

# ***ZTF Camera Paper Status***

**Marie Aubert**

ZTF-FRANCE 28/01/2026

# Paper plan

Paper status : Draft in progress

<b>1. Introduction</b> .....		To be written
<b>2. Instrument</b> .....		To be proofread
1. Telescope		
2. Camera		
3. Filters		
<b>3. Detrending pipeline</b> .....		Drafted
1. Pipeline Overview		
2. Overscan correction		
3. Linearity correction		
4. From calibration raw data to master calibration images .....		
1. Master bias		
2. Master flat		
5. Detrending and science image		
1. Defringing		
6. Forced aperture photometry and catalogues		
<b>4. Star-flat analysis</b> .....		Drafting
		General presentation & model.

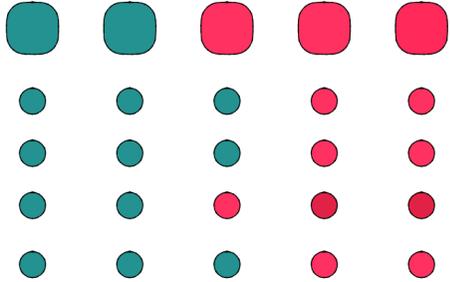
# Paper plan

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## 5. Synthetic Filters

- 1. Instrument model
- 2. tCBP campaign
- 3. Filter synthesis (pupil stitching ?)
- 4. Comparaison to Gaia

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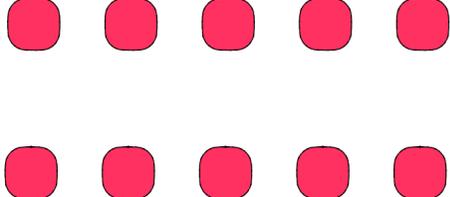


Drafted

## 6. On-sky performance (?)

→ Light curve & Ubercal mostly

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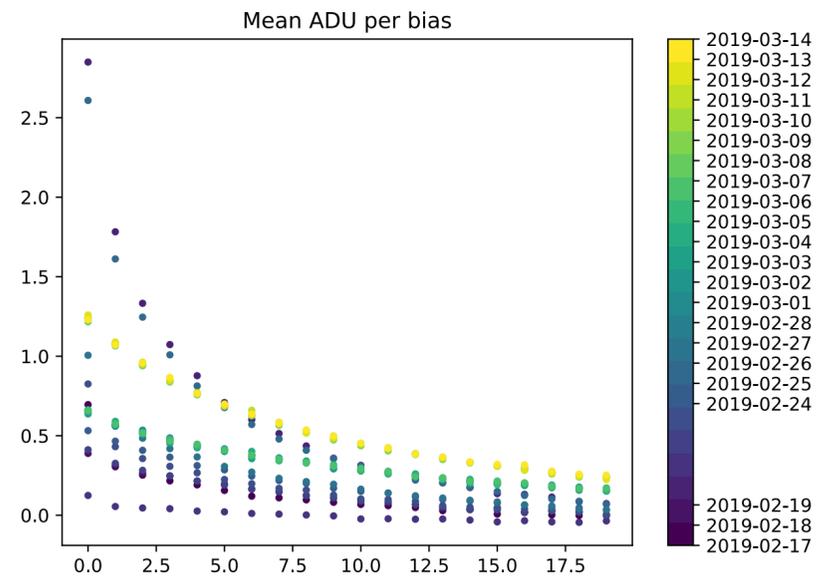
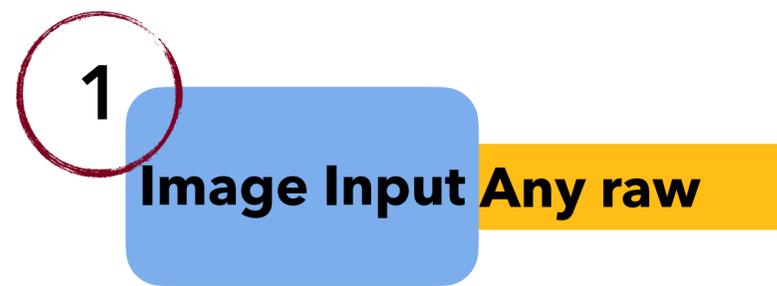
Unclear if this will disappear

To be written

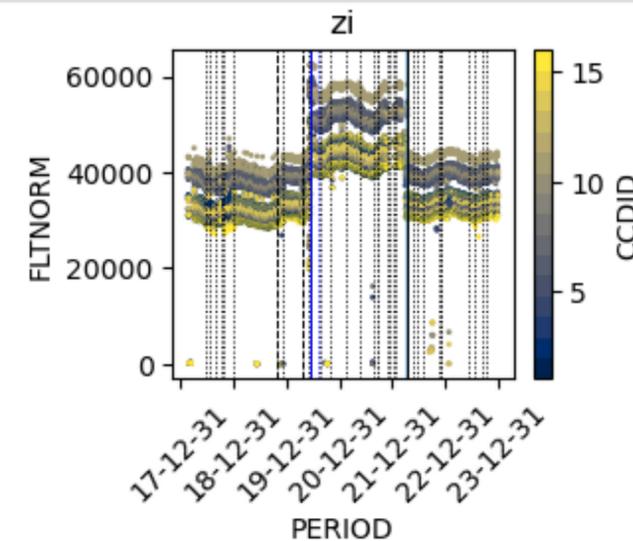
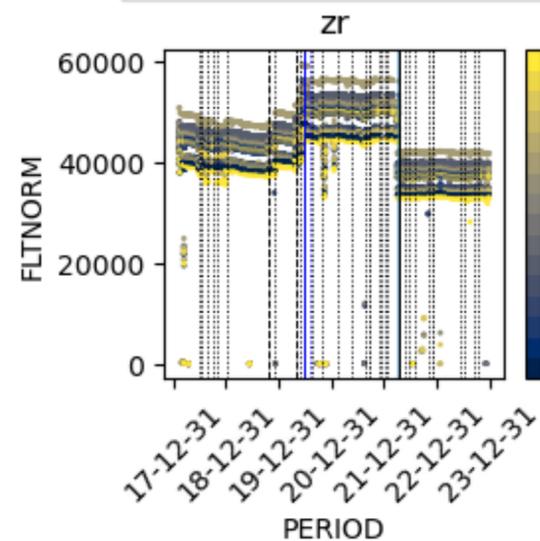
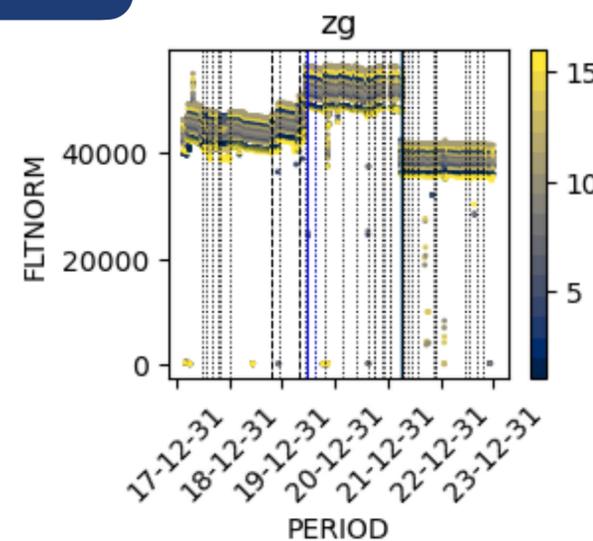
## 7. Conclusion

# Detrending pipeline

- Mostly drafted.
- To do : Add plots. (Bias dilation, Flat-field selection)



Dilation example



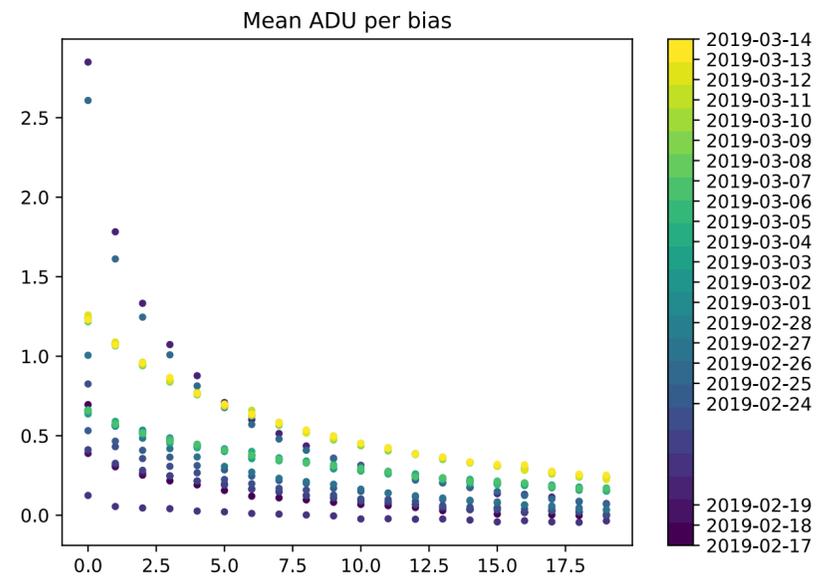
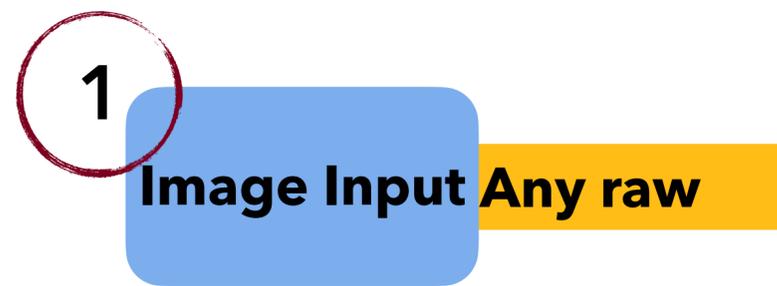
Outlier rejection

**UNCLEAR:** Overscan correction (5 last pix)

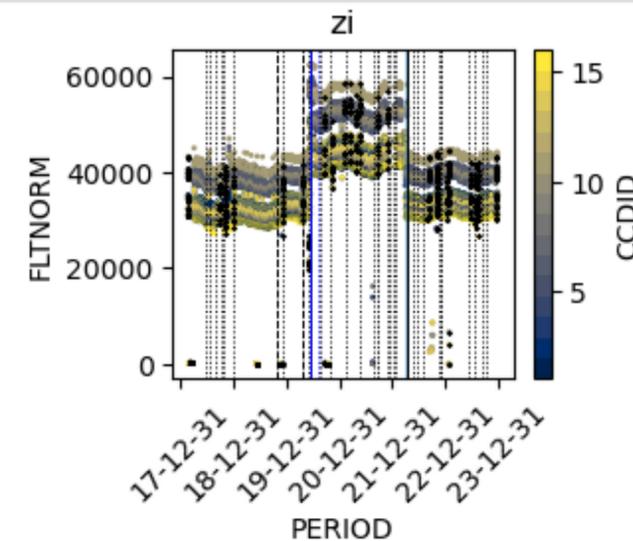
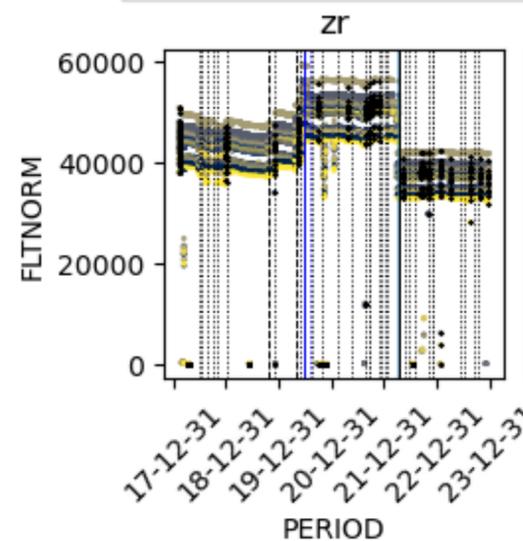
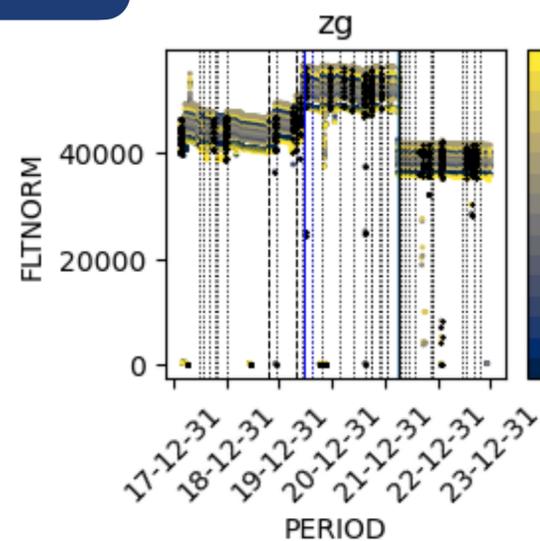
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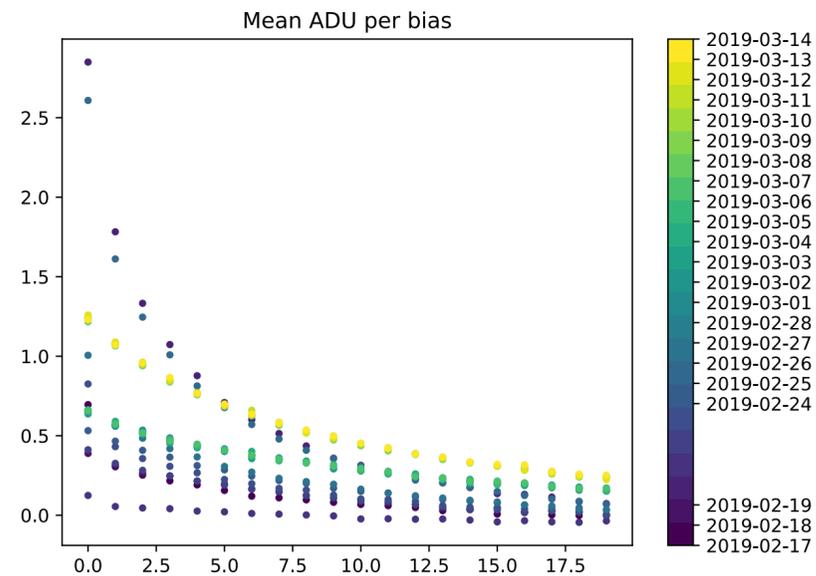
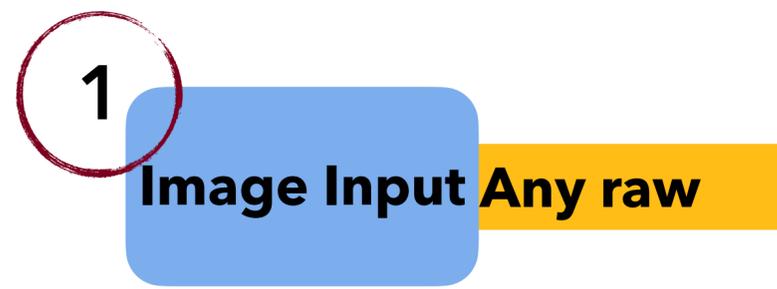
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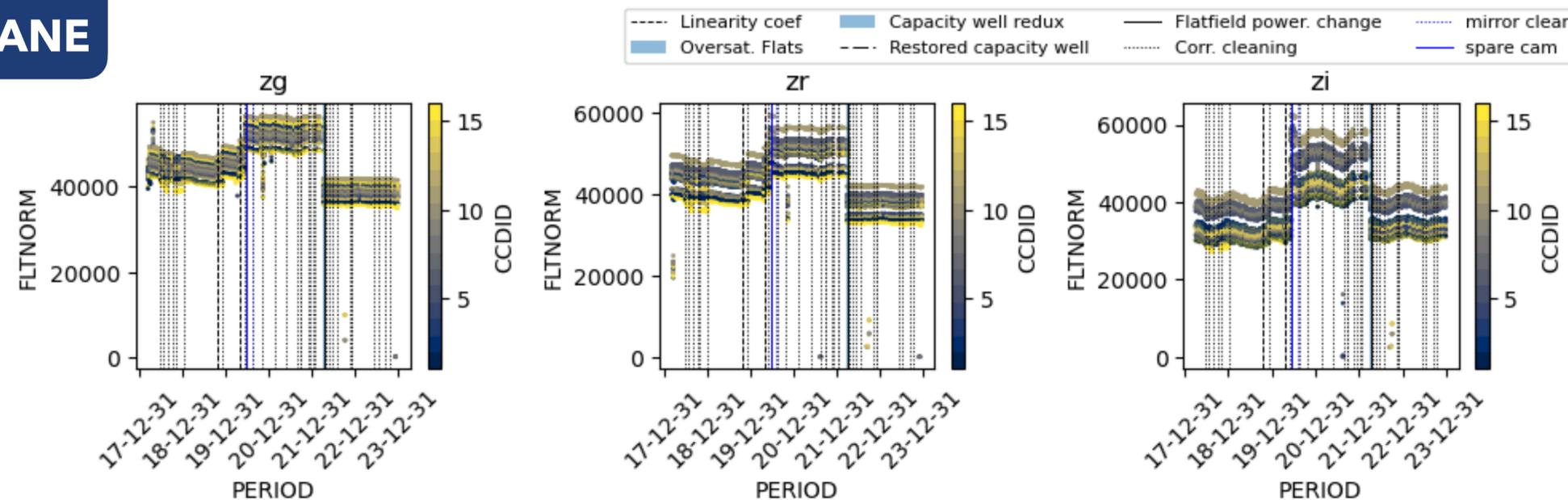
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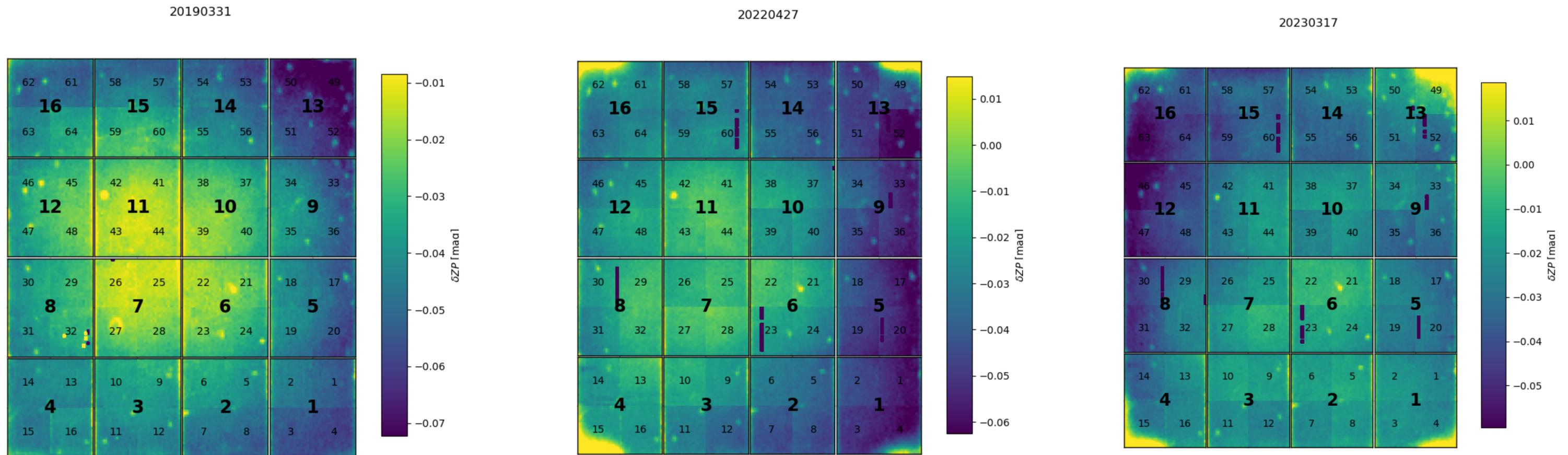
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# Starflat Analysis

- Finish interpreting the results and plug into paper (in a table).
- Issue : *Overlap with Ubercal to be defined.*

- Ran individually on each star flat acquisition.  
Should it be synthesized per year or per filter only ?
- Might need one last round of processing
- Should compare with IPAC starflat ?

G band.

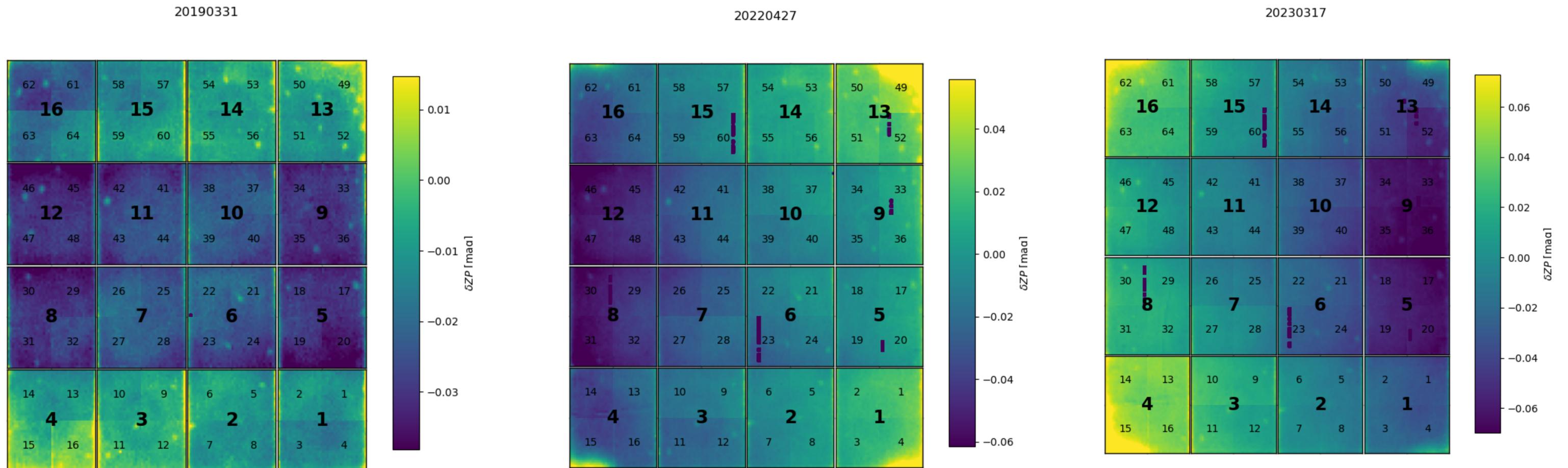


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R band.



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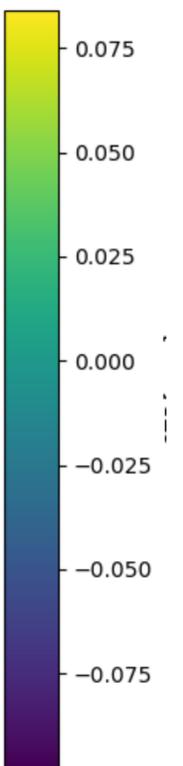
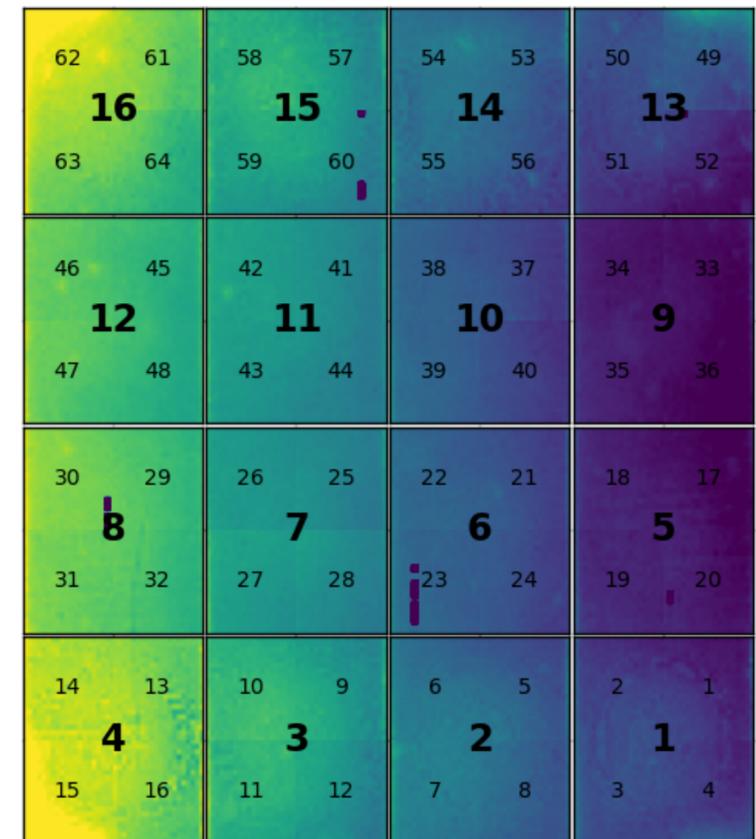
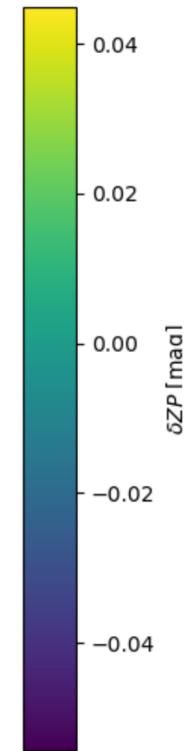
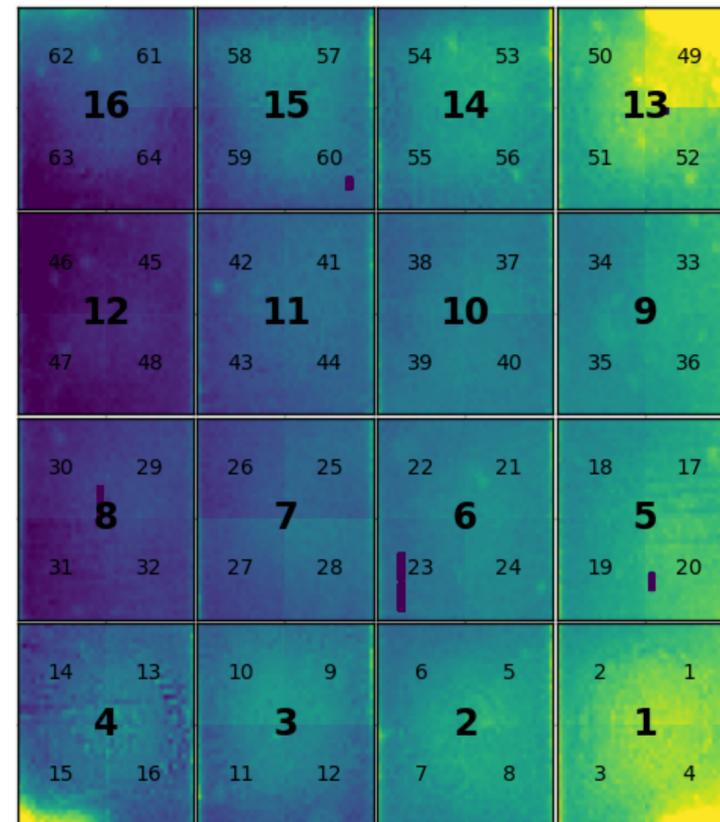
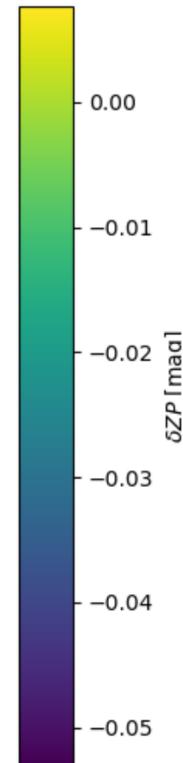
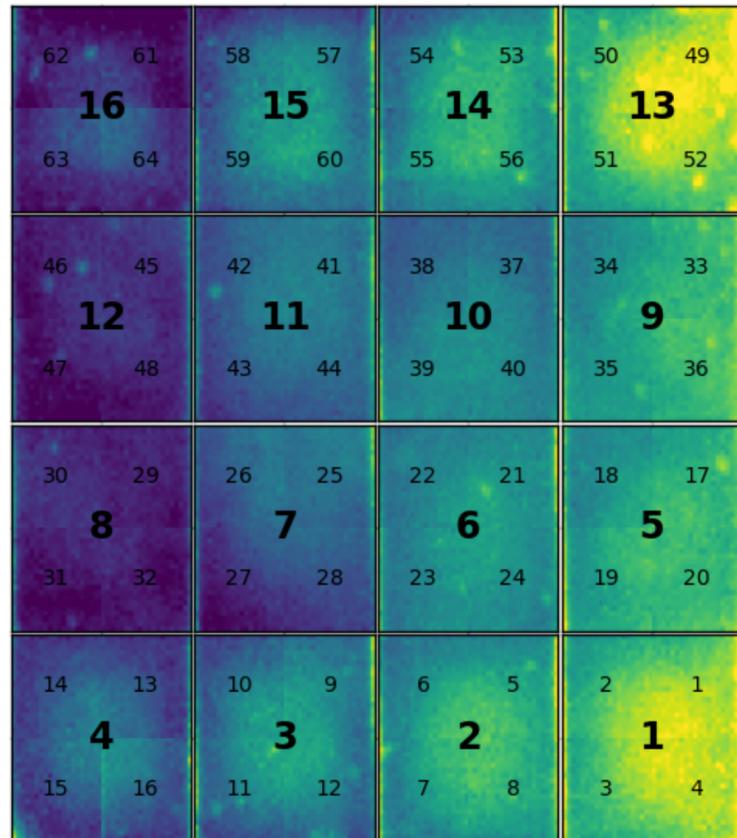
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I band.

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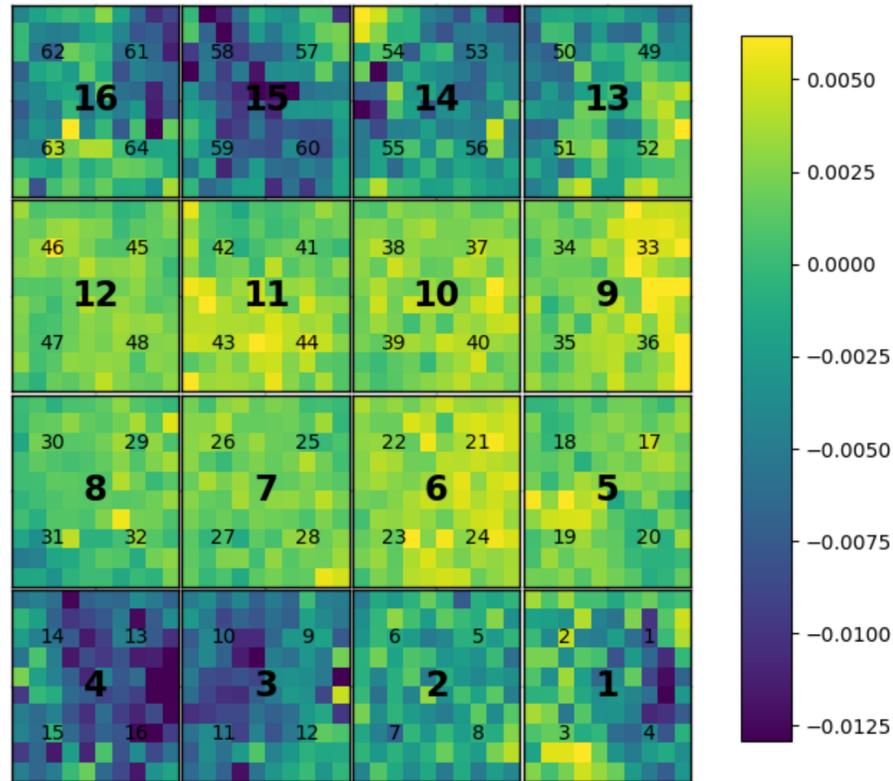
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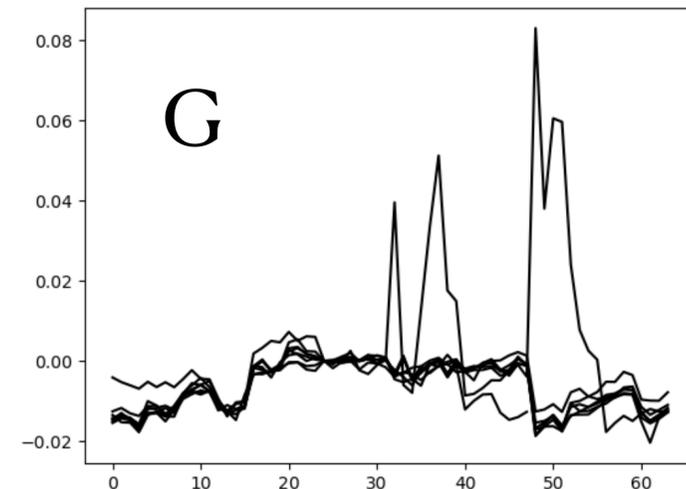
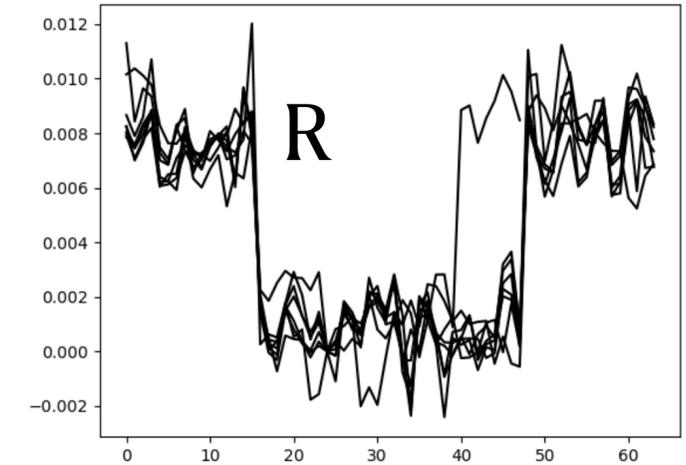
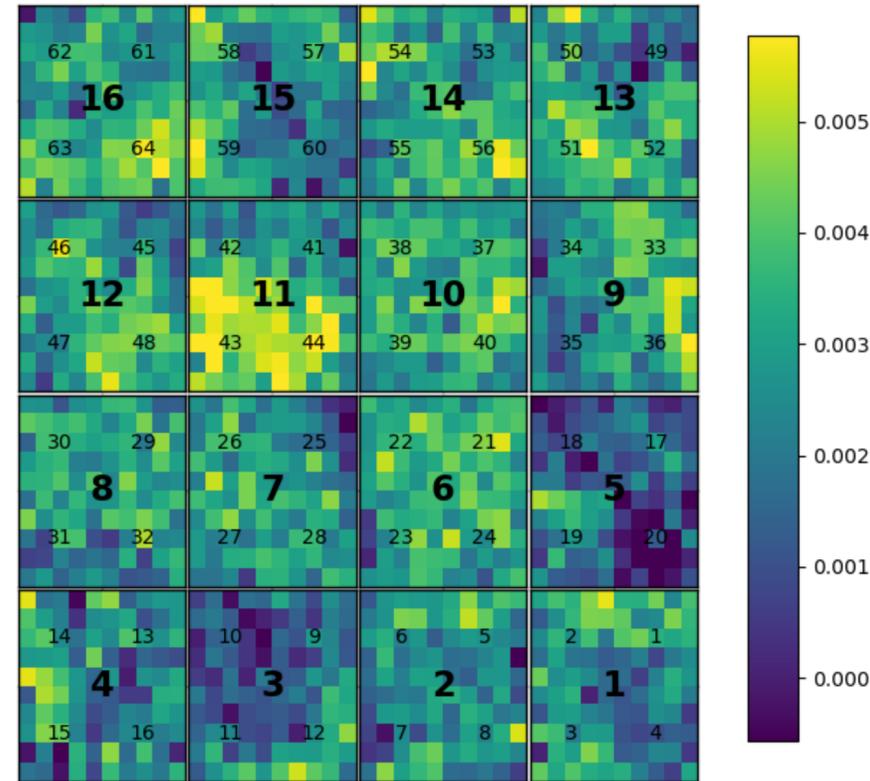
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COLOR

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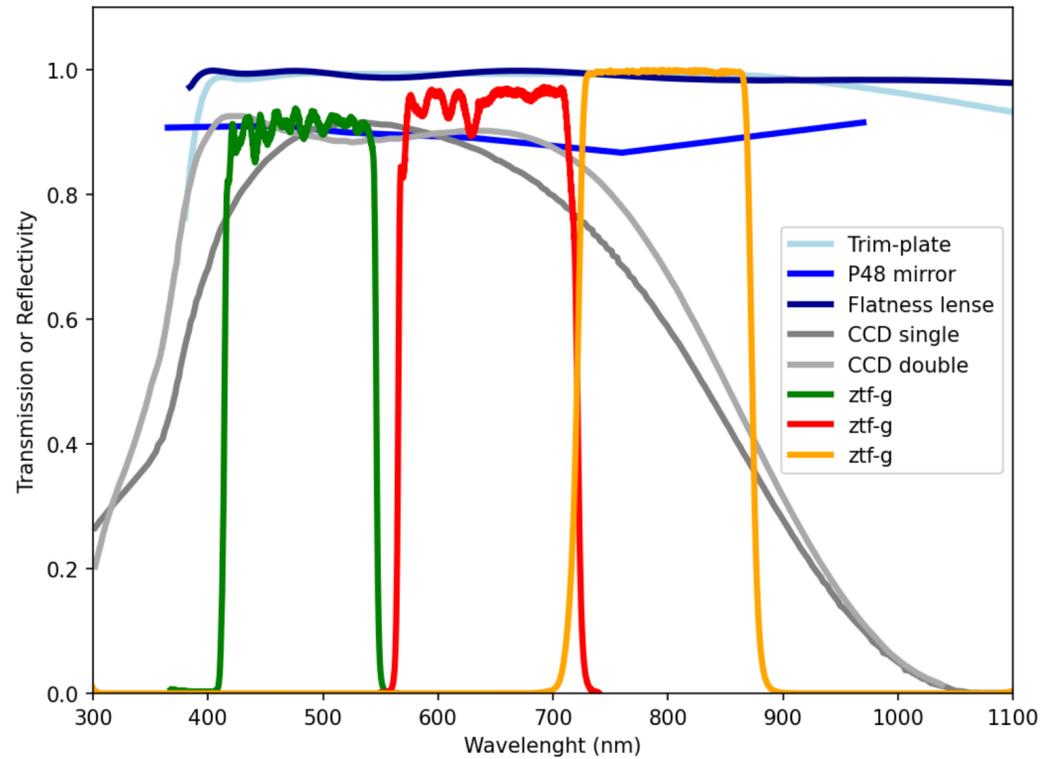
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Fringe correction change color in starflats.

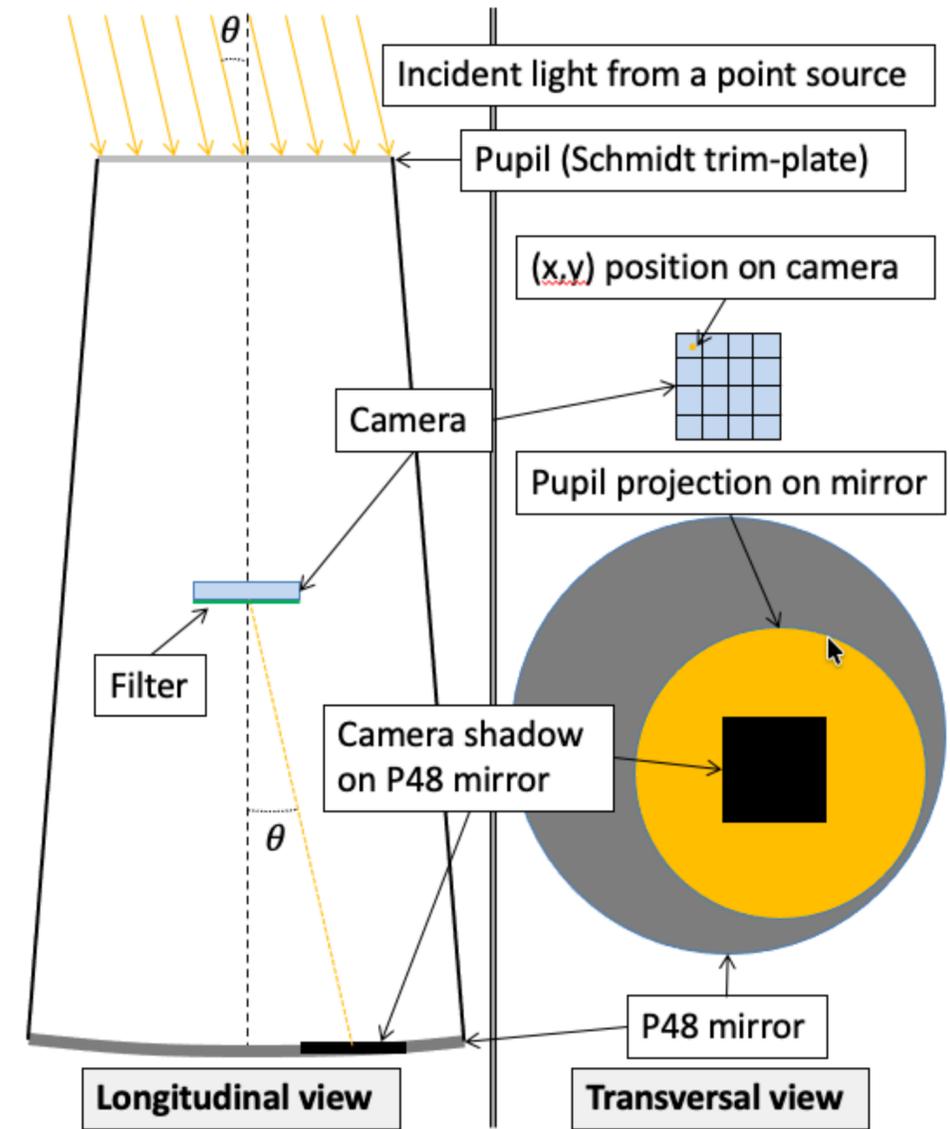
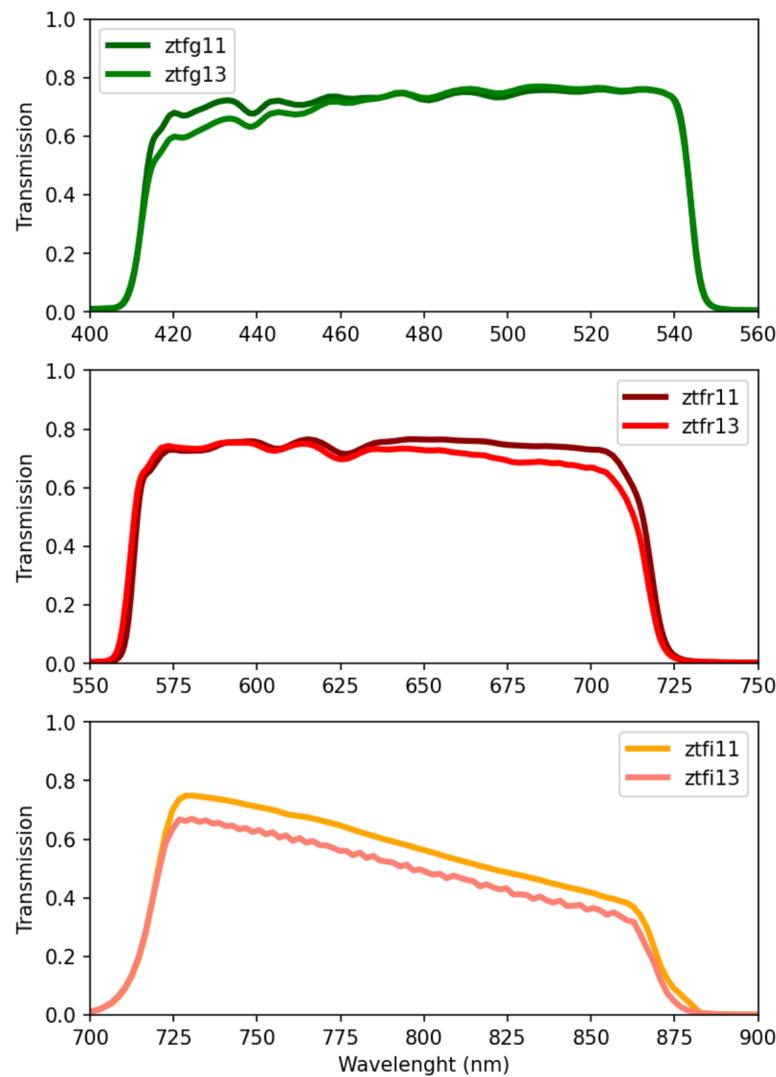
# Pupil stitching model

- New filter to be passed to BBF. (tCBP paper WIP)
- Do we need to applied a filtering of remaining fringing effects on i-filter ?



Input of instrument model

## Transmissions



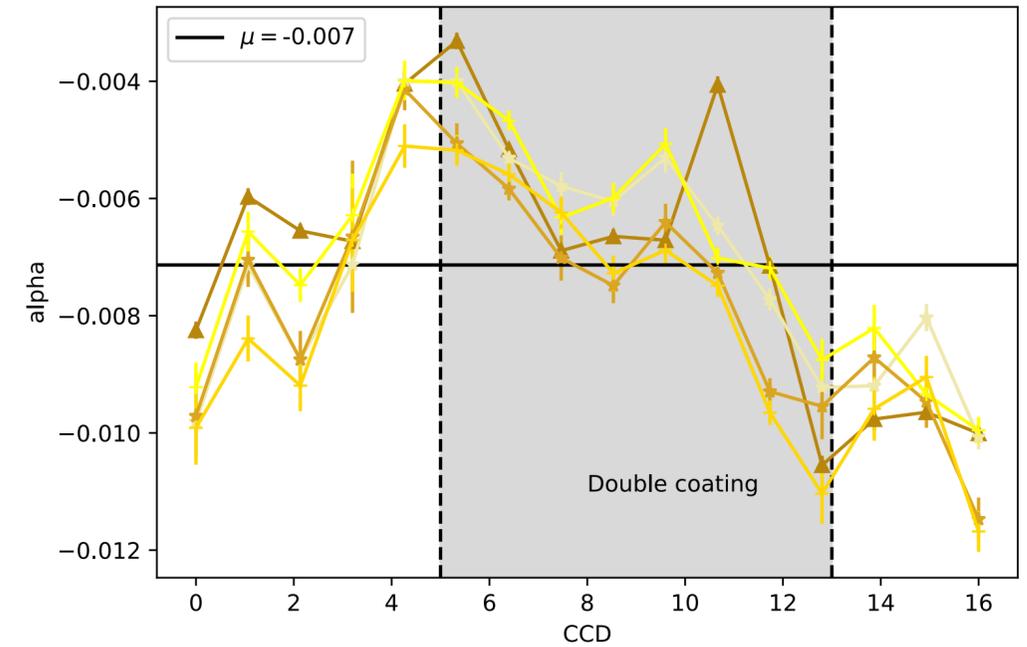
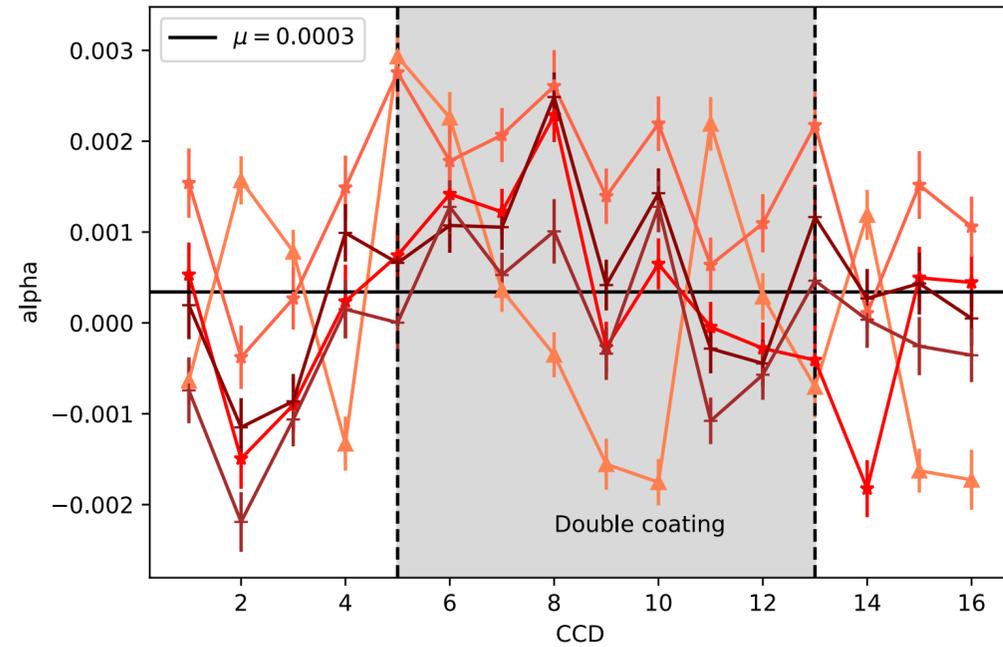
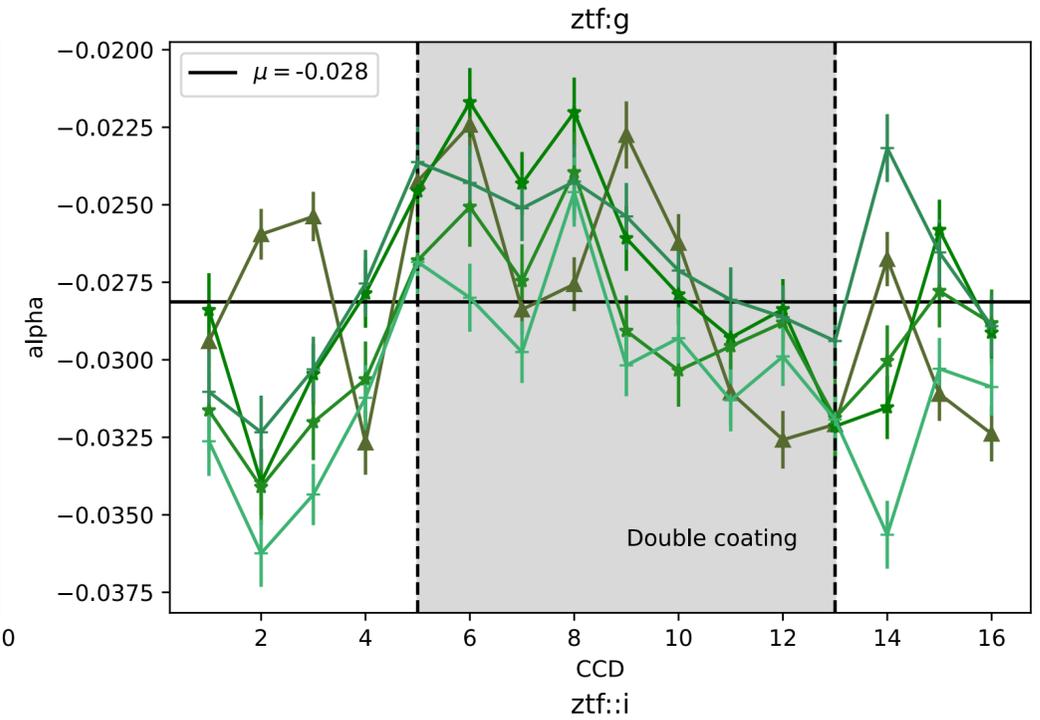
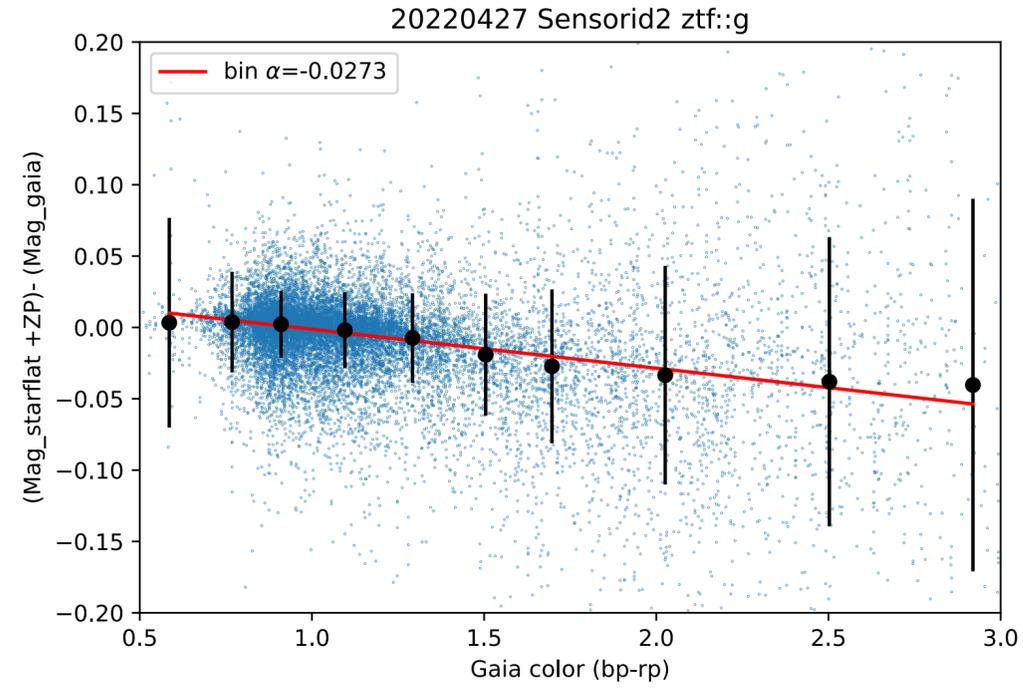
Instrument Model

# Synthetic Photometry

- BBF description missing

→ Try with higher density :

→ Same question as star flat use concatenated sequences / filter ?



# Conclusion / Notes

## What else

- Core of the paper *mostly* written. Some tweaking due to data analysis & interpretation.
- Proof-reading & harmonization.
- Conclusion & Intro to be written :  
Why & how we can do better in some aspects (re-training fringes etc) for DR<sub>3</sub><sup>+</sup>.

# On the to-do list

## Or questions

- Detrending :
  - Discussion & finalize.
  - Document the pipeline for « easy » usage.

- Starflat analysis :
  - Interpretation & write.
  - Year-round or individual
  - Versus IPAC or not ?
- Ubercal overlap to be discussed.

- Filter synthesis :
  - Refine fringe correction ?

*Blank space to add other things I didn't think about after discussion :*