

## BGO cosmics acquisition

29/08/25

# Reminder : data sets

After ALTO test beam several data sets for BGO are available :

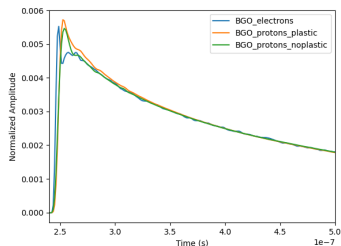
| Type                   | Wavecatcher conf. | Location   |
|------------------------|-------------------|------------|
| Electron source        | NOK               | IJCLab lab |
| Proton                 | OK                | ALTO       |
| Proton with attenuator | OK                | ALTO       |

- ▶ ALTO data has to be de-convoluted from the cable/amplifier effects

- ▶ Simultaneous fit :  
 $\tau_1 = 341 \text{ ns}$   $\tau_2 = 69 \text{ ns}$   
( $\chi^2 = 1.05$ )

| Dataset                   | $\tau_1$ % | $\tau_2$ % |
|---------------------------|------------|------------|
| Electron                  | 5.6        | 94.4       |
| Proton without attenuator | 5.3        | 94.6       |
| Proton with attenuator    | 6.4        | 93.4       |

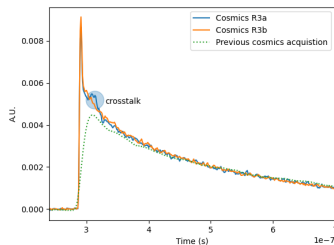
- ▶ Decision was taken to re-do cosmics muons with the new detector geometry



Comparison between the various data set mean decays (After applying the deconvolution)

# New acquisition

- ▶ Wavecatcher directly connected to the PM
- ▶ Coincidence with 2 external scintillators
- ▶ Annoying crosstalk solved with the BGO PM signal

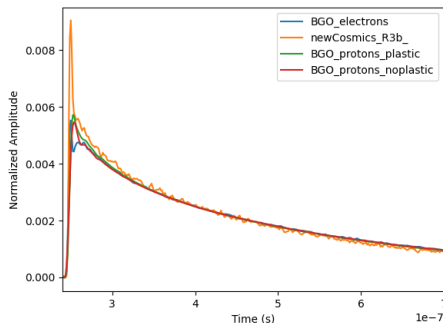


Comparison between the various cosmic acquisitions)

Difference with the previous acquisition :

- ▶ The trigger was not correctly applied (self-trigger) polluting the muons sample
- ▶ The amplifier has a too low bandwidth ( $\approx 50$  MHz)
- ▶ Issue with the wavecatcher config. (spotted at ALTO)

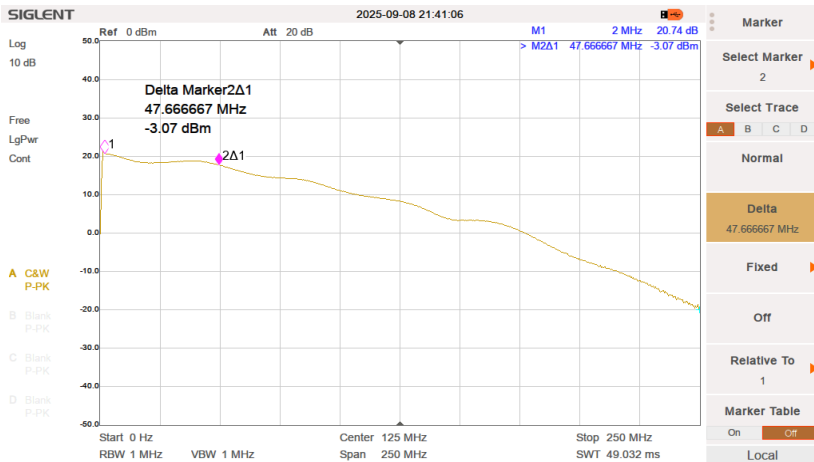
# Comparison to the ALTO data



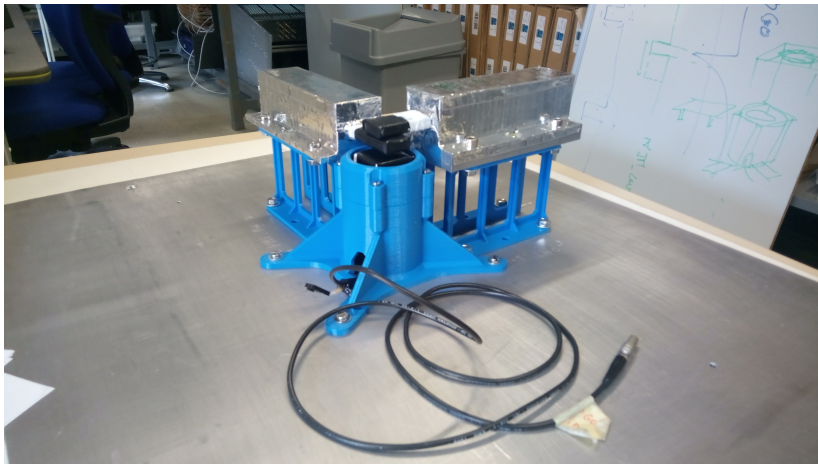
- Simultaneous fit :  $\tau_1 = 342$  ns  $\tau_2 = 71$  ns

| Dataset                   | $\tau_1$ % | $\tau_2$ % |
|---------------------------|------------|------------|
| Electron                  | 5.9        | 94.1       |
| Proton without attenuator | 5.5        | 94.5       |
| Proton with attenuator    | 6.5        | 93.5       |
| Muons                     | 9.3        | 90.7       |

# SPARE SLIDES



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