

# Revealing the progenitors of gravitational waves mergers with neural networks

Courty Théo - 1st year PhD  
Biennale 2026

Group: Gravitation

Supervisors: Sylvain Chaty & Eric Chassande-Mottin

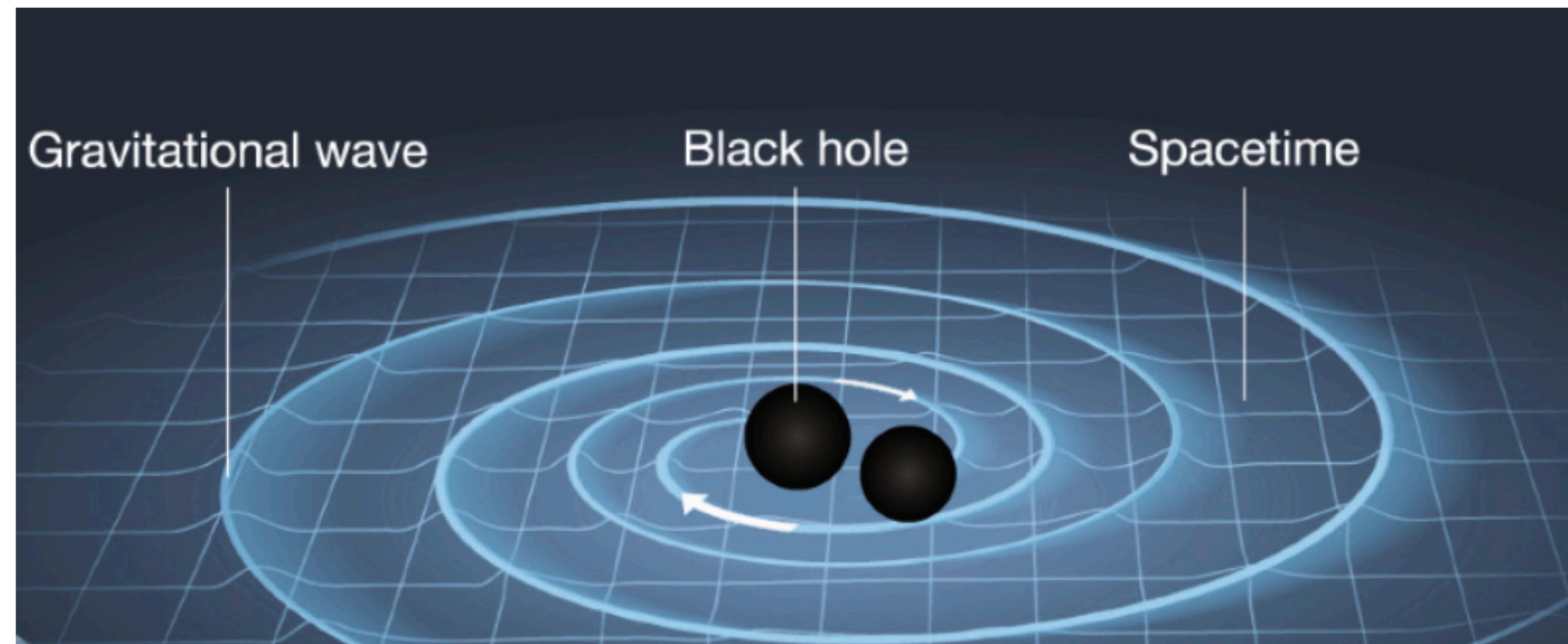
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Université  
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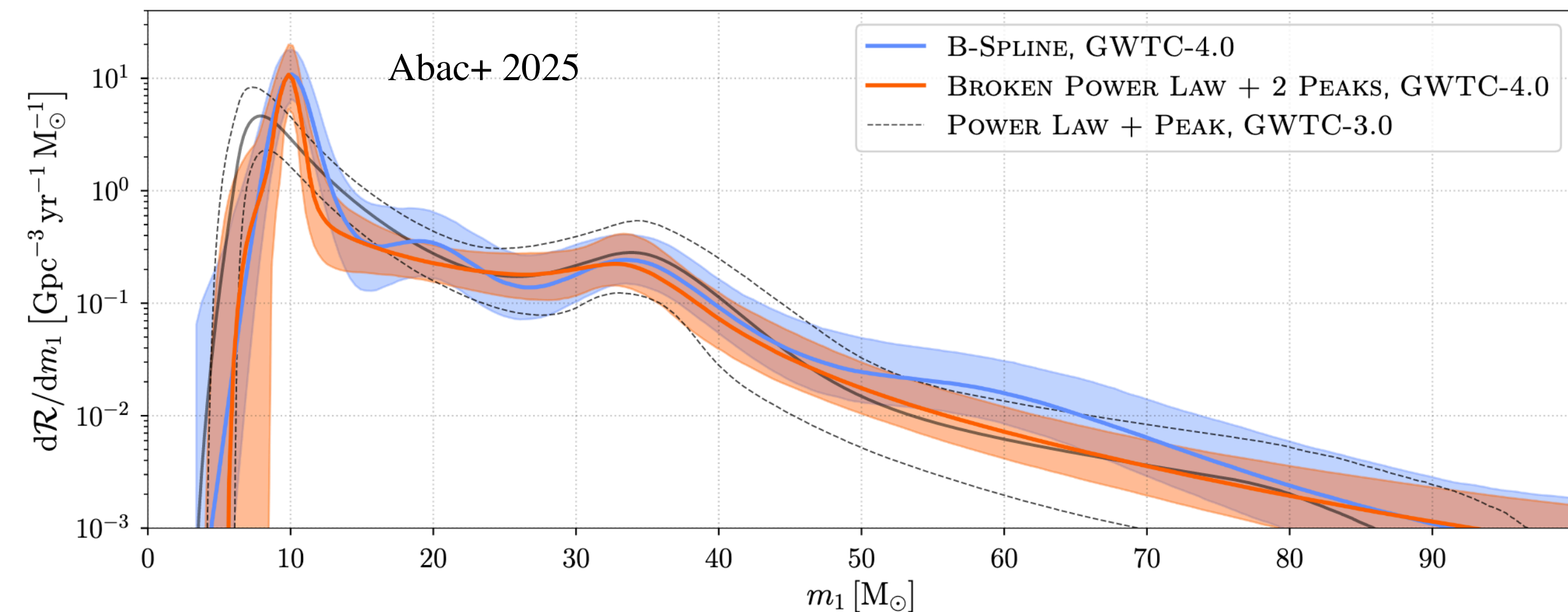


# Gravitational waves & black holes population



What are the **formation channels** of these black holes?

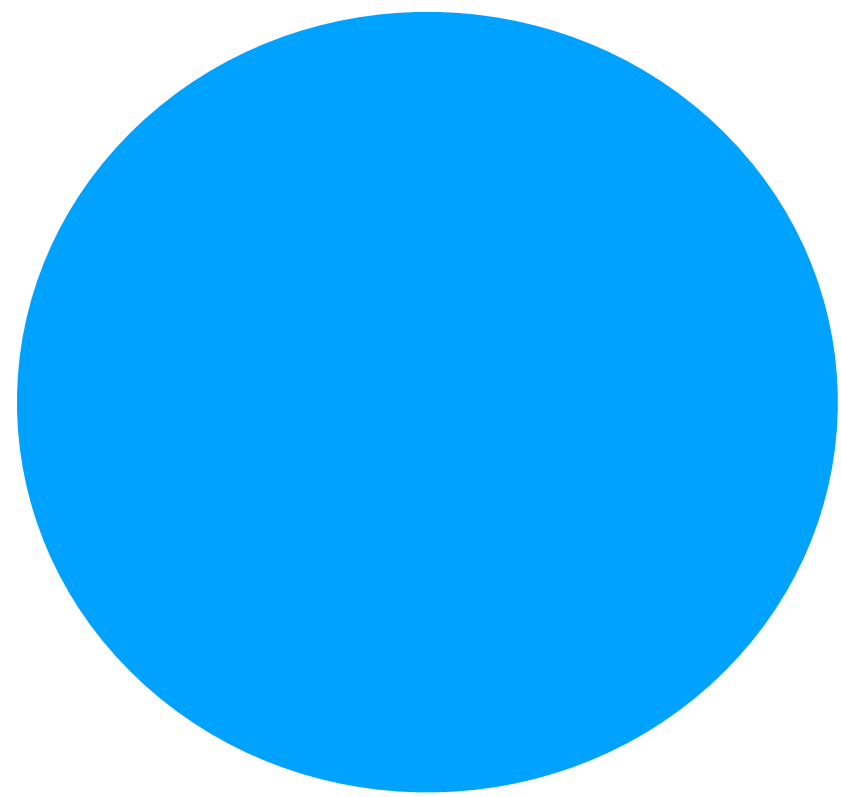
What are the links between binary black holes population and stellar astrophysics?





# Compact object formation from stellar evolution

Initial star



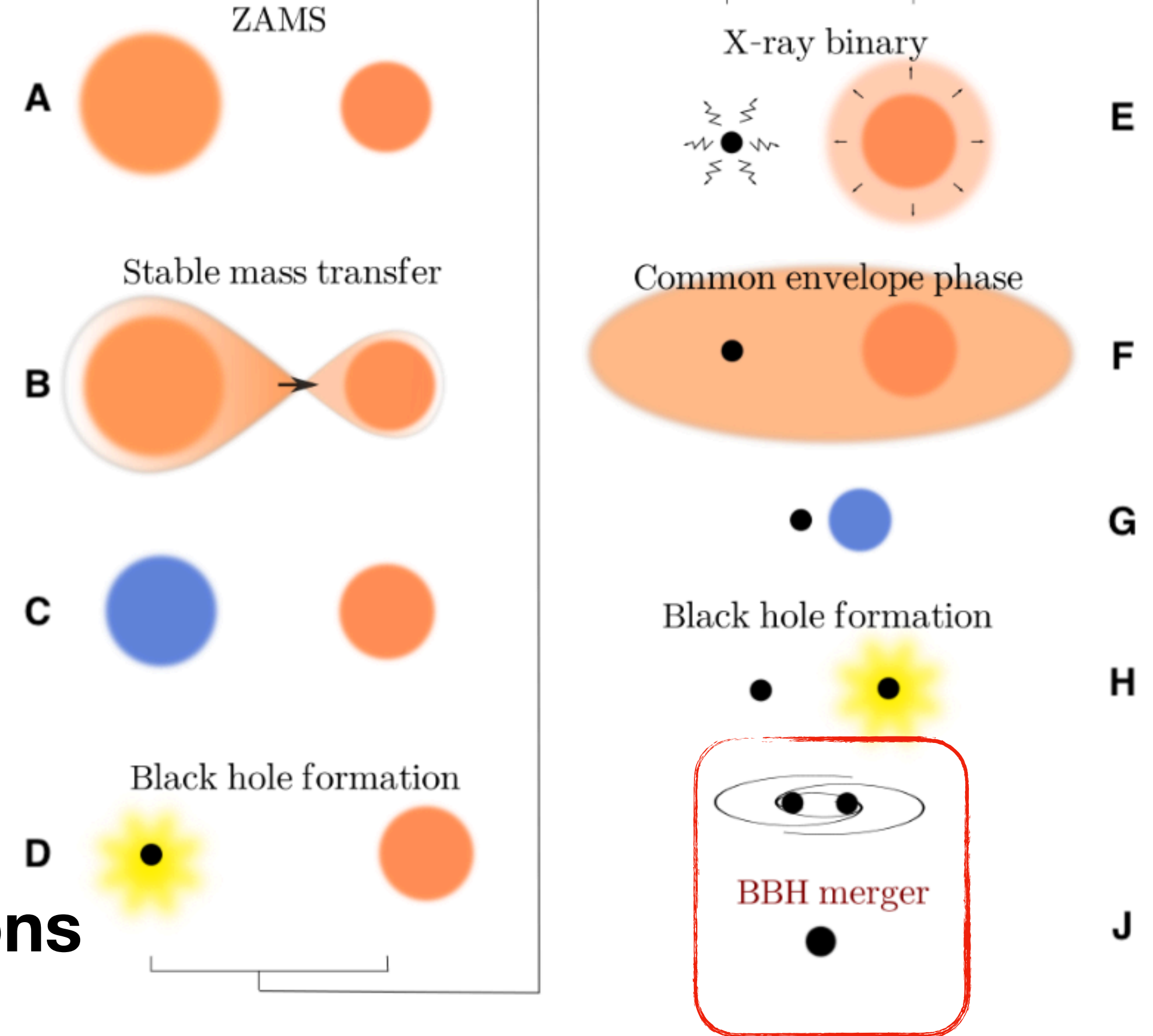
Remnant



Black hole



García+ 2021



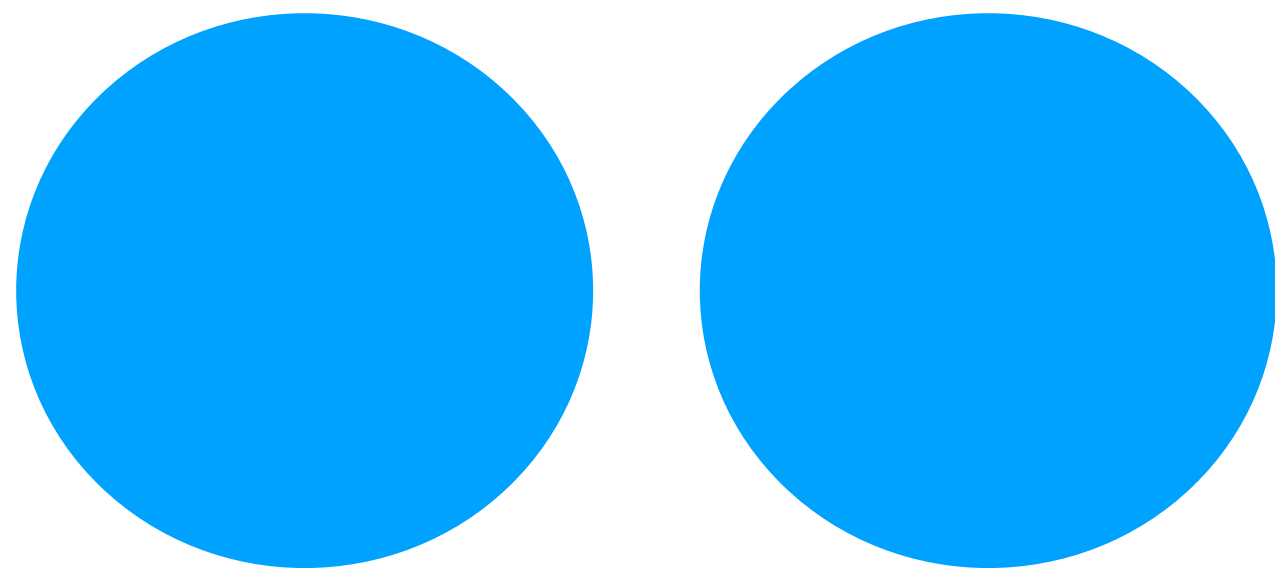
- Black holes are end of massive star evolution
- Binary systems = strong **interactions!**
- Detectable mergers need very **specific conditions**
- **Complex** and poorly constrained scenario

Gravitational waves observations

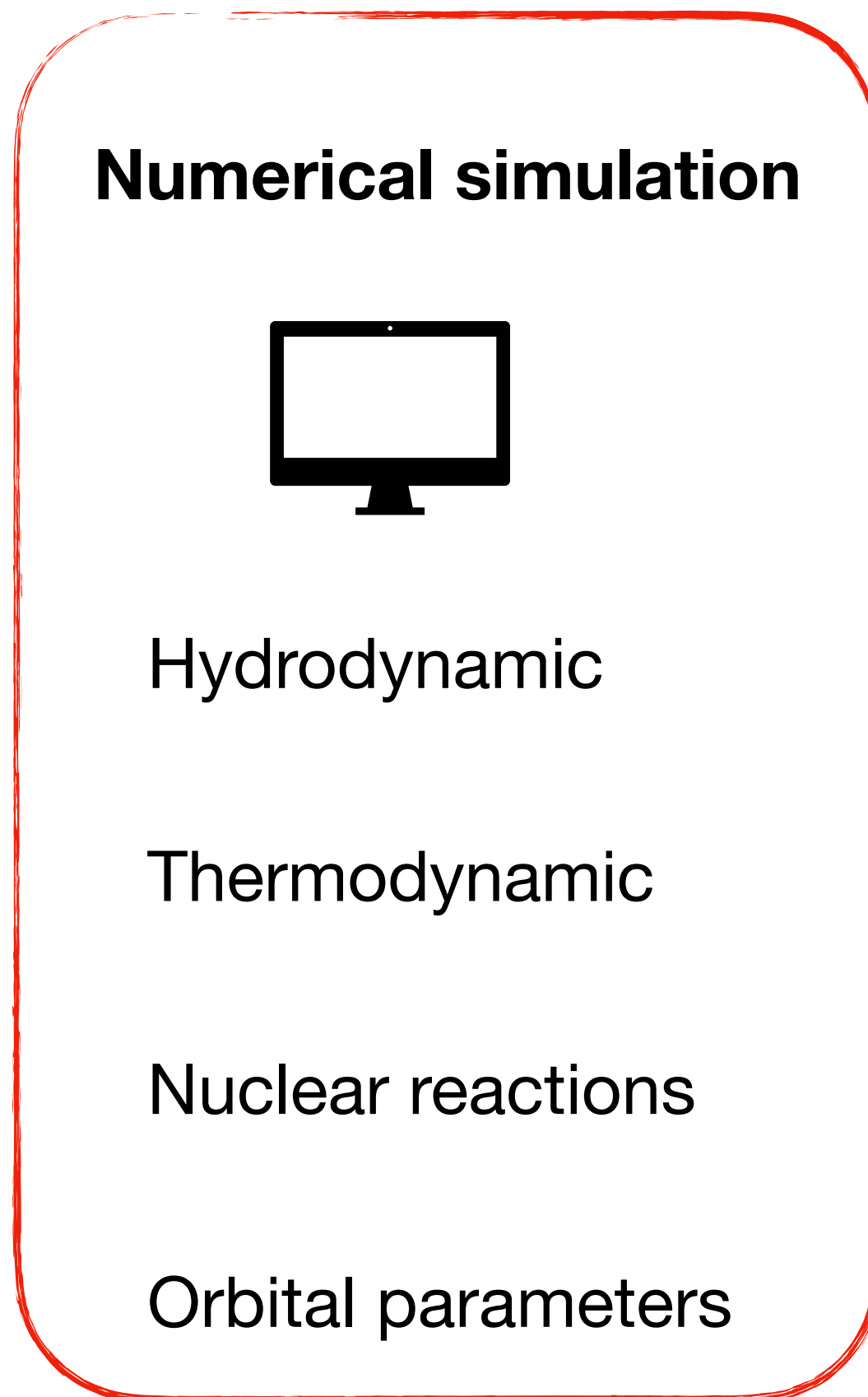


# My PhD in a nutshell

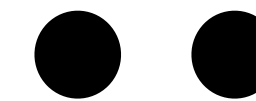
Initial binary properties X



X = (Masses, metallicity, orbital separation, eccentricity, rotations)



Final binary black hole Y



Y = (Masses, spins, orbital parameters)

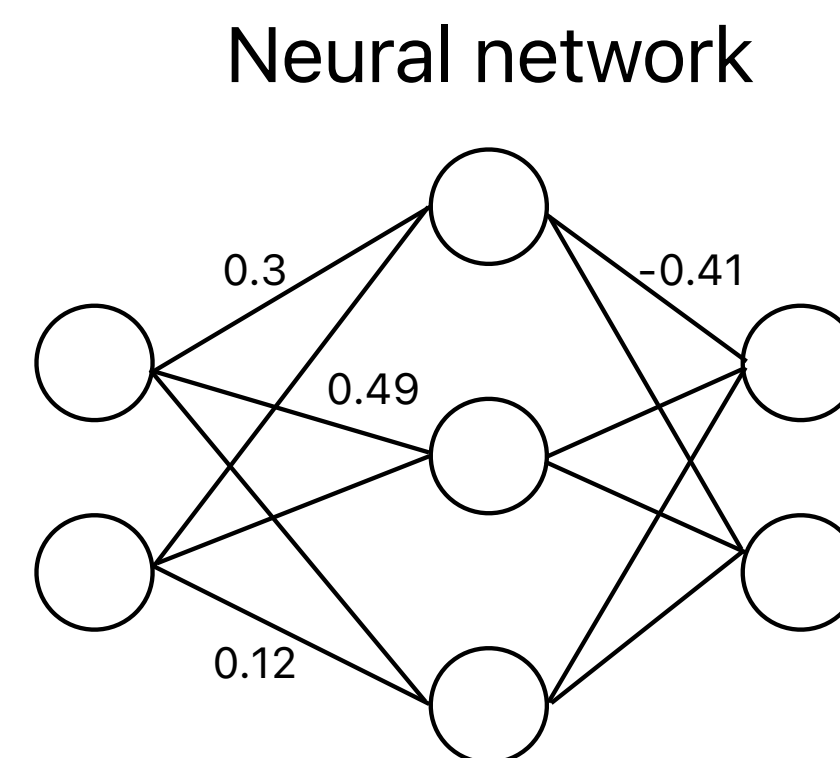
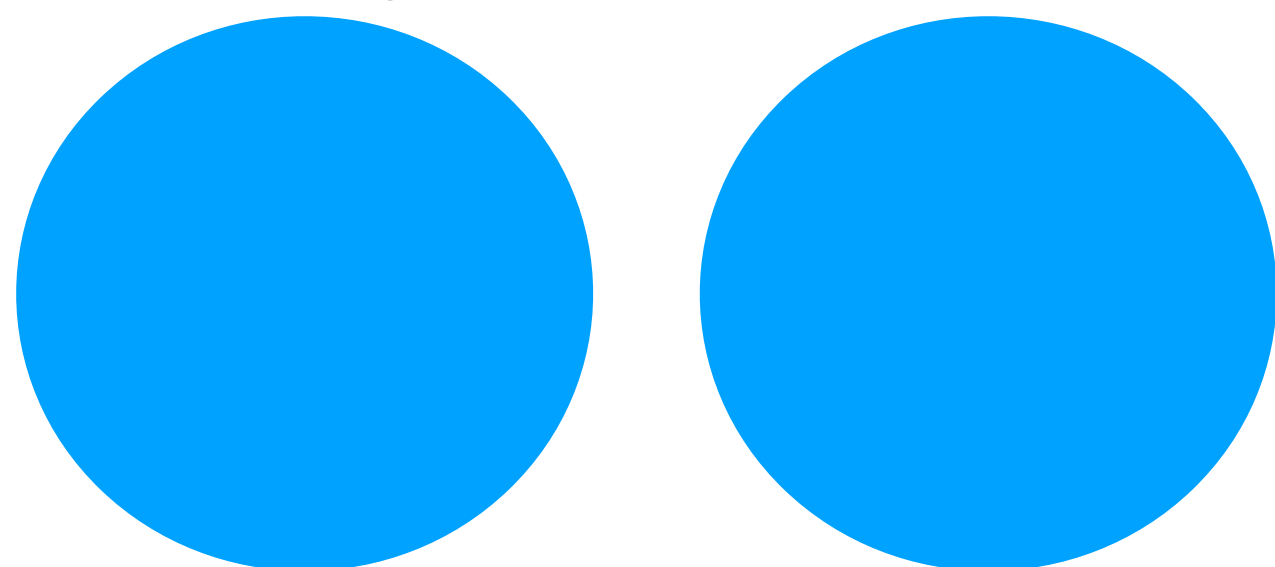
**Prohibitive computational time for population studies!**



# My PhD in a nutshell

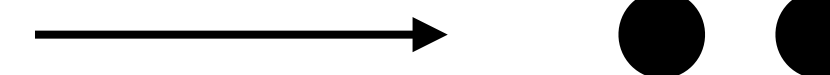
## Initial binary properties X

X = (Masses, metallicity, orbital separation, eccentricity, rotations)



## Final binary black hole Y

Y = (Masses, spins, orbital parameters)



- « Black box » method
- Needs a training dataset anyway



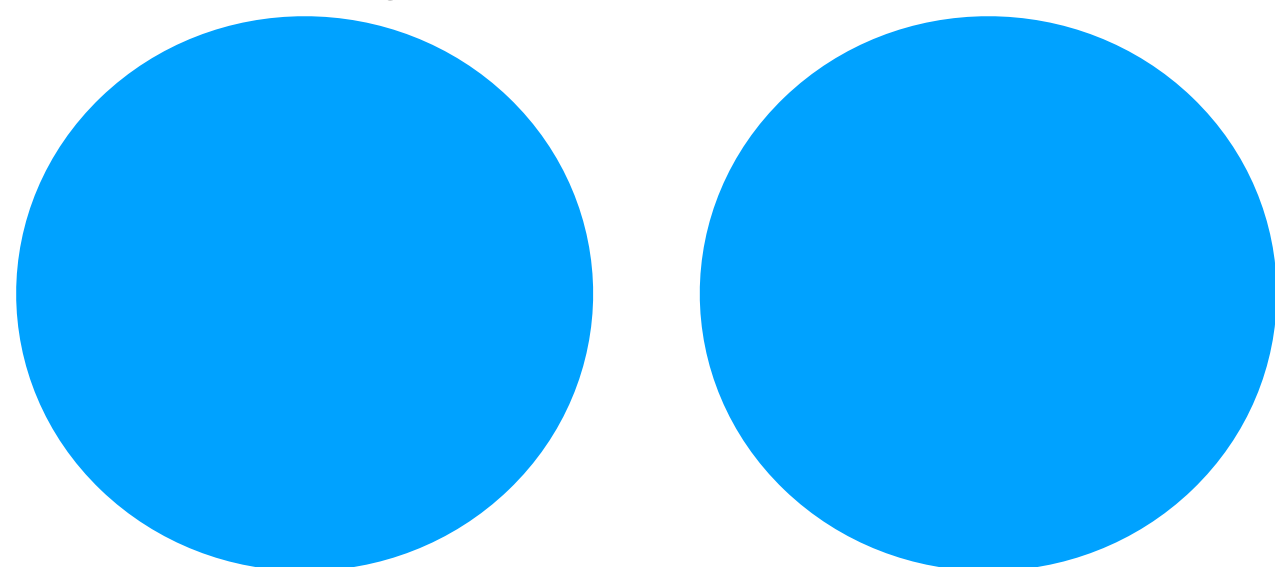
- Efficient to evaluate once trained
- Differentiable by design
- Suited for Bayesian inference pipelines



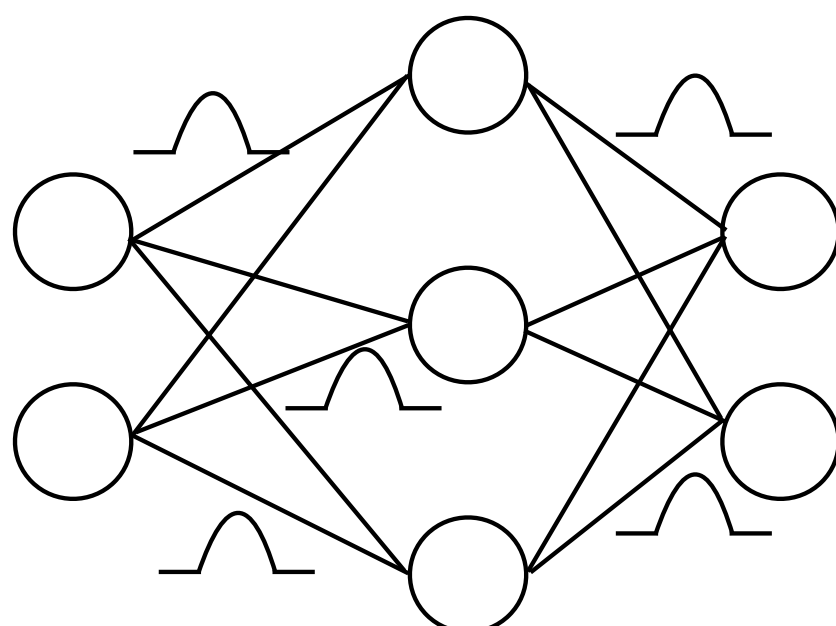
# My PhD in a nutshell

Initial binary properties X

X = (Masses, metallicities, orbital separation, eccentricity, rotations)



Bayesian neural network



Final binary black hole Y

Y = (Masses, spins, orbital parameters)



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## Bayesian deep active learning

- Predictive uncertainty quantification
- Choose training dataset based on low confidence samples (Active learning)
- Automatic process

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