

# **SPARTAI: Solar PArticles Radiation Tool by AI**

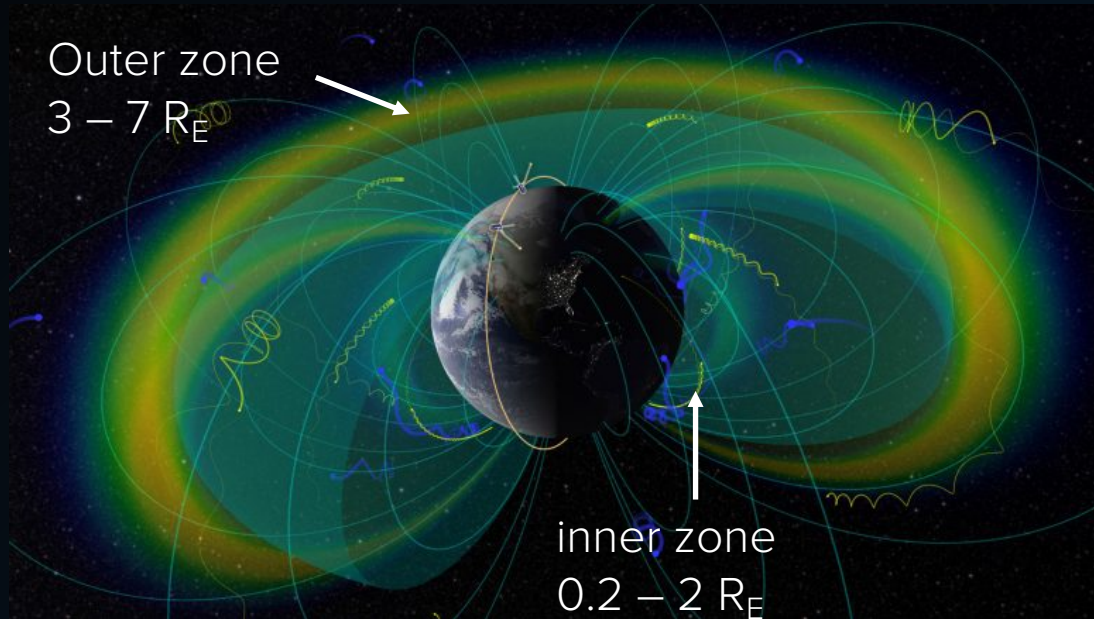
An AI-based forecasting pipeline for  
energetic electrons in Earth's radiation belts

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Research conducted at IRAP, UT, as part of the Augura Space project supported by Inria

# Van Allen Radiation Zone

Region of space filled with energetic charged particles ( $e^-$ ,  $p^+$ ) that are captured by Earth's magnetic field.

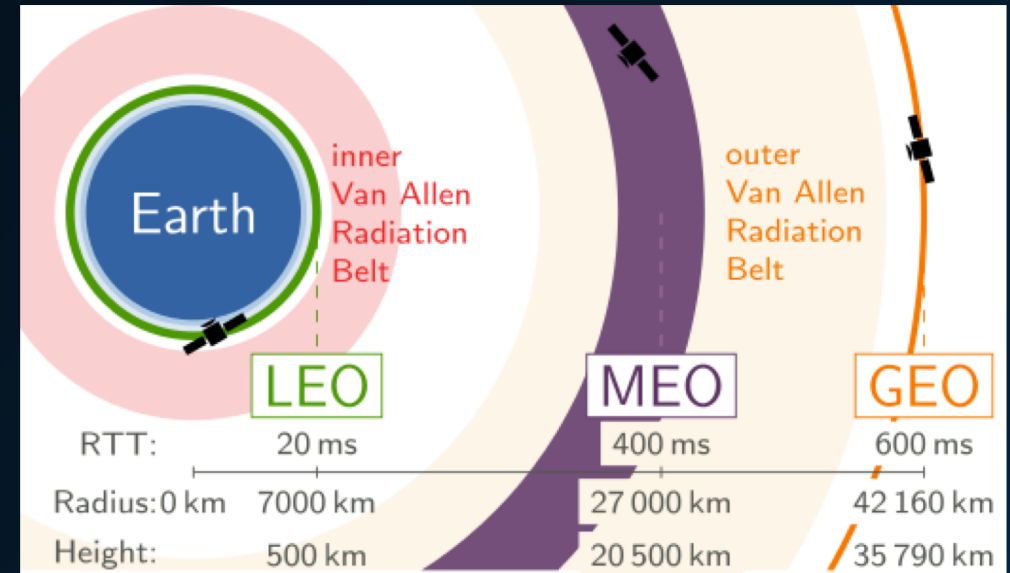


Its properties vary according to the solar activity.

**Challenge:** How the charged particles are accelerated to very high energies of several million e<sup>-</sup> volt (MeV) ?

**Radiation hazard** for manmade space systems :

- Long-term electron dose effects
- Surface and deep-dielectric charging

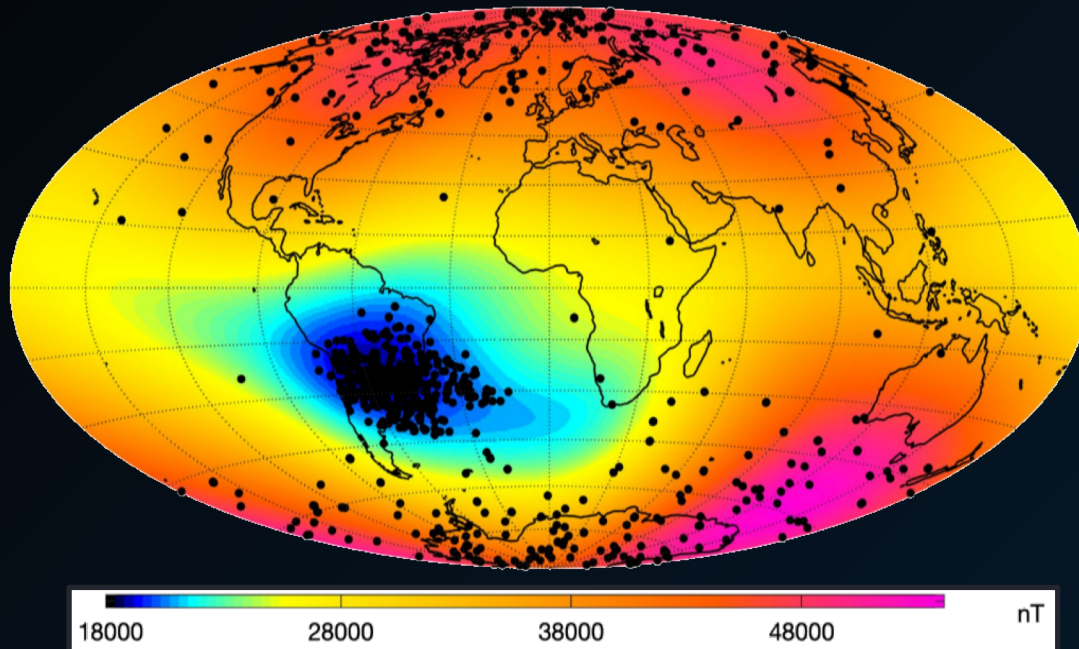


As of Oct 2025 (<https://orbit.ing-now.com/>)

~14000 space objects in orbit including 6961 in LEO, 521 in GEO, and 191 in MEO.

# Satellite operational anomalies

**Black dots:** spacecraft anomalies registered onboard SWARM in 2013 – 2019. Colour: magnetic field strength



**South Atlantic Anomaly:** weakened magnetic field region pushing radiation belt particles towards the Earth.

ASTRONOMY

## Software error doomed Japanese Hitomi spacecraft

Space agency declares the astronomy satellite a loss.

Nature, 2016

News

## Orbital Blames Galaxy 15 Failure on Solar Storm

Intelsat, 2010

**Predicting the variations of radiation belt energetic particle fluxes are thus critical**

- NOAA: Linear Prediction Filter Analysis of Relativistic Electron at GEO orbit (Baker et al. 1990)
- ESA space weather demo product: Radiation Belt Forecast And Nowcast (RB-FAN) by ONERA.



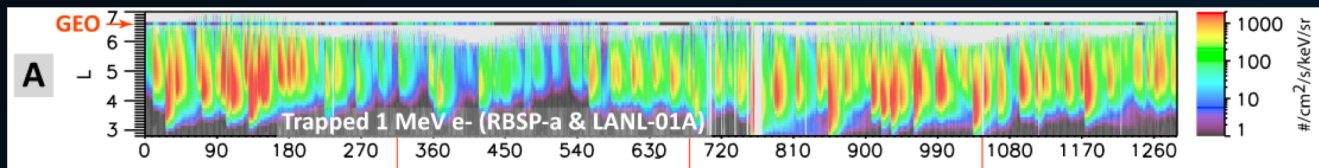
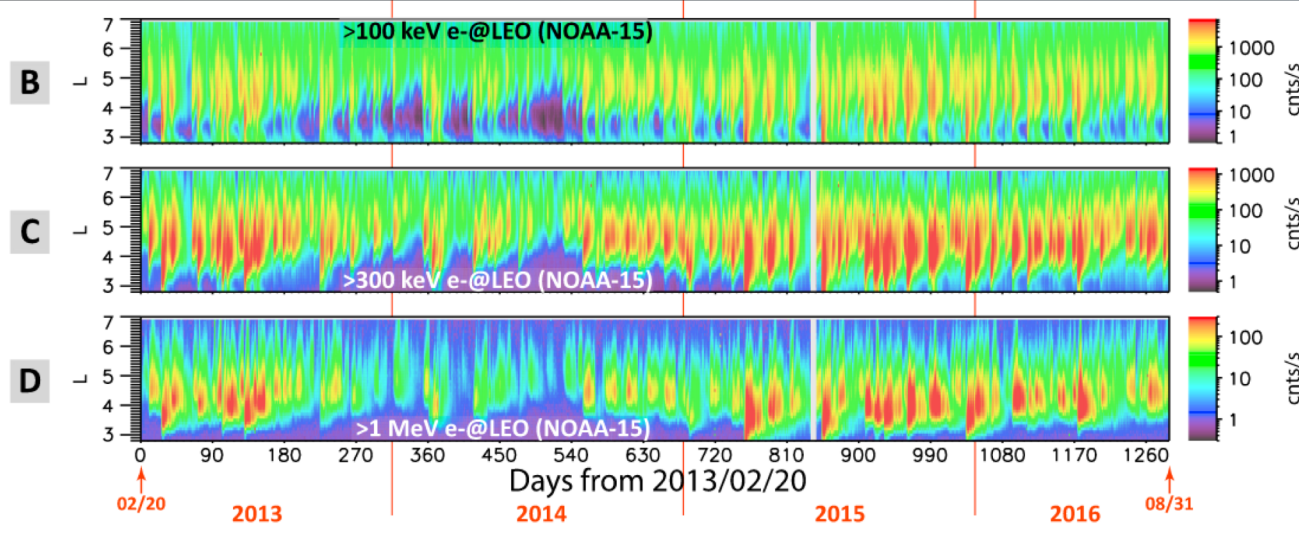
# Predicting energetic radiation belt electrons with AI

## Space Weather

RESEARCH ARTICLE  
10.1029/2018SW002095

### PreMevE: New Predictive Model for Megaelectron-Volt Electrons Inside Earth's Outer Radiation Belt

Modeling MeV electron fluxes from precipitating electrons



## PreMeVE model series:

- ← Chen+2019: linear prediction filters
- Pires de Lima+2020: linear regression; neural networks: FNN, 1D-CNN, LSTM.
- Feng+2024: CNN with Transformer layers

## Data

- NOAA-15 Polar Operational Environmental Satellite (POES):
  - e<sup>-</sup> flux measurements
- + solar wind speed upstream of Earth

Benchmarked with the **Van Allen probe** (2013-2016) and **LANL's GEO** satellite.




# SPARTAI – developing our own pipeline

(similar to PreMeVE)

POES at  $t-\delta t, \dots t-1$

(equivalent to LANL s/c)

GOES at  $t-\delta t, \dots t-1$

  $V_{sw}, \dots, \text{at } t-\delta t, \dots t-1$

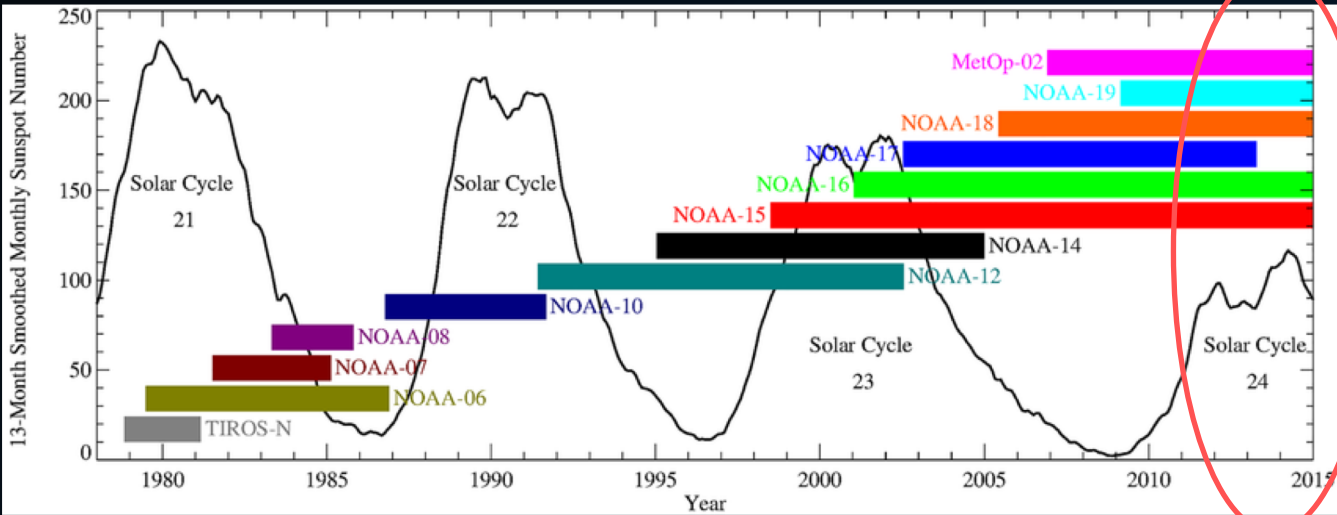
ML algorithms

POES 1-MeV  
at  $t, t+\Delta t$

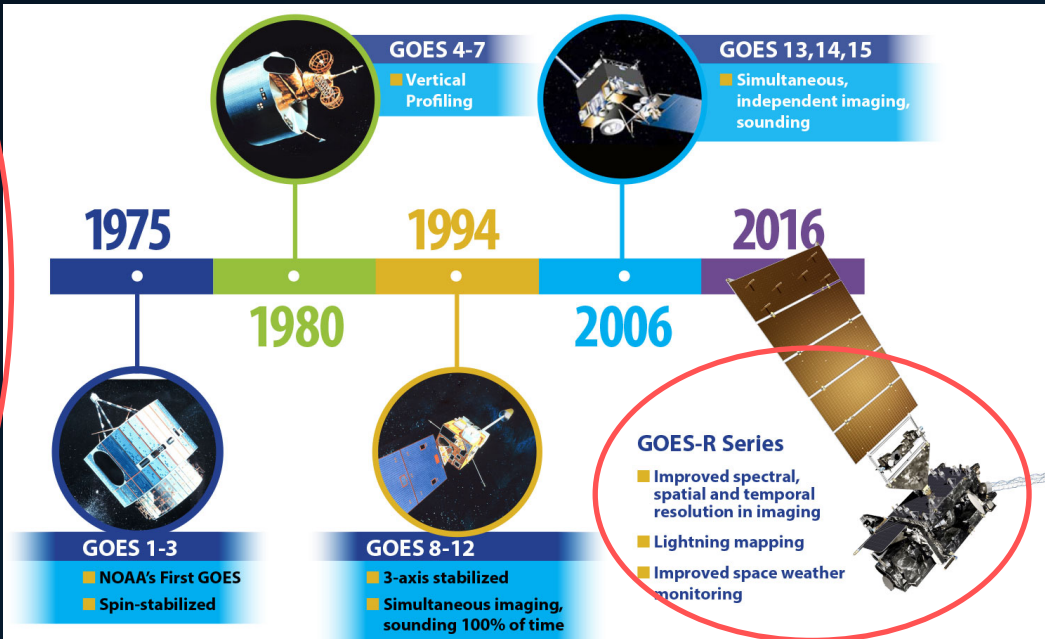
(equivalent to the  
Van Allen probe)

➤ Taking advantage of large datasets from **Space Environment Monitor (SEM)** onboard weather satellites

## Polar Operational Environmental Satellite (POES) / MetOp



## Geostationary Operational Environmental Satellite: GOES

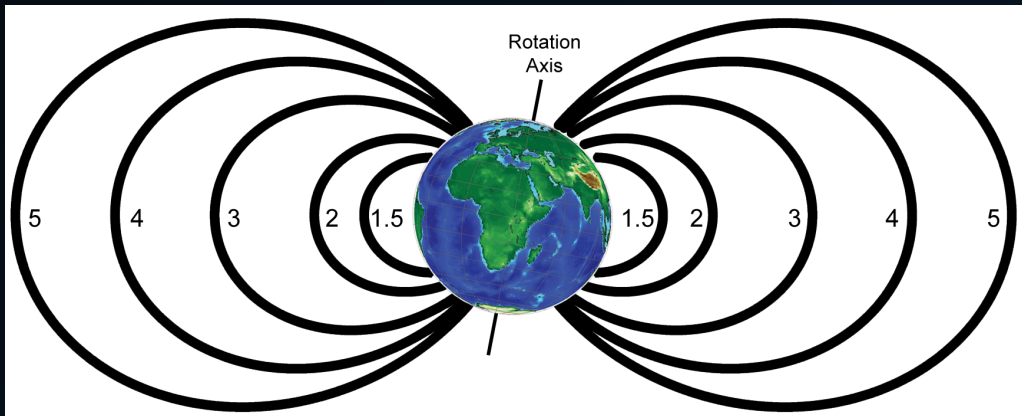


# SPARTAI – data preparation

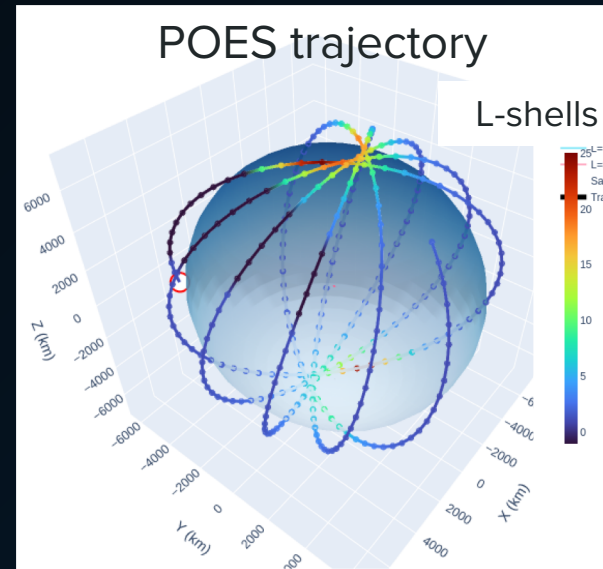
## Inputs (inspired by PreMeVE2.0):

- Solar wind speed (**V**) upstream of Earth
- GOES-15 SEM:
  - **E1** ( $> 0.8$  MeV)  $e^-$  flux
- POES-15 SEM-2:
  - **E2** ( $> 100$  keV) & **E3** ( $> 300$  keV)  $e^-$  fluxes
  - **P6** ( $> 1$  MeV) flux, designed for  $p^+$  but mostly contaminated by 1-MeV  $e^-$  flux (Chen+2019).

Coordinates L-value: the set of magnetic field lines which cross the magnetic equator at a number of  $R_E$



L-shells between 2.8 and 6.0, then at 6.6



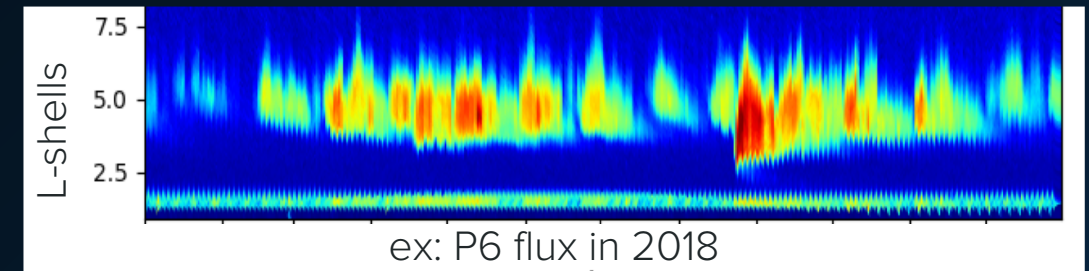
## Data processing

Original SEM 2s data:  
**F** (time, lat, lon, ..., fluxes)



Round, group by time and  
L-value using average

Tabular data: temporal  
flux variation for each L



## Benchmarking against PreMeVE2.0

- Modeling 1-MeV  $e^-$  flux, **L by L**, using linear regression and neural networks
- **Train** Feb 2013 – June 2015 (67%); **Validate** July – Dec 2015 (14%); **Test**: Jan – Aug 2016 (19%).

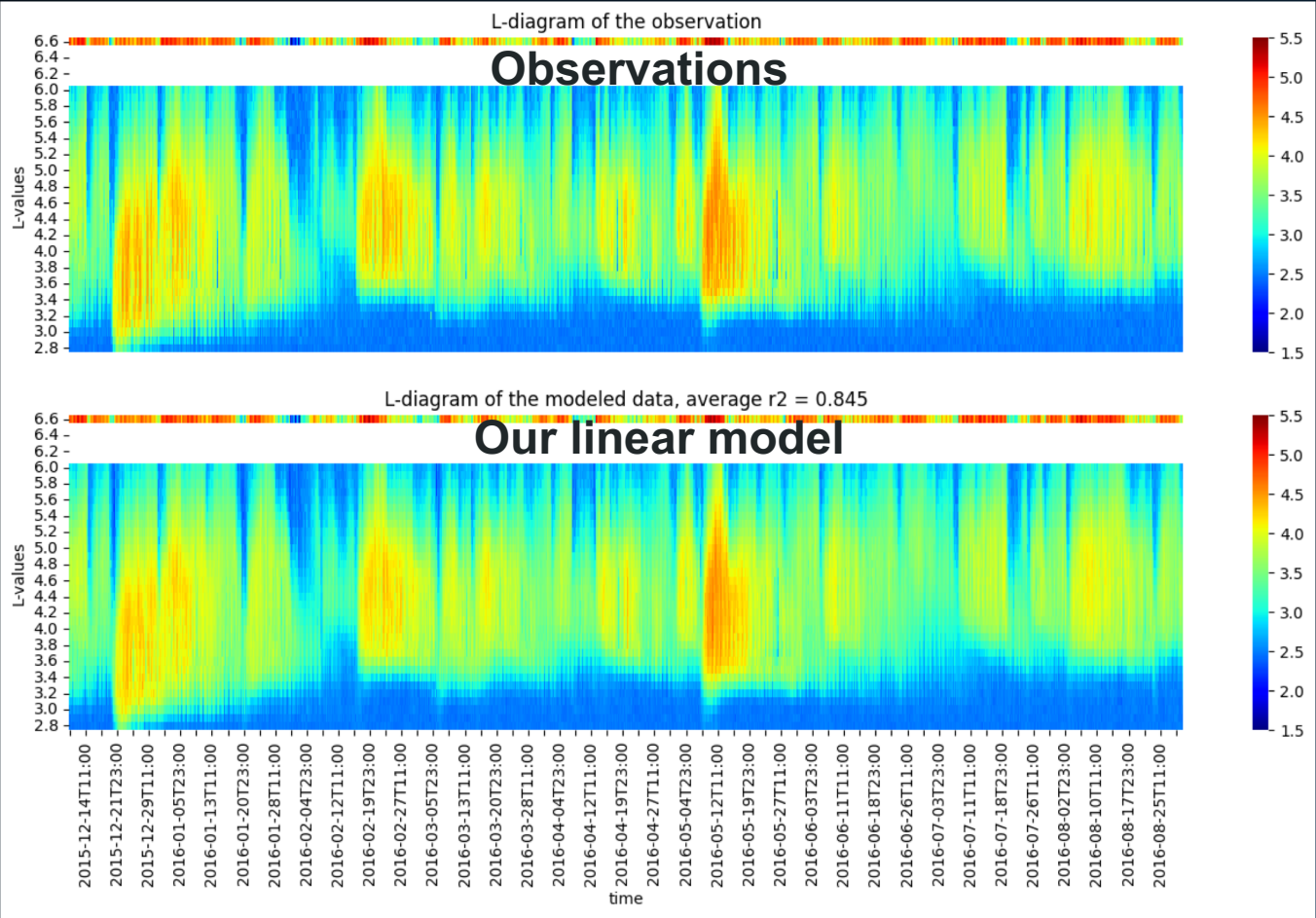
# SPARTAI – benchmarking results

Average R<sup>2</sup> of the test data

Model	PreMeV-E 2.0 $2.8 \leq L \leq 6.0$	<b>SPARTAI</b> $2.8 \leq L \leq 6.0$
Linear Reg	0.887	<b>0.845</b>
FNN-64-32-elu	0.879	<b>0.841</b>
LSTM-128	0.886	<b>0.844</b>
CNN-64-relu	0.890	<b>0.828</b>

Model	PreMeV-E 2.0 $L = 6.6$ (GEO)	<b>SPARTAI</b> $L = 6.6$ (GEO)
Linear Reg	0.587	<b>0.861</b>
FNN-64-32-elu	0.524	<b>0.821</b>
LSTM-128	0.589	<b>0.768</b>
CNN-64-relu	0.363	<b>0.601</b>

*Our models are about 5% worse on average, but they are 40% better for GEO.*



The 1-MeV e<sup>-</sup> flux variation during geomagnetic storms (e<sup>-</sup> penetration to low L) are reproduced, despite some different details.



# SPARTAI – development

Data Train + validate: 2013 – 2023 (11 years); test: Jan – June 2024

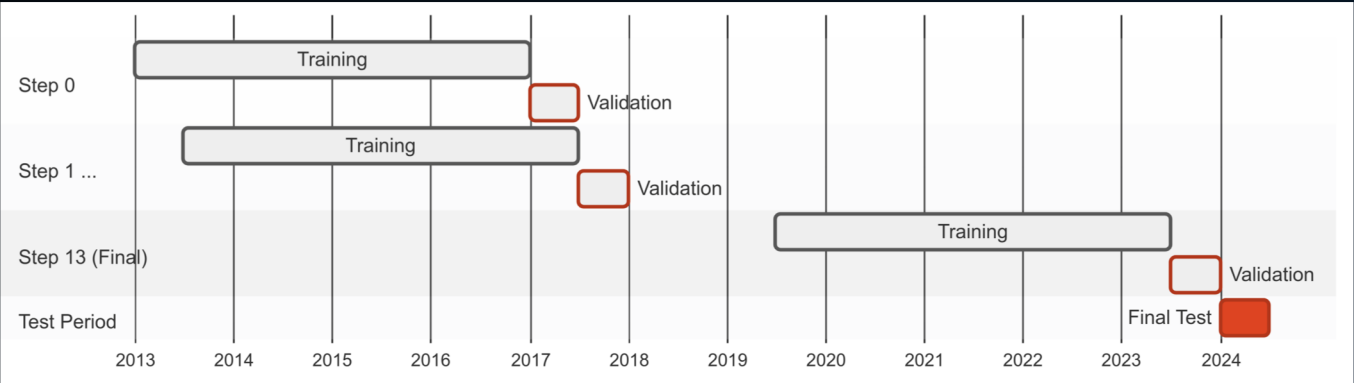
Missions	Monitor	2012	2013	2014	2015	2016	2017	2018	2019	2020	2021	2022	2023	2024	2025
POES n15	LEO														
MetOp-B	LEO														
GOES-16	GEO														
GOES-15	GEO														
OMNI	near L1														

(solar max)

(solar min)

(solar max)

**Training strategy:** Walk forward validation



- i. A model is trained from scratch and then validated.
- ii. The training is **resumed** from previous step **with the shifted window**, and then validated with the adjacent data.
- iii. Repeat until arriving at the test data.

*A model is saved at every step. The best model is chosen for “test”.*

**More solar & geomagnetic activity indicators:**

SW speed (V), density (N), magnetic field (Bz), storm disturbance index (Dst), Kp

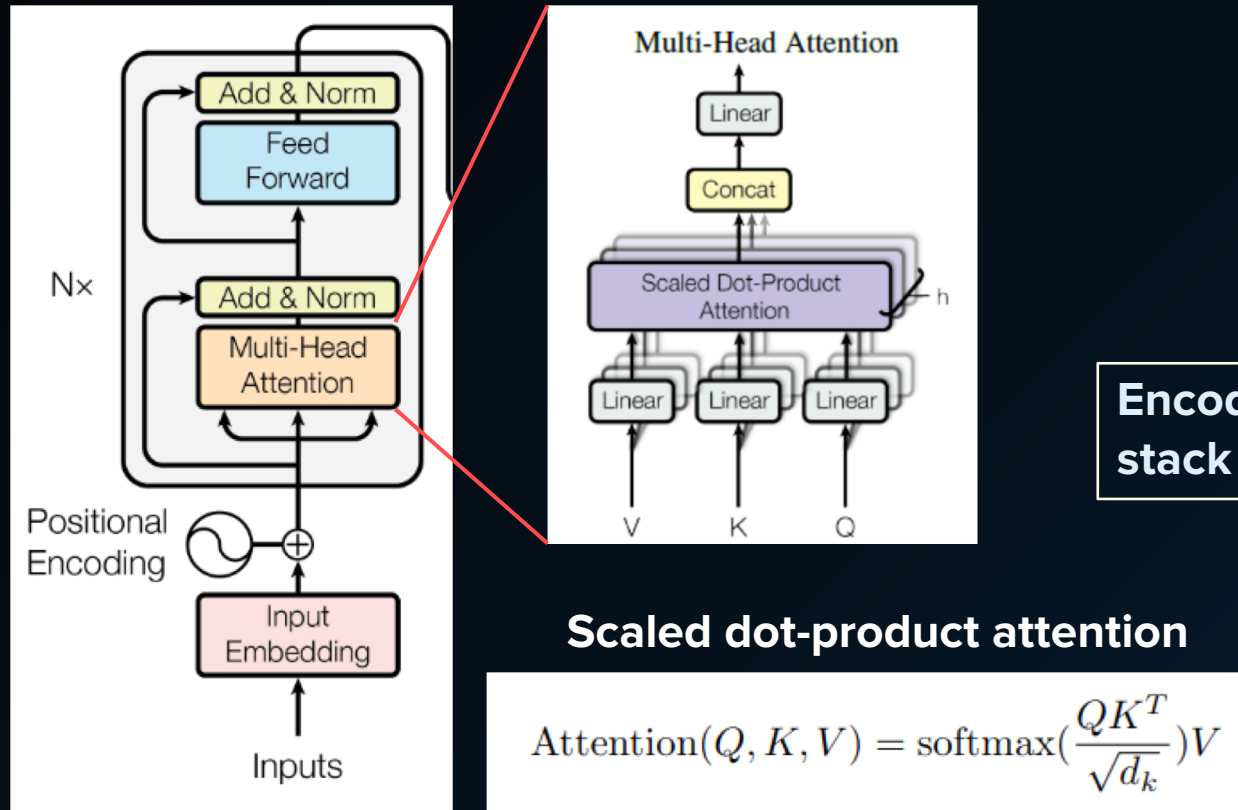
## More AI models

- Ridge regression
- XGBoost
- Transformer Encoder
- Time series foundation model (decoder only).

# The Transformer – model architecture

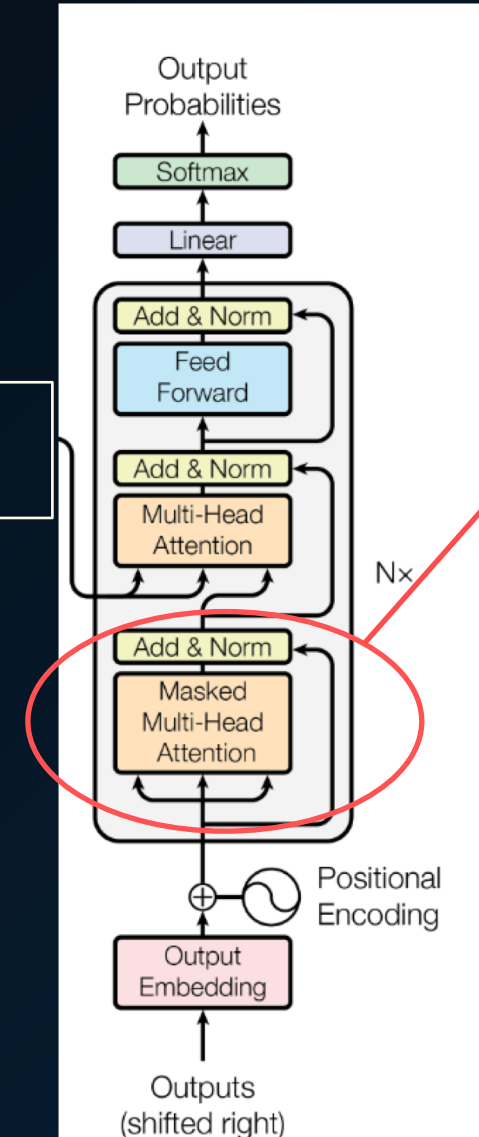
[Vaswani et al. 2017: Attention is All You Need]

**Encoder** maps an input sequence  $(x_1, \dots, x_n)$  to  $(z_1, \dots, z_n)$



**Decoder**

“Generation of the outputs”



This masking ensures that the predictions for position  $i$  can depend only on the known outputs at positions  $< i$ .

The output embeddings are offset by one position

**Query (Q):** the element seeking info, e.g., each word

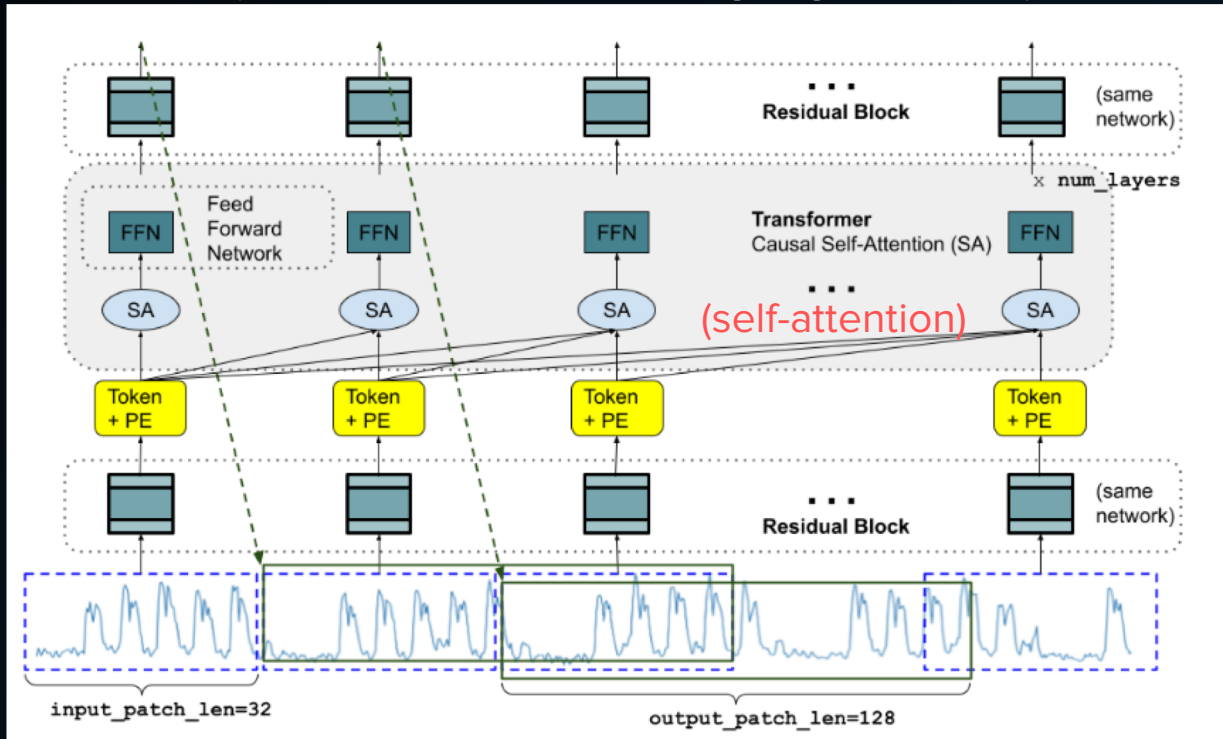
**Key (K):** signposts to locate important elements

**Value (V):** values carry the info, determining importance

# Time series foundation model

A DECODER-ONLY FOUNDATION MODEL FOR TIME-SERIES  
FORECASTING *Das+2024, ArXiv*

Decoder style attention model with **input patching**  
(similar to a token in language models)



**Pretrained using a large time-series corpus**, e.g., Google Trends (<https://trends.google.com>), Wiki pageviews, traffic, weather, electricity, synthetic, M4, ..., with different granularities from hourly to monthly.

**SPARTAI - our application** (TimesFM 2.4)

- Zero-shot (default) forecast
- Linear regression on covariates (our input variables, shifted in time), then we apply the foundation model on the residual



# SPARTAI – our development results

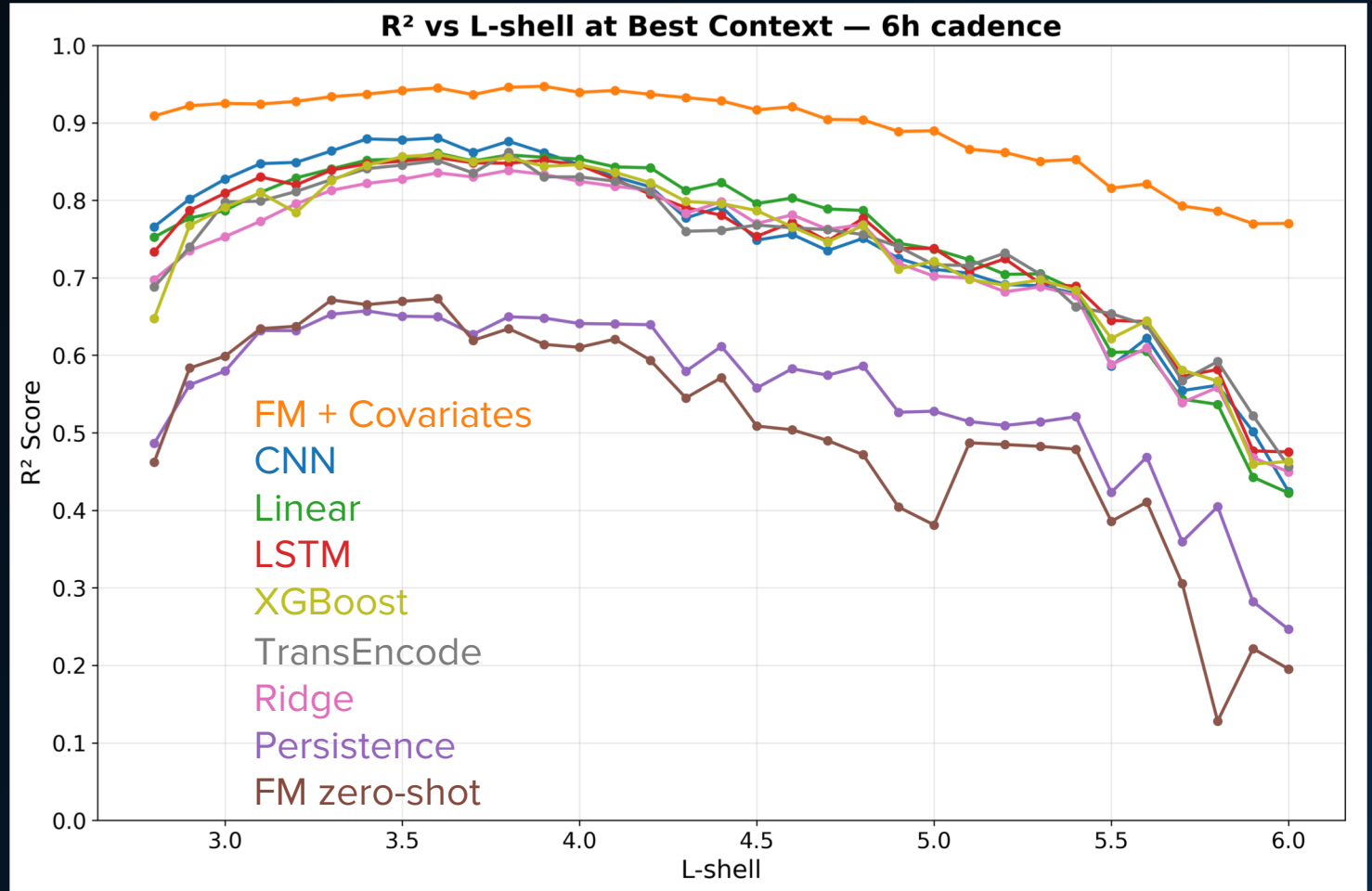
Linear, CNN, LSTM are identical to the benchmarking experiment.

Persistence: use last value as prediction.

Other models: simplest working architectures with no fine-tuning.

We performed experiments to find best context length (history length used for prediction). No  $R^2$  improvement for  $> 24$  hours.

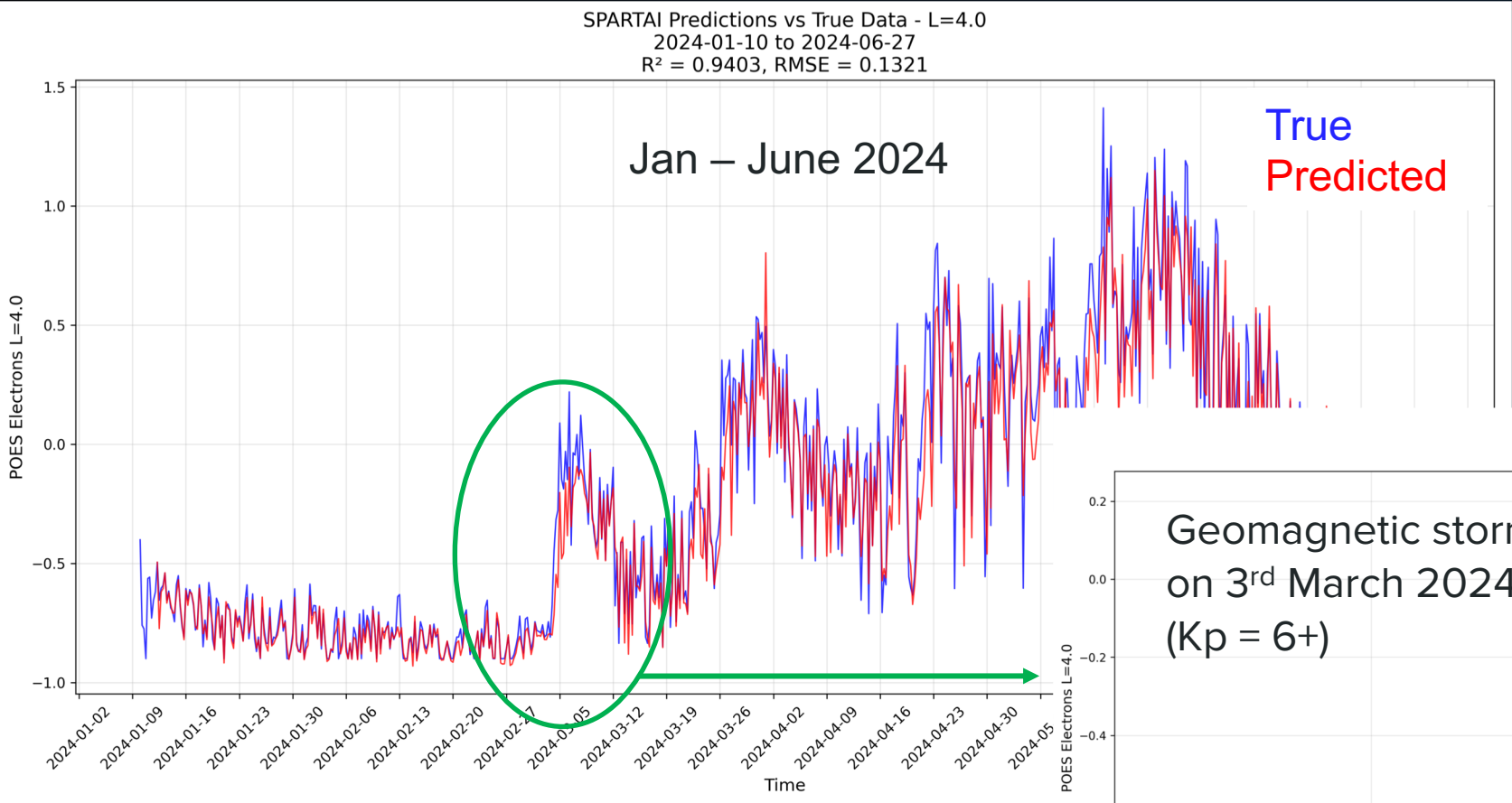
$R^2$  score for 6h forecast (Jan – June 2024)



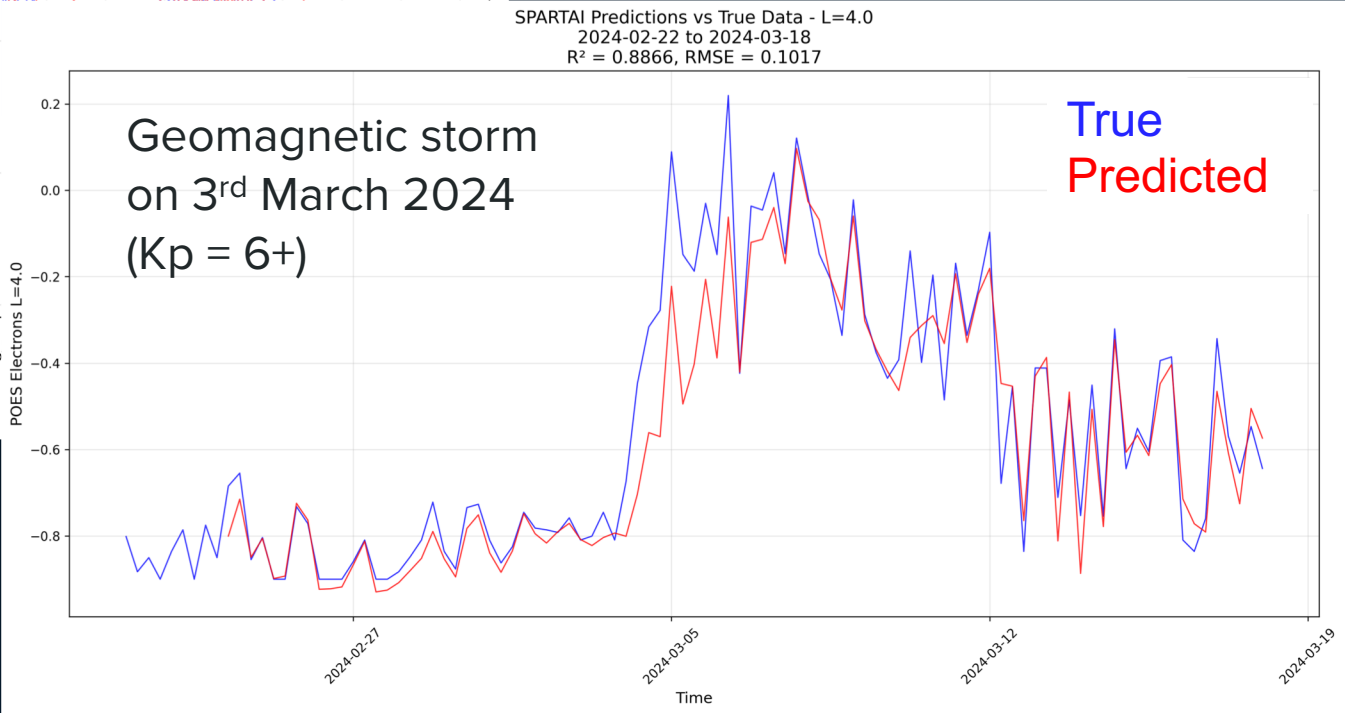
*The TimesFM + covariates outperforms all models for all L-shells, reaching  $R^2$  over 0.9 for  $L < 4.6$ .*

# SPARTAI – inference example

## TimesFM + covariates 1-MeV e<sup>-</sup> flux prediction for L=4.0



Some underestimated peaks and delay during 1-MeV e<sup>-</sup> flux enhancement during the storm



Large-scale variations including 1-MeV e<sup>-</sup> flux enhancements during storms are reproduced

# Classic ML approaches vs FM application

Similar to PreMeVE, we find that the **linear regression** model is the most successful. => The relationship between dynamics of **trapped 1-MeV e<sup>-</sup>** and the **precipitation e<sup>-</sup>** is dominated by **linear components**.

## Classic approach



- 1: Model  $M$  fit once on ( context , targ ) pairs in the entire train dataset
- 2: Fitted model makes forecast on last context LinForecast =  $M$  ( context\_b )

- The training is done on the 11 years of data (with walk forward validation). Model is used with the latest context for a prediction.

## Introduction of FM model

- Zero-short prediction yields very poor results



- We handle the linear part first, then compute the residuals

$$R(i) = M(\text{context}_i) - \text{targ}_i \quad \text{for } i < b$$

- The pretrained model is used to estimate the residuals (nonlinear).



# Summary and perspectives

**We develop a prototype pipeline for 1-MeV electron flux prediction for the outer radiation belt.**

## Benchmarking:

- We develop our own equivalent models (Linear, FNN, CNN, LSTM) to PreMeVE 2.0.
- We train, validate, and test with our defined (equivalent) dataset
  - Slightly poorer results to PreMeVE 2.0 (poorer data quality?)
  - Linear regression works best.

## Development:

- We add more models (Ridge, XGBoost, Transformer Encoder), as well as apply TimesFM.
- We train and validate with more data (1 solar cycle), and test in Jan – June 2024 (solar max)
- ✓ Linear fitting on latest context + nonlinear fitting with TimesFM works best

*How do we explain this physically?*

Perspectives: improve time cadence and forecast horizons, produce derived products.

Back-up slides

# Data processing

**Original POES data** (2s): geographical coordinates (lat, lon, alt), L-value, magnetic local time, flux measurements

**Processed POES data:** hourly averaged flux at each L for each energy channel ( $191 * 3 = 573$  columns)

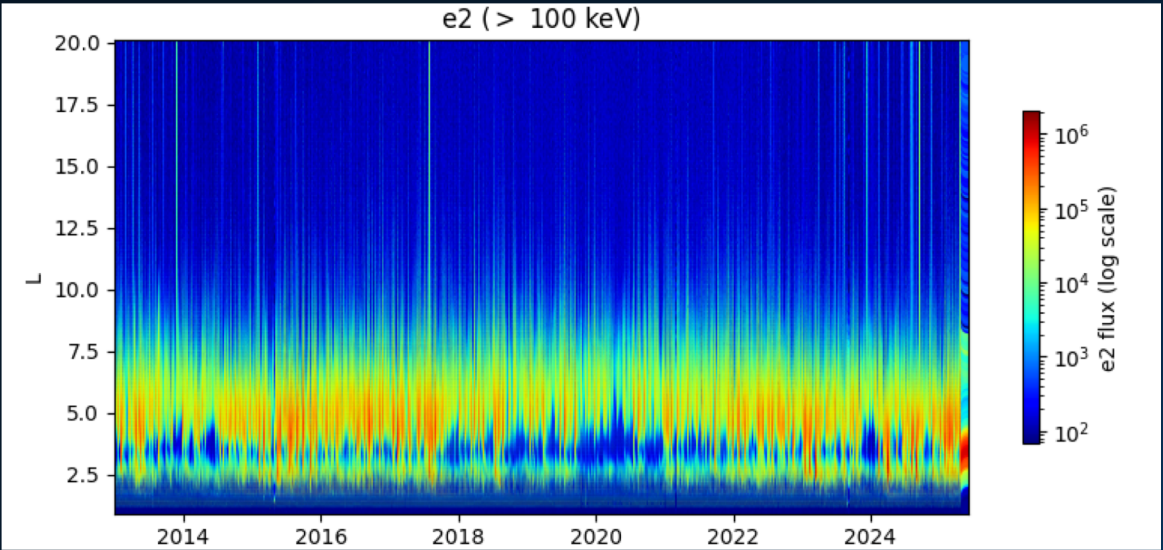
shape: (118\_953\_997, 9)

time	lat	lon	L	MLT	alt	e2	e3	p6
---	---	---	---	---	---	---	---	---
str	f64	f64	f64	f64	f64	f64	f64	f64
2013-01-01 00:02:39	48.373	80.639	2.07	5.34	819.2	69.44444	null	487.80487
2013-01-01 00:02:43	48.142	80.541	2.05	5.33	819.1	69.44444	null	731.70734
2013-01-01 00:02:45	48.027	80.493	2.05	5.33	819.1	69.44444	null	null
2013-01-01 00:02:53	47.564	80.301	2.01	5.31	818.9	69.44444	null	243.90244
2013-01-01 00:02:55	47.448	80.254	2.01	5.31	818.9	138.88889	133.33333	null
...	...	...	...	...	...	...	...	...
2025-06-03 07:22:39	81.07	105.081	18.78	14.12	811.7	486.11108	533.3333	null
2025-06-03 07:22:41	81.106	104.344	18.87	14.09	811.7	694.4444	533.3333	243.90244
2025-06-03 07:22:43	81.14	103.601	18.96	14.06	811.7	416.66666	400.0	null
2025-06-03 07:22:45	81.173	102.852	19.06	14.03	811.7	763.88885	1066.6666	null
2025-06-03 07:22:47	81.205	102.098	19.15	14.0	811.7	694.4444	1066.6666	null

shape: (107\_063, 574)

time	e2_2.1	e2_2.0	e2_1.9	...	p6_19.9	p6_20.0	p6_18.8	p6_19.3
---	---	---	---	---	---	---	---	---
datetime[μs]	f64	f64	f64	...	f64	f64	f64	f64
2013-01-01 00:00:00	77.16049	89.285711	92.59259	...	null	null	null	null
2013-01-01 01:00:00	1192.63282	515.232953	261.99494	...	null	null	null	null
2013-01-01 02:00:00	3780.864051	3671.874833	1436.965754	...	null	null	null	null
2013-01-01 03:00:00	154.671713	120.192305	217.803024	...	null	null	null	null
2013-01-01 04:00:00	146.604934	138.888885	123.456789	...	null	null	null	null
...	...	...	...	...	...	...	...	...
2025-06-03 03:00:00	47611.349072	48069.055736	50791.01229	...	null	null	null	null
2025-06-03 04:00:00	7165.798437	3281.249985	86.805553	...	null	null	487.80489	243.90244
2025-06-03 05:00:00	32385.543852	28527.92126	10938.977039	...	null	null	null	null
2025-06-03 06:00:00	102150.460167	156109.7208	191520.826	...	325.20325	null	243.90244	243.90244
2025-06-03 07:00:00	17253.924391	9978.718239	7723.765115	...	null	null	null	null

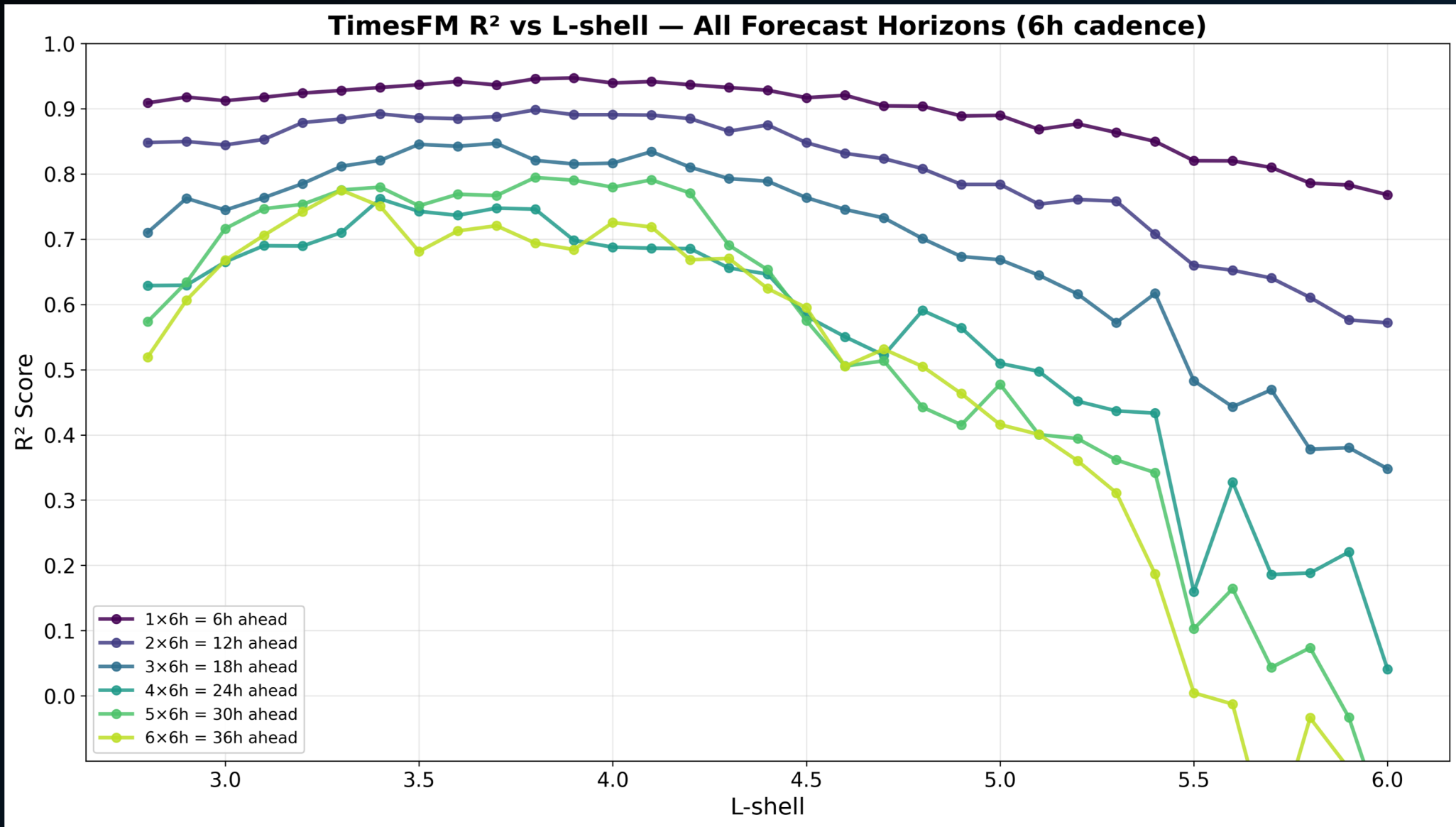
Visualization of L-time diagram of e2 flux



- Round from 2s to 1h; round L-value to one decimal place
- Limit  $1.0 \leq L \leq 20$ . Group by time and L using mean



# Performance of model forecasting with TimesFM + covariates



# Particles access to the Earth's magnetic field

Kallenrode, 1998  
D. A. Bryant 1993

**Rigidity** measures momentum of the particle. It refers to the fact that a higher momentum particle will have a higher resistance to deflection by a magnetic field.

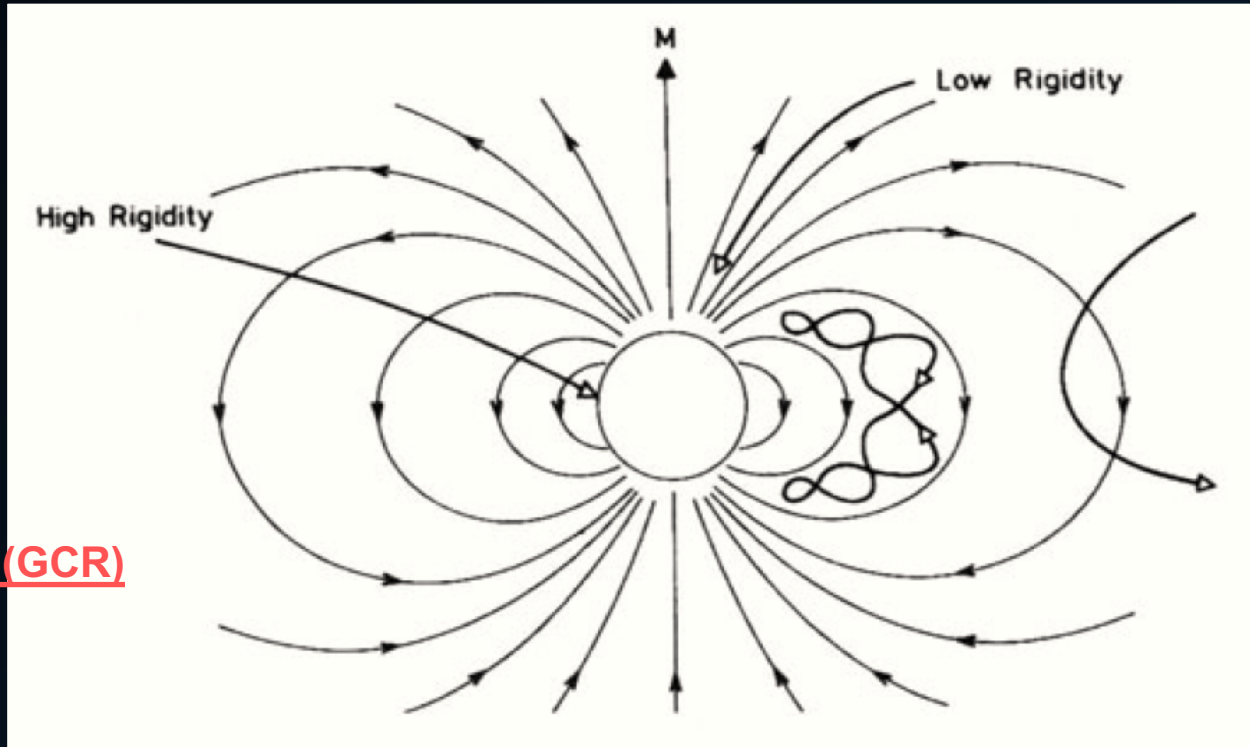
## High-rigidity particles

- able to traverse the magnetosphere
- no long-lived trapped particle components with high rigidities exists

## Galactic cosmic rays (GCR)

Depending on their direction of incidence, they either

- penetrate deep into the magnetosphere and interact with the upper atmosphere
- deflected back into space.



## Low-rigidity particles

Only at the polar cusps can these particles penetrate into the magnetosphere

Particles hitting the low-latitude magnetosphere from the outside perform half a gyro-orbit inside the magnetosphere and then are reflected back into space.

*Long-lived component: particles are trapped inside the radiation belts. Their motion is regulated by the adiabatic invariants.*