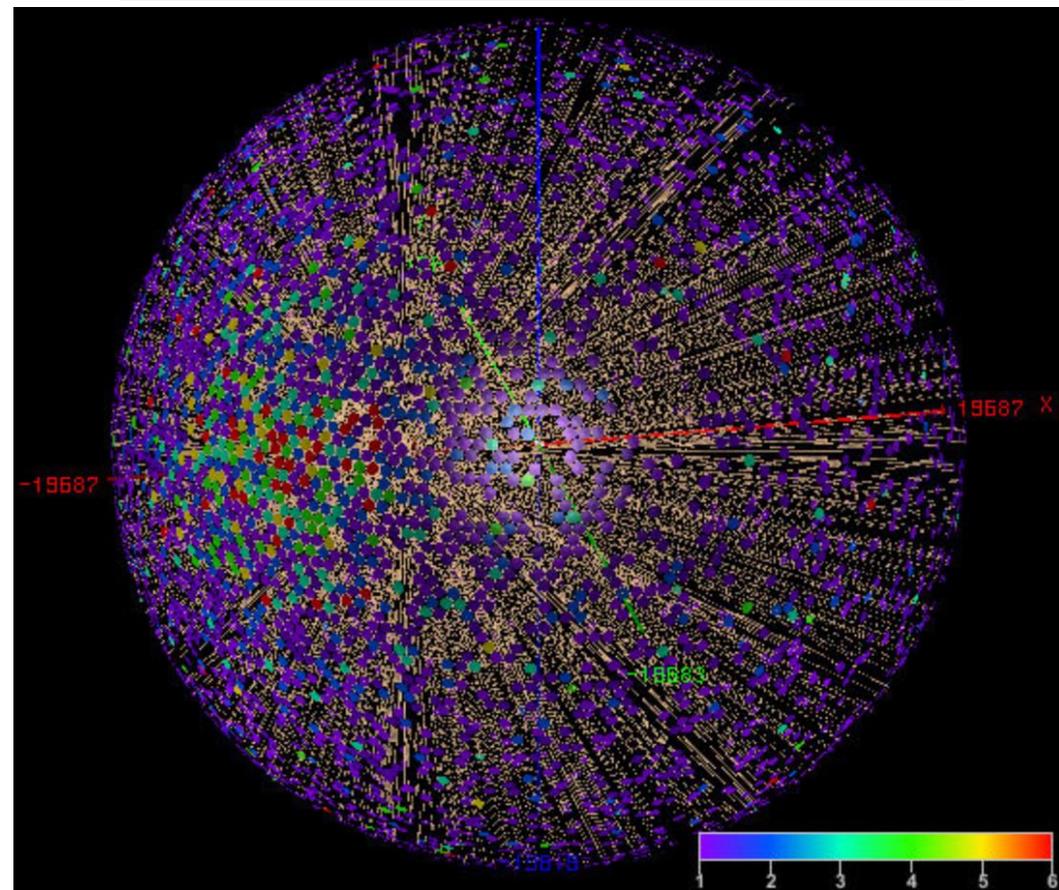
# A Golden Reactor Neutrino Event

Mon, 25 Aug 2025 22:50:45  
RecEnergy = 6.3 MeV  
RecVertex (-9458, -9707, 3820) mm



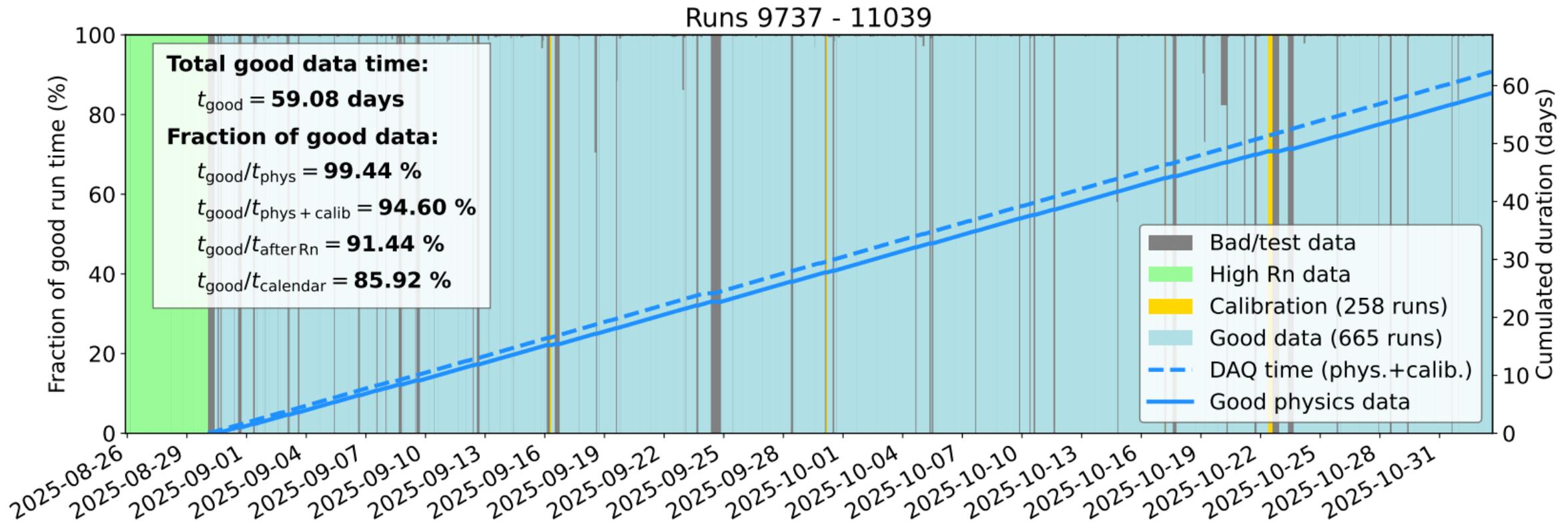
**Prompt e+ signal**

Mon, 25 Aug 2025 22:50:45  
RecEnergy = 2.4 MeV  
RecVertex (-10393, -9794, 4333) mm



**Delay neutron signal**

- **Data taking time larger than 98%**
  - Usually two times of calibration per week
  - Detector maintenance occupied with the calibration source moving
- **Physics data taking time larger than 92%**



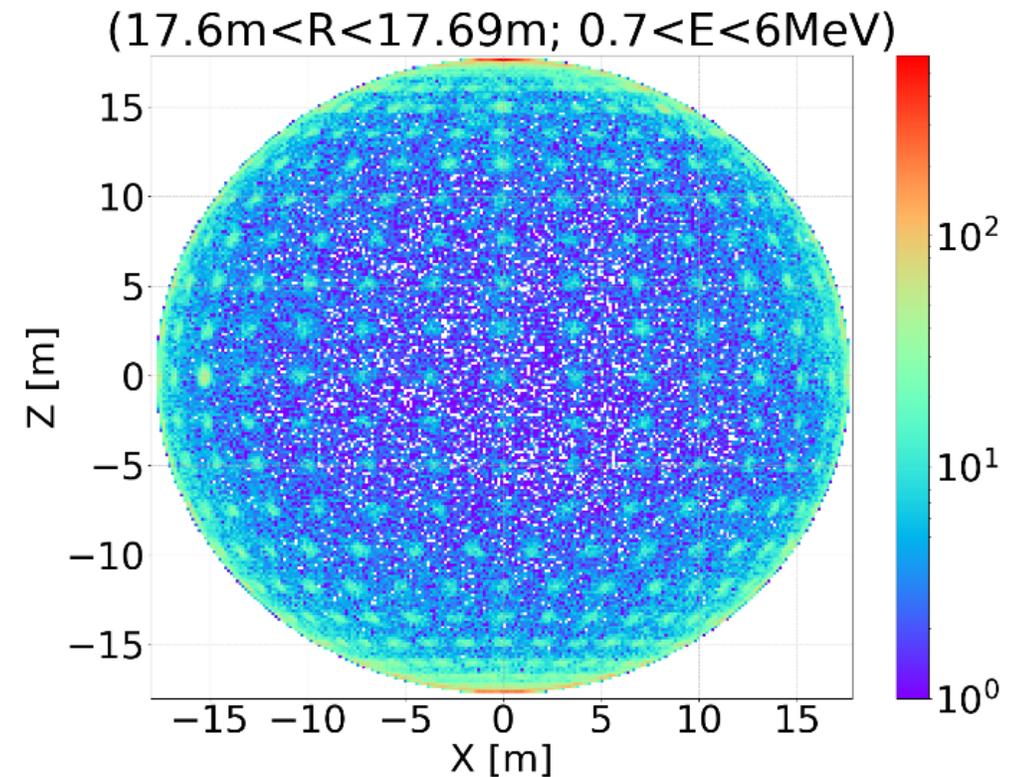
# Keys to Physics: clean detector

## Measured radioactivity in LS and water better than specification

- Radiopurity control of raw material, cleaning of inner surface,  $^{222}\text{Rn}$  control

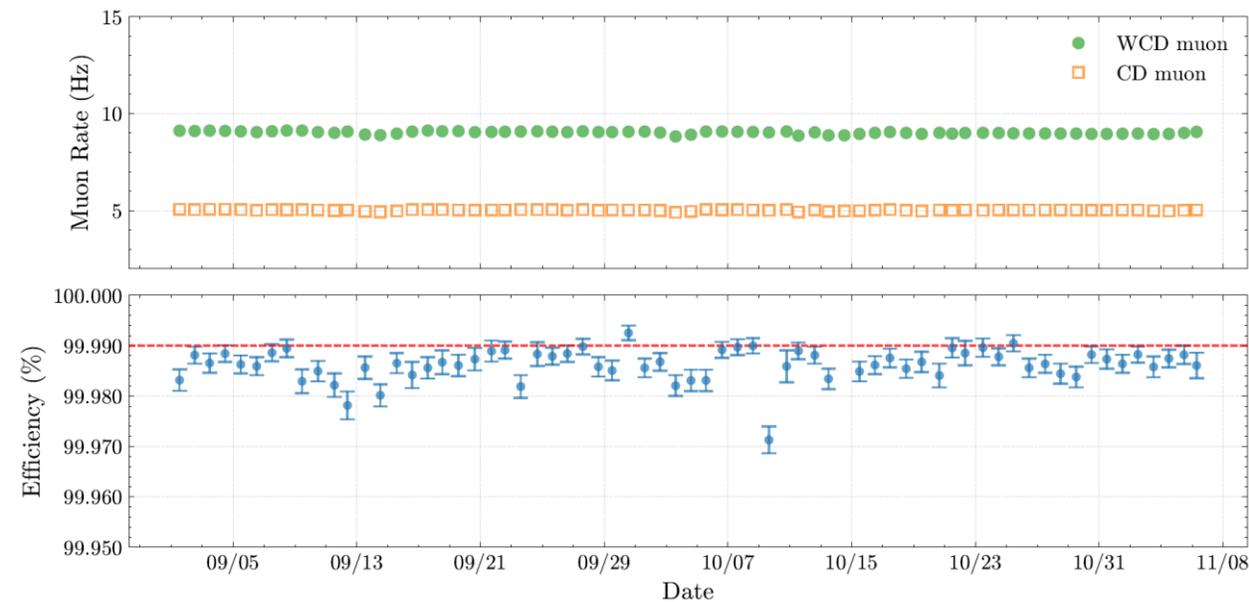
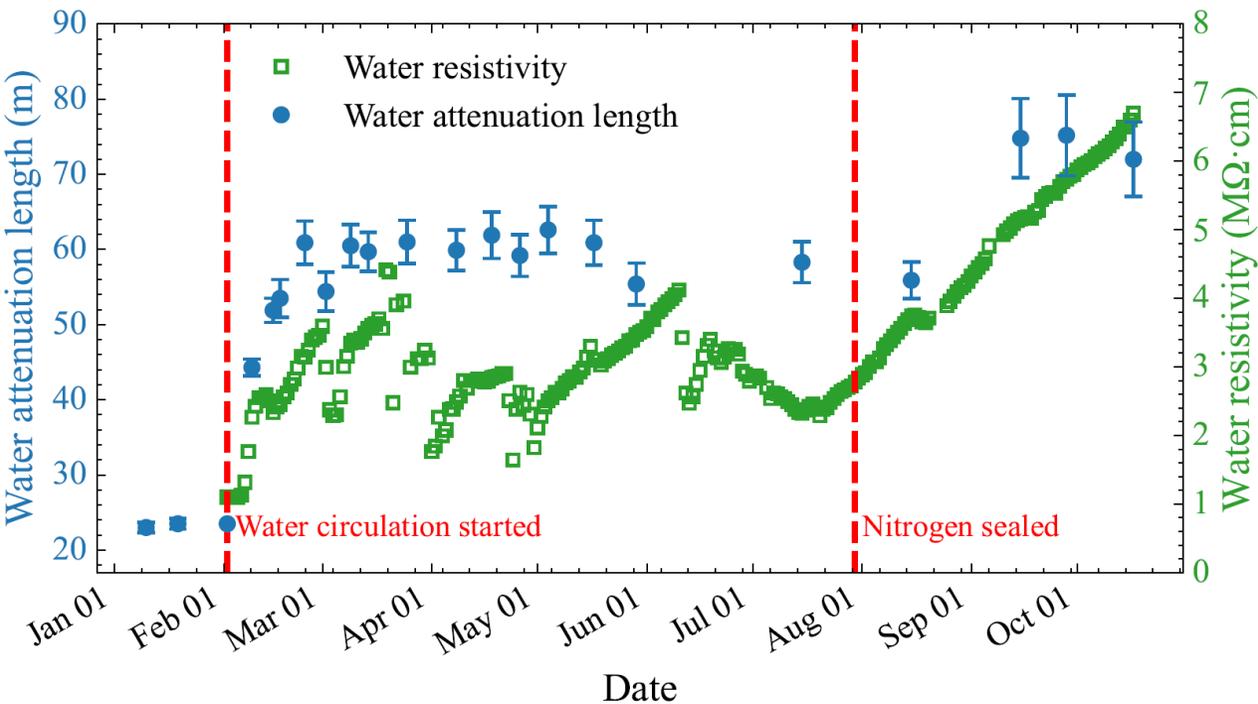
| Central Detector  | Specification | Measure |
|---|---------------|---------|
| Singles Rate (Hz)   | < 7.2 Hz      | < 7 Hz  |
| $^{238}\text{U}/^{232}\text{Th}$ ( $\times 10^{-15}$ g/g) | < 1           | < 0.1   |
| $^{210}\text{Po}$ ( $\times 10^4$ cpd/kt)                 | < 8           | < 5     |

| Muon Veto Detector                      | Specification | Measure |
|---|---------------|---------|
| $^{222}\text{Rn}$ (mBq/m <sup>3</sup> ) | < 10          | <10     |
| $^{226}\text{Ra}$ (mBq/m <sup>3</sup> ) | < 1           | < 0.01  |
| U/Th ( $\times 10^{-15}$ g/g)           | <10           | <0.4    |

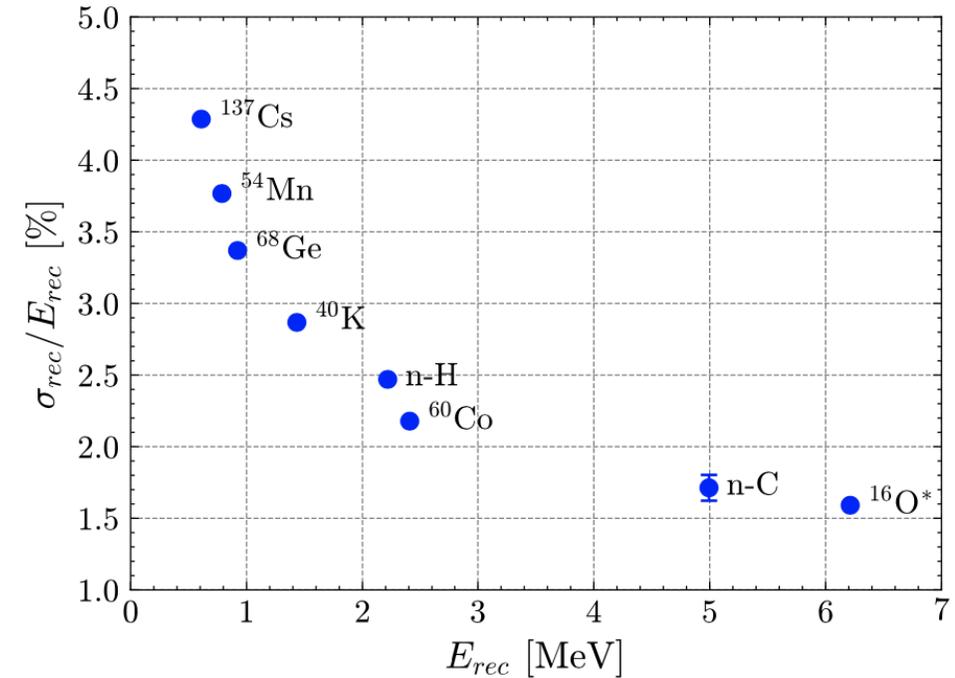
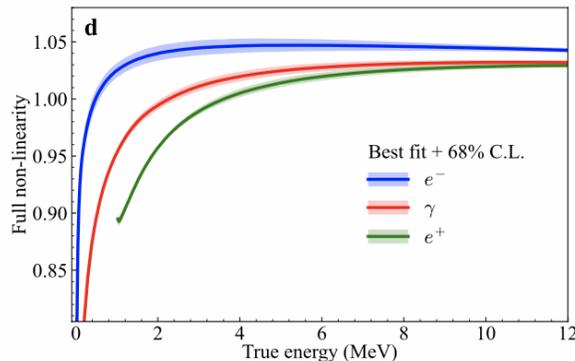
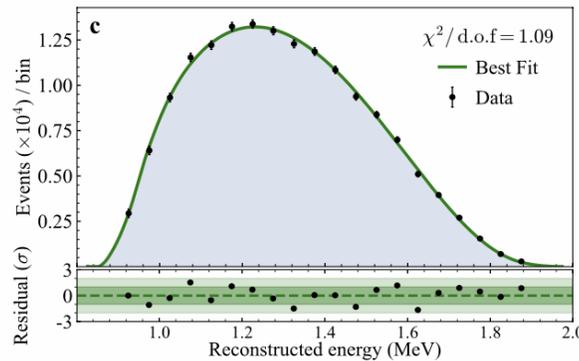
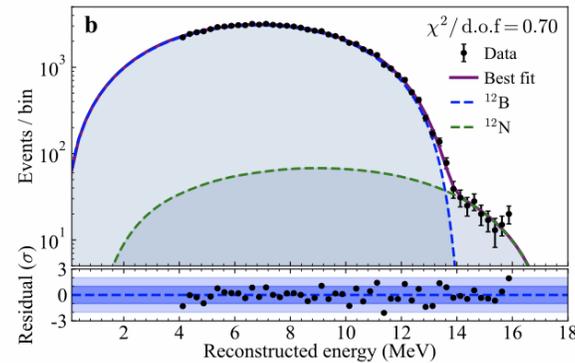
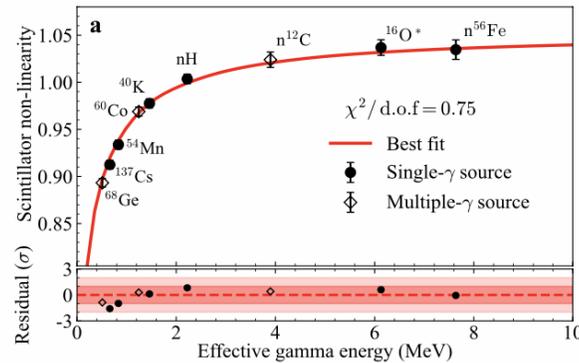


# Keys to Physics: muon tagging

- **Water attenuation length at 405nm wavelength longer than 75m**
- **Muon tagging efficiency of Water Cherenkov detector > 99.95%**
  - Untagged muons enter CD via chimney, Top Tracker detector helps to tag



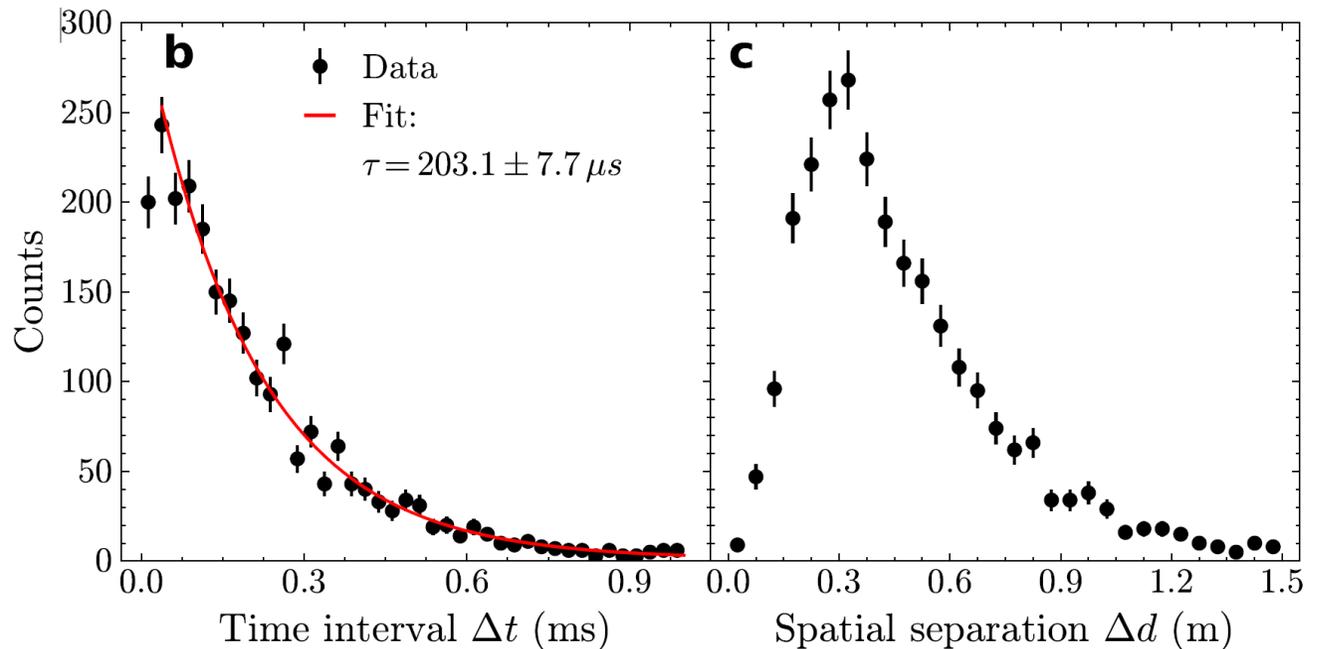
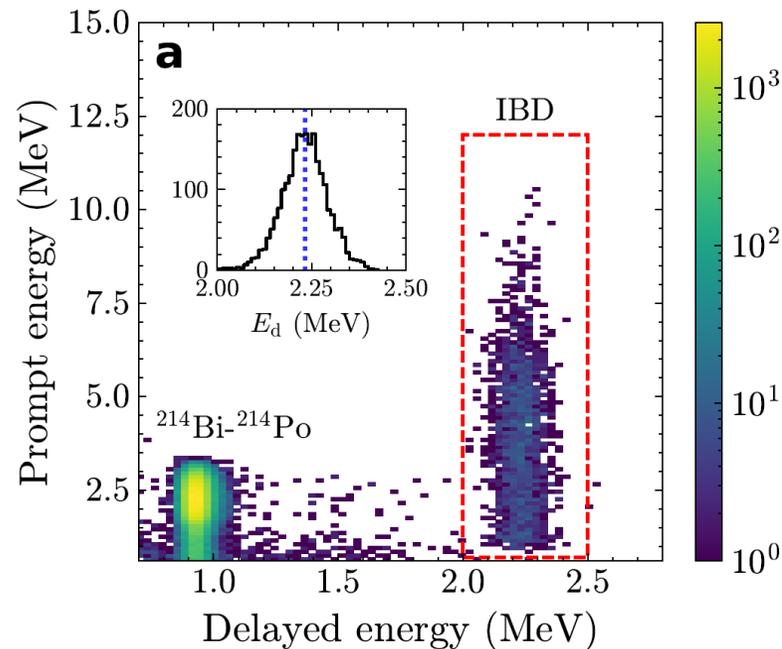
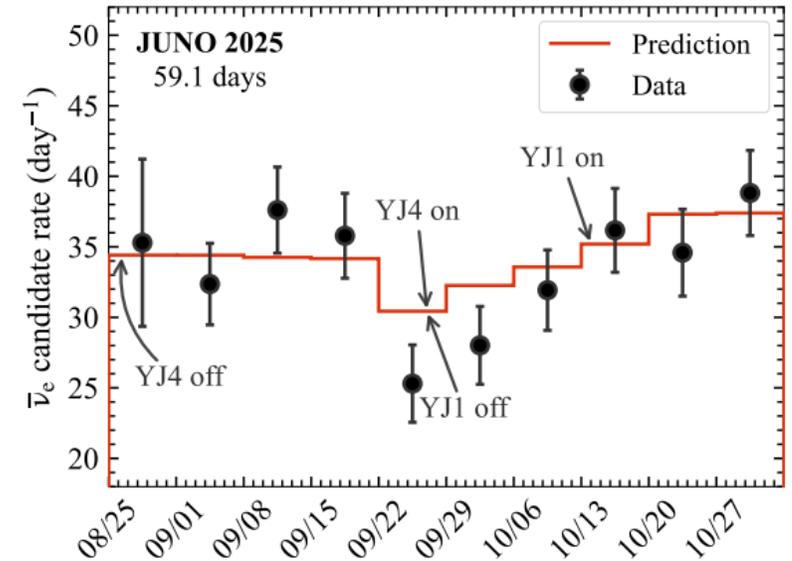
- Energy nonlinearity calibrated to better than 1% precision
- Energy resolution is near the design level
  - Keep improving from comprehensive calibration and better reconstruction algorithms



More details in Mingxia's flash talk

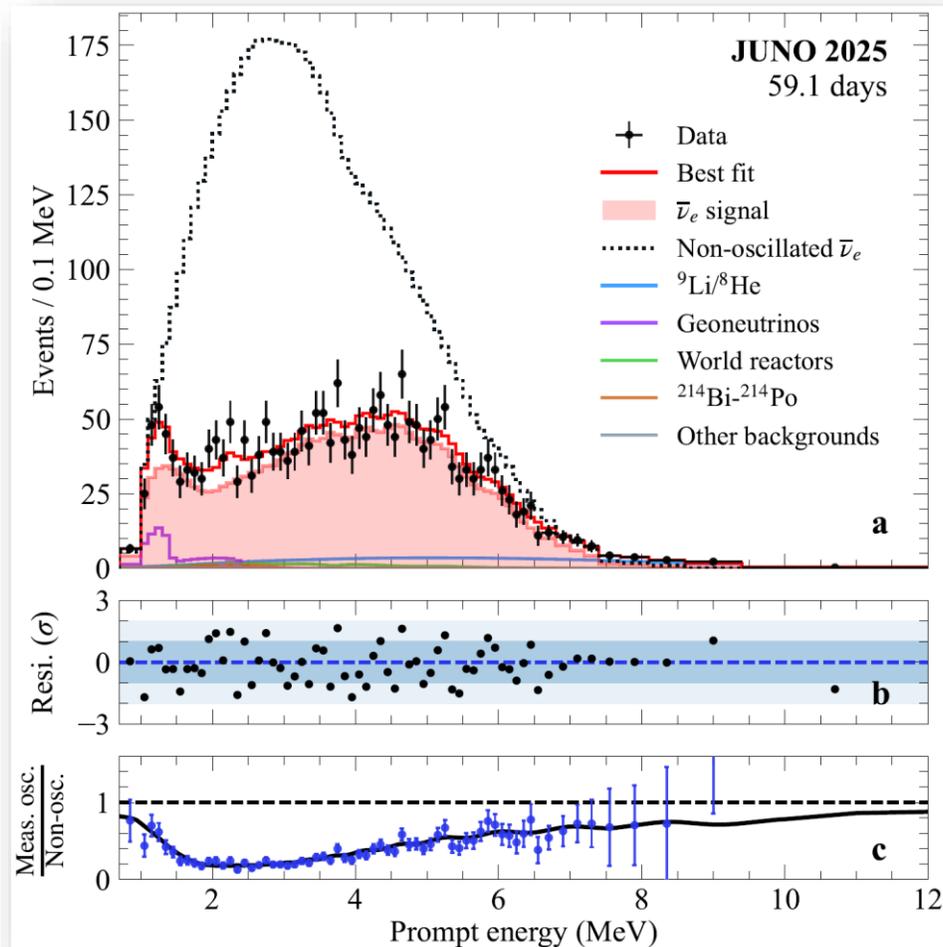
# Reactor neutrino sample

- Reactor neutrinos are selected according to the time, space and energy correlations
- Neutrino rates consistent with power prediction
- More details in Mingyuan's flash talk



# Oscillated spectrum

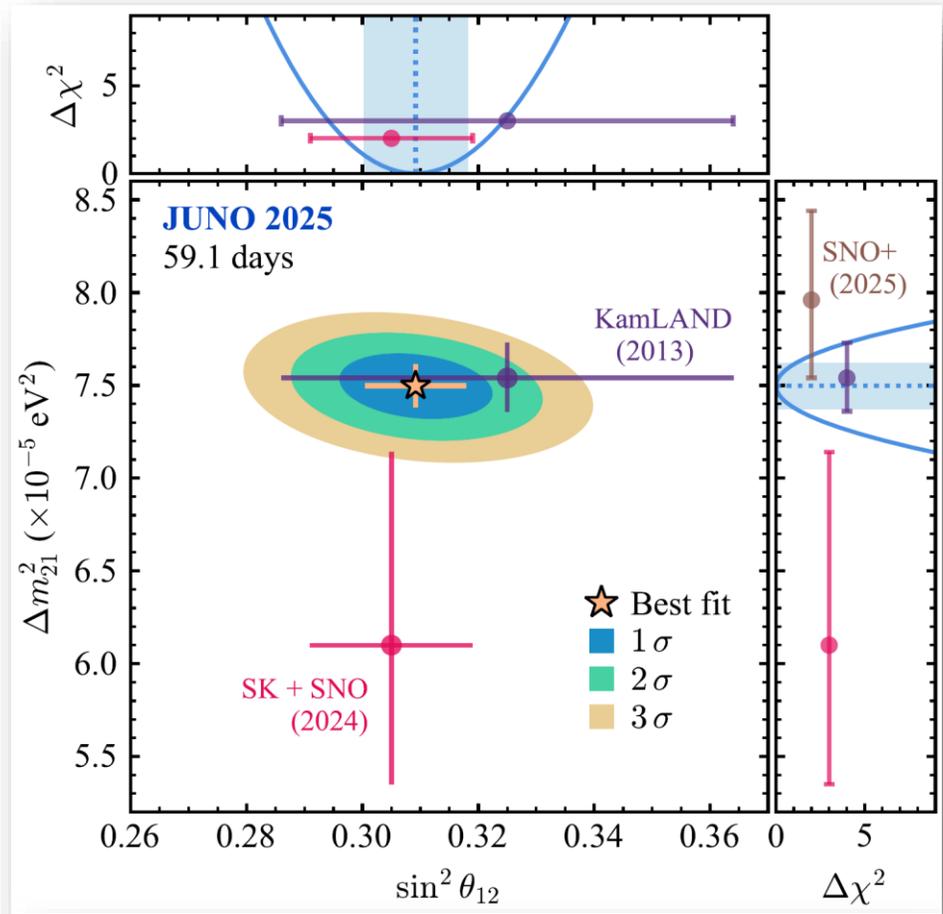
- Residual  $^9\text{Li}$  background will be reduced by veto along muon tracks
- Reasonable good efficiency, will be improved with larger fiducial volume



## Antineutrinos ( $\bar{\nu}_e$ ) Candidates Summary

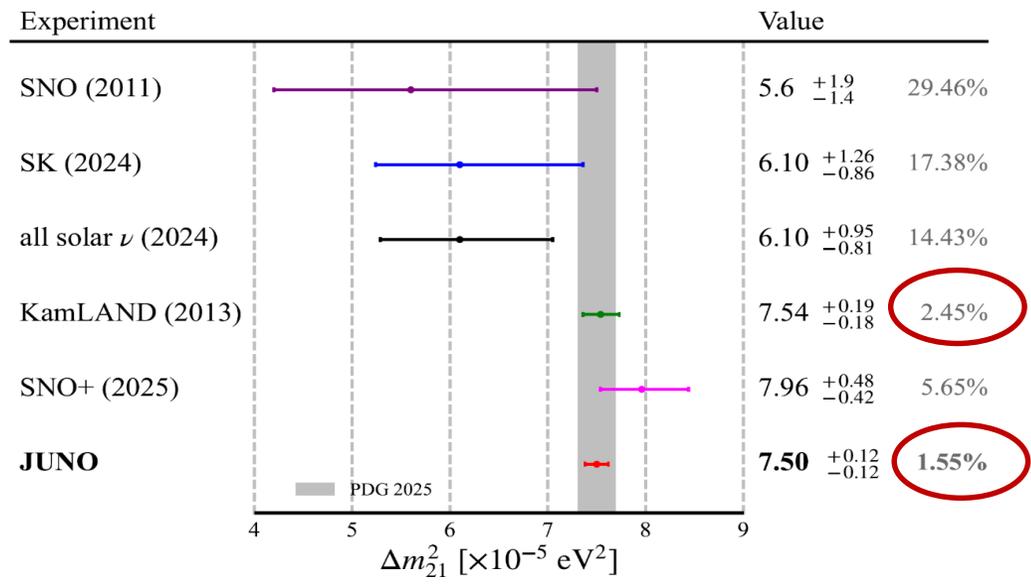
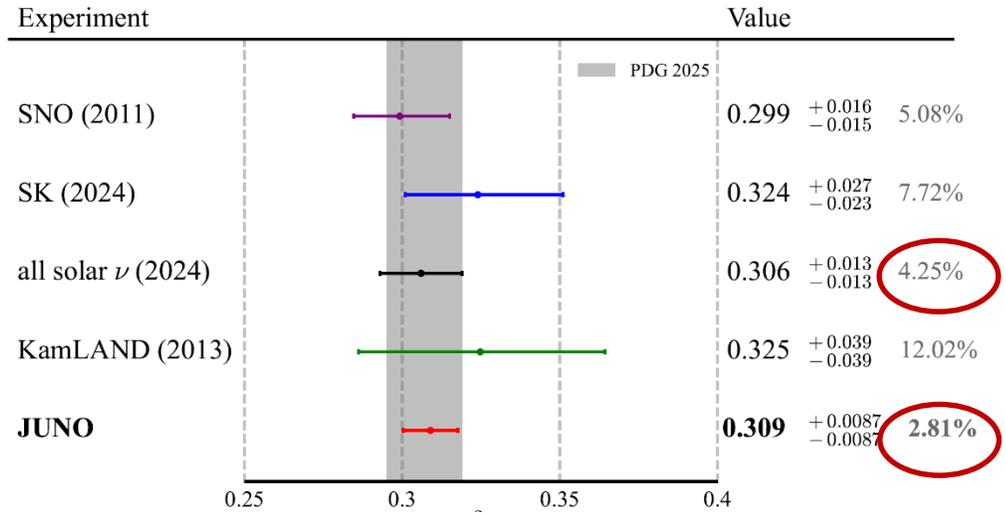
|  |                 |                       |
|--|-----------------|-----------------------|
| DAQ live time (days)                                     | 59.1            |                       |
| $\bar{\nu}_e$ candidates                                 | 2379            |                       |
| <b>Selection Efficiencies (%)</b>                        | $\varepsilon$   | $\sigma_{\text{rel}}$ |
| Fiducial volume  | 80.6            | 1.6                   |
| PMT flasher rejection                                    | >99.9           | negligible            |
| $\mu$ veto   | 93.6            | negligible            |
| Multiplicity   | 97.4            | negligible            |
| Prompt-delayed coinc.                                    | 95.1            | 0.13                  |
| Total efficiency ( $\varepsilon_{\text{tot}}$ )          | 69.9            | 1.6                   |
| <b><math>\bar{\nu}_e</math> signal (cpd<sup>1</sup>)</b> |                 |                       |
| w/o $\varepsilon_{\text{tot}}$ corrected                 | $33.5 \pm 1.7$  |                       |
| w/ $\varepsilon_{\text{tot}}$ corrected                  | $47.9 \pm 2.6$  |                       |
| Non-oscillated $\bar{\nu}_e$                             | $150.9 \pm 2.7$ |                       |
| <b>Backgrounds (cpd)</b>                                 | Pre-fit         | Best-fit              |
| $^9\text{Li}/^8\text{He}$                                | $4.3 \pm 1.4$   | $3.9 \pm 0.6$         |
| Geoneutrinos   | $1.2 \pm 0.5$   | $1.4 \pm 0.4$         |
| World reactors   | $0.88 \pm 0.09$ | $0.88 \pm 0.09$       |
| $^{214}\text{Bi}, ^{214}\text{Po}$                       | $0.18 \pm 0.10$ | $0.20 \pm 0.10$       |
| $^{13}\text{C}(\alpha, n)^{16}\text{O}$                  | $0.04 \pm 0.02$ | $0.04 \pm 0.02$       |
| Fast neutrons  | $0.02 \pm 0.02$ | $0.02 \pm 0.02$       |
| Double neutrons  | $0.05 \pm 0.05$ | $0.07 \pm 0.05$       |
| Atmospheric neutrinos                                    | $0.08 \pm 0.04$ | $0.07 \pm 0.04$       |
| Accidentals ( $\times 10^{-2}$ )                         | $4.9 \pm 0.3$   | $4.9 \pm 0.3$         |

# First physics result



$$\sin^2 \theta_{12} = 0.3092 \pm 0.0087,$$

$$\Delta m_{21}^2 = (7.50 \pm 0.12) \times 10^{-5} \text{ eV}^2$$



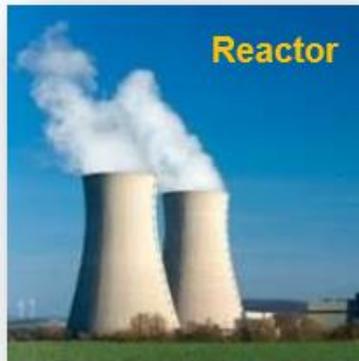
Consistent results from three independent analyses

[arXiv: 2511.14593](https://arxiv.org/abs/2511.14593)

# Stay tuned!

**Neutrino oscillations**, mass ordering ( $3-4 \sigma$ ), precision measurement ( $<0.5\%$ )

**Astro-particle physics**, supernovae neutrinos, geo and solar neutrinos



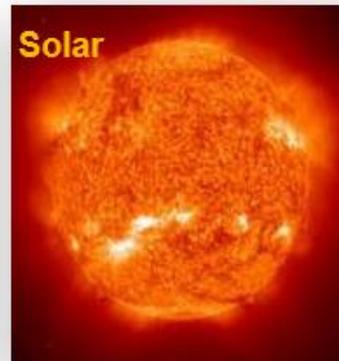
Reactor

~60 IBDs per day



Atmosphere

Several per day



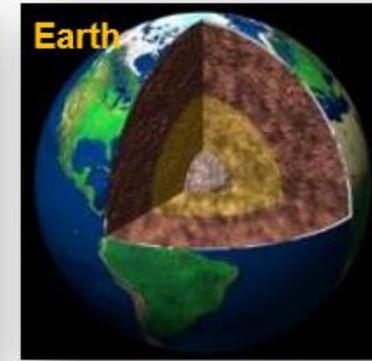
Solar

Hundreds per day



Supernova

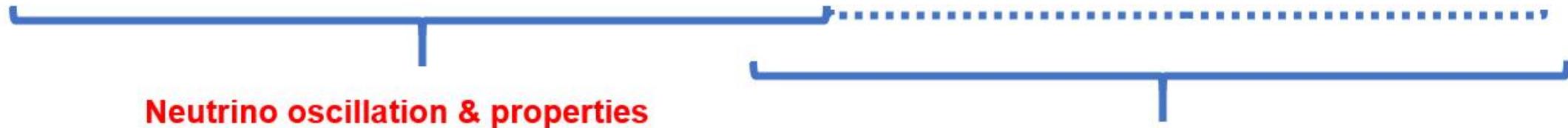
~5000 IBDs for  
CCSN @10 kpc



Earth

Several IBDs per  
day

+  
New  
physics



**Neutrino oscillation & properties**

**Neutrinos as a probe**

IBD: inverse beta decay  $\bar{\nu}_e + p \rightarrow e^+ + n$

CCSN: core-collapse supernova

Thanks!

