

# EW Input Schemes and NLO Corrections in POWHEG Z<sub>ew</sub>-BMNNPV

Practical study for measuring  $\sin^2\theta_W$  at High Mass

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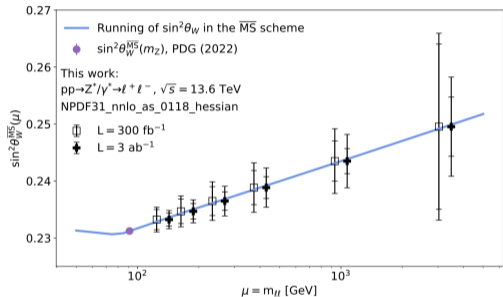
in collaboration with M. Boonekamp, E. Chapon, with input from A. Vicini

IRN Terascale

# Motivation: $\sin^2\theta_W$ Running at High Mass

- Neutral-current Drell-Yan ( $pp \rightarrow Z/\gamma^* \rightarrow \ell^+\ell^-$ ): pillar of precision EW physics at the LHC
- $\sin^2\theta_W$ : fundamental parameter, sensitive to BSM
- The SM predicts a **running** of  $\sin^2\theta_W^{\overline{\text{MS}}}(\mu)$ :
  - $\sim 0.2315$  at  $M_Z$
  - $\sim 0.25$  at  $\sim 3$  TeV
- Weak corrections grow from  $\sim 0.1\%$  at  $M_Z$  to  $\mathcal{O}(10\%)$  at TeV scale

The choice of EW input scheme directly impacts the size of these corrections and the precision of predictions



from arXiv:2302.10782

Running of  $\sin^2\theta_W^{\overline{\text{MS}}}(\mu)$  from  $M_Z$  to multi-TeV

# EW Input Parameter Schemes

## Schemes implemented in POWHEG Z<sub>ew</sub>-BMNPV

Scheme	Input parameters	Derived	POWHEG flags
$G_\mu$ scheme ( $M_W$ )	$G_\mu, M_W, M_Z$	$\sin^2\theta_W, \alpha$	scheme=2
$G_\mu$ scheme ( $\sin^2\theta_{\text{eff}}^\ell$ )	$G_\mu, \sin^2\theta_{\text{eff}}^\ell, M_Z$	$M_W, \alpha$	scheme=2 + use-s2effin
$\alpha_0$ scheme	$\alpha_0, G_\mu, M_Z$	$M_W, \sin^2\theta_W$	scheme=4 + azinscheme4=1

- $\overline{\text{MS}}$  scheme (arXiv:2302.10782):  
running  $\alpha_{\overline{\text{MS}}}(\mu), \sin^2\theta_W^{\overline{\text{MS}}}(\mu)$  as inputs
- $\sin^2\theta_{\text{eff}}^\ell$  as input (arXiv:1906.11569):  
enables **direct extraction** from  $A_{\text{FB}}$

### Key point

All schemes are **formally equivalent** at a given perturbative order.  
Numerical differences arise from **truncation** of the perturbative series.

Ref: Chiesa et al. 2024 (arXiv:2402.14659), Chiesa et al. 2019 (arXiv:1906.11569), Barzè et al. 2013 (arXiv:1302.4606)

# Interplay: Scheme Choice $\longleftrightarrow$ Theory Accuracy

## 1. Perturbative convergence

- $G_\mu / \alpha(M_Z^2)$  schemes:  
absorb  $\alpha$  running into LO  
 $\Rightarrow$  **smaller NLO corrections**  
 $\Rightarrow$  better convergence
- $\alpha_0$  scheme:  
NLO contains  $\log(m_f^2/Q^2)$   
 $\Rightarrow$  larger corrections
- **Spread between schemes** at a given order = estimate of missing higher orders

## 2. Experimental uncertainties

- $\alpha_0, G_\mu, M_Z$ :  
known with **excellent precision**  
 $\Rightarrow$  unc. negligible
- $M_W$ :  $\Delta M_W = \pm 10$  MeV  
 $\Rightarrow$  **non-negligible**

**Trade-off:**  $(\alpha_0, G_\mu, M_Z)$  has negligible parametric unc. but worst perturbative convergence

## 3. Free parameter for fits

- To extract  $\sin^2\theta_{\text{eff}}^\ell$  from  $A_{\text{FB}}$ :  
need  $\sin^2\theta_{\text{eff}}^\ell$  as input  
(Chiesa et al. 2019)
- To measure  $M_W$ :  
need  $M_W$  as free parameter
- Scheme must be **consistent with the target observable** of the measurement

## POWHEG-BOX-V2

MC event generator that computes the **hard scattering** at NLO accuracy (QCD and/or EW).

Key idea: NLO precision for the first (hardest) emission, then hands over to the parton shower

Module  $Z_{ew}$ -BMNNPV:  
adds **NLO EW corrections** to  
 $qq \rightarrow Z/\gamma^* \rightarrow \ell^+ \ell^-$

Barzè et al. 2013, Nason 2004

## Pythia8

### Parton shower + hadronisation

Takes POWHEG events and adds:

- Additional QCD radiation (softer gluons/quarks)
- Hadronisation: quarks  $\rightarrow$  hadrons
- Underlying event (multi-parton interactions)

## PHOTOS

### QED photon radiation

Adds photon emission off final-state leptons ( $\ell \rightarrow \ell + \gamma$ ):

- Soft photons resummed to all orders
- Multiple photon emission

# POWHEG Configuration: Key Flags

## EW content control

Flag	Role
scheme	EW renormalization scheme (0, 2, 4)
no_ew	Disable all EW corrections (1=off, 0=on)
weak-only	1: weak only at matrix-element -1: full EW (QED+weak)
ew_ho	Higher-order corrections ( $\Delta\alpha$ , $\Delta\rho$ )
use_s2thetaeff	Override $\sin^2\theta_{\text{eff}}^\ell$ as input

## Critical: weak-only flag

When using PHOTOS for QED FSR:

- weak-only = 1: weak corrections at matrix-element, QED via PHOTOS ✓
- weak-only = -1: QED+weak at matrix-element + PHOTOS  $\Rightarrow$  **QED double-counting!**

## Validation

POWHEG Z\_ew-BMNNPV validated against **HORACE** (independent EW generator with completely independent EW form factors)

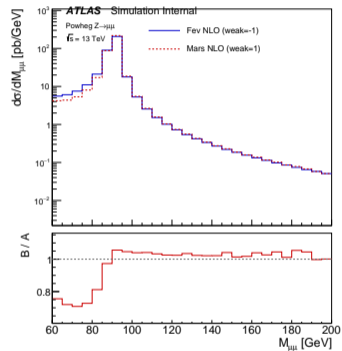
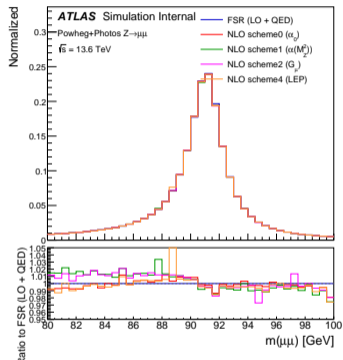
Barzè et al. 2013 (arXiv:1302.4606)

## Scan axes

- **3 EW levels:** Born  $\rightarrow$  NLO  $\rightarrow$  NLO+HO
- **3 schemes:**  $(G_\mu, M_W, M_Z) / (G_\mu, \sin^2\theta_{\text{eff}}, M_Z) / (\alpha_0, G_\mu, M_Z)$
- **Variations:**  $\Delta \sin^2\theta_{\text{eff}} = \pm 0.01$ , PHOTOS on/off
- $\sim 25\text{M}$  events per configuration

Config	scheme	no_ew	weak-only	ew_ho	s2eff
LO $(G_\mu, M_W, M_Z)$	2	1	—	0	-1
NLO $(G_\mu, M_W, M_Z)$	2	0	1	0	-1
NLO $(G_\mu, s_{\text{eff}}^2, M_Z)$	2	0	1	0	0.23154
NLO $(\alpha_0, G_\mu, M_Z)$	4	0	1	0	-1
NLO+HO (same 3 schemes)			<i>idem with ew_ho=1</i>		

# Impact of EW Corrections on $d\sigma/dM_{\ell\ell}$



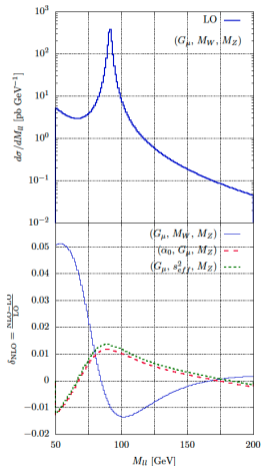
## Scheme comparison at $Z$ peak

• **Left:** EW corrections modify the  $Z$  peak shape — schemes give different corrections

• **Right:** Full EW vs weak-only — quantifies the QED component, the difference between the two being what we want to avoid double-counting when using PHOTOS !

## Full EW vs weak-only

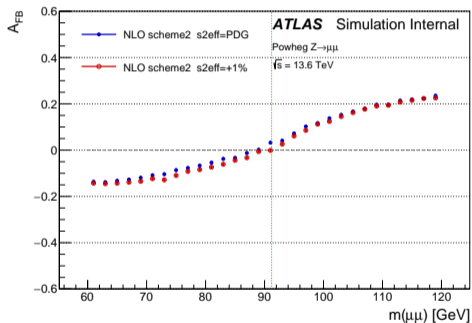
# NLO EW Corrections: $d\sigma/dM_{\ell\ell}$ — Scheme Dependence



## Key observations

- Corrections  $\sim$  few % at  $Z$  peak
- In  $G_\mu$  scheme, tree-level  $\sin^2\theta_W$  differs from measured  $s_{\text{eff}}^2 \Rightarrow$  compensation by NLO counterterms  $\Rightarrow$  visible %-level corrections
- In the two other schemes, LO  $\sin^2\theta_W$  is already close to  $s_{\text{eff}}^2 \Rightarrow$  smaller NLO corrections
- Spread between schemes is an estimator of the missing higher orders  $\Rightarrow$  built-in uncertainty estimate for theoretical predictions

$A_{FB} = \frac{\sigma_F - \sigma_B}{\sigma_F + \sigma_B}$ , with  $\sigma_F$  ( $\sigma_B$ ) the forward (backward) cross-sections in  $m_{\ell\ell}$ .



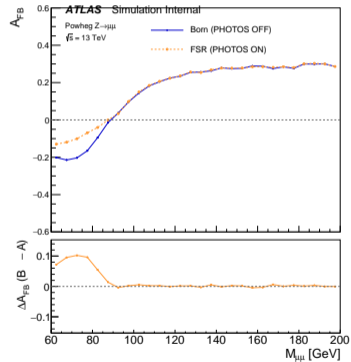
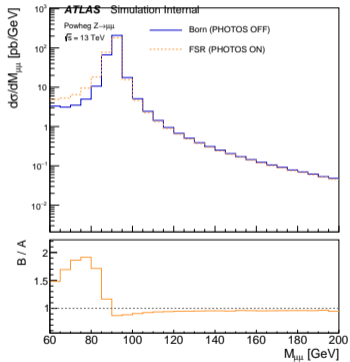
weak-only=1

## Key observations

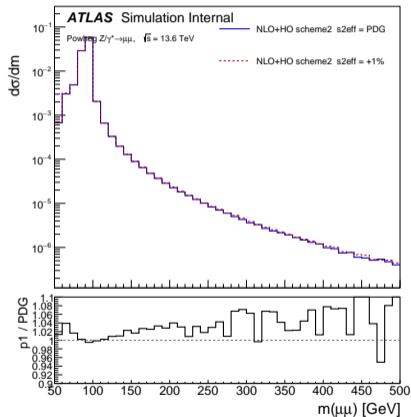
- $A_{FB}$  at  $Z$  peak: directly sensitive to  $\sin^2\theta_{\text{eff}}^\ell$
- Observable used to extract  $\sin^2\theta_{\text{eff}}^\ell$  at the peak

# Effect of PHOTOS (QED FSR)

Born without PHOTOS (OFF) vs with PHOTOS (ON). Only PHOTOS changes.



# From NLO EW Corrections to $\sin^2\theta_W$ Running



We saw that NLO EW corrections **grow with  $M_{\ell\ell}$**  and **depend on the scheme**.

This means high-mass observables are sensitive to EW parameters — including  $\sin^2\theta_{\text{eff}}^\ell$ :

- $\Delta \sin^2\theta_{\text{eff}} = 0.01$ :  $\sim 5\%$  shape effect at 110–500 GeV, growing at higher masses

## Connection to the running

The SM predicts that  $\sin^2\theta_W^{\overline{\text{MS}}}$  **evolves with the invariant mass**.

⇒ Measuring  $d\sigma/dM_{\ell\ell}$  in mass bins, with NLO EW corrections properly accounted for, directly probes  $\sin^2\theta_W(Q^2)$ .

Our POWHEG setup provides the tool to do this.

$$s_{\text{eff}}^2: 0.23148 \text{ vs } 0.24148$$

- 1 NLO EW corrections to NC DY **grow with**  $M_{\ell\ell}$  — from  $\sim 0.1\%$  at the  $Z$  peak to  $\mathcal{O}(10\%)$  at the TeV scale — and **depend on the EW input scheme**
- 2 The choice of scheme is guided by **3 criteria**: perturbative convergence, parametric precision, and coherence with the fit observable.  $G_\mu$  schemes offer better convergence than  $\alpha_0$ .
- 3 We provide and test a POWHEG setup:  
weak-only=1 + PHOTOS, 3 schemes  $\times$  3 EW levels.  
Scheme dependence quantified across the full mass range.
- 4 High-mass observables are **sensitive to**  $\sin^2\theta_W$ : this sensitivity, combined with the NLO EW corrections, enables a direct probe of the **scale evolution** of  $\sin^2\theta_W$ .
- 5 **Next**:  $\overline{\text{MS}}$  running implementation + template fits to measure  $\sin^2\theta_W^{\overline{\text{MS}}}(\mu)$  at different scales.

- S. Amoroso, M. Chiesa, C.L. Del Pio, K. Lipka, F. Piccinini, F. Vazzoler, A. Vicini, *“Probing the weak mixing angle at high energies at the LHC and HL-LHC”*, Phys. Lett. B (2023) — arXiv:2302.10782
- M. Chiesa, C.L. Del Pio, F. Piccinini, *“On electroweak corrections to neutral current Drell-Yan with the POWHEG BOX”*, Eur. Phys. J. C 84 (2024) 539 — arXiv:2402.14659
- M. Chiesa, F. Piccinini, A. Vicini, *“On the direct determination of  $\sin^2\theta_{eff}^\ell$  at hadron colliders”*, Phys. Rev. D 100 (2019) 071302 — arXiv:1906.11569
- L. Barzè, G. Montagna, P. Nason, O. Nicrosini, F. Piccinini, A. Vicini, *“Neutral current Drell-Yan with combined QCD and electroweak corrections in the POWHEG BOX”*, Eur. Phys. J. C 73 (2013) 2474 — arXiv:1302.4606

# Backup

# Backup: Full Configuration Table

Config	scheme	no_ew	weak-only	ew_ho	s2eff	bornonly	azin4
LO Born ( $G_\mu, M_W, M_Z$ )	2	1	-1	0	-1	1	—
LO Born ( $G_\mu, s_{\text{eff}}^2, M_Z$ )	2	1	-1	0	0.23154	1	—
LO ( $G_\mu, M_W, M_Z$ )	2	1	-1	0	-1	0	—
LO ( $G_\mu, s_{\text{eff}}^2, M_Z$ )	2	1	-1	0	0.23154	0	—
LO ( $\alpha_0, G_\mu, M_Z$ )	4	1	-1	0	-1	0	1
NLO ( $G_\mu, M_W, M_Z$ )	2	0	1	0	-1	0	—
NLO ( $G_\mu, s_{\text{eff}}^2, M_Z$ )	2	0	1	0	0.23154	0	—
NLO ( $\alpha_0, G_\mu, M_Z$ )	4	0	1	0	-1	0	1
NLO+HO ( $G_\mu, M_W, M_Z$ )	2	0	1	1	-1	0	—
NLO+HO ( $G_\mu, s_{\text{eff}}^2, M_Z$ )	2	0	1	1	0.23154	0	—
NLO+HO ( $\alpha_0, G_\mu, M_Z$ )	4	0	1	1	-1	0	1

$M_W = 80.385$  GeV,  $\sin^2\theta_{\text{eff}} = 0.23154$ , weak-only=1 (weak only)

# Definitions of $\sin^2\theta_W$ : On-shell vs Effective

## $\sin^2\theta_W^{\text{OS}}$ (on-shell)

:

$$\sin^2\theta_W^{\text{OS}} \equiv 1 - \frac{M_W^2}{M_Z^2}$$

- At tree-level

## $\sin^2\theta_{\text{eff}}^{\text{lept}}$ (effective)

Measured through **asymmetries** at the Z peak :

$$\sin^2\theta_{\text{eff}}^{\text{lept}} \approx 0.23148$$

- Includes radiative corrections at the  $Zf\bar{f}$  vertex
- Measured by  $A_{FB}$

## Relation between the two — the factor $\kappa$

$$\sin^2\theta_{\text{eff}}^{\text{lept}} = \kappa_{\ell} \times \sin^2\theta_W^{\text{OS}}, \quad \kappa_{\ell} \approx 1.037$$

$\kappa_{\ell} - 1 \approx 0.037$  : dominated by  $\Delta\rho \approx 3 G_{\mu} m_t^2 / (8\pi^2 \sqrt{2}) \approx 0.009$