

Reactor CE ν NS Detection with MIMAC

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Sino-French Collaboration Team

LPSC (Grenoble) : D. Santos, F. Naraghi; I. Scheinbein (Theory); O. Guillaudin, N. Sauzet (SDI); E. Lagorio, O. Bourrion (Electronics); T. Descombes (Data Acquisition); J-F. Muraz (COMIMAC)

CPPM (Centre de Physique des Particules de Marseille, Marseille): C. Tao

IRSN- LMDN (Cadarache): M. Petit, T. Vinchon (neutron spectroscopy)

SYSU (Sun Yat-sen University, Shenzhen): Tao Yi

THU (Tsinghua University, Beijing): Gao Fei (A prototype hosted in THU now)

USTC (University of Science and Technology of China, Hefei): Zhang Zhiyong

SJTU (Shanghai Jiao-Tong University): Wang Shaobo, Han Ke

SHAO (Shanghai Astronomical Observatory): Shan Huanyuan

Joint goal: Develop **directional** detection of **CEvNS** using the MIMAC gaseous TPC.



CEvNS (Coherent Elastic ν -Nucleus Scattering)



A long-awaited neutral-current process ...

- Proposed by **D. Freedman** in 1974 [PRD 9, 1389 \(1974\)](#)
- First observed by **COHERENT** in 2017 [Science 357, 1123-1126 \(2017\)](#)
- The neutrino **scatters coherently** from the entire nucleus when $qR \lesssim 1$

$$\nu + A \rightarrow \nu + A$$

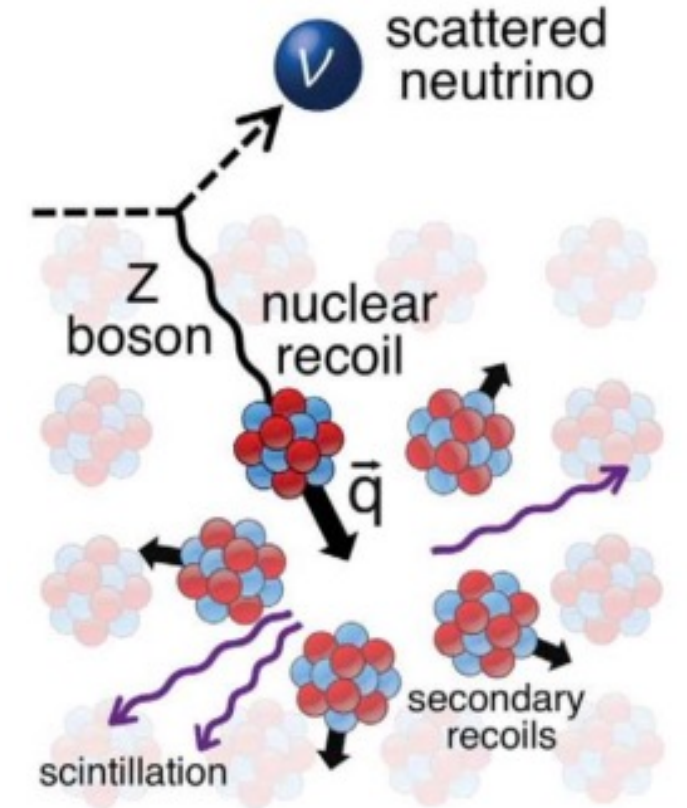
$$\frac{d\sigma}{dE_r} \propto Q_W^2 \simeq N^2$$

Why is CEvNS important?

- Largest neutrino-nucleus cross section at low energies
- Sensitive test of the Standard Model weak interaction
- Probe of light mediators and non-standard neutrino interactions
- Relevant to reactor monitoring and the neutrino background in dark-matter searches

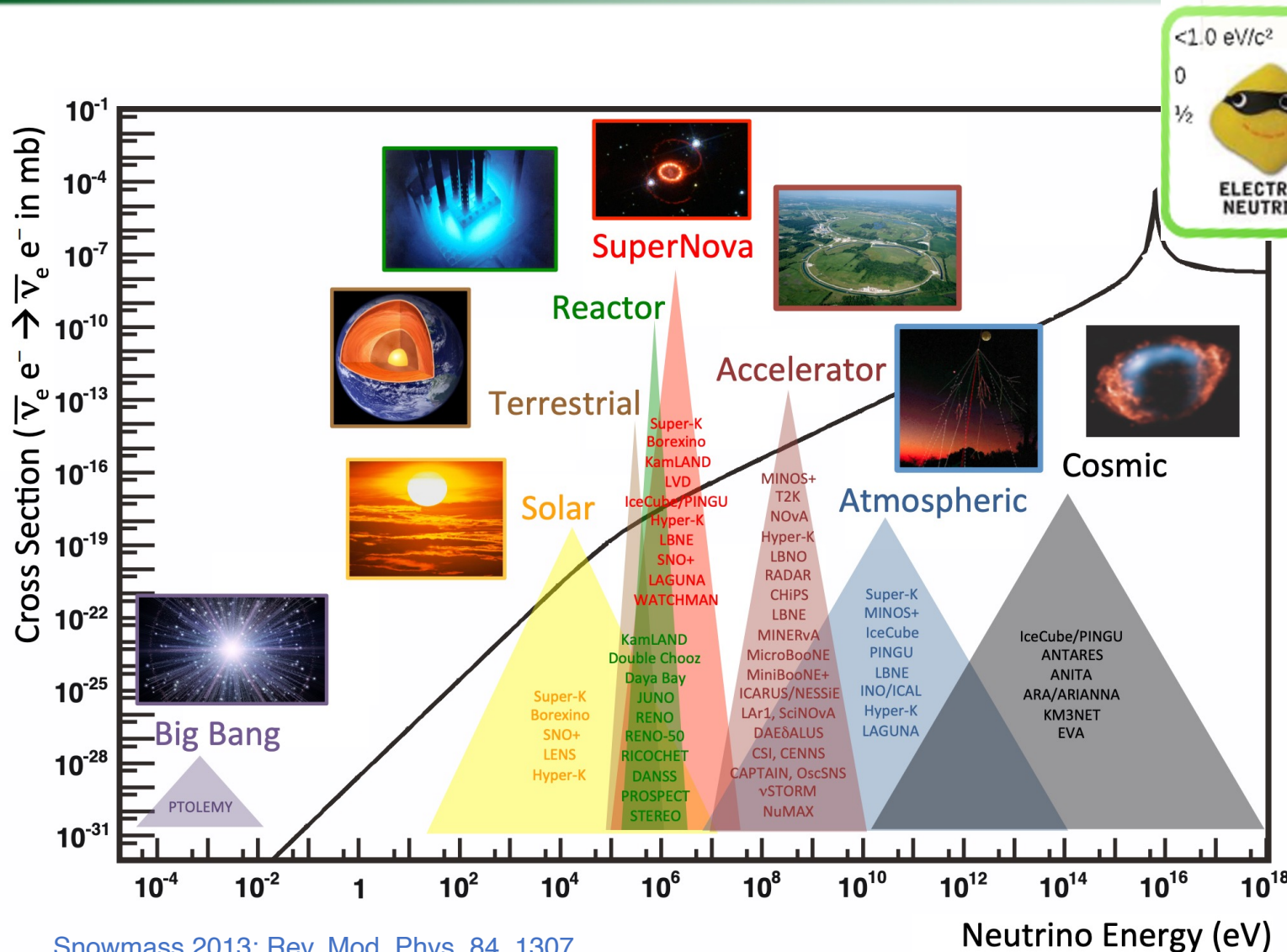
Experimental challenge

- Signal is a tiny nuclear recoil, typically at the sub-keV to few-keV scale.



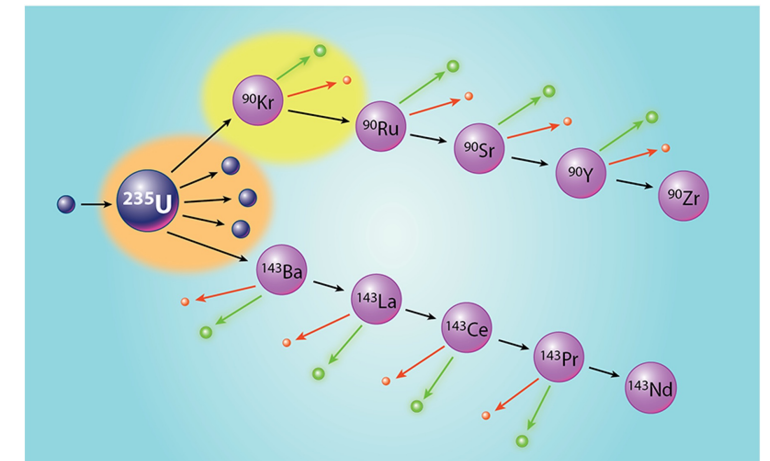
10.1126/science.aao0990

(Anti-)Neutrinos from Reactor



Snowmass 2013; Rev. Mod. Phys. 84, 1307

- Neutrinos come from many natural and artificial sources
- For CEvNS, reactor antineutrinos provide an intense, continuous, and controllable MeV-scale source.



Planned Experiment On-site



- Samen Nuclear Power Plant, Taizhou, China (share with RELICS)
- Reactor Power ~ 3.4 GW, Distance to Core ~ 25 m
- Neutrino flux $\sim 10^{13}$ $\nu/\text{cm}^2/\text{s}$

Detection Challenges



- **ER/NR discrimination:** reject X/ γ -ray and muon-induced electron recoils from the active volume
- **ν -NR / neutron-NR separation:** distinguish CEvNS NRs from reactor and atmospheric neutron recoils
- **Ionization quenching:** low-energy IQF strongly affects NR threshold and expected CEvNS yield

3D track topology and detector observables

Recoil directionality

Calibrated with IQF measurements

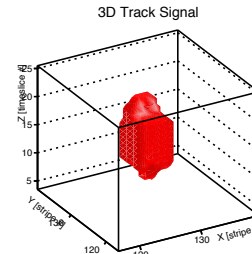
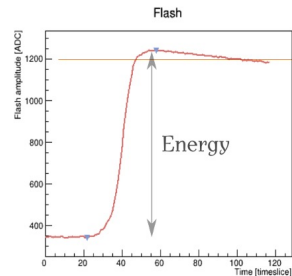
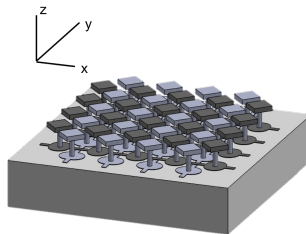
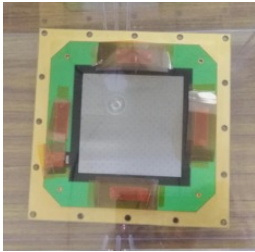
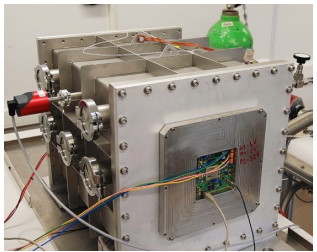
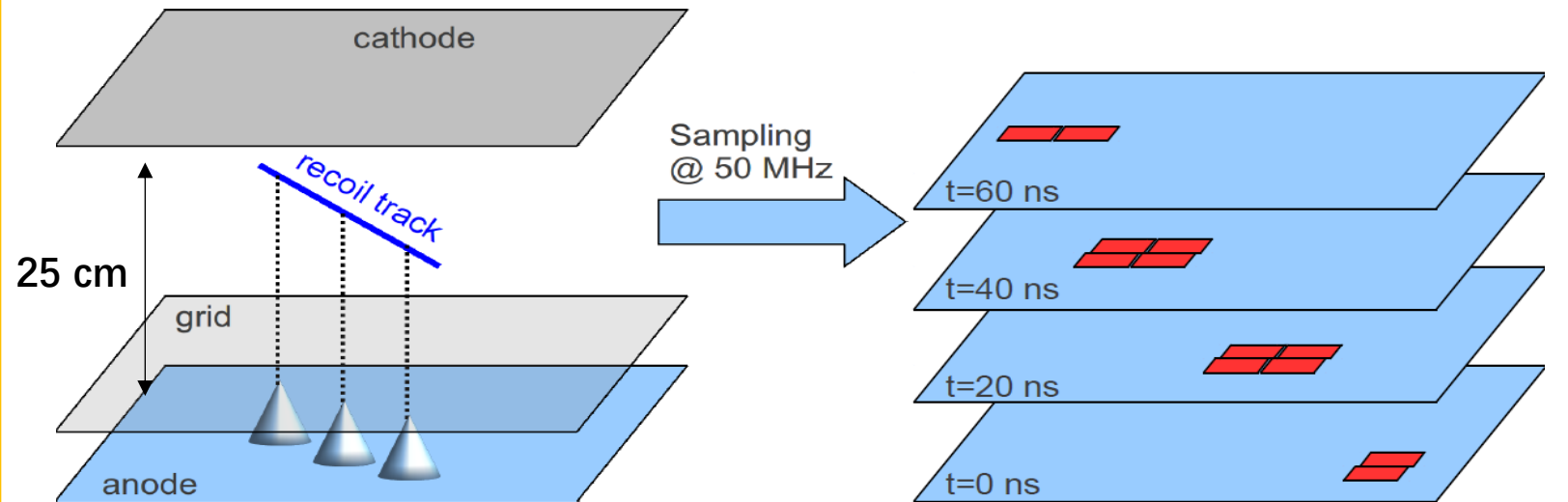
 **MIMAC detection strategy**

MIMAC Detector



MIMAC: Micro-TPC Matrix of Chambers = TPC + Micromegas

- ✓ TPC: Time Projection Chamber
- ✓ Micromegas: Micro Mesh Gaseous Structure



Advantages

- Energy Measurement
- 3D Track Reconstruction

Working Gas Mixture:

- $(80 - x) \% \text{ } ^4\text{He}$
- $(20 - y) \% \text{ CF}_4$
- $(x + y) \% \text{ C}_4\text{H}_{10}$

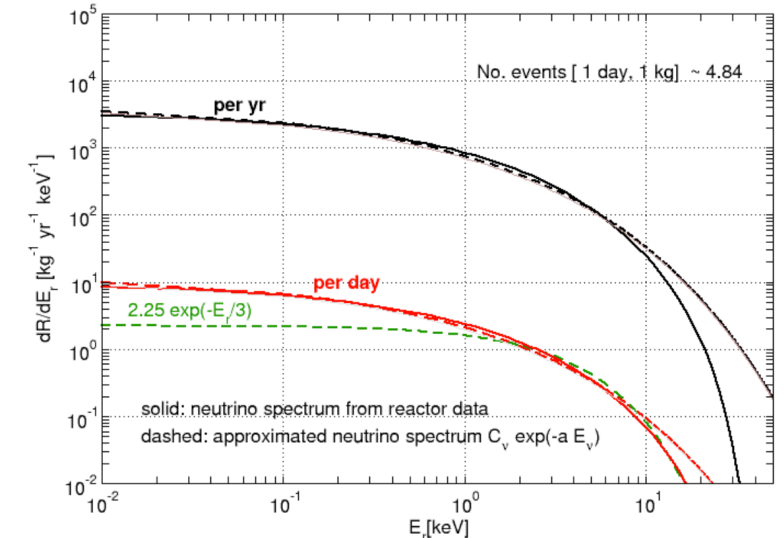
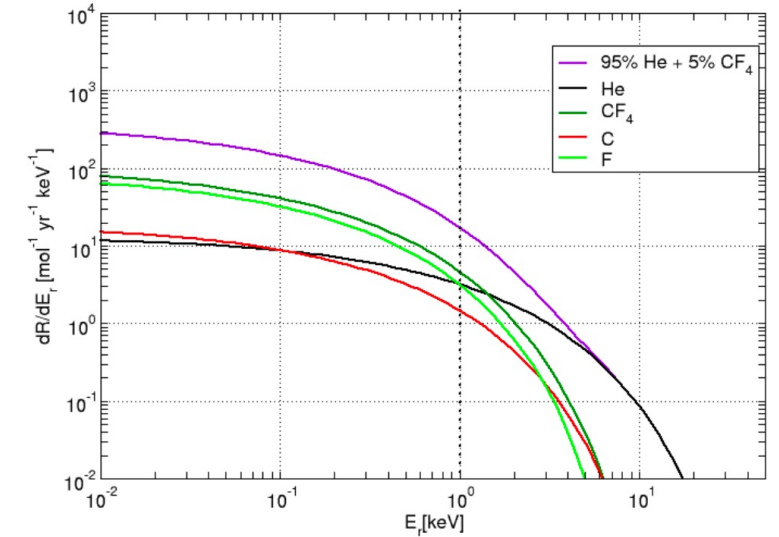
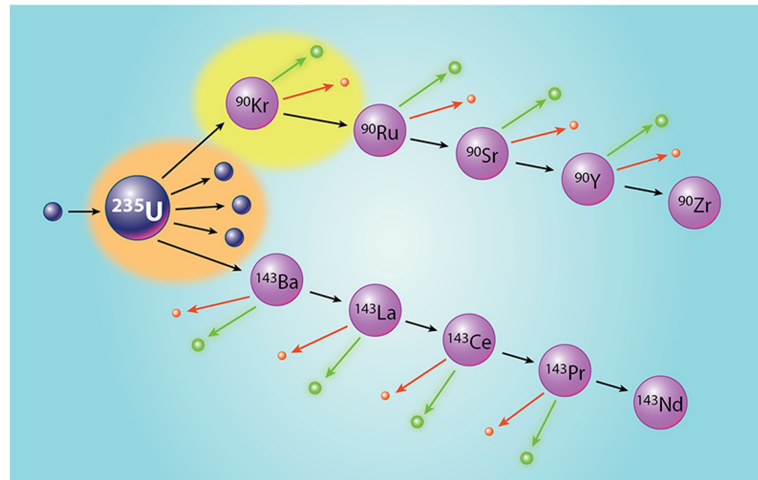
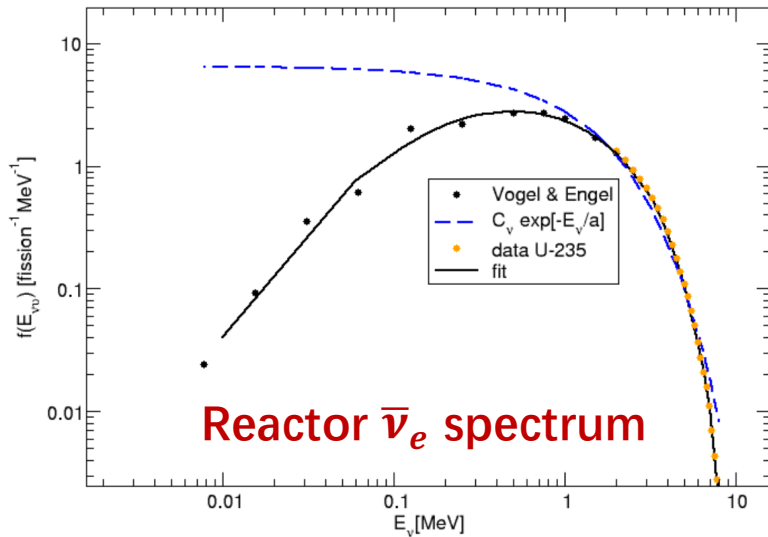
($x: \sim 0$; $y: \sim 1$)

- Goal: 16 bi-chambers to achieve 1 m^3 volume.

CEvNS Signals



- The empirical exponential approximation reproduces the main behavior well enough for rate estimates.
- Folding the spectrum with CEvNS cross section gives the He recoil spectrum, with high- E_r events strongly suppressed.
- In a 95% He + 5% CF_4 gas mixture: **He** dominates the signal, while **C** and **F** still add non-negligible events.
- Expected yield: **400-700 events/year** for 1 m^3 .

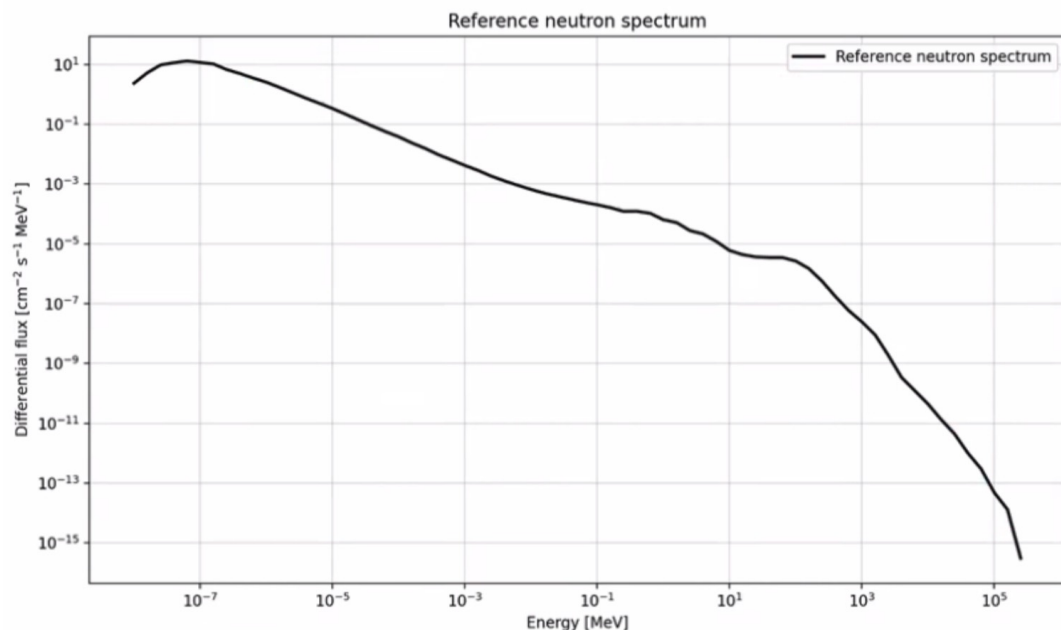


Dominant Background: Atmospheric Neutron

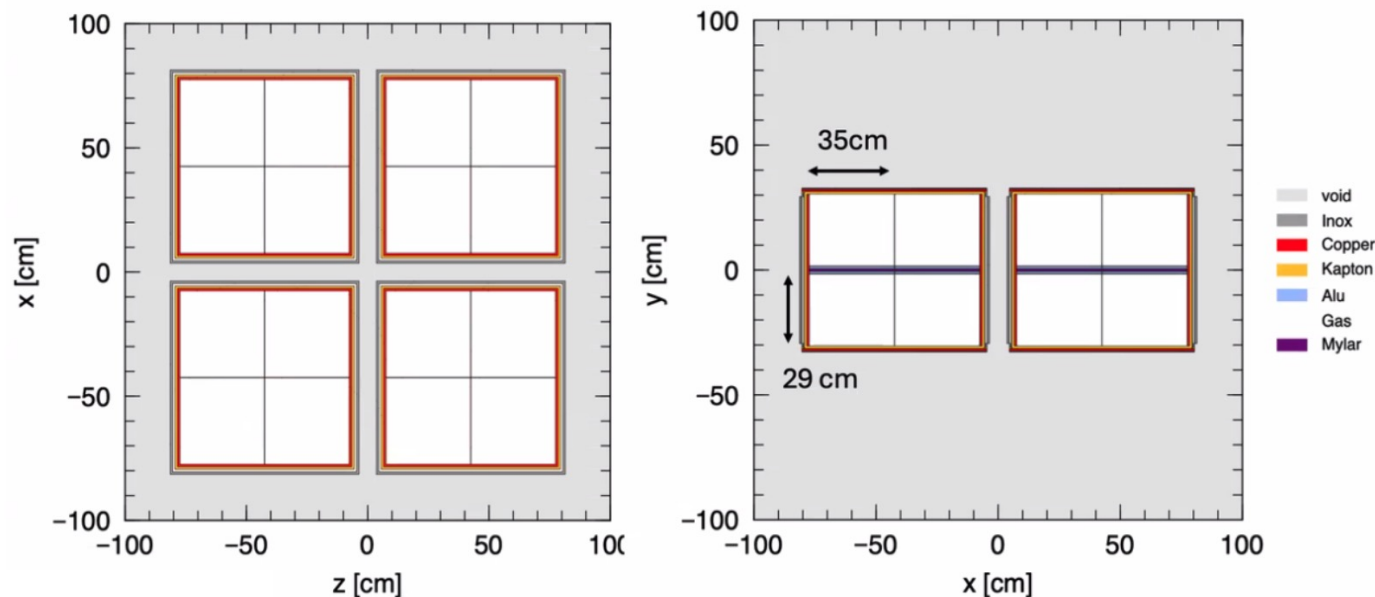


➤ Monte-Carlo Simulation (CRY + PHITS):

- CRY (Cosmic-ray Shower Library) → Generate atmospheric neutrons (w/ angular + energy distributions)
- PHITS (Particle and Heavy Ion Transport code System) → Detector response & background estimation & shielding optimization



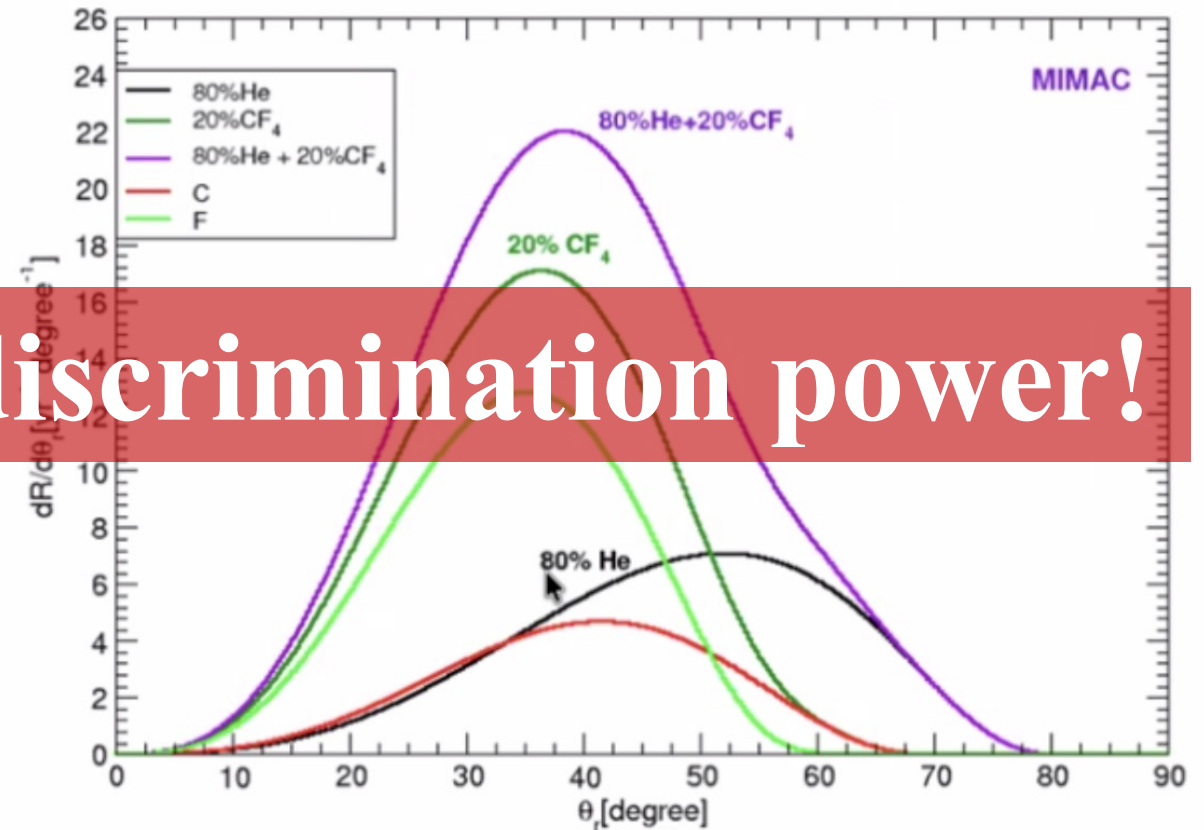
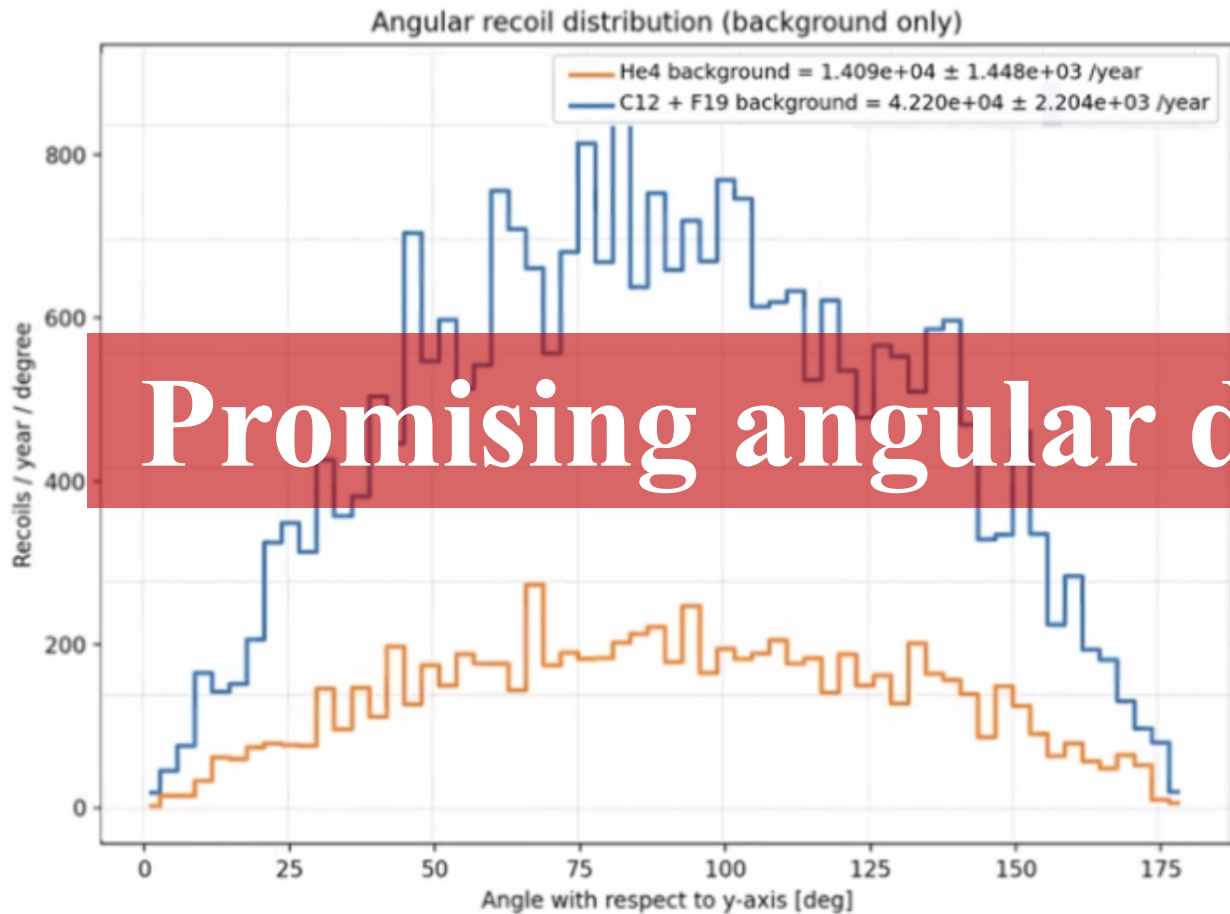
1 m³ detector 16 MIMAC bi-chambers



Dominant Background: Atmospheric Neutron



- Major advantage: Directionality!
- Preliminary results from MC: Background (L) vs. Signal (R)



Promising angular discrimination power!

How do we reduce such a significant background?



➤ Background study:

- With 1.5 m of HDPE shielding, $\sim 6 \times 10^4$ recoils/yr
- **Introduce hydrogen (H) as an internal neutron monitor**

Gas Mixture:

80 % ^4He + 19 % CF_4
+ 1 % C_4H_{10}

$$\frac{d\sigma}{dE_r} = \frac{G_F^2}{\pi} Q_W^2 M_A \left(1 - \frac{E_r}{E_r^{\max}} \right) F(q)^2 \propto Q_W^2 \simeq N^2$$

$$\text{where } Q_W = Z \left(2 \sin^2 \theta_W - \frac{1}{2} \right) + \frac{N}{2} \simeq 0.5N - 0.04Z$$

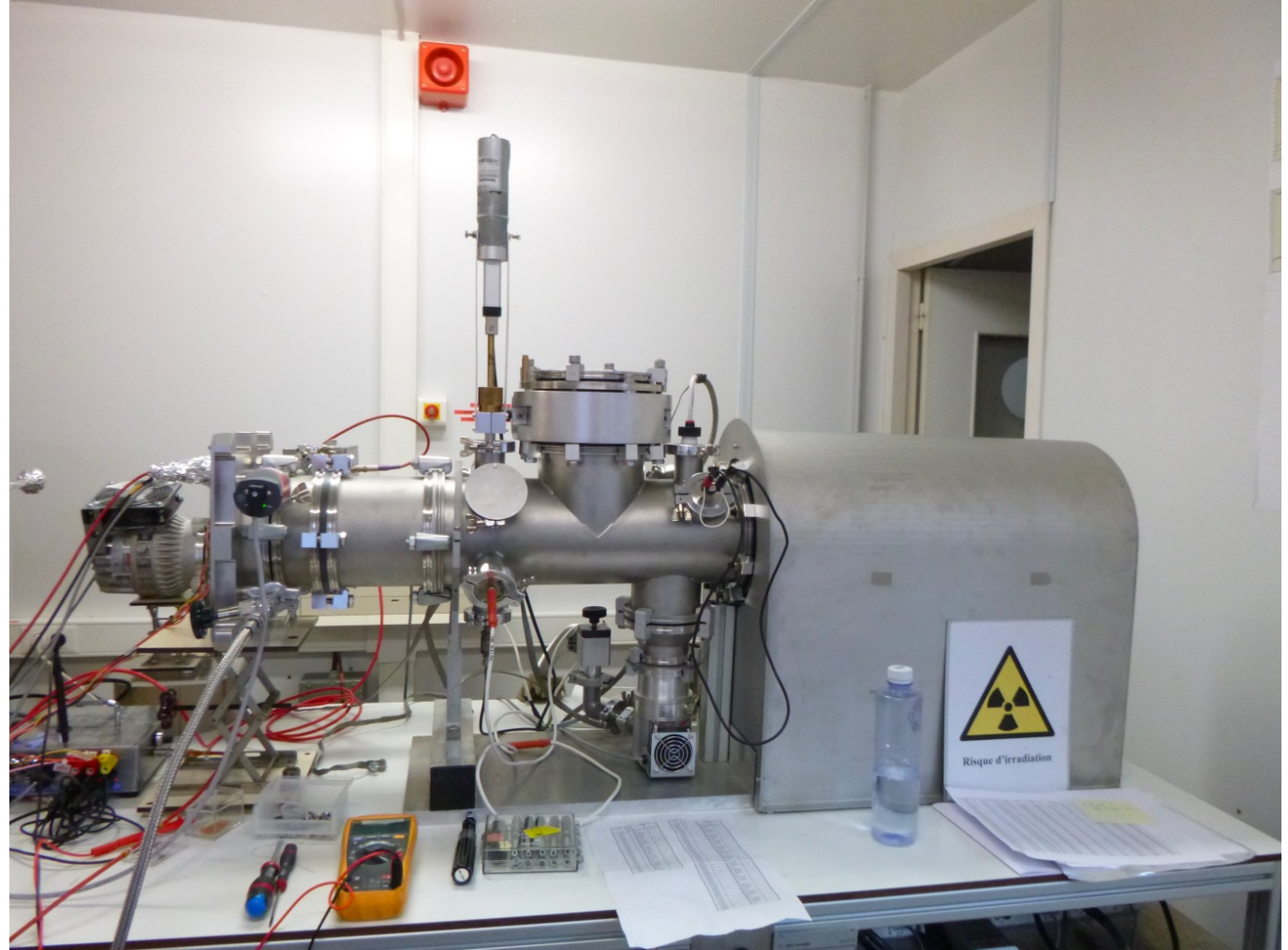
- Reactor-Off data could give direct validation

Ionization Quenching Measurement with COMIMAC



COMIMAC @ LPSC

- Portable table-top facility for low-energy detector calibration
- Delivers calibrated ion/electron beams up to **50 keV**
- Uses COMIC ECR source + Wien filter for particle selection
- Measures ion response and IQF in gas detectors
- Flexible choice of target particles, (including e^-)

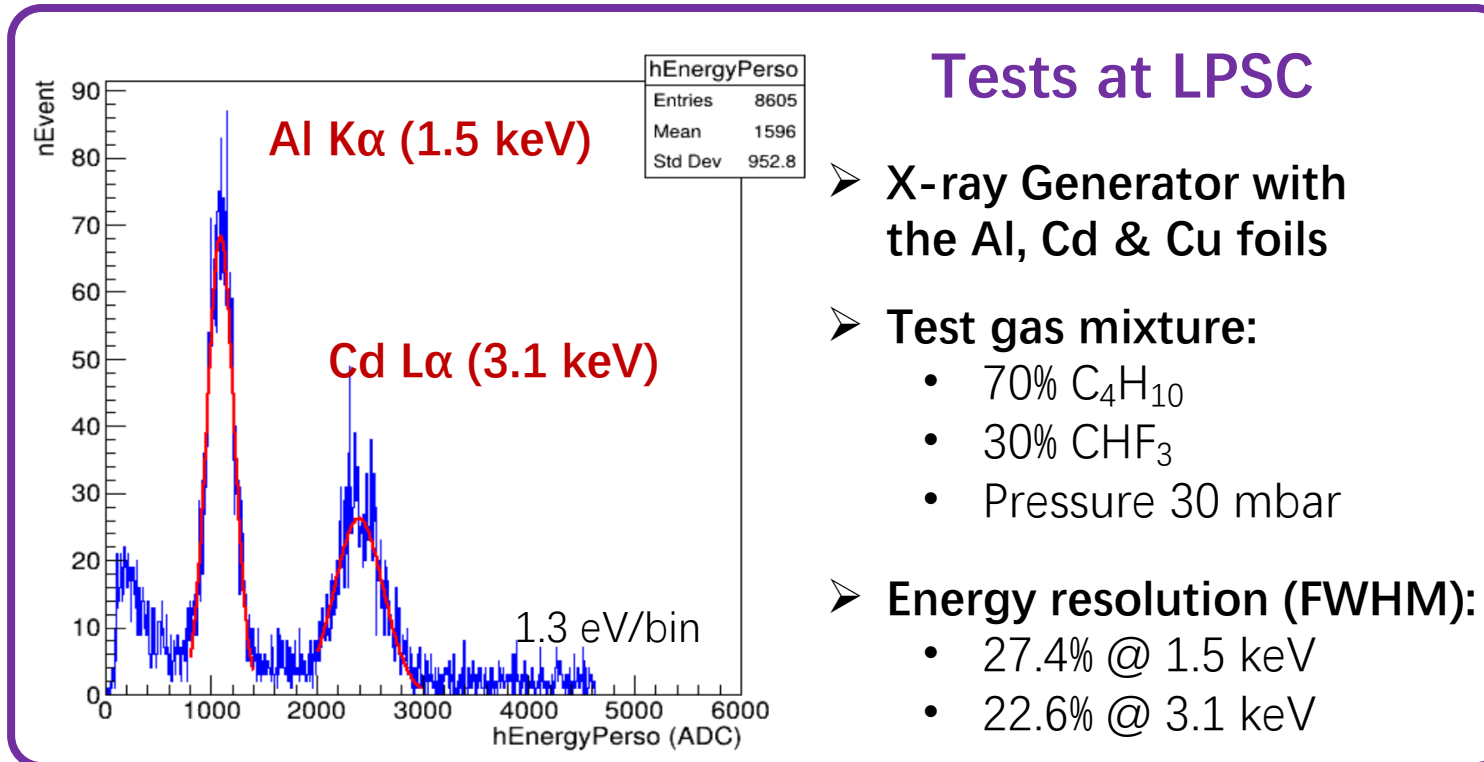


New Micromegas from USTC



➤ Several Key Parameters:

- Sensitive area: $35 \times 35 \text{ cm}^2$;
- Avalanche gas gap: $512 \text{ }\mu\text{m}$;
- Strip pitch: $200 \text{ }\mu\text{m}$;
- Withstands HV: $\sim 1800 \text{ V}$ @air pressure.



Tests at LPSC

- X-ray Generator with the Al, Cd & Cu foils
- Test gas mixture:
 - 70% C₄H₁₀
 - 30% CHF₃
 - Pressure 30 mbar
- Energy resolution (FWHM):
 - 27.4% @ 1.5 keV
 - 22.6% @ 3.1 keV



Credit: Zhang Zhiyong @USTC

New Micromegas from USTC

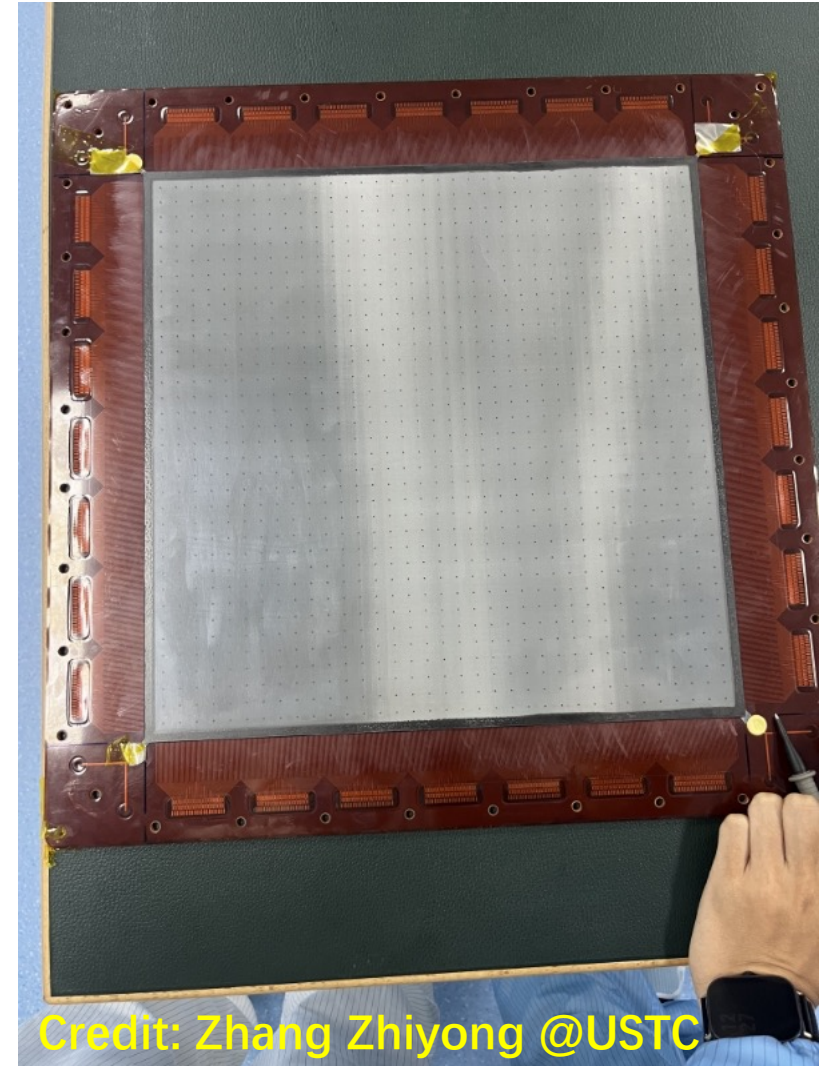
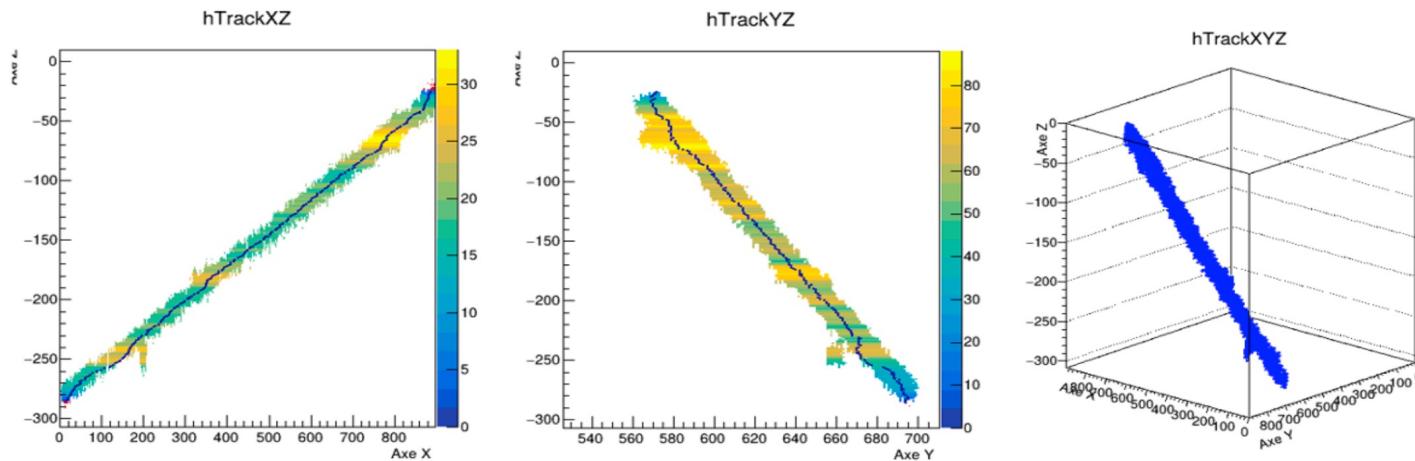


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Tests at LPSC

➤ Clear 3D alpha track reconstruction!



Credit: Zhang Zhiyong @USTC

Summary

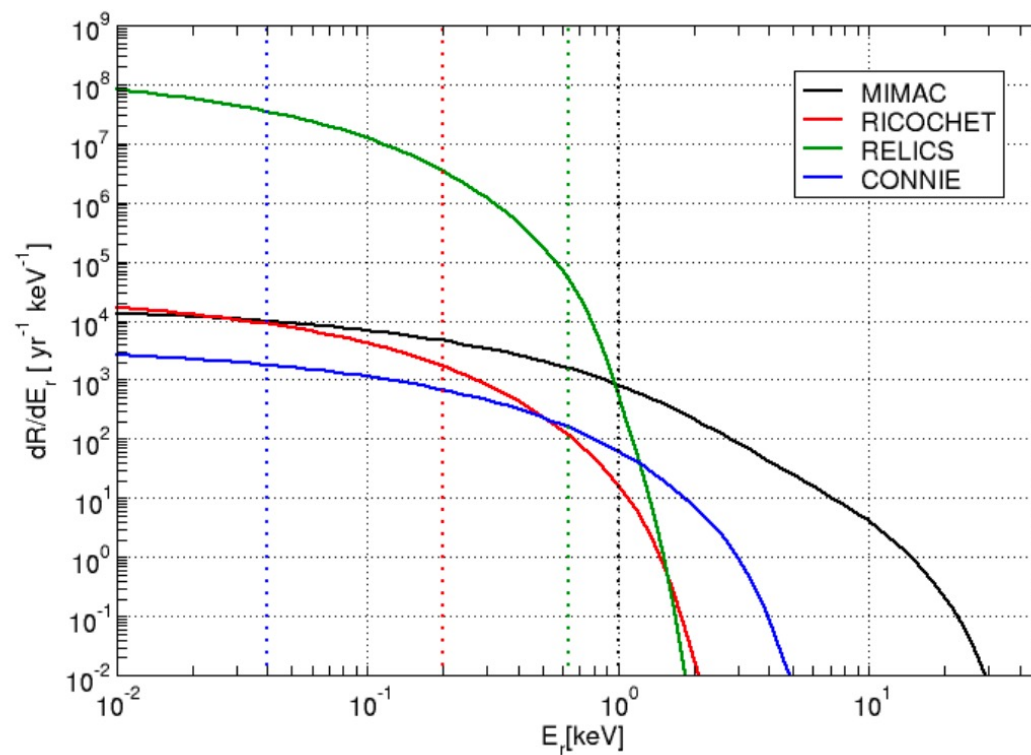
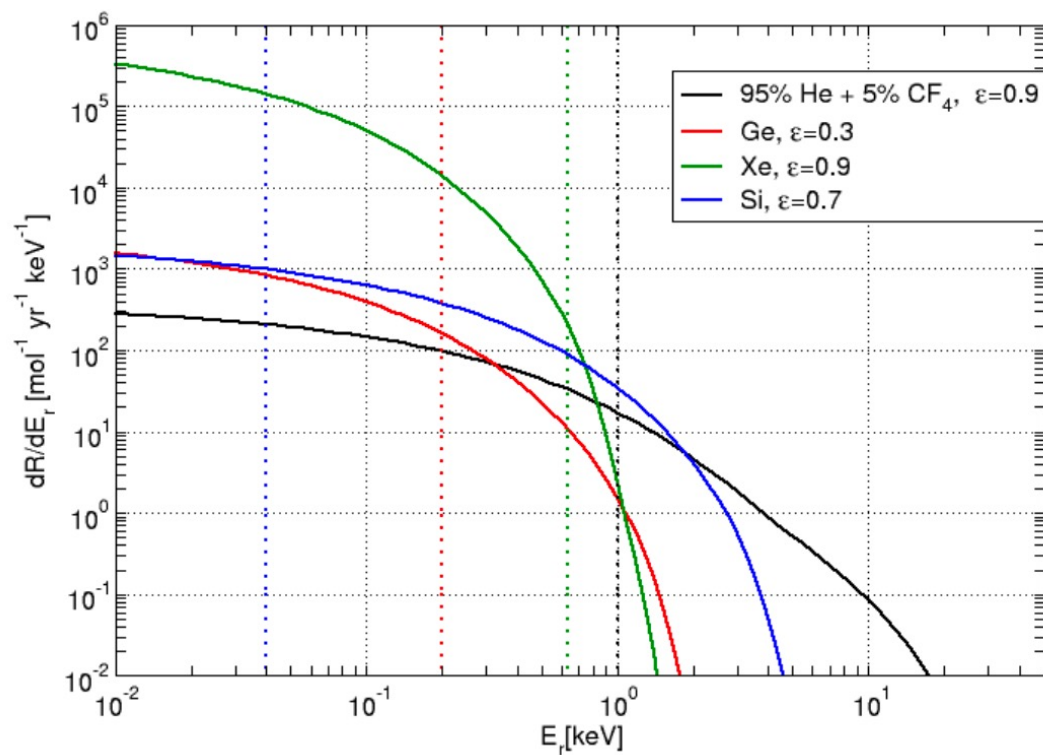


Thanks for your attention!

Backups



➤ Differential rate per mol per year



Backups



- 25 cm drift distance for each chamber

