

R&D towards a Water-Based Liquid Scintillator detector for ND280++



Daniel Ferlewicz
MSCA post-doctoral fellow

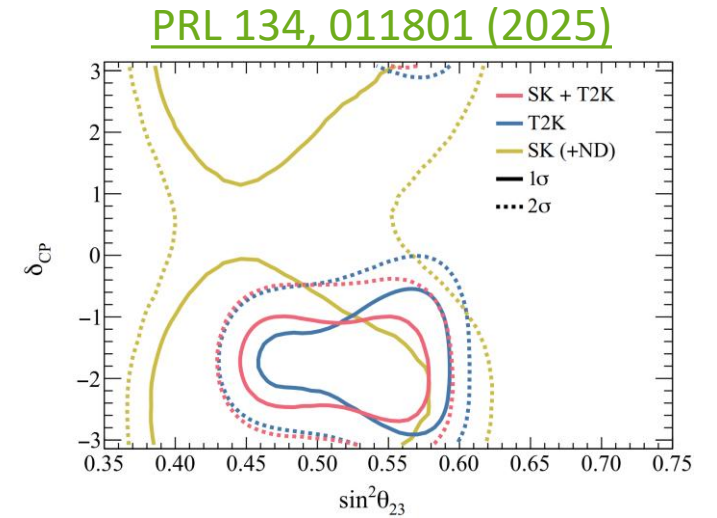
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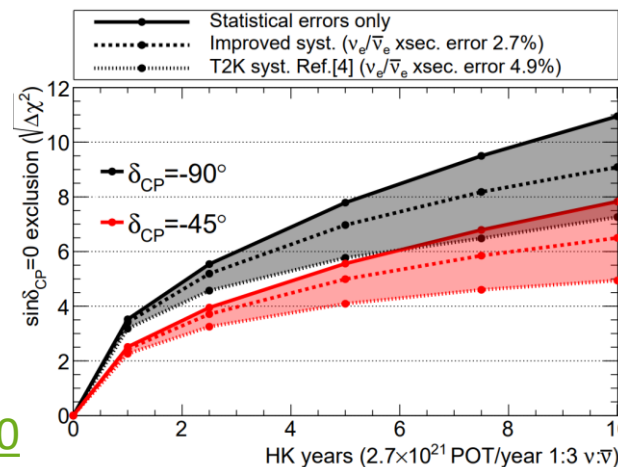
Physics motivations

- Long baseline neutrino experiments measure CP violation by comparing $P(\nu_\mu \rightarrow \nu_e)$ and $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$

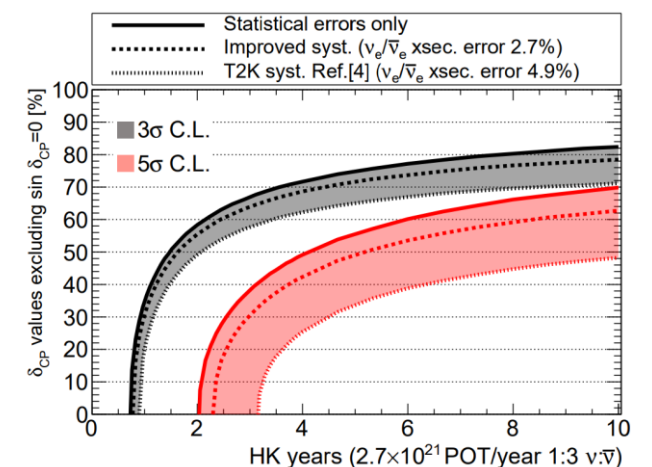
- Searches of CP violation in neutrinos in T2K, SK, Nova
 - [T2K+SK](#) exclude CP conservation at 1.9σ , hinting at maximal CPV
 - [T2K+Nova](#) also shows hints of $\delta_{CP} \neq 0$ depending on mass ordering



- Maximal CPV discovery at 5σ during the high-statistics HyperK era impacted by a **4% systematic uncertainty in $\sigma_{\nu_e}/\sigma_{\bar{\nu}_e}$**

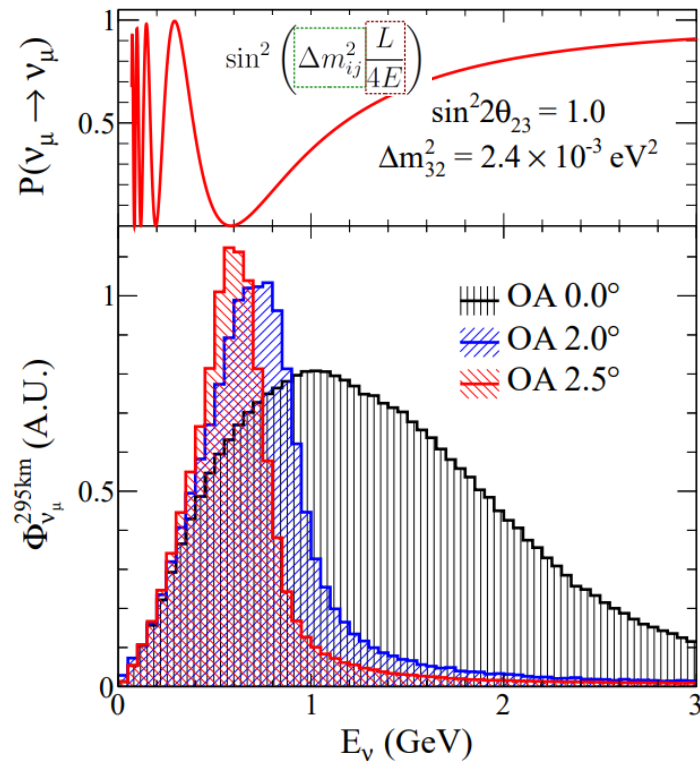
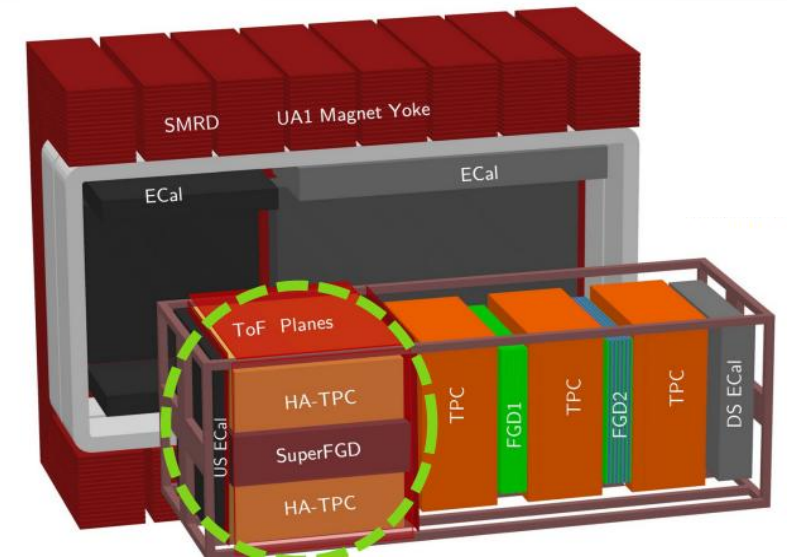


[Eur.Phys.J.C 86 \(2026\) 2, 170](#)

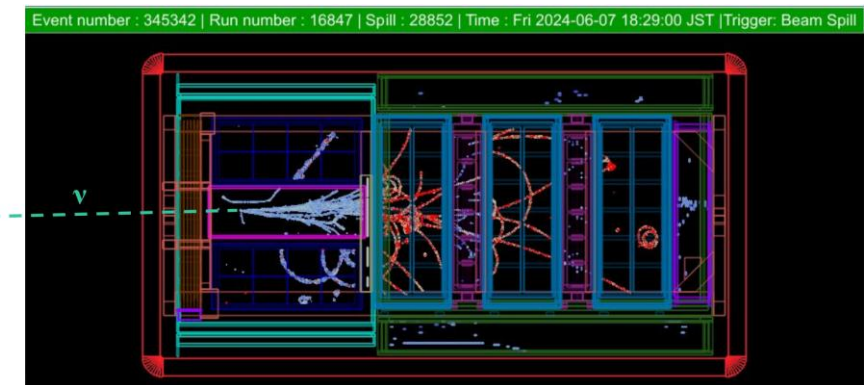


ND280

- Near detector 280m from the ν -beam origin at J-PARC
 - Installed in 2010 for the T2K experiment
 - To be inherited by HyperK experiment as its ND in ~ 2028
- Off-axis at 2.5° for narrow energy band at FD

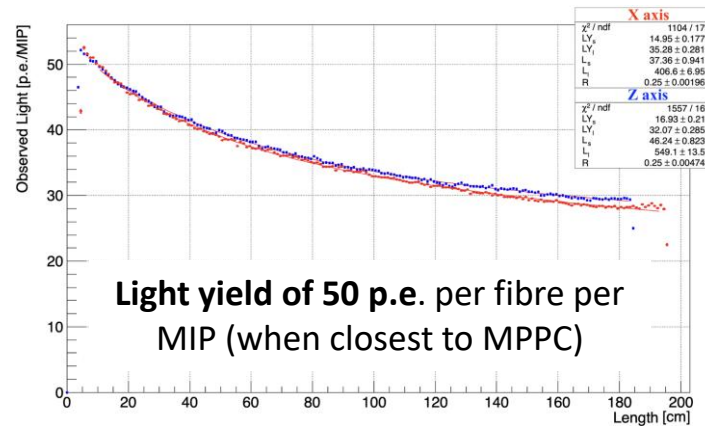
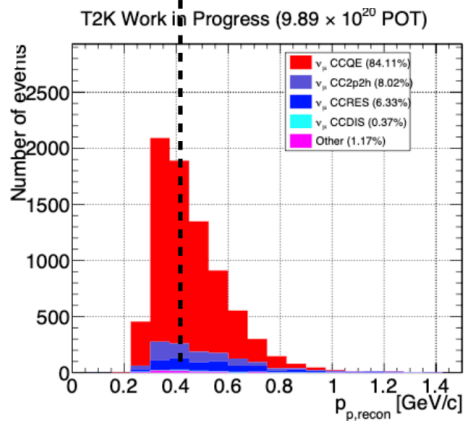
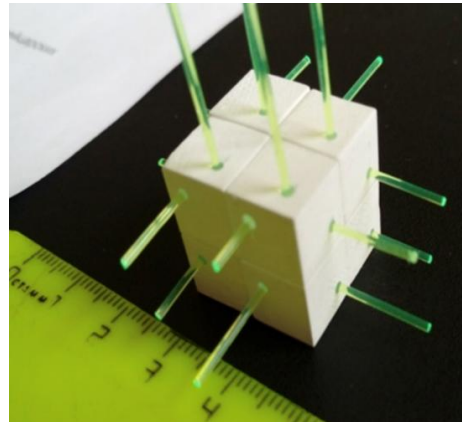
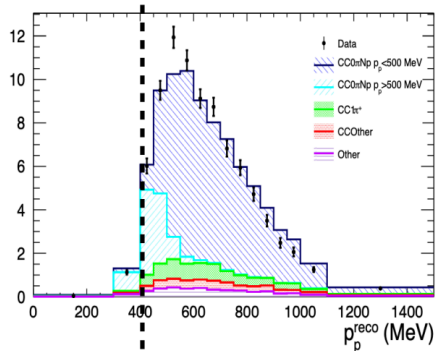


- Current configuration installed in April 2024 (“ND280+”)
- Downstream detectors: Time Projection Chambers (TPCs) for tracking and Fine Grain Detectors (FGDs) as targets
 - FGD2 has the only (small) water target in ND280
- Newer upstream detectors:
 - SuperFGD
 - Two High-angle TPCs
 - Six Time of Flight panels



SuperFGD

- 2 tonne target of optically separated plastic scintillator cubes
- Improved low momentum proton efficiency
- 4π angular acceptance



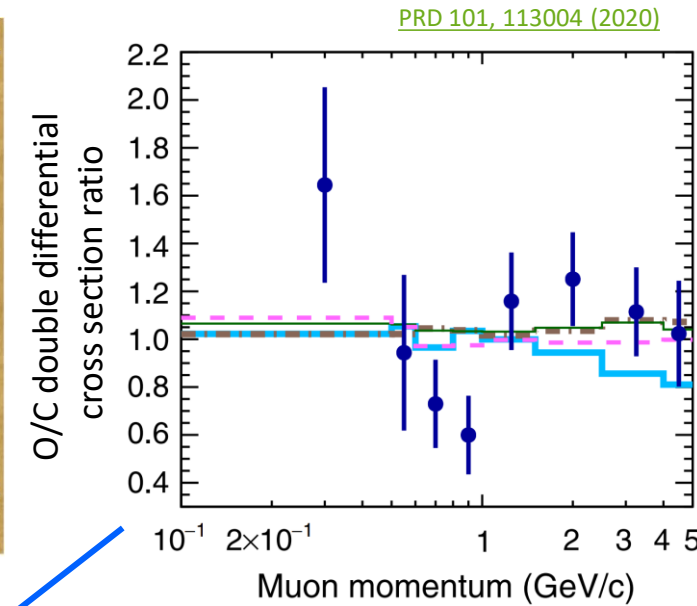
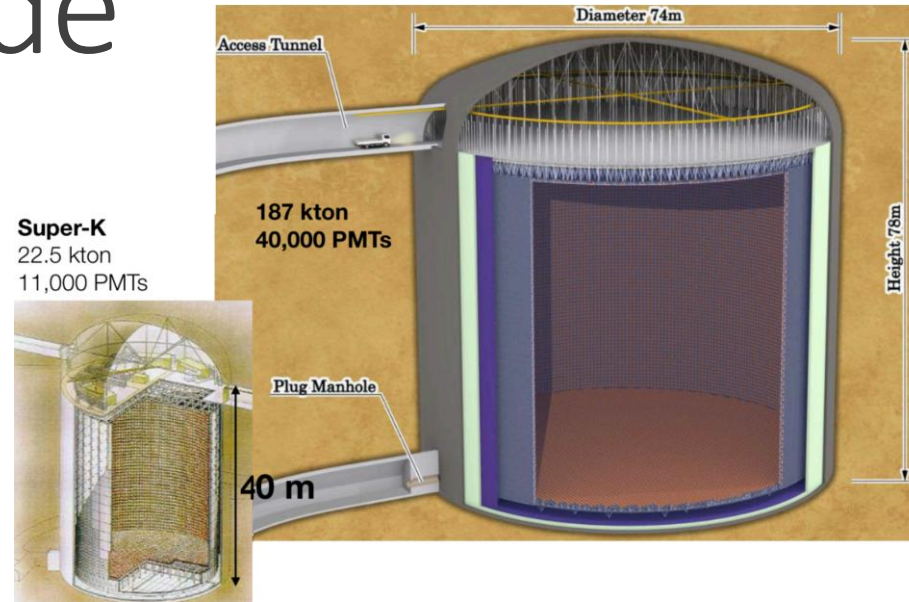
- ~ 2 million $1 \times 1 \times 1$ cm³ cubes
- 3D readout using WLS fibres, ~ 56 k channels

Paper under review with NIM

[arXiv: 2603.14921](https://arxiv.org/abs/2603.14921)

HyperKamiokande

- FD completion in 2028
 - Fiducial volume 8x larger than SK
 - Beam power: 750kW (T2K) → 1.3MW
 - 1 year HK data \approx 20 years of T2K
- Intermediate water Cherenkov detector under construction
 - See backup slides
- HK will inherit ND280:

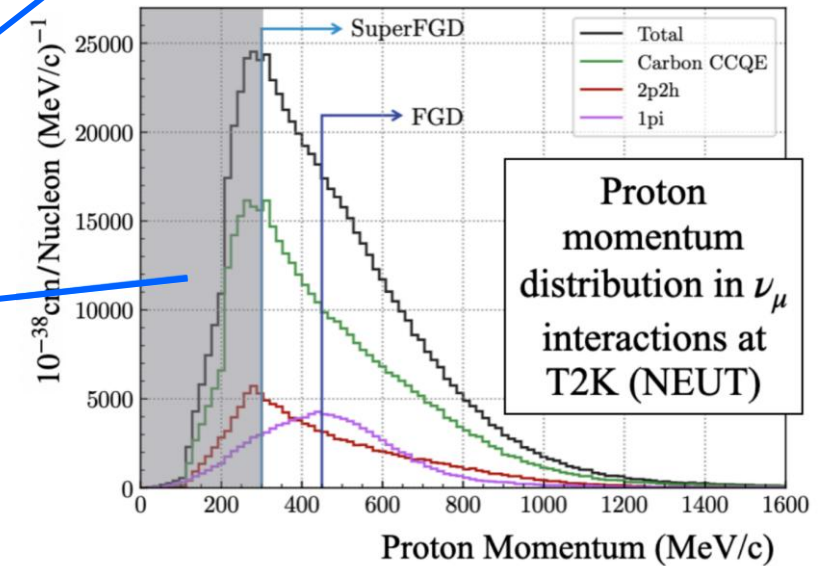


Neutron kinematics important for $\bar{\nu}$ interactions but many escape SFGD
 ↓
Need larger target

Scaling carbon interactions in ND to oxygen in FD will be a larger uncertainty
 ↓
More water target in ND280

δ_{CP} will be dominated by $\nu_e/\bar{\nu}_e$ in H_2O cross section uncertainties
 ↓
Need higher ND statistics

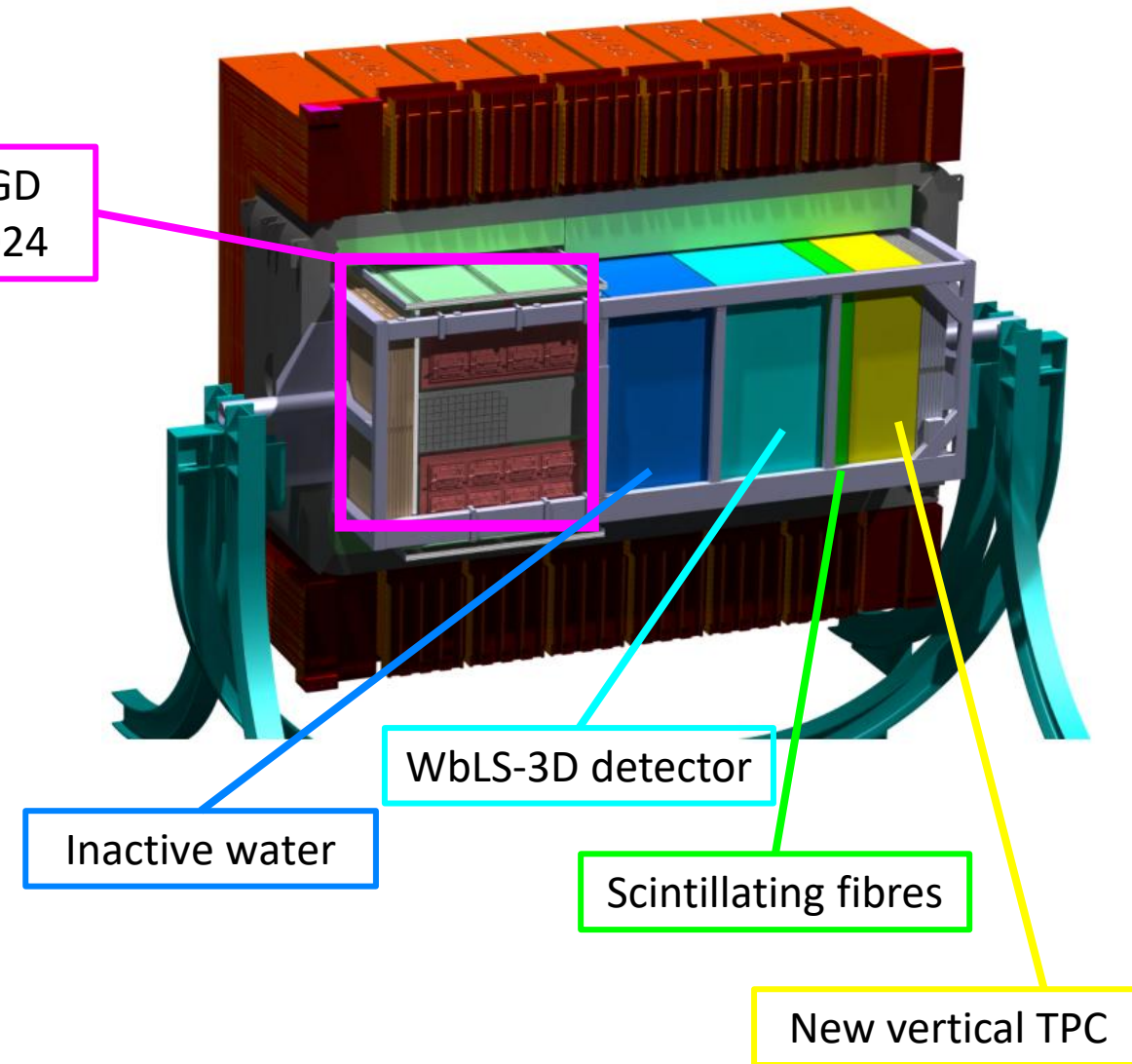
Proton threshold in the SFGD still misses half of tracks
 ↓
Need a target with better granularity



HK-ND280++

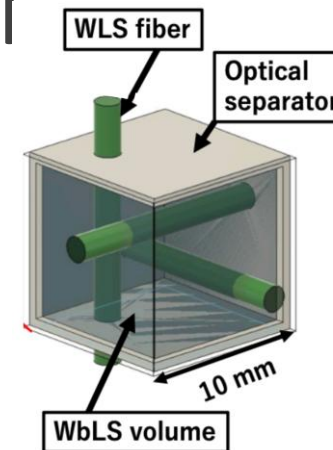
Reference for possible [ND280++ design](#) – not final

- R&D for ND upgrades underway:
 - Replace tracker region after 2030
 - New ND280+ detectors will remain
- **10 tonnes** available for new ideas
 - Large fraction for water
 - New vertical TPC
 - R&D for TPC optical readout at CEA Saclay
- Refurbished Electromagnetic Calorimeter
 - Replace MPPCs with latest available
- Expect $\sim 4000 \nu_e / 10^{21} \text{ POT}$ at 20% efficiency, comparable to IWCD



Water-based Liquid Scintillator

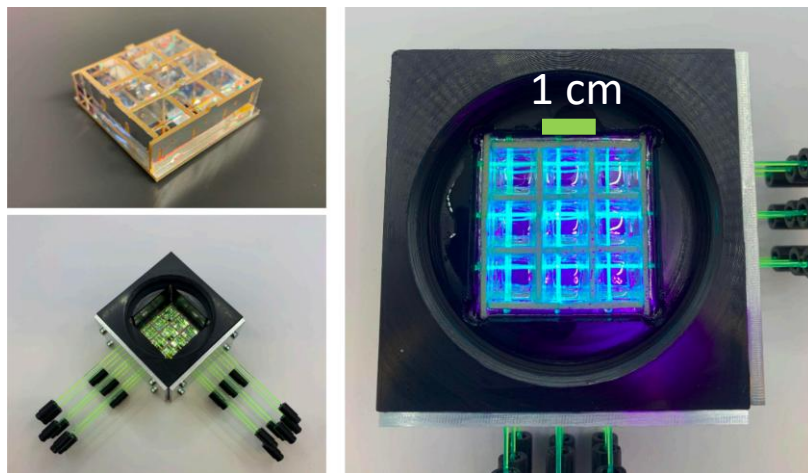
- SFGD-like highly segmented WbLS detector with 3D readout
 - Maximize H₂O content
 - Light Yield > 10 p.e.
- Developing prototypes at LPNHE, ETH Zurich and Kyoto U
 - In collaboration with WbLS inventor M. Yeh (BNL) [NIM A 660, 51–56 \(2011\)](#), [JINST 19 P01003 \(2024\)](#)



Neutron kinematics important for $\bar{\nu}$ interactions but many escape SFGD
 ☺ Need larger target

Scaling carbon interactions in ND to oxygen in FD will be a larger uncertainty
 ☺ More water target in ND280

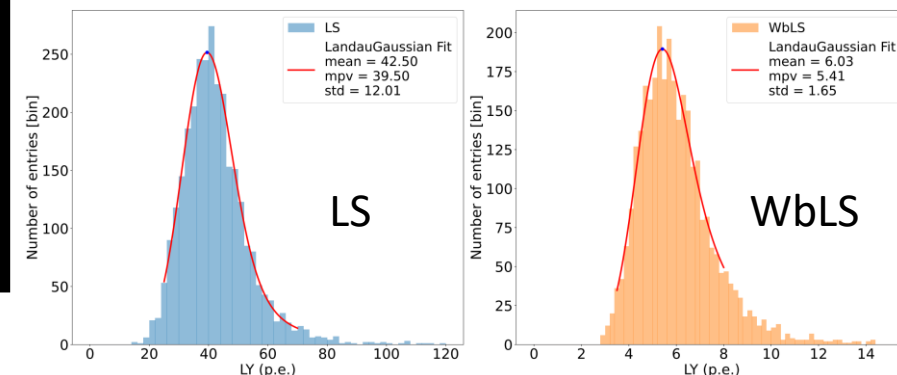
δ_{CP} will be dominated by $\nu_e/\bar{\nu}_e$ in H₂O cross section uncertainties
 ☺ Need higher ND statistics



ETH prototype, submitted to JINST
[JINST 21 \(2026\) 01, P01012](#)

- ETH prototype:
 - Composition of **90% H₂O**, 10% LS (LAB+PPO+MSB)
 - 3M and Divinycell for optical separation, >80% water by weight
- Average light yield over 6 channels for cosmic rays:

	<LY> per fiber	xtalk
LS	42.5 p.e.	2.3%
WbLS	6.0 p.e.	1.9%



- PID in organic scintillator can be done at lower light yield than the SFGD

WbLS studies at LPNHE

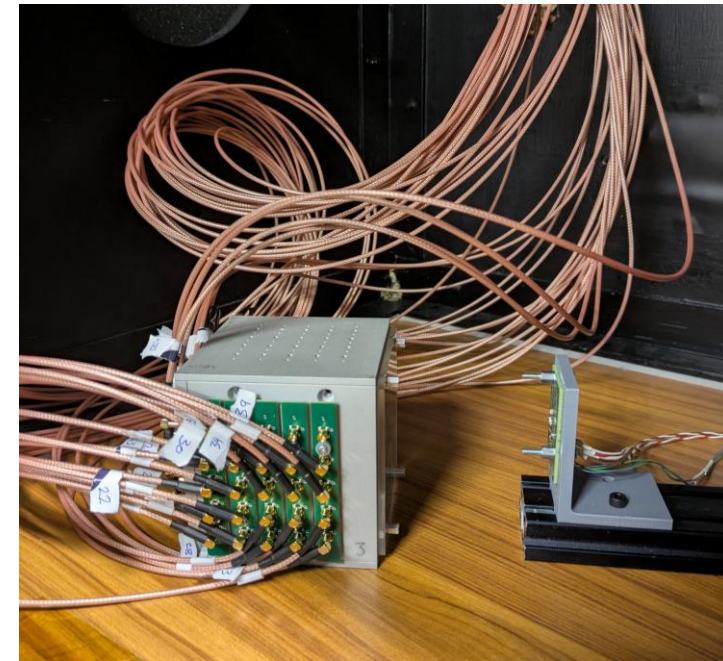
- Exploring detector design: WLS fibres, SiPMs, reflective materials
- Experimental setups include:
 - 5x5x5 microSFGD
 - Precision fibre and SiPM characterisation
 - 8cm³ SuperCube
- Designing a small liquid-tight vessel for WbLS tests

LPNHE team

C. Giganti
R. Gaio
E. Hily
X. Aubert
G. Daubard
D. Laporte
M. Antonelli
D.F.

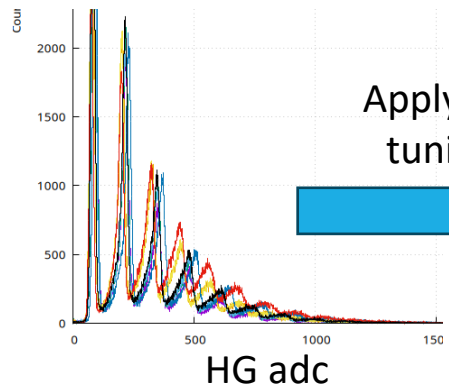
microSFGD

- 5x5x5 SFGD cubes, with additional gap in between
- Using CITIROC chips in a CAEN DT5202
- 3mm or 1mm SiPMs on a PCB with 3D printed guard
- Kuraray Y11 fibres polished on both ends
- Set up x- and y- sides with 20 channels each (40 total)

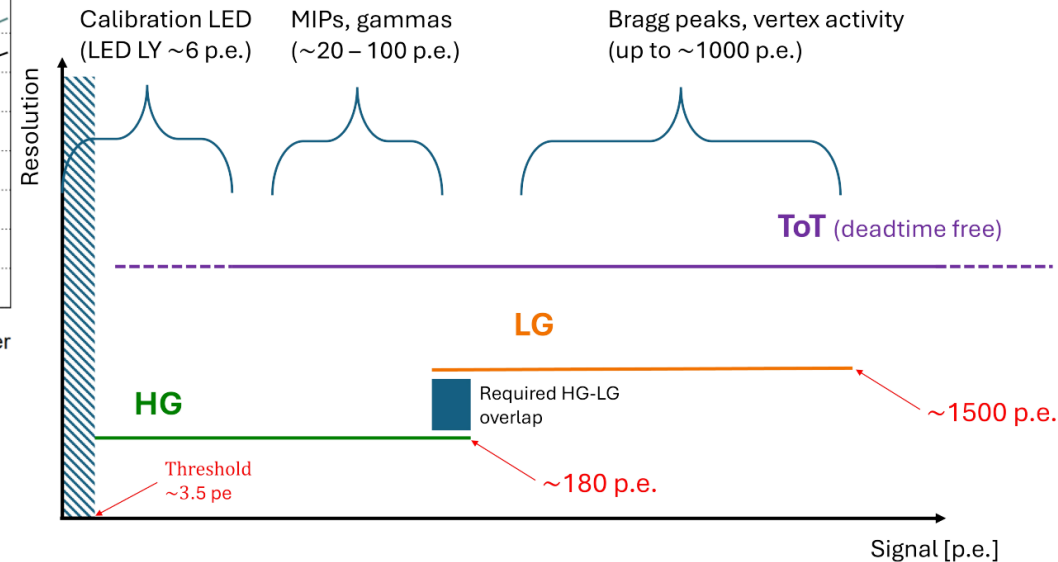
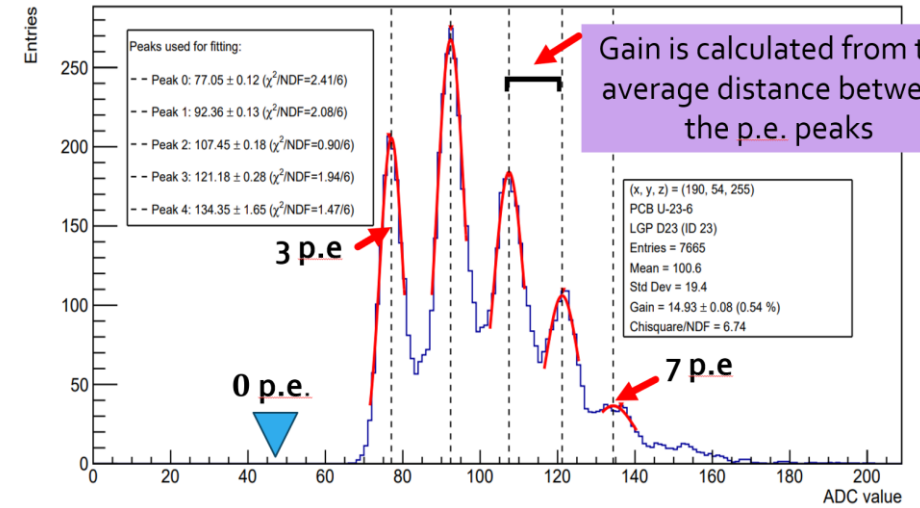
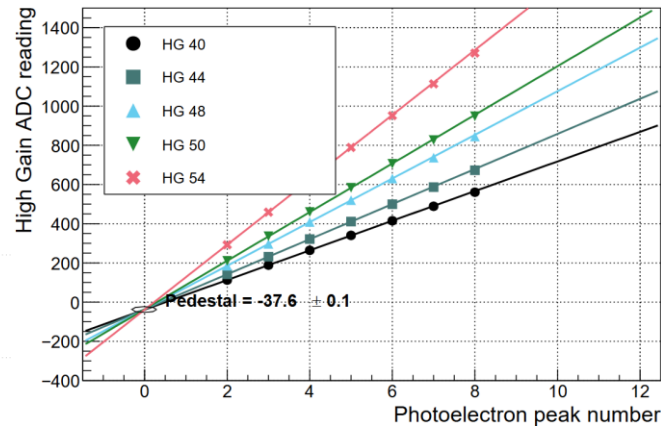
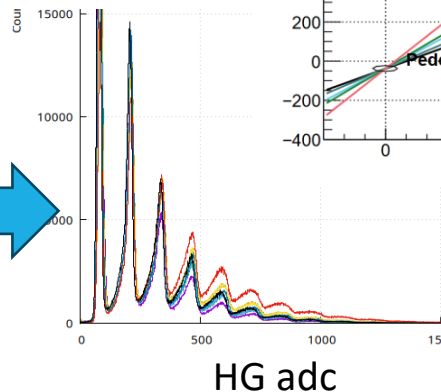


LED Calibration

- Low intensity LED pulses are used to characterise each MPPC, similar to the SFGD
- Tuned all bias voltages for similar channel responses
- For photoelectron charge deposit, need to measure high gain, pedestal and LG-HG conversion
 - Not considering ToT at this stage

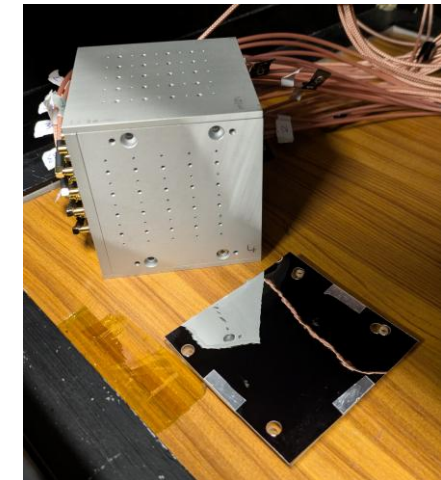
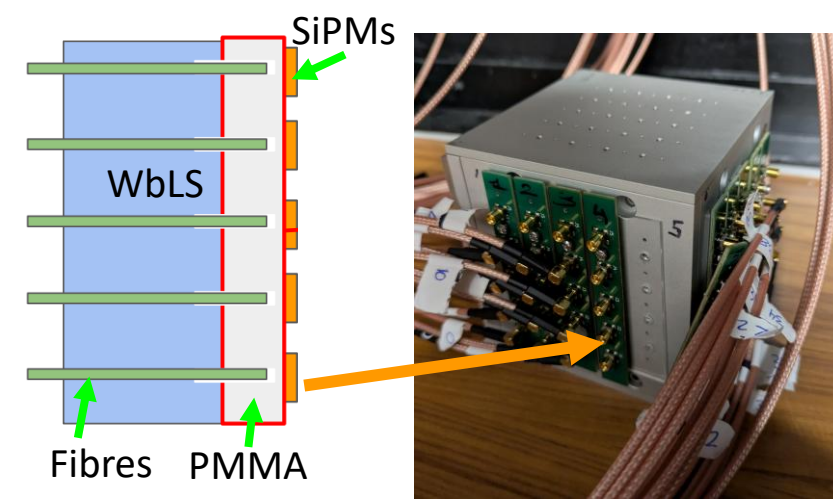
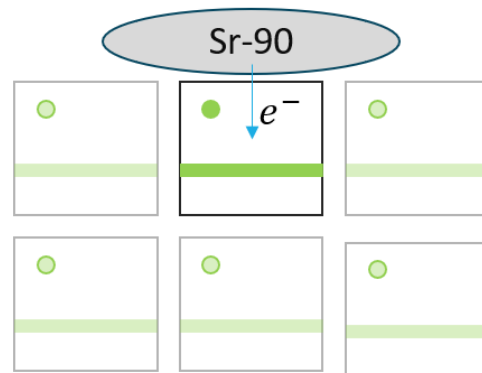


Apply HV tuning



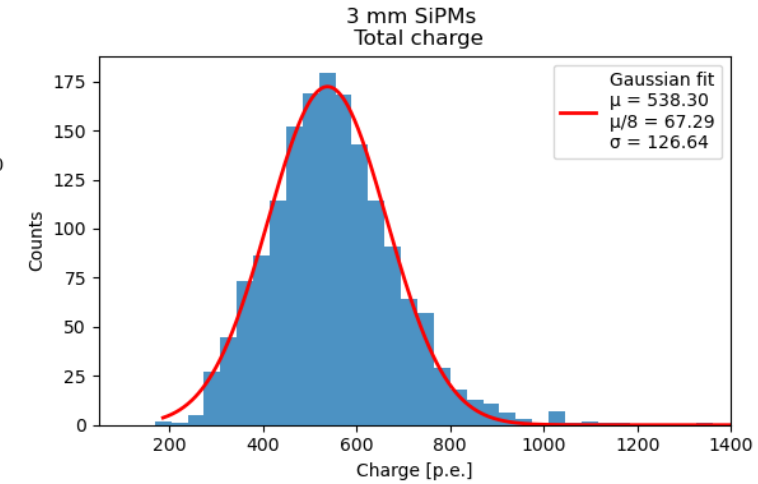
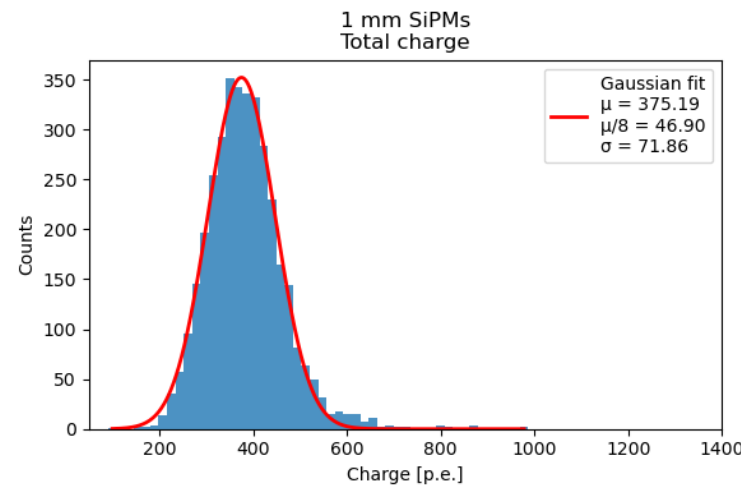
microSFGD studies

- Need a water-tight vessel in future – consider blind holes
 - Study effect of having PMMA between fibre end and SiPM
- Compare 1mm and 3mm SiPM performance
 - 1mm is Hamamatsu S13081-050CS
 - 3mm is Hamamatsu S13360-3075CS
- Study the effect of adding a reflective sheet on the opposite fibre end
- Took measurements with Sr-90 beta source and verify with cosmic data



microSFGD results

- In cosmics, took events that had vertical tracks
- 3mm SiPM has higher LY than 1mm by **factor 1.4**
- Adding a reflective sheet improves LY by **factor 1.15**
- Adding PMMA between fibre end and SiPM reduces LY by only a **factor of 0.91**



SiPM size	Reflector	PMMA	Mean L.Y/fibre [p.e.]
1mm	No	No	47
1mm	Yes	No	53
3mm	No	No	67
3mm	Yes	No	77
3mm	No	Yes	61
3mm	Yes	Yes	72

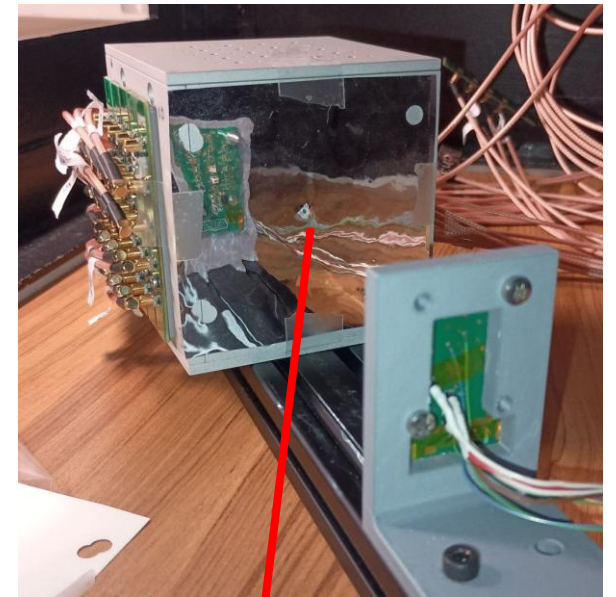
1.4x

0.91x

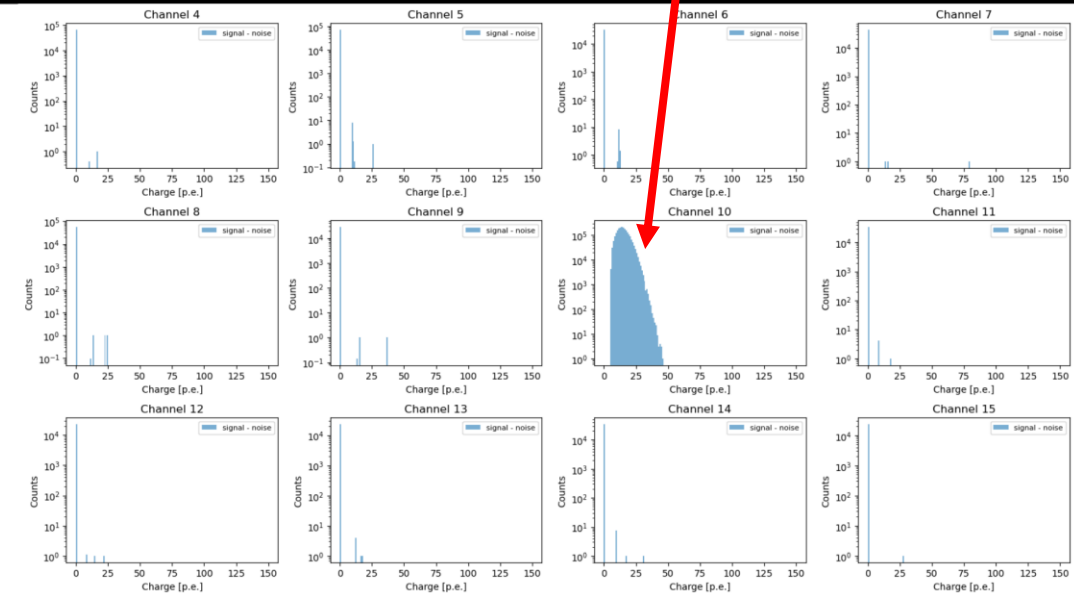
1.15x

microSFGD optical isolation

- PMMA between fibre and SiPM will spread light, and bigger SiPM surface may pick up signal from other channels
- Shine LED on only one fibre end and measure LY in surrounding channels
- Subtracted the noise and found no significant difference with and without PMMA

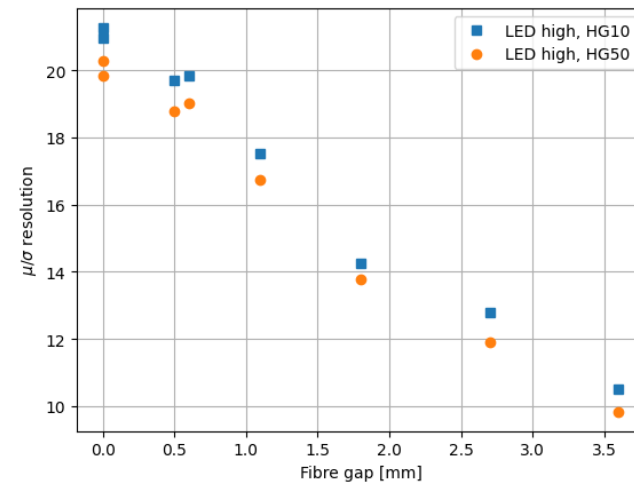
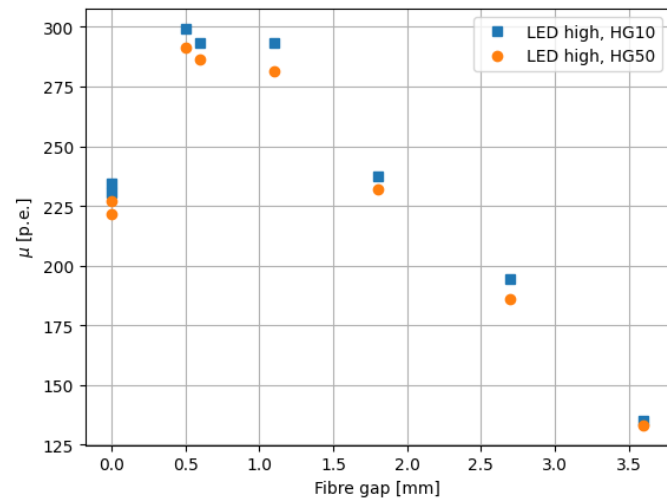
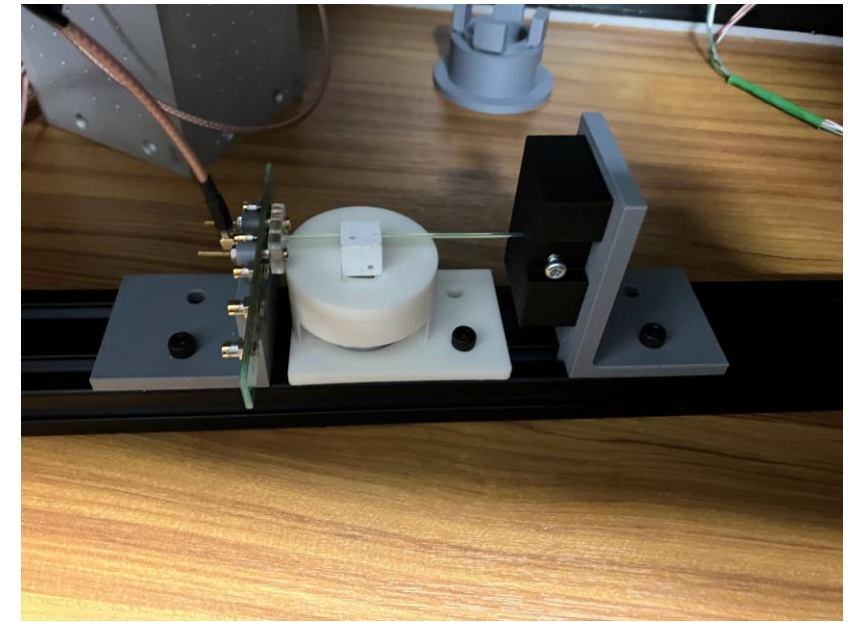


3 MM SIPMS + PMMA

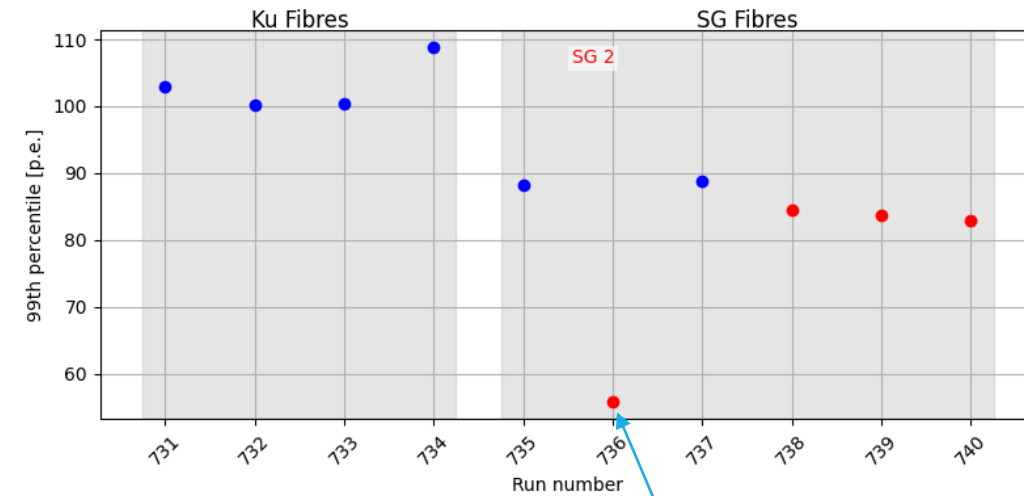


Fibre precision studies

- Fixed PCB and fibre+LED onto guided rails
- Measured mean light yield and resolution in 3mm SiPM as a function of fibre distance from SiPM



- Compared light yield from Kuraray Y11 and Saint Gobain BCF 92 fibres
 - Found Ku fibre had 15% higher light yield



Outlier result with one SG fibre. Repeated measurement and found consistent results

Conclusion

- Developing ND280++ to reduce ν_e cross section uncertainties
 - Planning construction in the 2030s
- Results at LPNHE are guiding the WbLS detector design
 - 3mm SiPMs: LY \nearrow by 40%
 - Reflective sheets on fibre ends: LY \nearrow by 15%
 - PMMA does not introduce significant LY loss or optical cross-talk between SiPMs
 - Larger prototype will be built in Japan to be exposed to particle beam
- Developing our own small prototype to use with WbLS

Back up

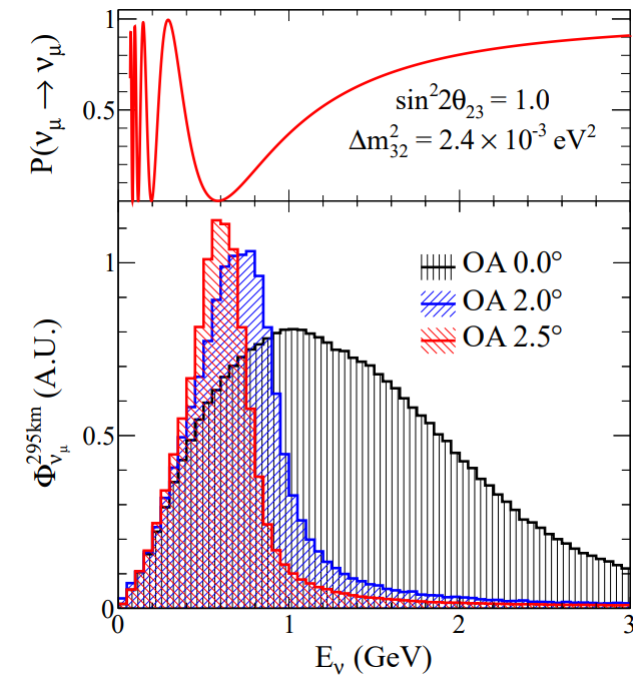
Acronyms

- LAB: Linear Alkyl Benzene
- PPO: Primary solute 2,5-diphenyloxazole
- bisMSB: p-bis-(o-methylstyryl)-benzene
- PC: pseudocumene
- BC: benzethonium chloride
- LAS: linear-alkylbenzene-sulfonate
- PTFE: polytetrafluoroethylene
- PMMA: Poly(methyl methacrylate)

T2K

- Tokai 2 Kamioka taking neutrino data since 2010
- J-PARC's 30GeV proton beam $\rightarrow \nu_\mu$ ($\bar{\nu}_\mu$) beam at $\sim 800\text{kW}$
- ND280 and SK 2.5° off-axis \rightarrow narrower E_ν

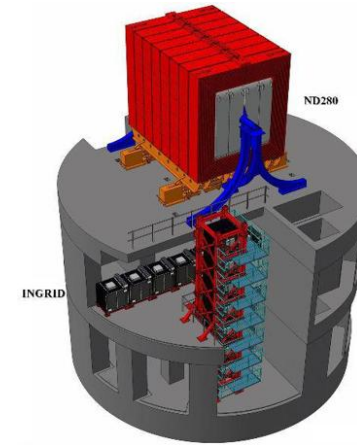
$$\sin^2 \left(\Delta m_{ij}^2 \frac{L}{4E} \right)$$



Super-Kamiokande

Mt. Noguchi-Goro
2,924 m

Mt. Ikeno-Yama
1,360 m



Near Detectors

J-PARC

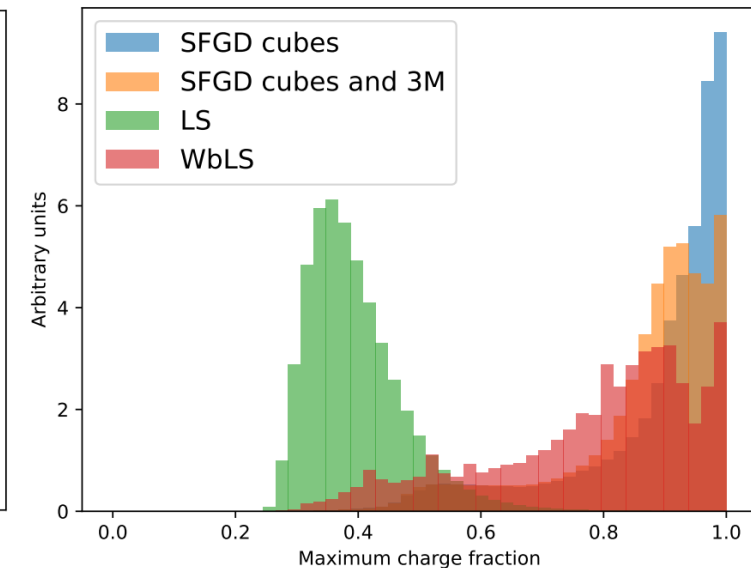
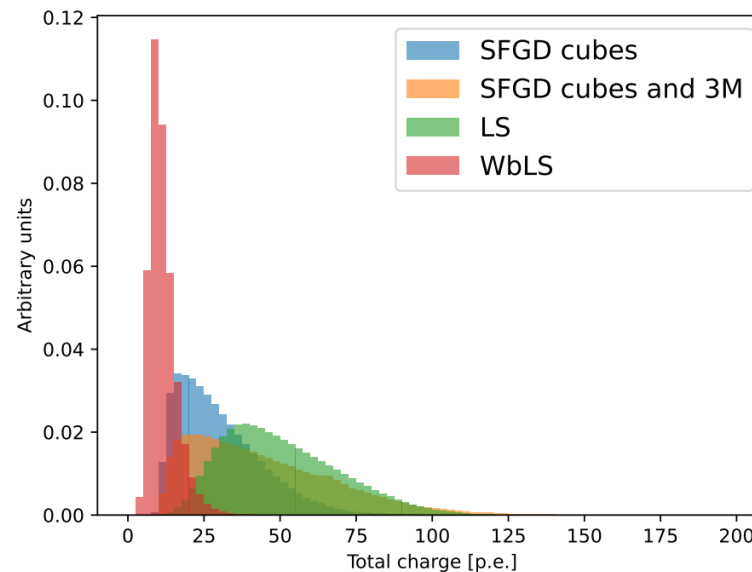
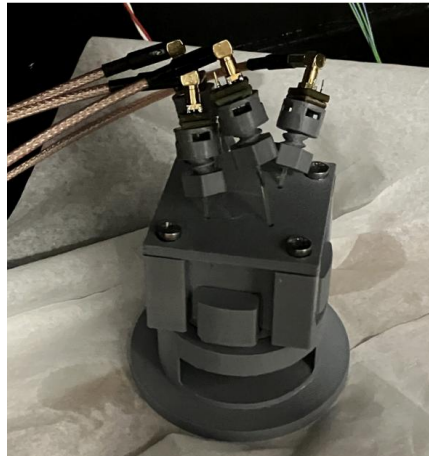
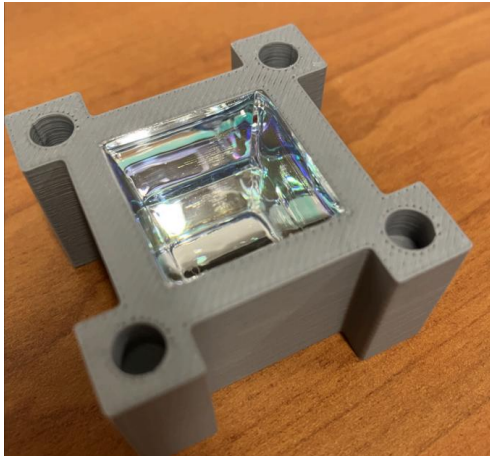
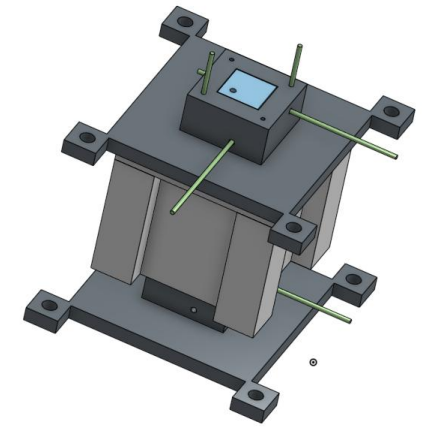
Neutrino Beam

1,700 m below sea level

295 km

Previous study: 8cm³ SuperCubes

- Set up with 4 channels, and 3M reflector on the edges
- Compared light yield from SFGD cubes, LS and WbLS
- Found WbLS to have LY 5x smaller than LS
- Design flaws included non-watertight lid, unstable SiPM placement, incorrect pedestal calibration algorithm (now updated), no segmentation
- Considering re-design for better precision



microSFGD results

