

Enhancing low energy reconstruction and classification in KM3NeT/ORCA with transformers

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On behalf of the KM3NeT collaboration



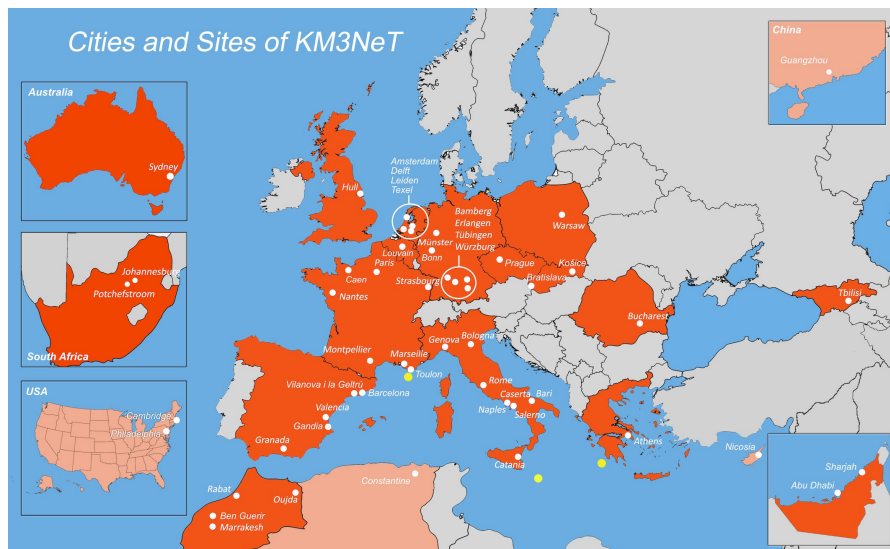
Overview

1. KM3NeT: neutrino telescopes
2. Reconstruction challenges in KM3NeT
3. Transformers for reconstruction
4. Multi-detector reconstruction in KM3NeT/ORCA
5. Conclusions

KM3NeT

KM3NeT is an **international collaboration**

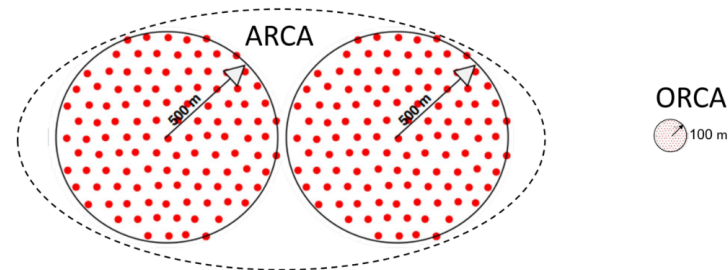
- 22 countries
- 65 partner institutes
- ~250 members



Two undersea neutrino telescopes

- **KM3NeT/ARCA**
 - Optimized for 1 TeV – 10 PeV
 - Identify high-energy neutrino sources in the Universe.
 - 36m vertical spacing and 90m horizontal spacing
- **KM3NeT/ORCA**
 - Optimized for 1 – 100 GeV
 - Determine the mass ordering of neutrinos.
 - 9m vertical spacing and 20m horizontal spacing

Currently under construction: **ORCA33 (29%), ARCA51 (22%)**



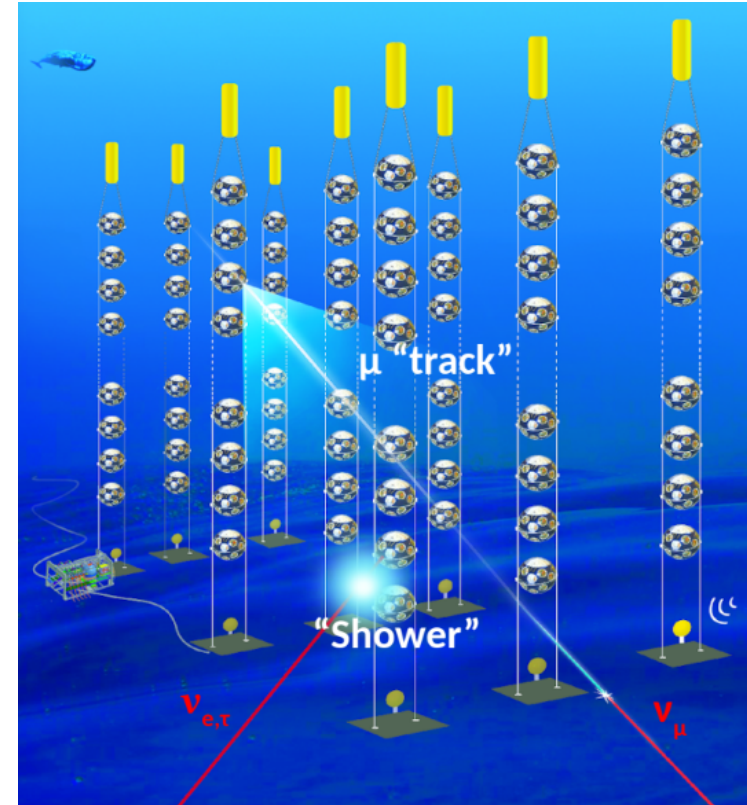
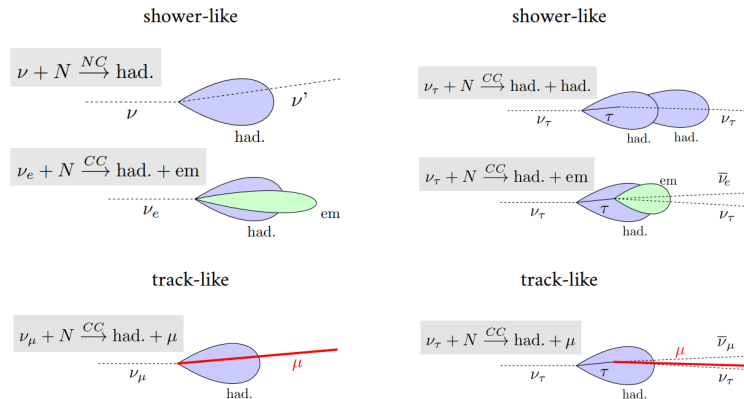
KM3NeT: neutrino telescopes

Same **technology**:

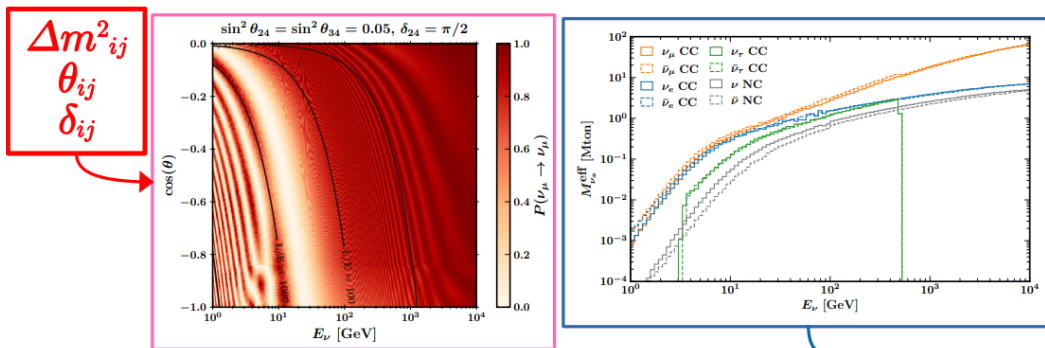
- 1 (2) building block(s) for ORCA (ARCA)
- 115 vertical detection units (DUs) per block
- 18 digital optical modules (DOMs) per DU
- 31" PMTs per DOM

Same **detection principle**:

*Light collection from **Cherenkov radiation** emitted by particles traveling faster than the speed of light in water*



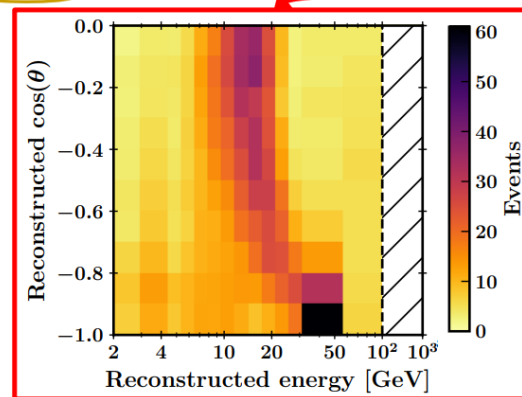
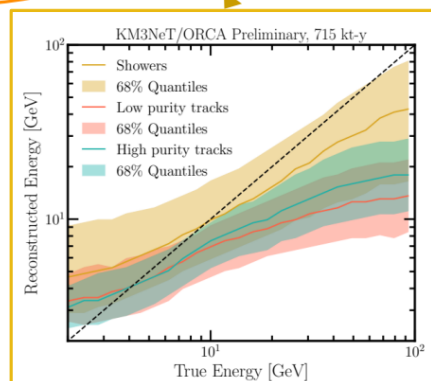
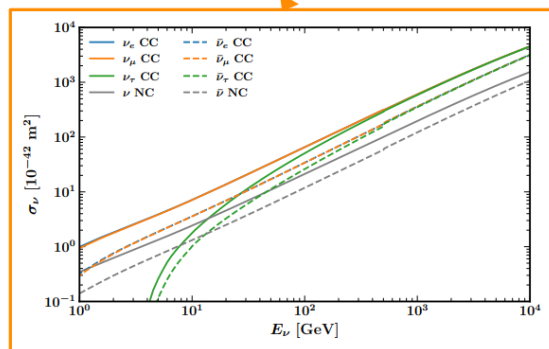
Importance of a good reconstruction



To produce good estimations of **predicted event rates** for an analysis, many pieces are considered: **fluxes**, **oscillations parameters**, **cross sections**, **effective masses** and **detector response**

One of the main sources of systematic errors in the analysis is the **detector response R_i** , modeled by the reconstruction algorithm.

$$\phi_{\text{atm}}^{\nu_y}(E_t, \theta_t) \times P_{\nu_y \rightarrow \nu_x}(E_t, \theta_t) \times \sigma_{\nu_x}(E_t) \times M_{\text{eff}}^{\nu_x}(E_t) \times R_i(E_t, \theta_t, \nu_x, E_r, \theta_r) \Rightarrow \mu_i(E_r, \theta_r)$$



Challenges of neutrino physics in ORCA

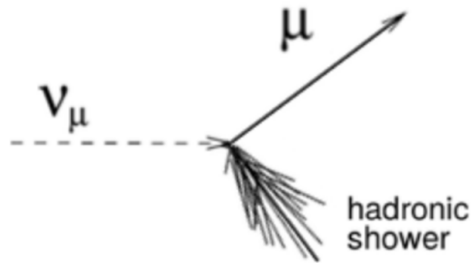
Reconstruct neutrinos from light

Maximum Likelihood Fit (MLF) algorithms

- Reconstruct under track or shower hypothesis
- Do not reconstruct the neutrino itself

nuT model beyond MLF

- Directly reconstruct neutrino
- Simultaneously reconstruct all hypothesis

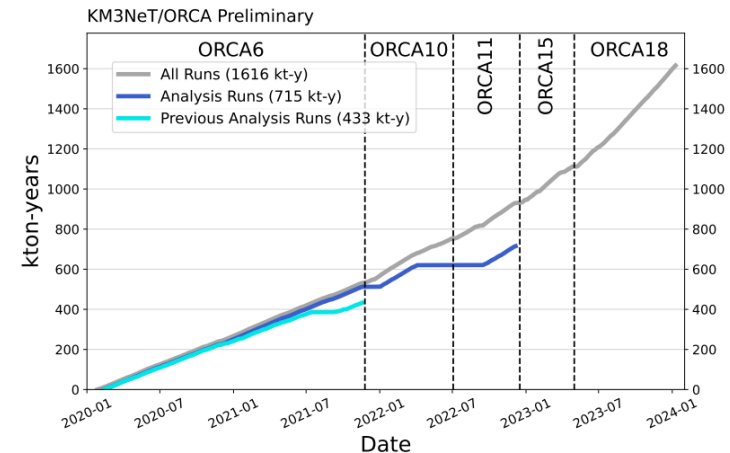


A growing detector

Realistic Monte Carlo samples are generated based on actual data-taking runs, capturing the complexities of deep-sea conditions

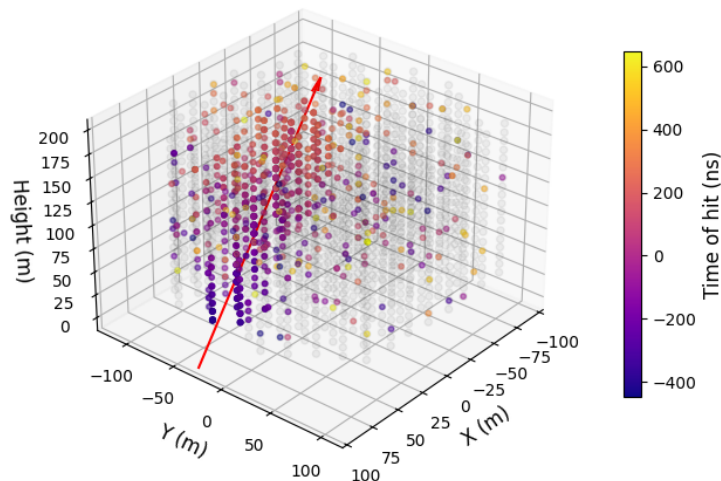
Need to use data and computing resources efficiently

Pre-trained models are leveraged to propagate information across configurations



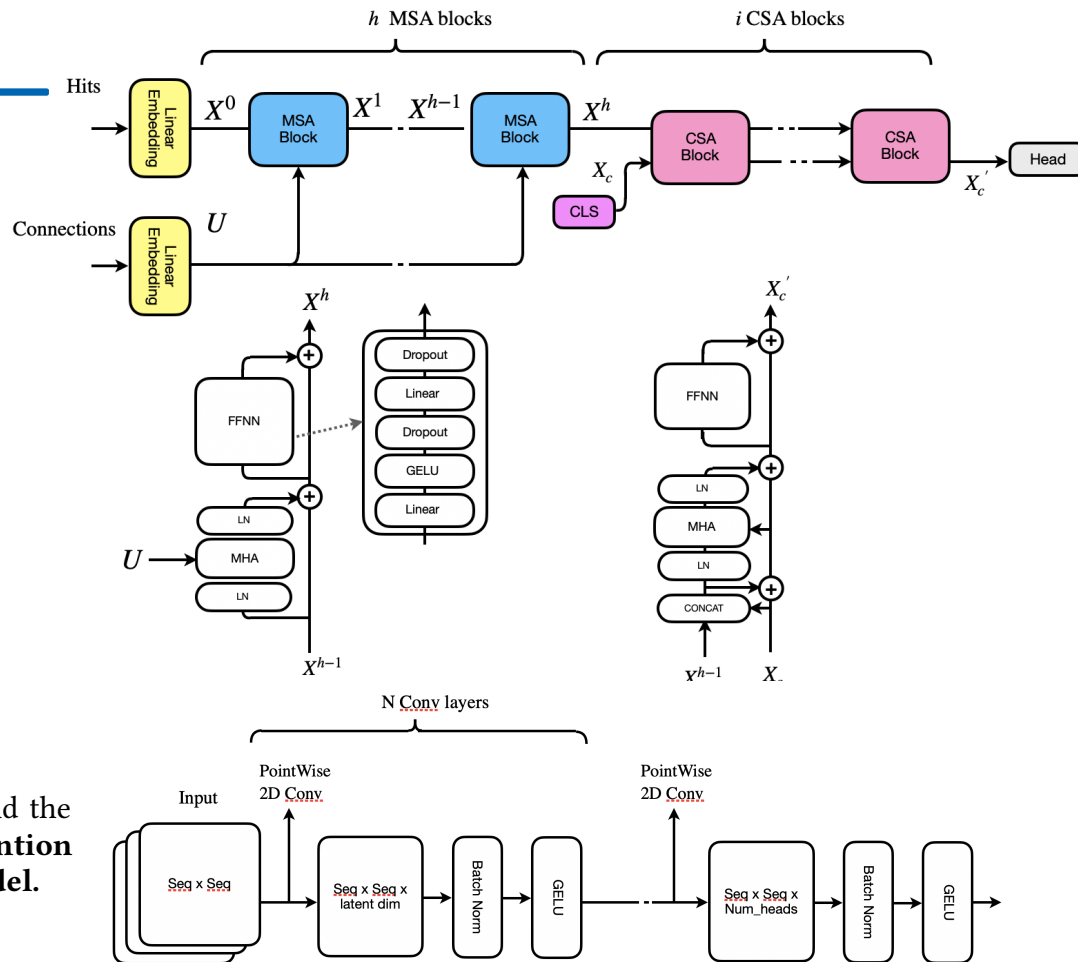
nuT architecture

The input data is the low-level hit information that composes the light pattern detected in the telescope, composed by **all type of hits**.



The information is processed in parallel by the transformer and the high-level information is extracted in the attention blocks. **Attention masks provide detector and physics knowledge to the model.**

The class token is an abstract representation of the event.

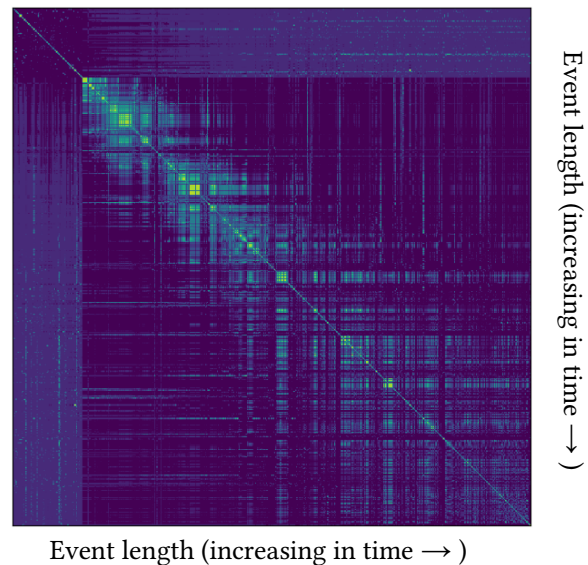
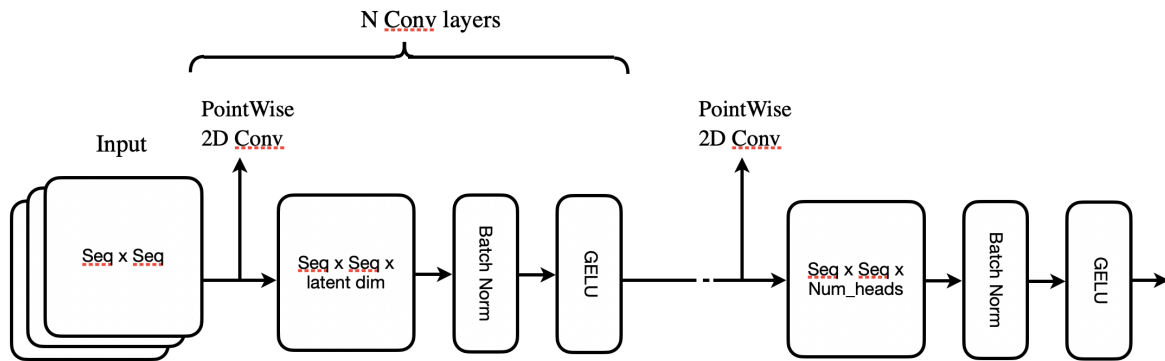


Attention masks

Detector and physics information introduced in the model through pairwise relationships

- **Causality:** space-time relativistic distance
- **Distance:** euclidean distance
- **Coincidence:** channel coincidence for DU, DOM and PMT in the **self-attention mechanism**

$$\text{Attn}(Q, K, V) = \text{SoftMax}\left(\frac{Q \cdot K^T}{\sqrt{d}} + U\right) \cdot V$$



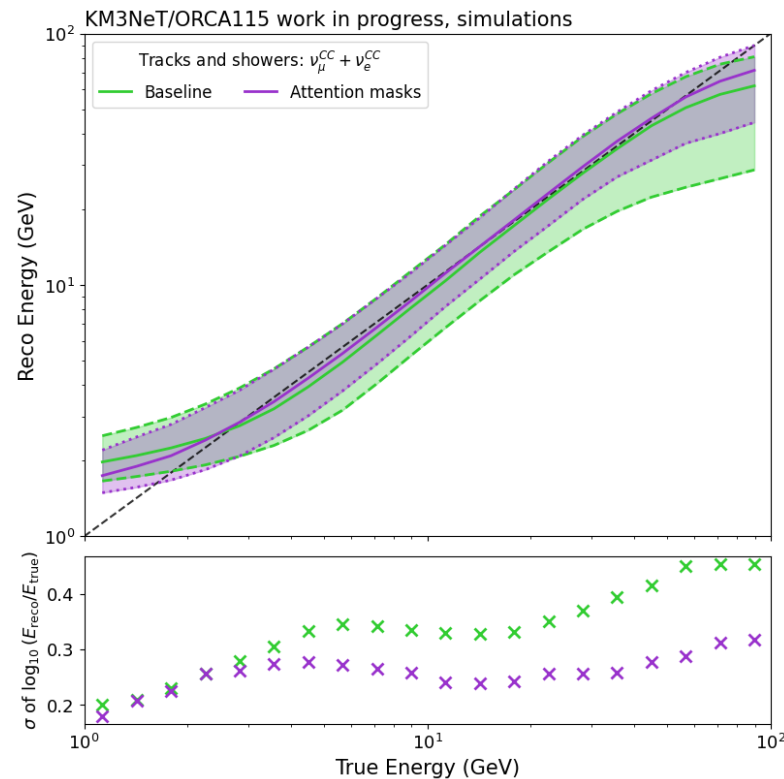
The attention mask show clear patterns of light pulses coming from a charged particle or from optical background → **The model understands the physics**

Attention masks

Detector and physics information introduced in the model through pairwise relationships in the **self-attention mechanism**.

Improvements in energy reconstruction:

- Better energy estimation at low energies.
- Better energy estimation at high energies.
- Less spreading of predicted energies.

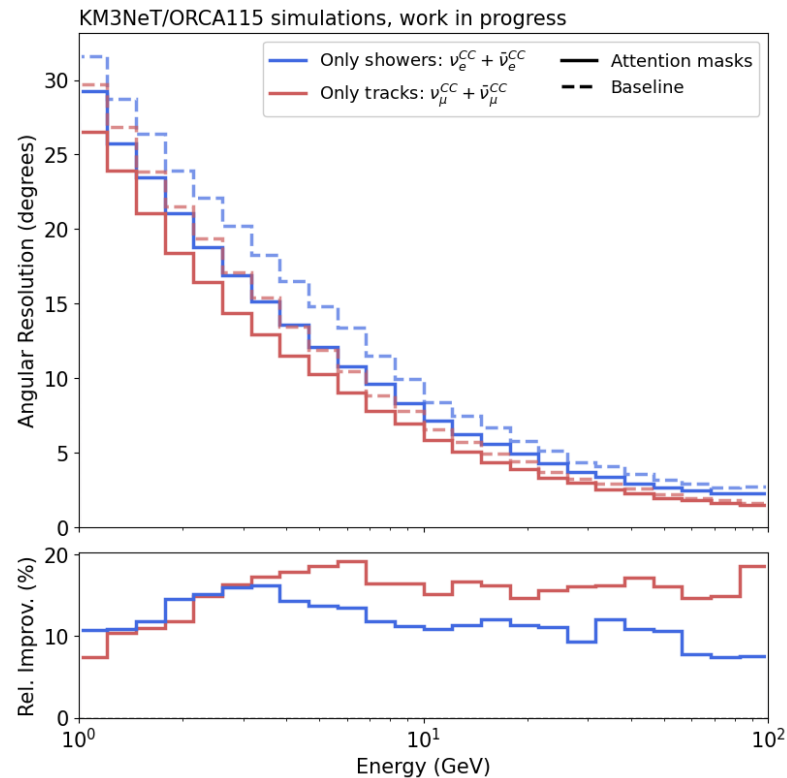


Attention masks

Detector and physics information introduced in the model through pairwise relationships in the **self-attention mechanism**.

Improvements in direction reconstruction:

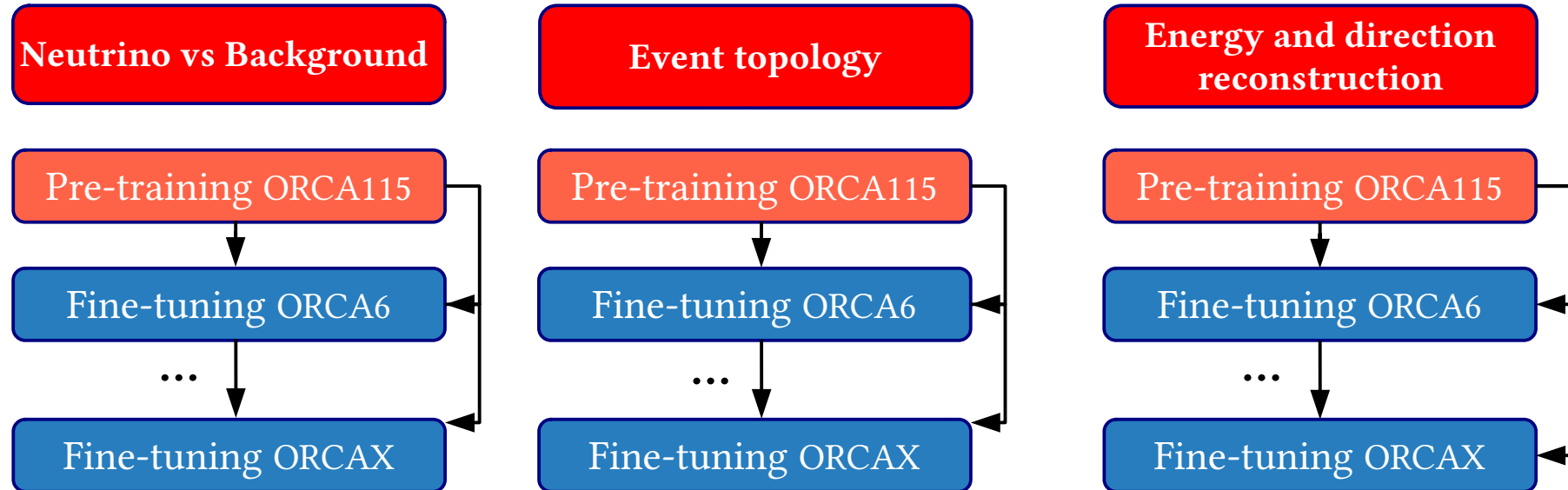
- Improvement of 10% for track-like events.
- ~10% improvement at low energies for showers.
- Above 15% improvement above 10 GeV for showers.



Learning approach for KM3NeT/ORCA

Motivation: the transformer is a language model

- KM3NeT/ORCA115 is the final detector, having all the possible neutrino physics encapsulated
- The information about KM3NeT/ORCA115 is used to understand our current detector
- All the knowledge will be retained when fine-tuning on other smaller configurations

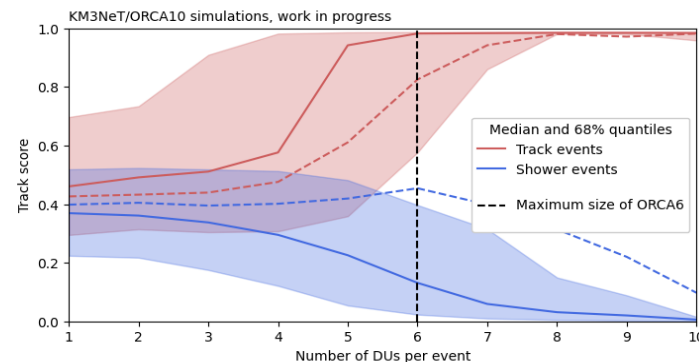
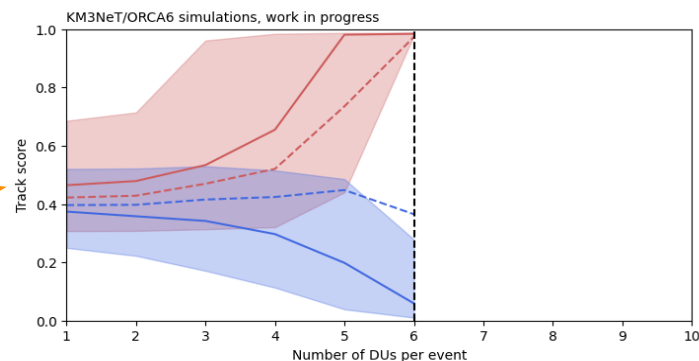
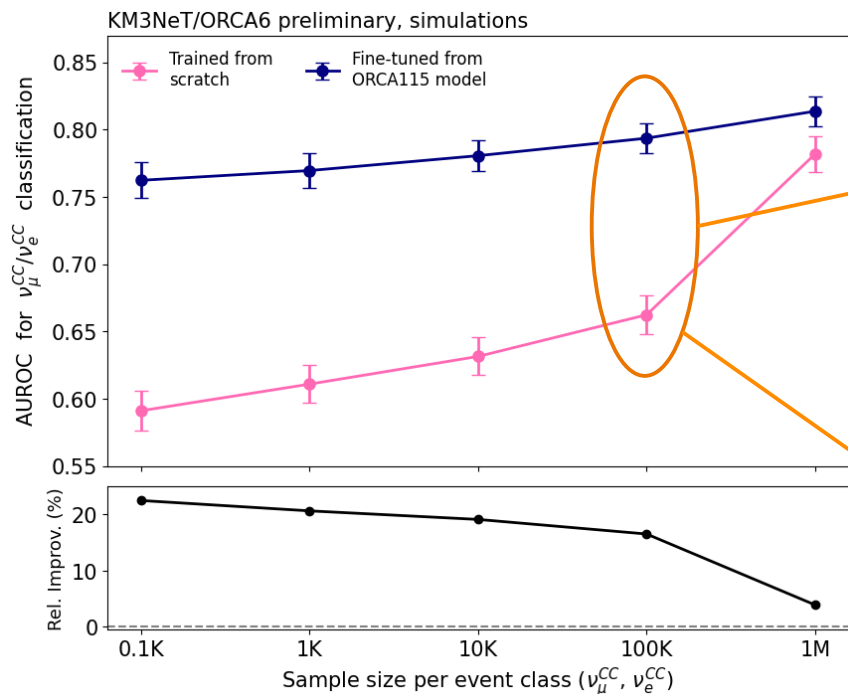


A growing detector

Realistic Monte Carlo samples are generated based on actual data-taking runs, capturing the complexities of deep-sea conditions

Pre-trained models are leveraged to propagate information across configurations w.r.t. **models trained from scratch**

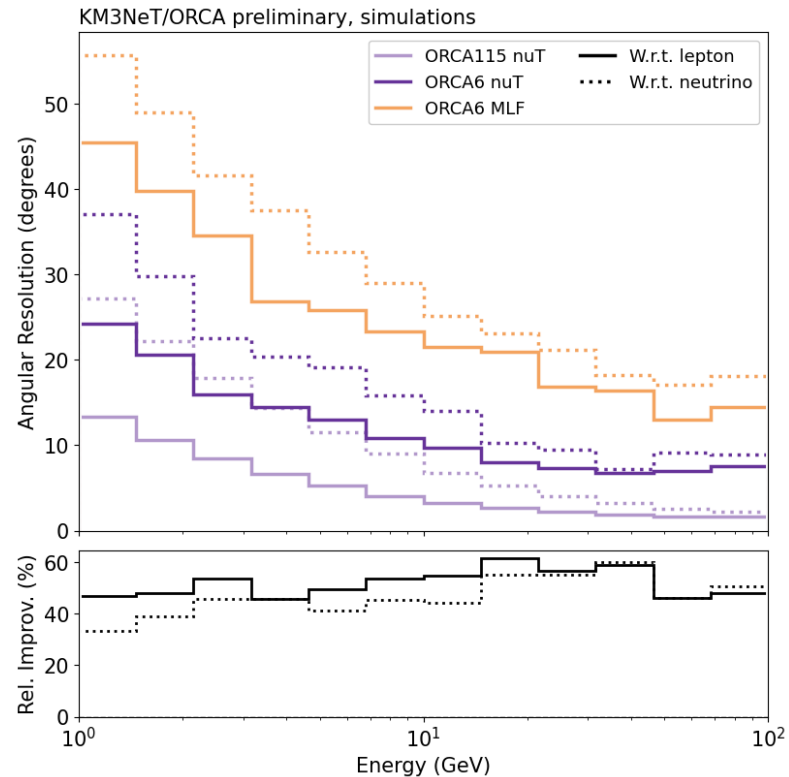
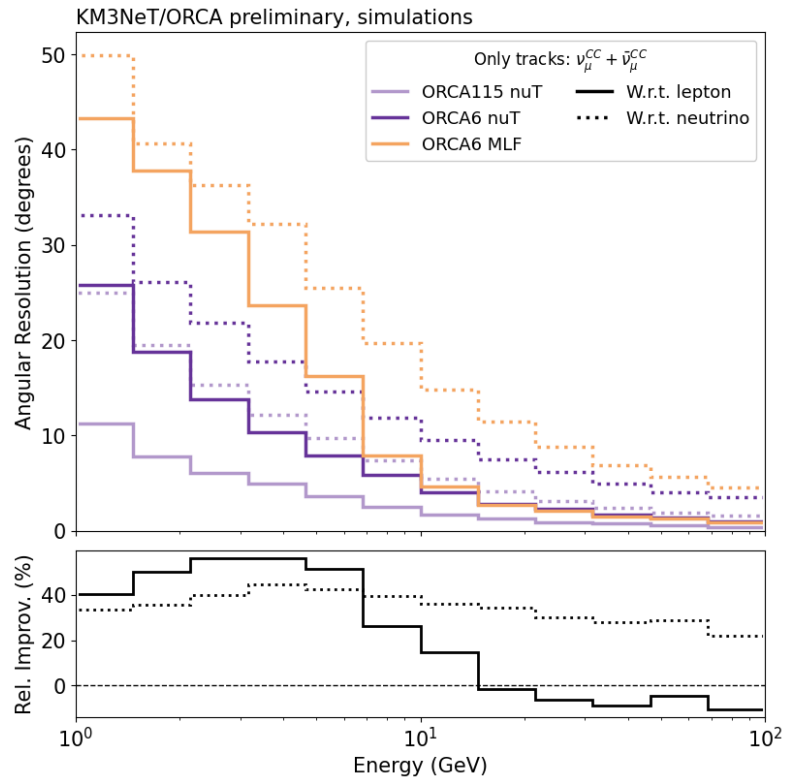
Better discrimination power when adding more lines



Reconstruction of neutrinos: angular resolution

MLF: $<10^\circ$ resolution above 10 GeV

nuT: improved resolution w.r.t. neutrino (+30% for tracks, +40% for showers)

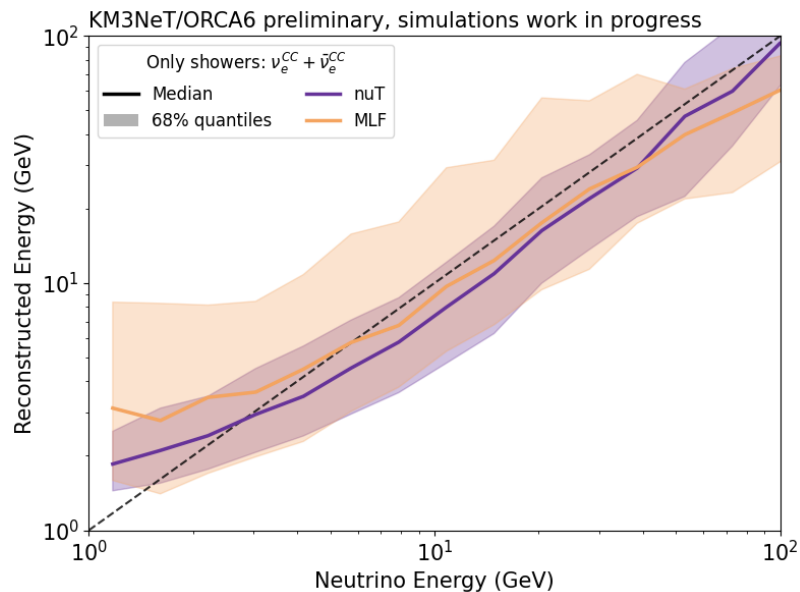
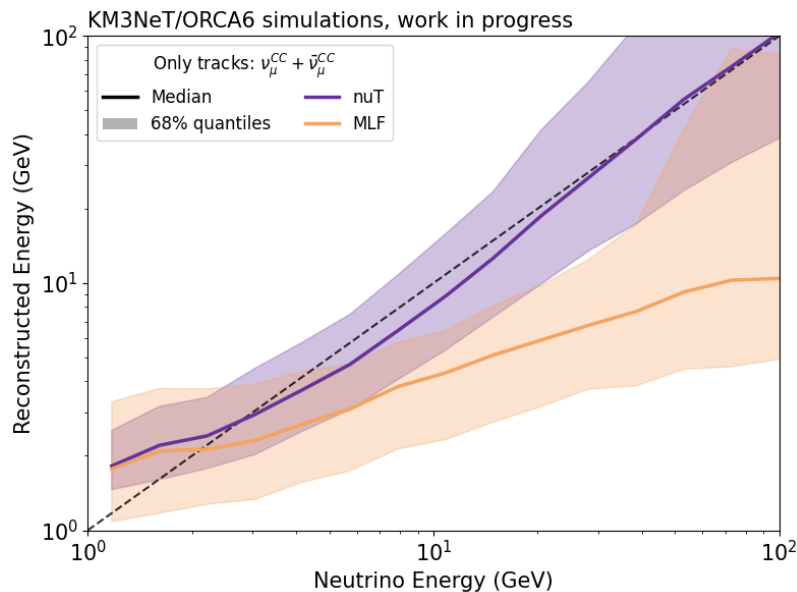


Reconstruction of neutrinos: energy estimation

MLF: underestimation in track reconstruction from missing hadronic component, compensated in pure shower events

nuT: accounts for visible and non-visible energy

Saturation effect at high energies due to event containment

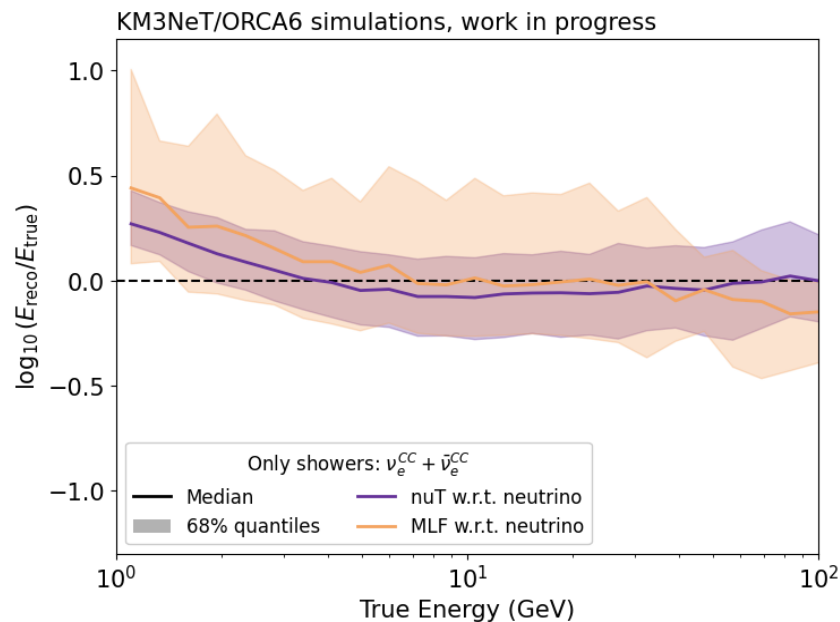
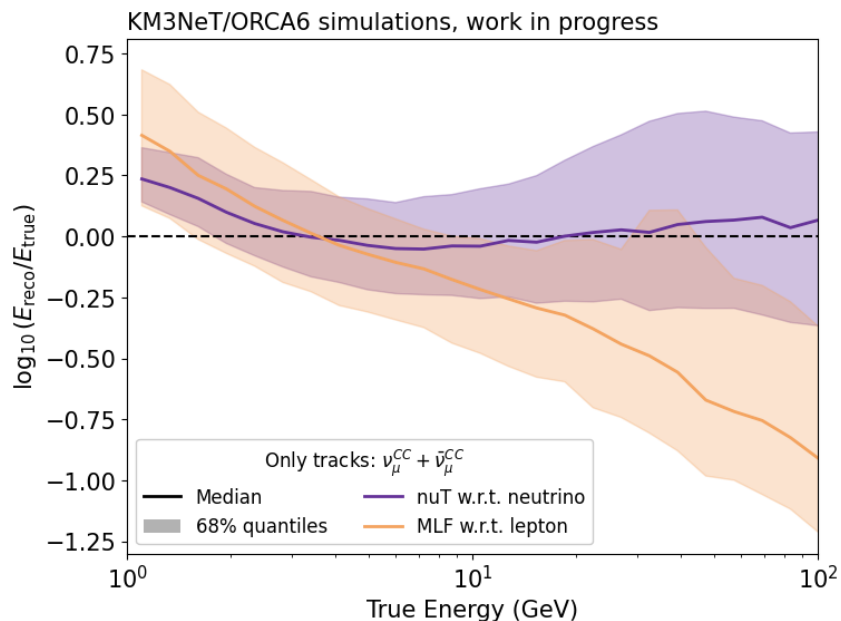


Reconstruction of neutrinos: energy resolution

MLF: underestimation in track reconstruction from missing hadronic component, compensated in pure shower events

nuT: compromise found when reconstructing tracks and showers simultaneously

Saturation effect at high energies due to event containment



Conclusions

Transformers in KM3NeT

- Reconstruct directly neutrino properties accounting simultaneously for track and shower hypothesis
- Attention masks:
 - Introduce physics and detector knowledge in the model
 - Boost training and increase performance

Propagate information across telescopes through fine-tuning

- Transformers are particularly effective to deal with small detectors and very limited data
- Information propagated across detectors
- Speeds up training and boosts model robustness

Thank you for your attention!