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→ Standard Model is based on the principle that — Any theory of fundamental nature, which explains physics at High Energy must be consistent with relativity & quantum theory.

→ The problem of single particle relativistic wave equation is

$$i\hbar \frac{\partial \psi}{\partial t} = H\psi = \sqrt{-\hbar^2 c^2 \nabla^2 + m^2 c^4} \psi$$

✓ in the above equation L.H.S $\frac{\partial \psi}{\partial t}$ & r.h.s $\sqrt{\nabla^2} \psi$. upon expansion it will give infinite terms containing ∇^2 on ψ . one can instead use

$$E = i\hbar \frac{\partial}{\partial t} \quad \& \quad \vec{p} = -i\hbar \vec{\nabla} \quad \text{in}$$

$$E^2 = \vec{p}^2 + m^2$$

which will give

$$\left(\frac{\partial^2}{\partial t^2} - \nabla^2 \right) \psi + m^2 \psi = 0$$

$$\Rightarrow \boxed{\square \psi + m^2 \psi = 0} \Rightarrow \text{Klein-Gordon equation}$$

Relativistic invariant where ψ is a wave-function.

→ However the problem is this kind of equation can not describe a process where new particles are produced or particle number not conserved.

slide on LHC/RHIC events

other problems

→ probability density associated with wave-function $\rho = \psi^* \psi$ in non relativistic approach.

$$\vec{j} = \frac{-i\hbar}{2m} (\psi^* \vec{\nabla} \psi - \psi \vec{\nabla} \psi^*)$$

Together they satisfy continuity equation

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0 \quad \checkmark$$

With relativistic approach, define

$$j^\mu = (\rho, \vec{j}) = \frac{i\hbar}{m} \psi^* \overleftrightarrow{\partial}^\mu \psi \quad \checkmark$$

definition $\overleftrightarrow{\partial}^\mu = \frac{1}{2} [A \partial^\mu B - (\partial^\mu A) B]$

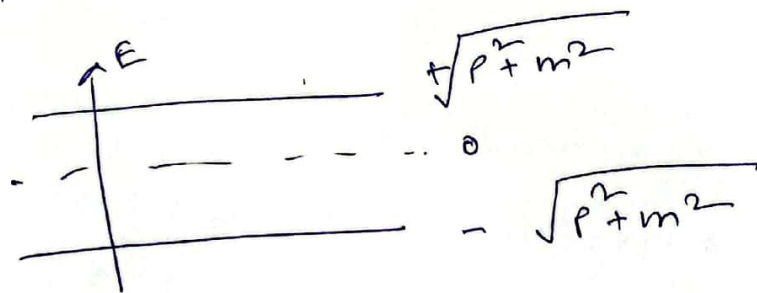
$$\text{So, } j^0 = \frac{i\hbar}{2m} \left[\psi^* \frac{\partial \psi}{\partial t} - \frac{\partial \psi^*}{\partial t} \psi \right]$$

$$j^i = \frac{-i\hbar}{2m} \left[\psi^* \nabla \psi - (\nabla \psi^*) \psi \right]$$

$$\partial_\mu j^\mu = 0$$

However ρ now can be negative.
So ρ as probability density not viable.

→ Another problem with Klein-Gordon equation is solution of $p^2 = p_\mu p^\mu = m^2$ gives both $E = \pm \sqrt{p^2 + m^2}$, i.e. along with positive energy also negative energy solutions.



For a free particle, energy is constant and once positive solution is chosen, it will be at $+\sqrt{p^2 + m^2}$. But for an interacting particle, there can be exchange of energy with other interacting states. So particle can cascade down to negative energy states with ground state energy infinitely negative.

→ The problem with causality.

✓ free particle to propagate from x_0 to x

$$\begin{aligned} u(t) &= \langle x | e^{-iHt} | x_0 \rangle \\ &= \langle x | e^{-i\hat{p}^2/2m t} | x_0 \rangle = \int \frac{d^3p}{(2\pi)^3} \langle x | e^{-i(\cdot)} | p \rangle \langle p | x_0 \rangle \\ &= \int \frac{d^3p}{(2\pi)^3} e^{-i p^2/2m t} e^{i p \cdot (x - x_0)} \\ \checkmark &= \left(\frac{m}{2\pi i t} \right)^{3/2} e^{i m (x - x_0)^2 / 2 t} \end{aligned}$$

The amplitude is non-zero for all $x \neq x_0$.
This indicates particle can propagate
between $x \neq x_0$ in $t \rightarrow 0$ time.

with $E = p^2/2m$

$$u(t) \sim e^{-m \sqrt{x^2 - t^2}} \quad \begin{array}{l} \text{still non-zero,} \\ \text{although} \\ \text{can be small.} \end{array}$$

Causality is violated.

We adhere to the field description,
rather than following single particle
relativistic wave equation.

Dynamics of fields

A field is a dynamical variable, $\phi(x^\mu)$.

→ Classical mechanics accompanied with finite number of generalised co-ordinates $q_a(t)$, a is index & conjugate moment p_a

$$\dot{q}_a(t)$$

→ In QM, we consider q_a & $\dot{q}_a$ as operators.

→ In field theory field is labelled as $\phi_a(\vec{x}, t)$, both a & $(\vec{x}, t)$ are labels.

→ We are working with infinite number of degrees of freedom, at least one in each point x^μ .

→ Common example is Electric & Magnetic field $\vec{E}(\vec{x}, t)$ & $\vec{B}(\vec{x}, t)$ $\{A^\mu(\vec{x}, t)\}$
 $= (\phi, \vec{A})$
 $\mu = 0, 1, 2, 3$

The field evolves (at any point x^μ)
respecting Eqs of motion.

→ $\mathcal{L}$ (Lagrangian) is a function of

$$\phi(\vec{x}, t) \text{ \& } \dot{\phi}(\vec{x}, t) \text{ \& } \vec{\nabla} \phi(\vec{x}, t)$$

[Analogous to classical
mechanics
 $\mathcal{L}(x, \dot{x})$]

→ Define Action $S = \int d^4x \mathcal{L}$
 $\mathcal{L}$ is Lagrangian density.

→ Remember in classical mechanics

$$L = T - V = \frac{1}{2} m \dot{q}^2 - V$$

→ In field theory $\mathcal{L} = f(\phi, \partial_\mu \phi)$

(a) $\mathcal{L}$ doesn't depend explicitly on x^μ , rather the dependency appears via $\phi(x^\mu) \text{ \& } \partial_\mu \phi(x^\mu)$

(b) We are not considering system $\mathcal{L} \rightarrow f(\phi, \partial_\mu \phi, \partial_\mu^2 \phi)$ etc, rather only $f(\phi, \partial_\mu \phi)$

Principle of least action

The field ϕ at ~~any~~ x^μ evolves in accordance to eqns of motion. The eqn of motion is derived from principle of least action.

$$\mathcal{L} \Rightarrow \mathcal{L}(\phi_\mu, \partial_\mu \phi)$$

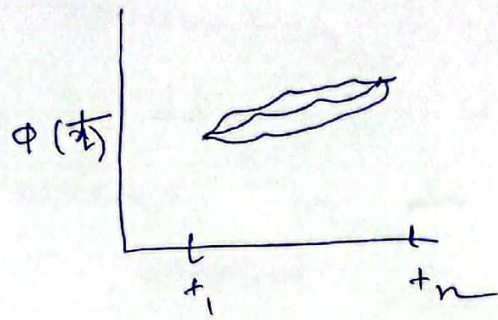
$$\begin{aligned} \delta S &= \int d^4x \left[\frac{\partial \mathcal{L}}{\partial \phi_a} \delta \phi_a + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \delta (\partial_\mu \phi_a) \right] \\ &= \int d^4x \left[\frac{\partial \mathcal{L}}{\partial \phi_a} \delta \phi_a - \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \right) \delta \phi_a \right] \\ &\quad + \underbrace{\int d^4x \partial_\mu \left[\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \delta \phi_a \right]} \end{aligned}$$

A total derivative term

By Gauss theorem $\int_V d^4x \partial_\mu F^\mu = \int_{\partial V} d\Omega F^\mu$

→ The total derivative term $\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \delta \phi_a$ is $\left. \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \delta \phi_a \right|_{x^\mu_{in}}^{x^\mu_{final}}$ Limits of surface integral

→ We consider $\delta \phi_a(\vec{x}, t_1) = \delta \phi_a(\vec{x}, t_2) = 0$
 & as $|\vec{x}| \rightarrow \infty$, $\delta \phi_a(\vec{x}, t) \rightarrow 0$



ϕ will evolve along the path for which action is an extremum. i.e. $\delta S = 0$

With this Euler-Lagrange Eqnⁿ

$$\partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \right) - \frac{\partial \mathcal{L}}{\partial \phi_a} = 0$$

$\Rightarrow$ This is relativistically co-variant

$\mathcal{L}$ is local, i.e. $\int d^3x d^3y \phi(x)\phi(y)$ not present $\rightarrow$ no $\vec{\nabla}^2 \phi$ term or $\dot{\phi}^2$ term $\rightarrow$ Hamiltonian unbounded from below $\rightarrow$ violates causality

(A) scalar field (Real)

$$x^\mu \rightarrow x'^\mu$$

$$x'^\mu = \Lambda^\mu_\nu x^\nu$$

$\phi(\vec{x}, t)$ is a scalar field.

$\rightarrow$ A Lorentz invariant field.

$$\phi(x') = \phi(x)$$

$$= \phi(\Lambda^{-1} x')$$

$$x' \rightarrow x$$

i.e. if $x^\mu \rightarrow \Lambda^\mu_\nu x^\nu$

$$\phi(x) \rightarrow \phi'(x) = \phi(\Lambda^{-1} x)$$

$$\phi(x) = \phi(\Lambda^{-1} x)$$

$$\mathcal{L} = \frac{1}{2} \eta^{\mu\nu} (\partial_\mu \phi)(\partial_\nu \phi) - \frac{1}{2} m^2 \phi^2$$

$$\eta^{\mu\nu} = \begin{pmatrix} +1 & & & \\ & -1 & & \\ & & -1 & \\ & & & -1 \end{pmatrix} = \frac{1}{2} (\partial^\mu \phi)(\partial_\mu \phi) - \frac{1}{2} m^2 \phi^2$$

$\rightarrow$ flat metric

Eqⁿ of motion

$$\frac{\partial \mathcal{L}}{\partial \phi} = -m^2 \phi$$

$$\begin{aligned} \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} &= \frac{1}{2} (\partial_\rho \phi) \eta^{\rho\sigma} \delta_{\sigma\mu} + \frac{1}{2} (\partial_\sigma \phi) \eta^{\rho\sigma} \delta_{\rho\mu} \\ &= (\partial^\mu \phi) \end{aligned}$$

Eqⁿ of motion becomes

$$\partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \right) - \frac{\partial \mathcal{L}}{\partial \phi} = 0$$

$$\Rightarrow \partial_\mu \partial^\mu \phi + m^2 \phi = 0$$

$$\Rightarrow \boxed{(\square + m^2) \phi = 0}$$

The difference w.r.t relativistic q.m is ϕ is not a wave function. Rather $\phi(x^\mu)$ is a field whose value at every x^μ can be different. Infinite number of degrees of freedom.

$$\rightarrow \text{In general } \boxed{\mathcal{L} = \frac{1}{2} (\partial_\mu \phi) (\partial^\mu \phi) - V(\phi)}$$

$$\begin{aligned} \rightarrow \text{for complex scalar field } \phi &= \frac{\phi' + i\phi''}{\sqrt{2}} \\ \mathcal{L} &= (\partial_\mu \phi) (\partial^\mu \phi^*) - m^2 |\phi|^2 \end{aligned}$$

Ex: 2

Prove that $\mathcal{L}$ with real Klein-Gordon field is Lorentz invariant.

Meaning

$$x^M \rightarrow \Lambda^M_{\ \nu} x^\nu \quad (\text{or } y^M = \Lambda^{-1 M}_{\ \nu} x^\nu)$$

for this transformation $\mathcal{L}$ remains the same.

→ The Hamiltonian

$$H = \pi \dot{\phi} - \mathcal{L}$$

$$\text{with } \pi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} \quad (\text{conjugate momentum to } \phi)$$

$$H = \frac{\pi^2}{2} + \frac{(\vec{\nabla} \phi)^2}{2} + \frac{1}{2} m^2 \phi^2$$

has the structure $T + V$

Hamilton's eqⁿ of motion

$$\left\{ \begin{array}{l} \dot{q} = \frac{\partial H}{\partial p} \quad \& \quad \dot{p} = -\frac{\partial H}{\partial q} \end{array} \right\} \text{ in classical mechanics}$$

$$\dot{\phi}(\vec{x}, t) = \frac{\partial H}{\partial \pi(\vec{x}, t)} \quad \& \quad \dot{\pi}(\vec{x}, t) = -\frac{\partial H}{\partial \phi(\vec{x}, t)}$$

A physical scalar particle with mass m is obtained by quantization of the scalar field.

A quick Guide to quantisation procedure [canonical quantization]

→ For real scalar field, we re-interpret the dynamical variables (i.e. $\phi(x^\mu)$ & $\pi(x^\mu)$) as operators that obey certain mathematical relation. ✓

→ Solve the theory by finding Eigenvalues & eigenstates of Hamiltonian
[Analogy is Harmonic Oscillator system]

→ Recall in QM, $[\hat{q}^i, \hat{p}^j] = i\delta^{ij}$
 $[\hat{q}^i, \hat{q}^j] = 0$ ✓
 $[\hat{p}^i, \hat{p}^j] = 0$

→ Generalise this

$[\phi(\vec{x}), \pi(\vec{y})] = i\delta^3(\vec{x} - \vec{y})$ ✓
 $[\phi(\vec{x}), \phi(\vec{y})] = [\pi(\vec{x}), \pi(\vec{y})] = 0$
equal time commutation relation.

$$\psi(\vec{r}) = \left\{ \frac{d^3 r}{(2\pi)^3} \tilde{\psi}(\vec{r}) e^{i\vec{p}\cdot\vec{r}} \right\} \quad \text{--- momentum eigenstates, } \vec{p} = \hbar \vec{k} \quad \vec{p} \cdot \vec{r} = \hbar \vec{k} \cdot \vec{r} = i\hbar \vec{k} \cdot \vec{r}$$

$$\psi(\vec{r}) = \left\{ \frac{d^3 r}{(2\pi)^3} \frac{1}{\sqrt{2\pi\hbar}} \left[\psi(\vec{r}) e^{i\vec{p}\cdot\vec{r}} + \psi^*(\vec{r}) e^{-i\vec{p}\cdot\vec{r}} \right] \right\}$$

$$\frac{d^3 r}{(2\pi)^3} \text{ can be } \dots$$

$$\psi(\vec{r}) = \int \frac{d^3 r}{(2\pi)^3} (-i) \sqrt{\frac{m}{2}} \left[\psi(\vec{r}) e^{i\vec{p}\cdot\vec{r}} + \psi^*(\vec{r}) e^{-i\vec{p}\cdot\vec{r}} \right]$$

The $\vec{r}$ & $\vec{p}$ are canonically conjugate variables satisfying

$$[\vec{r}_i, \vec{p}_j] = (2\pi)^3 \delta^3(\vec{r} - \vec{r}') \quad \text{---}$$

$$2 [\vec{r}_i, \vec{r}_j] = [\vec{p}_i, \vec{p}_j] = 0$$

The action of $a_{\vec{p}}$ is ground state
 $|0\rangle$ satisfies $a_{\vec{p}}|0\rangle = 0$ for all $\vec{p}$
 and $a_{\vec{p}}^\dagger|0\rangle = |\vec{p}\rangle$

→ The Hamiltonian is $\vec{p}^2/2m$

$$H = \int \frac{d^3 r}{(2\pi)^3} \psi_{\vec{p}}^\dagger (a_{\vec{p}}^\dagger a_{\vec{p}})$$

$$H|\vec{p}\rangle = \epsilon_{\vec{p}}|\vec{p}\rangle$$

$$\text{with } \epsilon_{\vec{p}} = \sqrt{|\vec{p}|^2 + m^2}$$

→ $|\vec{p}\rangle$ is an energy eigenstate of
 Hamiltonian with definite
 energy $\epsilon_{\vec{p}}$

$$\Phi(x) = \int \frac{d^3 p}{(2\pi)^3} \tilde{\Phi}(p) e^{i p \cdot x} \rightarrow \text{Harmonic oscillator system}$$

$$\Phi(\vec{x}) = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\sqrt{2\omega_{\vec{p}}}} \left[a_{\vec{p}} e^{i \vec{p} \cdot \vec{x}} + a_{\vec{p}}^\dagger e^{-i \vec{p} \cdot \vec{x}} \right]$$

$$\frac{\partial^2 \tilde{\Phi}(p)}{\partial t^2} + \omega^2 \tilde{\Phi}(p) = 0$$

$$\Pi(\vec{x}) = \int \frac{d^3 p}{(2\pi)^3} (-i) \sqrt{\frac{\omega_{\vec{p}}}{2}} \left[a_{\vec{p}} e^{i \vec{p} \cdot \vec{x}} - a_{\vec{p}}^\dagger e^{-i \vec{p} \cdot \vec{x}} \right]$$

The $a_{\vec{p}}^\dagger$ & $a_{\vec{p}}$ are creation and annihilation operators satisfying

$$[a_{\vec{p}}, a_{\vec{p}'}^\dagger] = (2\pi)^3 \delta^3(\vec{p} - \vec{p}')$$

$$\& [a_{\vec{p}}, a_{\vec{q}}] = [a_{\vec{p}}^\dagger, a_{\vec{q}}^\dagger] = 0$$

The action of $a_{\vec{p}}$ is ground state $|0\rangle$ satisfies $a_{\vec{p}}|0\rangle = 0$ for all $\vec{p}$ and $a_{\vec{p}}^\dagger|0\rangle = |\vec{p}\rangle$

→ The Hamiltonian $H = \Pi \dot{\Phi} - \mathcal{L}$

$$H = \int \frac{d^3 p}{(2\pi)^3} \omega_{\vec{p}} (a_{\vec{p}}^\dagger a_{\vec{p}})$$

$$\& H|\vec{p}\rangle = \omega_{\vec{p}}|\vec{p}\rangle$$

with $\omega_{\vec{p}} = \sqrt{|\vec{p}|^2 + m^2}$

→ $|\vec{p}\rangle$ is an energy eigenstate of Hamiltonian with definite energy $\omega_{\vec{p}}$

a Multiparticle state

$|\vec{p}_1, \dots, \vec{p}_n\rangle$ can be obtained as

$$|\vec{p}_1, \dots, \vec{p}_n\rangle = a_{\vec{p}_1}^\dagger \dots a_{\vec{p}_n}^\dagger |0\rangle$$

& for scalar particle $|\vec{p}, \vec{q}\rangle = |\vec{p}, \vec{p}\rangle$

Propagator

The framework of QFT can successfully describe particle interaction (cross-section, decay rates etc)

one of the main ingredient to understand how the interactions occur is via the concept of propagator.

- Meaning:

→ $\phi(x)|0\rangle$ implies $a_{\vec{p}}^\dagger|0\rangle$, i.e.
a particle is created at x .

→ $\langle 0|\phi(x')$ implies particle is annihilated at x' as only annihilation part $\langle 0|a_{\vec{p}}$ non zero.

→ $\boxed{\langle 0|\phi(x')\phi(x)|0\rangle}$
↳ Amplitude of propagation of a particle from x to x' .

Some Mathematics

We define $G(x-x')$, a green's function as

$$(\square_x + m^2) G(x-x') = -i\delta^4(x-x')$$

$$G(x-x') = \int \frac{d^4 p}{(2\pi)^4} e^{-ip \cdot (x-x')} \frac{i}{p^2 - m^2 + i\epsilon}$$

The Feynman propagator is defined as $\frac{i}{p^2 - m^2 + i\epsilon}$

$$\begin{aligned} \rightarrow i\Delta_F(x-x') &= \langle 0 | T(\underbrace{\phi(x)\phi(x')}) | 0 \rangle \\ &\begin{cases} \phi(x)\phi(x') & \text{if } x^0 > x'^0 \\ \phi(x')\phi(x) & \text{if } x'^0 > x^0 \end{cases} \end{aligned}$$

Connection with causality? Assignment
-I

$$\phi(x) = -\int d^4 x' G(x-x') J(x') + \phi_0$$

⑦ for complex scalar field

$$\phi = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2\omega_{\vec{p}}}} \left[b_{\vec{p}} e^{i\vec{p} \cdot \vec{x}} + c_{\vec{p}}^\dagger e^{-i\vec{p} \cdot \vec{x}} \right]$$

$$\text{i.e. } \phi^\dagger = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2\omega_{\vec{p}}}} \left[b_{\vec{p}}^\dagger e^{-i\vec{p} \cdot \vec{x}} + c_{\vec{p}} e^{i\vec{p} \cdot \vec{x}} \right]$$

$$[b_{\vec{p}}, b_{\vec{q}}^\dagger] = (2\pi)^3 \delta^3(\vec{p} - \vec{q})$$

$$[c_{\vec{p}}, c_{\vec{q}}^\dagger] = (2\pi)^3 \delta^3(\vec{p} - \vec{q})$$

All other combination $\Rightarrow 0$

$$[\phi(\vec{x}), \pi(\vec{y})] = [\phi^\dagger(\vec{x}), \pi^\dagger(\vec{y})] \\ = i\delta^3(\vec{x} - \vec{y})$$

→ Two creation operators
 $b_{\vec{p}}^\dagger$ & $c_{\vec{p}}^\dagger$

→ one of them for particle $\rightarrow b_{\vec{p}}^\dagger$ for particle

→ And another for antiparticle $\rightarrow c_{\vec{p}}^\dagger$ for Anti-particle

→ for a real scalar particle
particle & anti-particle same

$$\rightarrow H \text{ is } (b_{\vec{p}}^\dagger b_{\vec{p}} + c_{\vec{p}}^\dagger c_{\vec{p}})$$

Propagator for complex scalar field

$$\langle 0 | \phi^\dagger(x) \phi(y) | 0 \rangle \text{ for } x^0 > y^0$$

→ particle is created at y

This will give the combination

$$\langle 0 | b_p b_p^\dagger | 0 \rangle \rightarrow \text{This implies particle is created at } y \text{ by the operation of } \phi^\dagger(y) | 0 \rangle$$

→ propagates to x and annihilates at x by the operation $\langle 0 | \phi(x)$

→ Antiparticle travels from x to y .

→ complex scalar field propagator for particle or antiparticle

$$\langle 0 | T(\phi(x) \phi^\dagger(y)) | 0 \rangle$$

$$= \theta(x^0 - y^0) \langle 0 | \phi(x) \phi^\dagger(y) | 0 \rangle + \theta(y^0 - x^0) \langle 0 | \phi^\dagger(y) \phi(x) | 0 \rangle$$

$\downarrow$ particle propagation from y to x $\downarrow$ Antiparticle propagation from x to y .

propagator $\frac{i}{p^2 - m^2 + i\epsilon}$

So far, we were working with Schrödinger picture, where $\phi(x) \neq \phi(x, t)$. Given $\phi(x)$, one can find out how $\phi(x)$ changes in Heisenberg's picture, where $\phi(x)$ also depend on t .

$$\phi_H(x) = e^{iHt} \phi(x) e^{-iHt}$$

$$= \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\sqrt{2\omega_p}} \left(a_{\vec{p}} e^{-i p_m x^m} + a_{\vec{p}}^\dagger e^{+i p_m x^m} \right)$$