

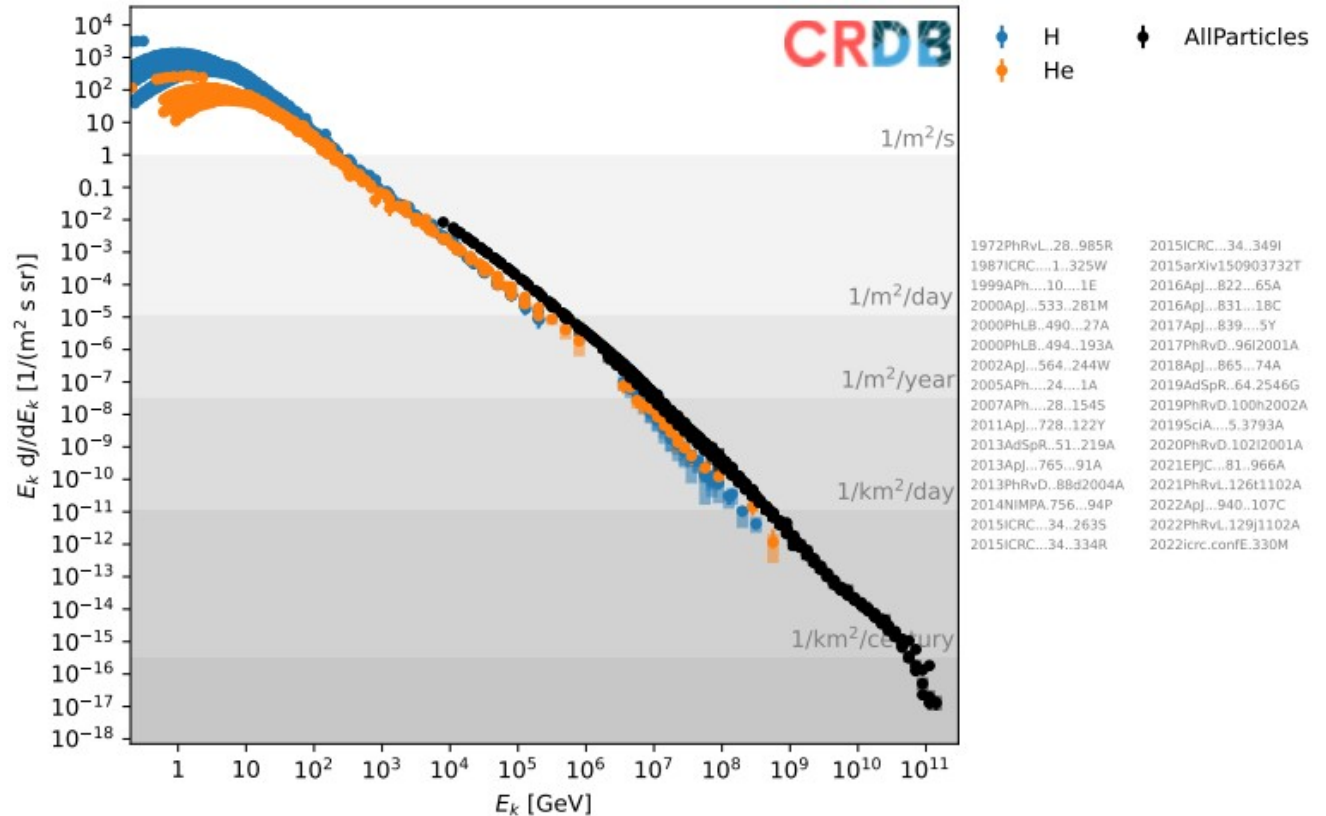
KPAP : Field line & particle transport in synthetic turbulence

Matthieu Bouchet with Y. Génolini, A. Marcowith, S. S.
Cerri and P. Mertsch



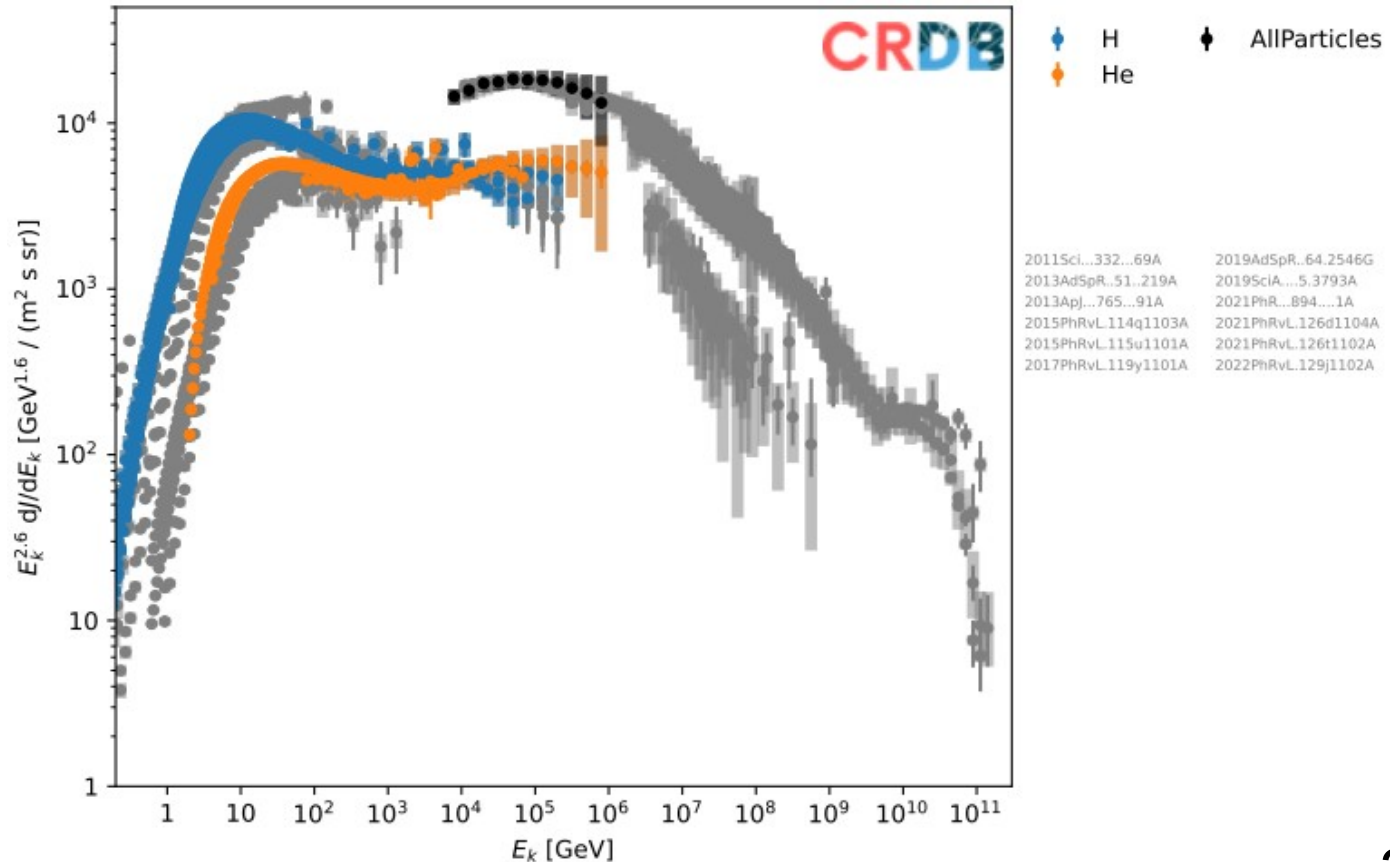
Introduction: phenomenology

- Composition
 - Protons and He nuclei (99%), heavy nuclei and antiparticles
- Observations
 - Energy spectrum follows a power law : injection, diffusive transport



Introduction: phenomenology

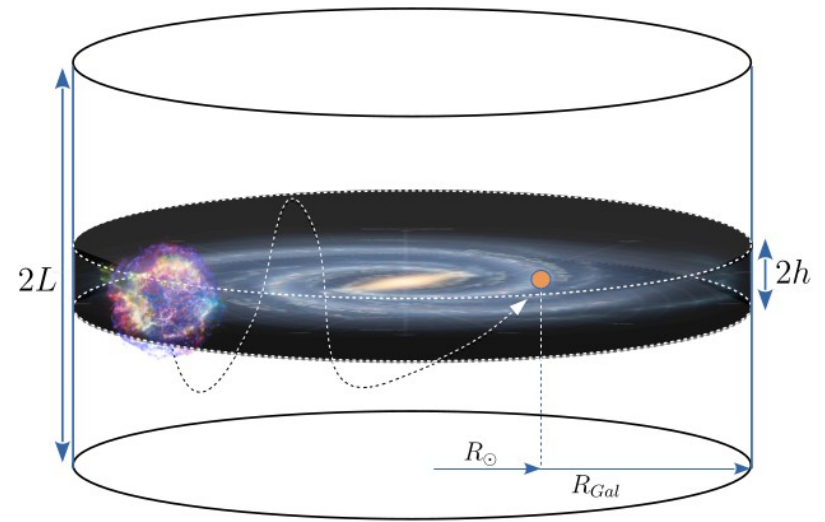
- Composition
 - Protons and He nuclei (99%), heavy nuclei and antiparticles
- Observations
 - Energy spectrum follows a power law : injection, diffusive transport
 - Features : knee, ankle >> change of sources or diffusion regime
- *Galactic confinement*
 - Spallation
 - Radioisotopes
 - Secondary-to-primary ratio



Microphysics of CR transport

Spherical cow galactic diffusion model
→ Quasi-Linear-Theory (QLT)

- Cylindrical diffusion box
- Homogeneous & isotropic diffusion tensor
- Pheno model reproduces the data but does not explain : rigidity dependence of diffusion transport, CR small scales anisotropies, CR spectral hardening towards GC...

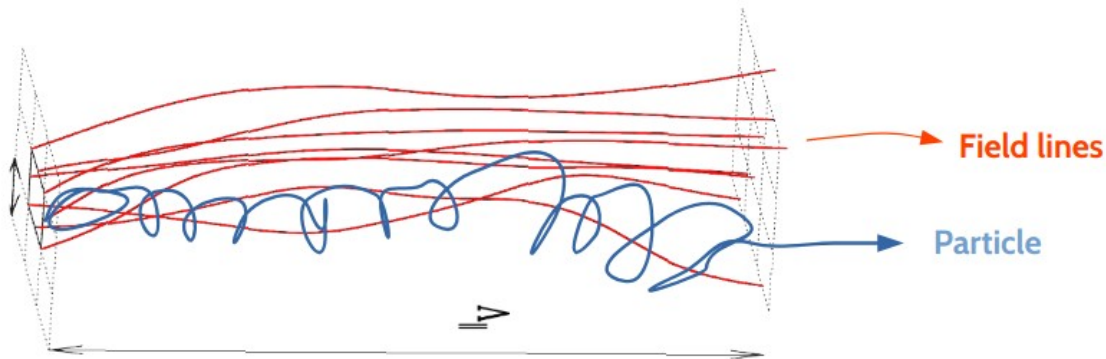
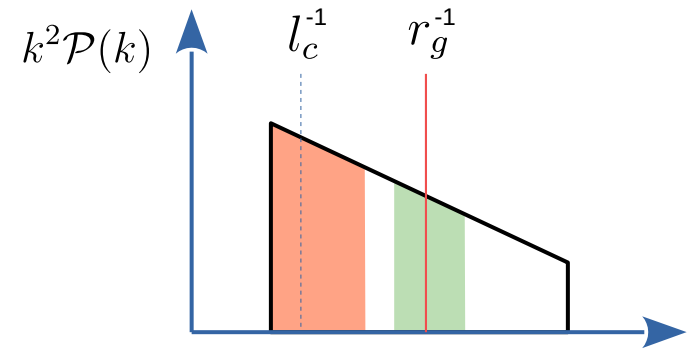


Microphysics of CR transport

Physics of perpendicular transport

$$D = \begin{bmatrix} D_{\parallel} & 0 & 0 \\ 0 & D_{\perp} & 0 \\ 0 & 0 & D_{\perp} \end{bmatrix}$$

- Parallel transport to the bckg field B driven by pitch angle scattering
- Perpendicular transport driven by :
 - Field line random walk
 - Parallel transport
 - Transverse complexity



Studied in synthetic turbulence simulations

Synthetic turbulence simulations

How to generate synthetic turbulence ?

Stored in memory

Main pros

Main cons

Harmonic method *(see Giacalone 1994)*

$$\delta \mathbf{B}(\mathbf{r}) = \sqrt{2} \sum_{n=0}^{N-1} A_n \hat{\boldsymbol{\xi}}_n \cos(k_n \hat{\mathbf{k}}_n \cdot \mathbf{r} + \beta_n)$$

$$\{A_n, \hat{\boldsymbol{\xi}}_n, \hat{\mathbf{k}}_n, \beta_n\}$$

Dynamical range

Computing time

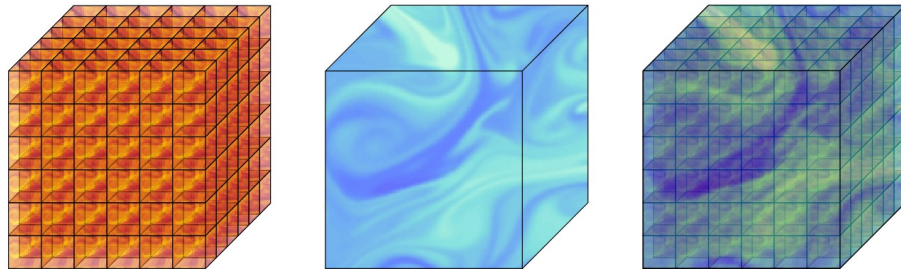
Nested grid method *(see Mertsch 2020)*

$$\delta B_j(\mathbf{r}_{n_1, n_2, n_3}) = \frac{1}{N^3} \sum_{l_1, l_2, l_3=0}^{N-1} e^{-2\pi i(\frac{l_1 n_1}{N} + \frac{l_2 n_2}{N} + \frac{l_3 n_3}{N})} \delta \tilde{B}_j^{l_1, l_2, l_3}$$

$$\delta B_j(\mathbf{r}_{n_1, n_2, n_3})$$

Computing time

Memory



Total B field : $\mathbf{B} = \delta \mathbf{B} + B_0 \mathbf{z}$
 Kolmogorov spectrum \rightarrow

Turb. level: $\eta = \frac{\delta B^2}{\delta B^2 + B_0^2}$



We study field lines and particle transport in this turbulence

Synthetic polarized turbulence simulations

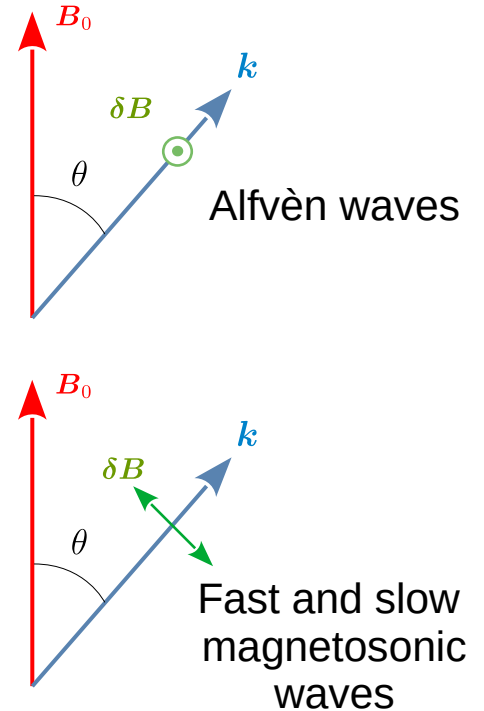
- Polarized synthetic turbulence
 - From ideal linearised MHD equations → eigenmodes

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

$$\rho \left(\frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v} \right) = -\nabla p + \frac{1}{\mu_0} (\nabla \times \mathbf{B}) \times \mathbf{B}$$

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{v} \times \mathbf{B})$$

$$\nabla \cdot \mathbf{B} = 0$$



- We generate synthetic isotropic polarised turbulence

$$\delta \mathbf{B}(\mathbf{r}) = \sum_{n=0}^{N-1} A_n \hat{\xi}_n \cos \left[k_n \hat{\mathbf{k}}_n \cdot \mathbf{r} + \beta_n \right]$$

3 cases :
Polarisation state

ALF :
Alfvén-like

MAG:
Magnetosonic-like

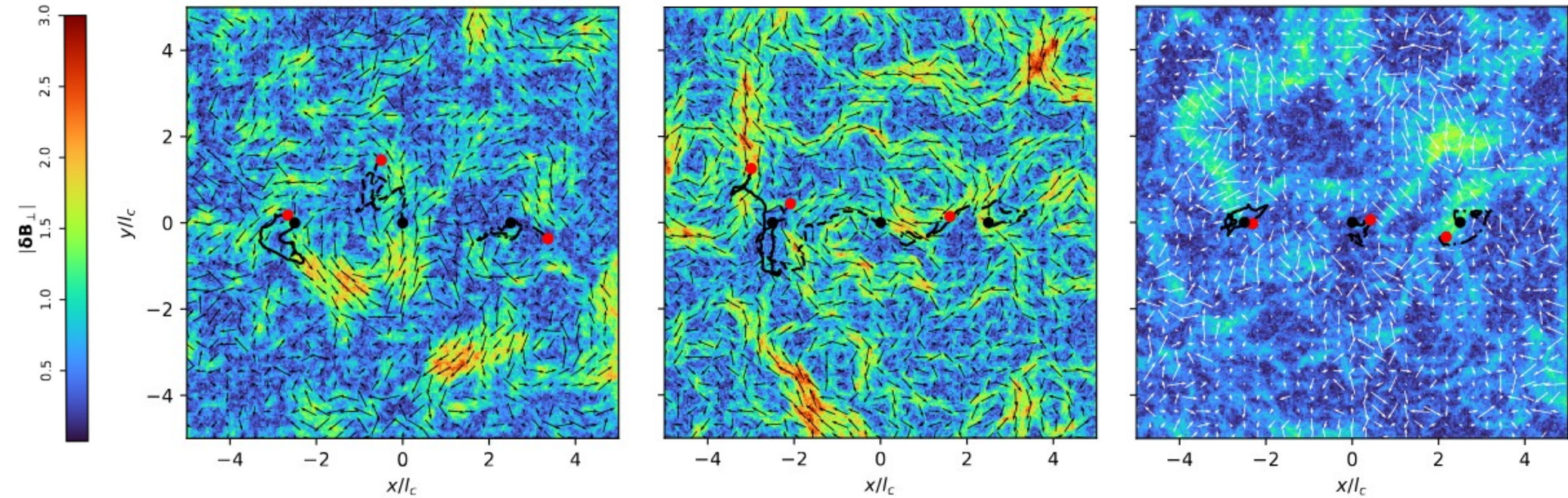
ISO:
Random

Synthetic polarized turbulence simulations

ISO
Random

ALF
Alfvén-like

MAG
Magnetosonic-like



→ We do see some patterns in such synthetic turbulence: filaments, different orientation of B vectors wrt these structures

Bouchet+ (to be sub.)

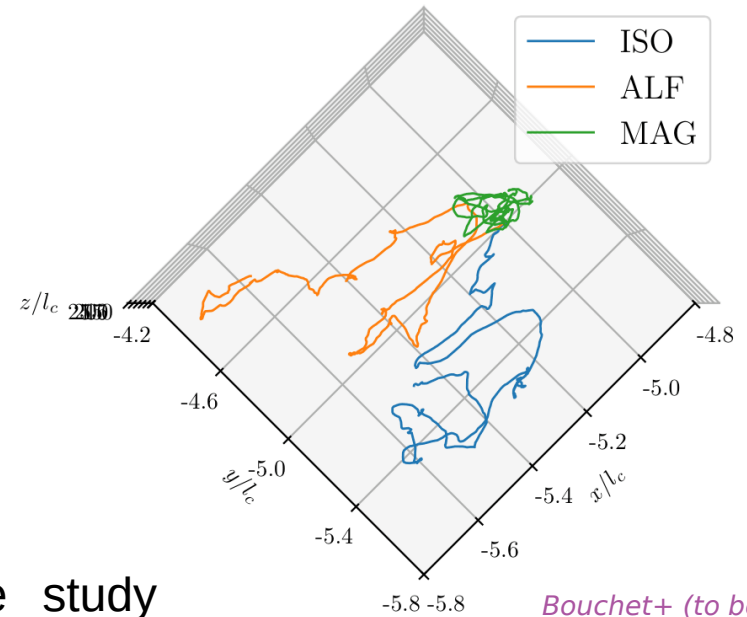
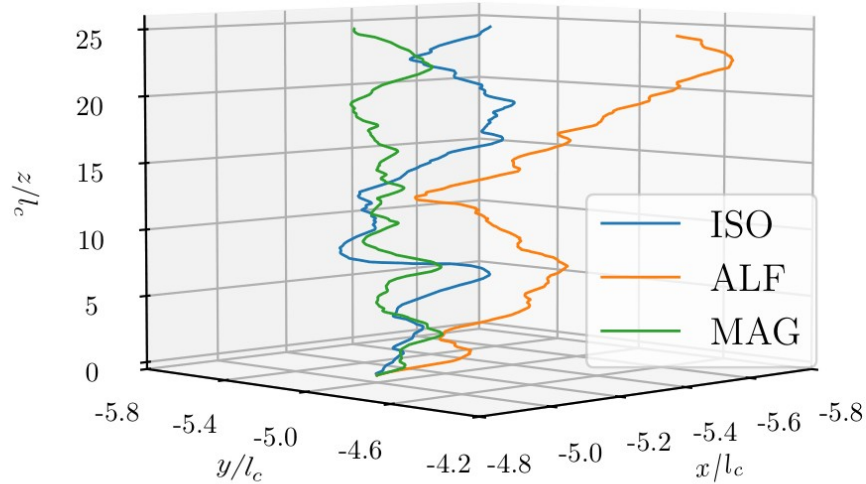
Field line transport

Solve field line equations :

$$\frac{d\mathbf{r}(s)}{ds} = \frac{\mathbf{B}(\mathbf{r}(s))}{|\mathbf{B}(\mathbf{r}(s))|} \equiv \mathbf{b}(\mathbf{r}(s)) \quad \text{or} \quad \frac{d\mathbf{r}(\tau)}{d\tau} = \mathbf{B}(\tau) \quad \tau = s/|\mathbf{B}|$$

Display of field lines for each polarisation config :

ISO ALF MAG
 Random Alfvén-like Magnetosonic-like



Bouchet+ (to be sub.)

- FL in MAG much more confined in (x, y) plane
- After reconstructing thousands of field lines, we study statistical behaviours

Field line transport

- Define parallel and transverse diffusion coefficients

$$\mathcal{D}_\perp = \frac{1}{4} \frac{d}{ds} \left(\langle (x(s) - x_0)^2 \rangle + \langle (y(s) - y_0)^2 \rangle \right)$$

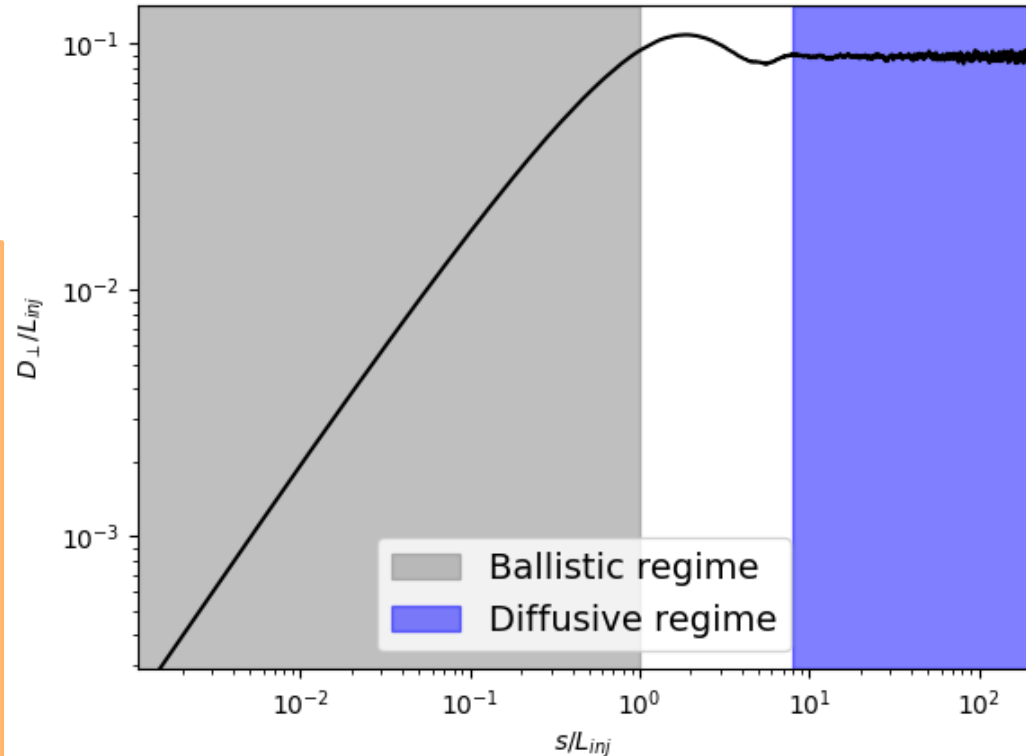
$$\mathcal{K}_\perp = \lim_{s \rightarrow \infty} \mathcal{D}_\perp$$

- Alternative parametrization: helpful for analytical modeling (especially at high η)

$$\begin{aligned} D_\perp &= \frac{1}{4} \frac{d}{d\tau} \left(\langle (x(\tau) - x_0)^2 \rangle + \langle (y(\tau) - y_0)^2 \rangle \right) \\ &= \frac{1}{2} \left(\langle \delta B_x(\mathbf{r}(\tau))(x(\tau) - x_0) \rangle \right. \\ &\quad \left. + \langle \delta B_y(\mathbf{r}(\tau))(y(\tau) - y_0) \rangle \right), \end{aligned}$$

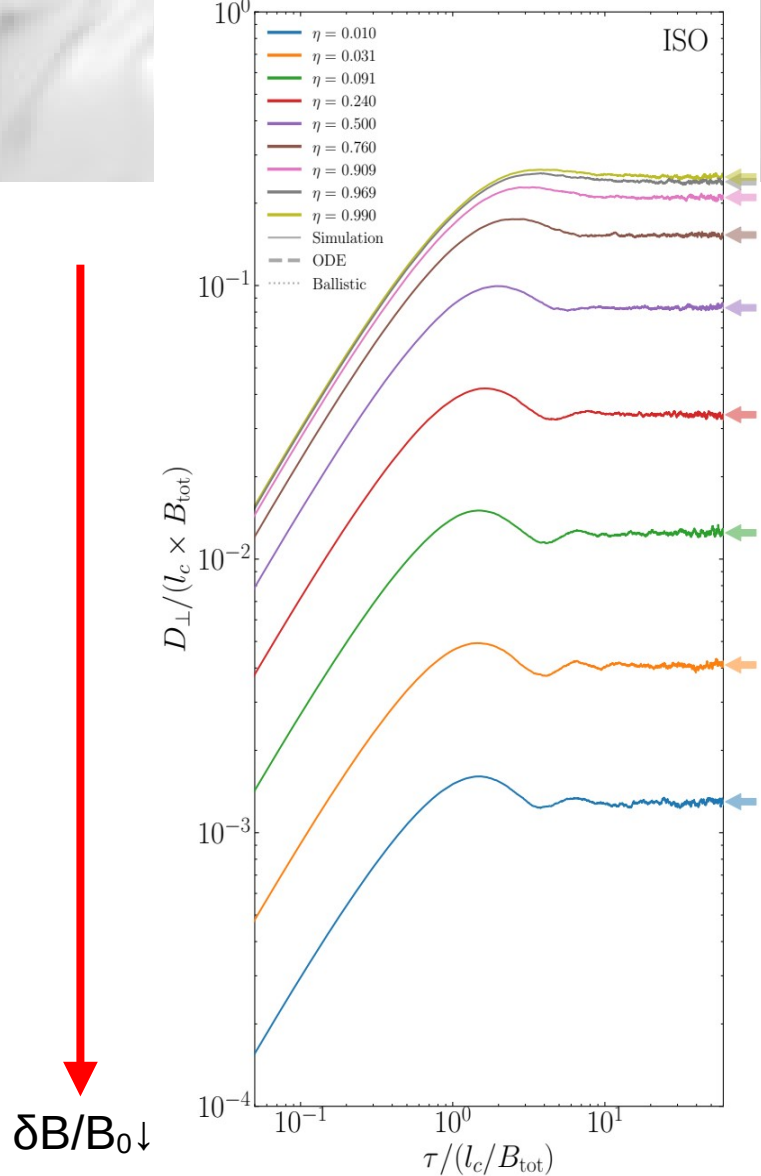
$$D_\parallel = \frac{1}{2} \frac{d}{d\tau} \langle (z - \langle z \rangle)^2 \rangle = \langle \delta B_z(\tau)(z - \langle z \rangle) \rangle.$$

$$K_\perp = \lim_{\tau \rightarrow \infty} D_\perp$$



Here we study FL running diffusion coefficients wrt to turbulence level $\eta = \frac{\delta B^2}{\delta B^2 + B_0^2}$ & polarisation config .

Field line transport



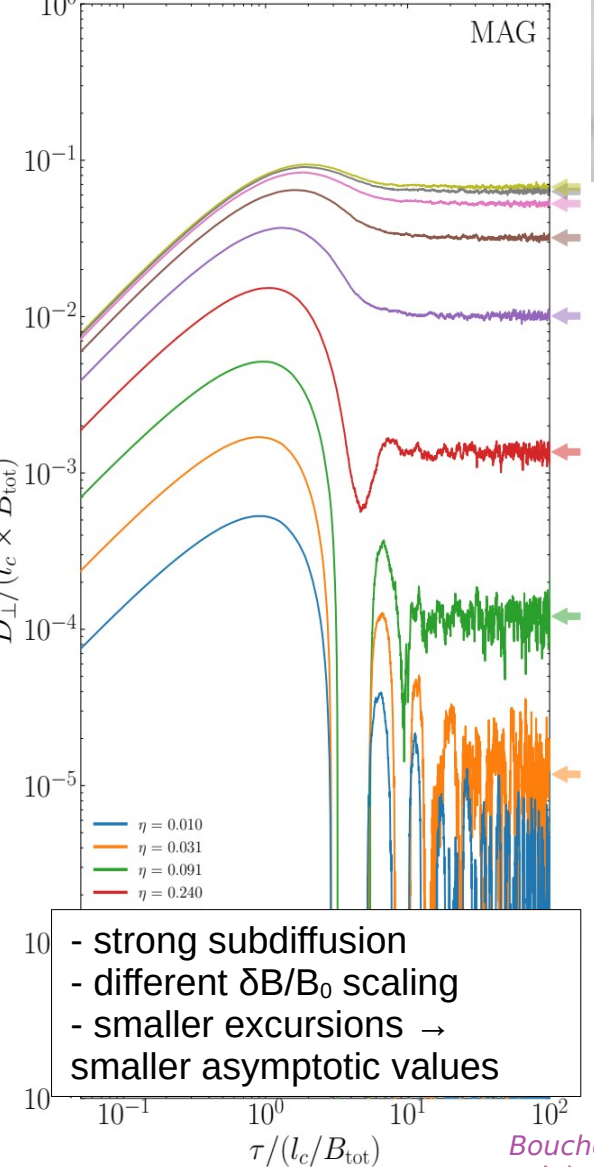
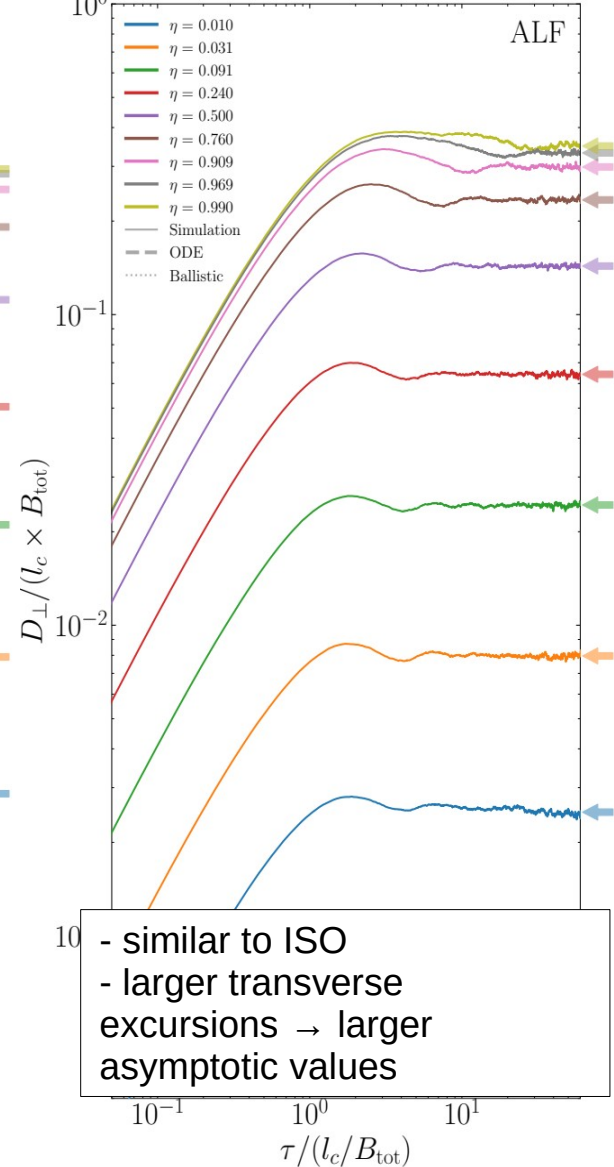
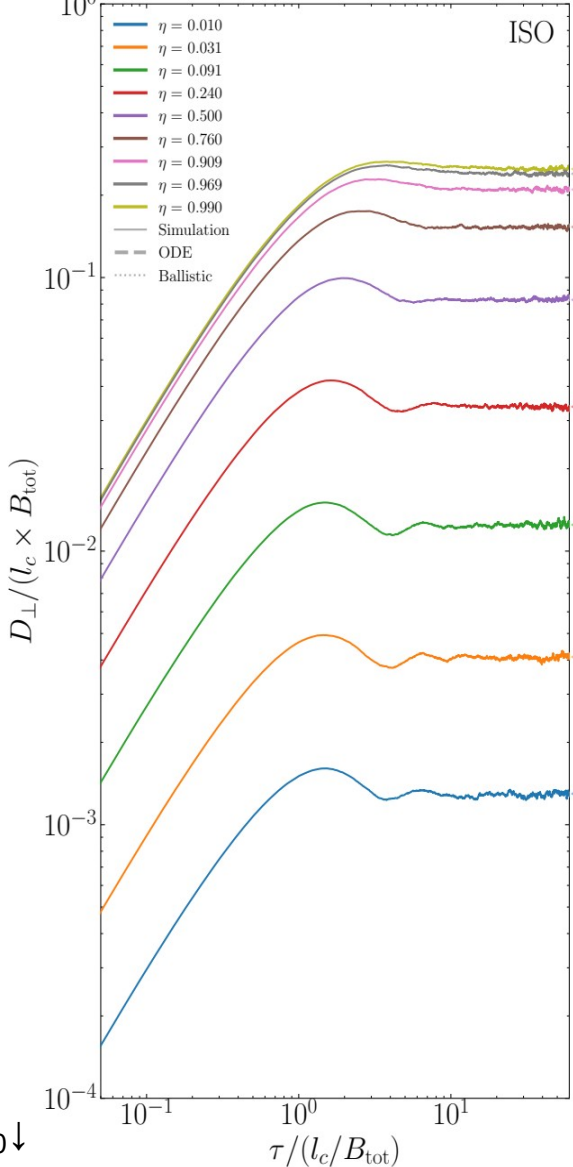
Results from simulations :

- Three phases : ballistic ($\alpha=1$), transition through subdiffusion ($\alpha<0$), diffusive plateau ($\alpha=0$)
- Particular scaling with B_0

Fig

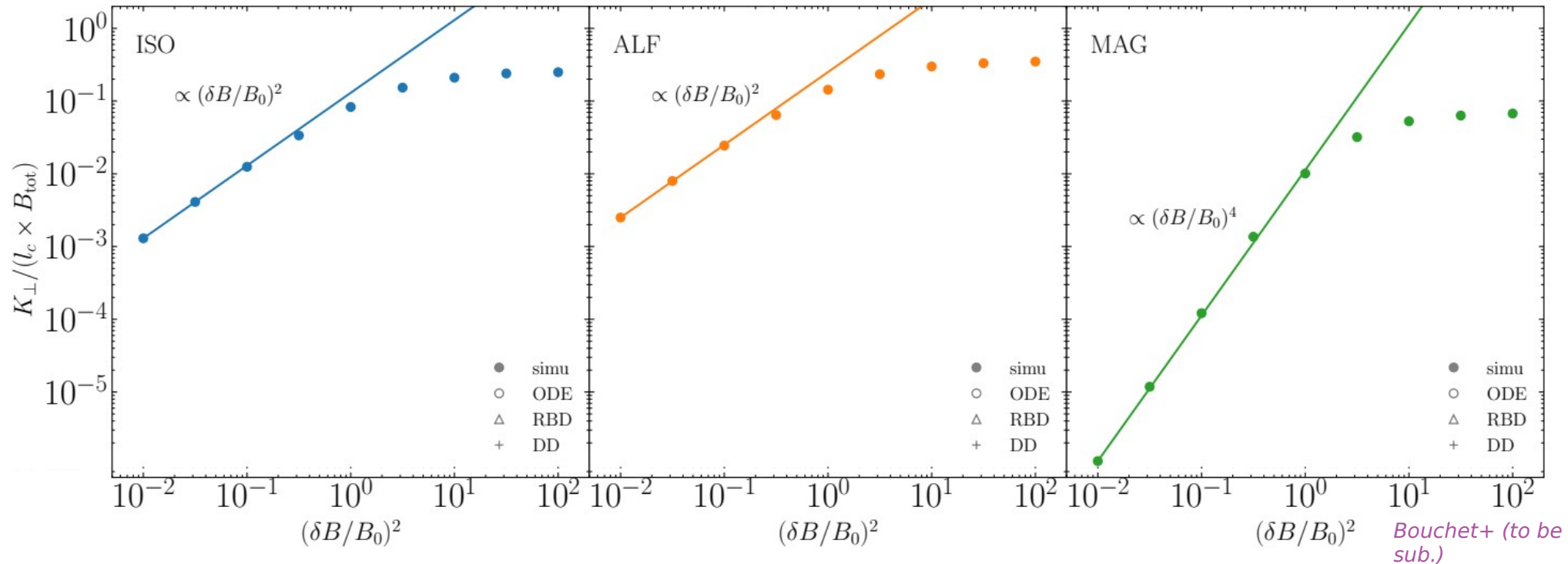


$\delta B/B_0 \downarrow$



Field line transport

Field line asymptotic diffusion coefficients



→ Different scaling in the MAG case

Field line transport

Correlation length of the total Bfield :

$$l_c = \frac{2}{\delta B^2} \int_0^\infty dl \langle \delta \mathbf{B}(\mathbf{r}) \delta \mathbf{B}(\mathbf{r} + \mathbf{l}) \rangle$$

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Correlation length of the total transverse Bfield :

$$l_{B_\perp}(\theta_*) = \frac{2}{\delta B_\perp^2} \int_0^\infty \langle \delta \mathbf{B}_\perp(\mathbf{r}) \delta \mathbf{B}_\perp(\mathbf{r} + \mathbf{l}(\theta_*)) \rangle dl$$

Field line transport

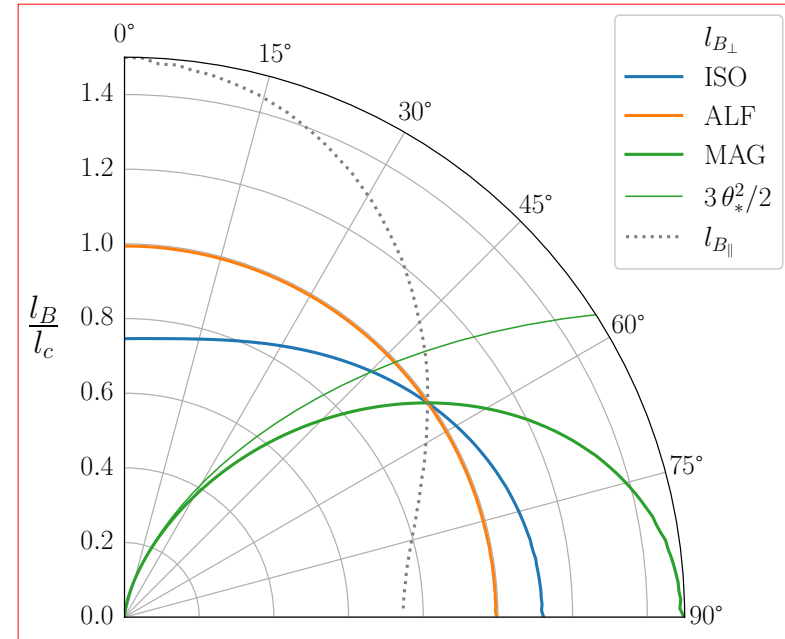
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In the MAG case : $l_{B_\perp}^{\text{MAG}}(\theta_* \rightarrow 0) \approx \frac{3}{2} l_c \theta_*^2 \approx \frac{3}{2} l_c \left(\frac{\delta B_\perp}{B_0} \right)^2$



Bouchet+ (to be sub.)

Field line transport

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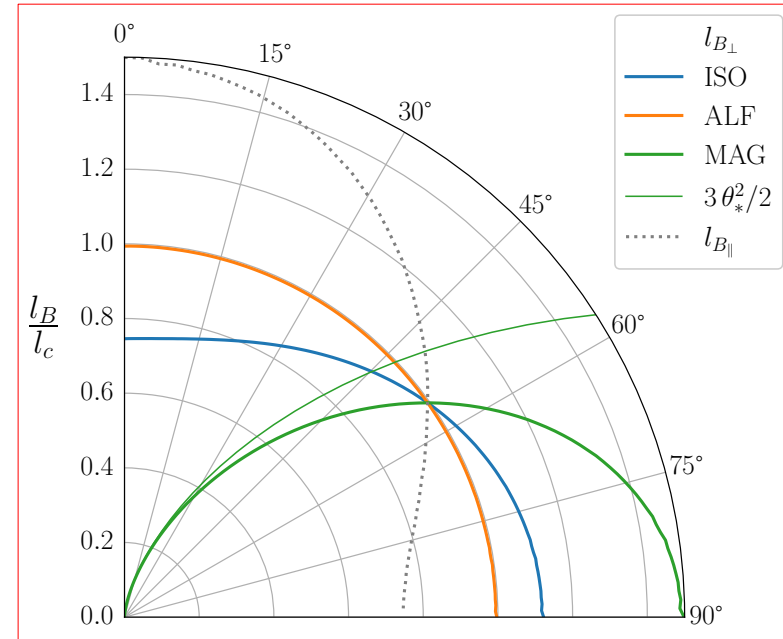
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Heuristic derivation of the scalings in the regime : $\delta B/B_0 \ll 1$

→ Transverse displacement : $\Delta r_\perp \approx \theta_* l_{B_\perp}(\theta_* \rightarrow 0)$

→ Scaling : $K_\perp \approx \frac{(\Delta r_\perp)^2}{l_{B_\perp}(\theta_* \rightarrow 0)} \approx \theta_*^2 l_{B_\perp}(\theta_* \rightarrow 0)$



Bouchet+ (to be sub.)

ISO & ALF

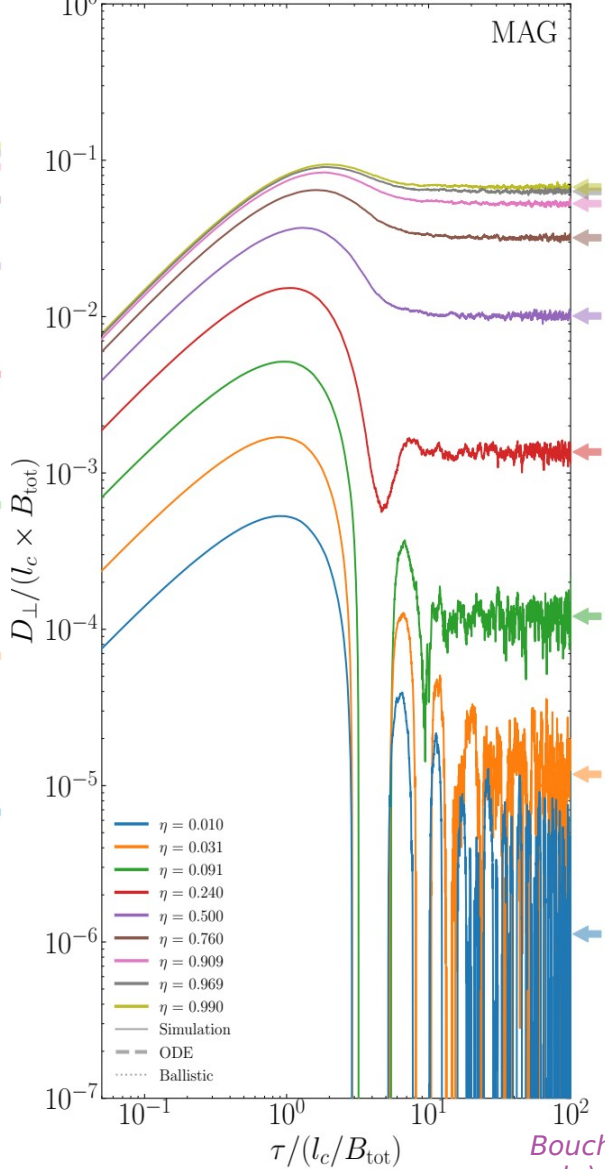
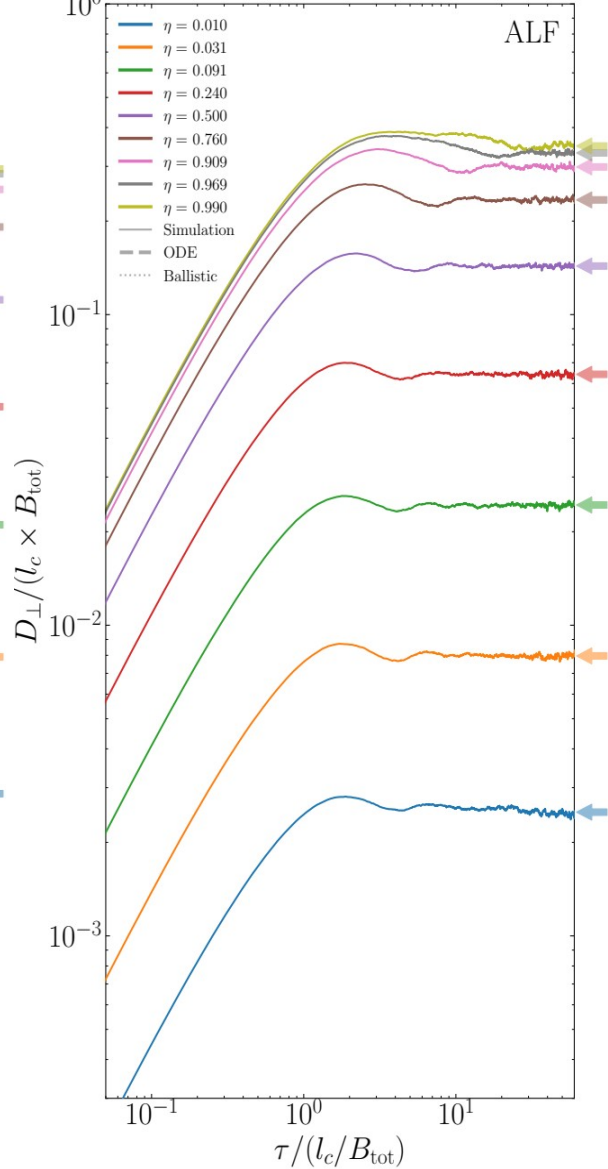
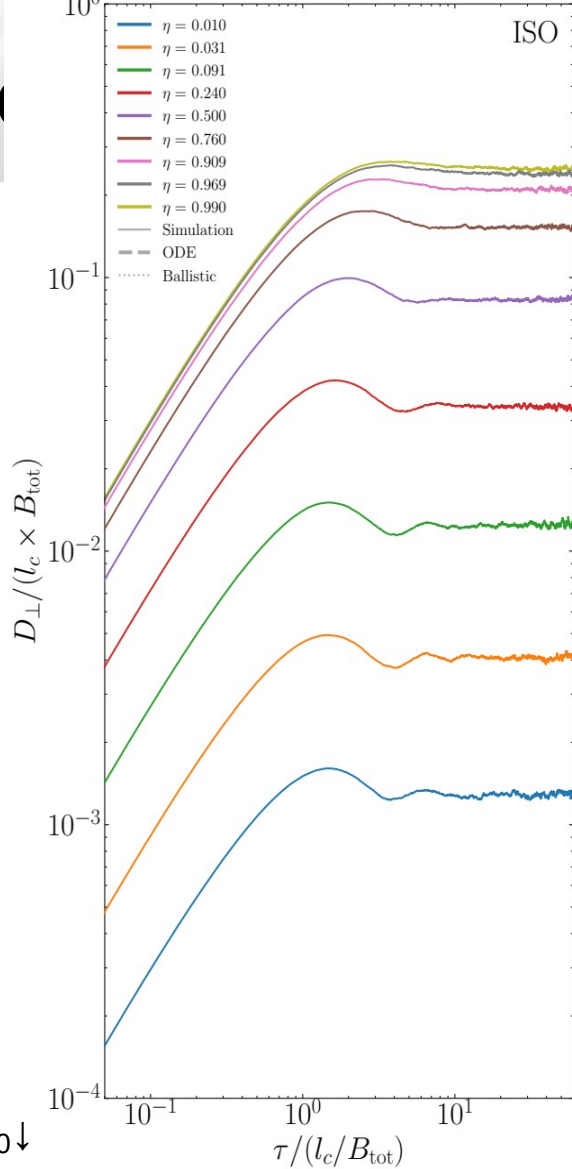
$$K_\perp \propto \theta_*^2 \propto (\delta B/B_0)^2$$

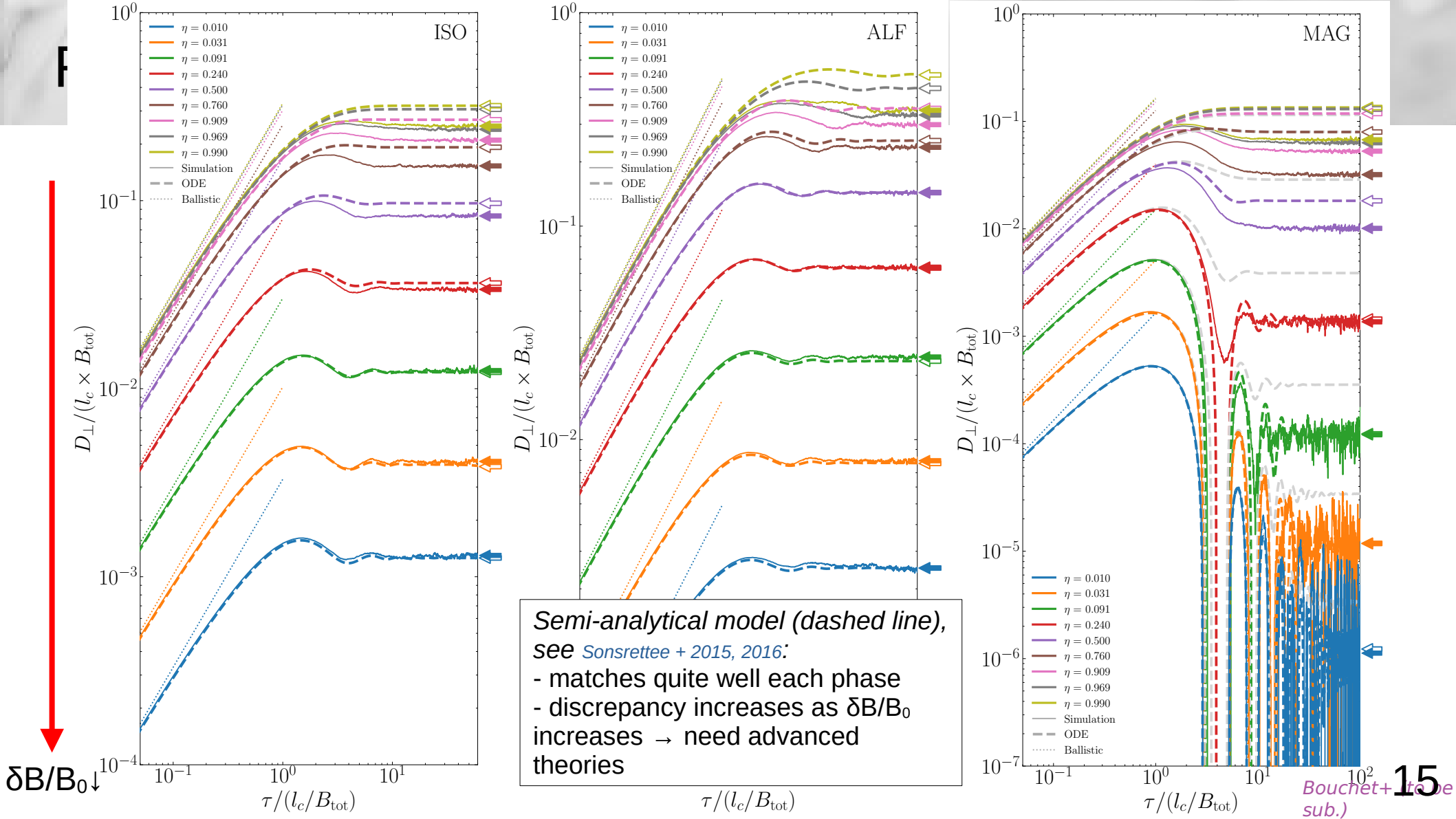
MAG

$$K_\perp \propto \theta_*^4 \propto (\delta B/B_0)^4.$$

Fig

$\delta B/B_0 \downarrow$





Particle transport

Now, we propagate particles in these magnetostatic nested grid turbulence configurations :

- Solve Lorentz force + Boris pusher
- Compute running diffusion coefficients

$$d_{\parallel}(t) \equiv \frac{1}{2} \frac{d}{dt} \langle (z(t) - z(0))^2 \rangle,$$

$$d_{\perp}(t) \equiv \frac{1}{4} \frac{d}{dt} (\langle (x(t) - x(0))^2 \rangle + \langle (y(t) - y(0))^2 \rangle)$$

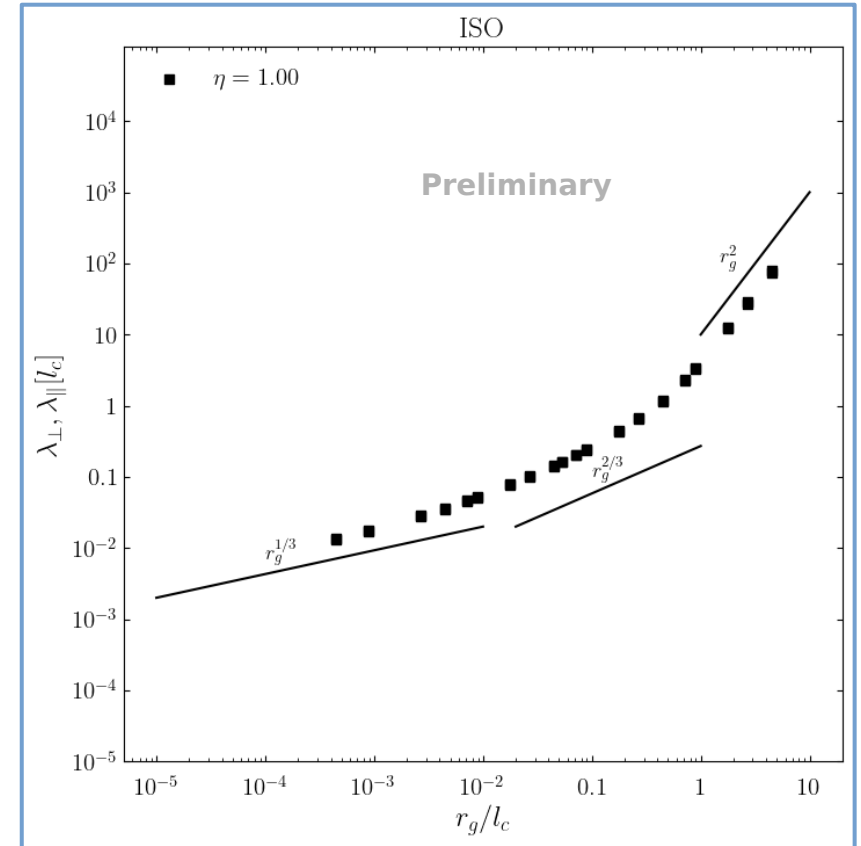
- Extract asymptotic values

$$\kappa_{\parallel} \equiv \lim_{t \rightarrow \infty} d_{\parallel}(t) \propto \lambda_{\parallel}$$

$$\kappa_{\perp} \equiv \lim_{t \rightarrow \infty} d_{\perp}(t) \propto \lambda_{\perp}$$

λ_{\parallel} Transition from QLT limit to small scattering regime

λ_{\perp} ISO : - small r_g recover QLT scaling
 - steeper transition before flattening
 - is it linked to FL subdiffusion ?



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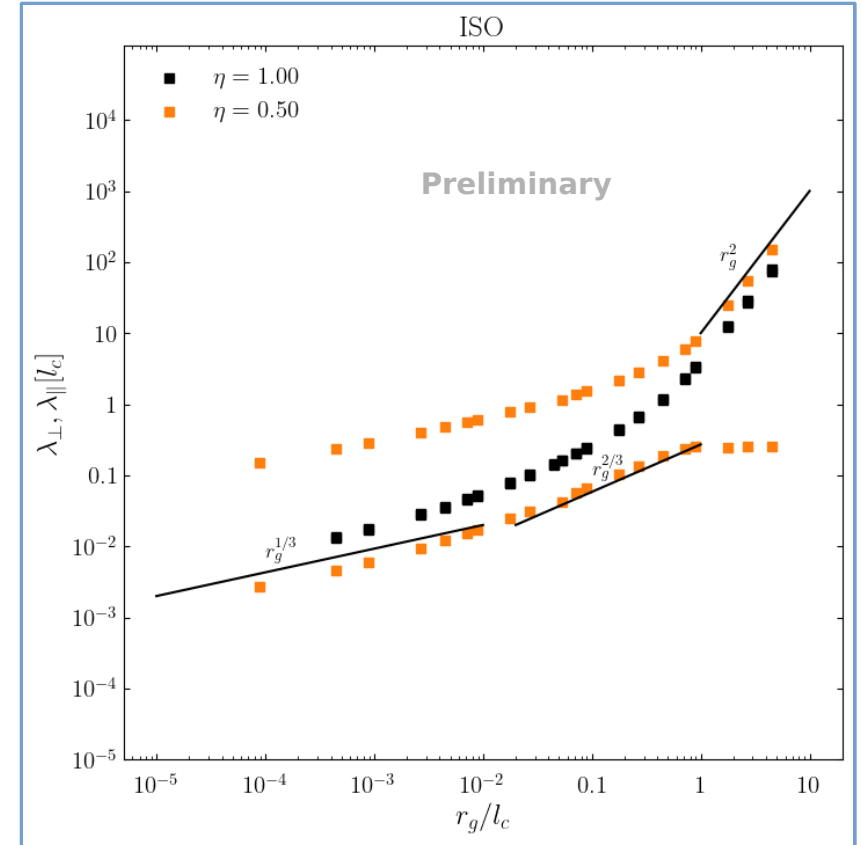
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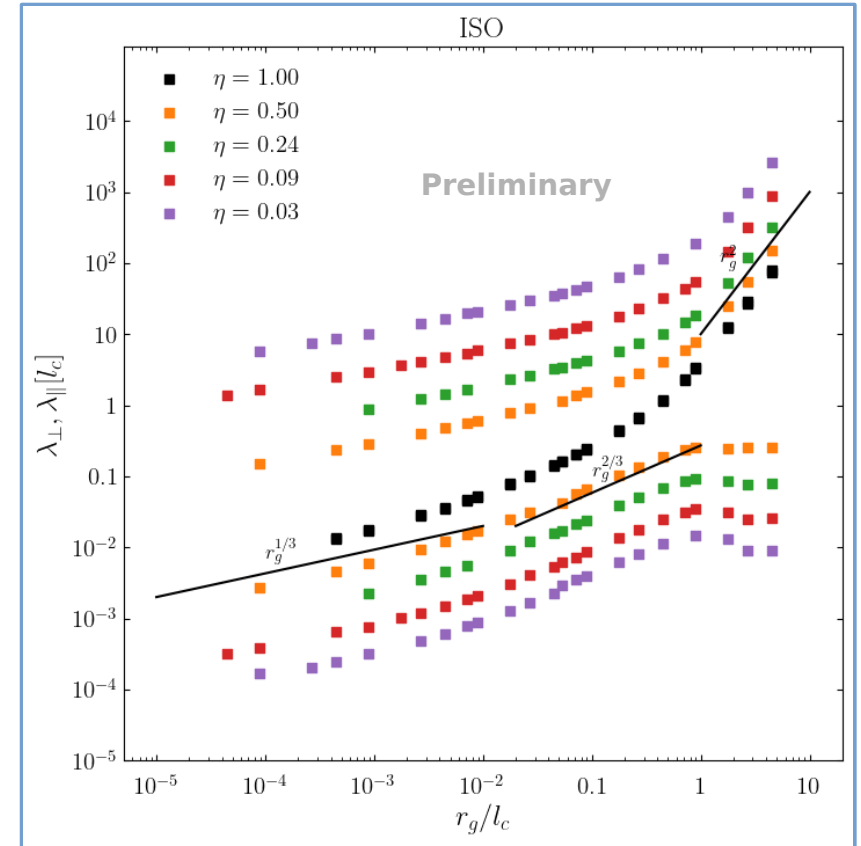
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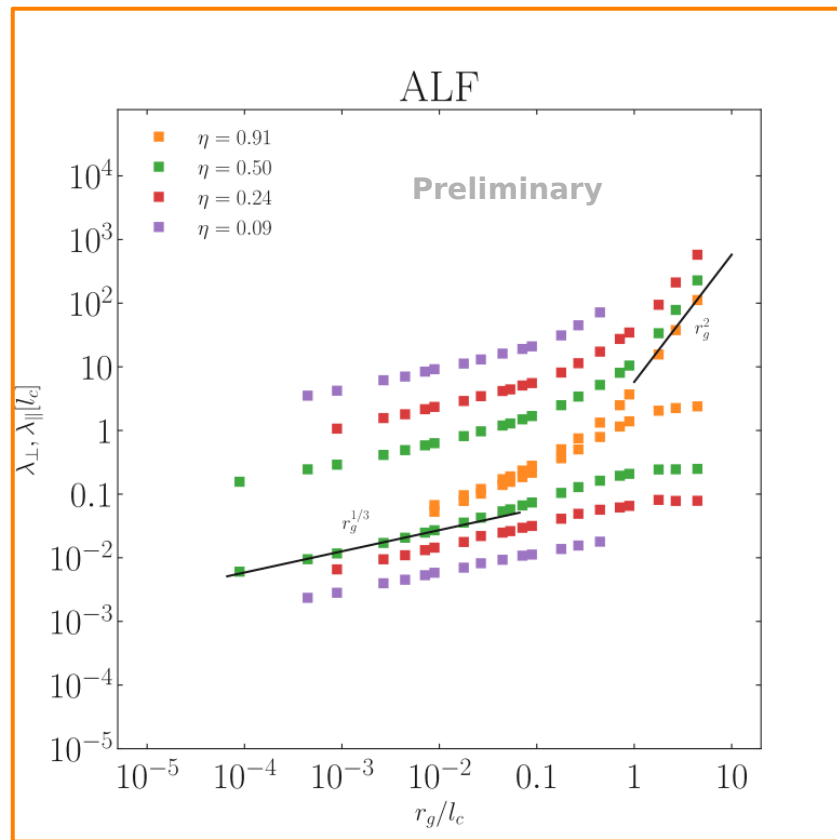
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ALF : - QLT scaling and faster transition before flattening



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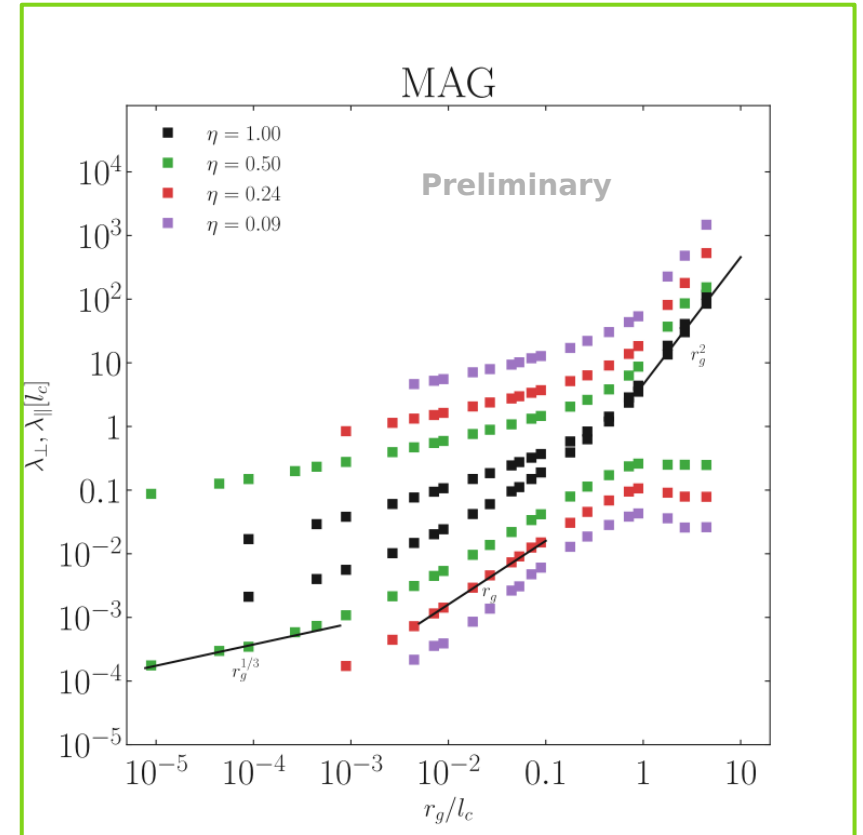
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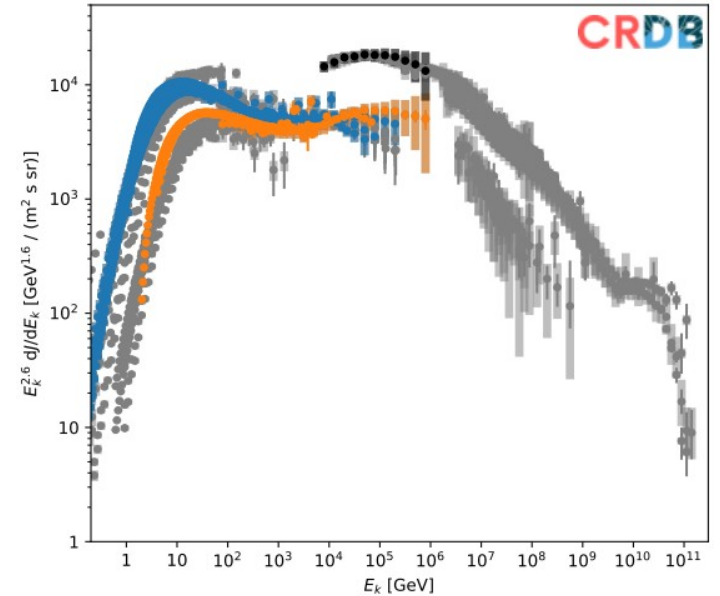
MAG : - intermediate regime propto r_g

We are currently working on predicting these values from FL transport



Conclusion

- Context : cosmic ray data from high precision direct and indirect observations → from phenomenology to microphysics
- Goal : understand and quantify CR scattering in « simple » synthetic turbulence, still ongoing !
- We found new behaviours for FL & particle transport in magnetosonic-like polarized turbulence : two papers coming soon...
- These findings are small steps towards a general understanding of CR scattering in MHD turbulence.



Thank you !