

Towards a better understanding of Giant Molecular Clouds' evolution

I L \wedge N C E



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Giant Molecular Cloud (GMC) / 巨大分子雲



Birthplaces of stars

How : by collapse of gas

Virial equilibrium : $2E_{\text{therm}} + E_{\text{grav}} = 0$

$$E_{\text{therm}} = \frac{3Mk_B T}{2\mu m_H} \quad E_{\text{grav}} = \frac{-3GM^2}{5R}$$

Jeans mass :
$$M_J = \left(\frac{5k_B T}{G\mu m_H} \right)^{\frac{3}{2}} \left(\frac{3}{4\pi\rho} \right)^{\frac{1}{2}}$$

Fragmentation until ρ high enough to form a star₂

NGC 253

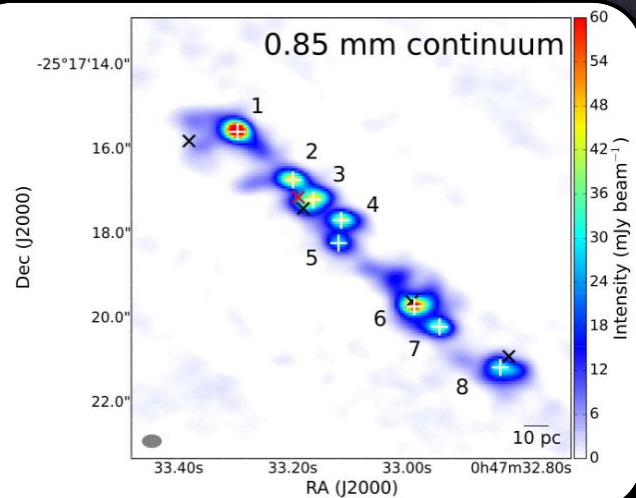
Starburst galaxy

3.5Mpcs

R.A. $00^{\text{h}}43^{\text{m}}33.1^{\text{s}}$
Dec $-25^{\circ}17'18''$

Radio observation

<https://arxiv.org/pdf/1710.01432> (Ando et al. 2017)

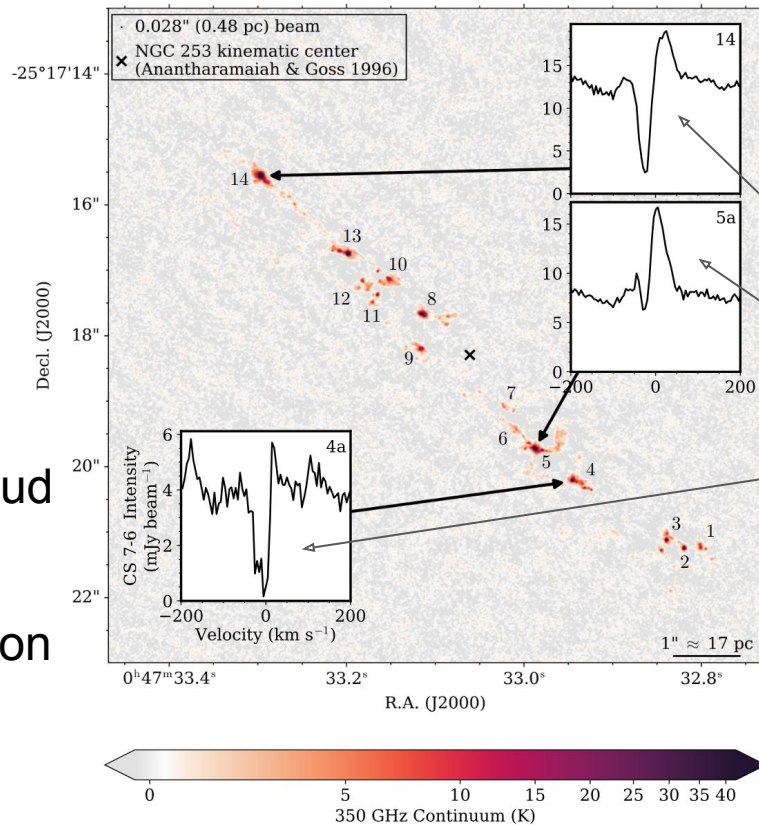


Optical observation

Mount Lemmon SkyCenter/University of Arizona

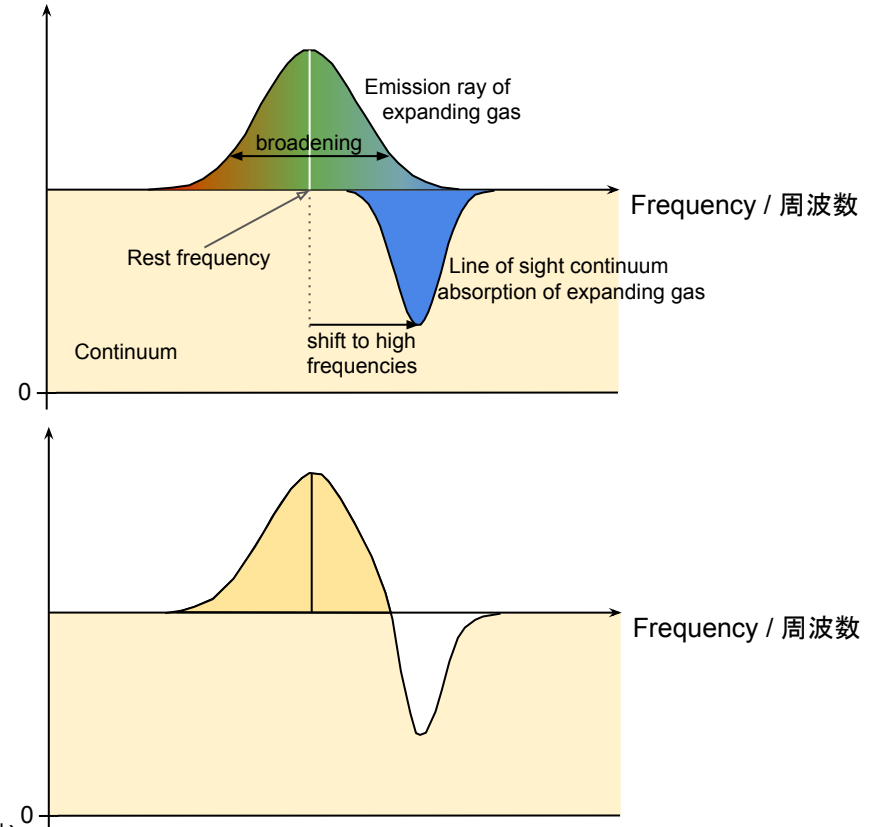
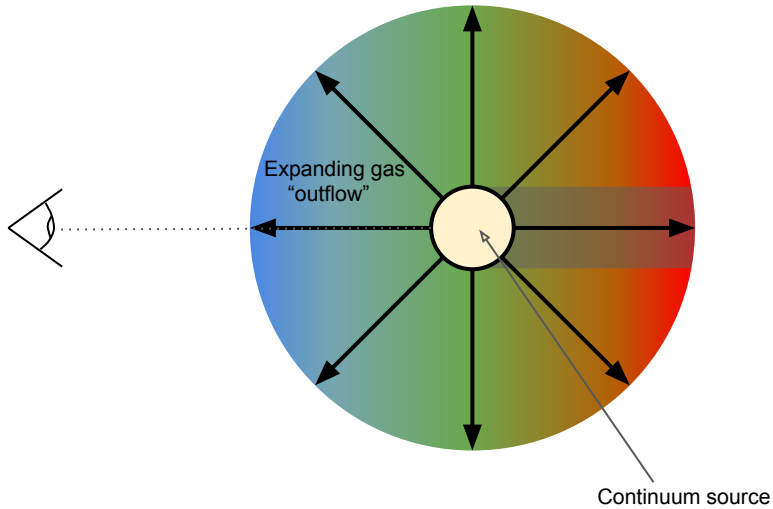
Super Star Clusters (SSC) / 超星団

fast outflows
↓
gas depletion in the cloud
↓
mitigation of star formation



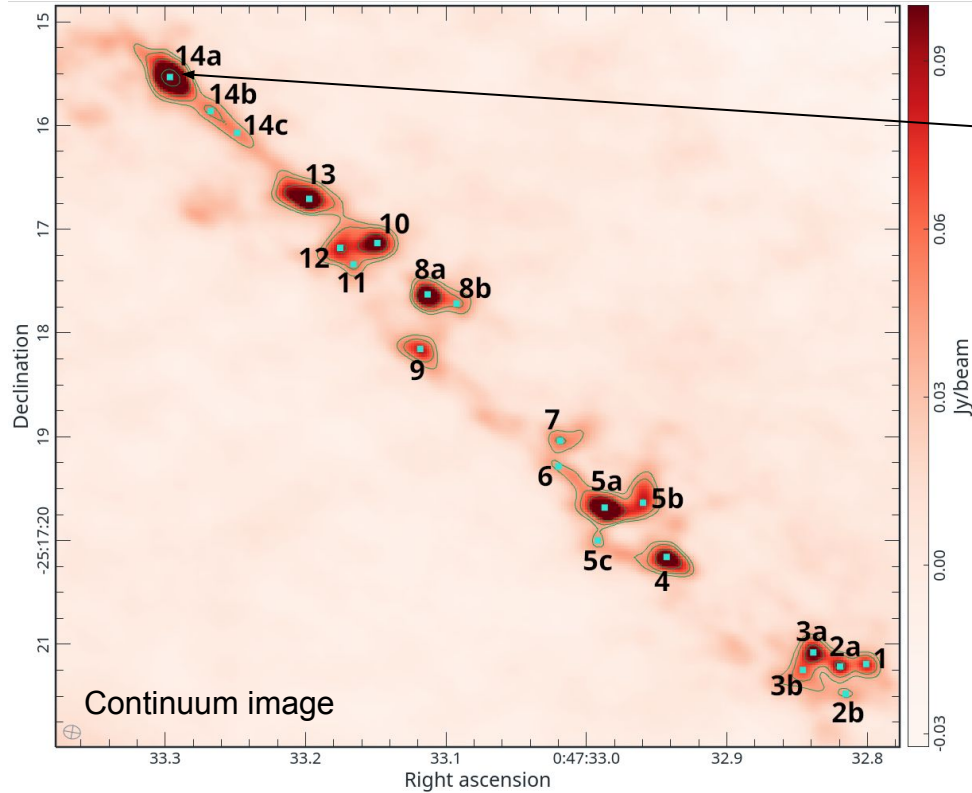
“P-Cygni profile”, signature of outflows

P-Cygni profile, signature of outflows

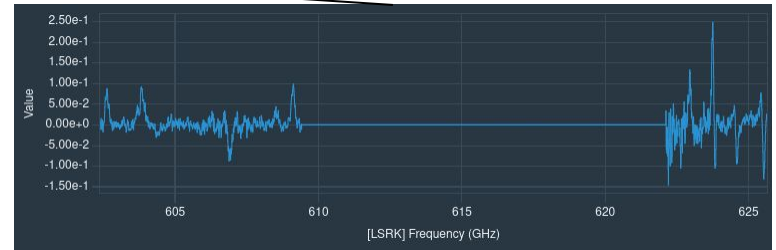


another good sign : presence of species such as ionized water (H_2O^+)

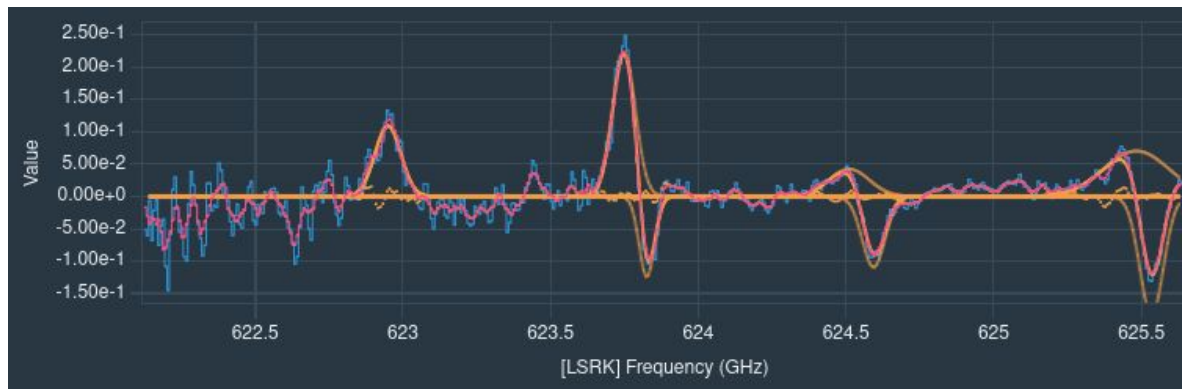
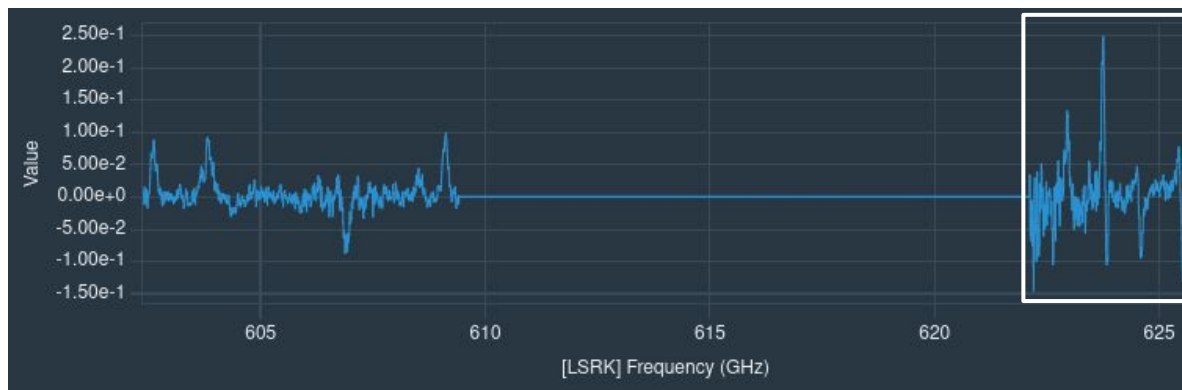
First Results 1/2



ALMA Band 9 (Radio observation)
602.38~609.40 GHz + 622.12~625.64GHz



First Results 2/2

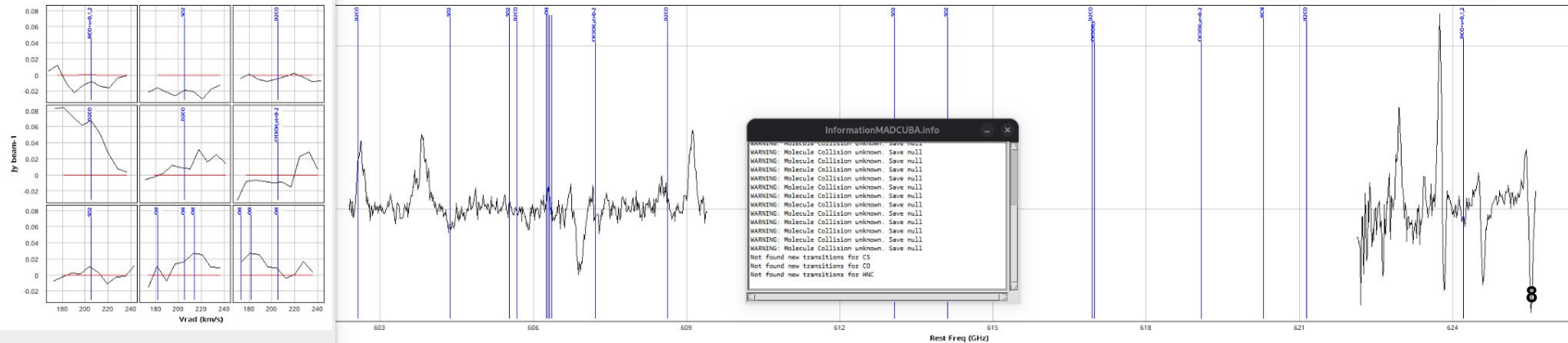


TODO List



- Line identification (currently)
- Find less clean P-Cygni profile if any
- Explain declination offset with Levy et al. 2021
- Better understanding of those outflows
- Write a paper

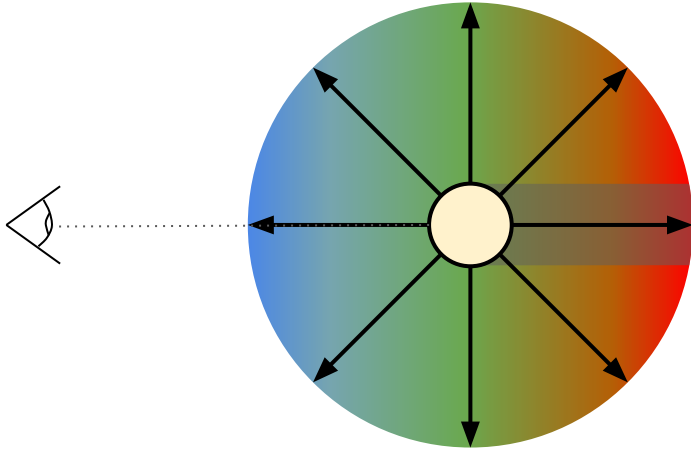
ionized water (H_2O^+)
604.68 GHz
607.23 GHz



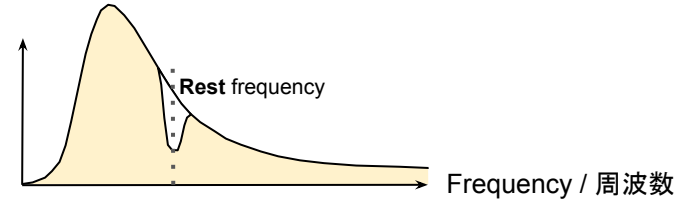


Thank you !

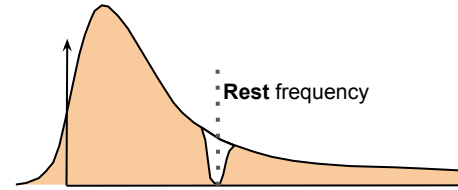
P-Cygni profile, signature of outflows (blueshifted absorption)



case of **motionless** gas

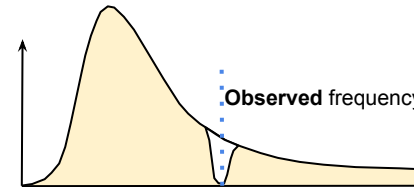


From **gas'** point of view



- Continuum is redshifted as the source is moving away
- Absorption frequency is intrinsic characteristic of the gas

From **observer's** point of view



- Continuum source is not moving
- Absorption seems blueshifted