

Detailed treatment of uncertainties and correlations for jet energy calibration, jet cross section measurements, for the HVP contribution to  $(g-2)_\mu$  and related phenomenological studies

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*HDR Defence*  
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# Areas of interest and topics for today

$(g-2)_\mu$  & Precision QCD



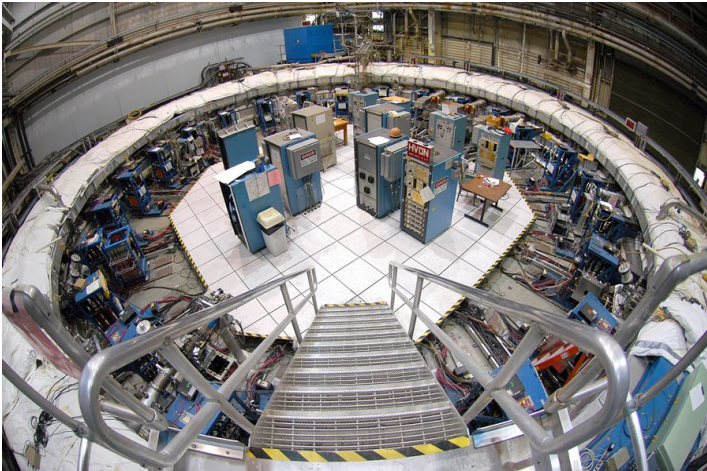
Jet Physics & Performance

Statistics & Machine Learning techniques for:  
Unfolding; Calibration; Anomaly detection

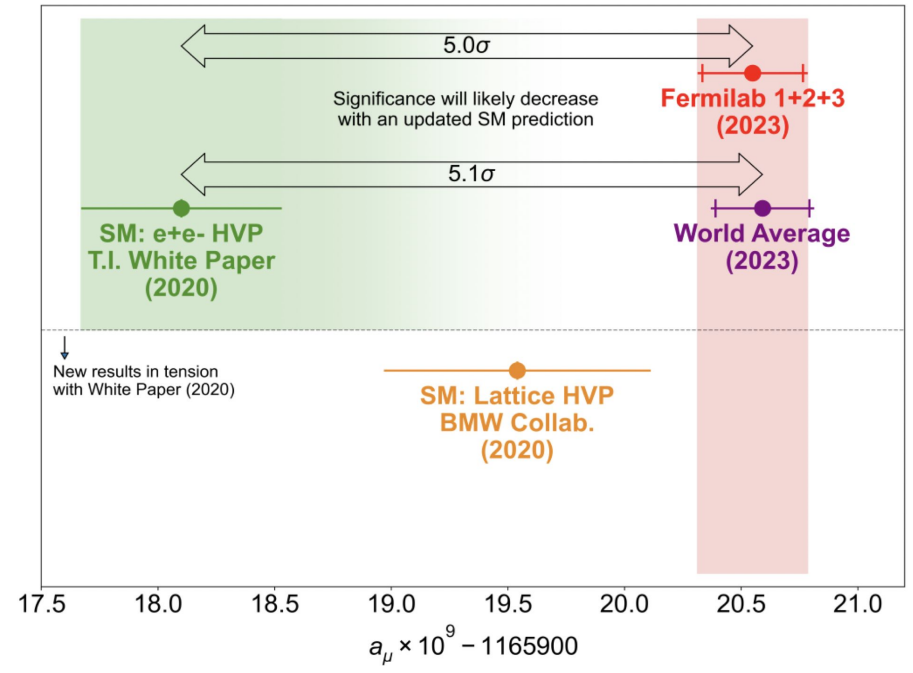
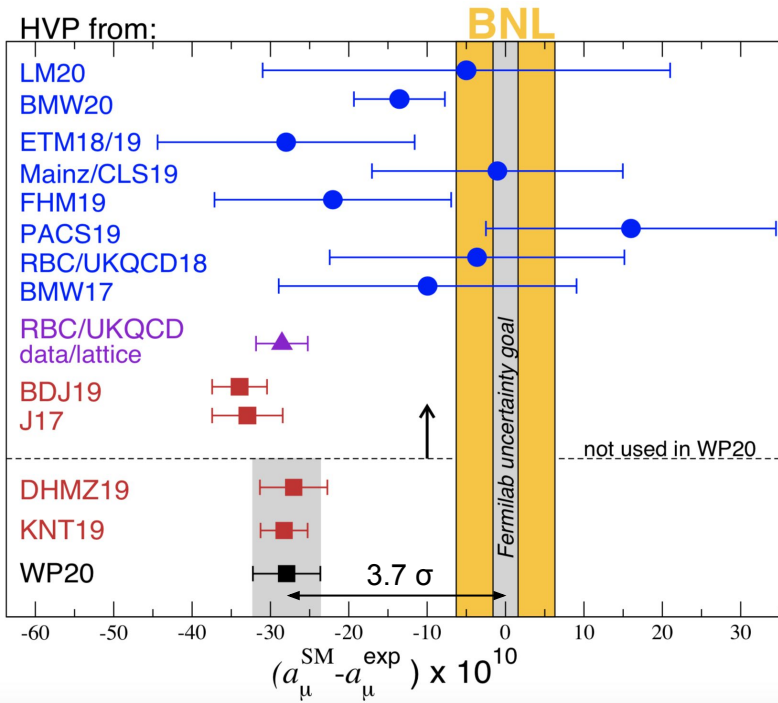
QCD studies @ FCC-ee

$$(g-2)_{\mu}$$

BNL  $\rightarrow$  Fermilab



# Status of $a_\mu$ before 1st / with 2nd Fermilab result

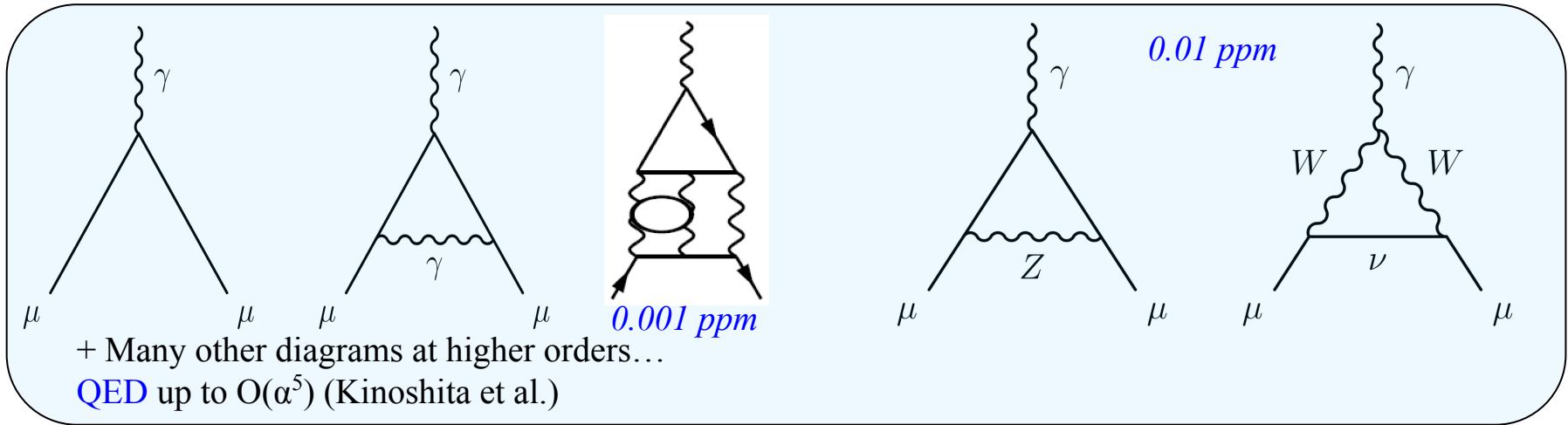


*We have an interesting, long standing, multifaceted problem to solve...*

*Focusing here on the situation for the dispersive approach and the comparison with lattice QCD*

# Theoretical prediction

Why is it (so) complicated to compute one number ? (*very precisely*)

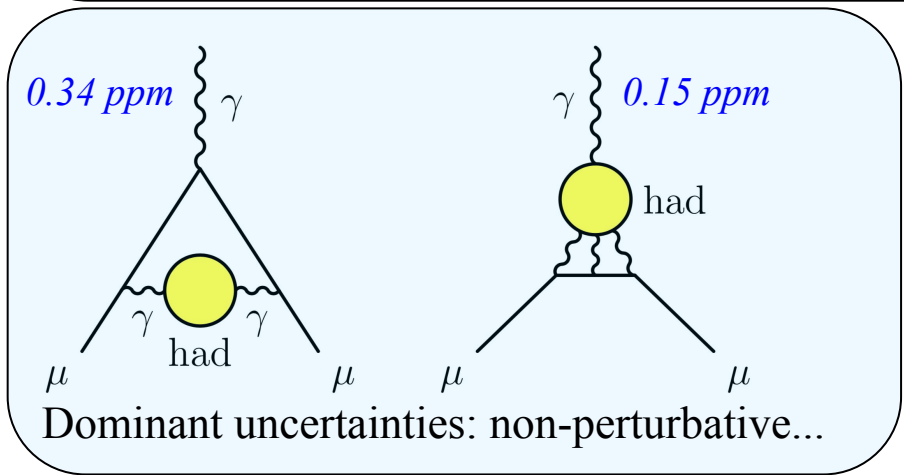


0.001 ppm

0.01 ppm

+ Many other diagrams at higher orders...  
QED up to  $O(\alpha^5)$  (Kinoshita et al.)

Detailed description: This panel shows five Feynman diagrams for the muon g-2 calculation. From left to right: 1) A tree-level diagram with an incoming photon  $\gamma$  and two outgoing muons  $\mu$ . 2) A diagram with a photon  $\gamma$  loop between the muon lines. 3) A diagram with a photon  $\gamma$  loop and a photon  $\gamma$  exchange between the muon lines. 4) A diagram with a Z boson exchange between the muon lines, labeled '0.001 ppm'. 5) A diagram with W boson exchange between the muon lines, labeled '0.01 ppm'. The text below indicates that many other diagrams exist at higher orders and that QED calculations go up to  $O(\alpha^5)$  (Kinoshita et al.).

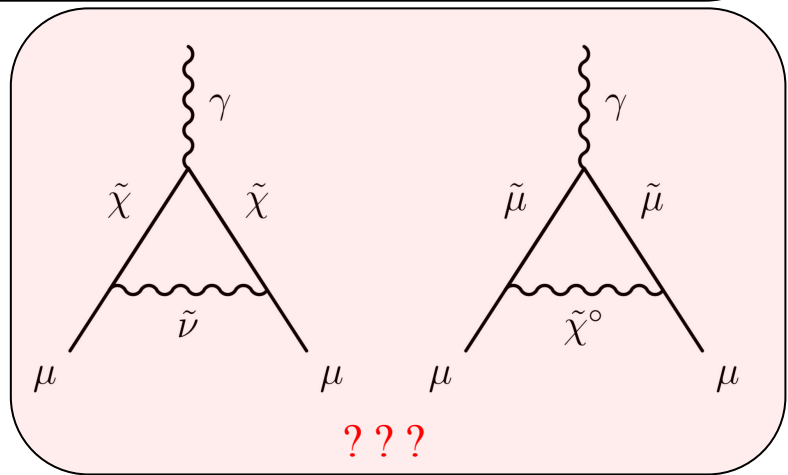


0.34 ppm

0.15 ppm

Dominant uncertainties: non-perturbative...

Detailed description: This panel shows two Feynman diagrams representing hadronic uncertainties. The left diagram shows a photon  $\gamma$  loop with a hadronic vacuum polarization insertion (yellow circle) on the internal photon line, labeled '0.34 ppm'. The right diagram shows a photon  $\gamma$  loop with a hadronic vacuum polarization insertion on the muon line, labeled '0.15 ppm'. The text below states that these are dominant non-perturbative uncertainties.



???

Detailed description: This panel shows two Feynman diagrams representing supersymmetric uncertainties. The left diagram shows a photon  $\gamma$  loop with a stop squark  $\tilde{t}$  loop, labeled '???' in red. The right diagram shows a photon  $\gamma$  loop with a chargino  $\tilde{\chi}^\pm$  loop, also labeled '???' in red.

# Hadronic Vacuum Polarization and Muon $(g-2)_\mu$

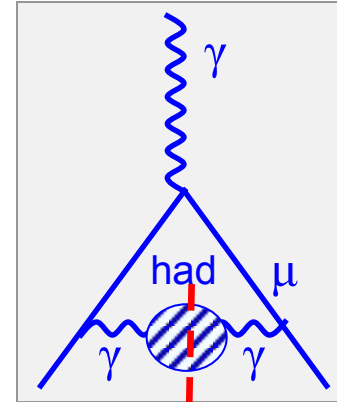
Dominant uncertainty for the theoretical prediction: from lowest-order HVP piece

Cannot be calculated from pQCD (low E-scale), but one can use experimental data on  $e^+e^- \rightarrow$  hadrons cross section

Born:  $\sigma^{(0)}(s) = \sigma(s)(\alpha/\alpha(s))^2$

$$12\pi \operatorname{Im}\Pi_\gamma(s) = \frac{\sigma^0 [e^+e^- \rightarrow \text{hadrons} (\gamma_{FSR})]}{\sigma_{pt}} \equiv R(s)$$

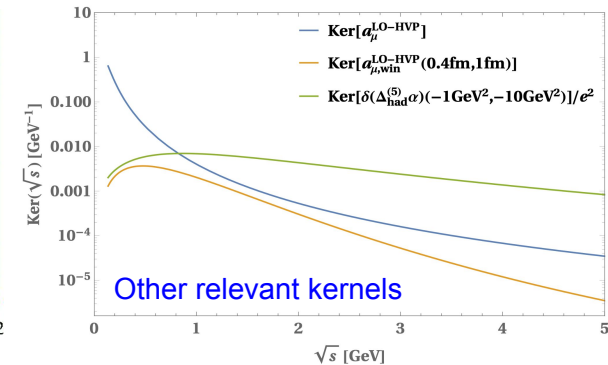
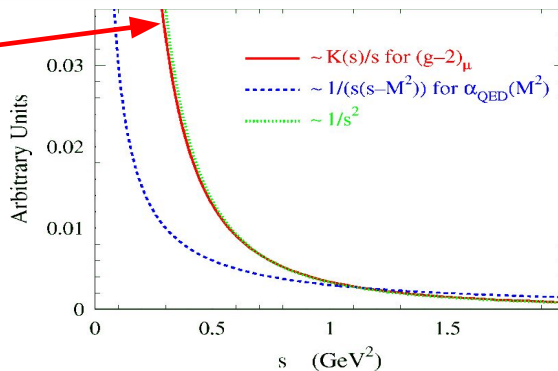
$\operatorname{Im}[ \text{diagram} ] \propto | \text{diagram} \text{ hadrons} |^2$



$$a_\mu^{\text{had}} = \frac{\alpha^2}{3\pi^2} \int_{4m_\pi^2}^{\infty} ds \left( \frac{K(s)}{s} \right) R(s)$$

Dispersion relation

Bouchiat and Michel, 1961

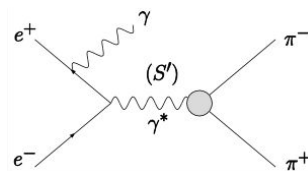
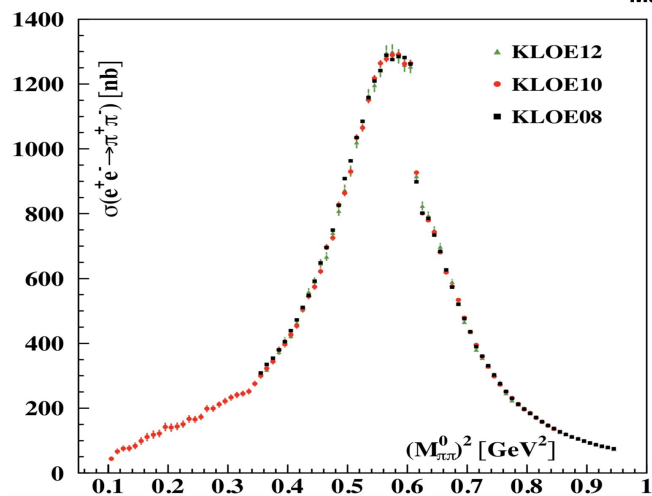
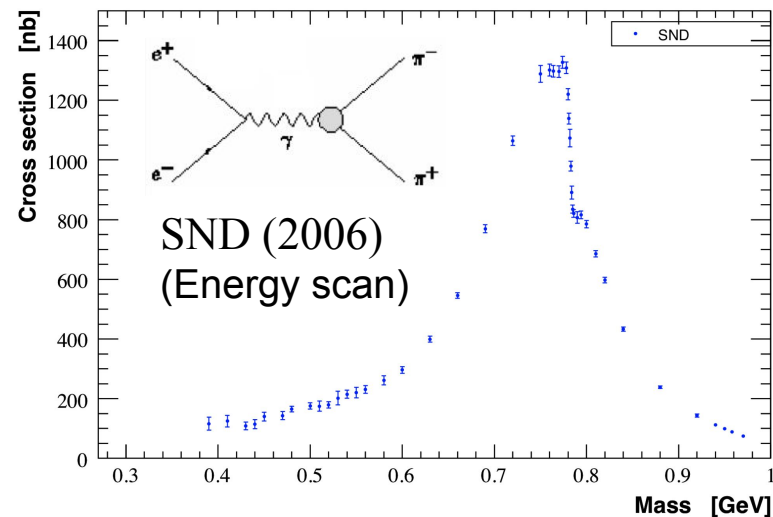
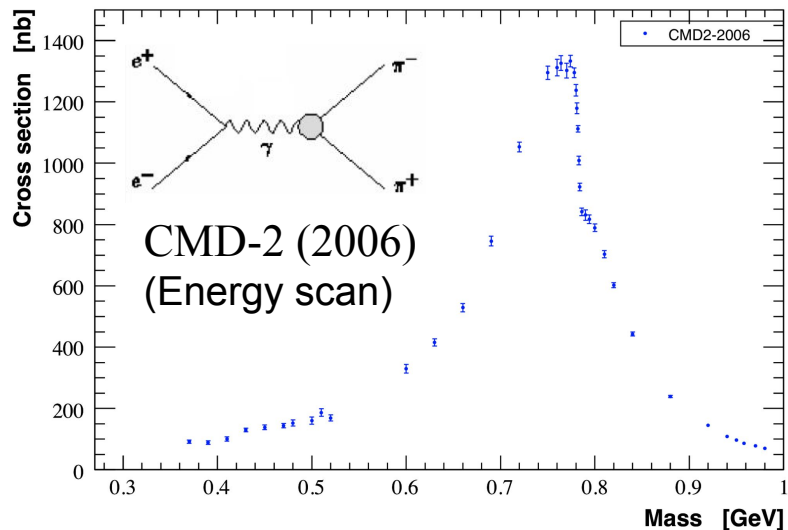


→ Precise  $\sigma(e^+e^- \rightarrow \text{hadrons})$  measurements at low energy are very important

$\pi\pi$  channel: 73% (70%) of HVP contribution to  $a_\mu$  (uncertainty  $^2$ )

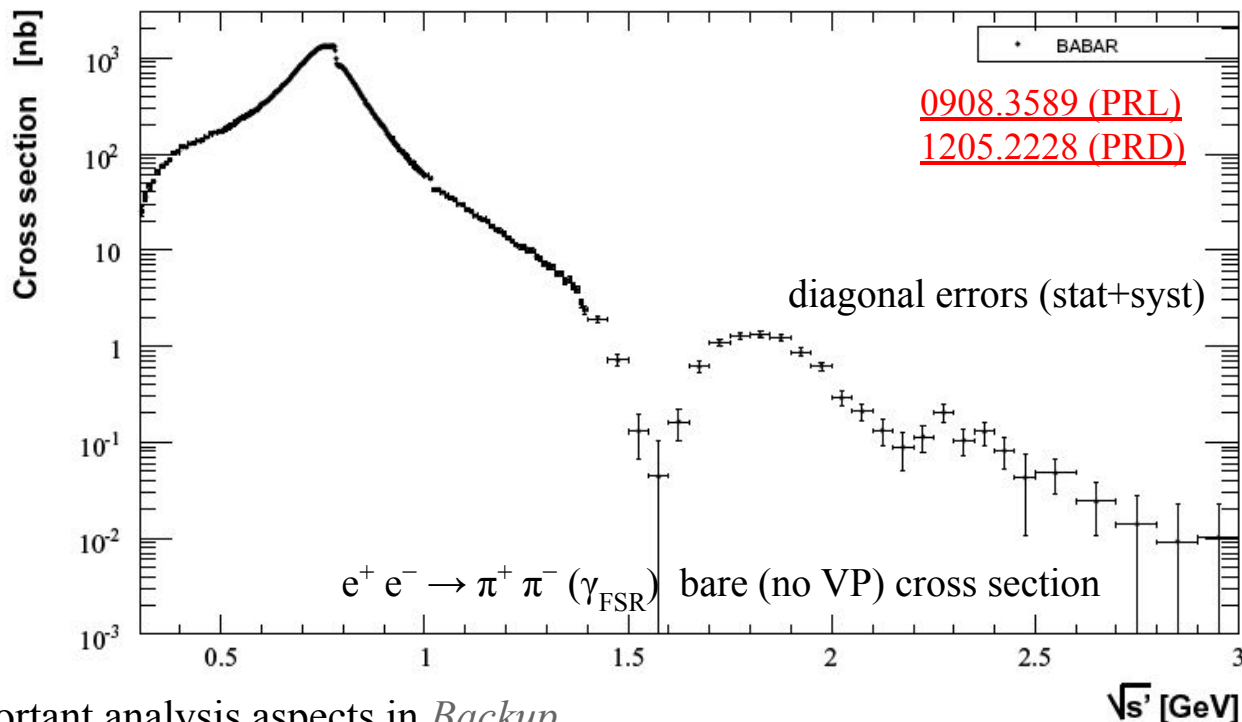
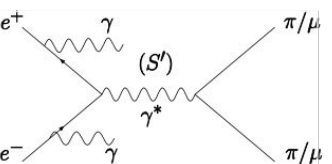
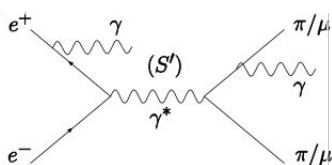
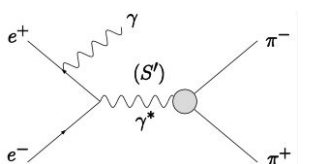
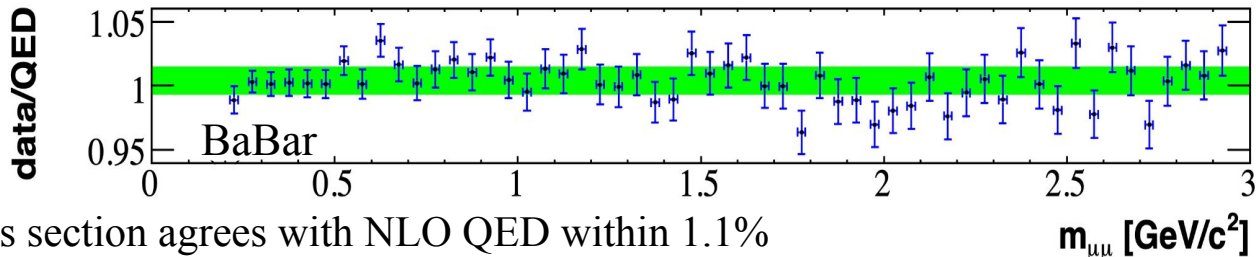
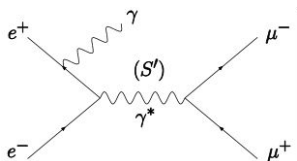
→ Alternatively, one can use hadronic  $\tau$  decays data + Isospin Breaking corrections

# HVP: Data on $e^+e^- \rightarrow \text{hadrons}$



KLOE (08&10) +  $\mu\mu$  (12) (ISR)

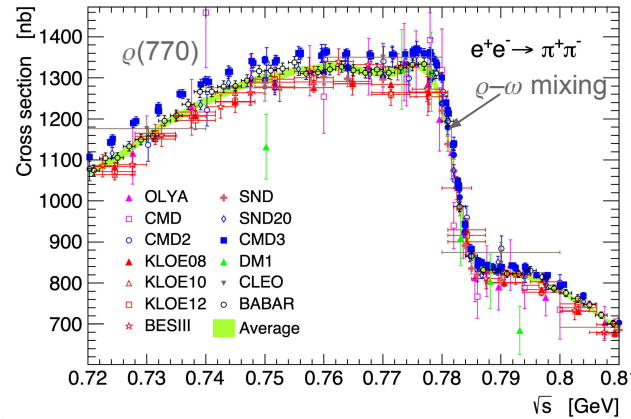
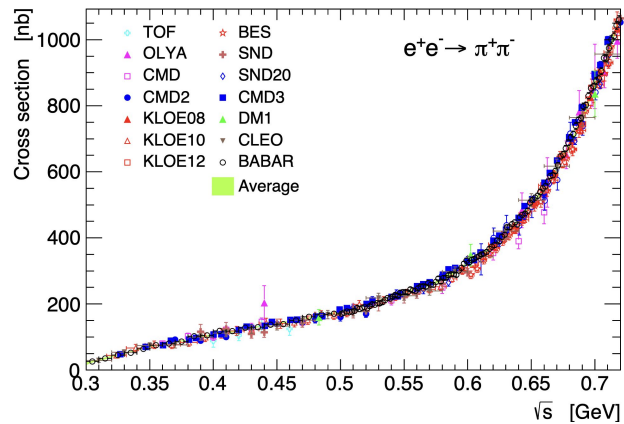
# BaBar results: “NLO” cross-section



→ More details on important analysis aspects in [Backup](#)

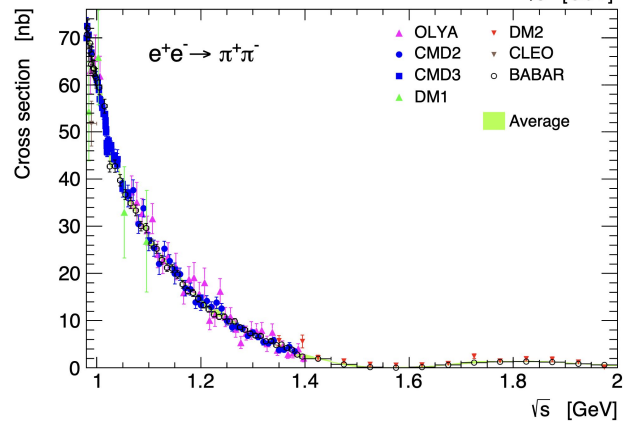
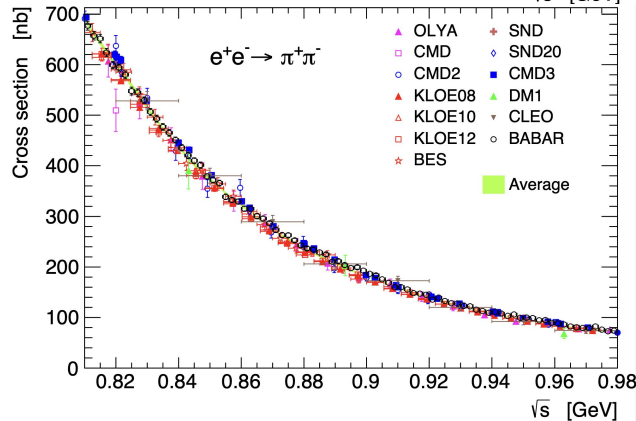


# Experimental data combination (Example: $e^+e^- \rightarrow \pi^+\pi^-$ channel)



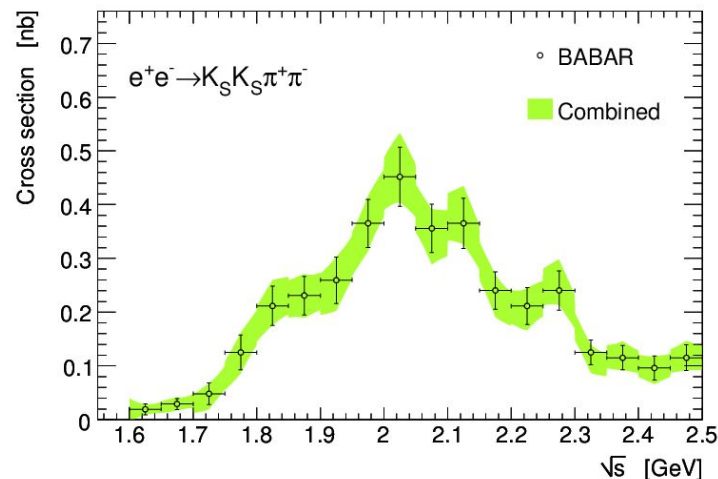
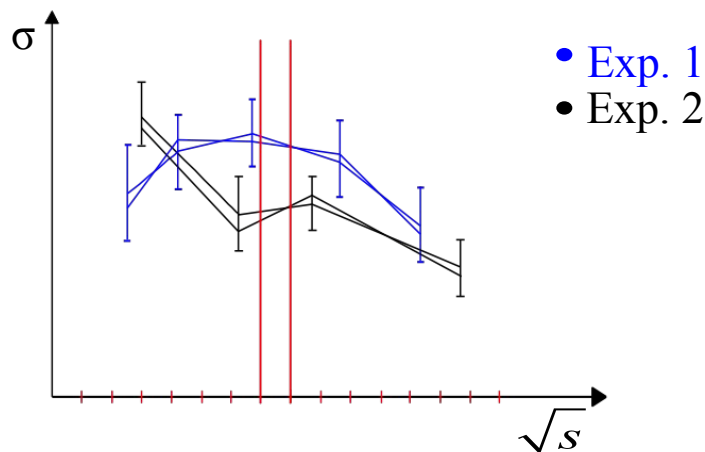
[2312.02053](#)

→ *New since g-2 Theory Initiative White Paper: Large tensions, especially between KLOE & CMD3, which provide the smallest / largest cross-sections in the  $\rho$  region*



Procedure and software (*HVPTools* - Since 2009) for combining cross section data with arbitrary point spacing/binning → Validated through closure test. Featuring full & realistic (i.e. not too optimistic) treatment of uncertainties and correlations (between measurements (data points/bins) of a given experiment, b. experiments, b. different channels), fully accounting for systematic tensions between experiments. (*1<sup>st</sup> motivation for DHMZ uncertainties = “baseline” in g-2 TI WP*)

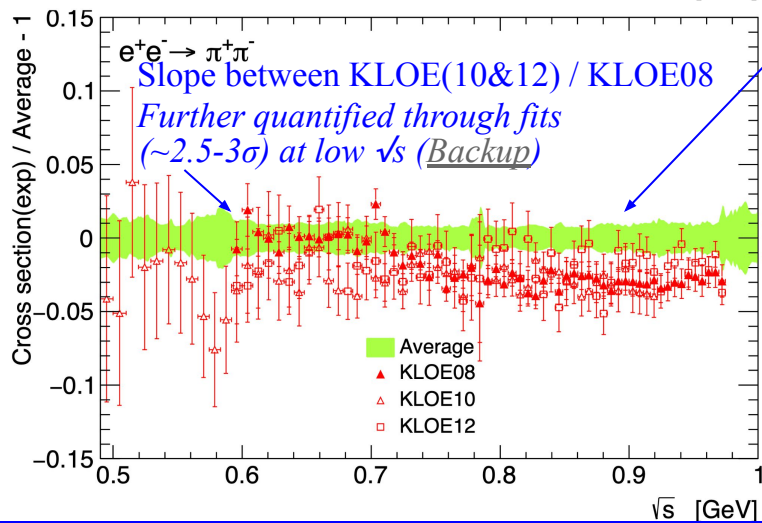
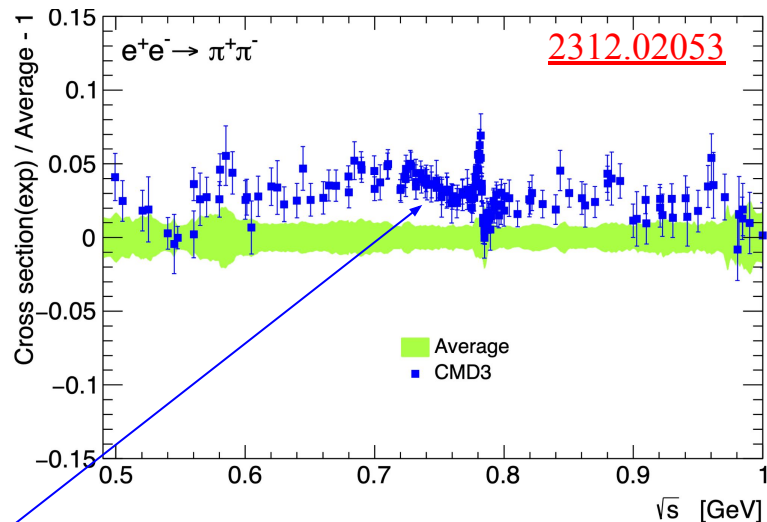
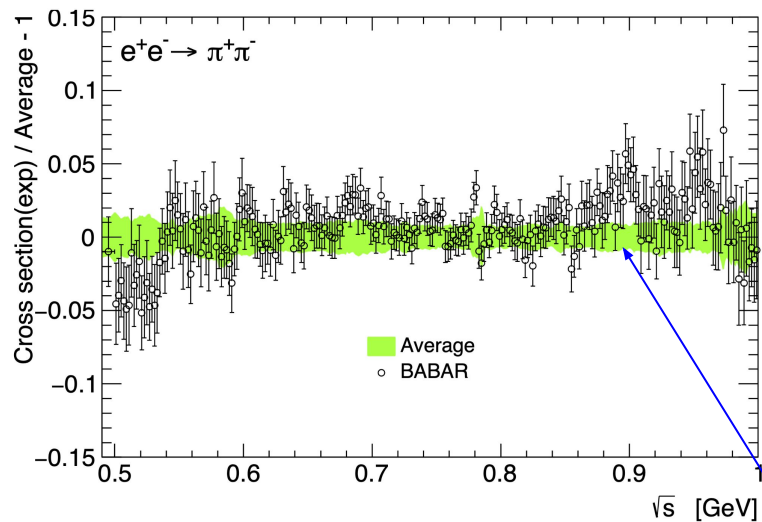
# Combination procedure implemented in HVPTools software



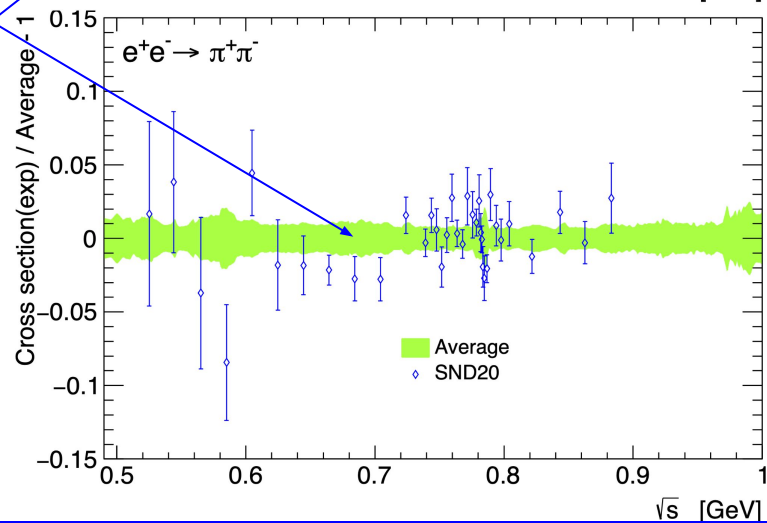
- Define a (fine) final binning (to be filled and used for integrals etc.)
- Linear/quadratic splines to interpolate between the points/bins of each experiment
  - for binned measurements: preserve integral inside each bin
  - closure test: replace nominal values of data points by Gounaris-Sakurai model and re-do the combination
    - (non-)negligible bias for (linear)quadratic interpolation
- Fluctuate data points taking into account correlations & re-do the splines for each (pseudo-)experiment
  - each uncertainty fluctuated coherently for all the points/bins that it impacts
  - eigenvector decomposition for (statistical) covariance matrices
- In each final bin, compute: average value for each measurement & its uncertainty; correlation matrix between experiments

Backup

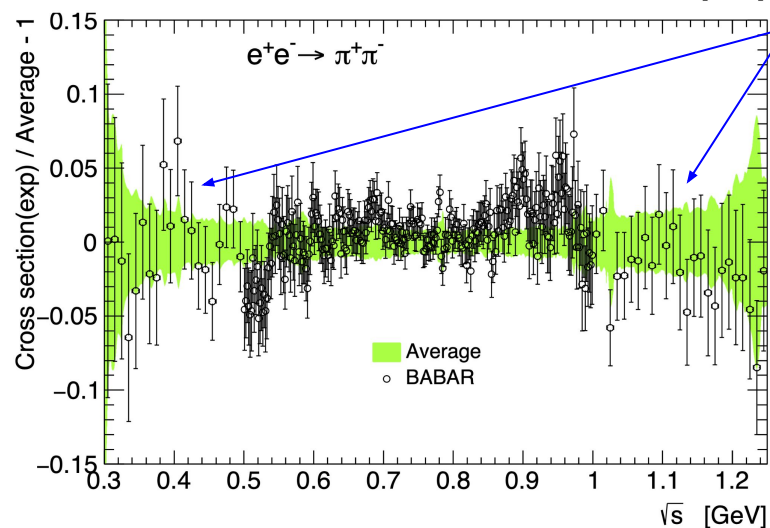
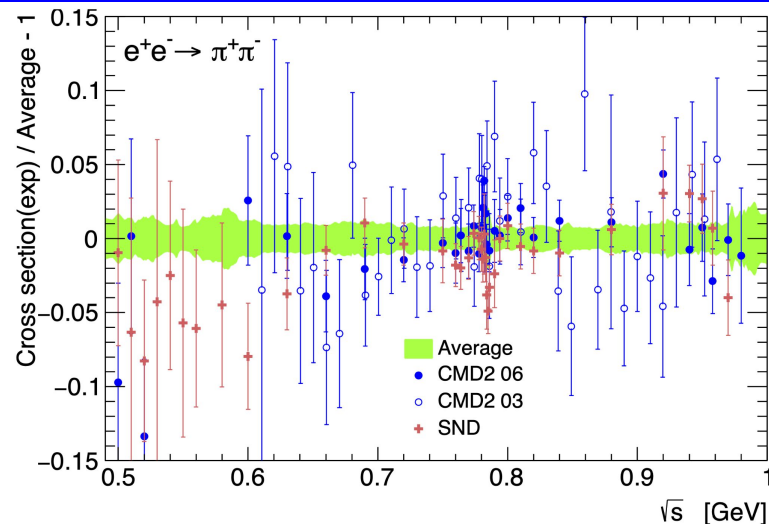
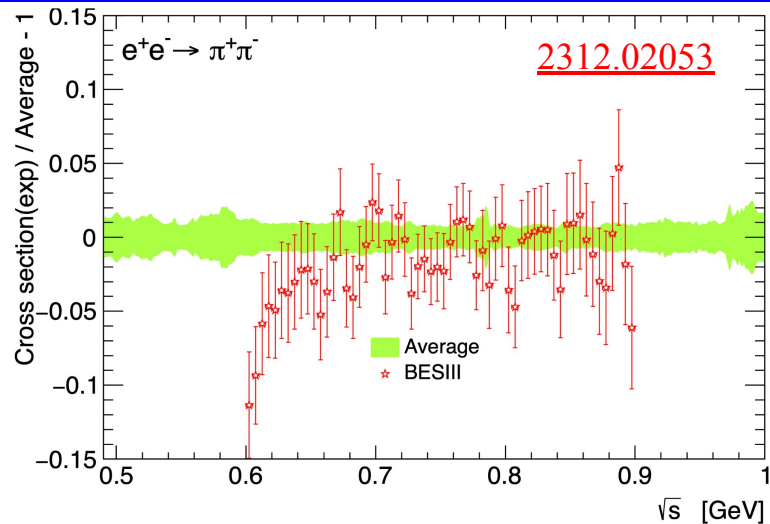
# Combining the $e^+e^- \rightarrow \pi^+\pi^-$ data: relative differences



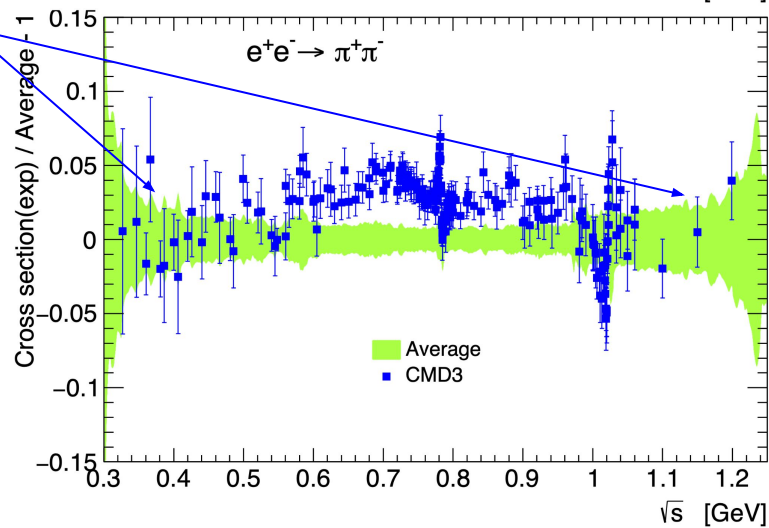
Systematic tensions



# Combining the $e^+e^- \rightarrow \pi^+\pi^-$ data: relative differences

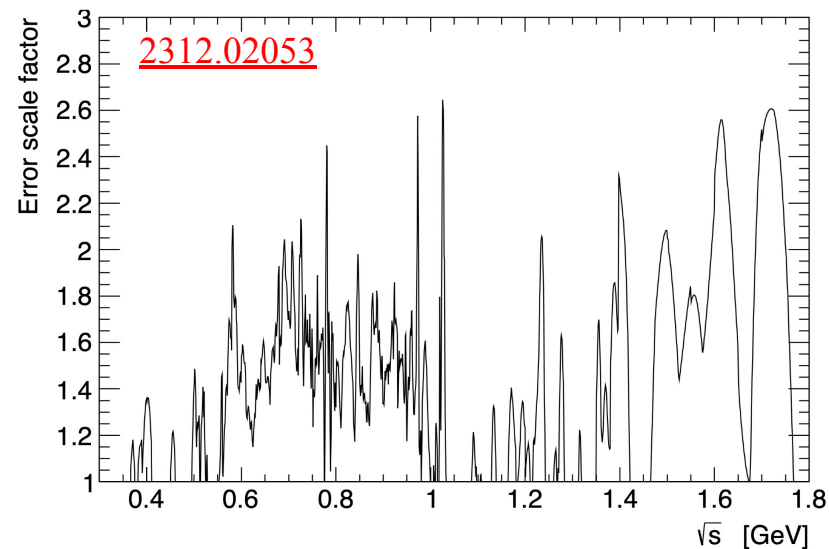
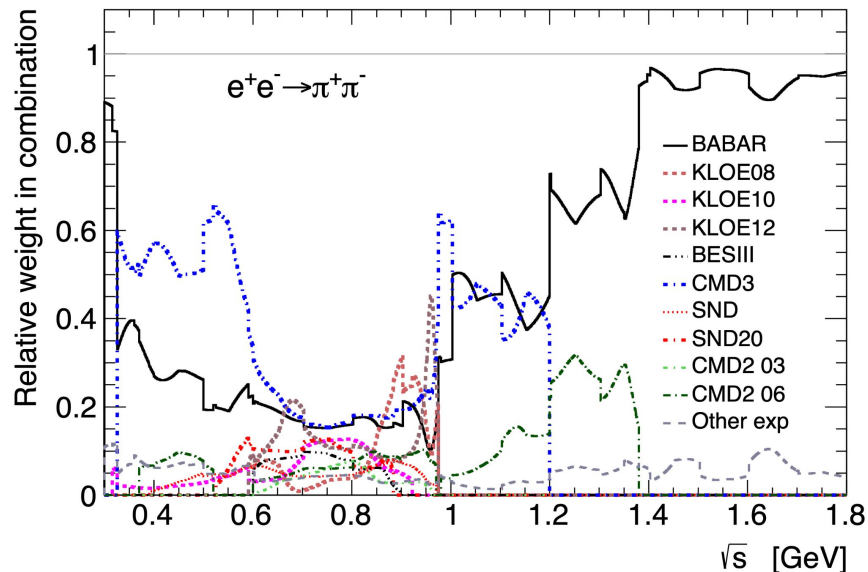


Reasonable  
BABAR/CMD3  
agreement at  
low & high E



# Combination procedure: weights and tension

- For each narrow final bin minimize  $\chi^2$  to get average coefficients test locally the level of agreement
- Average weights account for bin sizes/point-spacing of measurements: compare precisions on same footing

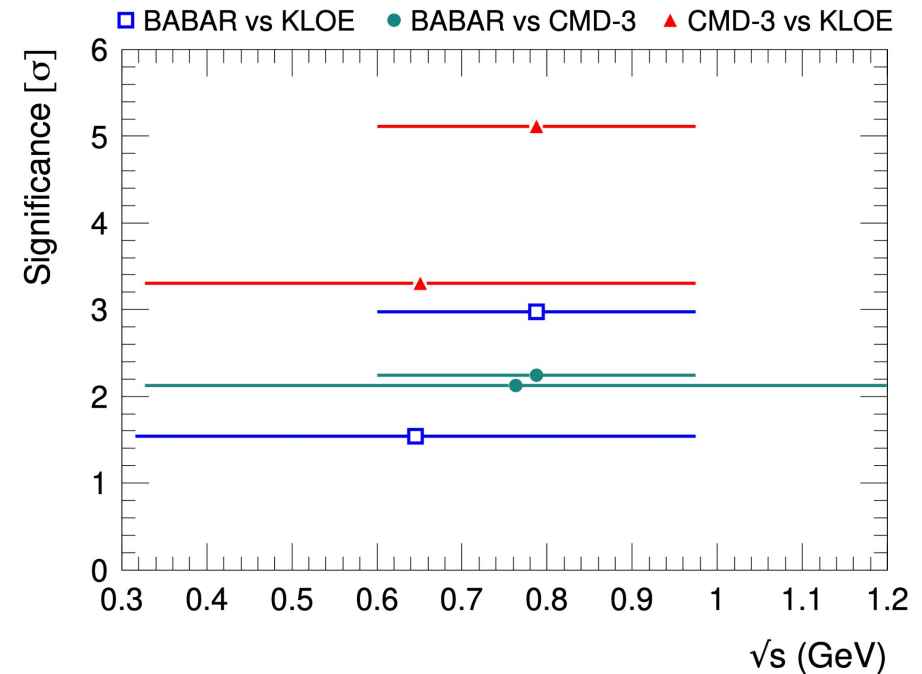
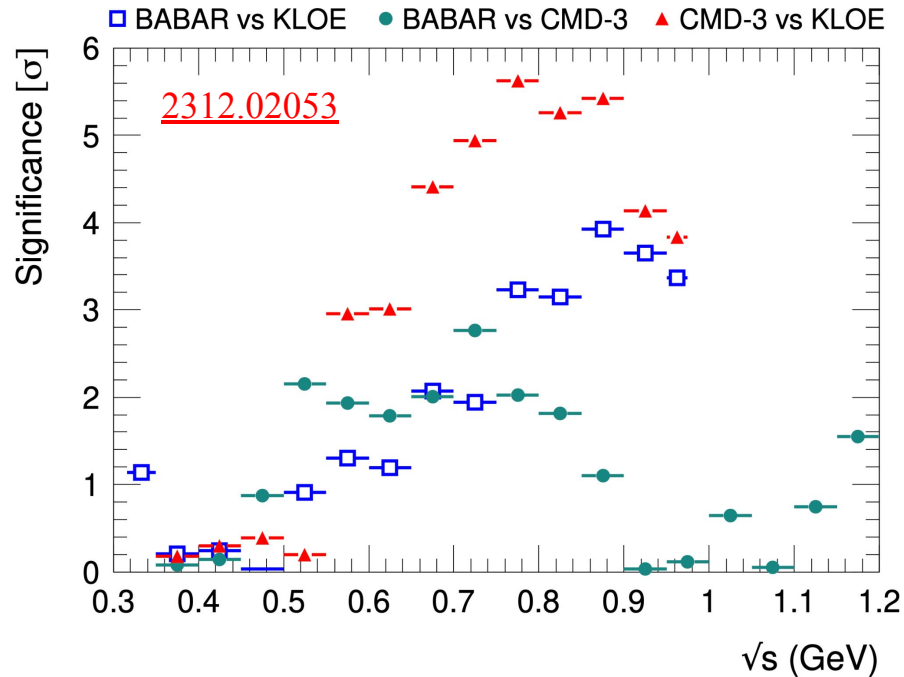


- Average dominated by BaBar, CMD3, KLOE, SND20; BaBar covers full energy range
- Enhanced tensions, especially between KLOE & CMD3, which provide the smallest / largest cross-sections in the  $\rho$  region: *clear indication of underestimated uncertainties*
- *Calls for conservative uncertainty treatment in combination fit* (fits / evaluation of weights)
- *Systematic effects beyond the local  $\chi^2$  /ndof rescaling*: had already motivated the inclusion of the dominant BABAR-KLOE systematic by DHMZ since 2019 (*2<sup>nd</sup> motivation for DHMZ uncertainties = “baseline” in g-2 TI White Paper*), but *tensions are larger now*

Backup

# Quantitative comparisons for $a_{\mu}^{\text{HVP}}$

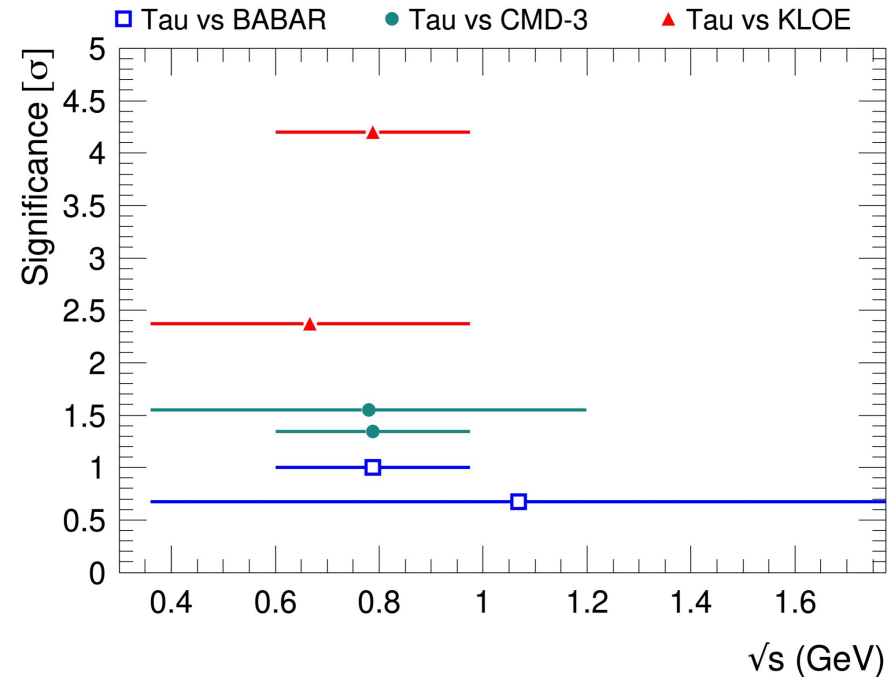
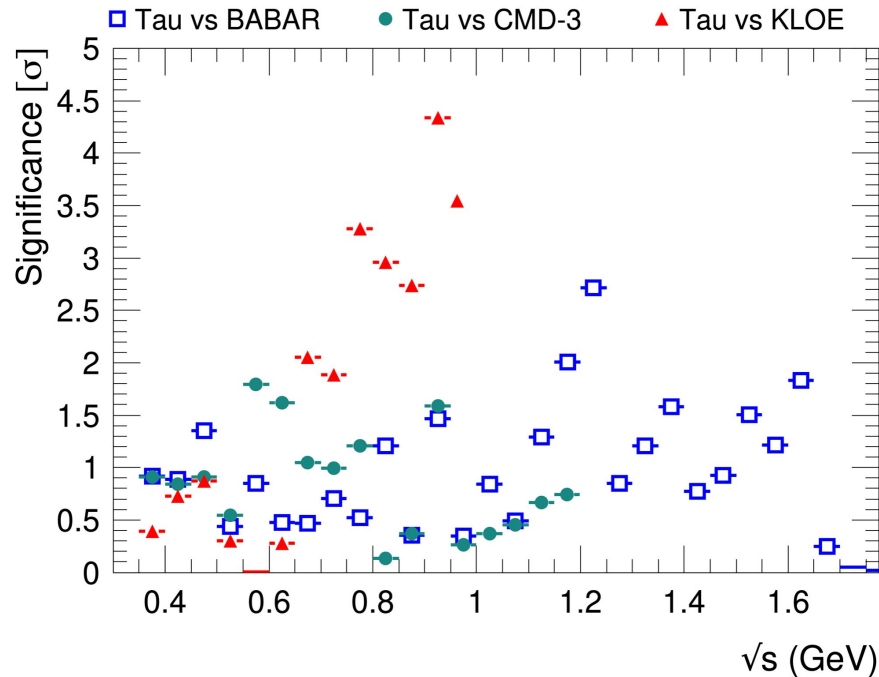
→ Comparison of integrals computed in various restricted energy ranges, for individual  $e^+e^-$  experiments: significance of the difference between different experiments, taking into account correlations



→ Largest tensions between CMD3 and KLOE

# Quantitative comparisons for $a_{\mu}^{\text{HVP}}$

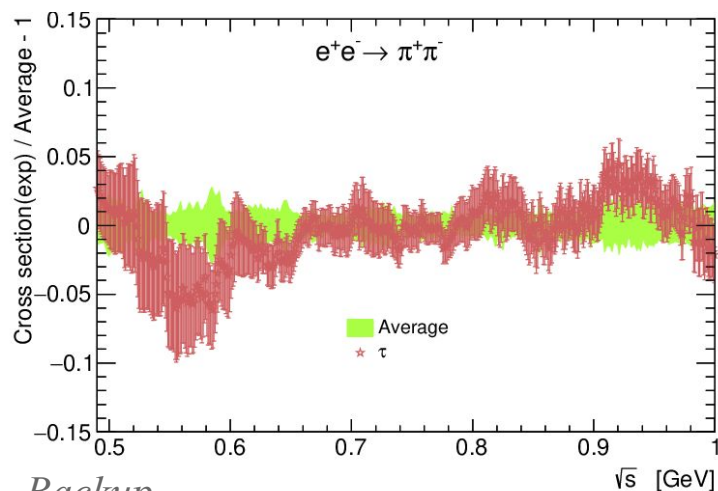
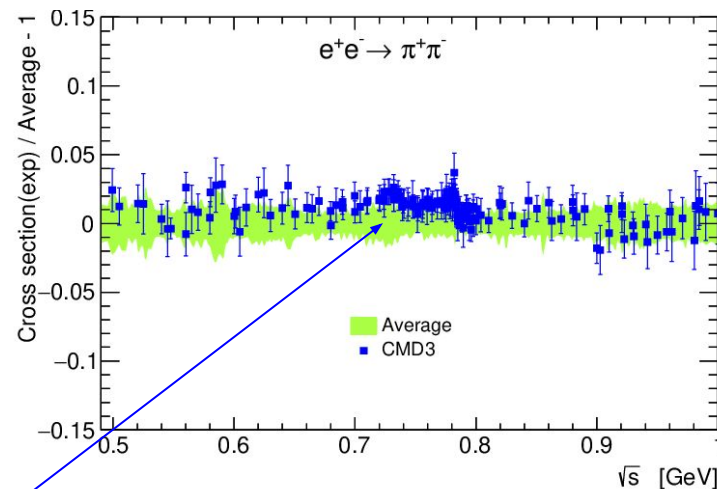
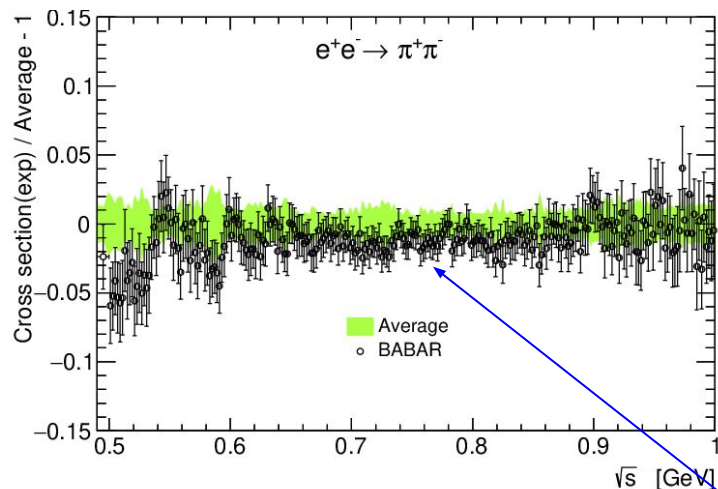
→ Comparison of integrals computed in various restricted energy ranges, for  $\tau$  / individual  $e^+e^-$  experiments: significance of the difference between different experiments, taking into account correlations



→ Largest tensions between Tau and KLOE

→ Good agreement among the Tau measurements ([Backup](#))

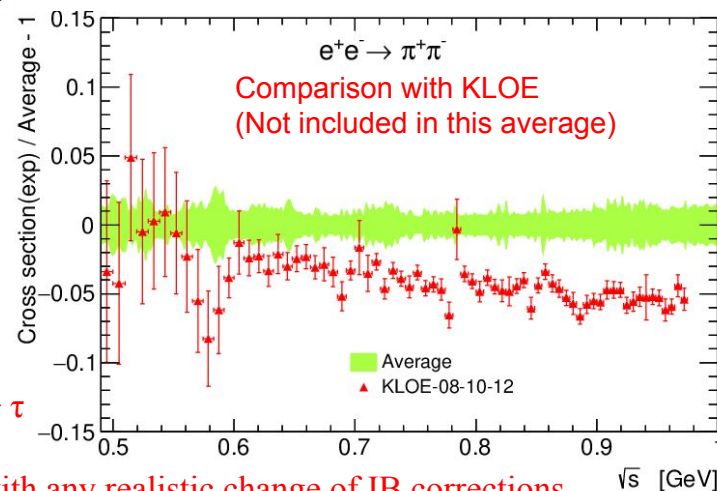
# Combining the $e^+e^- \rightarrow \pi^+\pi^-$ data, BaBar & CMD3 & Tau(+IB)



Some (reduced)  
systematic  
tensions

Much larger tension  
(slope and shift) for  
KLOE vs.  
BABAR + CMD-3 +  $\tau$   
combination

→ Not compatible with any realistic change of IB corrections



Backup

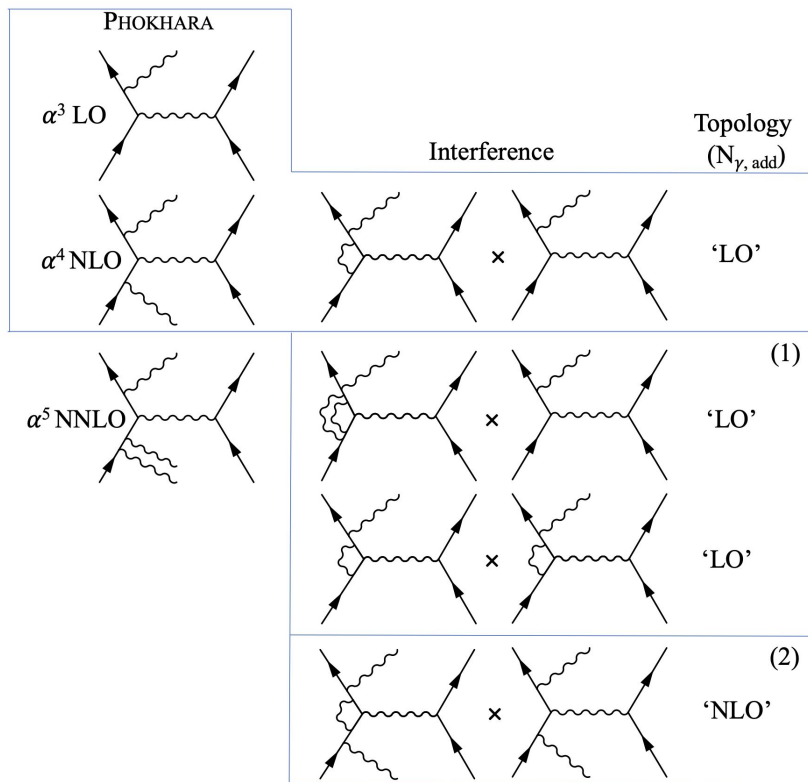


# Impact of higher order photon emissions: *Unique '(N)NLO' BaBar study*

→ Studied in-situ in BaBar data, using kinematic fits: test the most frequently used Monte Carlo generators

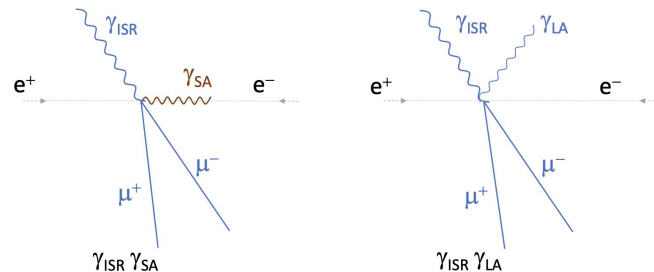
- PHOKHARA: full NLO matrix element for ISR and FSR
- AFKQED: NLO and NNLO, with collinear approximation for additional ISR

Generic ISR diagrams

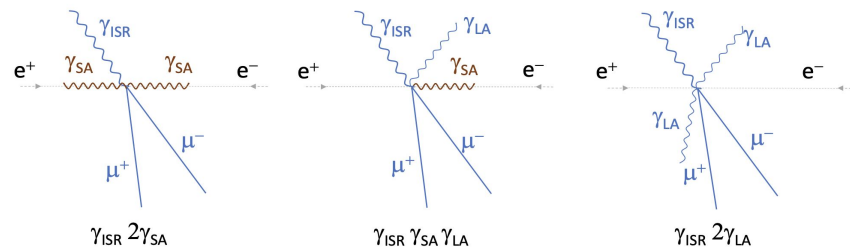


2308.05233

'NLO'

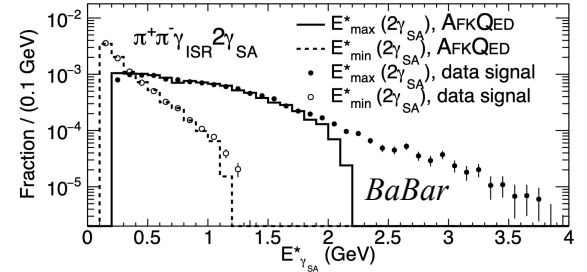
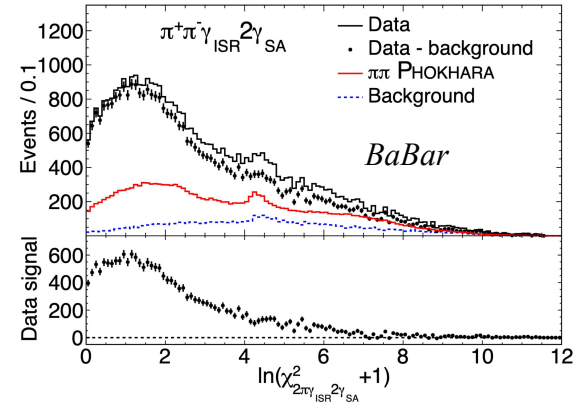
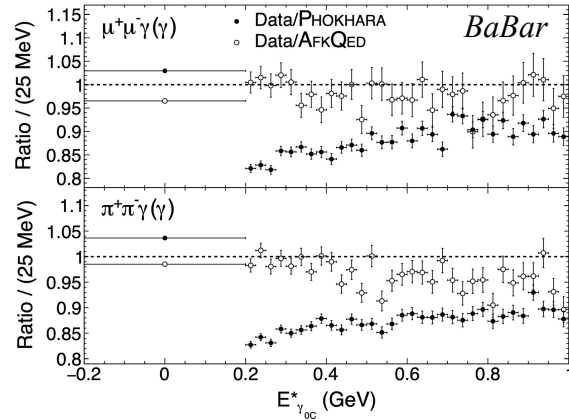
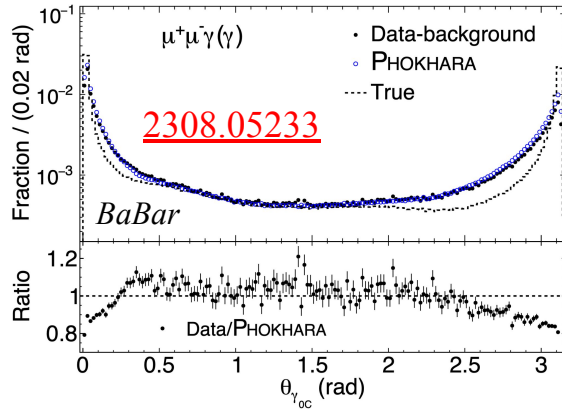


'NNLO'



'(N)NLO':  $\gamma_{\text{add}}$  counting ( $E_{\gamma_{\text{add}}} > E_{\text{min}}$ )

# BaBar results on higher order photon emissions



→ NLO small-angle ISR in PHOKHARA higher than in data; large-angle ratios consistent with unity

→ Independent Belle II confirmation of PHOKHARA problem

→ AFKQED: reasonable description of rate and energy distributions for ‘(N)NLO’ data

→ NNLO contributions clearly observed in data

→ BaBar measurements: loose selection incorporates NLO and HO radiation, minimising MC-dependence

→ Other ISR measurements select ‘LO’ topology and rely on PHOKHARA for hard NLO (but with no NNLO)

→ Aspects further studied with fast simulation, questioning the KLOE systematic uncertainties ([2312.02053](#))

Backup

# Sum of hadronic contributions

Channel	$a_{\mu}^{\text{had,LO}} [10^{-10}]$	$\Delta\alpha_{\text{had}}(m_Z^2) [10^{-4}]$
$\pi^0\gamma$	$4.41 \pm 0.06 \pm 0.04 \pm 0.07$	$0.35 \pm 0.00 \pm 0.00 \pm 0.01$
$\eta\gamma$	$0.65 \pm 0.02 \pm 0.01 \pm 0.01$	$0.08 \pm 0.00 \pm 0.00 \pm 0.00$
$\pi^+\pi^-$	$507.85 \pm 0.83 \pm 3.23 \pm 0.55$	$34.50 \pm 0.06 \pm 0.20 \pm 0.04$
$\pi^+\pi^-\pi^0$	$46.21 \pm 0.40 \pm 1.10 \pm 0.86$	$4.60 \pm 0.04 \pm 0.11 \pm 0.08$
$2\pi^+2\pi^-$	$13.68 \pm 0.03 \pm 0.27 \pm 0.14$	$3.58 \pm 0.01 \pm 0.07 \pm 0.03$
$\pi^+\pi^-2\pi^0$	$18.03 \pm 0.06 \pm 0.48 \pm 0.26$	$4.45 \pm 0.02 \pm 0.12 \pm 0.07$
$2\pi^+2\pi^-\pi^0$ ( $\eta$ excl.)	$0.69 \pm 0.04 \pm 0.06 \pm 0.03$	$0.21 \pm 0.01 \pm 0.02 \pm 0.01$
$\pi^+\pi^-3\pi^0$ ( $\eta$ excl.)	$0.49 \pm 0.03 \pm 0.09 \pm 0.00$	$0.15 \pm 0.01 \pm 0.03 \pm 0.00$
$3\pi^+3\pi^-$	$0.11 \pm 0.00 \pm 0.01 \pm 0.00$	$0.04 \pm 0.00 \pm 0.00 \pm 0.00$
$2\pi^+2\pi^-2\pi^0$ ( $\eta$ excl.)	$0.71 \pm 0.06 \pm 0.07 \pm 0.14$	$0.25 \pm 0.02 \pm 0.02 \pm 0.05$
$\pi^+\pi^-4\pi^0$ ( $\eta$ excl., isospin)	$0.08 \pm 0.01 \pm 0.08 \pm 0.00$	$0.03 \pm 0.00 \pm 0.03 \pm 0.00$
$\eta\pi^+\pi^-$	$1.19 \pm 0.02 \pm 0.04 \pm 0.02$	$0.35 \pm 0.01 \pm 0.01 \pm 0.01$
$\eta\omega$	$0.35 \pm 0.01 \pm 0.02 \pm 0.01$	$0.11 \pm 0.00 \pm 0.01 \pm 0.00$
$\eta\pi^+\pi^-\pi^0$ (non- $\omega$ , $\phi$ )	$0.34 \pm 0.03 \pm 0.03 \pm 0.04$	$0.12 \pm 0.01 \pm 0.01 \pm 0.01$
$\eta 2\pi^+2\pi^-$	$0.02 \pm 0.01 \pm 0.00 \pm 0.00$	$0.01 \pm 0.00 \pm 0.00 \pm 0.00$
$\omega\eta\pi^0$	$0.06 \pm 0.01 \pm 0.01 \pm 0.00$	$0.02 \pm 0.00 \pm 0.00 \pm 0.00$
$\omega\pi^0$ ( $\omega \rightarrow \pi^0\gamma$ )	$0.94 \pm 0.01 \pm 0.03 \pm 0.00$	$0.20 \pm 0.00 \pm 0.01 \pm 0.00$
$\omega 2\pi$ ( $\omega \rightarrow \pi^0\gamma$ )	$0.07 \pm 0.00 \pm 0.00 \pm 0.00$	$0.02 \pm 0.00 \pm 0.00 \pm 0.00$
$\omega$ (non- $3\pi$ , $\pi\gamma$ , $\eta\gamma$ )	$0.04 \pm 0.00 \pm 0.00 \pm 0.00$	$0.00 \pm 0.00 \pm 0.00 \pm 0.00$
$K^+K^-$	$23.08 \pm 0.20 \pm 0.33 \pm 0.21$	$3.35 \pm 0.03 \pm 0.05 \pm 0.03$
$K_S K_L$	$12.82 \pm 0.06 \pm 0.18 \pm 0.15$	$1.74 \pm 0.01 \pm 0.03 \pm 0.02$
$\phi$ (non- $K\bar{K}$ , $3\pi$ , $\pi\gamma$ , $\eta\gamma$ )	$0.05 \pm 0.00 \pm 0.00 \pm 0.00$	$0.01 \pm 0.00 \pm 0.00 \pm 0.00$
$K\bar{K}\pi$	$2.45 \pm 0.05 \pm 0.10 \pm 0.06$	$0.78 \pm 0.02 \pm 0.03 \pm 0.02$
$K\bar{K}2\pi$	$0.85 \pm 0.02 \pm 0.05 \pm 0.01$	$0.30 \pm 0.01 \pm 0.02 \pm 0.00$
$K\bar{K}\omega$	$0.00 \pm 0.00 \pm 0.00 \pm 0.00$	$0.00 \pm 0.00 \pm 0.00 \pm 0.00$
$\eta\phi$	$0.33 \pm 0.01 \pm 0.01 \pm 0.00$	$0.11 \pm 0.00 \pm 0.00 \pm 0.00$
$\eta K\bar{K}$ (non- $\phi$ )	$0.01 \pm 0.01 \pm 0.01 \pm 0.00$	$0.00 \pm 0.00 \pm 0.01 \pm 0.00$
$\omega 3\pi$ ( $\omega \rightarrow \pi^0\gamma$ )	$0.06 \pm 0.01 \pm 0.01 \pm 0.01$	$0.02 \pm 0.00 \pm 0.00 \pm 0.00$
$7\pi$ ( $3\pi^+3\pi^-\pi^0$ + estimate)	$0.02 \pm 0.00 \pm 0.01 \pm 0.00$	$0.01 \pm 0.00 \pm 0.00 \pm 0.00$
$J/\psi$ (BW integral)	$6.20 \pm 0.11$	$7.00 \pm 0.13$
$\psi(2S)$ (BW integral)	$1.56 \pm 0.05$	$2.48 \pm 0.08$
$R$ data [3.7 – 5.0] GeV	$7.29 \pm 0.05 \pm 0.30 \pm 0.00$	$15.79 \pm 0.12 \pm 0.66 \pm 0.00$
$R_{\text{QCD}} [1.8 – 3.7 \text{ GeV}]_{uds}$	$33.45 \pm 0.28 \pm 0.65_{\text{dual}}$	$24.27 \pm 0.18 \pm 0.28_{\text{dual}}$
$R_{\text{QCD}} [5.0 – 9.3 \text{ GeV}]_{udsc}$	$6.86 \pm 0.04$	$34.89 \pm 0.18$
$R_{\text{QCD}} [9.3 – 12.0 \text{ GeV}]_{uds cb}$	$1.20 \pm 0.01$	$15.53 \pm 0.04$
$R_{\text{QCD}} [12.0 – 40.0 \text{ GeV}]_{uds cb}$	$1.64 \pm 0.00$	$77.94 \pm 0.13$
$R_{\text{QCD}} [ > 40.0 \text{ GeV}]_{uds cb}$	$0.16 \pm 0.00$	$42.70 \pm 0.05$
$R_{\text{QCD}} [ > 40.0 \text{ GeV}]_t$	$0.00 \pm 0.00$	$-0.72 \pm 0.01$
<b>Sum</b>	$694.0 \pm 1.0 \pm 3.5 \pm 1.6 \pm 0.1_{\psi} \pm 0.7_{\text{QCD}}$	$275.29 \pm 0.15 \pm 0.72 \pm 0.23 \pm 0.15_{\psi} \pm 0.55_{\text{QCD}}$

→ 32 exclusive channels are integrated up to 1.8 GeV

Relative contributions to  $a_{\mu}$  from missing channels (estimated based on isospin symmetry)

→  $0.87 \pm 0.15$  % (DEHZ 2003)

→  $0.69 \pm 0.07$  % (DHMZ 2010)

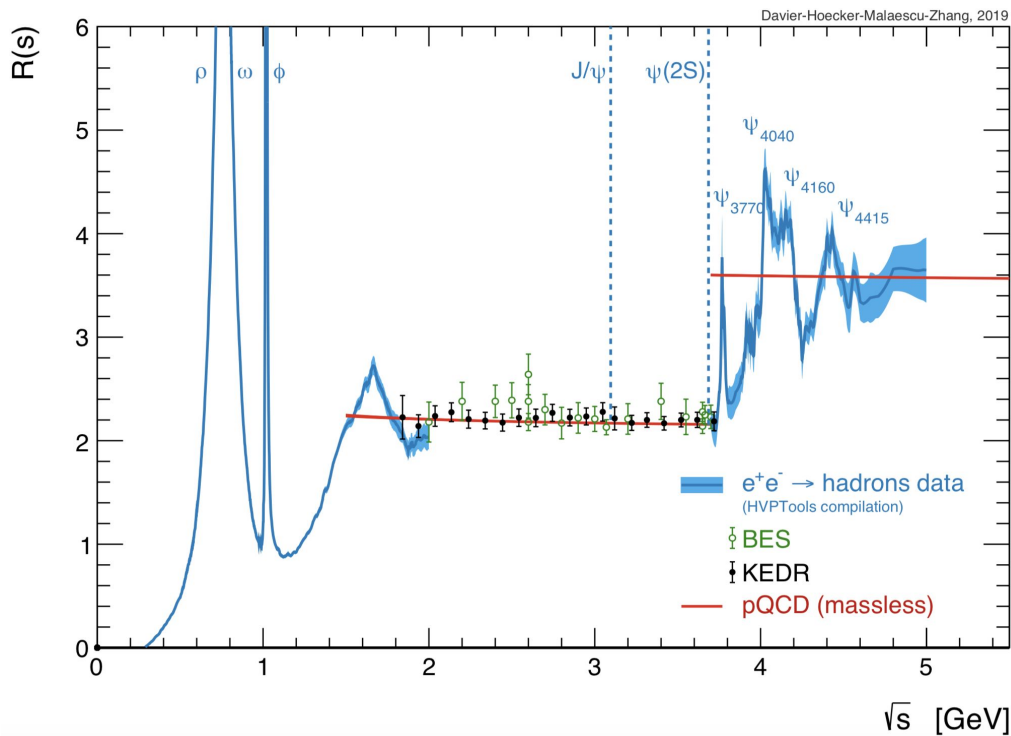
→  $0.09 \pm 0.02$  % (DHMZ 2017)

→  $0.016 \pm 0.016$  % (DHMZ 2019)

(Nearly complete set of exclusive measurements from BABAR)

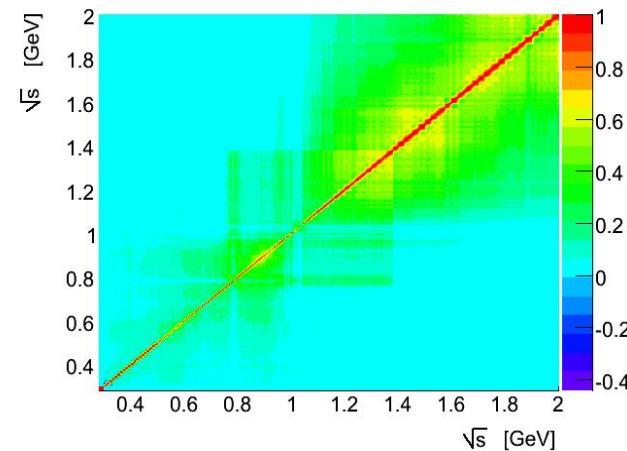
Backup

# $R_{e^+e^-} \rightarrow \text{Hadrons}$

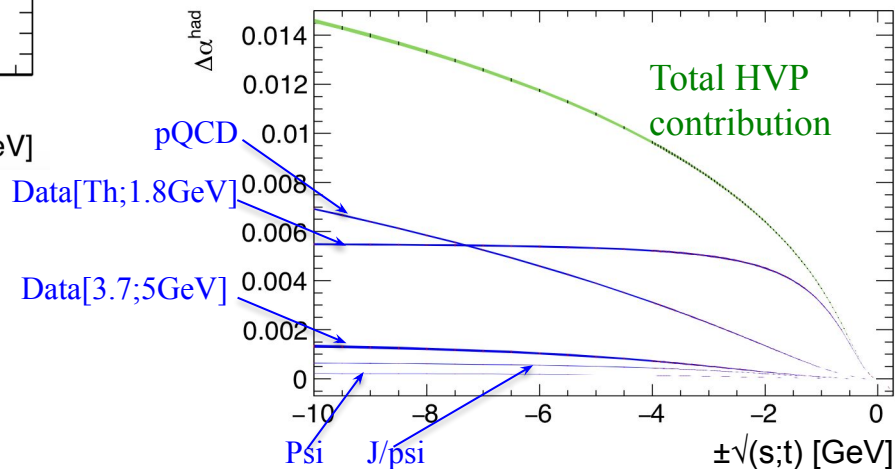


→ Performed non-trivial check:  $a_\mu$  and  $\Delta\alpha_{\text{had}}$  from sum of individual channels and from  $\text{Re}e$  integral  $< 1.8 \text{ GeV}$

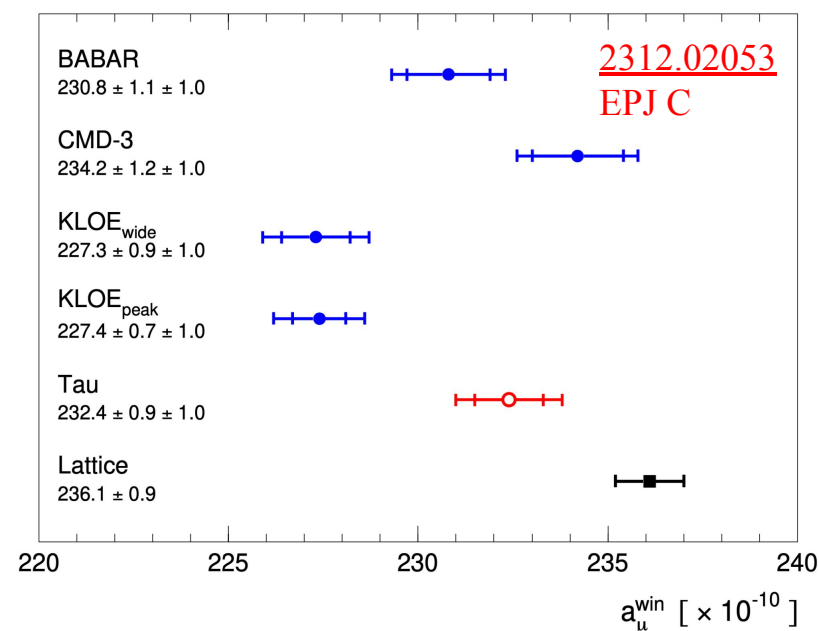
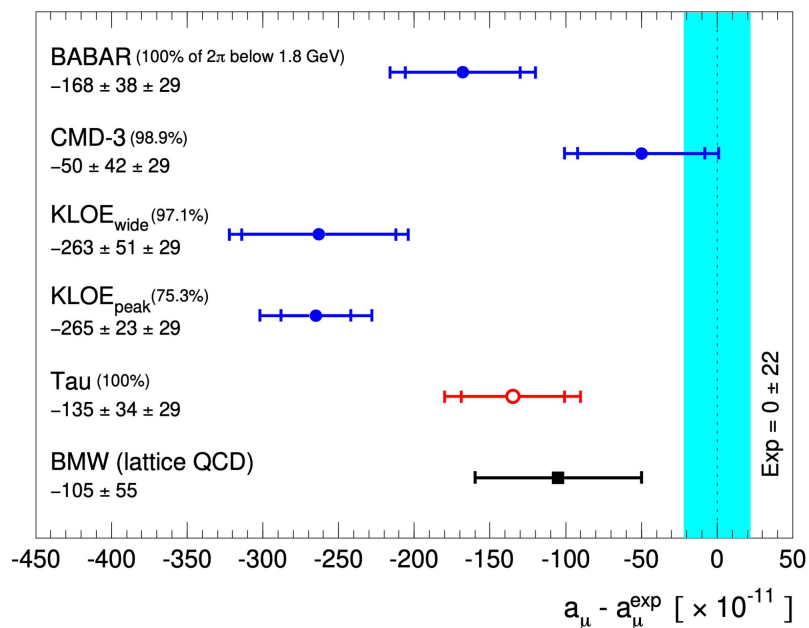
→ Enables the determination of the various HVP contributions to the “running” of  $\alpha_{\text{QED}}$ , evaluation of  $\alpha_s$  and test of RGE with different correlation scenarios for theory uncertainties (*Backup*)



Sum of 32 *exclusive channels* with *full propagation of correlations*



# A new perspective on $a_\mu$ (HVP)



- The  $\tau$ -based HVP contribution close to the values provided by BABAR and CMD-3
- Their combination ( $3.8\sigma > \text{KLOE}_{\text{peak}}$ ) is compatible with BMW for  $a_\mu$ , but a  $2.9\sigma$  tension persists for  $a_\mu^{\text{win}}$
- The BMW-based prediction is  $1.8\sigma$  below the experimental value; *not* incompatible with the EW fit (Backup)
- Combining BABAR, CMD-3,  $\tau$  (+BMW):  $2.5\sigma$  ( $2.8\sigma$ ) difference with experiment
- When including KLOE in the dispersive calculation:  $> 5\sigma$  w.r.t. experiment
- Tests of MC generators using KLOE data & in-situ studies of impact on the analysis are very much desirable

# Lattice calculations and comparisons w.r.t. dispersive

→ Lattice: employ simulations to compute electromagnetic-current two-point function

$$C(t) = \frac{1}{3e^2} \sum_{i=1}^3 \int d^3x \langle J_i(\vec{x}, t) J_i(0) \rangle$$

$$C(t) = \frac{1}{24\pi^2} \int_0^\infty ds \sqrt{s} R(s) e^{-|t|\sqrt{s}}$$

Based on BMW'20: All contributions to  $C(t)$ , with all limits taken:  $a \rightarrow 0$ ,  $L \rightarrow \infty$ ,  $M_\pi \rightarrow M_\pi^\phi$ , ...

→ Tensions between lattice (weighted sums of  $C(t)$  over  $t$ ) and pre-CMD3 data-driven (DD) HVP results

$$[\Delta a_\mu^{\text{LO-HVP}}]_{\text{lat-DD}} \sim 2.1\sigma \quad [\Delta a_{\mu, \text{win}}^{\text{LO-HVP}}]_{\text{lat-DD}} \sim 4.0\sigma \quad [\Delta \alpha_{\text{had}}^{(5)}(-1 \rightarrow -10 \text{GeV}^2)]_{\text{lat-DD}} \sim 1.4\sigma, \dots$$

→ Simultaneous comparisons with correlations:  $\sim$  dilution compared to  $a_{\mu, \text{win}}^{\text{LO-HVP}}$  alone, but still significant tension

→ *New in this study*: Correlations among lattice HVP observables and uncertainties on these correlations

→ Differences could be explained by: a  $C(t)$  that is enhanced in  $t \sim [0.4, 1.5]$  fm

→ Lattice  $\rightarrow$  R-ratio: inverse Laplace transform (ill-posed problem)

→ *Developed dedicated statistical methods, benefiting from experience with experimental measurements*

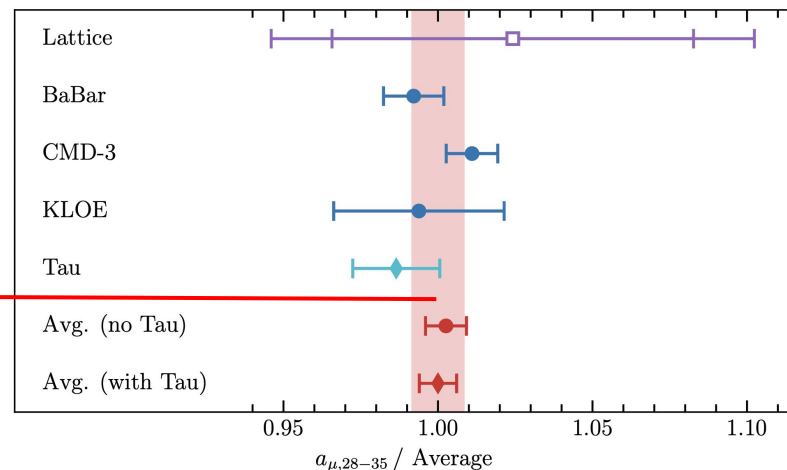
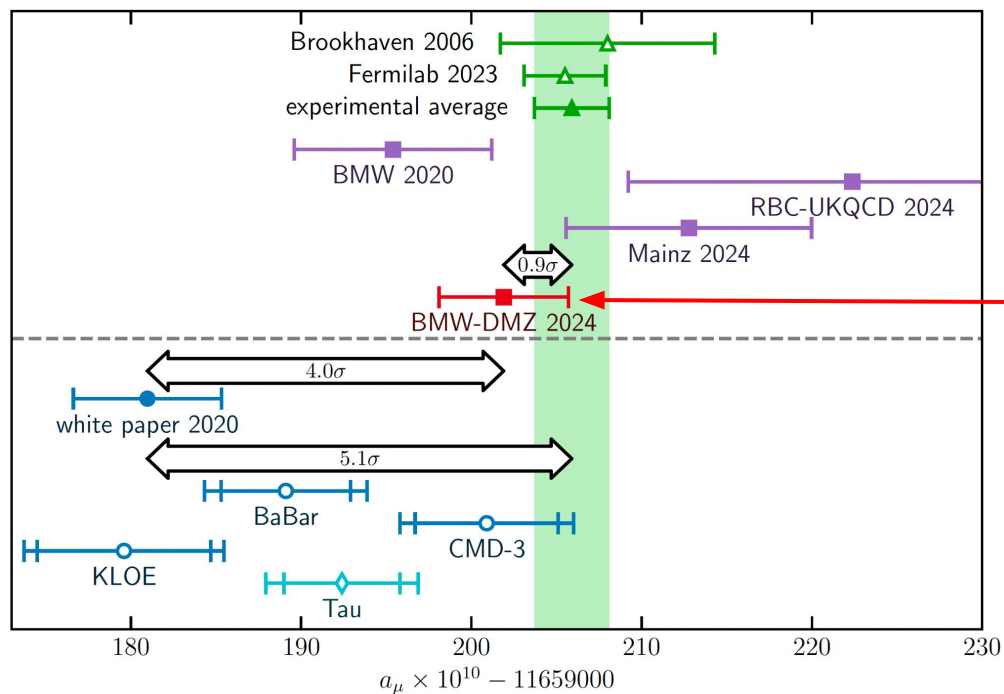
→ Differences could be explained by enhancing measured R-ratio around (/any larger interval including)  $\rho$ -peak, but rescalings beyond the uncertainties of  $\text{Re}^+e^-$

[2308.04221](#)

→ Outcome of the studies stable within stat. and syst. uncertainties on lattice covariance matrices

*Backup*

# Merging dispersive and lattice HVP calculations



[2407.10913](https://arxiv.org/abs/2407.10913)

→ *Currently most precise prediction, based on improved lattice QCD (BMW) + data-driven inputs (DMZ) at large- $t$  (input data in good agreement at low energy)*

*Backup*

# Uncertainties on uncertainties and correlations

*Numerous indications of uncertainties on uncertainties and on correlations, with a direct impact on combination fits*

- Shapes of systematic uncertainties *evaluated* in  $\sim$ -wide mass ranges with sharp transitions
- One standard deviation is statistically not well defined for systematic uncertainties
- Systematic uncertainties like acceptance, tracking efficiency, background etc. not necessarily fully correlated between low and high mass
- Are all systematic uncertainty components fully independent between each-other? (e.g. tracking / trigger)
- *Yield uncertainties on uncertainties and on correlations*
- Tensions between measurements (BABAR/KLOE/CMD3; 3 KLOE results etc.):  
*experimental indications of underestimated uncertainties*
- *Statistical methods ( $\chi^2$  with correlations, likelihood fits, ratios of measured quantities etc.) should not over-exploit the information on the amplitude and correlations of uncertainties*

*Topic of general interest, in other fields too (e.g. ATLAS JES and Jet Xsec studies)*

*Backup*



# Remarks and conclusions on $(g-2)_\mu$

*We have an interesting, long standing, multifaceted problem...*

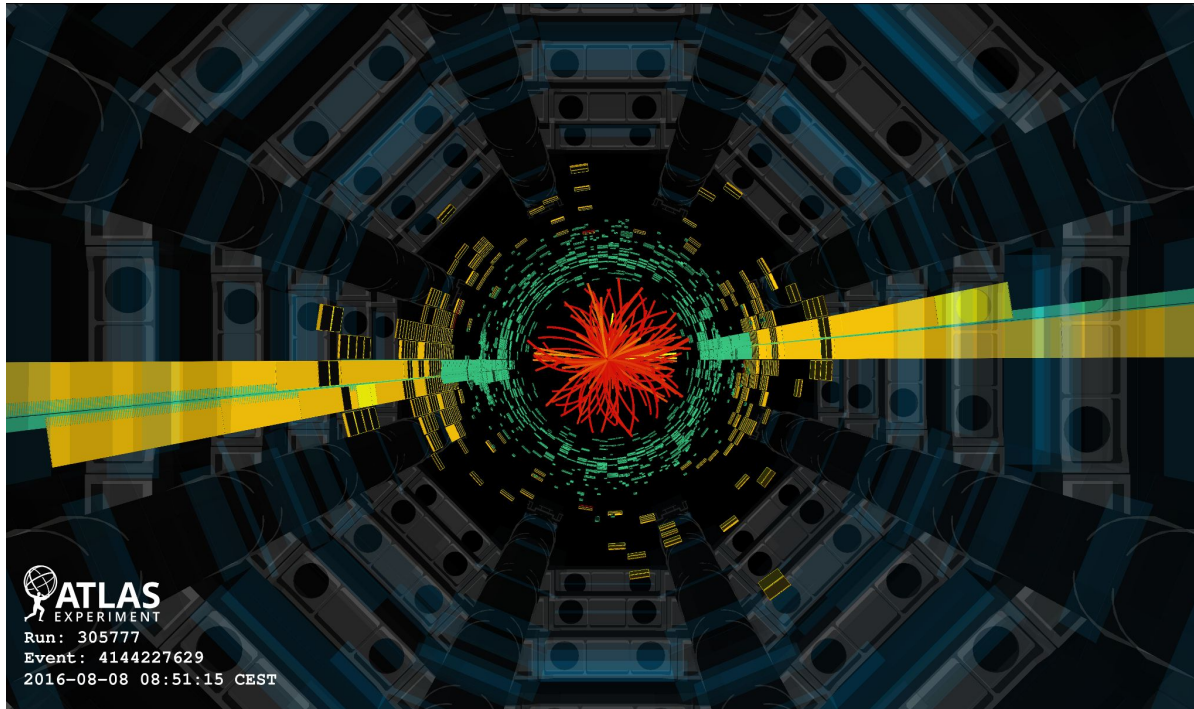
*... And very important elements to solve the puzzle started to become available !*

- Future  $e^+e^-$  measurements very important: independent  $2\pi$  measurement from BaBar w/o PID this year  
Long-term collaboration with M. Davier, A.-M. Lutz, Z. Zhang  
Intense work by Leonard Polat & Andres Pinto (*PostDocs*)

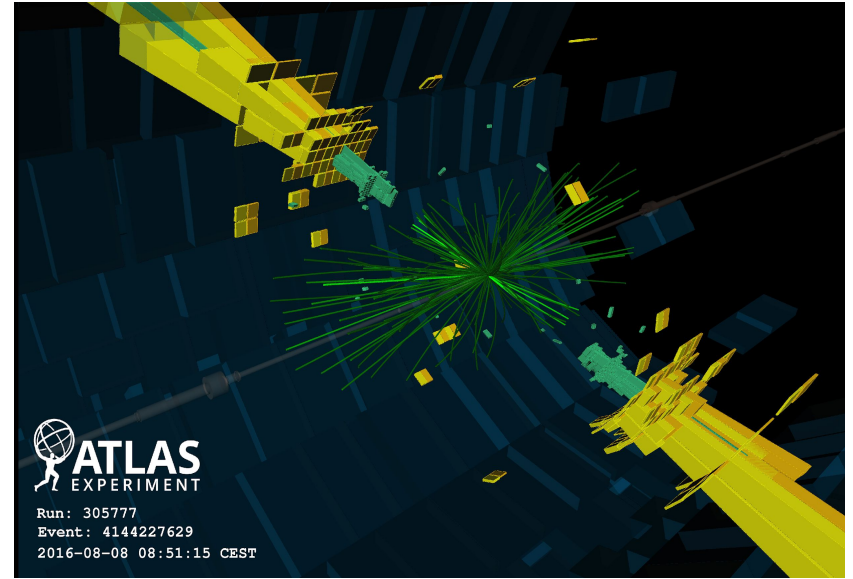
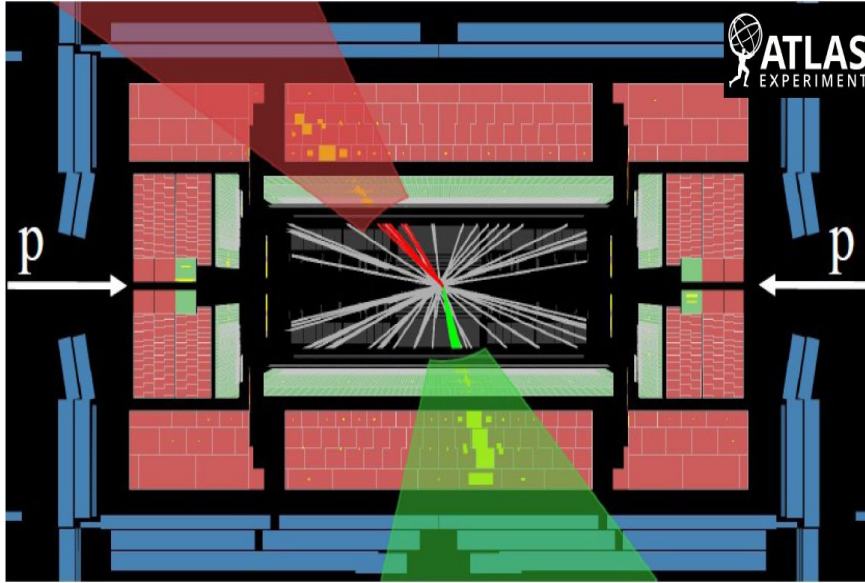
## Guiding ideas:

- Need *rigorous and realistic* treatment of uncertainties and correlations at all levels  
(Underestimated uncertainties do not bring scientific progress & can put studies on wrong path)
- Caution about significance:  
statistics-dominated measurement; prediction uncertainty limited by non-Gaussian systematic effects
- Studies for understanding differences between data-driven and Lattice QCD approaches need to follow similar standards as the  $g-2$  experiment: *double-blinding*

# Jet studies with ATLAS @ LHC



# Jets @ LHC

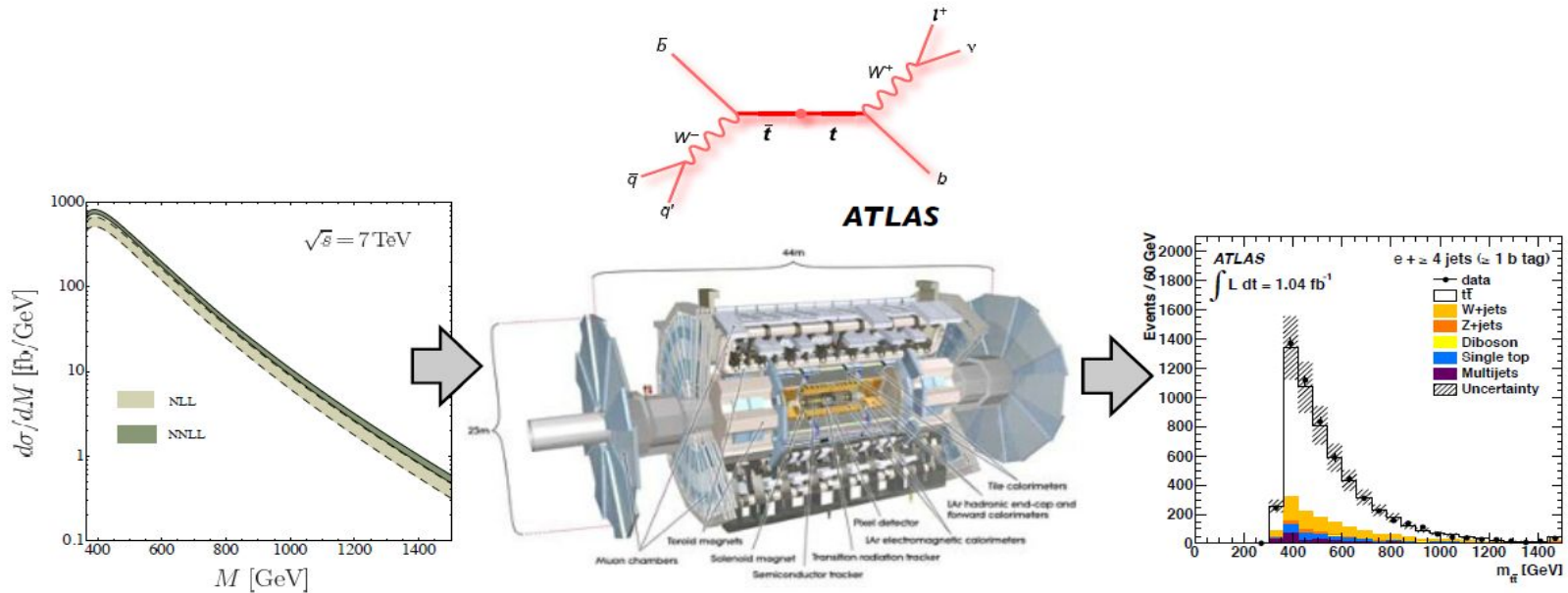


**Jets:** “sprays” of (quasi-)stable particles called hadrons, observed in the ATLAS detector

- Proxy to fundamental interactions in Nature, probing the smallest scales accessible in laboratory:  
Test SM on wide phase-space; important ingredients to  $\alpha_s$  and PDF fits; sensitivity to New Physics
- Precise and robust definition necessary for any quantitative studies:  
anti- $k_t$  - infrared and collinear safe

*Backup*

# Example of (un)folding problem @ LHC



Folding

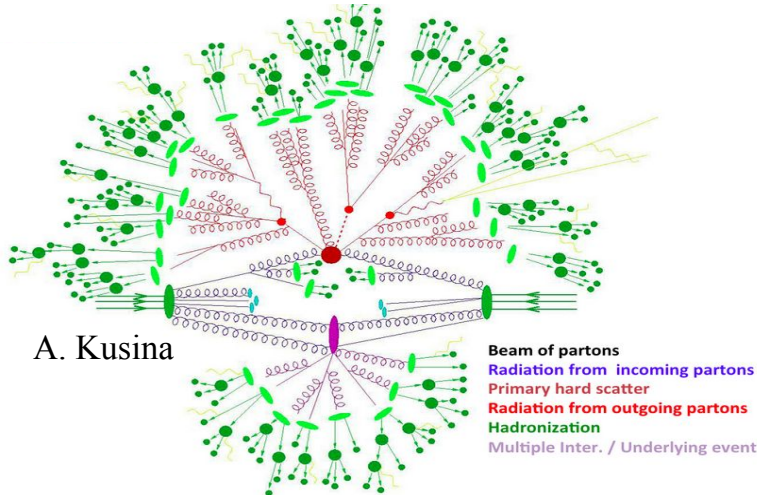


Calibration + Unfolding



# Environment and unfolding strategy for jet studies @ LHC

Typical proton-proton collision: a complex process in a difficult environment

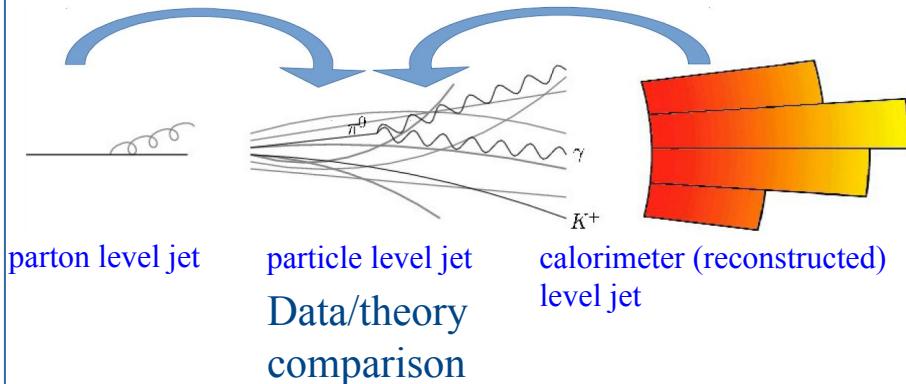


Pile-up



NP corrections  
Hadronization & UE

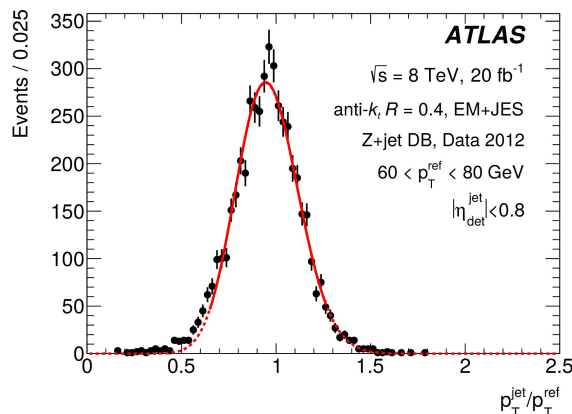
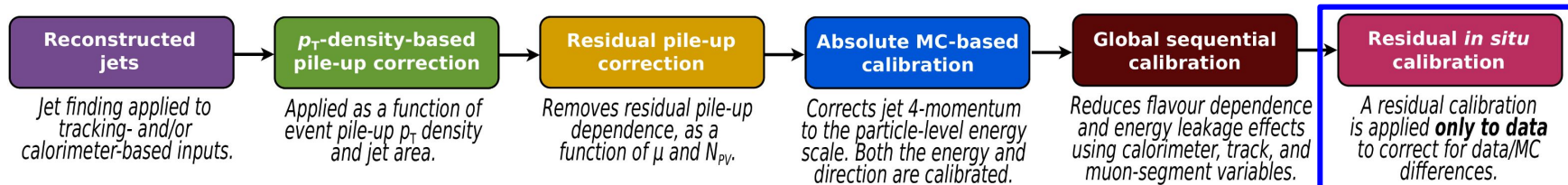
Calibration+Unfolding  
Jet energy response & resolution



Goal: *publish data “corrected for detector effects” (on average, in the sense of an estimator), with minimal bias and minimal model dependence, with the full information needed for comparisons with theory predictions*

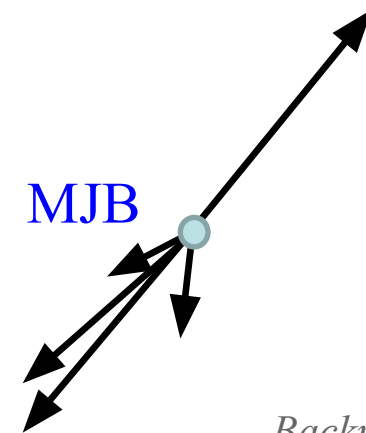
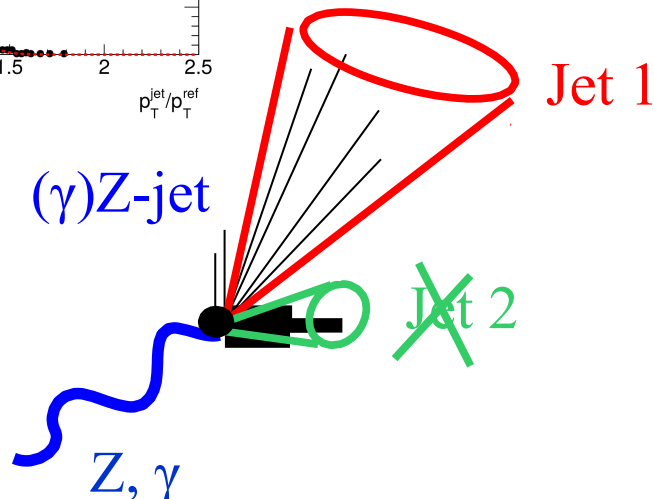
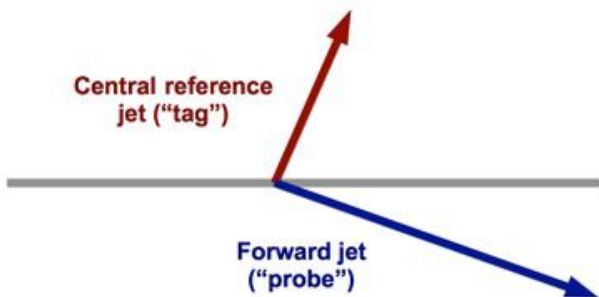
→ Typically implies unfolding to hadron level although there are cases where one can unfold to parton level

# Jet reconstruction and calibration procedure



$$c = \frac{\mathcal{R}_{in-situ}^{data}}{\mathcal{R}_{in-situ}^{MC}} \equiv \frac{\left( \frac{p_T^{jet}}{p_T^{reference}} \right)_{data}^{reco}}{\left( \frac{p_T^{jet}}{p_T^{reference}} \right)_{MC}^{reco}}$$

## Dijet balance



Backup

# In-situ $\eta$ – intercalibration calibration method

$$\mathcal{A} = \frac{p_T^{\text{left}} - p_T^{\text{right}}}{p_T^{\text{avg}}} \quad \mathcal{R} = \frac{c^{\text{right}}}{c^{\text{left}}} = \frac{2 + \langle \mathcal{A} \rangle}{2 - \langle \mathcal{A} \rangle} \approx \frac{\langle p_T^{\text{left}} \rangle}{\langle p_T^{\text{right}} \rangle}$$

$$S(c_{1x}, \dots, c_{Nx}) = \sum_{j=2}^N \sum_{i=1}^{j-1} \left( \frac{1}{\Delta \langle \mathcal{R}_{ijx} \rangle} (c_{ix} \langle \mathcal{R}_{ijx} \rangle - c_{jx}) \right)^2 + X(c_{ix})$$

→  $p_T$  balance in dijet *in-situ* events to achieve homogeneity of the calibration in  $\eta$ : using multiple combinations of central-forward bins to obtain a better statistical precision

→ Over-constrained system: gained a factor  $\sim 1000$  in speed by using analytic solution

(Robert Hankache - PhD)

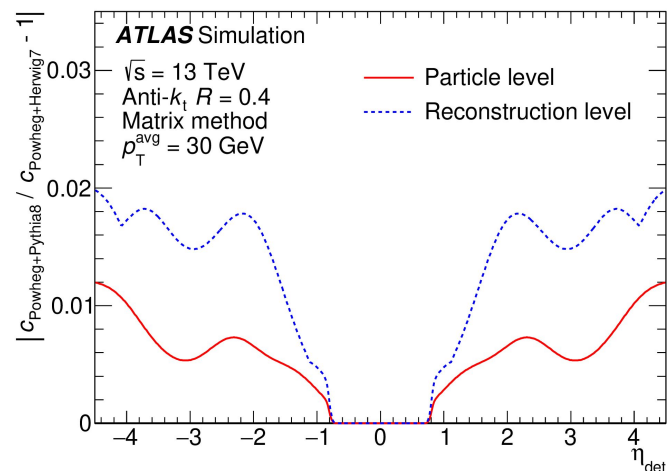
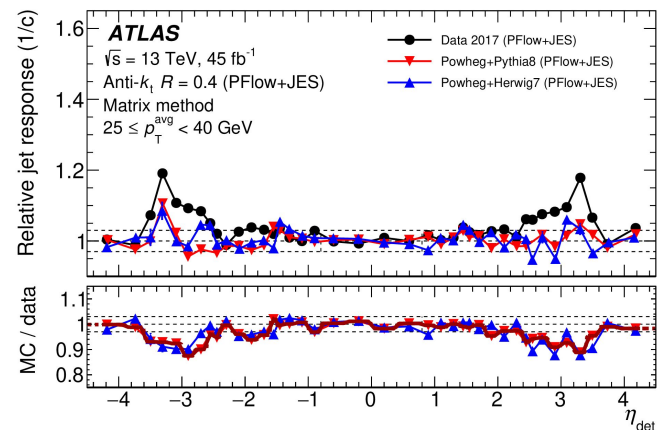
→ Reduced modeling uncertainty by a factor  $\sim 2$ , avoiding double-counting of detector effects

(Louis Ginabat - PhD)

→ Detailed study of the compatibility of constraints, improved statistical uncertainties and correlations

(Line Delagrangue - PhD)

→ Input for ML-based calibration (Laura Boggia - PhD)



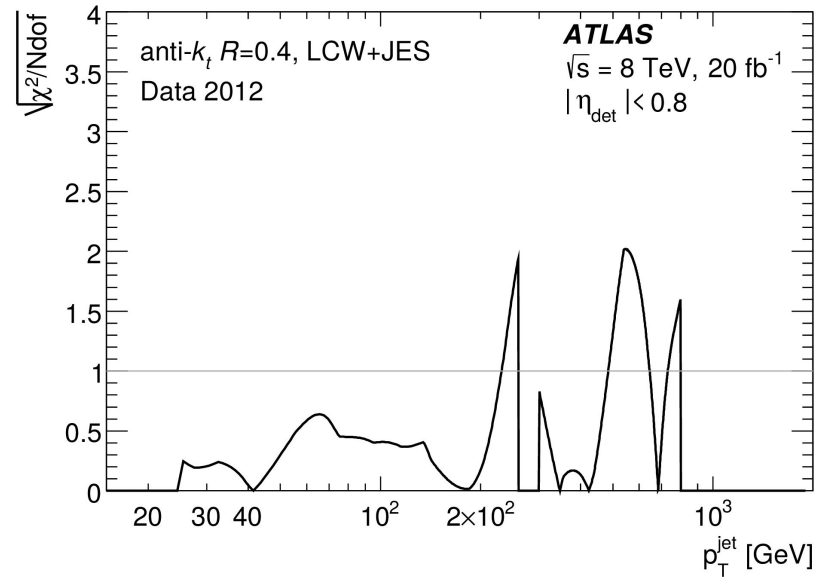
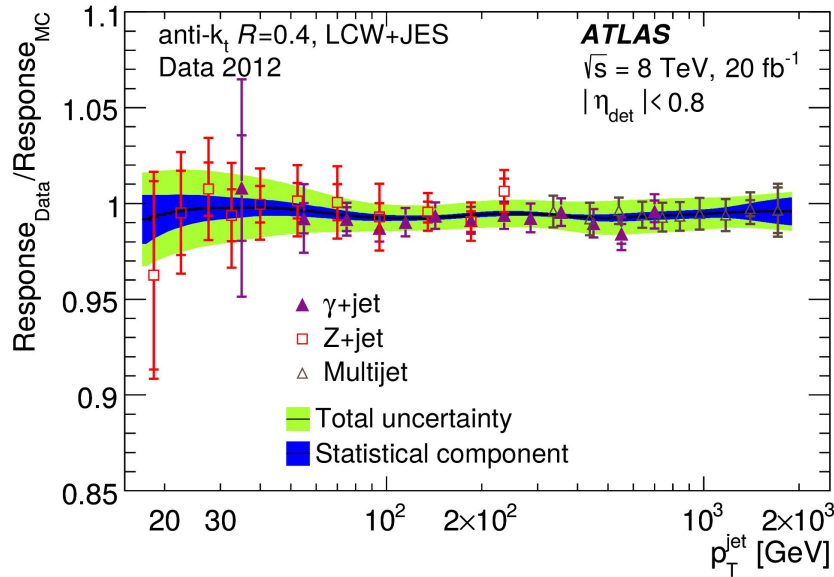
Backup

# $\gamma$ -jet + Z-jet + MJB combination

→ Combination of in-situ results with spline-based interpolations + weighted averages

→ Testing also compatibility of in-situ inputs

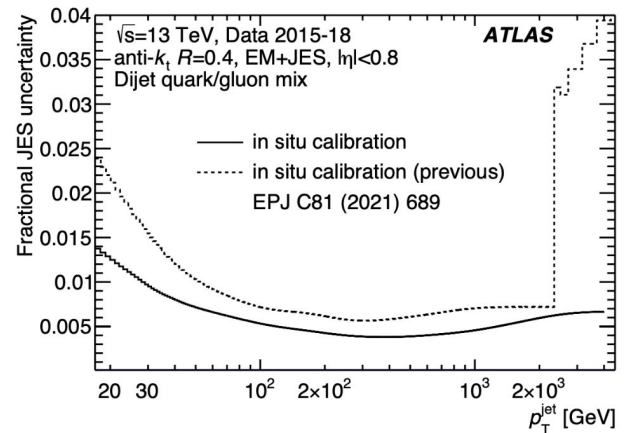
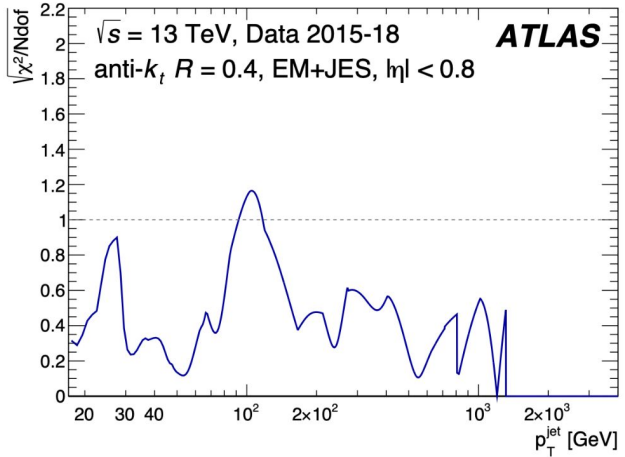
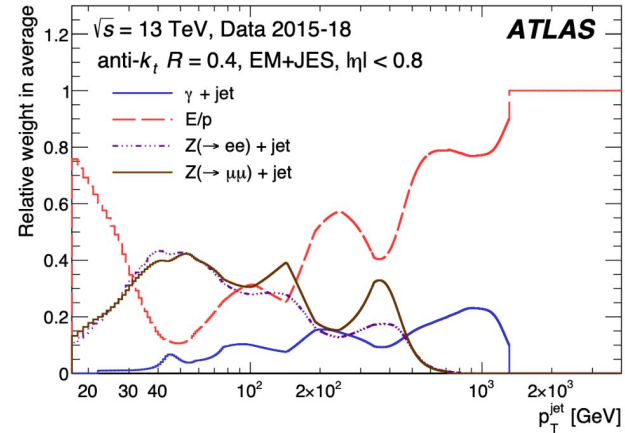
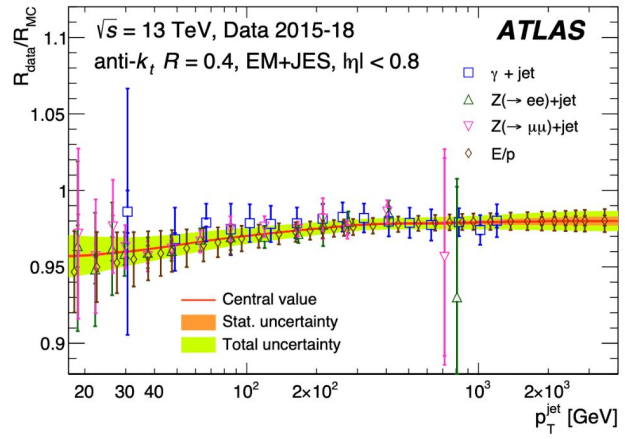
(Inspired by methodology employed for hadronic spectra)



→ Includes Z+jet method for jet calibration and resolution (*Guillaume Lefebvre - PhD*)



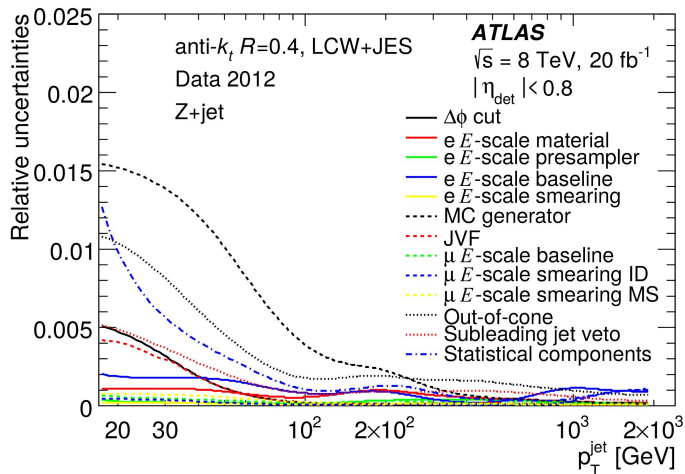
# $\gamma$ -jet + Z-jet + E/p combination



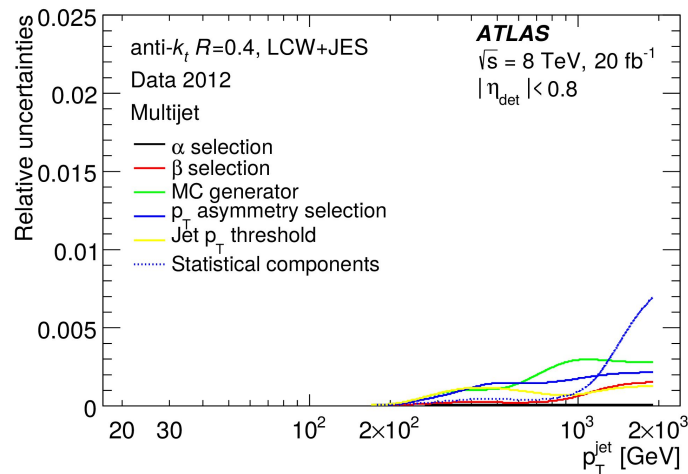
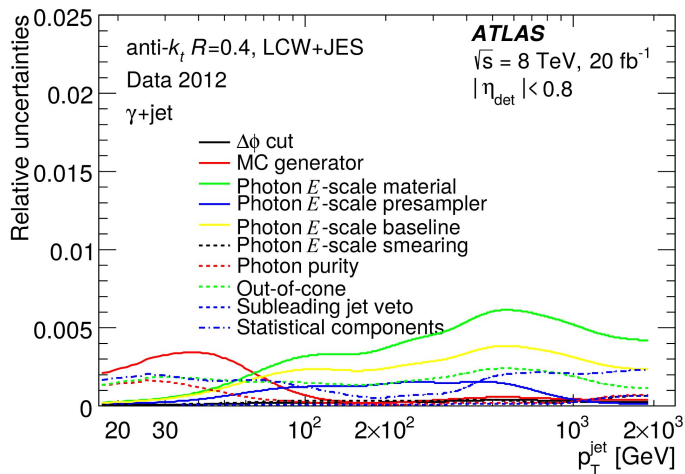
→ Important improvement due to E/p method (*Lata Panwar - PostDoc*)

Backup

# In-situ uncertainties affecting the combination result



→ 47 in-situ uncertainty components (NPs) :  
full propagation of information on uncertainty & correlations



# Using a diagonalization procedure to reduce the number of NPs

$$C = S \cdot D \cdot S^T$$

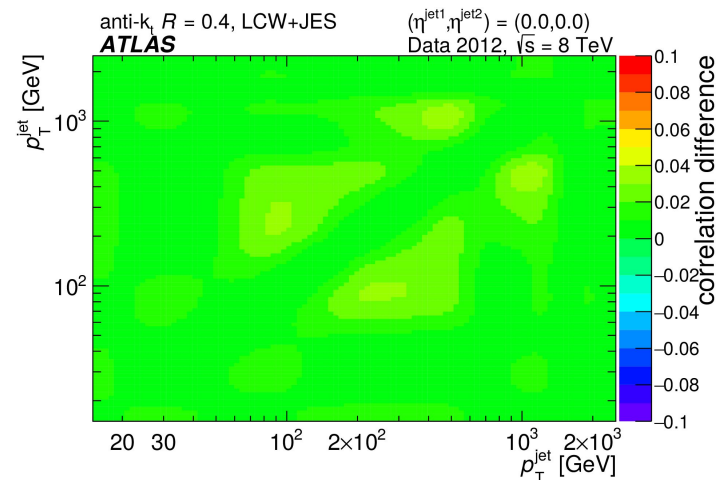
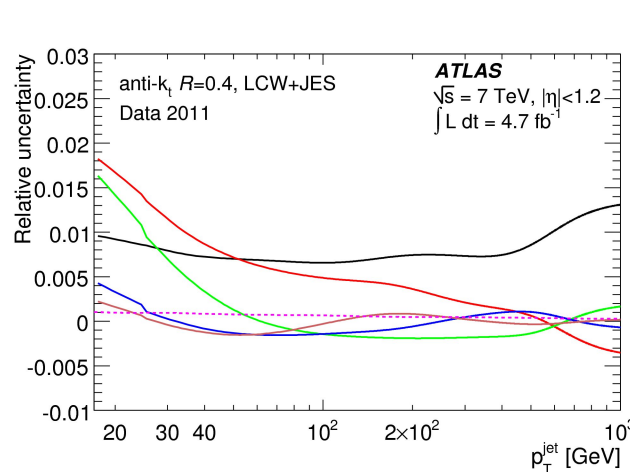
$$D = \begin{pmatrix} \dots & 0 & 0 \\ 0 & \sigma_i^2 & 0 \\ \dots & 0 & \dots \end{pmatrix}$$

$$S = \begin{pmatrix} V_1 & \dots & V_n \\ \vdots & & \vdots \end{pmatrix}$$

- $V_i$  = eigen vectors
- $\sigma_i^2$  = eigen values

Going from 47 in-situ NPs to 6 NPs

$$C = \sum_{i=1}^{N_{bins}} \sigma_i^2 C(V_i) \approx \sum_{i=1}^{N_{ev}-1} \sigma_i^2 C(V_i) + C'(\text{other e. v.})$$



→ Use part ( $N_{ev} - 1$ ) of the important (large) eigenvalues, plus an effective contribution for the others (rest term), to approximate the covariance matrix: percent-level precision on correlations (difference between original and approximate matrices)

→ Keep track of uncertainty origin using a reduction by category:

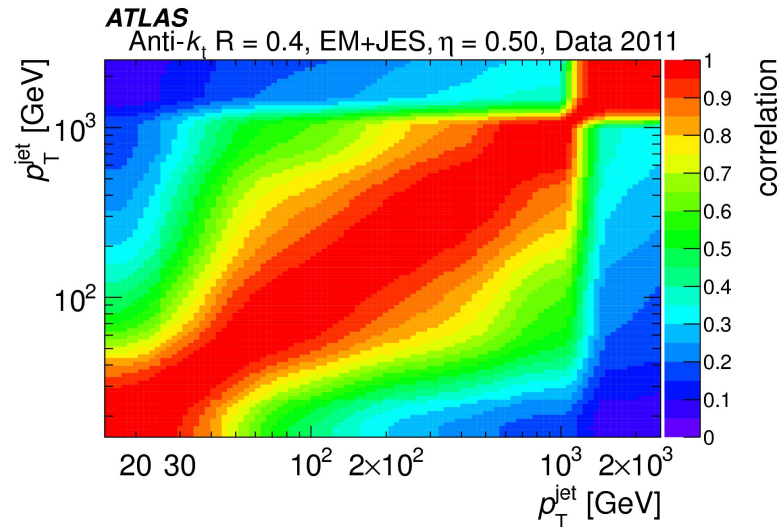
“statistical”, “modeling”, “detector”, “mixed (modeling/detector)”, “special”

# Uncertainties on JES correlations

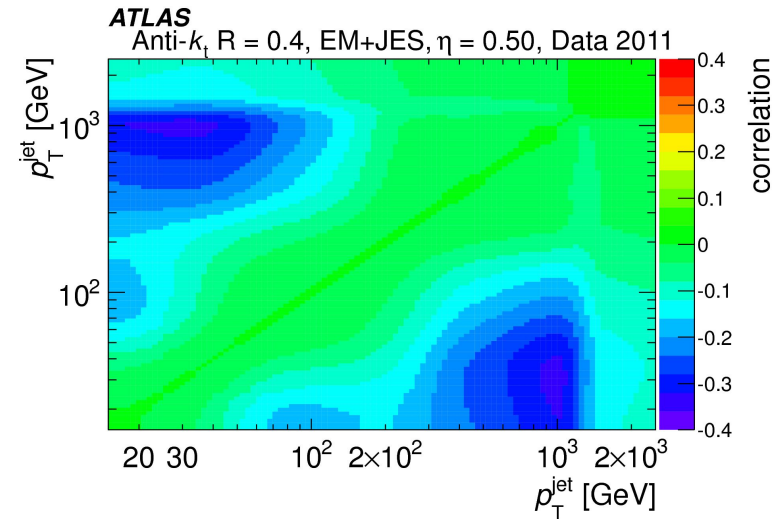
→ Derived two alternative configurations with stronger/weaker correlations w.r.t. nominal

*(Inspired remarks about uncertainties on uncertainties for combinations of hadronic spectra)*

*Nominal correlation matrix*



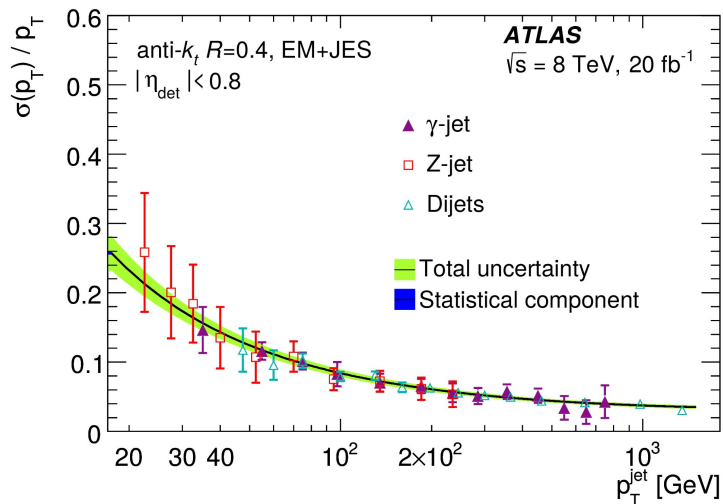
*Strong - Weak correlation scenarios*



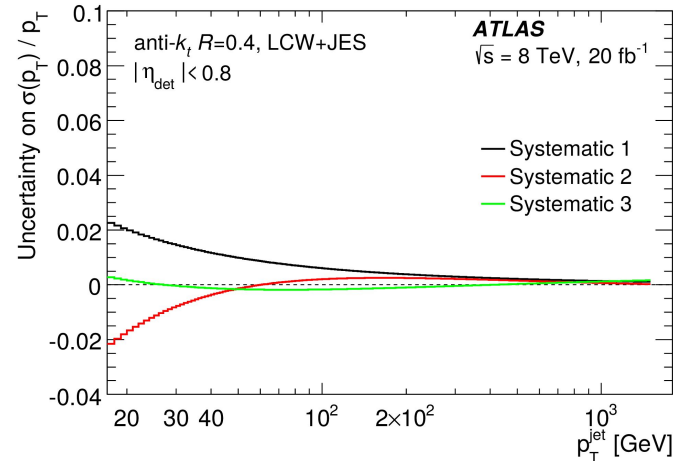
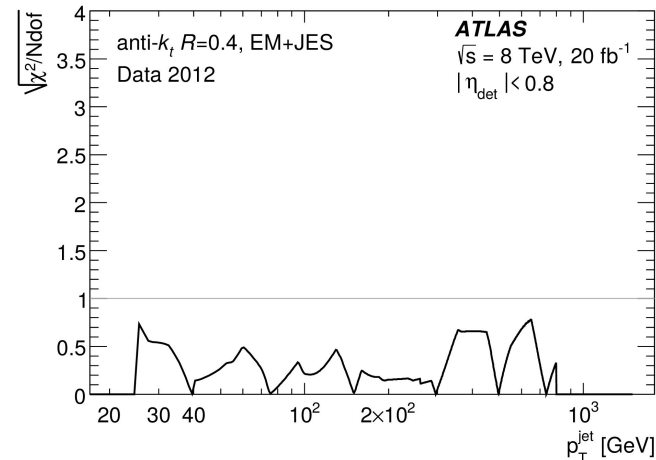
# First $\gamma$ -jet + Z-jet + Dijets + Zero Bias JER fit

- Good agreement (local  $\chi^2$  with correlations) between in-situ methods
- In-situ combination based on global fit of (N, S, C)
- Full propagation of uncertainties & correlations:  
3 eigenvectors are enough for  $\sim 10^{-3}$  precision on correlations

$$\frac{\sigma(p_T)}{p_T} = \frac{N}{p_T} \oplus \frac{S}{\sqrt{p_T}} \oplus C$$

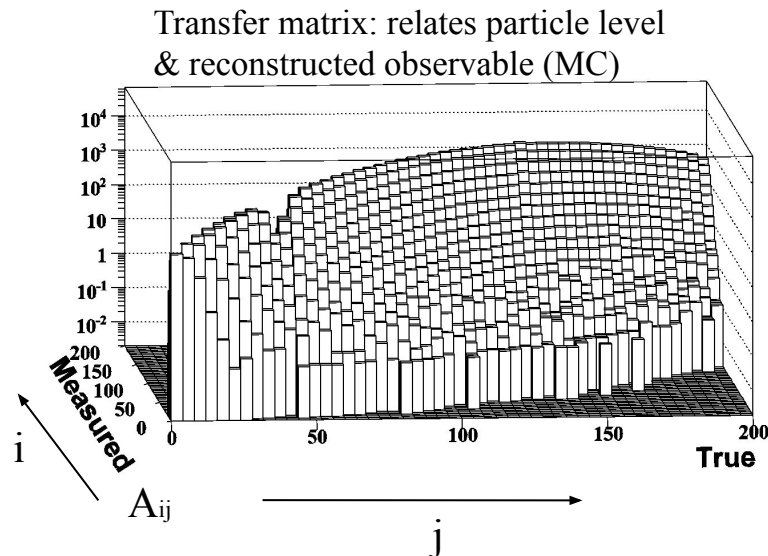
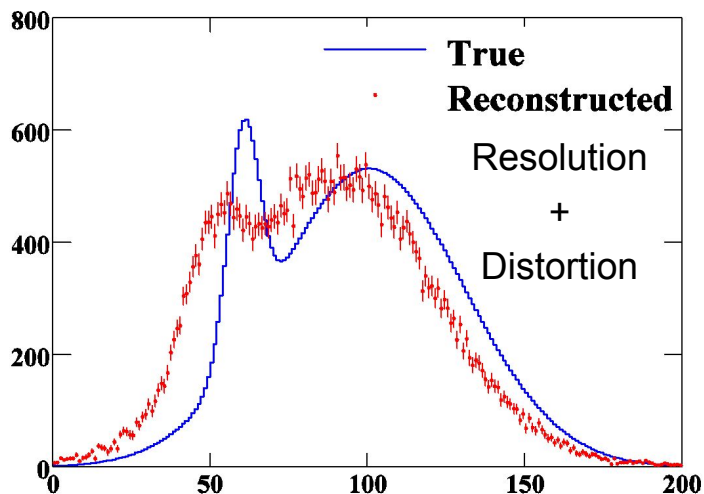


- Coherent propagation of uncertainties & correlations in physics analyses (*Dimitris Varouchas - PostDoc*)



*Backup*

# Detector effects, folding and unfolding



→ Folding:  $f_{\text{meas}}(y) = \int R(y|x) f_{\text{true}}(x) dx$

$$P_{ij} = \frac{A_{ij}}{\sum_{k=1}^{n_d} A_{kj}} ; d = P \cdot t$$

→ Unfolding of detector smearing effects is generally not a simple numerical problem

Regularization methods are often necessary

*Backup*

# Data unfolding: methodology & studies

- Measurements corrected back to truth particle level using a matrix-based unfolding method
  - transfer matrix relating the true & reconstructed observable (MC): matching needed
  - 3 steps procedure: 1) matching (in)efficiency correction at reco level;
    - 2) IDS / SVD / bin-by-bin / IBU unfolding for jets with matching;
    - 3) matching (in)efficiency correction at true level.
- Numerous aspects studied in this context, among which:
  - choice of the phase-space
  - optimization of the choice of the binning
  - intrinsic unfolding uncertainties and choice of the regularization
  - propagation of statistical and systematic (JES, JER etc.) uncertainties, with their correlations
  - non-linear effects in the uncertainty propagation
  - statistical noise in the uncertainty propagation
  - modeling uncertainties and “hidden variables”
- Developments of ML-based methods: will enable qualitative improvements of such measurements

# Data-driven closure test: motivation, procedure, example

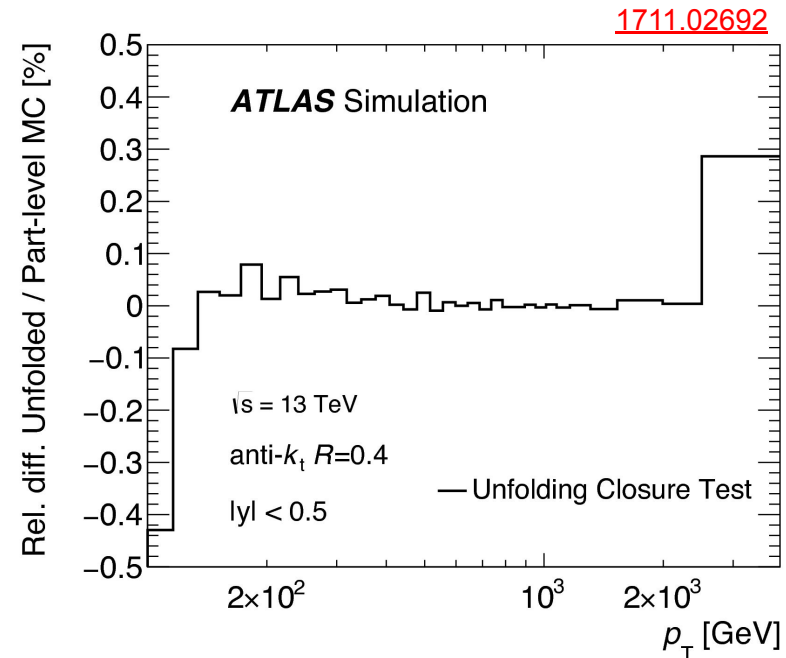
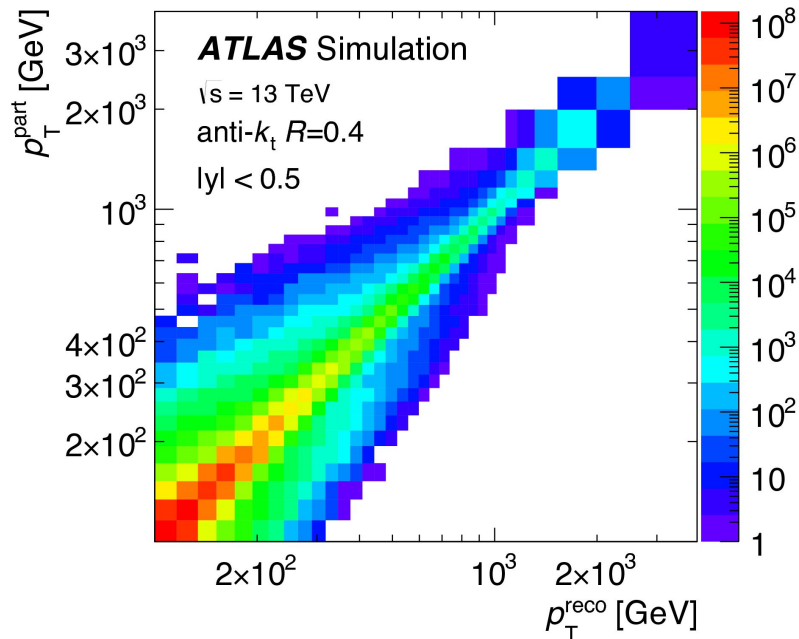
→ In-situ (i.e. *realistic*) determination of the unfolding uncertainty related to the data/MC shape difference and to the regularization (performed for several unfolding methods; choosing the most precise)

- reweight true MC by (smooth) function: improved data/recoMC agreement

Reweighting performed within fine bins / event-by-event

- unfold the reweighted reconstructed MC

- compare with reweighted true MC





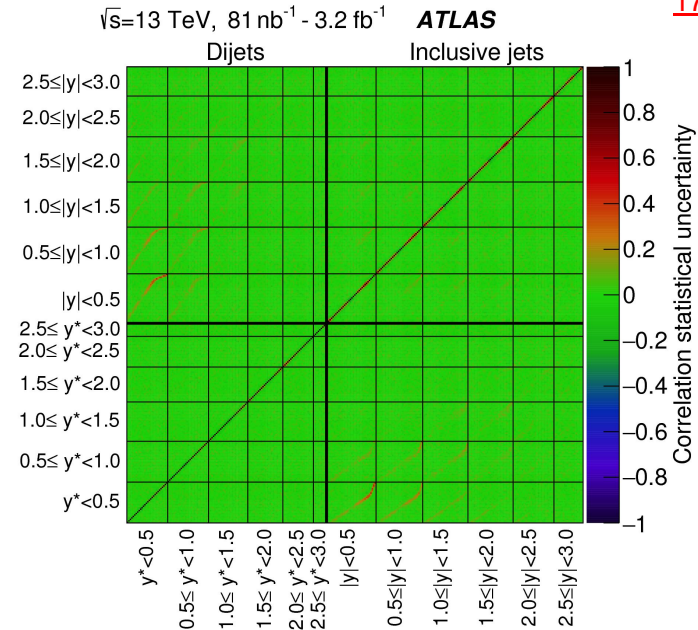
# Statistical uncertainties

- Due to data and MC
- Propagated using pseudo-experiments done separately/simultaneously for data and MC

## → Bootstrap method

- multiply event weights  
by random number:  $\text{Poisson}(1)$
- seed given by event number
- allows to correlate measurements  
with overlapping samples

[ATL-PHYS-PUB-2021-011](#)

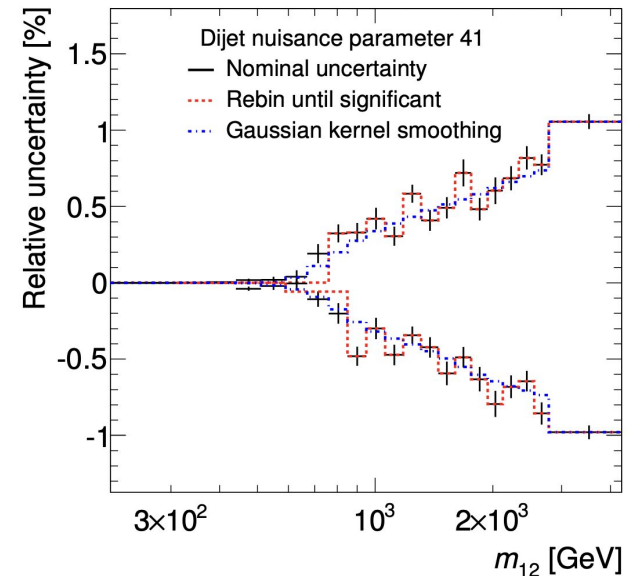
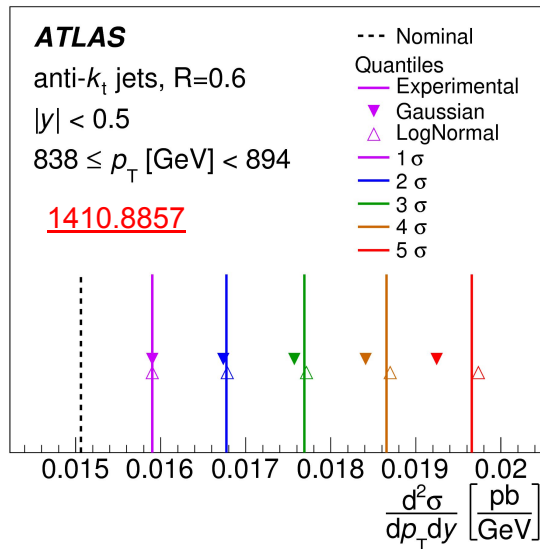


- Publish covariance matrix and/or a series of results based on each pseudo-experiment (i.e. Bootstrap replicas)

# Propagation of systematic uncertainties from inputs

→ Modify input (pseudo-)data spectrum by  $\pm 1\sigma$  of the (asymmetric) uncertainty, re-do unfolding and compare with nominal result

→ Can also use 1...5 $\sigma$  scans or pseudo-experiments



→ Bootstrap method to evaluate statistical uncertainties on the propagated systematics + rebinning/smoothing

Evaluate statistical significance to avoid multiple-counting of statistical uncertainties ([1312.3524](#))

→ For resolution uncertainties, perform smearing of the transfer matrix: smearing factor given by quadratic difference between resolution enhanced by  $1\sigma$  and nominal resolution

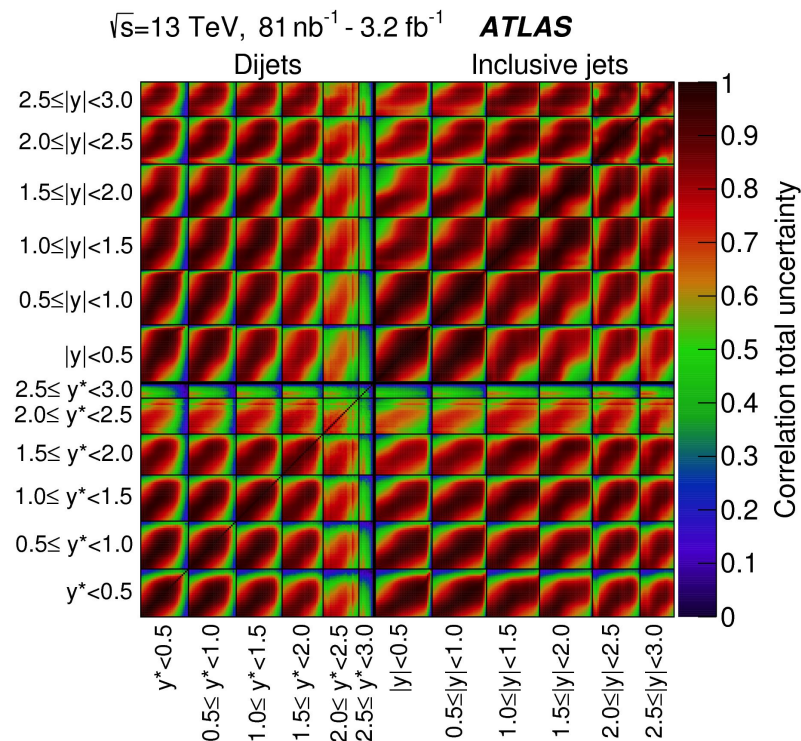
# Propagation of systematic uncertainties from inputs

→ Split of systematics in sub-components (fully correlated in phase-space, independent between each-other) allows to evaluate correlations between different phase-space regions and between different measurements

Relevant when effectively merging uncertainty components in ML-based methods

→ Information made available in HEPData tables (<http://hepdata.cedar.ac.uk/>)

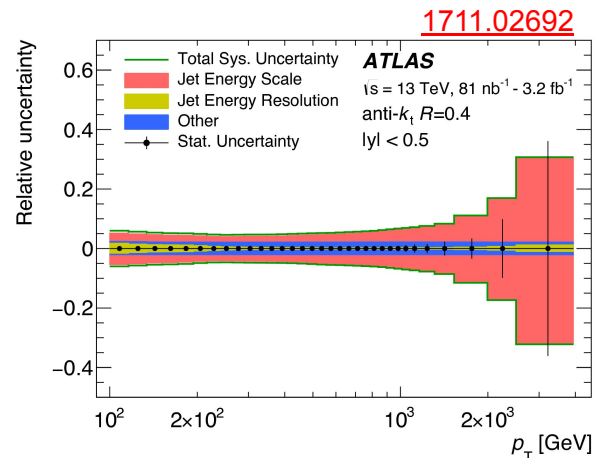
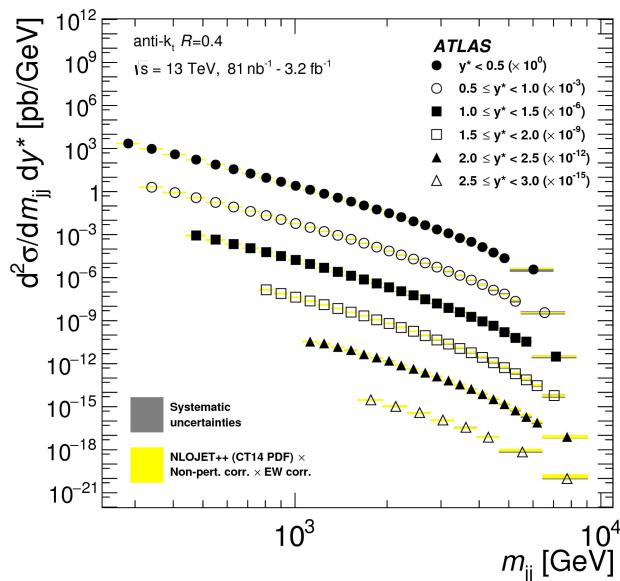
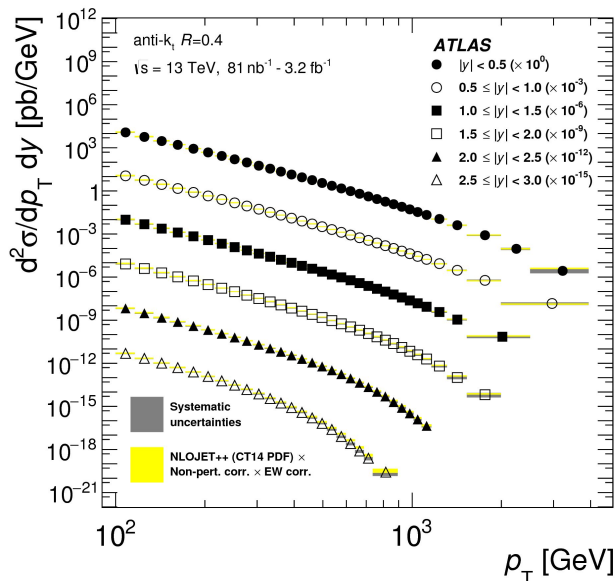
$$Cov_{ij} = \sum_{k=1}^{N_{syst}} s_i^k \cdot s_j^k$$



[1711.02692](#)

# Inclusive jet and dijet cross sections - ATLAS

→ Double-differential measurements for anti- $k_T$  jets with  $R=0.4$ ,  $\sqrt{s}=13$  TeV,  $L=3.2\text{fb}^{-1}$   
 ( $p_T^{\text{jet}} ; |y|$ ) ( $m_{\text{jj}} ; y^*$ ) compared to NLO pQCD + Non-pert. & EW corrections



- At least 2 jets:  $p_T^{\text{jet}} > 75$  GeV,  $|y| < 3$
- $p_T^{\text{jet } 1} + p_T^{\text{jet } 2} > 200$  GeV

→ Uncertainties (  $\sim 5\%$  on wide range, sub-% statistical → *precision era* )

Backup

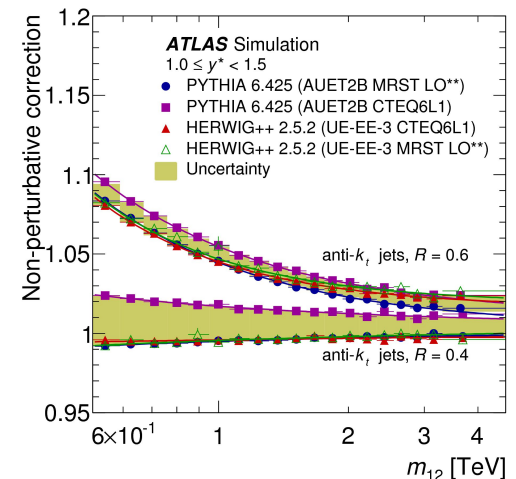
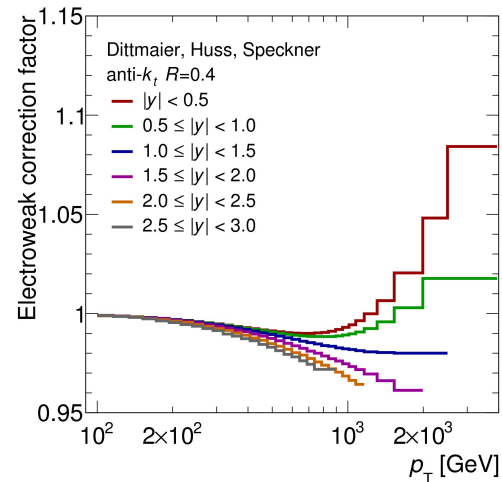
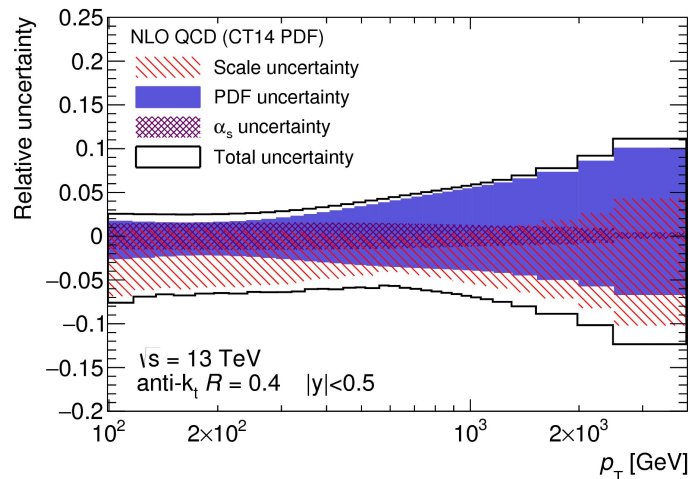
# Theoretical predictions and uncertainties

→ Perturbative QCD predictions from NLOJET++

- Uncertainties: renormalization & factorization scales (0.5 / 2 variations +  $p_T^{\text{jet}}$  vs.  $p_T^{\text{max}}$  scale choice), PDFs and  $\alpha_s$  via APPLGRID
- NNLO prediction (APPLfast)

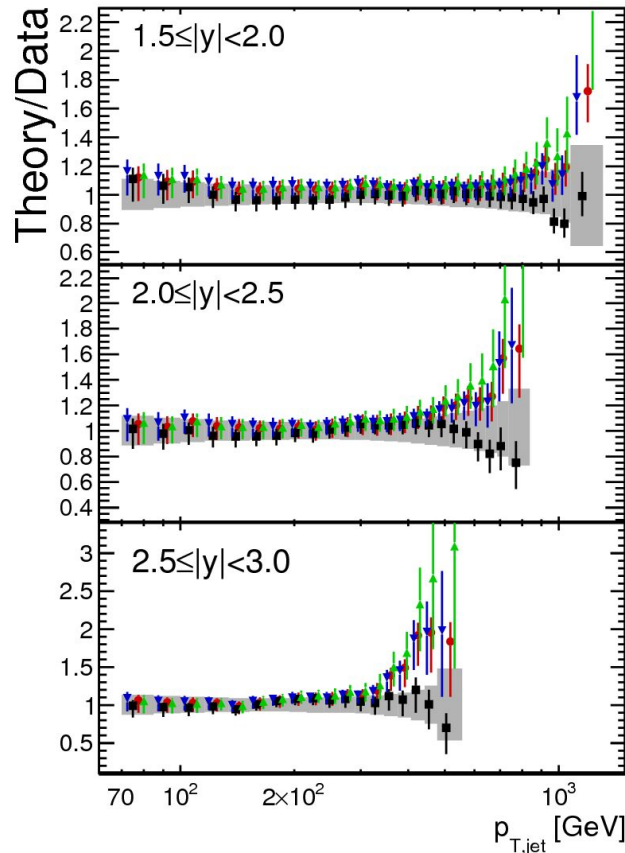
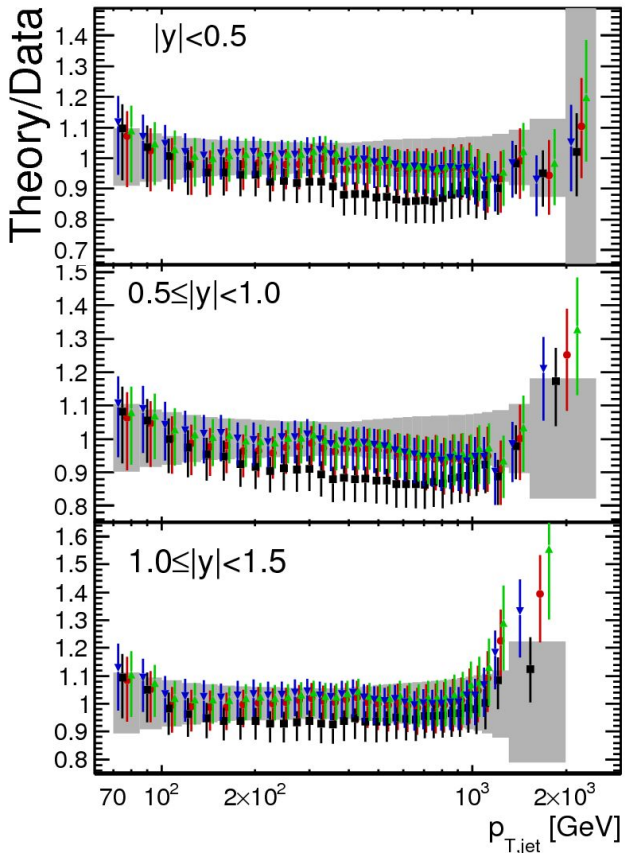
→ EW corrections

→ Non-perturbative corrections (accounting for hadronization and UE) and uncertainties: various Pythia tunes + different MC generators(Herwig++); **strong dependence on R**  
- additional comparisons to Powheg (NLO ME + PS)



# Inclusive jet cross sections at $\sqrt{s}=8$ TeV: Theory/Data

→ Good data/theory agreement within uncertainties observed for most PDF sets



**ATLAS**

$L = 20.2 \text{ fb}^{-1}$

$\sqrt{s} = 8 \text{ TeV}$

anti- $k_t$   $R = 0.6$

■ Data

NLO QCD

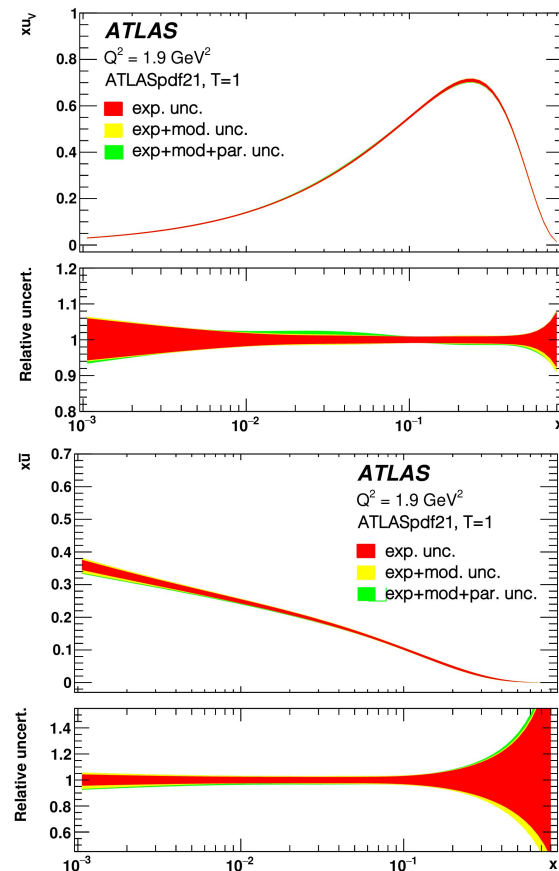
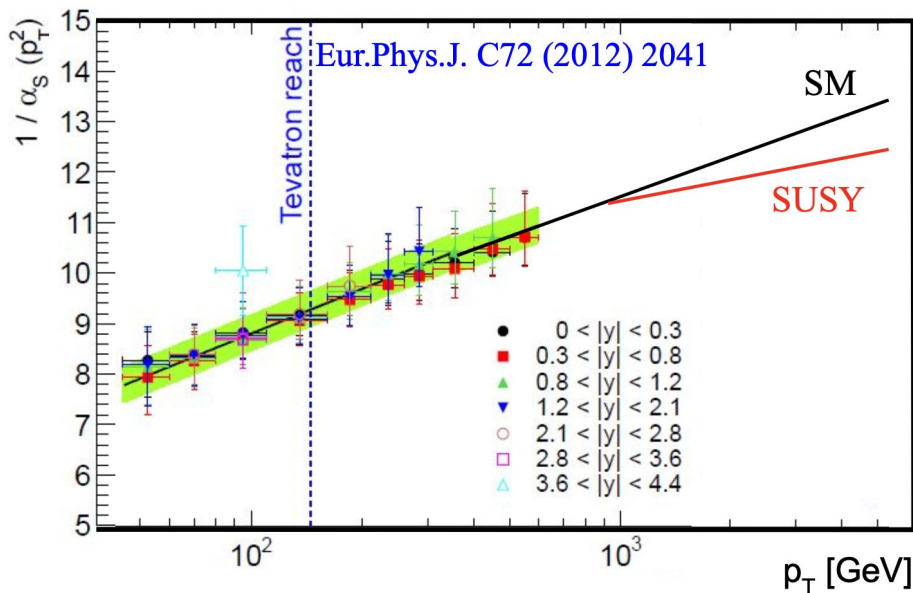
⊗  $k_{EW}$  ⊗  $k_{NP}^{Pythia8 \text{ AU2CT10}}$

$\mu_R = \mu_F = p_{T,jet}^{\max}$

● CT14  
 ■ HERAPDF2.0  
 ▲ NNPDF3.0  
 ▼ MMHT2014

*Backup*

# Extraction of Physics information from measurements



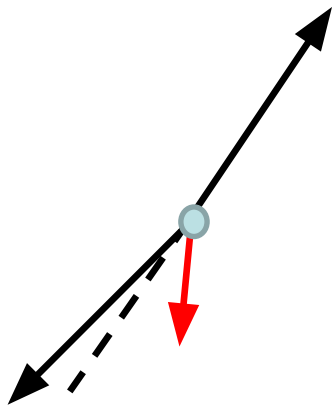
→ Involves using information on uncertainties and their correlations (between various measurement bins), keeping in mind that there are uncertainties impacting them too

*Backup*

# $\alpha_s$ from jet Xsec ratios: energy range for the RGE test and PDF sensitivity

→ Observables like  $R_{3/2}$ ,  $R_{\Delta\phi}$  and (A)TEEC non-trivial due to events that are *not* back-to-back dijets: sensitivity to  $\alpha_s$  originates from probability of emission of extra radiation (3<sup>rd</sup> jet etc.)

→ Relevant scale for RGE test related to  $p_{T,3}$  (low), not to event-level observables (e.g.  $p_T^{\text{lead. jet}}$ ,  $p_T^{\text{(all jets)}}$ ,  $(p_{T,1}+p_{T,2})/2$ ,  $H_T/2$ )



Observable [Ref.]	$\alpha_s(M_Z^2)$	Range PDF variations
$R_{32}$ [109]	$0.111 \pm 0.006$ (exp) $^{+0.016}_{-0.003}$ (PDF, NP, scale)	0.109 – 0.116
$R_{32}$ [110]	$0.1148 \pm 0.0014$ (exp) $\pm 0.0018$ (PDF) $\pm 0.0050$ (theory)	0.1135 – 0.1148
3-jet mass [108]	$0.1171 \pm 0.0013$ (exp) $\pm 0.0024$ (PDF) $\pm 0.0008$ (NP) $^{+0.0069}_{-0.0040}$ (scale)	0.1143 – 0.1183
2-jets [111]	$0.1159 \pm 0.0025$ (exp, PDF, NP)	0.1159 – 0.1183
3-jets [111]	$0.1161 \pm 0.0021$ (exp, PDF, NP)	0.1159 – 0.1179
2- & 3-jets [111]	$0.1161 \pm 0.0021$ (exp, PDF, NP)	0.1161 – 0.1188
$R_{32}$ [111]	$0.1150 \pm 0.0010$ (exp) $\pm 0.0013$ (PDF) $\pm 0.0015$ (NP) $^{+0.0050}_{-0.0000}$ (scale)	0.1139 – 0.1184
TEEC [112]	$0.1162 \pm 0.0011$ (exp) $\pm 0.0018$ (PDF) $\pm 0.0003$ (NP) $^{+0.0076}_{-0.0061}$ (scale)	0.1151 – 0.1177
ATEEC [112]	$0.1196 \pm 0.0013$ (exp) $\pm 0.0017$ (PDF) $\pm 0.0004$ (NP) $^{+0.0061}_{-0.0013}$ (scale)	0.1185 – 0.1206
$R_{\Delta\phi}$ [113]	$0.1127^{+0.0019}_{-0.0018}$ (exp) $\pm 0.0006$ (PDF) $^{+0.0003}_{-0.0001}$ (NP) $^{+0.0052}_{-0.0019}$ (scale)	0.1127 – 0.1156

→ PDF uncertainties non-negligible for ( $\alpha_s$  from) cross-section ratio measurements & (A)TEEC:

- probability of extra radiation sensitive to the type of partons in the initial state
- both  $\alpha_s$  & PDF sensitivities of the observables reduced for ratios: both relevant for the  $\alpha_s$  evaluation

Backup

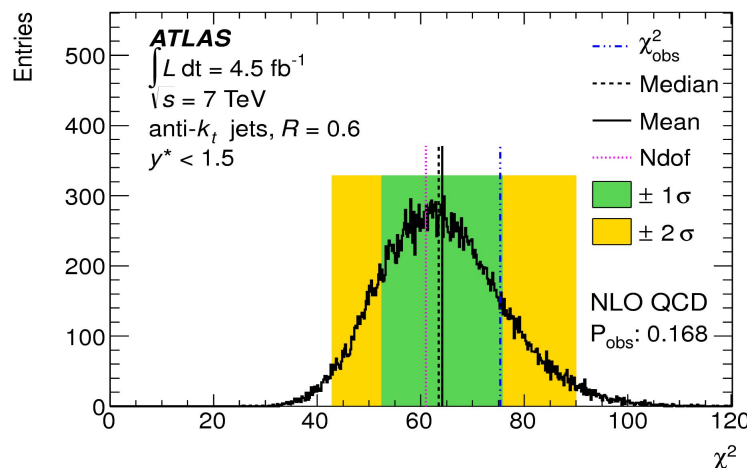
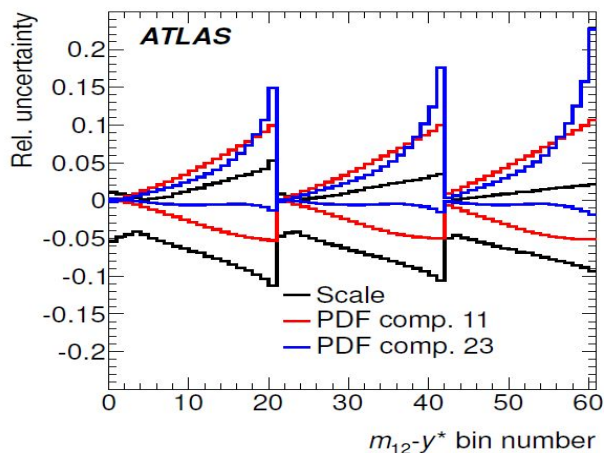


# Quantitative data/theory comparisons

→ Generalized

$$\chi^2(\mathbf{d}; \mathbf{t}) = \min_{\beta_a} \left\{ \sum_{i,j} \left[ d_i - \left( 1 + \sum_a \beta_a \cdot (\epsilon_a^\pm(\beta_a))_i \right) t_i \right] \cdot [C_{\text{su}}^{-1}(\mathbf{t})]_{ij} \right. \\ \left. \cdot \left[ d_j - \left( 1 + \sum_a \beta_a \cdot (\epsilon_a^\pm(\beta_a))_j \right) t_j \right] + \sum_a \beta_a^2 \right\},$$

Accounts for correlations and asymmetries of experimental and theoretical uncertainties (stat. & syst.)



→ Using frequentist method to compute p-value:

- pseudo-experiments from theory prediction, with the full information on the uncertainties:  
build the generalized  $\chi^2$  distribution (no assumption needed)
- observed  $\chi^2$  from the data/theory comparison

# Quantitative comparison between data and NLO QCD+NP+EW

*Comparisons performed for a large number of configurations:*

→ PDFs: ABM11(as for 7TeV), CT14, MMHT 2014, NNPDF 3.0, HERAPDF 2.0, ABMP16

→ Phase-space regions:

$p_T$  ranges:

- “wide”:  $> 70$ ;  $> 100$ ;  $100 - 900$ ;  $100 - 400$  GeV

- “narrow”:  $70 - 100$ ;  $100 - 240$ ;  $240 - 408$ ;  $408 - 642$ ;  $642 - 952$ ;  $> 952$  GeV

$|y|$  ranges:

- “individual bins”:  $|y| < 0.5$ ;  $0.5 - 1$ ;  $1 - 1.5$ ;  $1.5 - 2$ ;  $2 - 2.5$ ;  $2.5 - 3$

- “full range”:  $|y| < 3$

- “pairs of consecutive bins”:  $|y| < 1$ ;  $0.5 - 1.5$ ;  $1 - 2$ ;  $1.5 - 2.5$ ;  $2 - 3$

- “central-forward pairs”:  $|y| < 0.5$  &  $2.5 - 3$ ;  $< 0.5$  &  $2 - 2.5$ ;  $< 0.5$  &  $1.5 - 2$

→  $R=0.4$  and  $R=0.6$ ;  $p_T^{\text{leading jet}}$  and  $p_T^{\text{jet}}$  scale choices

→ Generally good agreement for inclusive jets for individual & pairs of  $|y|$  bins

→ Tension when including all  $|y|$  bins for inclusive jets

→ Sensitive to treatment of correlations for “2-point” systematic uncertainties

→ Good data/theory agreement for dijets

Backup

# Testing realistic alternative correlation assumptions

*Inclusive jets - nominal  $\chi^2/ndf$  for CT14 with  $p_T^{\text{leading jet}}$  scale:  
321 – 360/159 (8 TeV); 419/177(13 TeV)*

*Splitting a single systematic: some  $\chi^2$  reduction, but still small p-values.*

*Splitting simultaneously several uncertainties:*

*→ JES Flavour Response, JES MJB Fragmentation, JES Pile-up Rho Topology:  
 $\chi^2$  reduction by up to 51 units (8 TeV)*

*→ Scale variations, alternative scale choice, non-perturbative correction:  
 $\chi^2$  reduction by up to 87 units (8 TeV)  
– more work needed on the correlations of theory uncertainties*

*→ Splitting both the experimental and theoretical uncertainties:  
 $\chi^2$  reduction by up to 96 units (8 TeV); 58 units (13 TeV)*

*→ Possible (extra) motivation for including scale uncertainties in PDF fits - in progress*

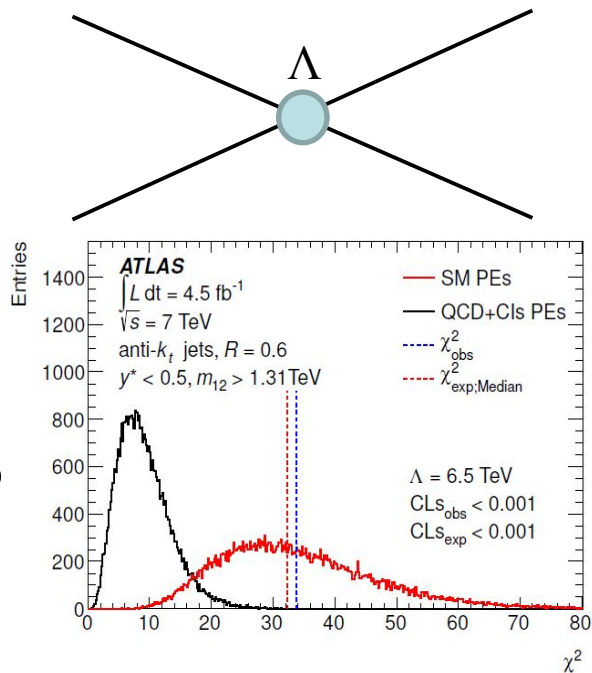
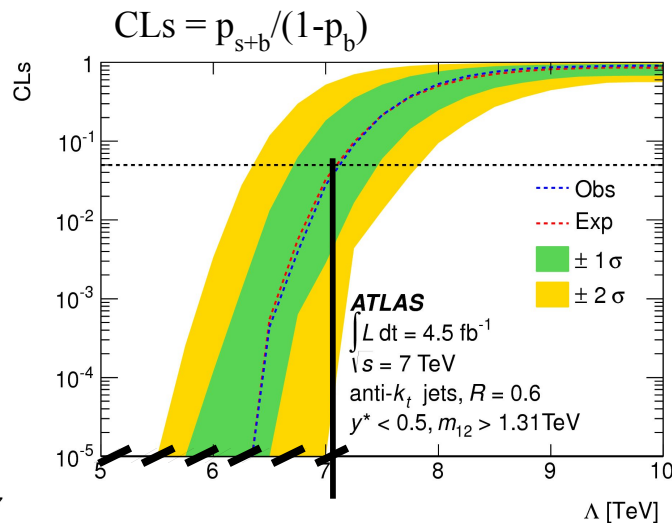
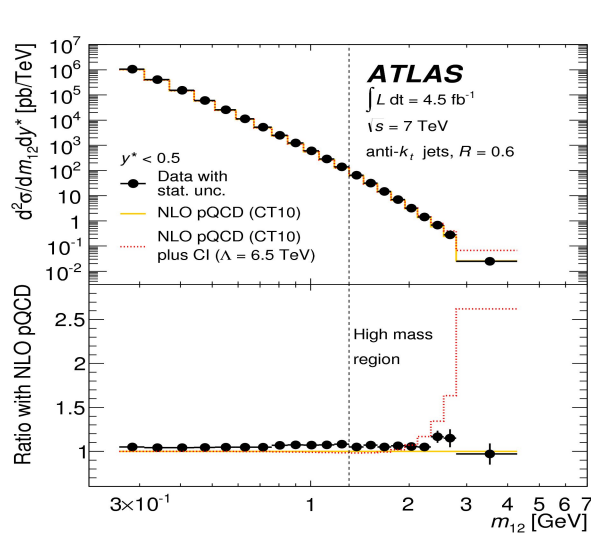
*Note: there is also an uncertainty on the phase-space dependence for the size of 2-point systematics  
→ may explain part of the observed tension*

# Limits on New Physics using unfolded distributions

→ Explore BSM physics directly at particle level

Contact Interaction Model (CI)

New force mediated by heavy particle



→ Full frequentist analysis (CLs), with generalized  $\chi^2$  as test statistic

→ Accounts for correlations and asymmetries of uncertainties (stat. & syst.)

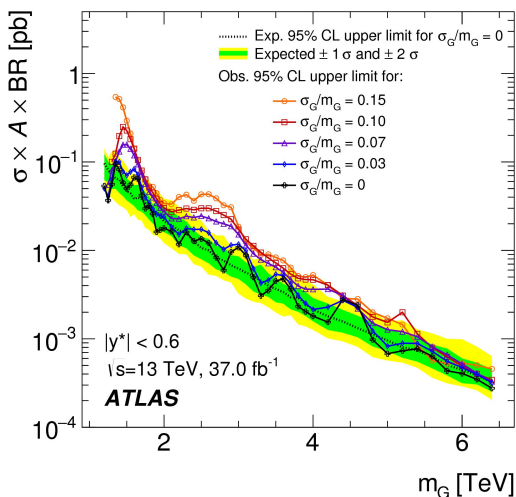
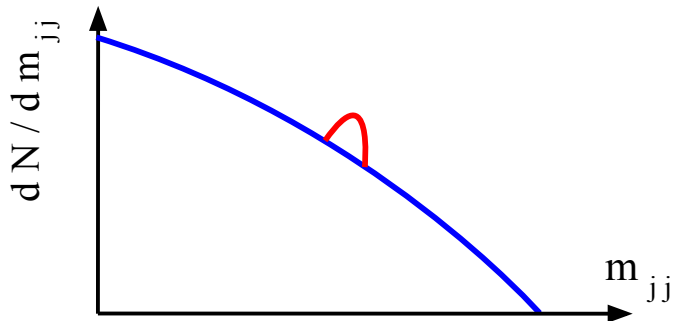
→ Limits similar to the ones obtained by dedicated searches

(comparing reconstructed-level data with theory predictions folded with detector effects)

Backup

# Generic Gaussian signals: folding-based method

- Limits on generic Gaussian signals can be re-interpreted in terms of various signal models
- Previously studied at reconstructed-level – hadron-level preferable (limits more straightforward to use)
- Folding method using MC-based transfer matrix allows to factorize physics & detector effects



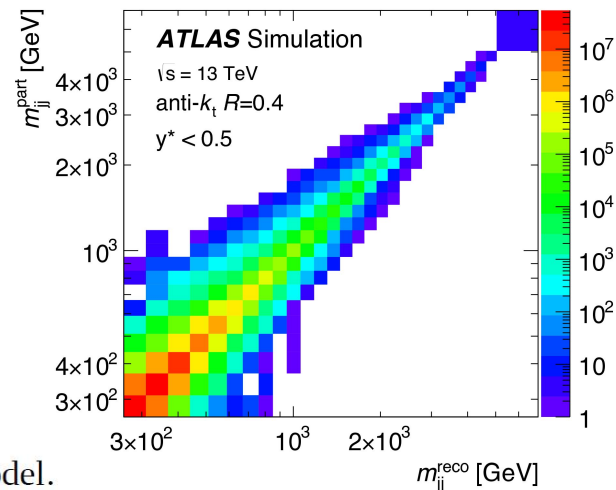
(Robert Hankache - PhD)

$f_y(M^{truth})$  = truth entries for a given model.

$F'_x(M^{reco})$  = the expected reco entries.

$$f_y(M^{truth}) \xrightarrow{\text{Folding}} F'_x(M^{reco}) = \sum_y f_y * \underbrace{E_y^T * A_{xy} / E_x^R}_{\tilde{A}}$$

→ For resonance width ~ resolution: differences between folding result and reconstructed-level limits of up to 20% (different interpretation)



Backup

# Conclusion

- *Interesting QCD / New Physics-related questions to address both at the energy frontier, with jets, and in precision low-energy studies employing hadronic spectra*
- *Developed several methodologies relevant for both areas*
- *Potential for multiple improvements of such measurements and their phenomenological interpretation*

*Thank you !!!*

# Backup

# The $(g-2)_\mu$ : definition & experimental measurement

- Magnetic dipole moment of a charged lepton:  $\vec{\mu} = g \frac{e}{2m} \vec{s}$
- “anomaly” = deviation w.r.t. Dirac’s prediction:  $a = \frac{g-2}{2}$

- Experimental “ingredients” to measure  $a_\mu$ :

→ Polarised muons from pion decays (parity violation)

→ “Anomalous frequency”

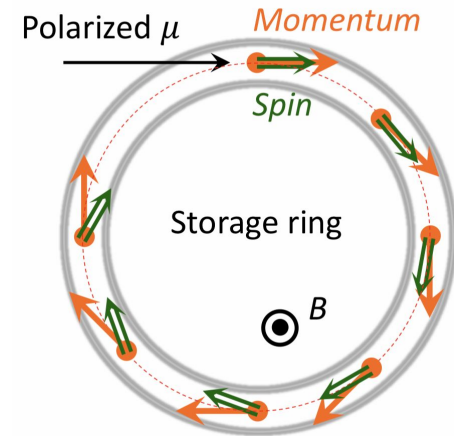
(difference between spin precession and cyclotron frequency)

proportional to  $a_\mu$  for the “magic  $\gamma$ ”

$$\vec{\omega}_a = \frac{e}{m_\mu c} \left[ a_\mu \vec{B} - \left( a_\mu - \frac{1}{\gamma^2 - 1} \right) \vec{\beta} \times \vec{E} \right] \approx \frac{e}{m_\mu c} a_\mu \vec{B}$$

→ Parity violation in muon decays

(electron emitted in the direction opposite to the muon spin)

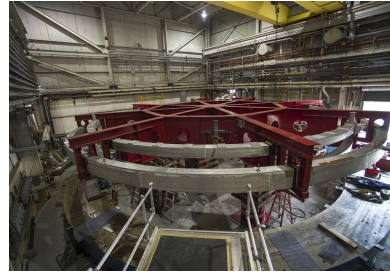
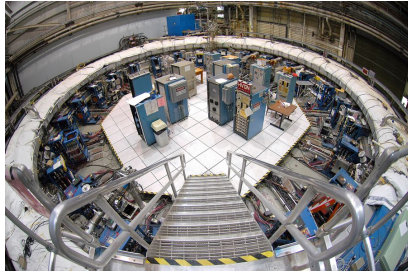


$$\mu^-_{\text{polarised}} \rightarrow e^- + \bar{\nu}_e + \nu_\mu$$



# From BNL to Fermilab

BNL → 1 month long trip for the g-2 storage ring



→ Fermilab  
July 26, 2013

*This is NOT an UFO !!!*



$$a_{\mu}^{\text{Exp}}(\text{BNL}): (11\,659\,208.9 \pm 6.3) \cdot 10^{-10}$$

$$a_{\mu}^{\text{Exp}}(\text{Fermilab runs 1-3 + BNL}): (11\,659\,205.9 \pm 2.2) \cdot 10^{-10} \quad (0.19 \text{ ppm}) \rightarrow \text{One of the most precise quantities ever measured}$$

- Expectation for final publication: another factor 2 improvement for the statistical precision

# The $(g-2)_\mu$ experiment

$$a_\mu^{\text{Exp}}(\text{BNL}): (11\,659\,208.9 \pm 6.3) \cdot 10^{-10}$$

→ Expected uncertainty reduction by a factor 4 with the experiment at Fermilab

- improved apparatus and enhanced statistics: more intense (x20) and pure muon beam; B-field mapped every 3 days with special trolley with probes pulled through beampipe (homogeneity ~ ppm); tracking system for electron detectors etc.

- 1st publication: similar precision & good agreement with BNL (7th of April 2021) PRL 126, 141801 (2021)

$$a_\mu^{\text{Exp}}(\text{Fermilab}): (11\,659\,204.0 \pm 5.1 \pm 1.8) \cdot 10^{-10} \rightarrow 6\% \text{ of total data}$$

$$a_\mu^{\text{Exp}}(\text{Fermilab} + \text{BNL}): (11\,659\,206.1 \pm 4.1) \cdot 10^{-10} \text{ (0.35 ppm)}$$

- 2nd publication: uncertainty reduction by a factor ~2 (10th of August 2023) PRL 131, 161802 (2023) (+Run 2 & 3 data)

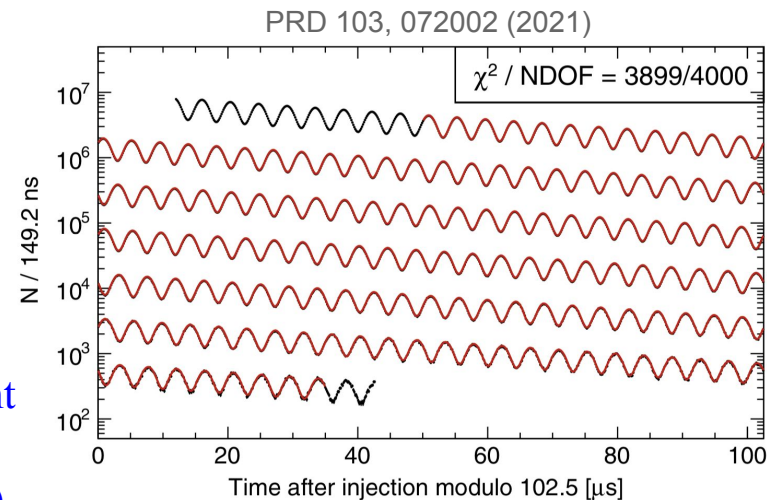
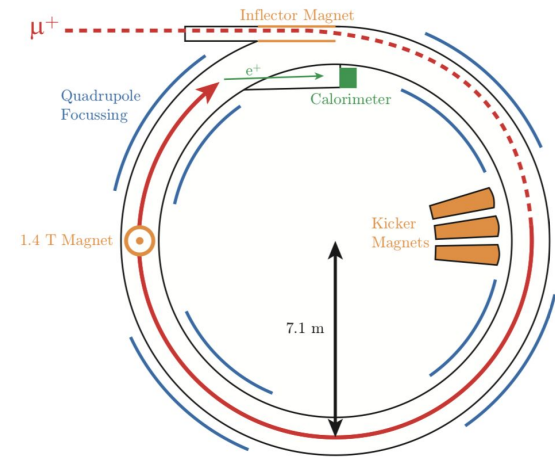
$$a_\mu^{\text{Exp}}(\text{Fermilab}): (11\,659\,205.5 \pm 2.4) \cdot 10^{-10} \text{ (0.20 ppm)}$$

$$a_\mu^{\text{Exp}}(\text{Fermilab} + \text{BNL}): (11\,659\,205.9 \pm 2.2) \cdot 10^{-10} \text{ (0.19 ppm)}$$

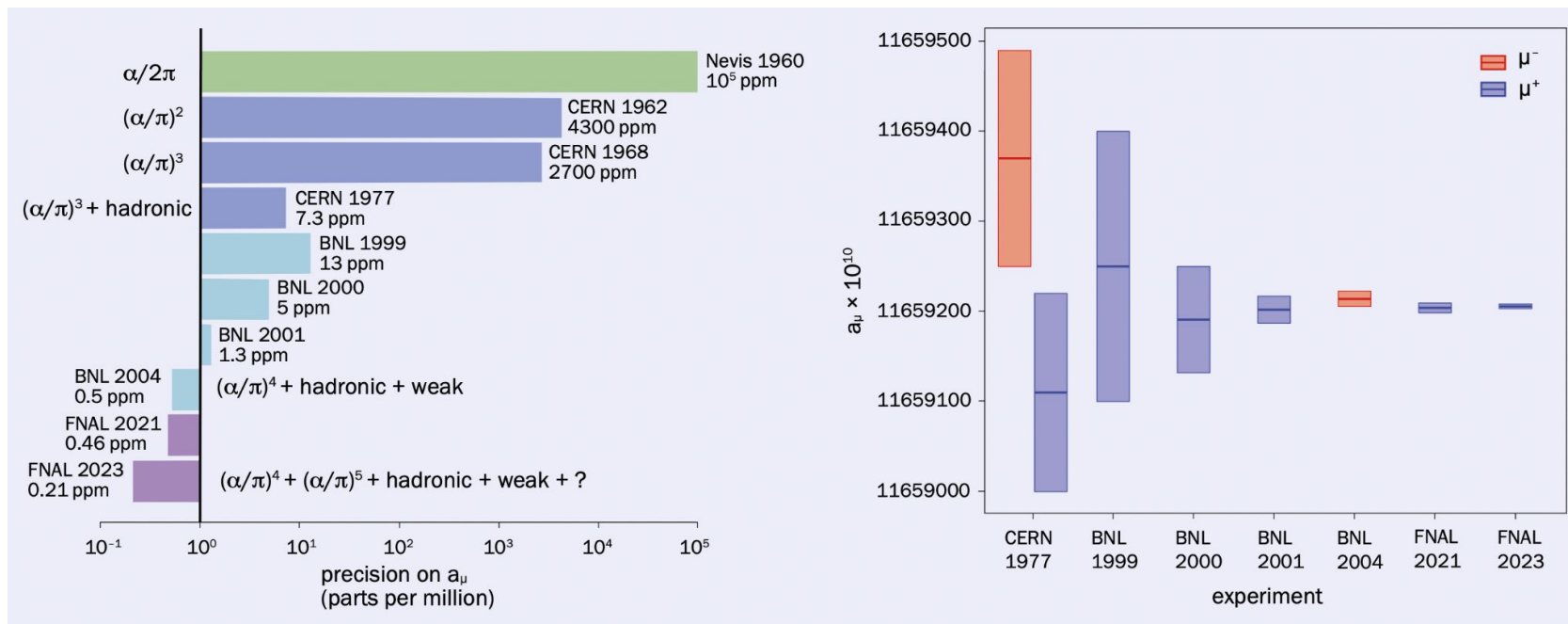
→ One of the most precise quantities ever measured

- Expectation for final publication: another factor 2 improvement for the statistical precision

→ Initiative for a measurement using slow muons (KEK, Japan)



# Precision of the $(g-2)_\mu$ experiments



[CERN Courier March-April '25](#)

# Lepton Magnetic Anomaly: from Dirac to QED

- Magnetic dipole moment of a charged lepton:  $\vec{\mu} = g \frac{e}{2m} \vec{s}$   
Dirac (1928)  $g_e=2$   $a_e=0$
- “anomaly” = deviation w.r.t. Dirac’s prediction:  $a = \frac{g-2}{2}$

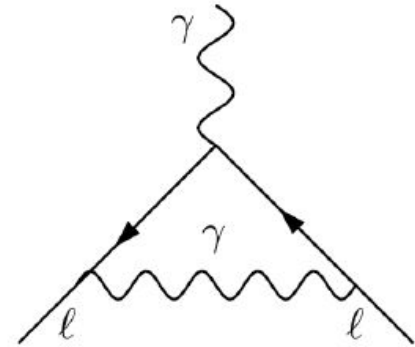
anomaly discovered:

Kusch-Foley (1948)  $a_e = (1.19 \pm 0.05) 10^{-3}$

and explained by  $O(\alpha)$  QED contribution:

Schwinger (1948)  $a_e = \alpha/2\pi = 1.16 10^{-3}$

first triumph of QED



⇒  $a_e$  sensitive to quantum fluctuations of fields

# More Quantum Fluctuations

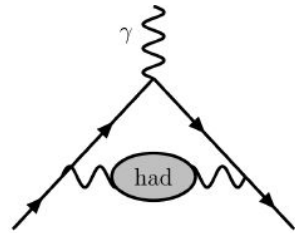
$$a = a^{\text{QED}} + a^{\text{had}} + a^{\text{weak}} + ? a^{\text{new physics}} ?$$

typical contributions:

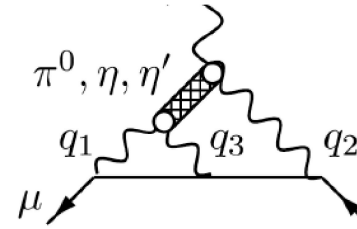
QED up to  $O(\alpha^5)$  (Kinoshita et al.)



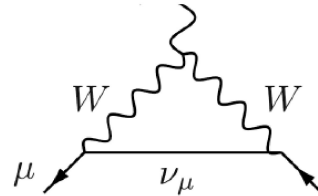
Hadrons vacuum polarization



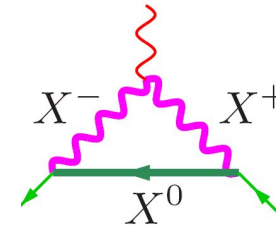
light-by-light (dispersive & lattice QCD)



Electroweak



new physics at high mass scale



$$\delta a_l \propto \frac{m_l^2}{M^2} \Rightarrow a_\mu \text{ much more sensitive to high scales}$$

# Theory initiative white paper executive summary & new results

Contribution	Section	Equation	Value $\times 10^{11}$	References
Experiment (E821)		Eq. (8.13)	116 592 089(63)	Ref. [1]
HVP LO ( $e^+e^-$ )	Sec. 2.3.7	Eq. (2.33)	6931(40)	Refs. [2–7]
HVP NLO ( $e^+e^-$ )	Sec. 2.3.8	Eq. (2.34)	−98.3(7)	Ref. [7]
HVP NNLO ( $e^+e^-$ )	Sec. 2.3.8	Eq. (2.35)	12.4(1)	Ref. [8]
HVP LO (lattice, $udsc$ )	Sec. 3.5.1	Eq. (3.49)	7116(184)	Refs. [9–17]
HLbL (phenomenology)	Sec. 4.9.4	Eq. (4.92)	92(19)	Refs. [18–30]
HLbL NLO (phenomenology)	Sec. 4.8	Eq. (4.91)	2(1)	Ref. [31]
HLbL (lattice, $uds$ )	Sec. 5.7	Eq. (5.49)	79(35)	Ref. [32]
HLbL (phenomenology + lattice)	Sec. 8	Eq. (8.10)	90(17)	Refs. [18–30, 32]
QED	Sec. 6.5	Eq. (6.30)	116 584 718.931(104)	Refs. [33, 34]
Electroweak	Sec. 7.4	Eq. (7.16)	153.6(1.0)	Refs. [35, 36]
HVP ( $e^+e^-$ , LO + NLO + NNLO)	Sec. 8	Eq. (8.5)	6845(40)	Refs. [2–8]
HLbL (phenomenology + lattice + NLO)	Sec. 8	Eq. (8.11)	92(18)	Refs. [18–32]
Total SM Value	Sec. 8	Eq. (8.12)	116 591 810(43)	Refs. [2–8, 18–24, 31–36]
Difference: $\Delta a_\mu := a_\mu^{\text{exp}} - a_\mu^{\text{SM}}$	Sec. 8	Eq. (8.14)	279(76)	

→ Dominant uncertainty: HVP LO → Merging of model independent results: DHMZ and KNT (and CHHKs for  $\pi^+\pi^-$  &  $\pi^+\pi^-\pi^0$ ) Central value from simple average; BABAR-KLOE tension & correlations between channels from DHMZ; Max(DHMZ & KNT uncertainties) in each channel

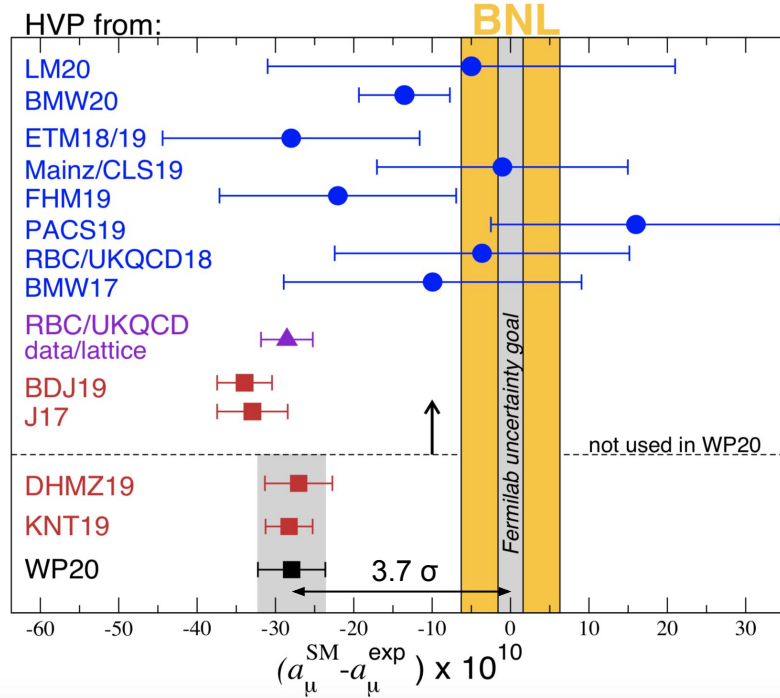
→ HLbL also has an important uncertainty

→ Lattice QCD (+QED) results become more and more interesting; Precision of BMW20 (to be cross-checked by other lattice groups) became similar to the one of dispersive approaches; Good agreement using Euclidean time windows (related to HVP with suppression of very low and high energies) for which various groups achieved similar precision; If BMW20 result is fully confirmed, the difference w.r.t. dispersive results to be understood.

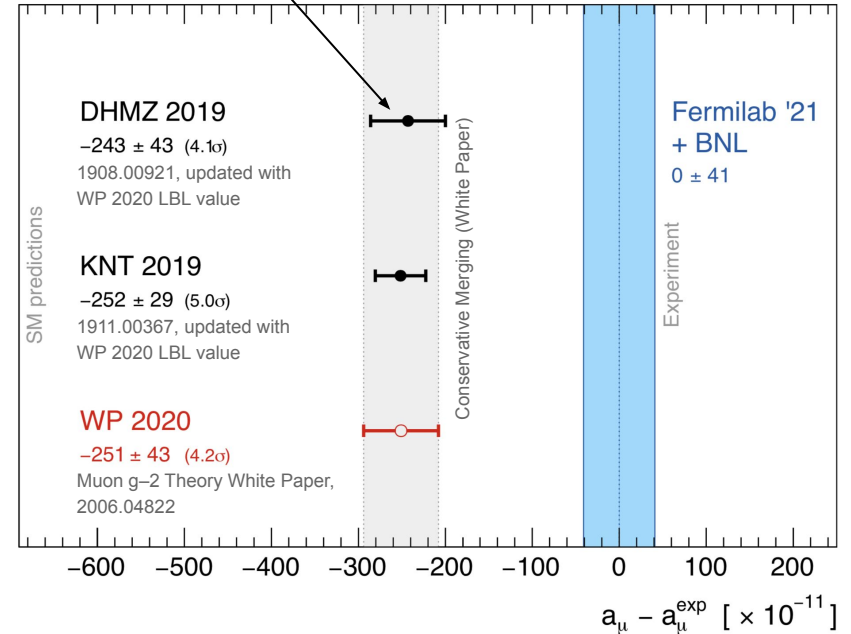
→ A tension between the BNL measurement and the reference SM prediction:  $\sim 3.7 \sigma$  ( $\sim 4.2 \sigma$  including FNAL)

→ Tension significantly smaller when using BMW20 for the LO HVP

# Status of $a_\mu$ before/with 1st Fermilab result

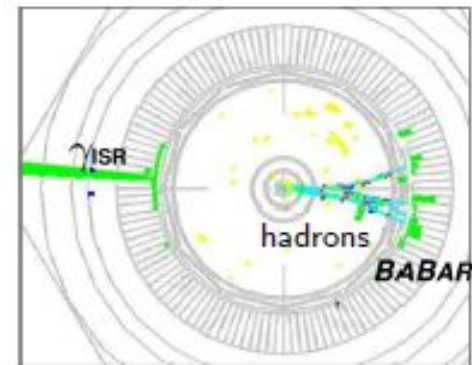
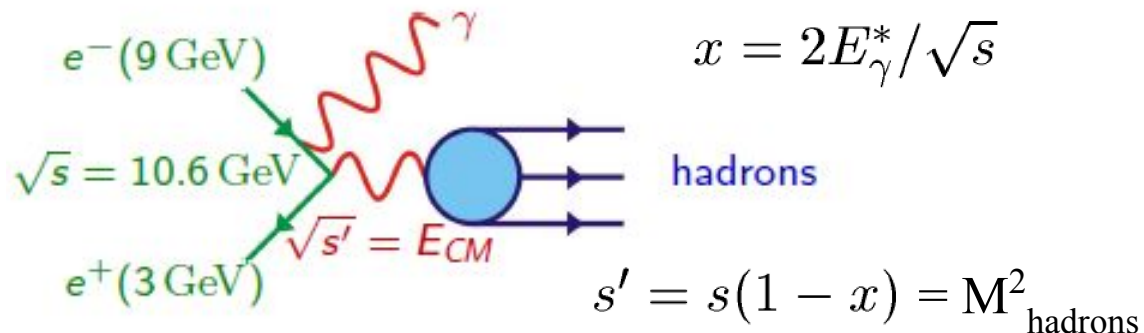


Important to account for BABAR-KLOE diff. & inter-channel correlations



- Caution about significance: statistics-dominated measurement; prediction uncertainty limited by non-Gaussian systematic effects
- Nevertheless, large discrepancy between measurement and reference SM prediction (to be significantly improved in view of the forthcoming updates of the Fermilab measurement)
- Tension significantly smaller when using BMW20 for the LO HVP (TBC by other lattice groups), *not* incompatible with the EW fit (*see below*)

# The ISR method for the $e^+e^- \rightarrow \pi^+\pi^-$ channel at BaBar



- *High energy ISR photon ( $E_\gamma^* > 3 \text{ GeV}$ ) detected at large angle, back-to-back to hadrons*

→ defines  $\sqrt{s'}$  and provides strong background rejection

→ high acceptance, large boost to hadrons (start @ threshold; easier PID)

- *Final state can be hadronic or leptonic (QED)*

→  $\mu^+\mu^-\gamma_{\text{ISR}}(\gamma_{\text{FSR}})$ ,  $\pi^+\pi^-\gamma_{\text{ISR}}(\gamma_{\text{FSR}})$  and  $K^+K^-\gamma_{\text{ISR}}(\gamma_{\text{FSR}})$  measured simultaneously

→  $\mu^+\mu^-\gamma(\gamma)$  used for ISR luminosity: *add. ISR almost cancels for  $\pi\pi\gamma(\gamma)/\mu\mu\gamma(\gamma)$*

- *Kinematic fit including ISR photon (+ additional ISR/FSR)*

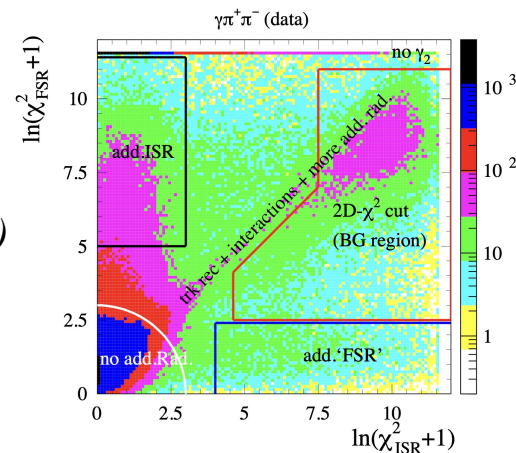
→ removes multihadronic background

→ improves mass resolution (a few MeV)

- *Data/MC corrections for efficiencies and acceptance*

- *Continuous measurement from threshold to 3-5 GeV*

→ *reduced systematic uncertainties* compared to multiple data sets with different colliders and detectors

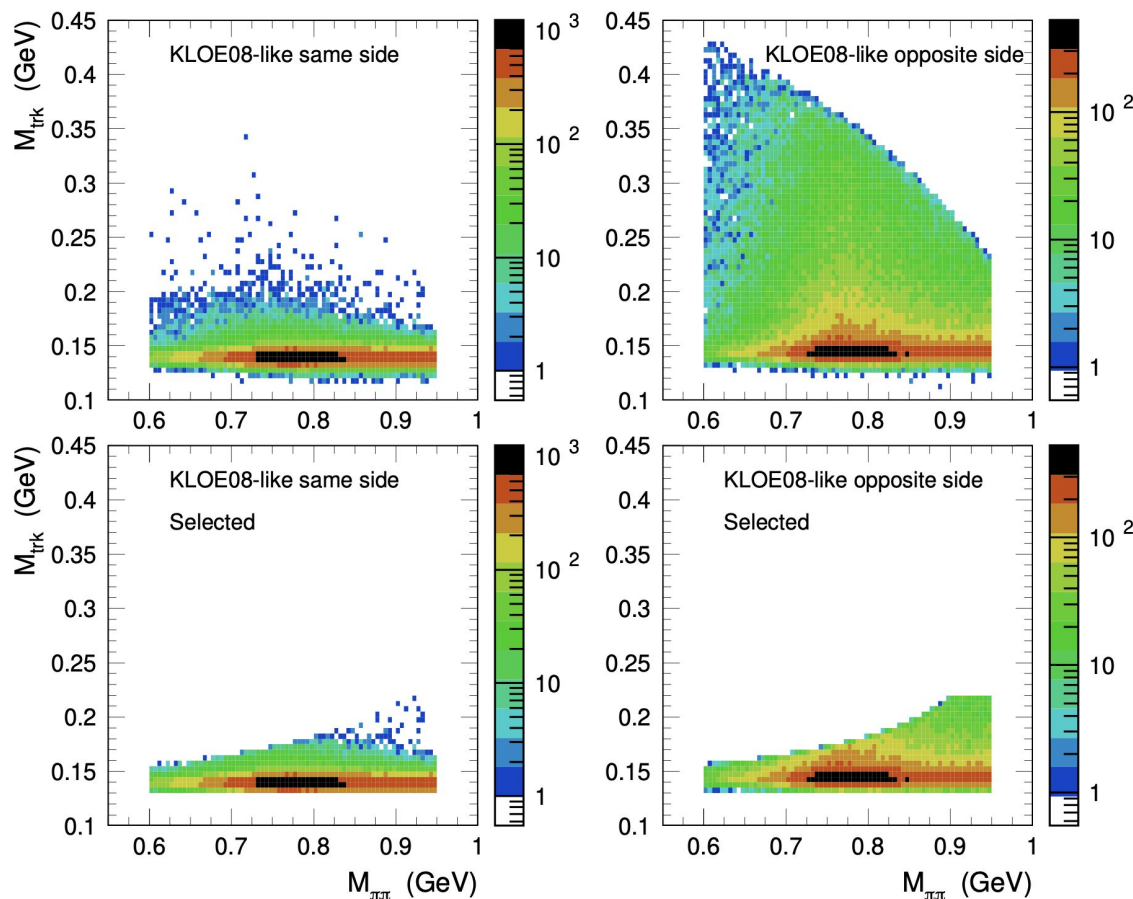


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# Impact of higher order photon emissions studied with fast simulation

[2312.02053](#)



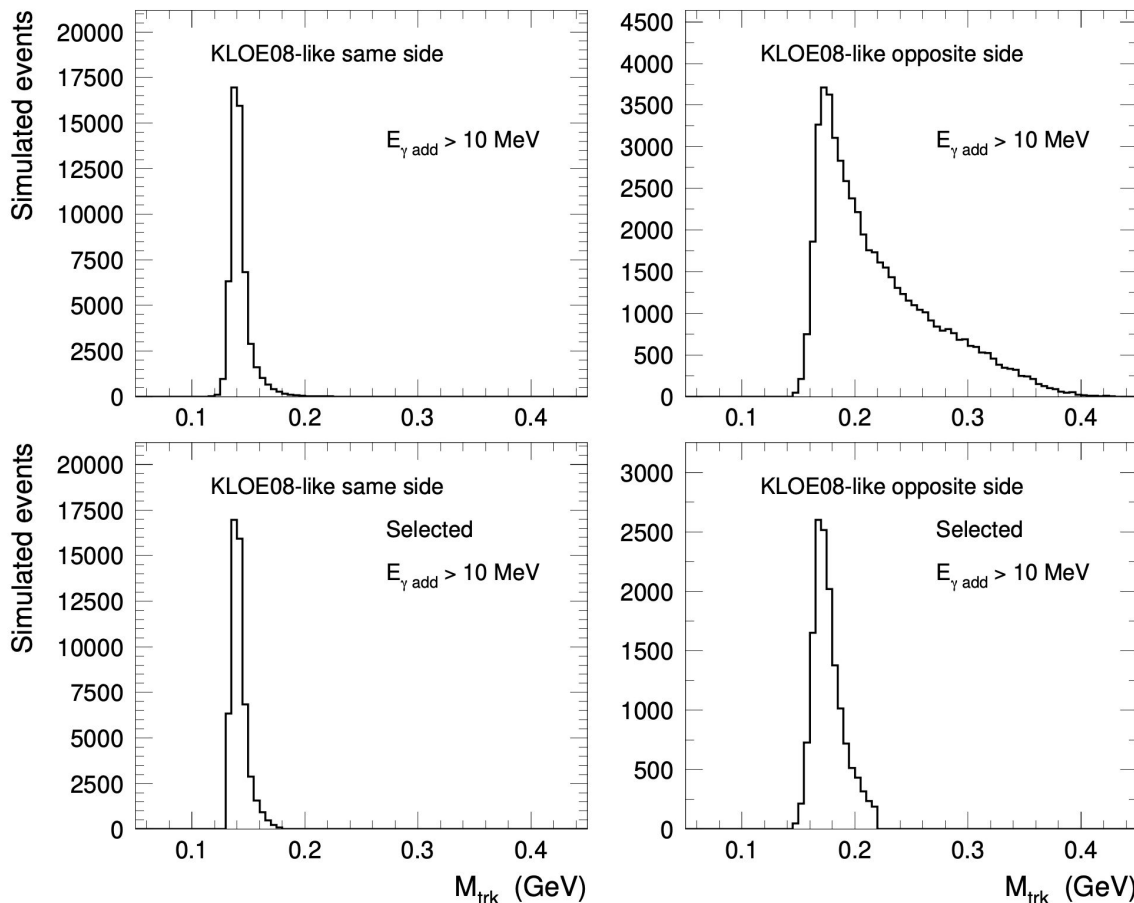
$M_{\text{trk}}$  : common track mass of the two charged particles, computed under the LO assumption

→ Sensitivity of the KLOE selection to additional ‘NLO’ photon emissions ( $E > 5$  MeV) in the same/opposite hemispheres w.r.t. the hard ISR photon

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# Impact of higher order photon emissions studied with fast simulation

[2312.02053](#)

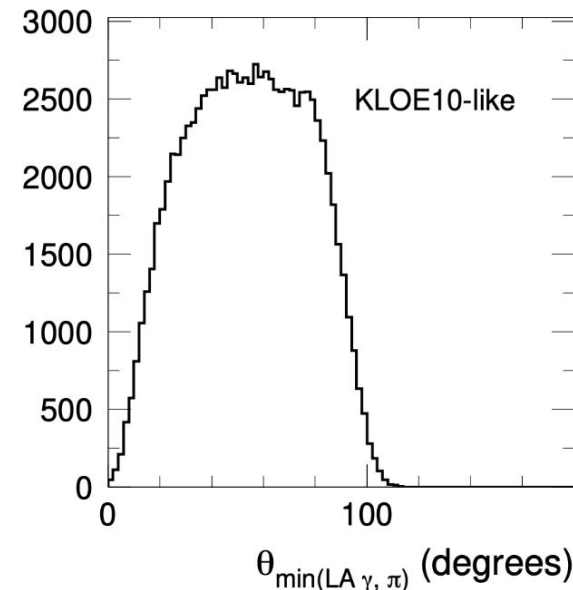
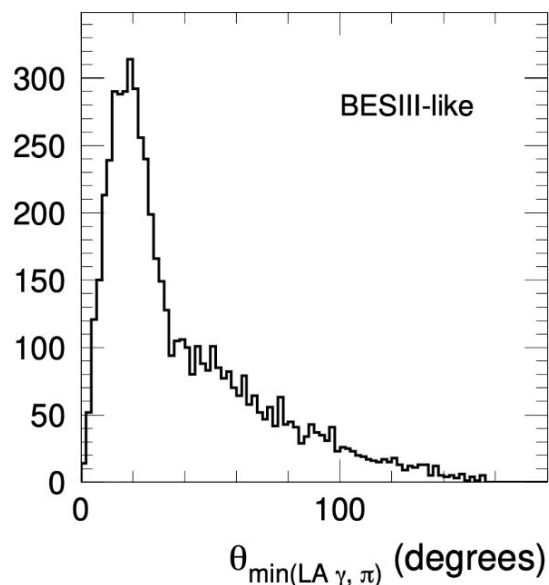
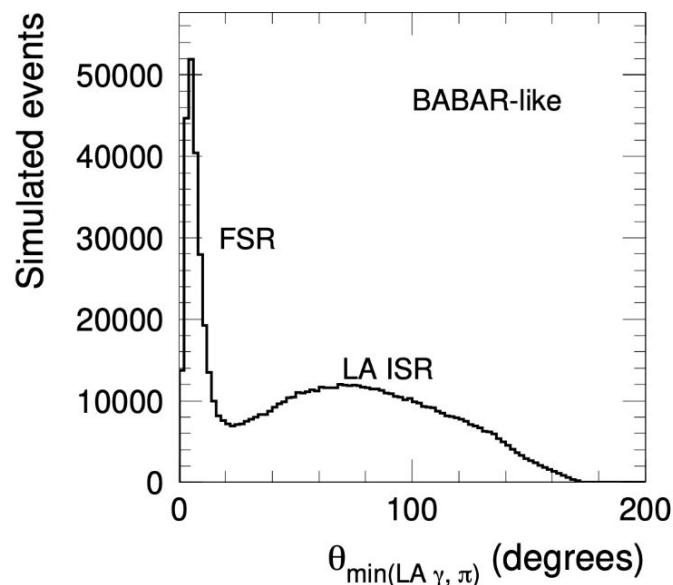


$M_{\text{trk}}$  : common track mass of the two charged particles, computed under the LO assumption

→ Sensitivity of the KLOE selection to additional ‘NLO’ photon emissions ( $E > 10$  MeV) in the same/opposite hemispheres w.r.t. the hard ISR photon

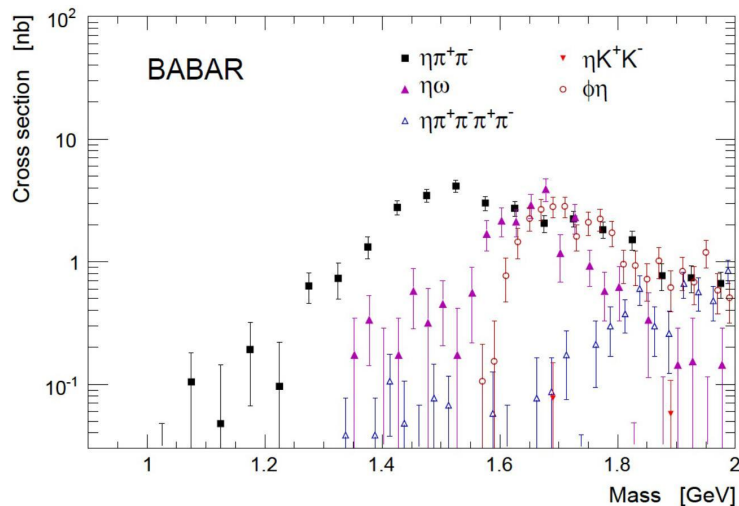
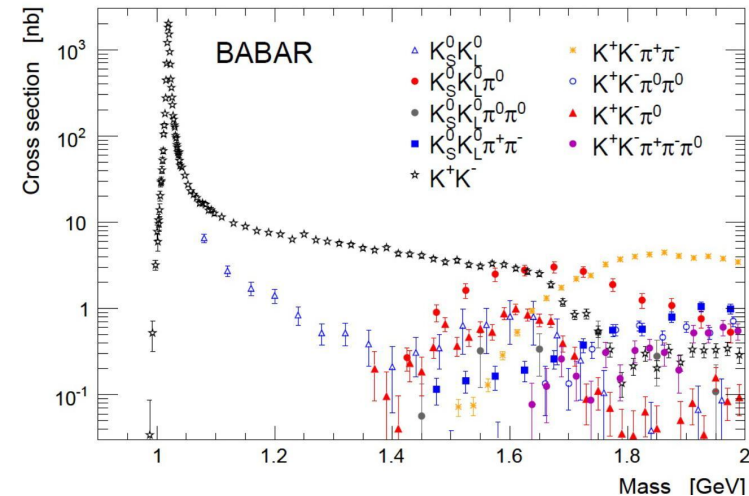
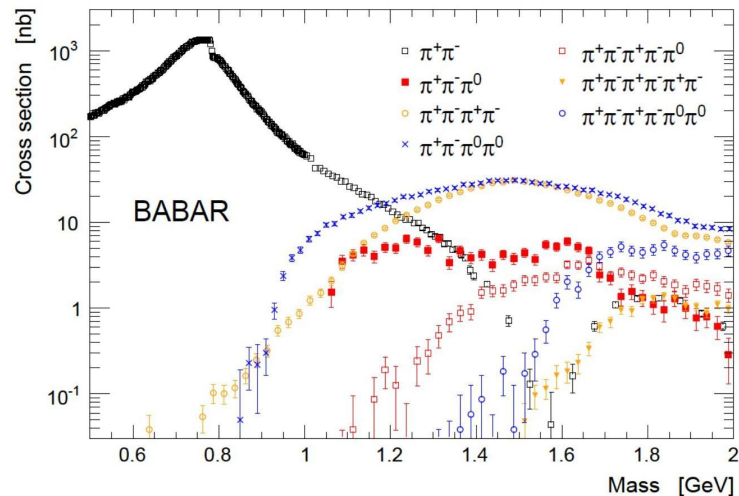
# Impact of higher order photon emissions studied with fast simulation

[2312.02053](#)



→ The angular separation between FSR and LA ISR events is pronounced at high CM energy (BABAR), still visible at intermediate CM energy (BESIII), and vanishes at low CM energy (KLOE)

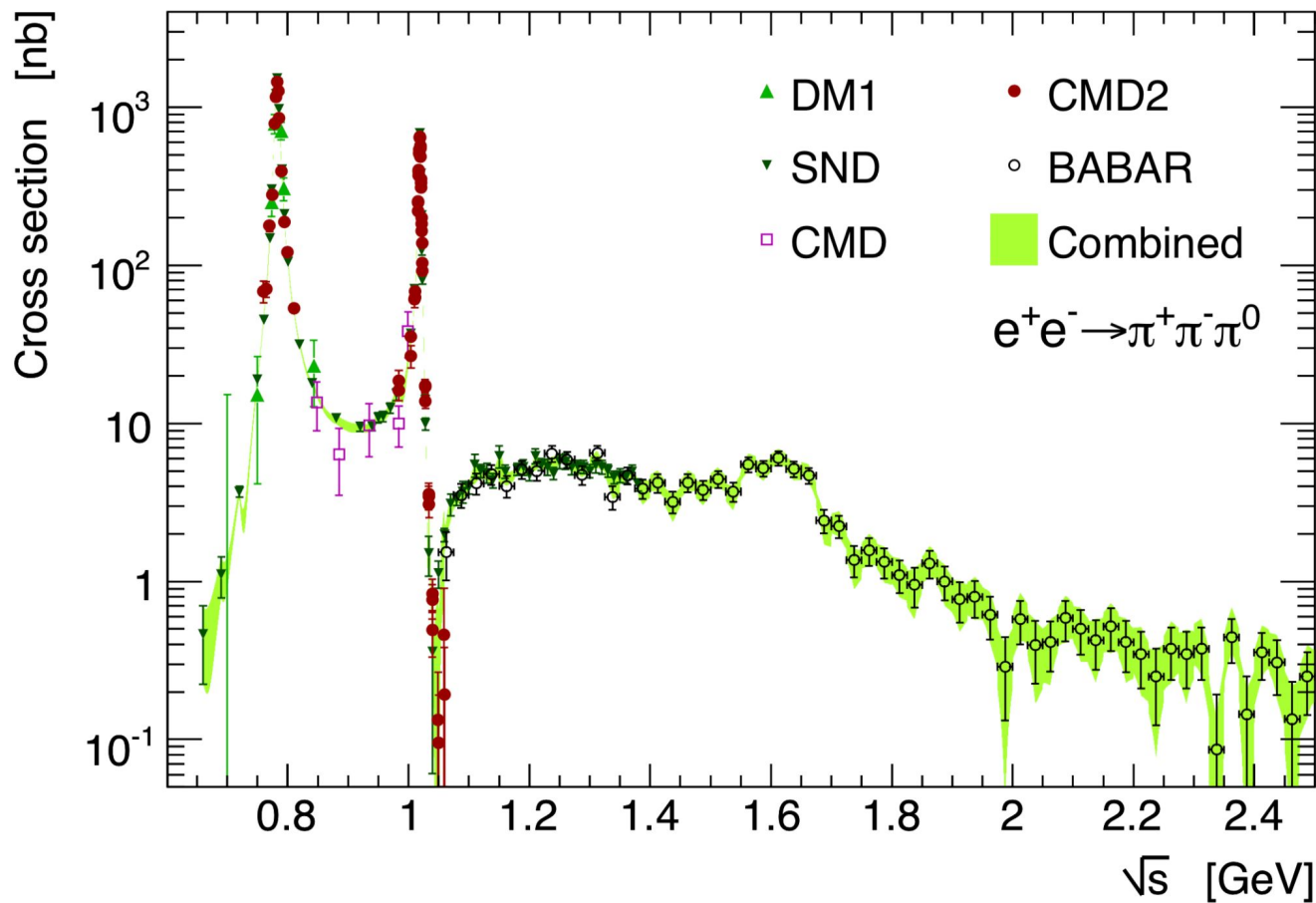
# The BaBar ISR program for differential hadronic Xsec measurements



→ Large effort invested from many groups (Frascati, Mainz, Novosibirsk, Orsay - Paris, ...)

→ Developed *innovative methods* allowing to obtain a large number of *robust and precise measurements*

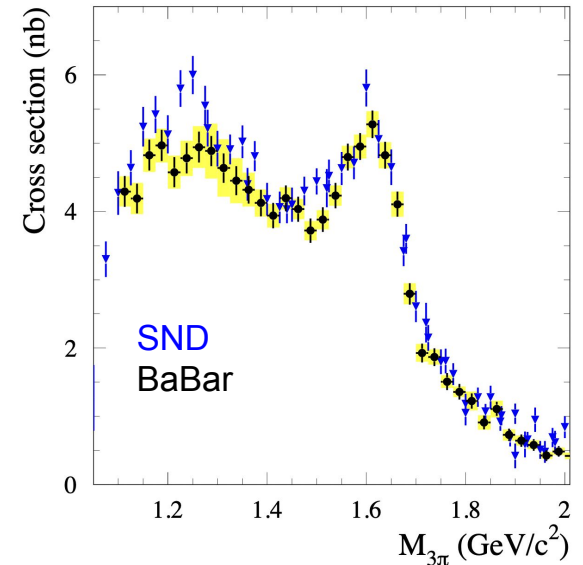
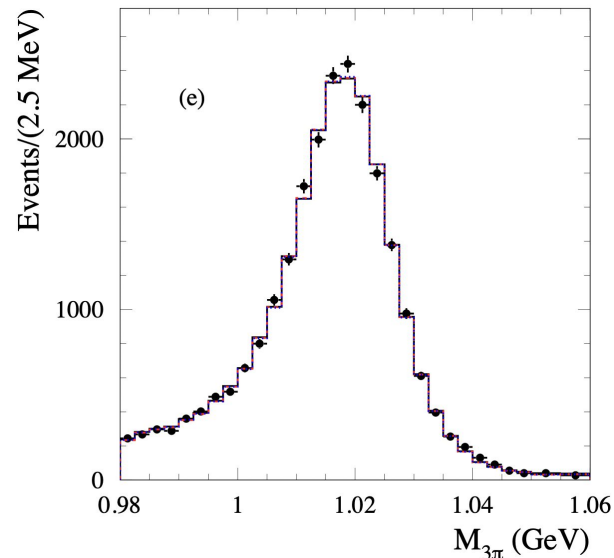
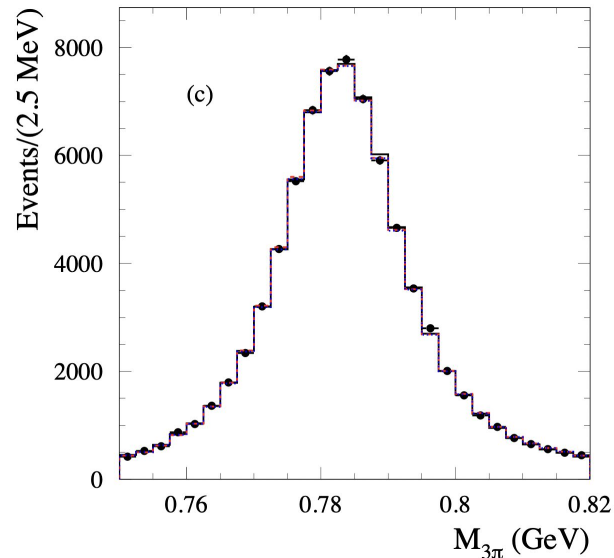
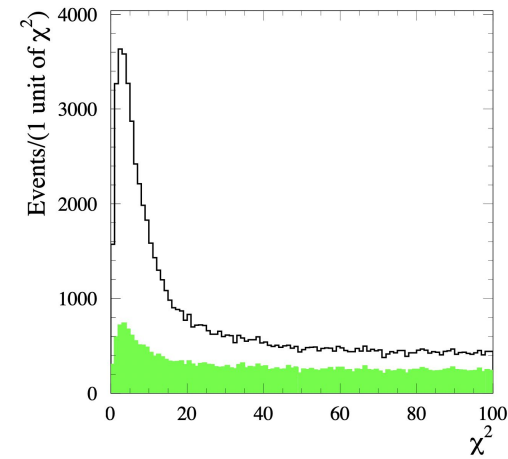
# Combination for the $e^+e^- \rightarrow \pi^+\pi^-\pi^0$ channel



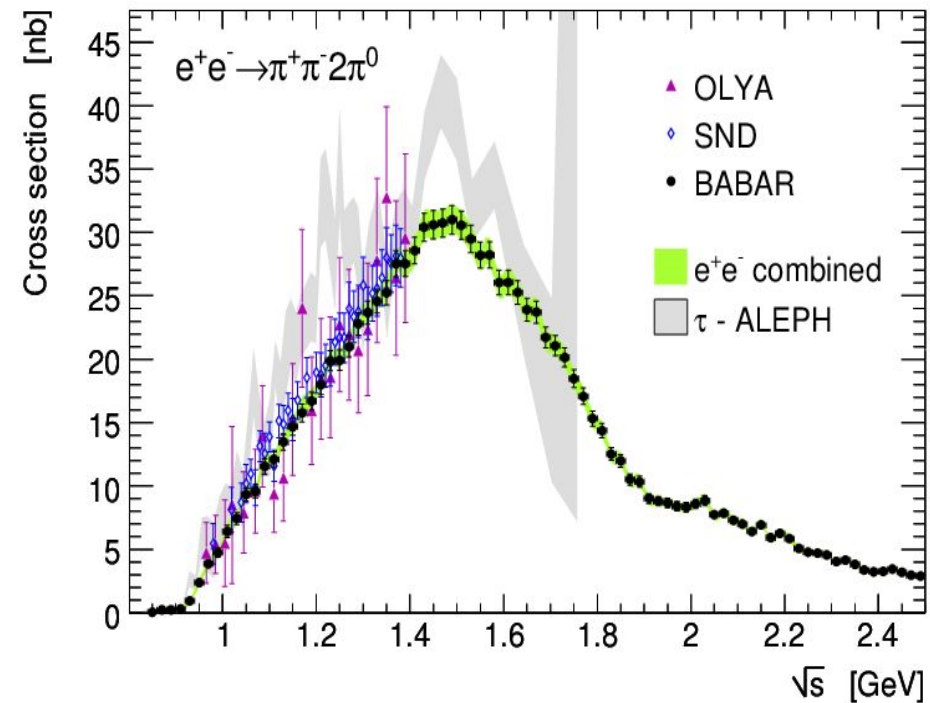
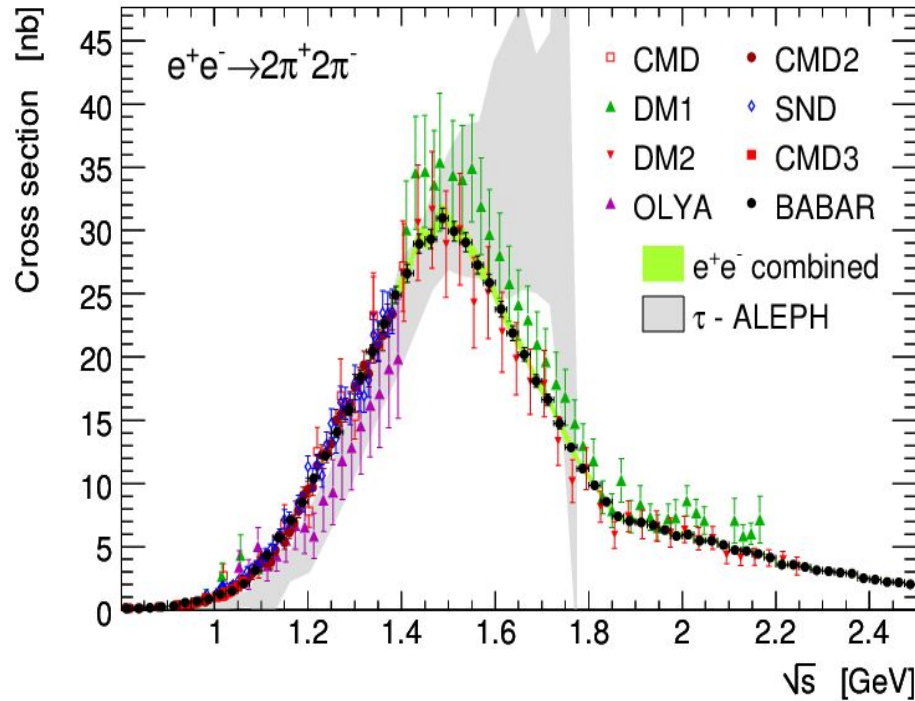
# New data for the $e^+e^- \rightarrow \pi^+\pi^-\pi^0$ channel

[2110.00520](#)

- LO kinematic fits for event selection and  $\sqrt{s}$ ' reconstruction
- Unfolded measurement (0.62-3.5 GeV) to correct for detector resolution
- ISR lumi' derived from total luminosity base on Bhabha (and di-muon) events

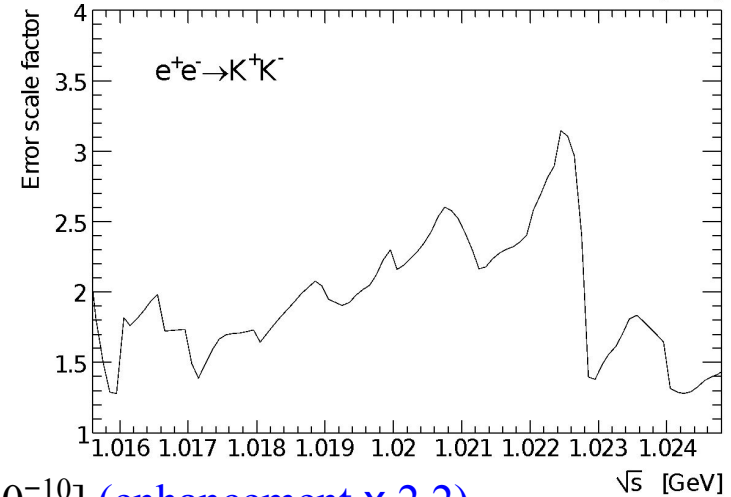
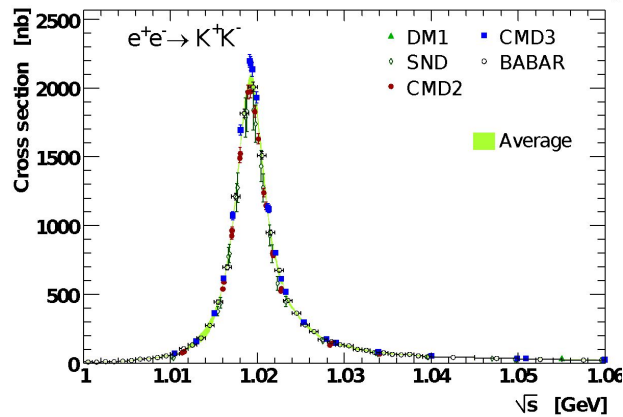
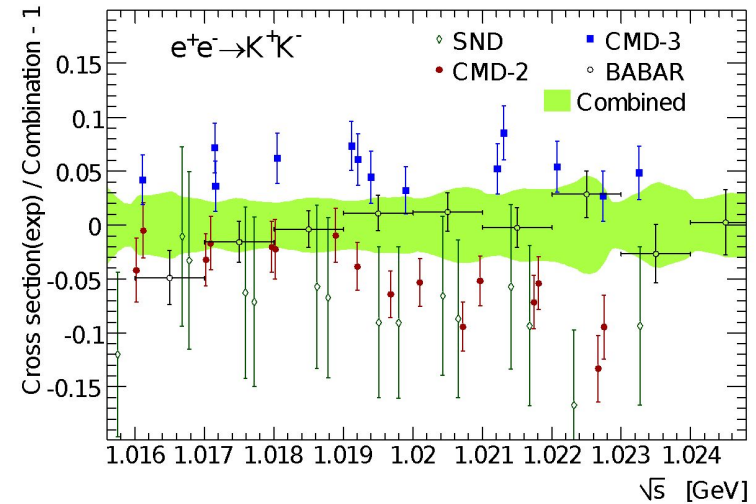
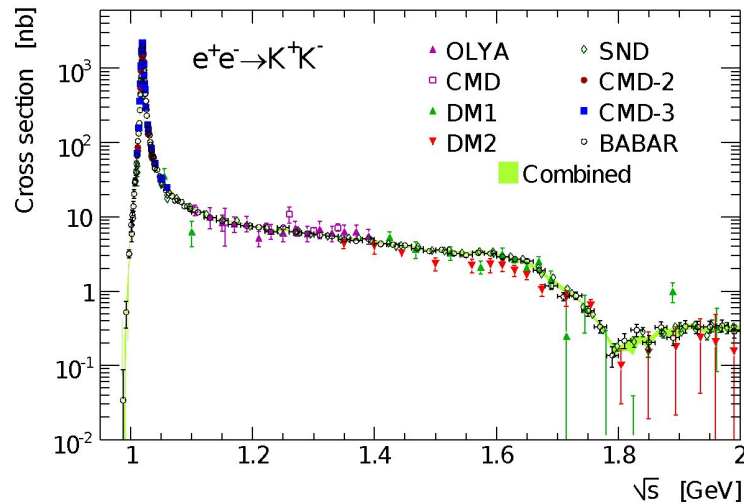


$$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-, e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$$



→ Essentially normalization differences w.r.t.  $\tau$  data: *cross-checks very desirable*

# Combination for the $e^+e^- \rightarrow K^+K^-$ channel

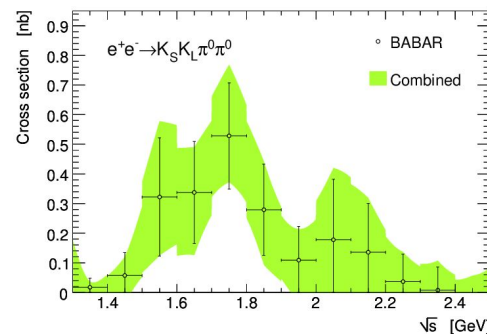
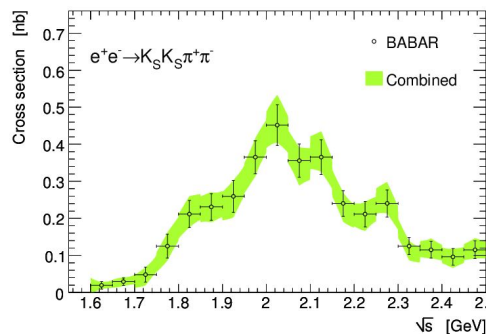
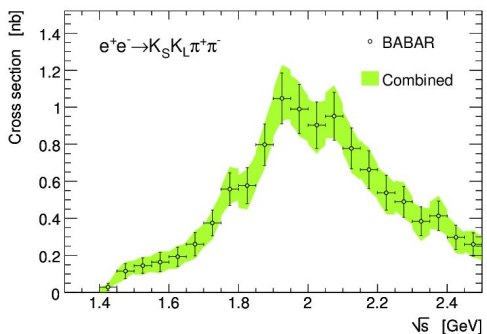
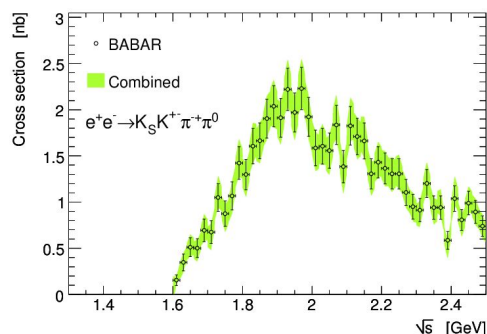
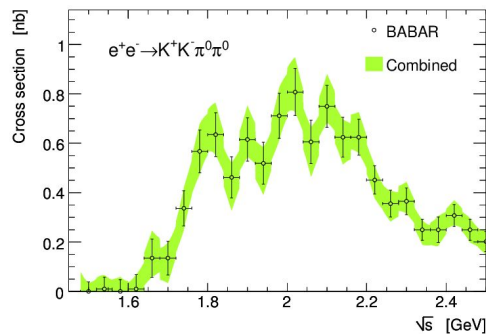
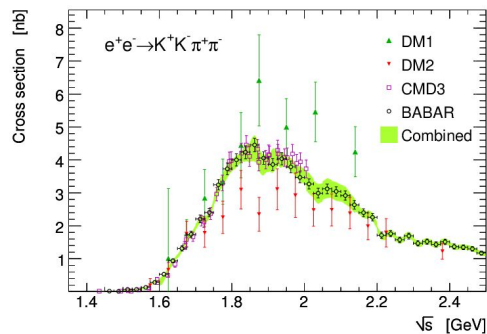
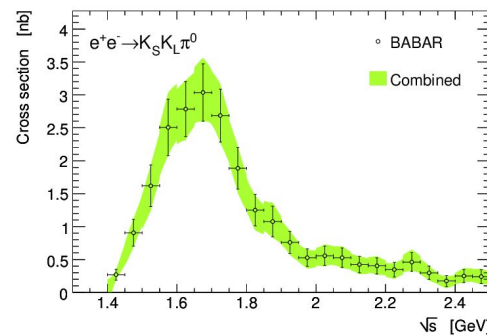
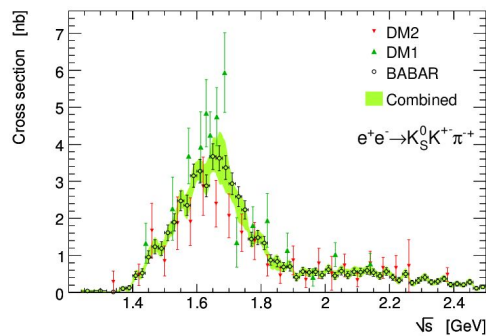
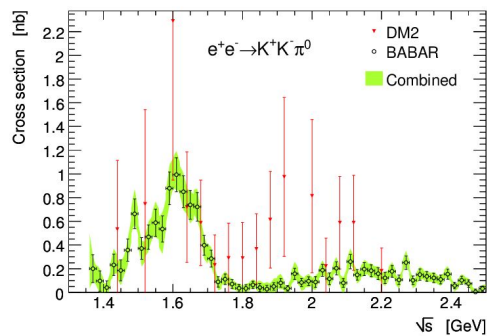


→ Tension between measurements

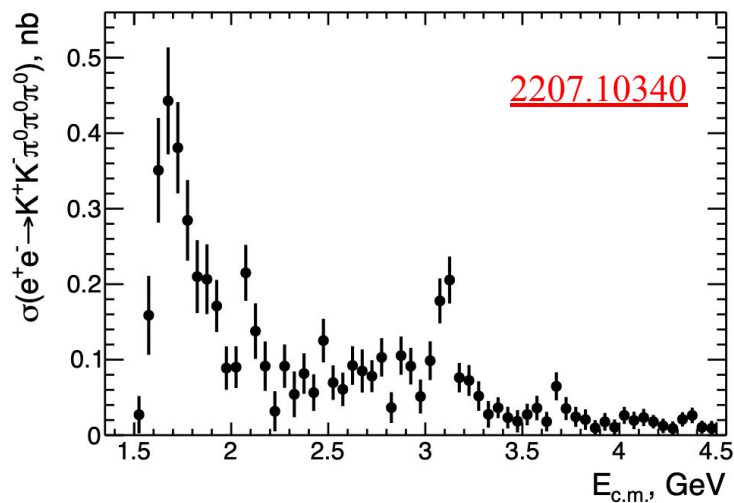
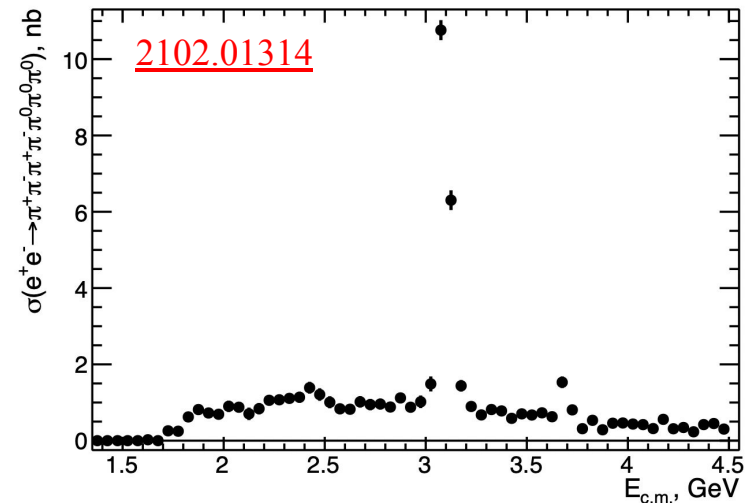
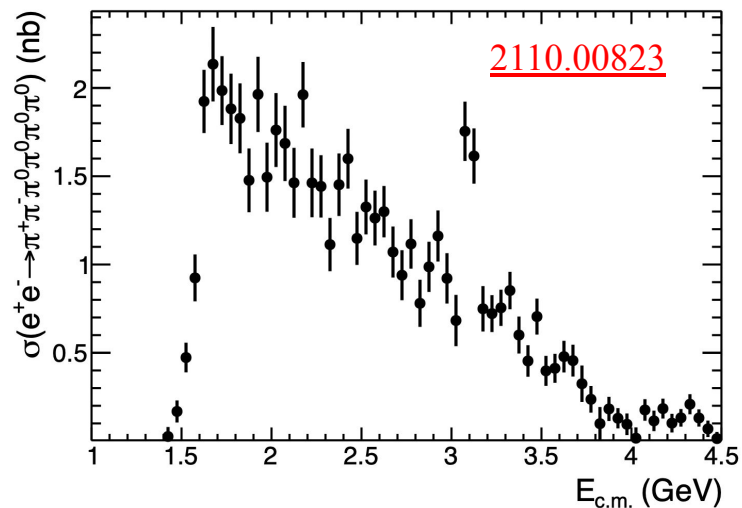
→  $a_\mu[\rightarrow 1.8\text{GeV}]$ :  $23.08 \pm 0.20$  (stat.)  $\pm 0.40$  (syst.) [ $10^{-10}$ ] (enhancement x 2.2)



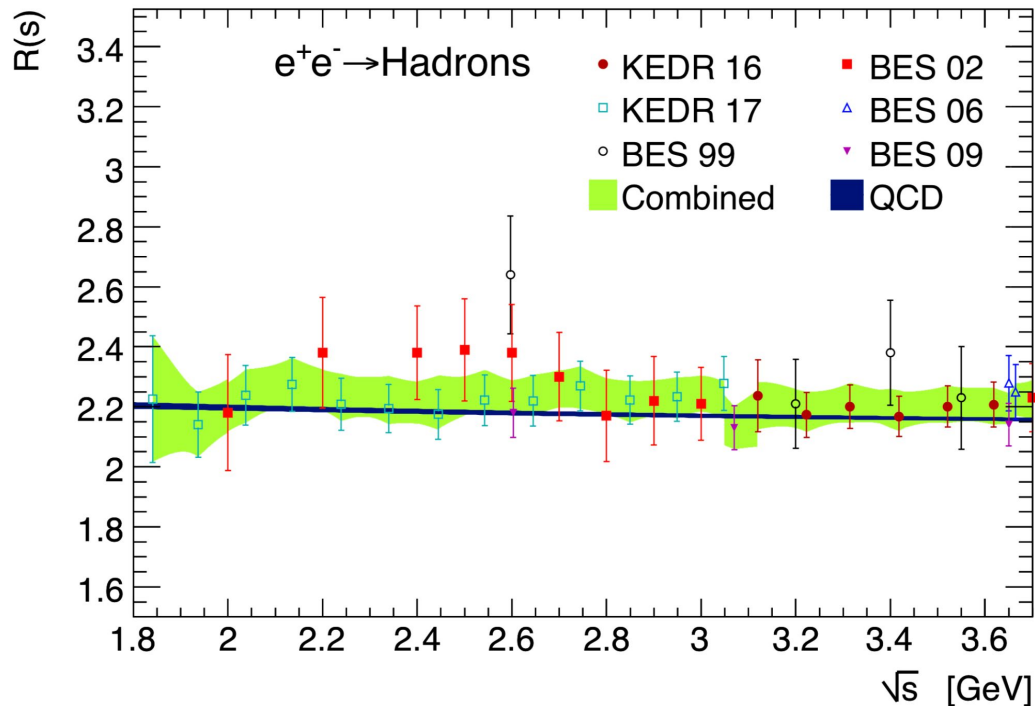
# Combination for the $e^+e^- \rightarrow KK\pi$ and $KK2\pi$ channels



# Examples of $\sim$ new high-multiplicity measurements



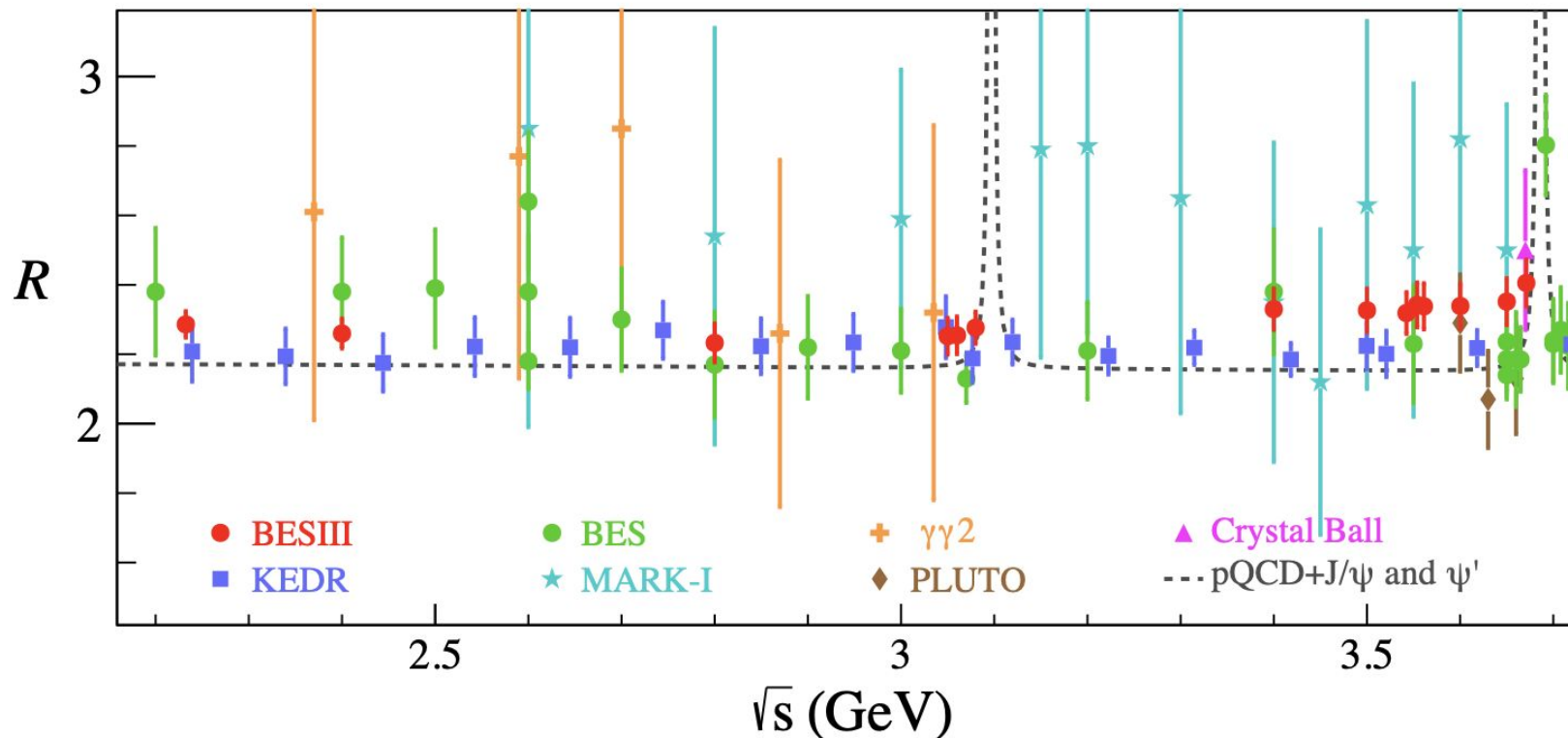
# Contributions from the 1.8 – 3.7 GeV region



- Contribution evaluated from pQCD (4 loops) +  $O(\alpha_s^2)$  quark mass corrections
- Uncertainties:  $\alpha_s$ , truncation of perturbative series, CIPT/FOPT,  $m_q$
- 1.8-2.0 GeV:  $7.65 \pm 0.31$  (data excl.);  $8.30 \pm 0.09$  (QCD); added syst.  $0.65 [10^{-10}]$
- 2.0-3.7 GeV:  $25.82 \pm 0.61$  (data);  $25.15 \pm 0.19$  (QCD); agreement within  $1\sigma$
- BES III results to be included:  $\sim$ tension with pQCD and with KEDR 16 (*next slide*)

# Comparison of inclusive measurements with pQCD

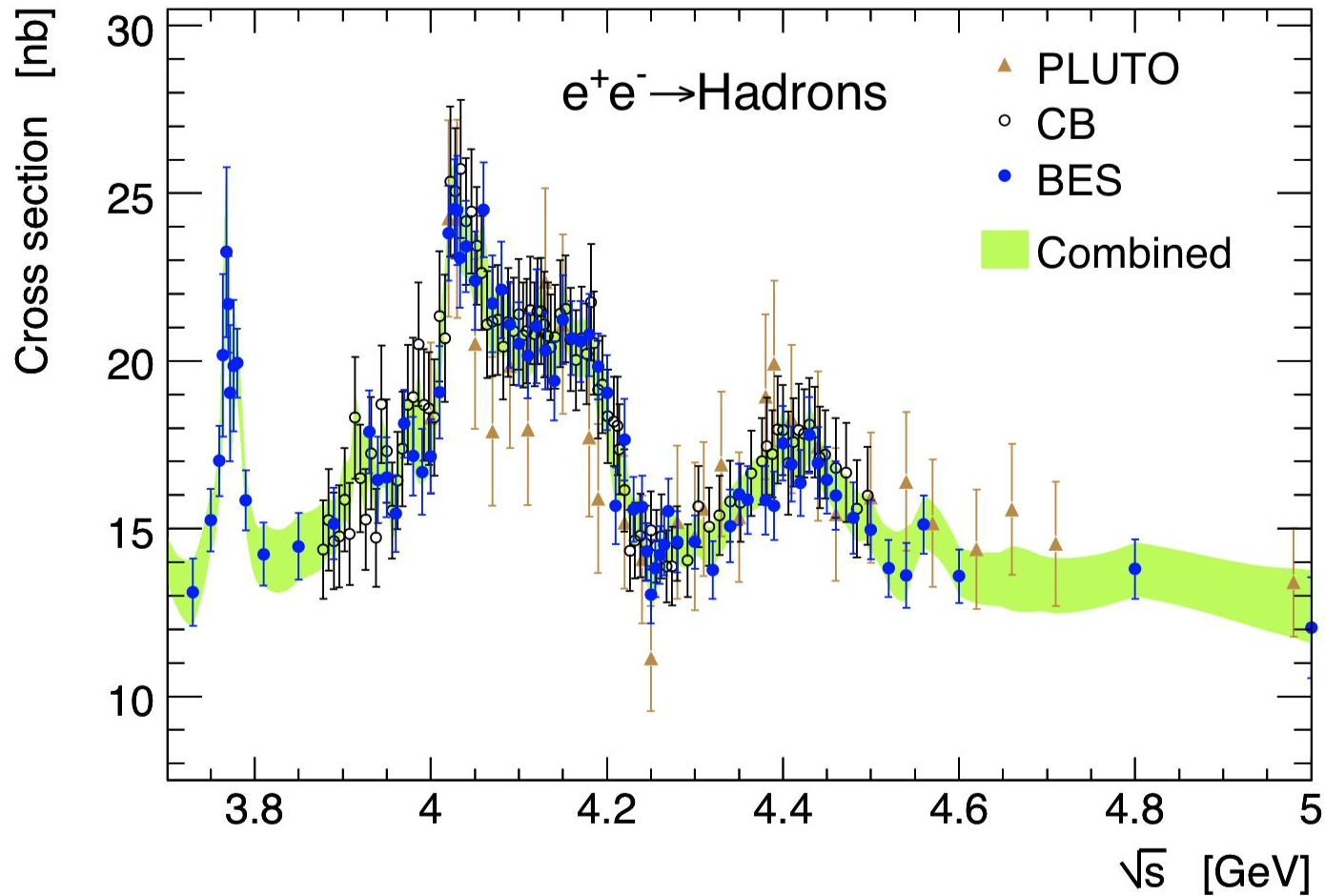
2112.11728



→ BES III results to be included:  $\sim$ tension with pQCD and with KEDR 16

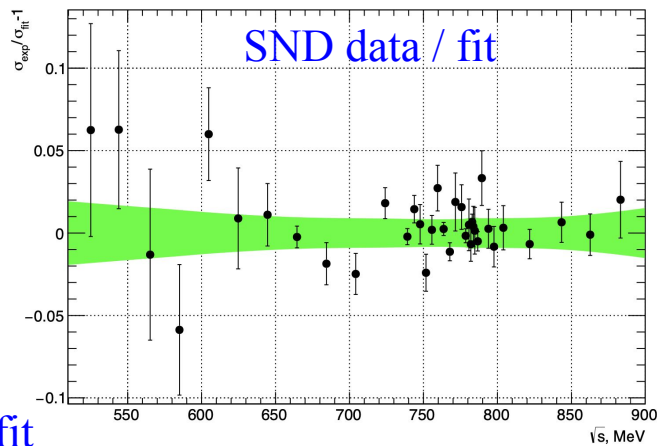
→ Another example of “*uncertainties on the uncertainties*” / *systematic effects to be understood* at the level of precision that is claimed

# Contributions from the charm resonance region



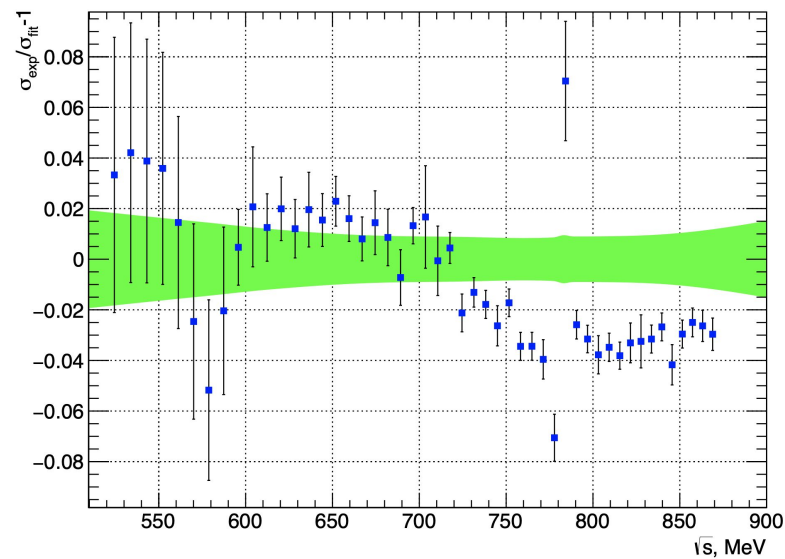
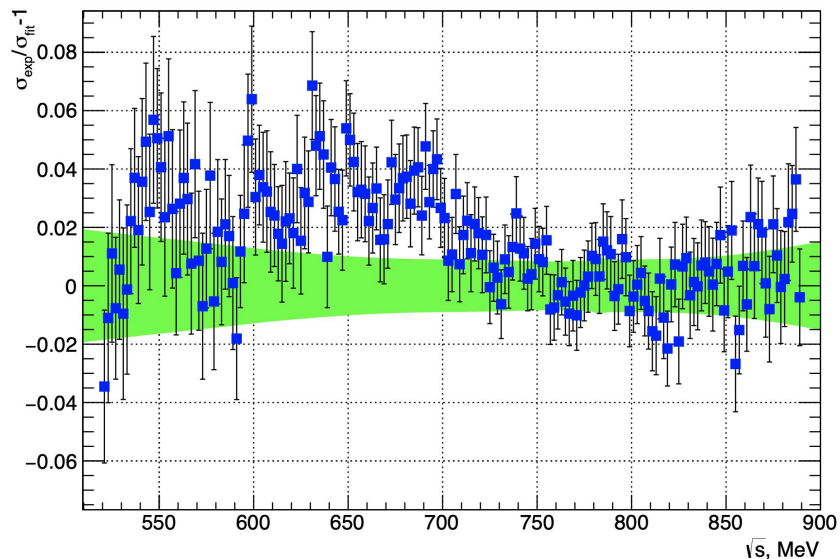
# Comparison of SND measurement with BABAR and KLOE

[2004.00263](#)



BABAR / SND fit

KLOE / SND fit



# Combine cross section data: goal and requirements

→ Goal: combine experimental spectra with arbitrary point spacing / binning

→ Requirements:

- Properly propagate uncertainties and correlations

- *Between measurements (data points/bins) of a given experiment*

- (covariance matrices and/or detailed split of uncertainties in sub-components)

- *Between experiments (common systematic uncertainties, e.g. VP)*

- based on detailed information provided in publications

- *Between different channels* – motivated by understanding of the meaning of systematic uncertainties and **identifying the common ones**

- BABAR luminosity (ISR or Bhabha), efficiencies (photon, Ks, Kl, modeling);

- BABAR radiative corrections;  $4\pi^2\pi^0-\eta\omega$

- CMD2  $\eta\gamma - \pi^0\gamma$ ; CMD2/3 luminosity; SND luminosity;

- FSR; hadronic VP (old experiments)

- (1<sup>st</sup> motivation for using DHMZ uncertainties as “baseline” in the g-2 TI White Paper)*

- Minimize biases

- Optimize g-2 integral uncertainty

- (without overestimating the precision with which the uncertainties of the measurements are known)*

# Combination procedure implemented in HVPTools software

## For each final bin:

- Compute an average value for each measurement and its uncertainty
- Compute correlation matrix between experiments
- Minimize  $\chi^2$  and get average coefficients (weights)
- Compute average between experiments and its uncertainty

## Evaluation of integrals and propagation of uncertainties:

- Integral(s) evaluated for nominal result and for each set of toy pseudo-experiments; uncertainty of integrals from RMS of results for all toys
- The pseudo-experiments also used to derive (statistical & systematic) covariance matrices of combined cross sections → Integral evaluation
- Uncertainties also propagated through  $\pm 1\sigma$  shifts of each uncertainty:
  - allows to account for correlations between different channels (for integrals and spectra)
- *Checked consistency between the different approaches*



# Combination procedure: weights of various measurements

For each final bin:

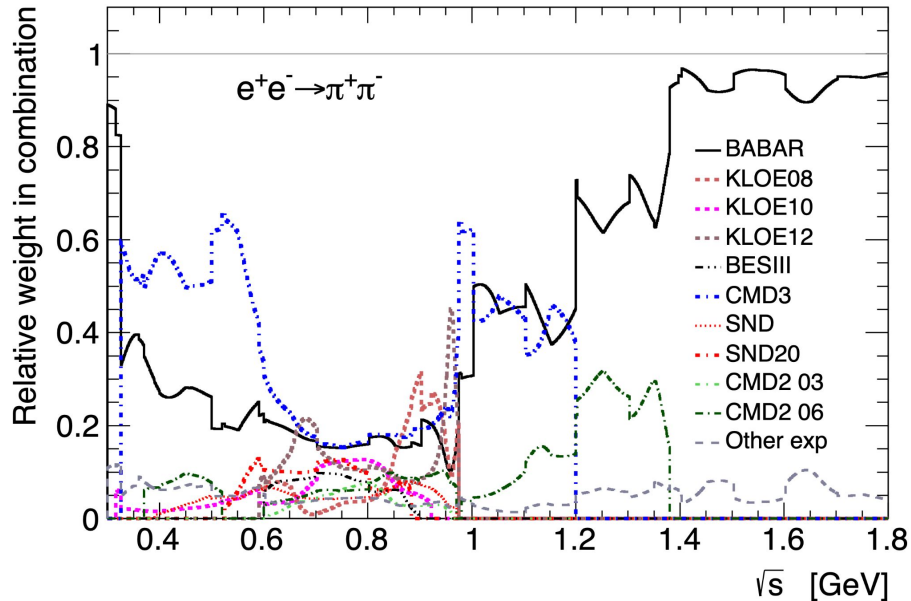
→ Minimize  $\chi^2$  and get average coefficients

Note: average weights must account for bin sizes / point spacing of measurements

(do not over-estimate the weight of experiments with large bins)

→ Weights in fine bins evaluated using a common (large) binning for measurements + interpolation

→ Compare the precisions on the same footing



→ Bins used by KLOE larger than the ones by BABAR in  $\rho$ - $\omega$  interference region (factor  $\sim 3$ )

→ Average dominated by BaBar, CMD3, KLOE, SND20  
BaBar covering full range

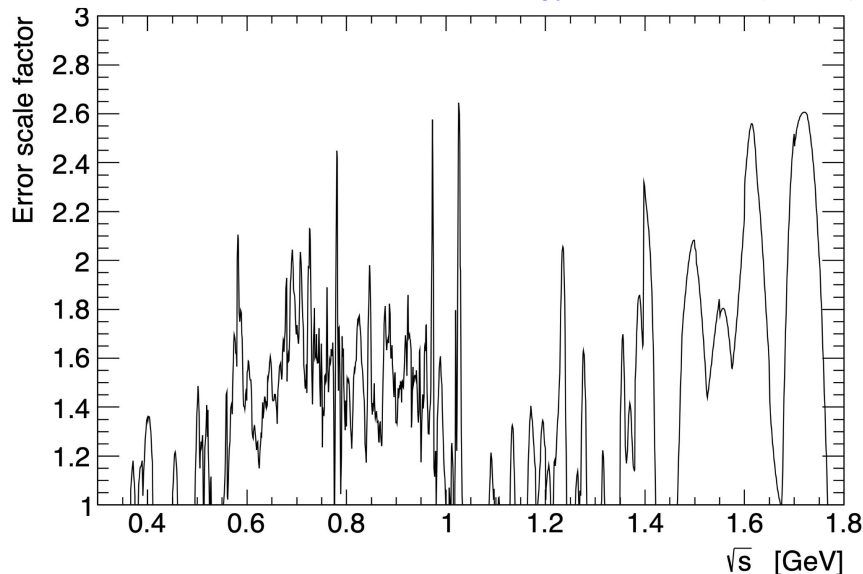
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# Combination procedure: compatibility between measurements

For each final bin:

→  $\chi^2/\text{ndof}$ : test locally the level of agreement between input measurements, *taking into account correlations*

→ Scale uncertainties in bins with  $\chi^2/\text{ndof} > 1$  (PDG): *locally conservative*; Adopted by KNT since '17



→ Tension between measurements, especially between KLOE & CMD3, which provide the smallest / largest cross-sections in the  $\rho$  region:

*Indication of underestimated uncertainties*

Motivates conservative uncertainty treatment

in combination fit (evaluation of weights / fits based on analyticity & unitarity to constrain uncertainties at low  $\sqrt{s}$  - *below*)

→ Observed (systematic) tension between measurements, beyond the local  $\chi^2/\text{ndof}$  rescaling

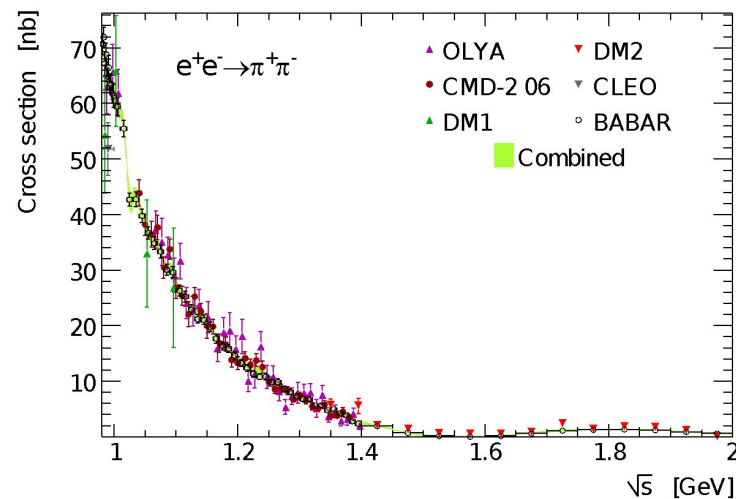
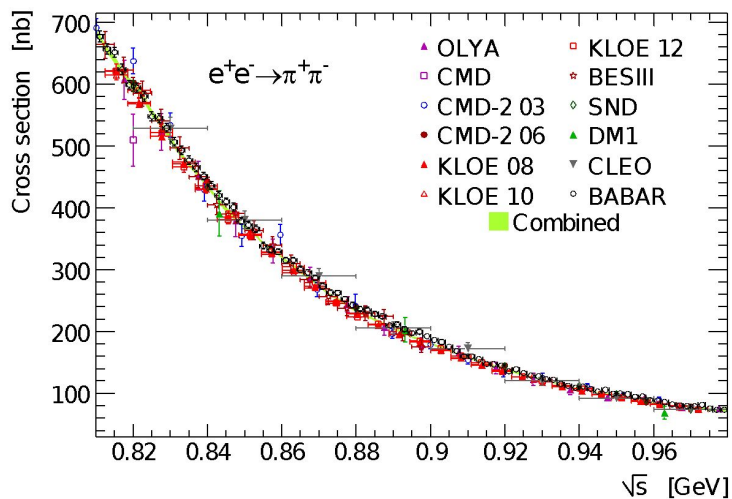
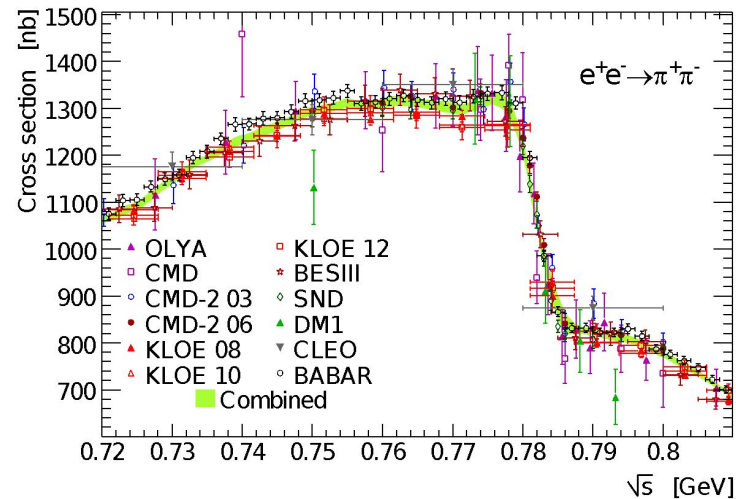
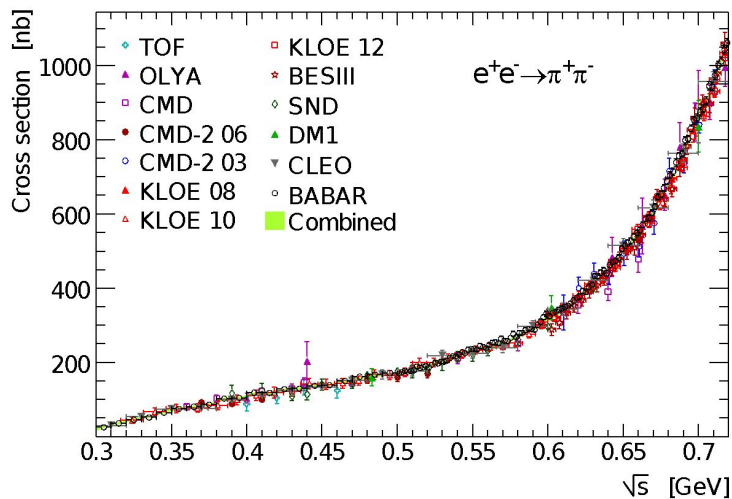
→ (Since 2019) Included extra (dominant) uncertainty: 1/2 difference between integrals w/o either BABAR or KLOE ( *2<sup>nd</sup> motivation for using DHMZ uncertainties as “baseline” in the TI WP* )

Extra uncertainty started to be adopted in other studies (2205.12963)

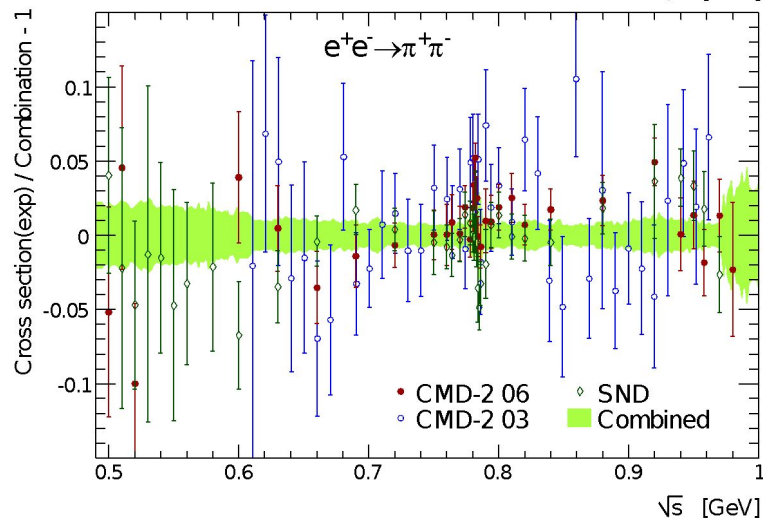
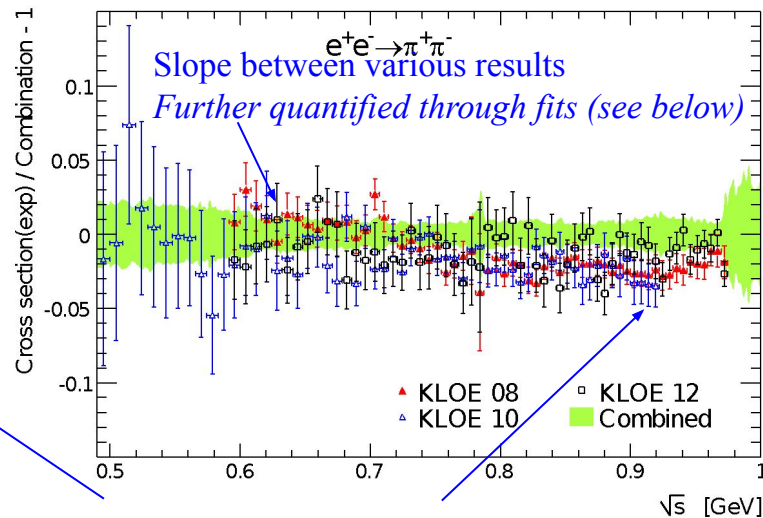
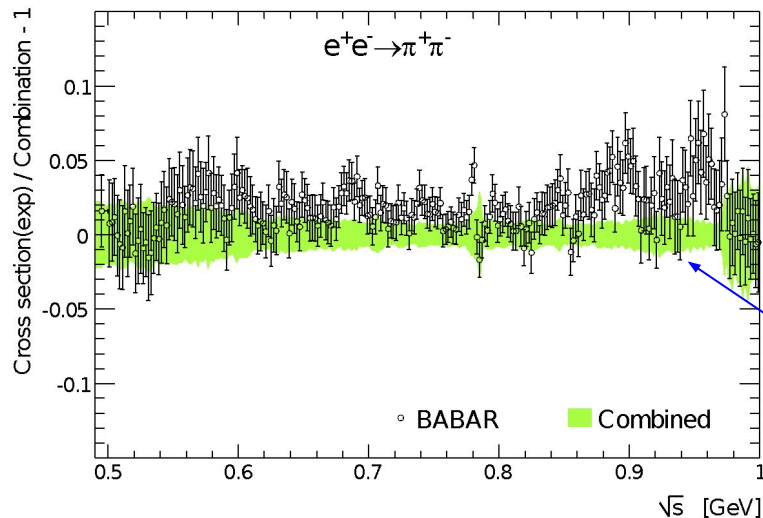
However, *tensions are larger now and we need to understand their source!*

Two panel TI discussions with 49 questions addressed to CMD3 did not allow to identify any major problem. CMD2 / CMD3 tension still open question!

# Combination for the $e^+e^- \rightarrow \pi^+\pi^-$ channel (DHMZ '19)



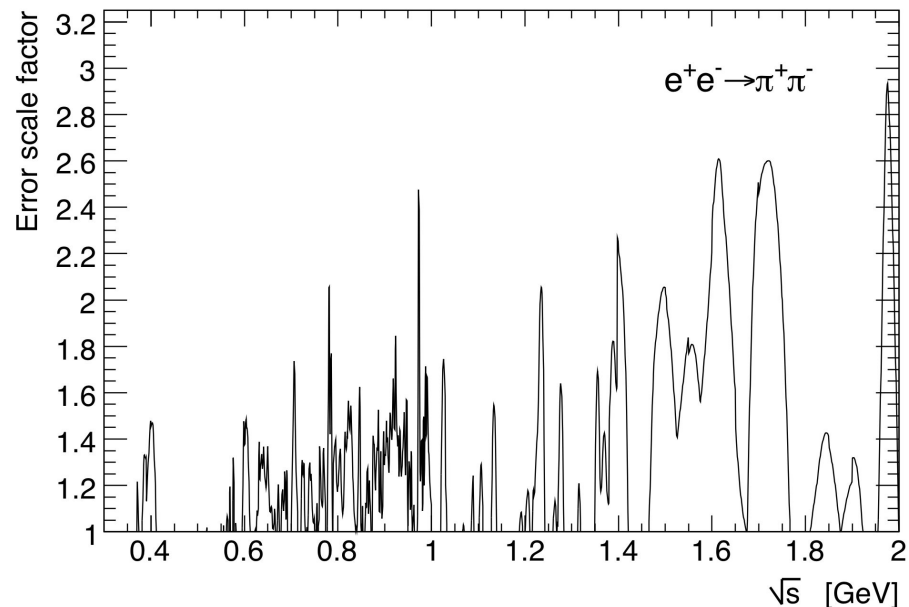
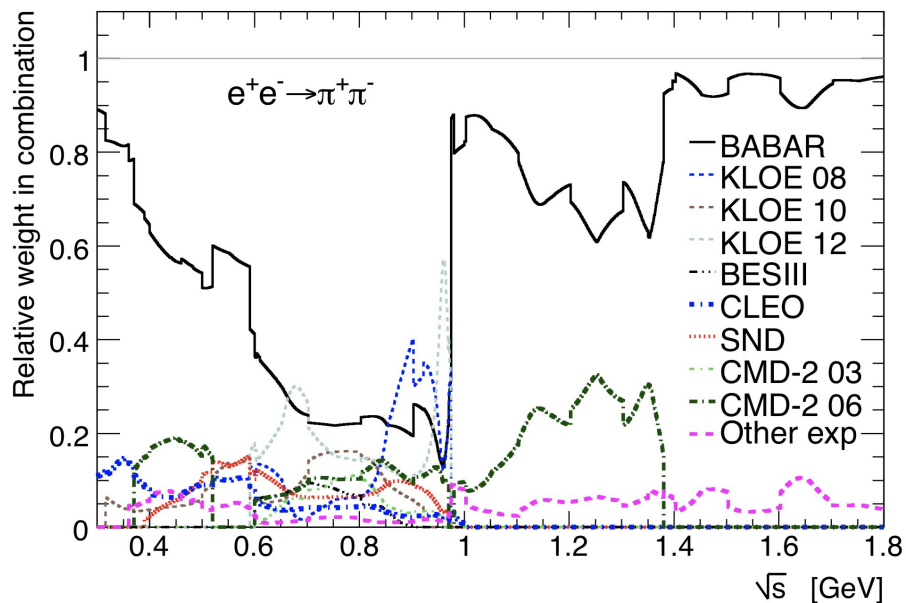
# More on the combination for the $e^+e^- \rightarrow \pi^+\pi^-$ channel (DHMZ '19)



Local tension & systematic trends  
 Indication of “uncertainties on uncertainties”  
 (i.e. unaccounted biases)

Other experiments not yet precise enough  
 to discriminate  
 (see however update from SND: ~significant  
 tension with KLOE above 720 MeV)

# Combining the $e^+e^- \rightarrow \pi^+\pi^-$ data: weights and tension (DHMZ '19)



# Improving $a_\mu$ through fits for the $e^+e^- \rightarrow \pi^+\pi^-$ channel (*Since 2019*)

→ Fit bare form-factor using 6 param. model based on *analyticity* and *unitarity*

$$|F_\pi^0|^2 = |R(s) \times J(s)|^2$$

$$R(s) = 1 + \alpha_V s + \frac{\kappa s}{m_\omega^2 - s - im_\omega \Gamma_\omega} \quad (1611.09359, C. Hanhart et al.)$$

$$J(s) = e^{1 - \frac{\delta_1(s_0)}{\pi}} \left(1 - \frac{s}{s_0}\right)^{\left[1 - \frac{\delta_1(s_0)}{\pi}\right] \frac{s_0}{s}} \left(1 - \frac{s}{s_0}\right)^{-1} e^{\frac{s}{\pi} \int_{4m_\pi^2}^{s_0} dt \frac{\delta_1(t)}{t(t-s)}}$$

Omnès integral

(hep-ph/0402285, F.J. Yndurain et al.)

$$\cot \delta_1(s) = \frac{\sqrt{s}}{2k^3} (m_\rho^2 - s) \left[ \frac{2m_\pi^3}{m_\rho^2 \sqrt{s}} + B_0 + B_1 \omega(s) \right]$$

$$k = \frac{\sqrt{s - 4m_\pi^2}}{2}$$

(1102.2183, F.J. Yndurain et al.)

$$\omega(s) = \frac{\sqrt{s} - \sqrt{s_0 - s}}{\sqrt{s} + \sqrt{s_0 - s}} \quad \sqrt{s_0} = 1.05 \text{ GeV}$$

→ Conservative  $\chi^2$  (diagonal matrix) & local rescaling of input uncertainties

→ Full propagation of uncertainties & correlations using pseudo-experiments

DHMZ - 1908.00921

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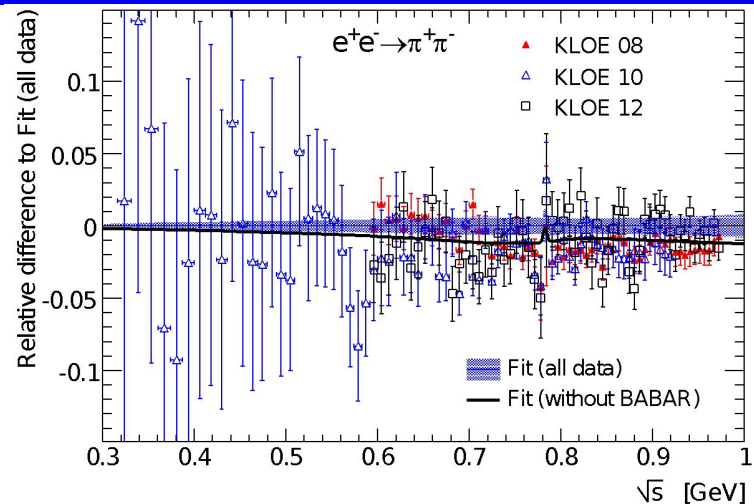
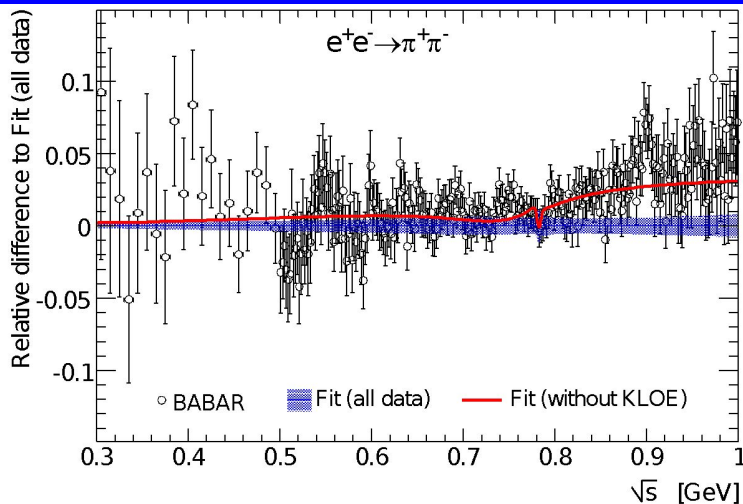
# Fit parameters, uncertainties and correlations $e^+e^- \rightarrow \pi^+\pi^-$

	$\alpha_V$	$\kappa[10^{-4}]$	$B_0$	$B_1$	$m_\rho$ [MeV]	$m_\omega$ [MeV]
$\alpha_V$	$0.133 \pm 0.020$	0.52	-0.45	-0.97	0.90	-0.25
$\kappa[10^{-4}]$		$21.6 \pm 0.5$	-0.33	-0.57	0.64	-0.08
$B_0$			$1.040 \pm 0.003$	0.40	-0.40	0.29
$B_1$				$-0.13 \pm 0.11$	-0.96	0.20
$m_\rho$ [MeV]					$774.5 \pm 0.8$	-0.17
$m_\omega$ [MeV]						$782.0 \pm 0.1$

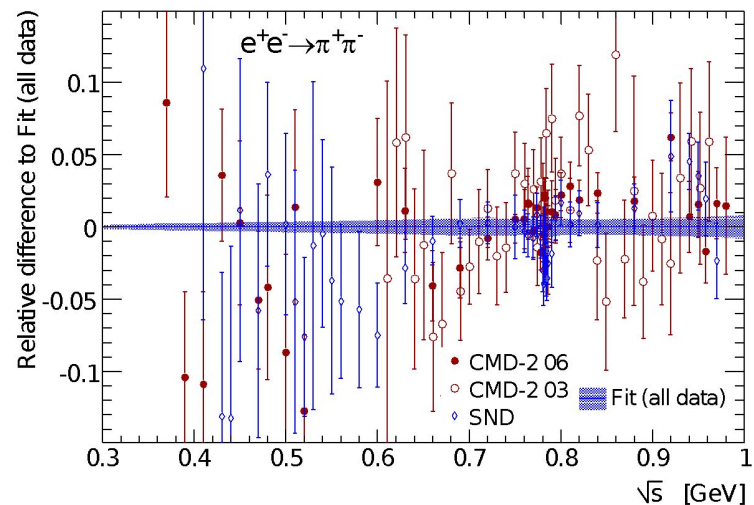
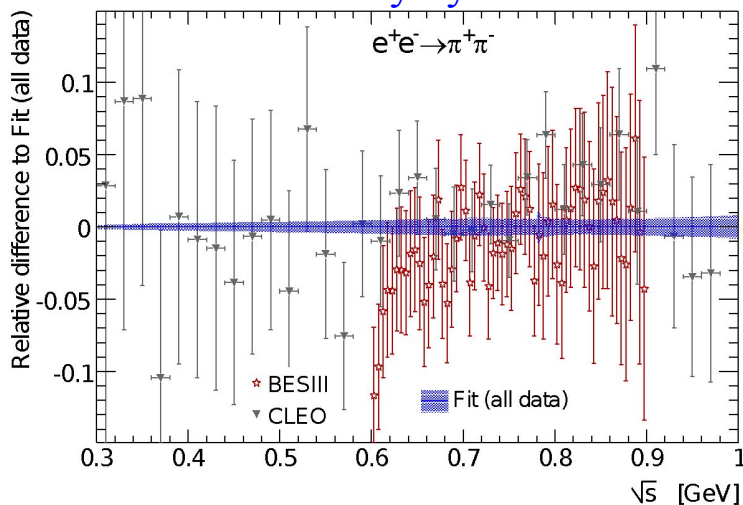
→  $\kappa$  corresponds to a Br ( $\omega \rightarrow \pi^+\pi^-$ ) of  $(2.09 \pm 0.09) \cdot 10^{-2}$ , in agreement with the result extracted from the fit of arXiv:1810.00007,  $(1.95 \pm 0.08) \cdot 10^{-2}$ . Both values disagree with the PDG average  $(1.51 \pm 0.12) \cdot 10^{-2}$ , dominated by the result of arXiv:1611.09359 which uses fits to essentially the same data.

→ The fitted  $\omega$  mass is found to be lower than the PDG average obtained from  $3\pi$  decays by  $(0.65 \pm 0.12 \pm 0.12_{\text{PDG}})$  MeV, in agreement with previous fits of the  $\rho - \omega$  interference in the  $2\pi$  spectrum (see e.g. arXiv:1205.2228 and arXiv:1810.00007).

# Fit performed up to 1 GeV: comparison with data

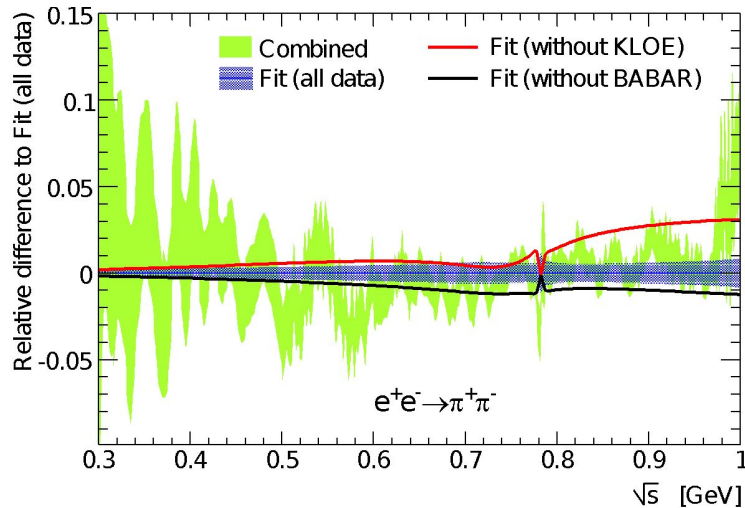


→ Fit constrained mainly by BABAR and KLOE measurements





# Fit performed up to 1 GeV, Result used up to 0.6 GeV



→ Use fit only below 0.6 GeV for  $a_\mu$  integral:

- where data is less precise and scarce
- less impacted by potential uncertainties of inelastic effects

$\sqrt{s}$ range [GeV]	$a_\mu^{\text{had}} [10^{-10}]$ Fit	$a_\mu^{\text{had}} [10^{-10}]$ Data Integration
0.3 - 0.6	$109.80 \pm 0.37_{\text{exp}} \pm 0.36_{\text{para}^*}$	$109.6 \pm 1.0_{\text{exp}}$

→ The difference  $0.2 \pm 0.9$   
(72% correlation accounted for)

→ The fit improves the precision by a factor  $\sim 2$

(\*) Parameter uncertainty corresponds to variations with/without the  $B_1$  term in the phase shift formula and  $\sqrt{s}_0$  varied from 1.05 GeV to 1.3 GeV (absolute values summed linearly), *checked to be statistically significant*

# Combined results: Fit [ $<0.6\text{GeV}$ ] + Data[ $0.6-1.8\text{GeV}$ ]

→ Full uncertainty propagation using the same pseudo-experiments as for the spline-based combination: 62% correlation among the two contributions

$\sqrt{s}$ range [GeV]	$a_{\mu}^{\text{had}}$ [ $10^{-10}$ ] All data	$a_{\mu}^{\text{had}}$ [ $10^{-10}$ ] All but BABAR	$a_{\mu}^{\text{had}}$ [ $10^{-10}$ ] All but KLOE
threshold - 1.8	$506.9 \pm 1.9_{\text{total}}$	$505.0 \pm 2.1_{\text{total}}$	$510.6 \pm 2.2_{\text{total}}$

→ The difference “All but BABAR” and “All but KLOE” = 5.6, to be compared with 1.9 uncertainty with “All data”

- The local error inflation is not sufficient to amplify the uncertainty
- Global tension (normalisation/shape) not previously accounted for
- Potential underestimated uncertainty in at least one of the measurements?
- Other measurements not precise enough to discriminate BABAR / KLOE

→ Given the fact we do not know which dataset is problematic, we decide to:

- Add half of the discrepancy ( $2.8 \times 10^{-10}$ ) as an uncertainty (corrected local PDG inflation to avoid double counting)
- Take (“All but BABAR” + “All but KLOE”) / 2 as central value

Channel	$a_{\mu}^{\text{had,LO}}$ [ $10^{-10}$ ]	$\Delta\alpha_{\text{had}}(m_Z^2)$ [ $10^{-4}$ ]
$\pi^+\pi^-$	$507.85 \pm 0.83 \pm 3.23 \pm 0.55$	$34.50 \pm 0.06 \pm 0.20 \pm 0.04$

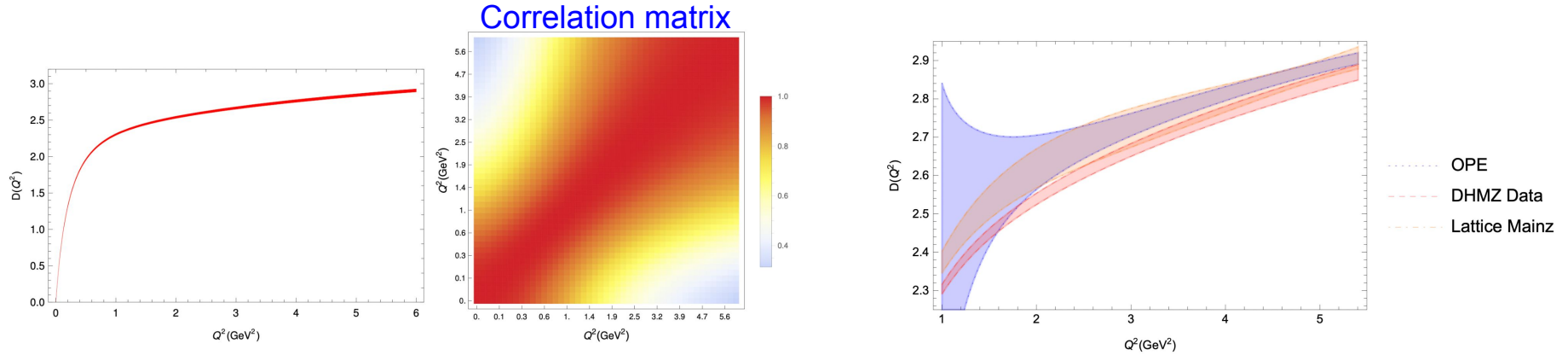
→ Potential precision improvement for  $a_{\mu}$ ; less important for  $\Delta\alpha_{\text{had}}(m_Z^2)$ , BABAR-KLOE syst. ~16% of total uncertainty

# $\alpha_s$ extraction from the Adler function and test of the RGE

2302.01359

$$D(Q^2) \equiv \sum_{i,j} Q_i Q_j D_{ij}(Q^2) = 3\pi Q^2 \frac{d\Delta\alpha_{\text{had}}(Q^2)}{\alpha dQ^2}$$

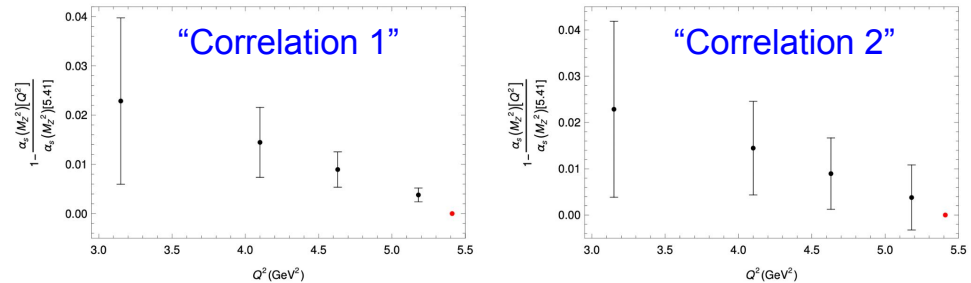
→ Experimental values with correlations; Theoretical predictions: perturbative + non-perturbative corrections (OPE)



→ Using W.Av.  $\alpha_s(M_Z)$  value  $\sim 0.118$ : OPE prediction in good agreement with Lattice QCD, above dispersive

→ Fit DHMZ data:  $\alpha_s^{(n_f=5)}(M_Z^2) = 0.1136 \pm 0.0025$

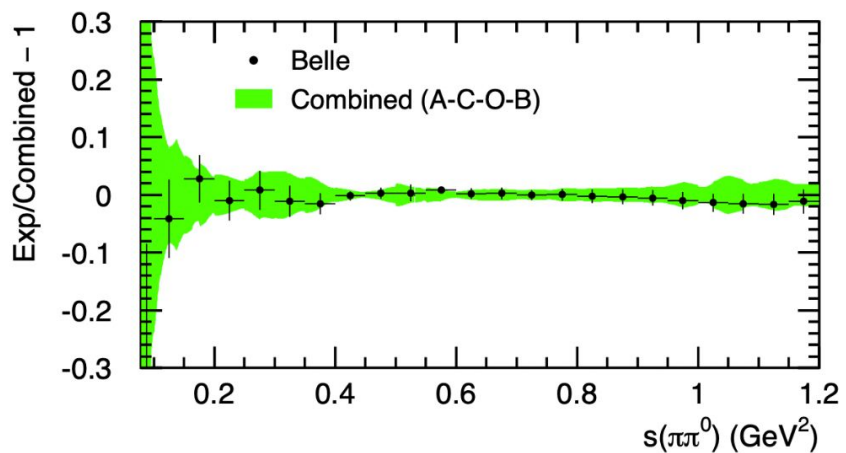
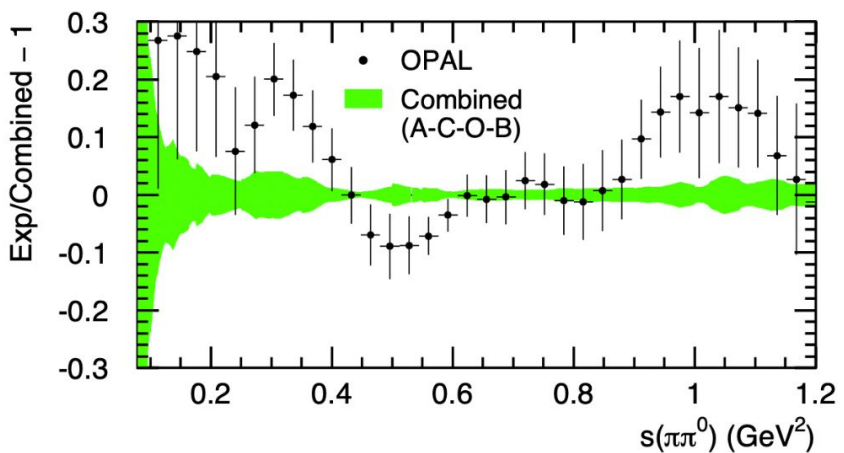
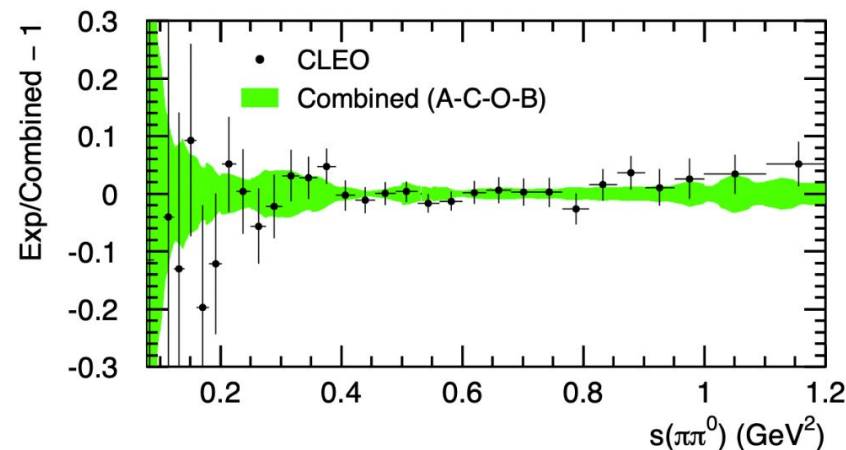
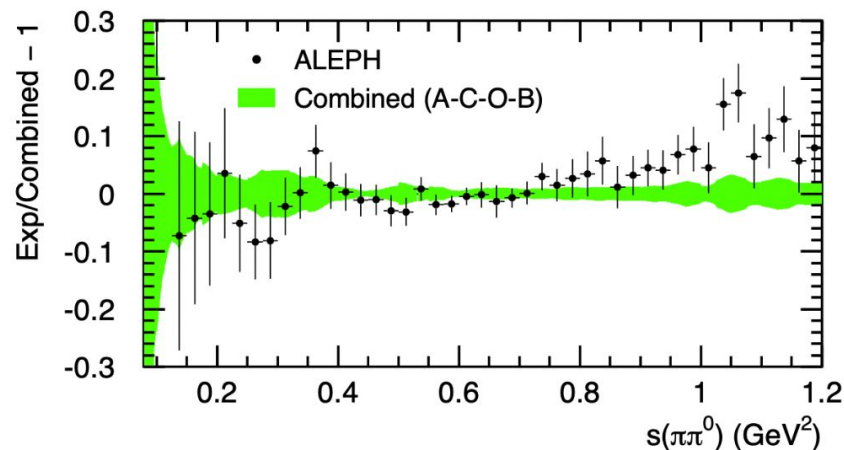
→ Performing RGE test and evaluating its precision, with different correlation scenarios for theory uncertainties



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# Combining the $\tau$ data in the $\pi\pi$ channel

1312.1501

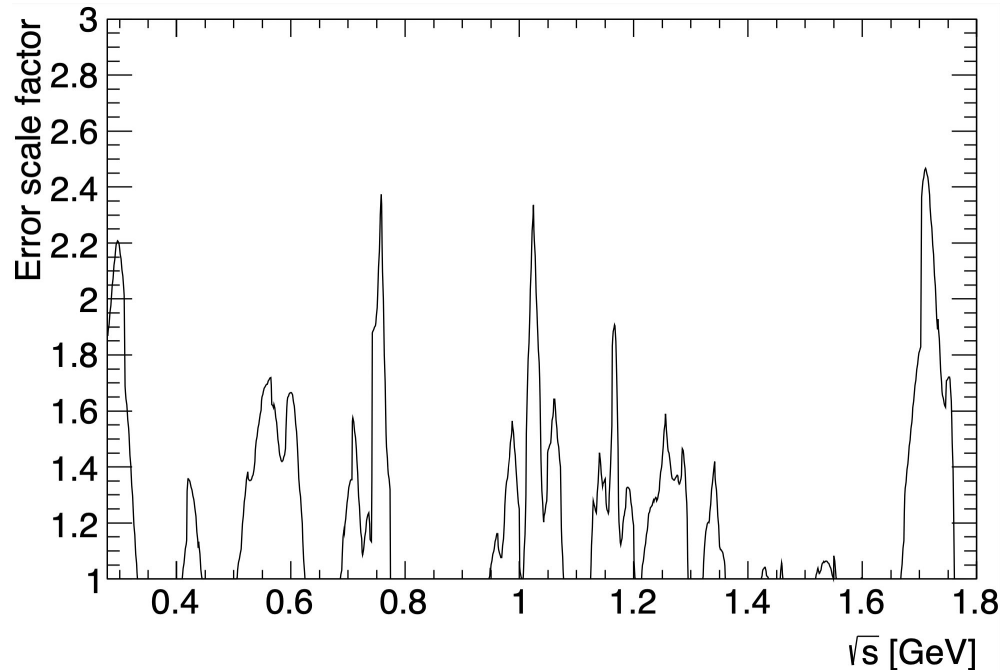


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# Combination: compatibility between measurements

For each final bin:

- $\chi^2/\text{ndof}$ : test locally the level of agreement between input measurements, *taking into account correlations*
- Scale uncertainties in bins with  $\chi^2/\text{ndof} > 1$  (PDG)



→ Level of agreement significantly better than the one observed for  $e^+e^- \rightarrow \pi^+\pi^-$  data

# Combination: weights of various measurements

For each final bin:

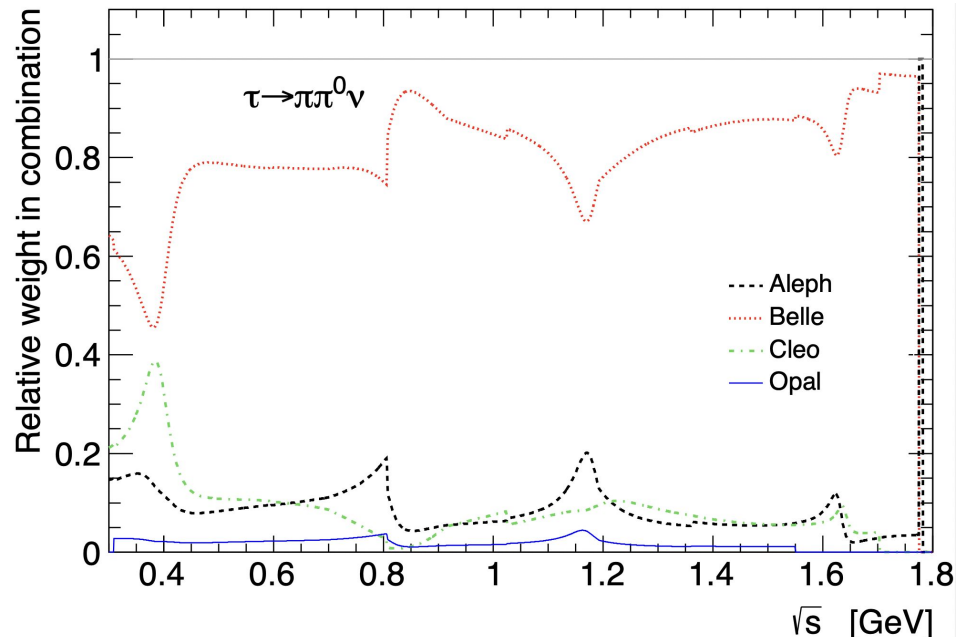
→ Minimize  $\chi^2$  and get average coefficients

Note: average weights must account for bin sizes / point spacing of measurements

(Compare the precisions on the same footing: do not over-estimate the weight of experiments with large bins)

→ Weights in fine bins evaluated using a common (large) binning for measurements + interpolation

→ Their determination also integrates bin-to-bin statistical & systematic correlations on moderate energy ranges



→ Shape information provided mainly by Belle (reflected by the weights from the combination of spectra)

# Combining the $\tau$ data in the $\pi\pi$ channel

→ Normalisation dominated by ALEPH (directly impacting and very relevant for the integrals)

Experiment	$a_\mu^{\text{had,LO}}[\pi\pi, \tau] (10^{-10})$	
	$2m_{\pi^\pm} - 0.36 \text{ GeV}$	$0.36 - 1.8 \text{ GeV}$
ALEPH	$9.80 \pm 0.40 \pm 0.05 \pm 0.07$	$501.2 \pm 4.5 \pm 2.7 \pm 1.9$
CLEO	$9.65 \pm 0.42 \pm 0.17 \pm 0.07$	$504.5 \pm 5.4 \pm 8.8 \pm 1.9$
OPAL	$11.31 \pm 0.76 \pm 0.15 \pm 0.07$	$515.6 \pm 9.9 \pm 6.9 \pm 1.9$
Belle	$9.74 \pm 0.28 \pm 0.15 \pm 0.07$	$503.9 \pm 1.9 \pm 7.8 \pm 1.9$
Combined	$9.82 \pm 0.13 \pm 0.04 \pm 0.07$	$506.4 \pm 1.9 \pm 2.2 \pm 1.9$

**Table 6.** The isospin-breaking-corrected  $a_\mu^{\text{had,LO}}[\pi\pi, \tau]$  (in units of  $10^{-10}$ ) from the measured mass spectrum by ALEPH, CLEO, OPAL and Belle, and the combined spectrum using the corresponding branching fraction values. The results are shown separately in two different energy ranges. The first errors are due to the shapes of the mass spectra, which also include very small contributions from the  $\tau$ -mass and  $|V_{ud}|$  uncertainties. The second errors originate from  $B_{\pi\pi^0}$  and  $B_e$ , and the third errors are due to the isospin-breaking corrections, which are partially anti-correlated between the two energy ranges. The last row gives the evaluations using the combined spectra.

Individual measurements with the corresponding uncertainties:

ALEPH:  $511.0 \pm 5.3$  ( $\pm 1.9$  common, from IB)

CLEO:  $514.2 \pm 10.1$

OPAL:  $526.9 \pm 12.3$

Belle:  $513.7 \pm 8.0$

→ Most precise determination from ALEPH, due to most precise Br

→ Uncertainty from combined spectra ( $\pm 2.9$ ) smaller than uncertainty from weighted average of integrals ( $\pm 3.8$ ):

Due to better use of the available information on the precision of the measurements ( Br and mass-dependent uncertainties)

$\chi^2$  : 1.45/3 dof, when averaging the 4 individual integrals

$\chi^2$  : 1.88/3-4 dof, when comparing the 4 individual integrals with the integral of the combined spectrum

→ Excellent agreement among the 4 measurements

# Moment integrals from $\tau$ data ( $2\pi$ channel) with IB corrections

$$\sigma_{e^+e^- \rightarrow \pi^+\pi^-}^{I=1} = \frac{4\pi\alpha^2}{s} v_{1,\pi^-\pi^0\nu_\tau}$$

$$a_\mu [0.36, 1.775 \text{ GeV}] = (507.51 \pm 1.86) \times 10^{-10}$$

uncertainties from *combined spectrum*

$$\pm 2.12 \times 10^{-10}$$

uncertainties from *normalisation (Be & B $\pi\pi^0$ )*

$$\pm 1.9 \times 10^{-10}$$

uncertainties from *IB uncertainties*

$$v_{1,X^-}(s) = \frac{m_\tau^2}{6|V_{ud}|^2} \frac{\mathcal{B}_{X^-}}{\mathcal{B}_e} \frac{1}{N_X} \frac{dN_X}{ds} \times \left(1 - \frac{s}{m_\tau^2}\right)^2 \left(1 + \frac{2s}{m_\tau^2}\right)^{-1} \frac{R_{IB}(s)}{S_{EW}}$$

$$a_\mu [0.36, 1.775 \text{ GeV}] = (507.51 \pm 3.41) \times 10^{-10}$$

uncertainties from *combined spectrum, normalisation (Be & B $\pi\pi^0$ ) and IB uncertainties*

→ Next slides:

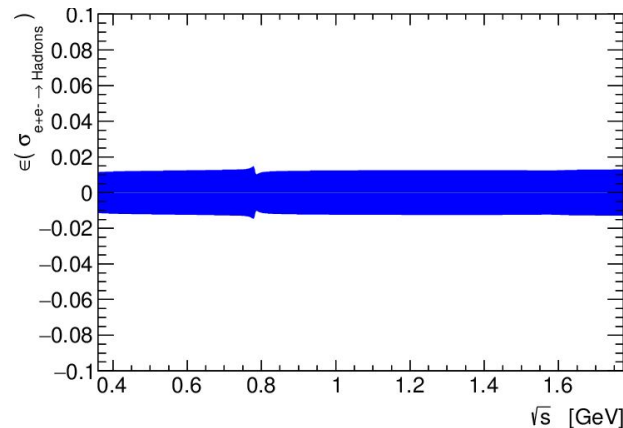
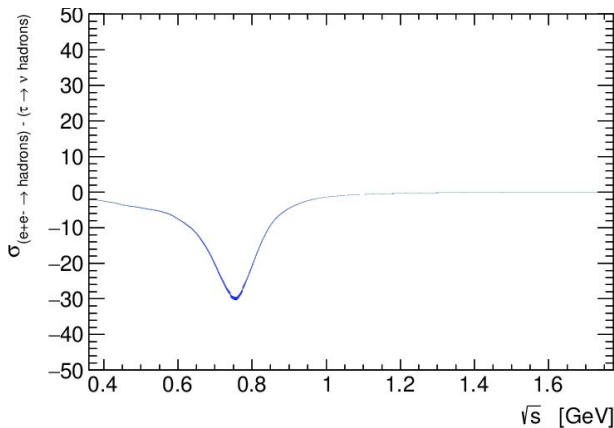
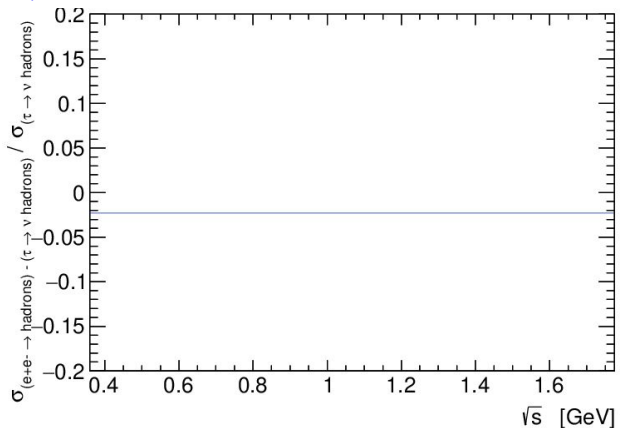
Display of energy dependence for IB corrections and uncertainties

$$R_{IB}(s) = \frac{\text{FSR}(s)}{G_{EM}(s)} \frac{\beta_0^3(s)}{\beta_-^3(s)} \left| \frac{F_0(s)}{F_-(s)} \right|^2$$

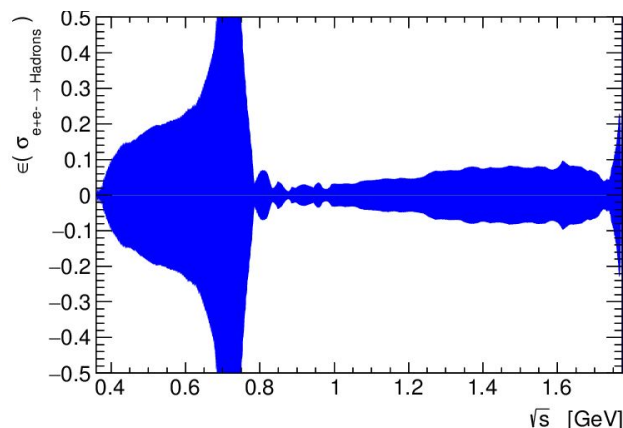
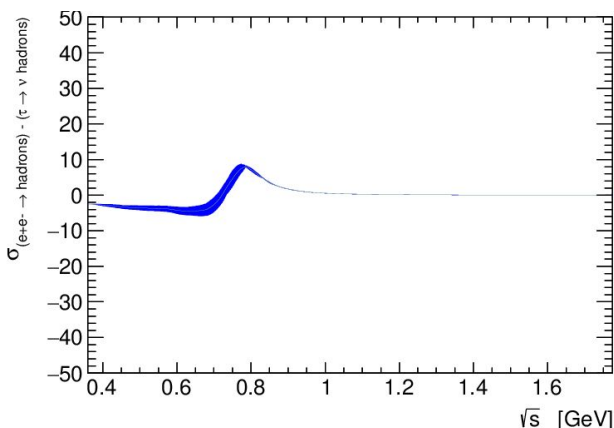
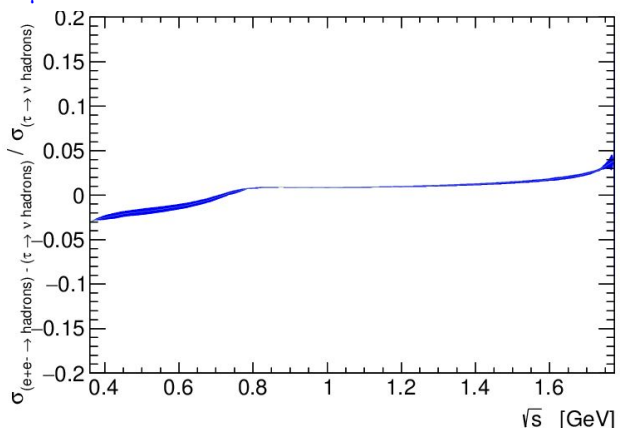


# Moment integrals from IB corrections for $\tau$ data ( $2\pi$ channel)

$a_\mu [0.36, 1.775 \text{ GeV}] = (-11.94 \pm 0.15) \times 10^{-10} \rightarrow \text{Corr. \& unc. from IB Sew (Short-distance EW radiative effects)}$

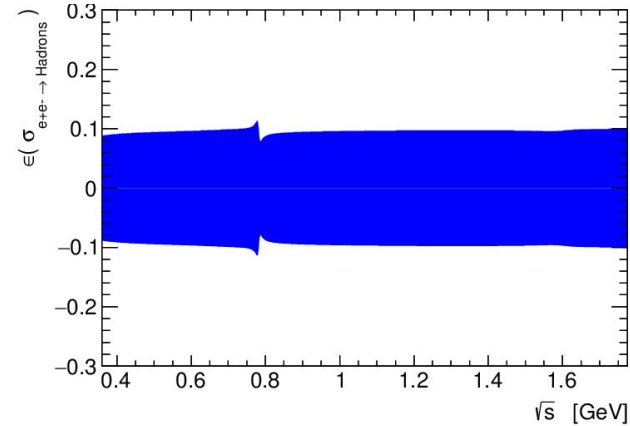
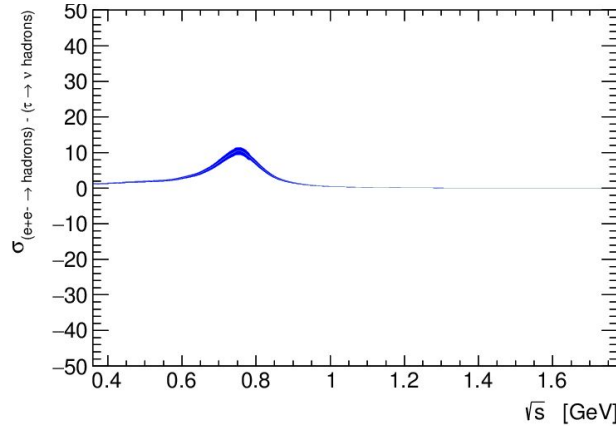
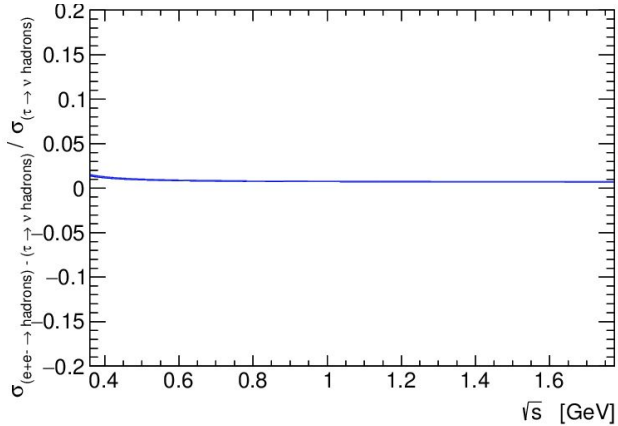


$a_\mu [0.36, 1.775 \text{ GeV}] = (-1.31 \pm 0.94) \times 10^{-10} \rightarrow \text{Corr. \& unc. from IB Gem (long-distance radiative corrections)}$

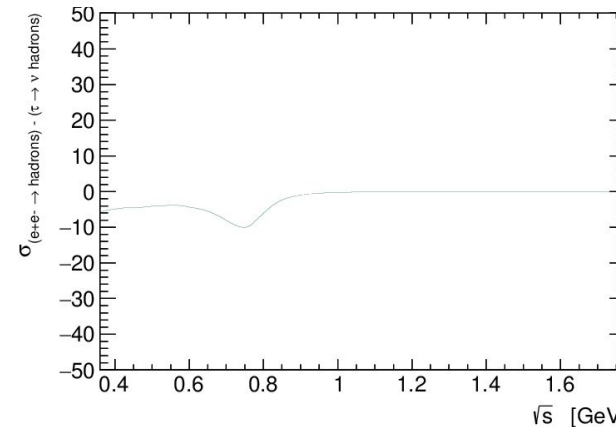
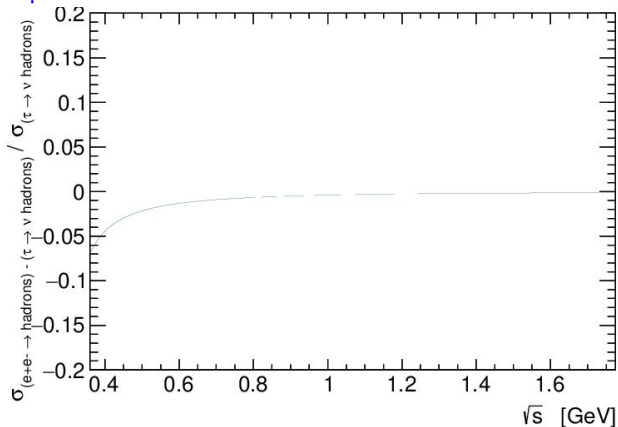


# Moment integrals from IB corrections for $\tau$ data ( $2\pi$ channel)

$a_\mu [0.36, 1.775 \text{ GeV}] = (4.41 \pm 0.43) \times 10^{-10} \rightarrow$  Corrections and uncertainties from *IB FSR*



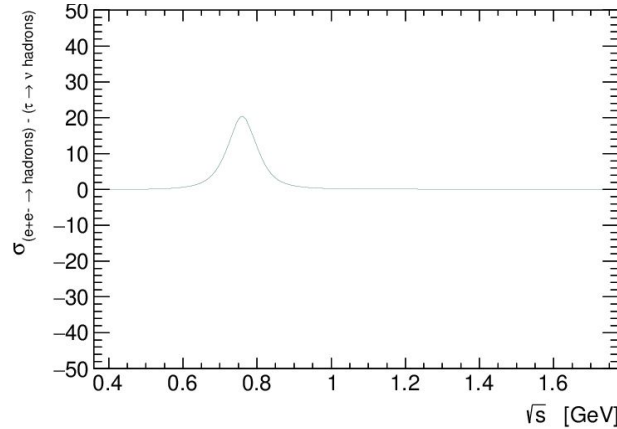
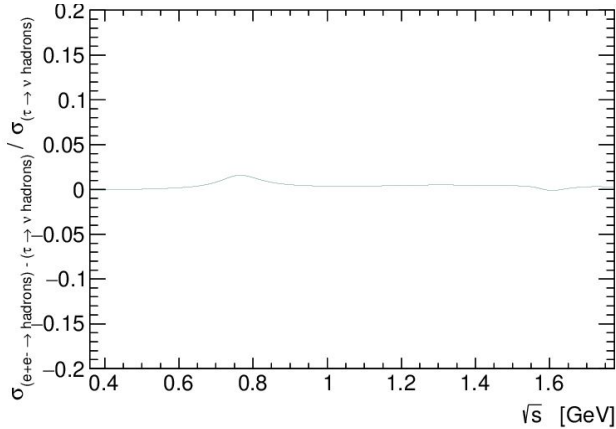
$a_\mu [0.36, 1.775 \text{ GeV}] = (-6.05 \pm 0) \times 10^{-10} \rightarrow$  Corrections and uncertainties from *IB beta* ( $\pi^\pm - \pi^0$  mass splitting)



$$\frac{\beta_0^3(s)}{\beta_-^3(s)}$$

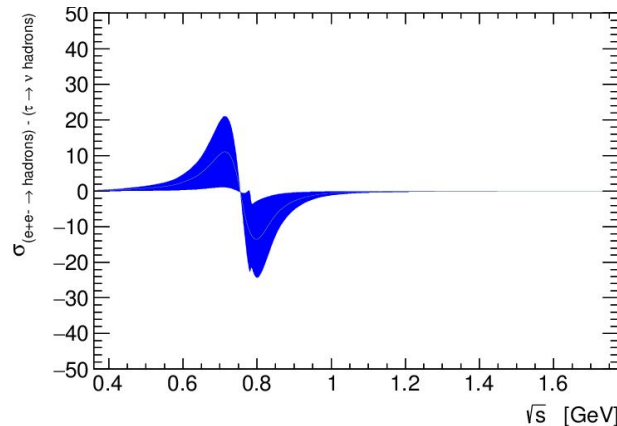
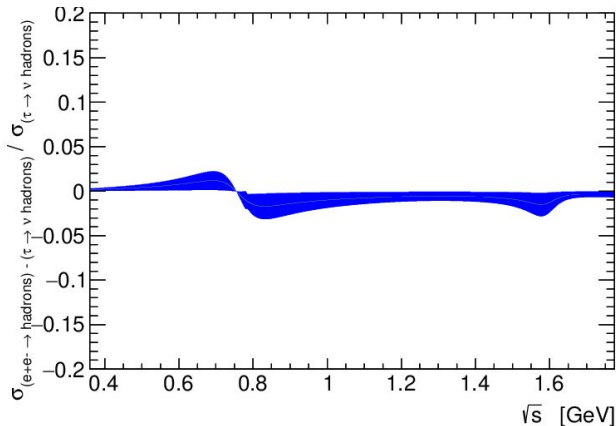
# Moment integrals from IB corrections for $\tau$ data ( $2\pi$ channel)

$a_\mu [0.36, 1.775 \text{ GeV}] = (4.11 \pm 0) \times 10^{-10} \rightarrow$  Corrections and uncertainties from  $IB$   $m\pi$  (impact on  $\rho$  width)



$$\left| \frac{F_0(s)}{F_-(s)} \right|^2$$

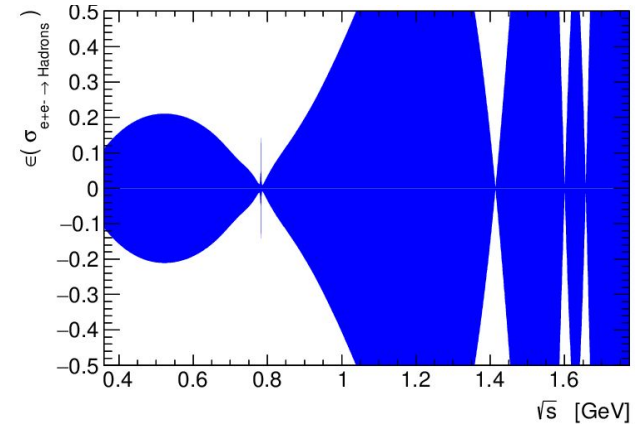
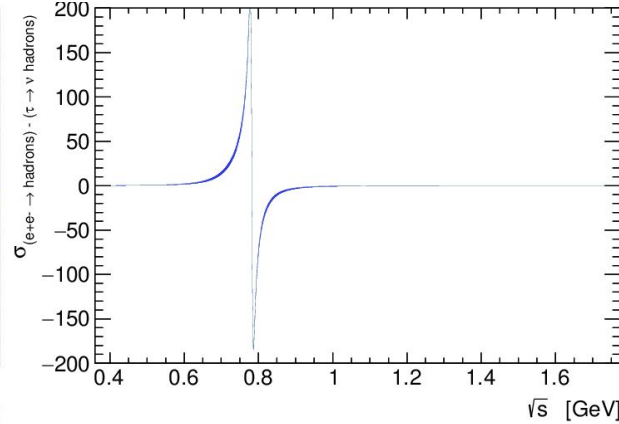
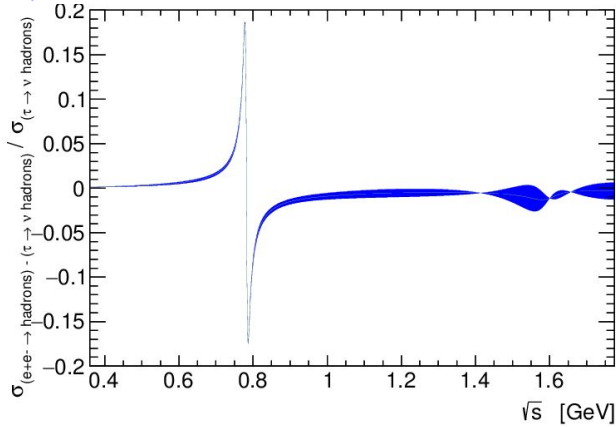
$a_\mu [0.36, 1.775 \text{ GeV}] = (0.20 \pm 0.27) \times 10^{-10} \rightarrow$  Corrections and uncertainties from  $IB$   $m\rho$



$$\left| \frac{F_0(s)}{F_-(s)} \right|^2$$

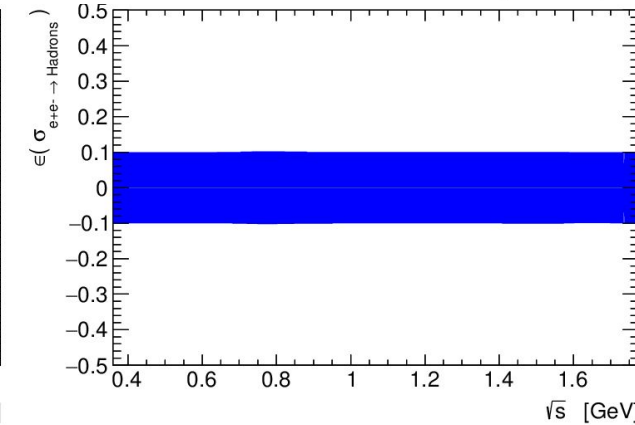
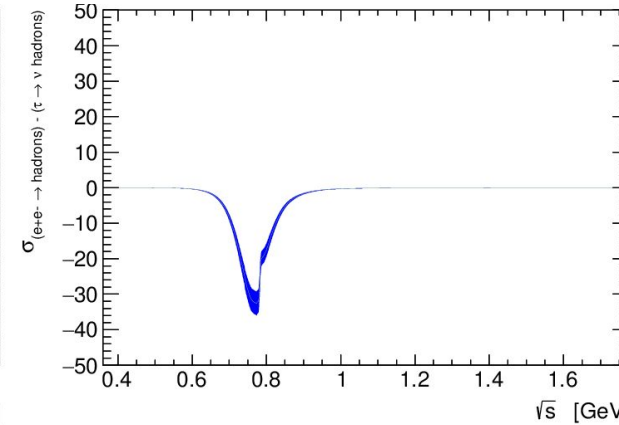
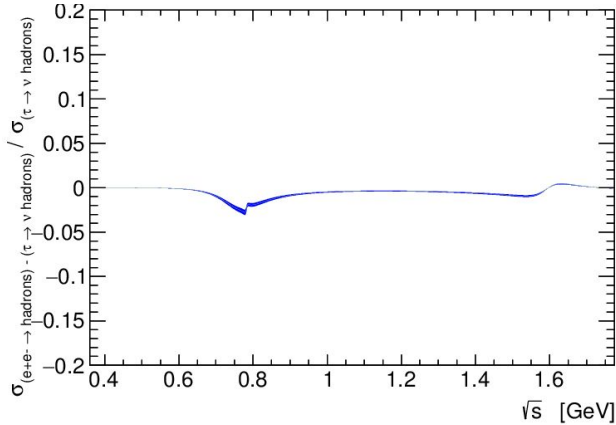
# Moment integrals from IB corrections for $\tau$ data ( $2\pi$ channel)

$a_\mu [0.36, 1.775 \text{ GeV}] = (3.99 \pm 0.98) \times 10^{-10} \rightarrow$  Corrections from *IB interference* and uncertainties from *IB KS-GS*



$a_\mu [0.36, 1.775 \text{ GeV}] = (-5.82 \pm 0.59) \times 10^{-10} \rightarrow$  *IB EM decay* corrections and uncertainties from *IB EM decay*

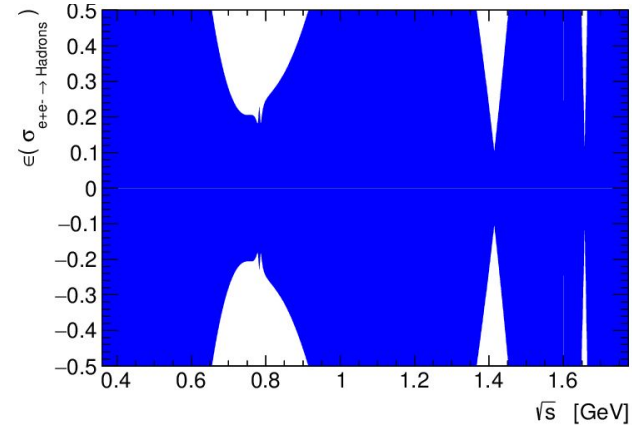
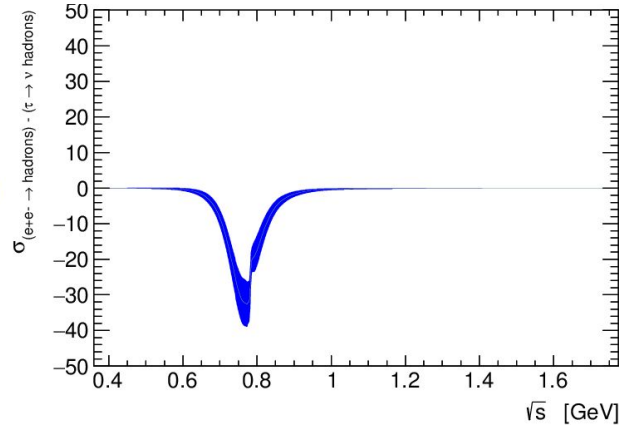
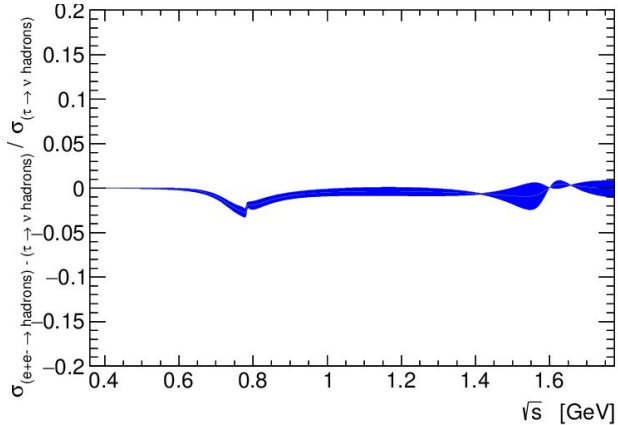
$$\left| \frac{F_0(s)}{F_-(s)} \right|^2$$



# Moment integrals from IB corrections for $\tau$ data ( $2\pi$ channel)

$$a_\mu [0.36, 1.775 \text{ GeV}] = (-5.82 \pm 1.57) \times 10^{-10}$$

→ *IB EM decay corrections and uncertainties from IB EM decay + KS-GS (conservative sum of uncertainties)*



→ Note:

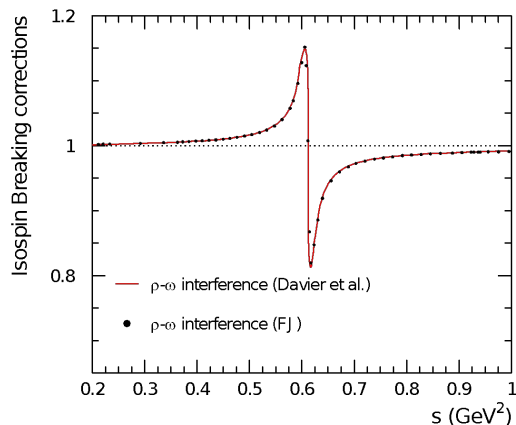
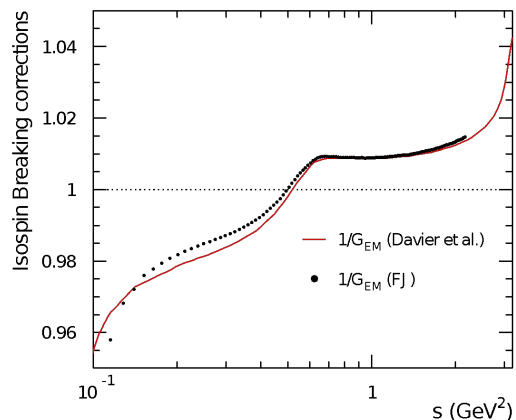
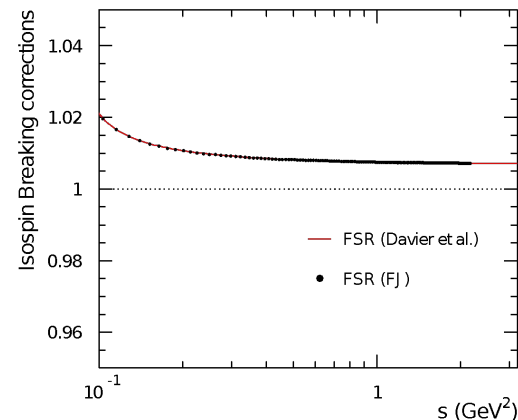
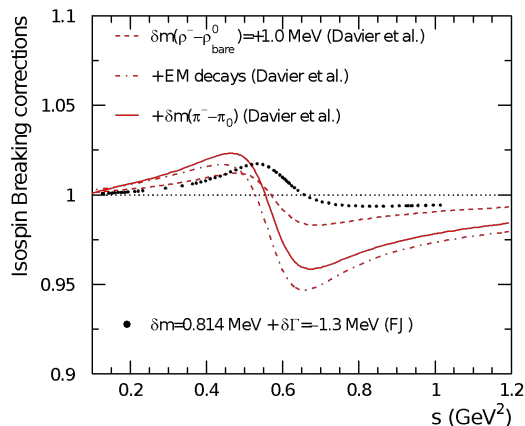
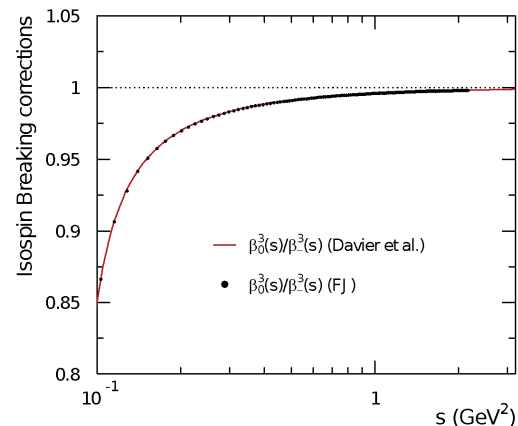
Various models for description of  $\rho$ - $\omega$  interference in IB corrections adjusted to the same  $e^+e^-$  data  
KS-GS uncertainty, using external parameters, conservatively covers this effect

# Comparison with IB-corrected $\tau$ data

$$\sigma_{e^+e^- \rightarrow \pi^+\pi^-}^{I=1} = \frac{4\pi\alpha^2}{s} v_{1,\pi^-\pi^0\nu_\tau}$$

$$v_{1,X^-}(s) = \frac{m_\tau^2}{6|V_{ud}|^2} \frac{\mathcal{B}_{X^-}}{\mathcal{B}_e} \frac{1}{N_X} \frac{dN_X}{ds} \times \left(1 - \frac{s}{m_\tau^2}\right)^2 \left(1 + \frac{2s}{m_\tau^2}\right)^{-1} \frac{R_{IB}(s)}{S_{EW}}$$

→ Comparing corrections used by Davier et al. with the ones by F. Jegerlehner

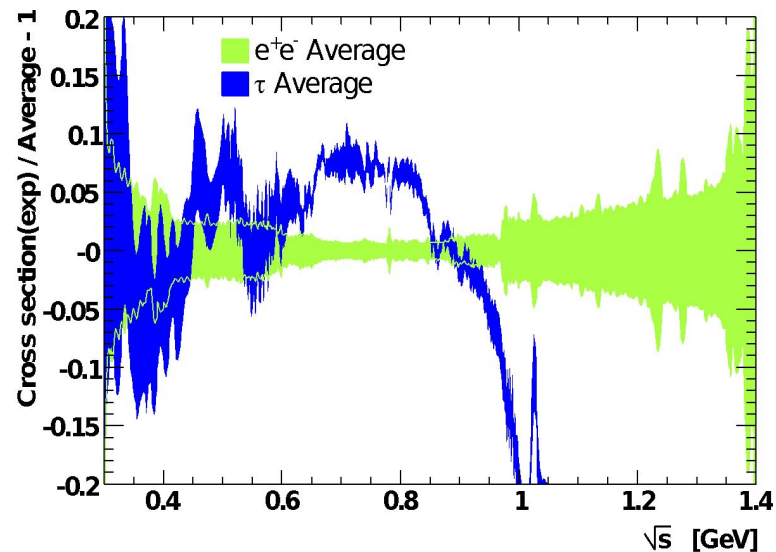
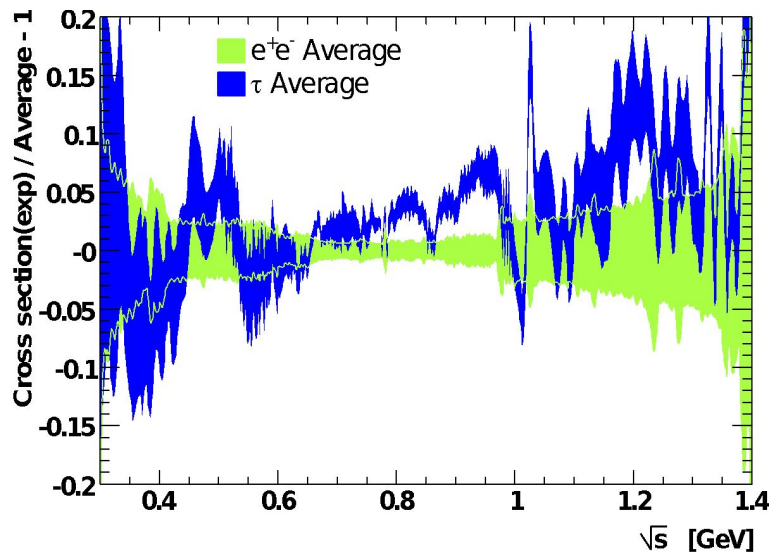
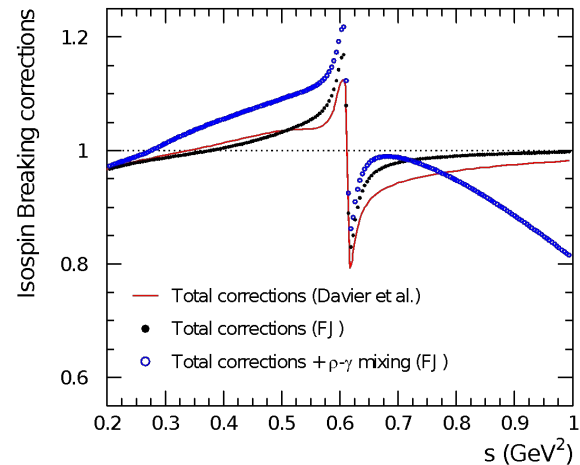


$$R_{IB}(s) = \frac{\text{FSR}(s) \beta_0^3(s)}{G_{EM}(s) \beta_-^3(s)} \left| \frac{F_0(s)}{F_-(s)} \right|^2$$

# Comparison with IB-corrected $\tau$ data

→ for a  $\mu^2$ ,  $e^+e^- - \tau$  difference of  $2.2 \sigma$   
(Davier et al.)

→ the  $\rho$ - $\gamma$  mixing correction proposed in  
arXiv:1101.2872 (FJ) seems to over-estimate  
the  $e^+e^- - \tau$  difference



# The new context for dispersive HVP since White Paper 2020 and Tau data

- At the time of WP 2020  $\Delta a_\mu^{\text{HVP LO}} (10^{-10})$ 

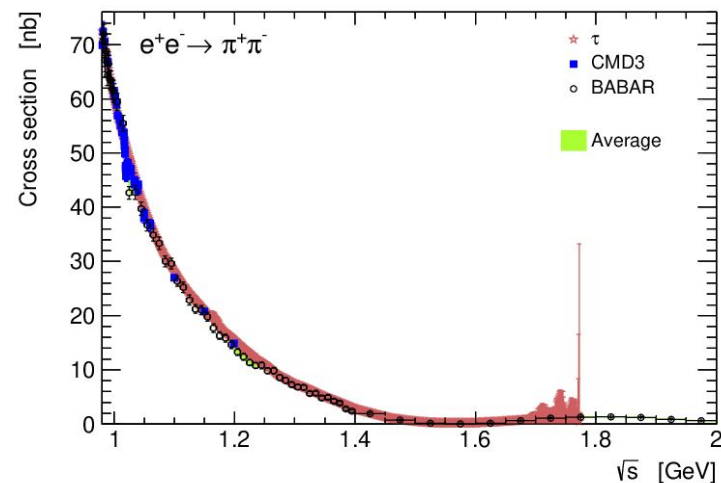
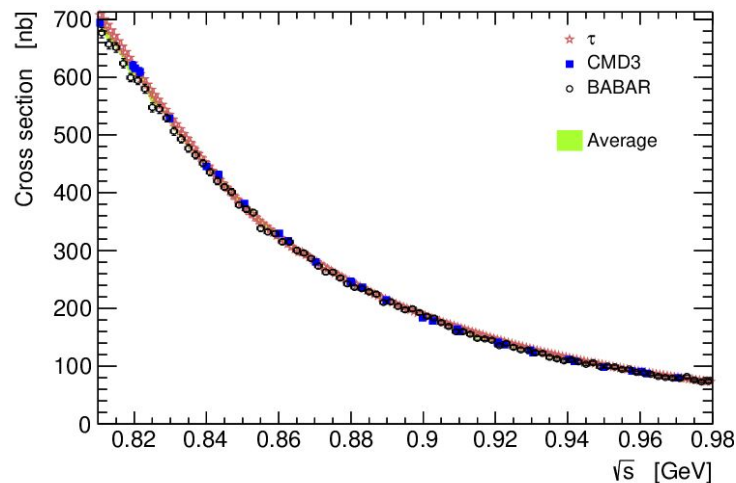
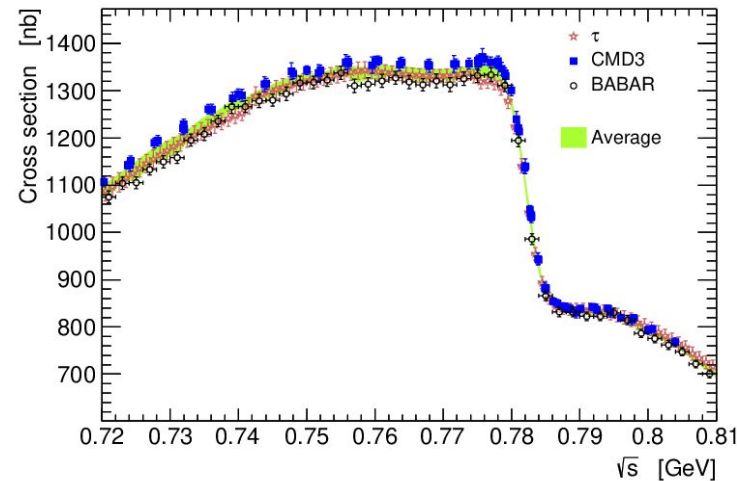
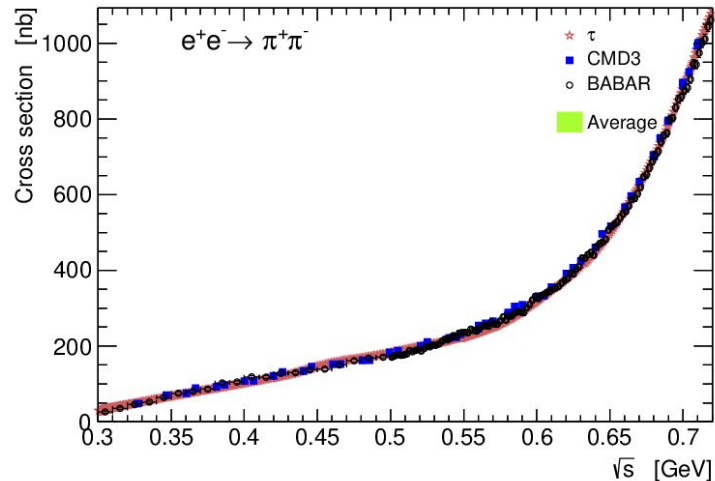
KLOE <sub>peak</sub> (0.6-0.9+comb)	2.3
BABAR	3.8
<b>BABAR – KLOE difference</b>	<b>9.8 (5.6 found with all-KLOE/all-BABAR)</b>
- Now
  - $\gamma$ - $\rho$  mixing not justified from theoretical point of view (discussions with several TI theorists)
  - CMD-3 4.2 *result changing e+e- data landscape*
  - CMD-3 – KLOE difference 21.6**
  - BABAR LO/NLO/NNLO study: *points to a necessary revisiting of KLOE analysis*
- Focusing on  $\tau$  for  $2\pi$  (competitive with best  $e^+e^- 2\pi$ ) +  $e^+e^-$  for the rest (non- $2\pi$  + I=0)

data	$1.9_{\text{spectrum}} \oplus 2.2_{\text{BR}} = 2.9$
IB correction	<b>-14.9 ± 1.9 uncertainty x11 smaller than CMD3-KLOE ≠</b>

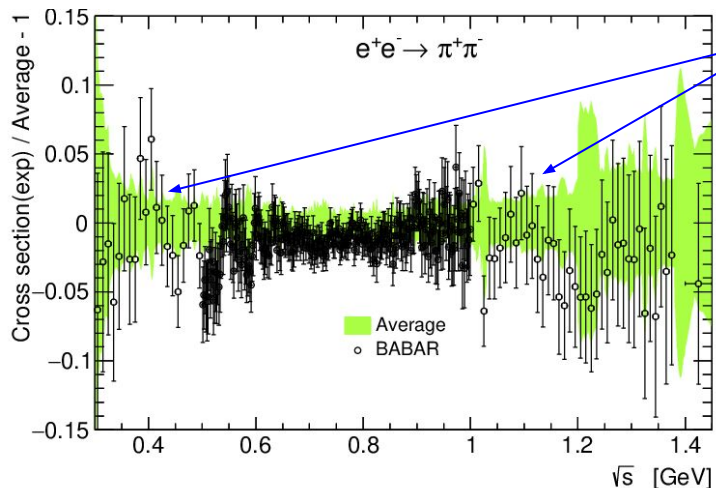


# Combining the $e^+e^- \rightarrow \pi^+\pi^-$ data, BaBar & CMD3 & Tau(+IB)

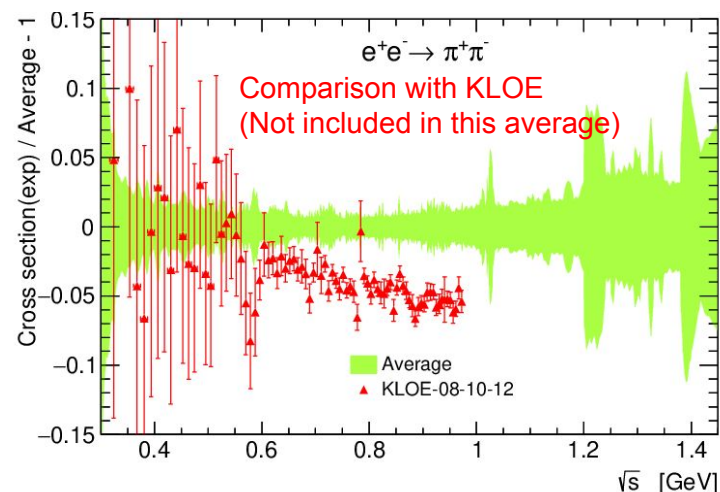
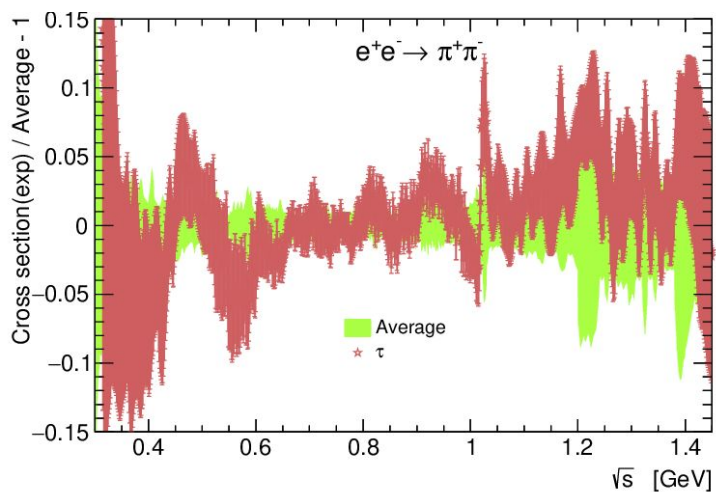
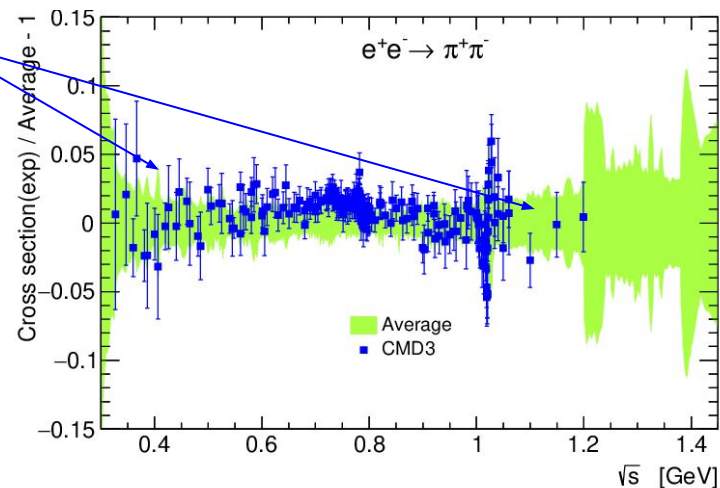
→ Motivated by the previous findings, combine  $\tau$ , BABAR and CMD-3 spectra



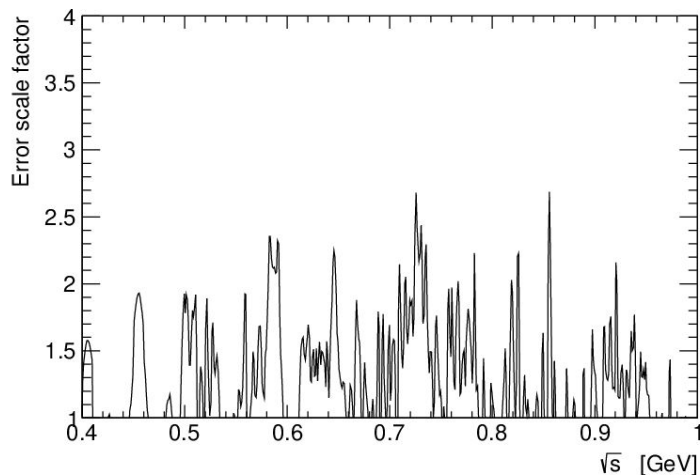
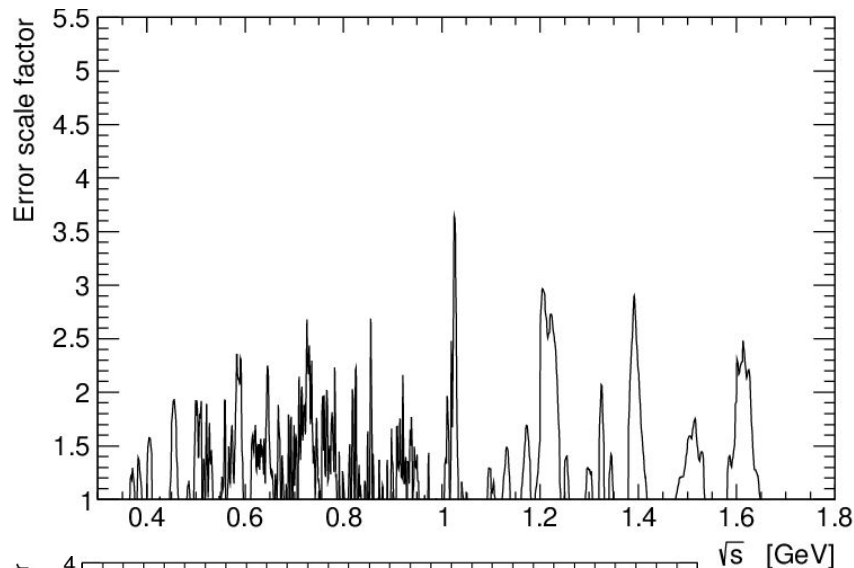
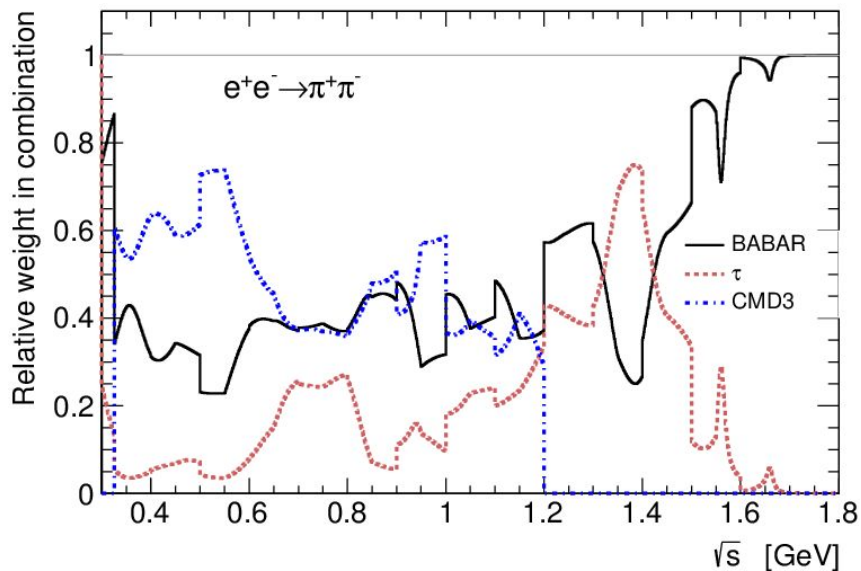
# Combining the $e^+e^- \rightarrow \pi^+\pi^-$ data, BaBar & CMD3 & Tau(+IB)



Reasonable  
BABAR/CMD3  
agreement at  
low & high E



# Combining the $e^+e^- \rightarrow \pi^+\pi^-$ data, BaBar & CMD3 & Tau(+IB)



- Average dominated by BaBar and CMD3;  
*BaBar and  $\tau$  cover full energy range*
- Some tension between BaBar & CMD3 in the  $\rho$  region
- Much larger tension (slope and shift) when comparing KLOE with the BABAR + CMD-3 +  $\tau$  combination

# Combining the $e^+e^- \rightarrow \pi^+\pi^-$ data, BaBar & CMD3 & Tau(+IB)

$\times 10^{-10}$

$$a_\mu [ 0.3 ; 1.8 \text{ GeV } ] = 519.8 \pm 3.3 ( \pm 1.3(\text{stat}) \pm 3.1 (\text{syst}) )$$

Without applying the  $\chi^2/\text{ndof}$  rescaling of uncertainties:

$$a_\mu [ 0.3 ; 1.8 \text{ GeV } ] = 519.8 \pm 2.5 ( \pm 1.0 (\text{stat}) \pm 2.3 (\text{syst}) )$$

→ Coherent with DHLMZ value ([2312.02053](#), EPJ C) obtained from average of BaBar, CMD3 and  $\tau$  integrals:

$$518.0 \pm 3.3 (\text{after uncertainty rescaling } \times 1.5)$$

(Different  $e^+e^-$  combination to complete CMD3 energy range;

Using fit of  $\tau$  data to complete their integral for [Thr.;0.36 GeV])

$$a_\mu [ 0.6 ; 0.9747 \text{ GeV } ] = 394.6 \pm 2.5 ( \pm 1.3 (\text{stat}) \pm 2.2 (\text{syst}) )$$

Without applying the  $\chi^2/\text{ndof}$  rescaling of uncertainties:

$$a_\mu [ 0.6 ; 0.9747 \text{ GeV } ] = 394.6 \pm 1.8 ( \pm 0.9(\text{stat}) \pm 1.6 (\text{syst}) )$$

$$a_\mu^{\text{win}} [ 0.3 ; 1.8 \text{ GeV } ] = 148.5 \pm 1.0 ( \pm 0.4(\text{stat}) \pm 0.9 (\text{syst}) )$$

Without applying the  $\chi^2/\text{ndof}$  rescaling of uncertainties:

$$a_\mu^{\text{win}} [ 0.3 ; 1.8 \text{ GeV } ] = 148.5 \pm 0.7 ( \pm 0.3 (\text{stat}) \pm 0.6 (\text{syst}) )$$

$$a_\mu^{\text{win}} [ 0.6 ; 0.9747 \text{ GeV } ] = 127.4 \pm 0.8 ( \pm 0.4 (\text{stat}) \pm 0.7 (\text{syst}) )$$

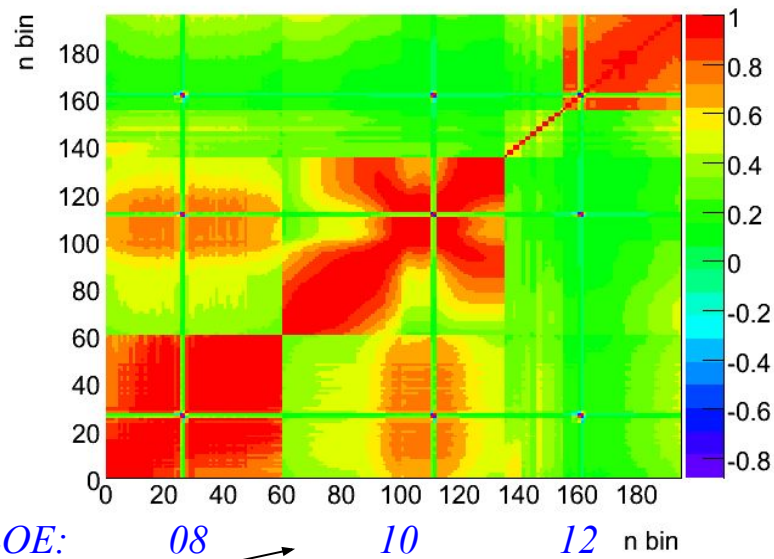
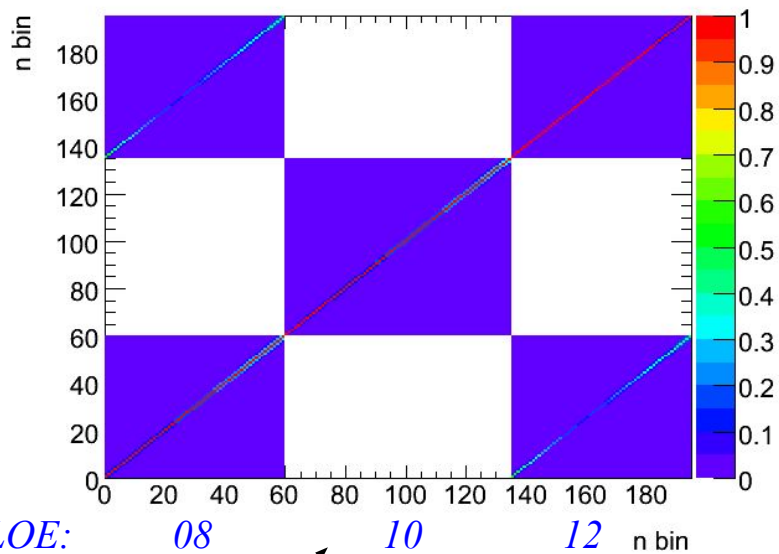
Without applying the  $\chi^2/\text{ndof}$  rescaling of uncertainties:

$$a_\mu^{\text{win}} [ 0.6 ; 0.9747 \text{ GeV } ] = 127.4 \pm 0.6 ( \pm 0.3(\text{stat}) \pm 0.5 (\text{syst}) )$$

→ Still non-negligible effect of uncertainty enhancement through the local  $\chi^2/\text{ndof}$  rescaling;

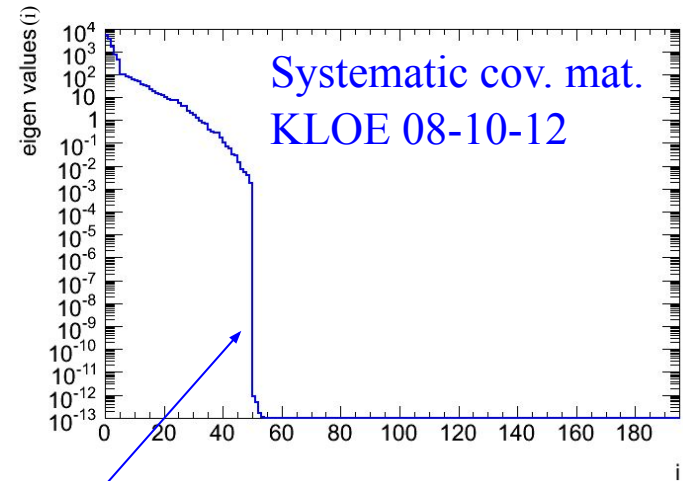
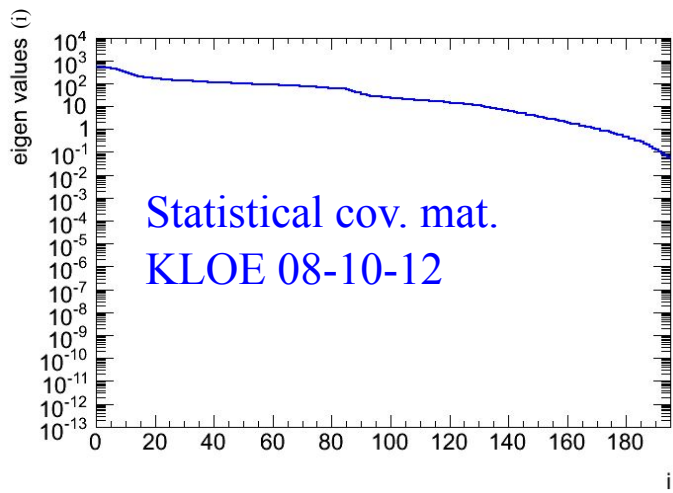
In addition, an extra uncertainty accounting for systematic deviations between measurements has to be added, as done for DHMZ'19

# Treatment of the KLOE correlation matrices



→ Statistical and systematic correlation matrices among the 3 measurements

# Treatment of the KLOE data – eigenvector decomposition

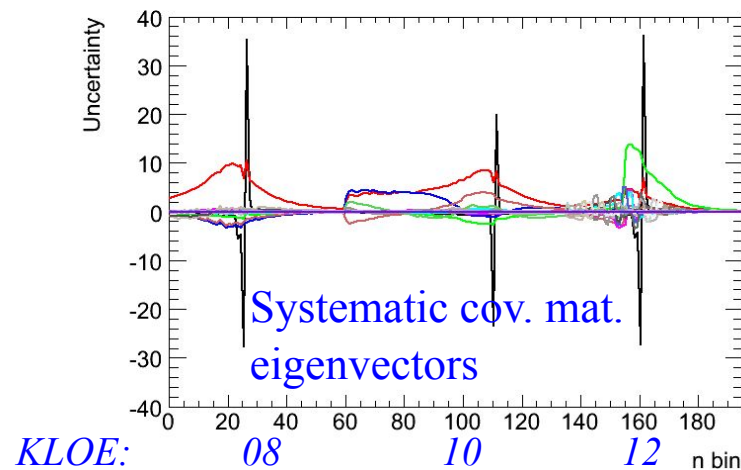
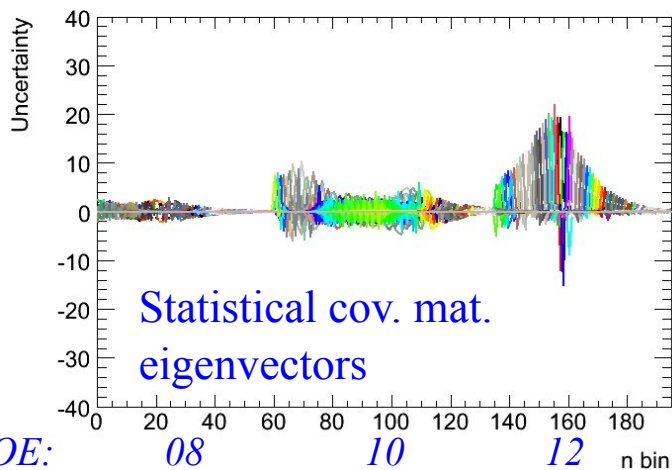


→ “counting” the number of independent components (50) used to build the covariance matrix

$$C = S \cdot D \cdot S^T$$
$$D = \begin{pmatrix} \diagdown & 0 & 0 \\ 0 & \sigma_i^2 & 0 \\ 0 & 0 & \diagdown \end{pmatrix}$$
$$S = \begin{pmatrix} V_1 & \dots & V_n \\ \vdots & & \vdots \end{pmatrix}$$

→ Problem of negative eigenvalues for previous systematic covariance matrix solved (informed KLOE collaboration about the problem in summer 2016)

# Treatment of the KLOE data – eigenvector decomposition

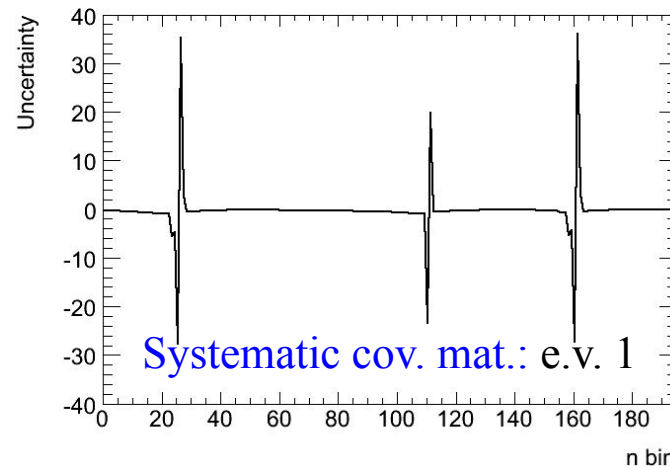
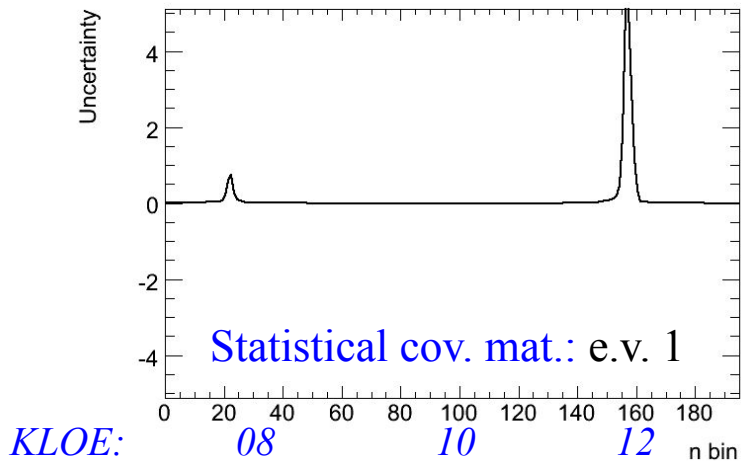


- Each normalized eigenvector ( $\sigma_i \cdot V_i$ ) treated as an uncertainty fully correlated between the bins
- All these uncertainties are independent between each-other

$$C = \sum_{i=1}^{N_{bins}} \sigma_i^2 \cdot C(V_i)$$

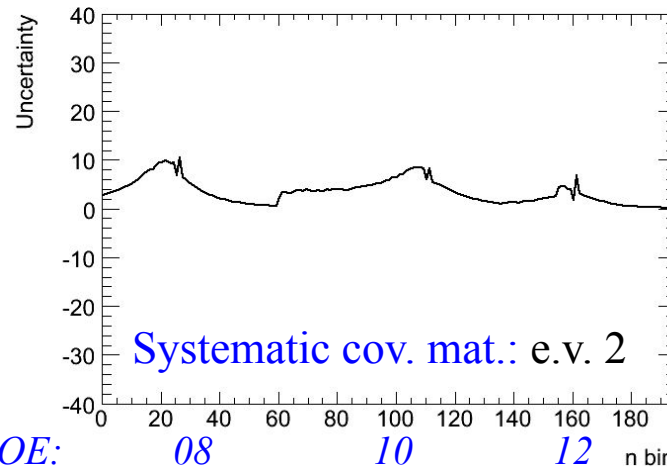
- Checked exact matching with the original matrices + with all  $a_\mu$  integrals and uncertainties published by KLOE

# Treatment of the KLOE data – eigenvector decomposition



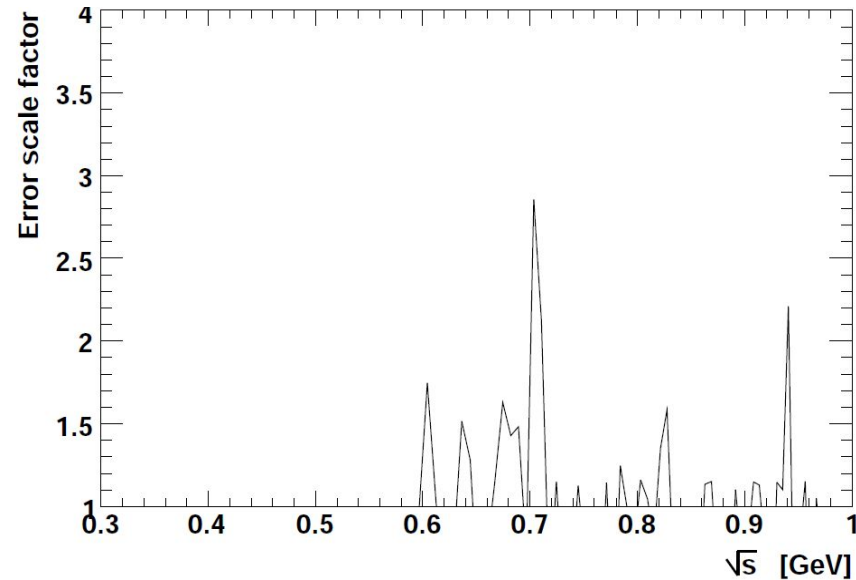
→ Eigenvectors carry the general features of the correlations:

- long-range for systematics
- ~short-range for statistical uncertainties + correlations between KLOE 08 & 12





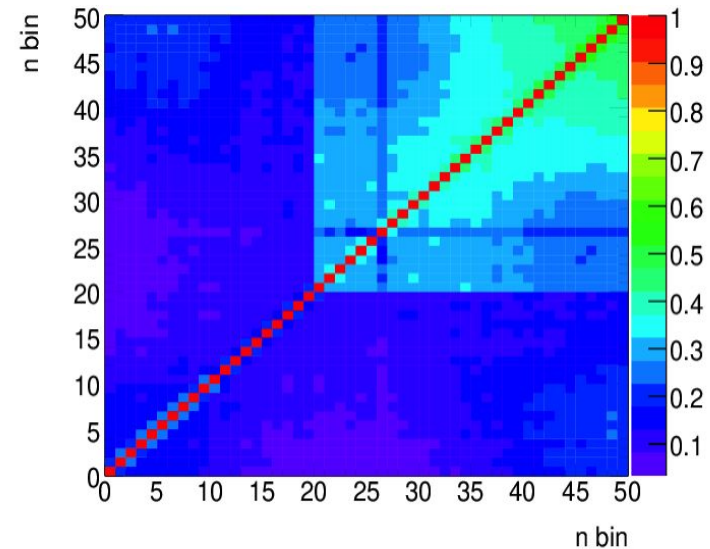
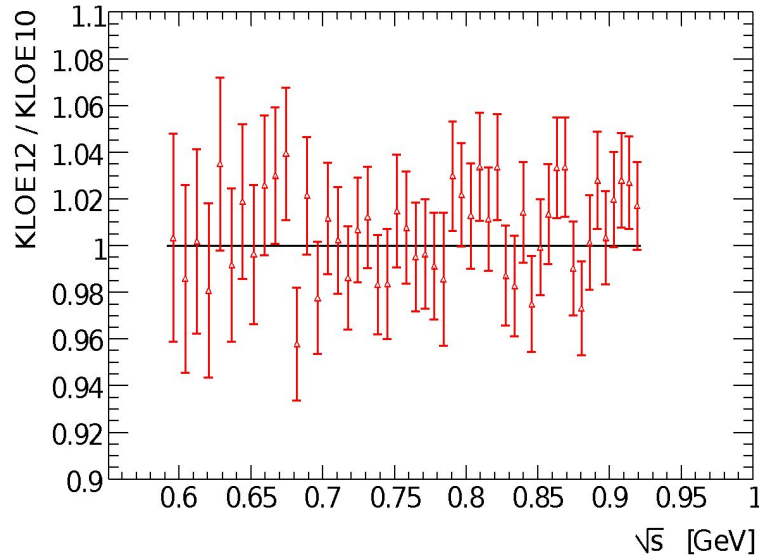
# Local comparison of the 3 KLOE measurements



- Local  $\chi^2$ /ndof test of the local compatibility between KLOE 08 & 10 & 12, taking into account the correlations: some tensions observed
- Does not probe general trends of the difference between the measurements (e.g. slopes in the ratio)

# Ratios between measurements

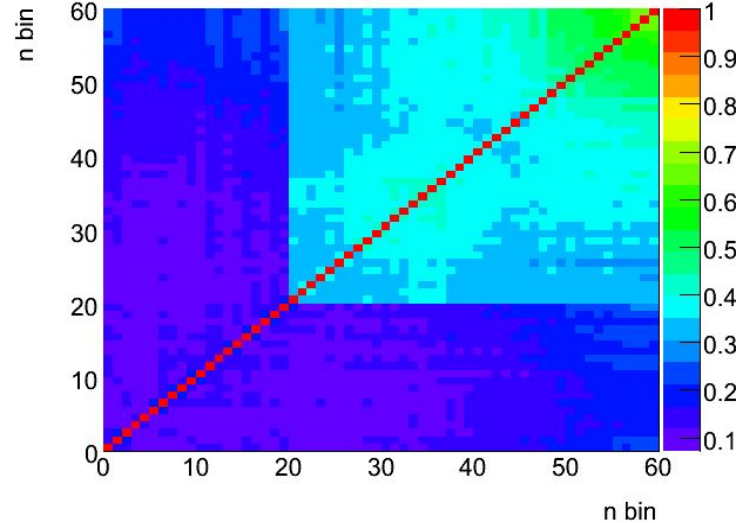
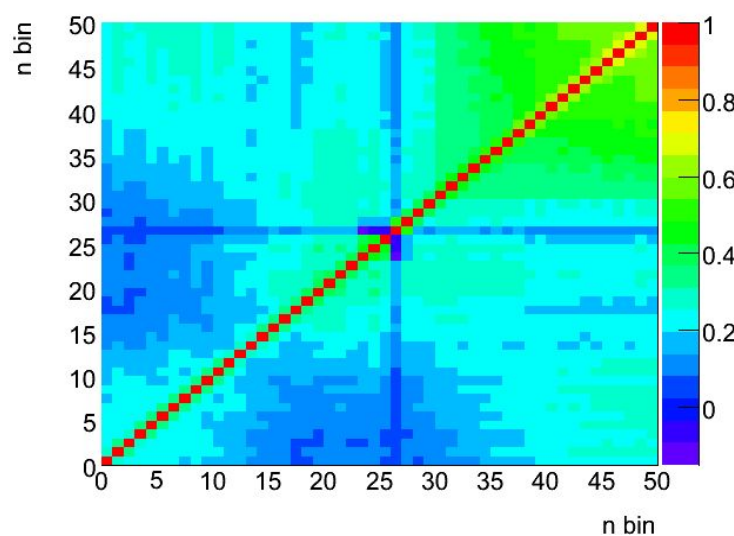
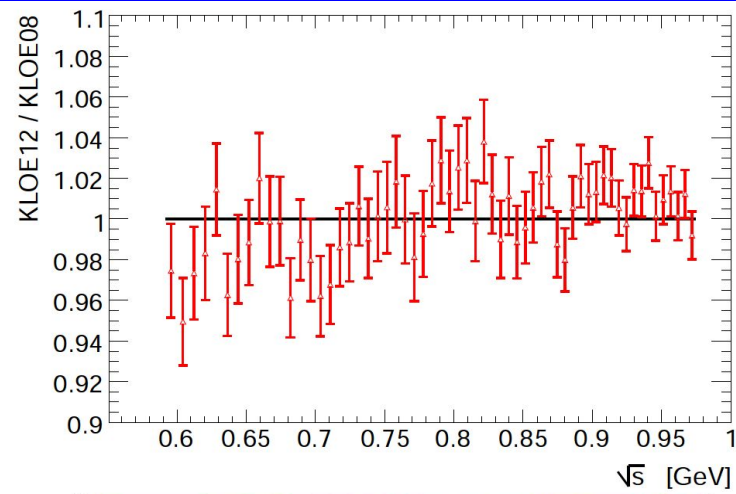
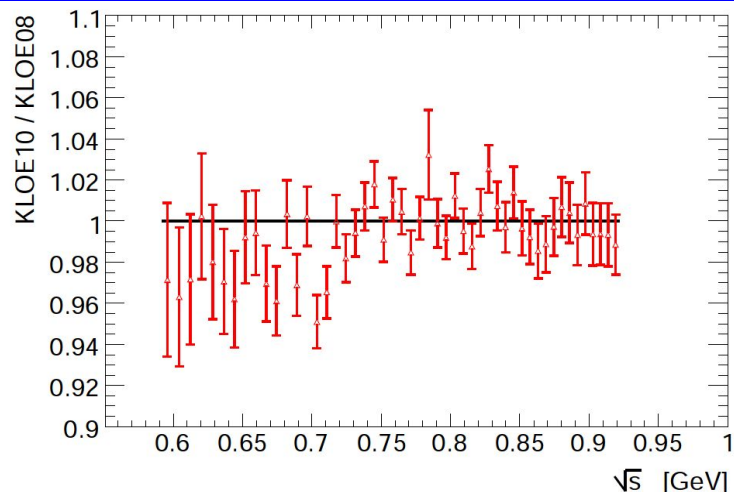
- Compute ratio between pairs of KLOE measurements
- Full propagation of uncertainties and correlations using pseudo-experiments (agreement with analytical linear uncertainty propagation)



- Good agreement between KLOE 10 and KLOE 12

[Back](#)

# Ratios between measurements



# Direct comparison of the 3 KLOE measurements

→ Quantitative comparison between the ratios and unity, taking into account correlations

## KLOE 10 / KLOE 08

$\chi^2 [0.35;0.85] \text{ GeV}^2 : 79.0 / 50(\text{DOF})$   
p-value= 0.0056

$\chi^2 [0.35;0.58] \text{ GeV}^2 : 46.2 / 23(\text{DOF})$   
p-value= 0.0028

$\chi^2 [0.58;0.85] \text{ GeV}^2 : 29.7 / 27(\text{DOF})$   
p-value= 0.33

$\chi^2 [0.64;0.85] \text{ GeV}^2 : 20.7 / 21(\text{DOF})$   
p-value= 0.47

## KLOE 12 / KLOE 08

$\chi^2 [0.35;0.95] \text{ GeV}^2 : 73.7 / 60(\text{DOF})$   
p-value= 0.11

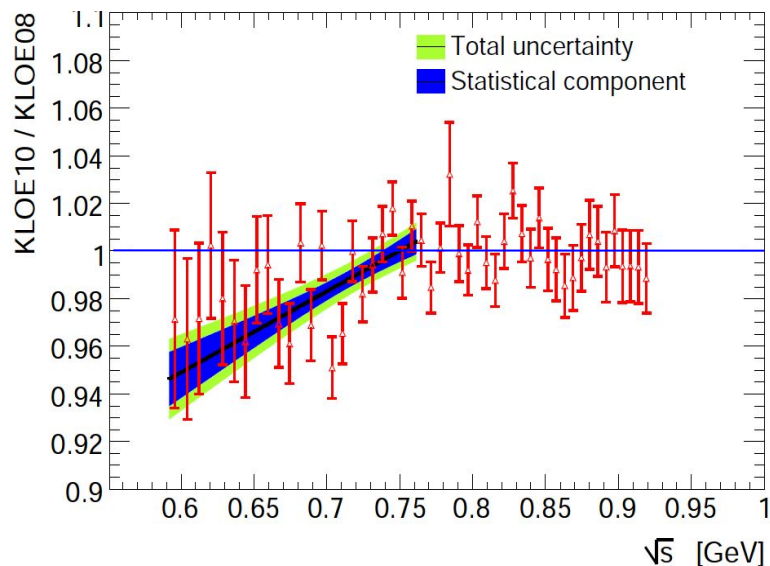
$\chi^2 [0.35;0.58] \text{ GeV}^2 : 21.8 / 23(\text{DOF})$   
p-value= 0.53

$\chi^2 [0.35;0.64] \text{ GeV}^2 : 27.5 / 29(\text{DOF})$   
p-value= 0.55

$\chi^2 [0.64;0.95] \text{ GeV}^2 : 39.4 / 31(\text{DOF})$   
p-value= 0.14

# Quantitative comparisons of the KLOE measurements

- Quantitative comparison between the ratios and unity, taking into account correlations
- Fitting the ratio taking into account correlations
- Full propagation of uncertainties and correlations – 3 methods yielding consistent results:  $\pm 1\sigma$  shifts of each uncertainty, pseudo-experiments and fit uncertainties from Minuit



Comparison with Unity:

$\chi^2 [0.35;0.85] \text{ GeV}^2 : 79.0 / 50(\text{DOF})$

p-value= 0.0056

$\chi^2 [0.35;0.58] \text{ GeV}^2 : 46.2 / 23(\text{DOF})$

p-value= 0.0028

$\chi^2 [p_0 + p_1\sqrt{s}] : 36.1 / 21(\text{DOF})$

p-value= 0.02

$p_0 : 0.745 \pm 0.085$

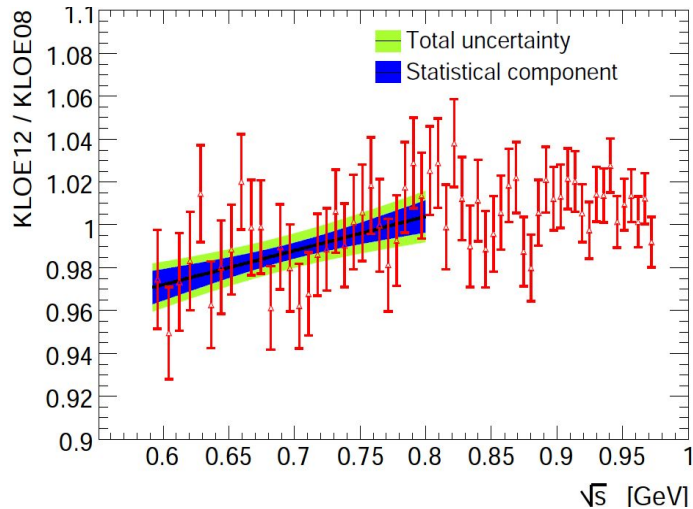
$p_1 : 0.341 \pm 0.117$

- Significant shift & slope ( $\sim 2.5\text{-}3\sigma$ ) at low  $\sqrt{s}$ , no significant shift at high  $\sqrt{s}$   
Similar shift & slope for KLOE 12 / KLOE 08 (*see below*)
- Should motivate conservative treatment of uncertainties and correlations in combination

# Direct comparison of the 3 KLOE measurements

→ Fitting the ratio taking into account correlations

→ Full propagation of uncertainties and correlations – 3 methods yielding consistent results:  
 $\pm 1\sigma$  shifts of each uncertainty, pseudo-experiments and fit uncertainties from Minuit

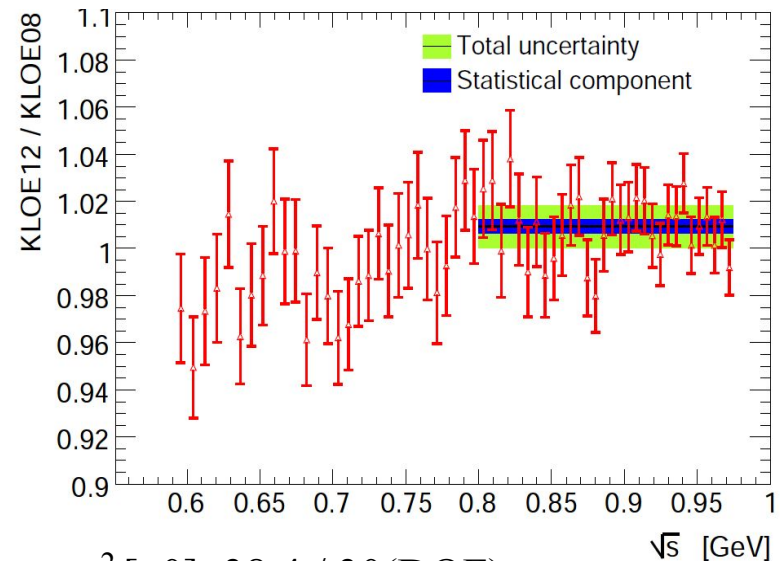


$$\chi^2 [p0 + p1\sqrt{s}]: 20.7 / 27(\text{DOF})$$

$$p\text{-value} = 0.80$$

$$p0 : 0.876 \pm 0.056$$

$$p1 : 0.159 \pm 0.081$$



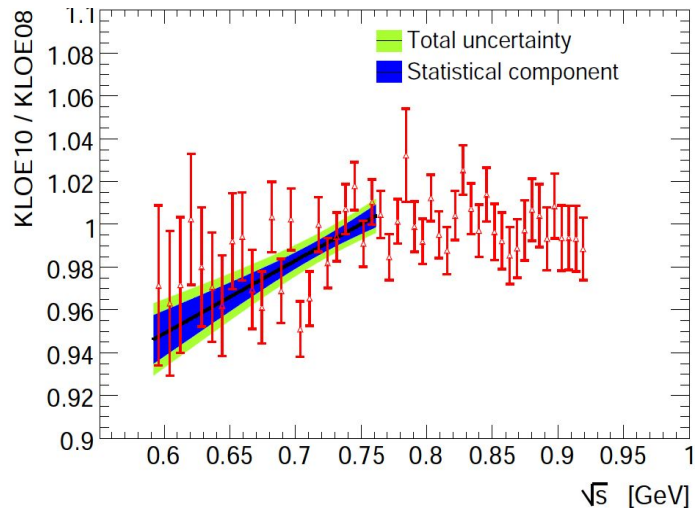
$$\chi^2 [p0]: 38.4 / 30(\text{DOF})$$

$$p\text{-value} = 0.14$$

$$p0 : 1.009 \pm 0.009$$

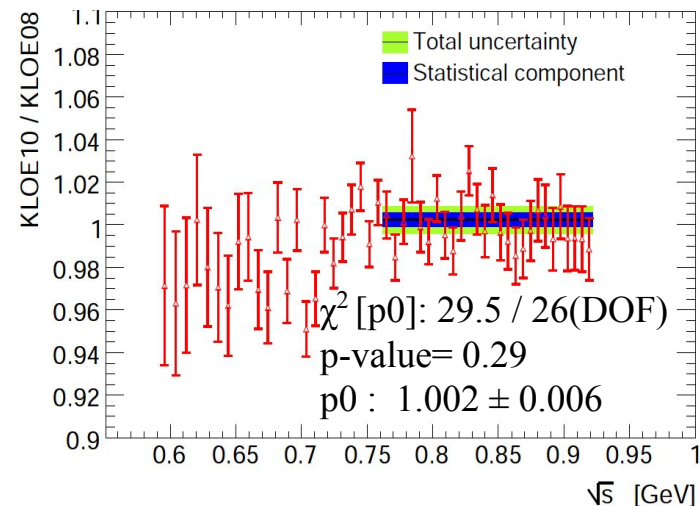
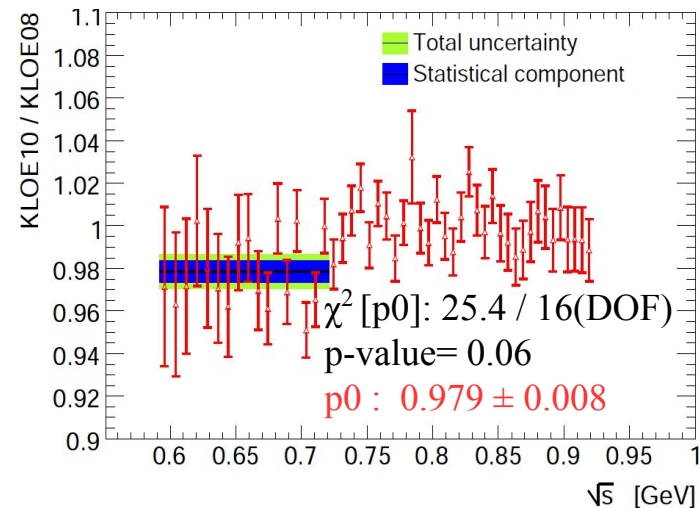
→ Significant shift and slope ( $\sim 2\sigma$ ) at low  $\sqrt{s}$ , no significant shift at high  $\sqrt{s}$

# Direct comparison of the 3 KLOE measurements

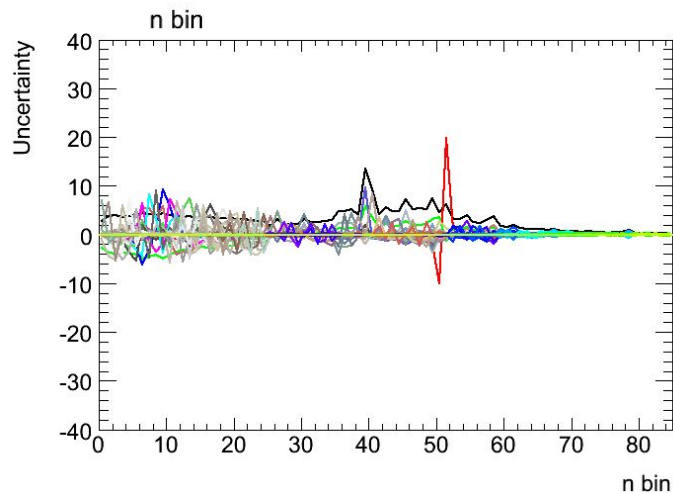
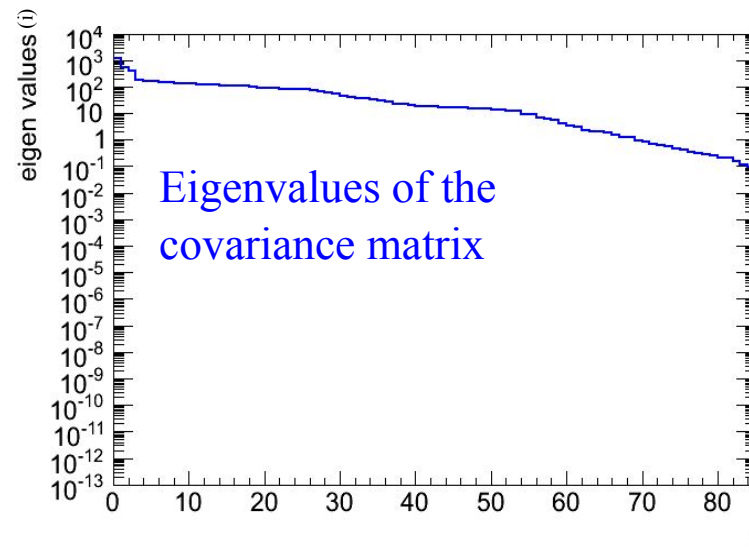
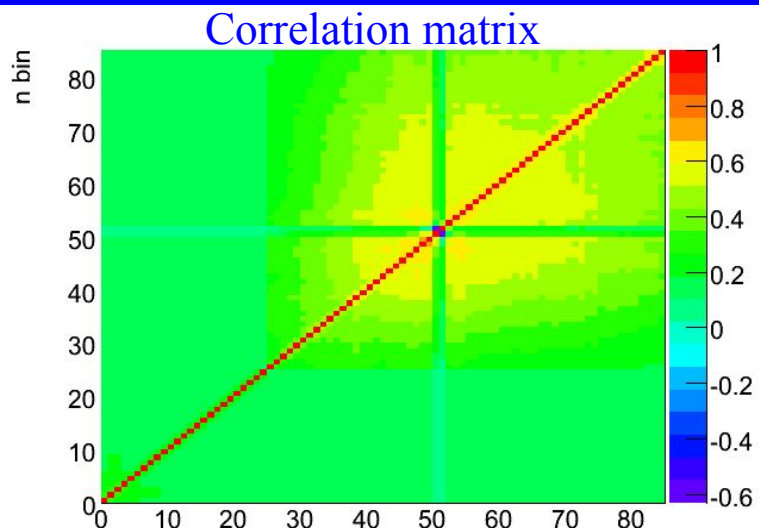


$\chi^2 [p0 + p1\sqrt{s}]$ : 36.1 / 21(DOF)  
 p-value= 0.02  
 $p0 : 0.745 \pm 0.085$   
 $p1 : 0.341 \pm 0.117$

→ Significant shift and slope ( $\sim 2.5-3\sigma$ ) at low  $\sqrt{s}$ ,  
 no significant shift at high  $\sqrt{s}$

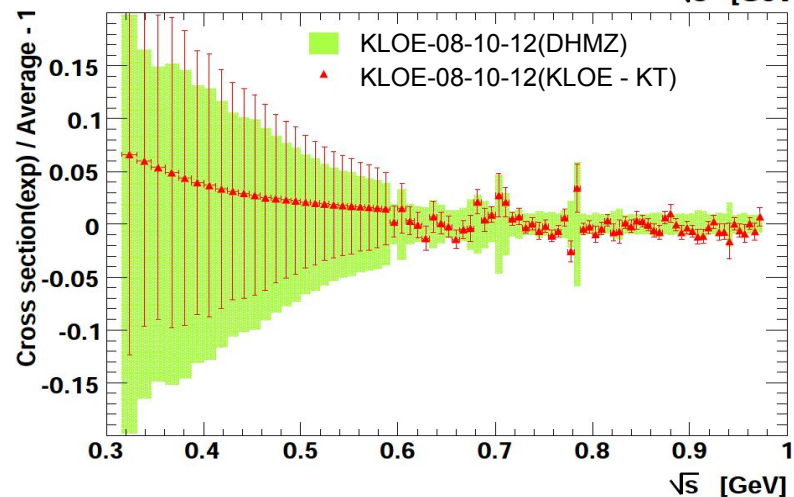
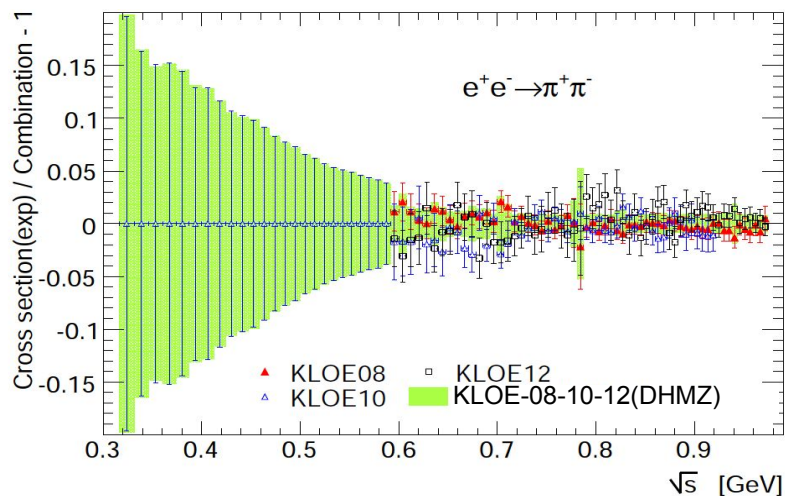
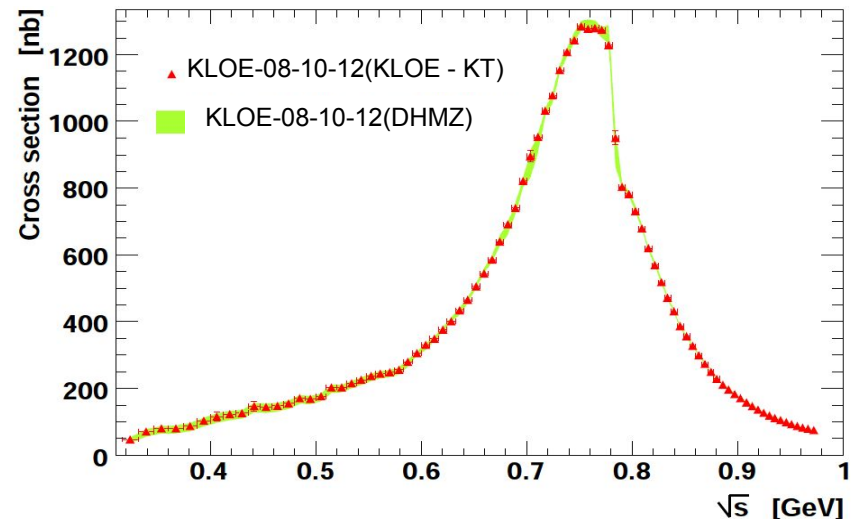
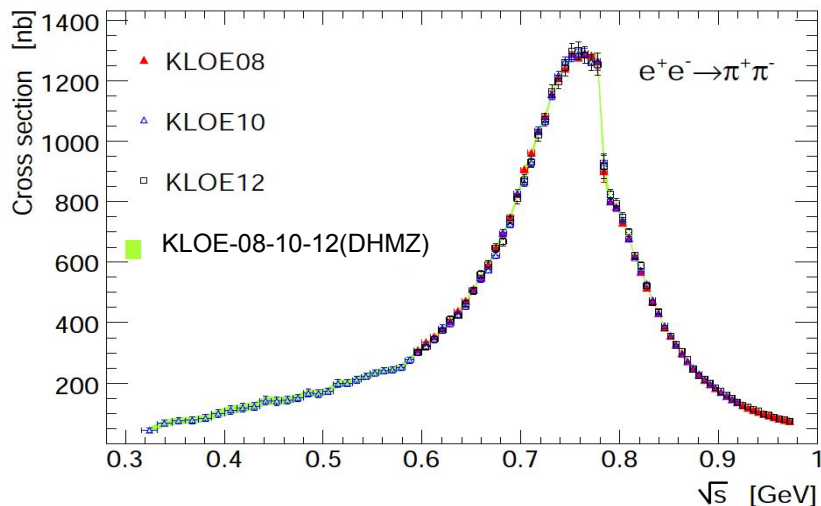


# Treatment of the combined KLOE data





# Combining the 3 KLOE measurements



# Comparison of / consequences for combination methods

Analysis aspect	DHMZ	KNT
Blinding	Not necessary (No ad-hoc choices to make)	Included for upcoming update
Binning	Fine ( $\leq 1$ MeV) final binning for average and integrals. Large ( $O(100)$ MeV) or less) common binning @ intermediate step: compare statistics of experiments coherently for deriving weights in fine bins.	Re-bin data into "clusters". Scans over cluster configurations for optimisation.
Closure test	Using model for spectrum: negligible bias. (since 2009)	Not performed
Additional constraints	Analyticity constraints for $2\pi$ channel.	None
Fitting	$\chi^2$ minimisation with correlated uncertainties incorporated locally (in fine & large bins), for deriving weights. Full propagation of uncertainties & correlations.	$\chi^2$ minimisation with correlated uncertainties incorporated globally.
Integration / interpolation	Av. of quadratic splines (3 <sup>rd</sup> order polynomial), integral preservation in bins of measurements. Analyticity-based function for $2\pi$ ( $< 0.6$ GeV).	Trapezoidal for continuum, quintic for resonances.
Uncertainty inflation	Local $\chi^2$ uncertainty inflation. (since 2009) Extra BABAR-KLOE systematic. (since 2019)	Local $\chi^2$ uncertainty inflation. (adopted since 2017)
Inter-channel correlations	Taken into account. (since 2010)	Not included.
Missing channels	Estimated based on isospin symmetry. (since 1997 - ADH)	Adopted in subsequent updates

→ Large DHMZ/KNT differences for the resulting uncertainties, as well as for the shapes of the combined spectra

→ CHS approach for  $2\pi$  and  $3\pi$ : Analyticity and global  $\chi^2$  fit

WP TI	DHMZ19	KNT19
$a_\mu^{\text{HVP,LO}} \times 10^{10}$	694.0(4.0)	692.8(2.4)

# $a_{\mu}^{\pi\pi}$ contribution [0.28; 1.8] GeV – spline-based (2018)

→ Updated result:

$$506.70 \pm 2.32 ( \pm 1.01 \text{ (stat.)} \pm 2.08 \text{ (syst.)} ) [10^{-10}]$$

(after uncertainty enhancement by  $\sim 14\%$  caused by the tension between inputs, taken into account through a local rescaling)

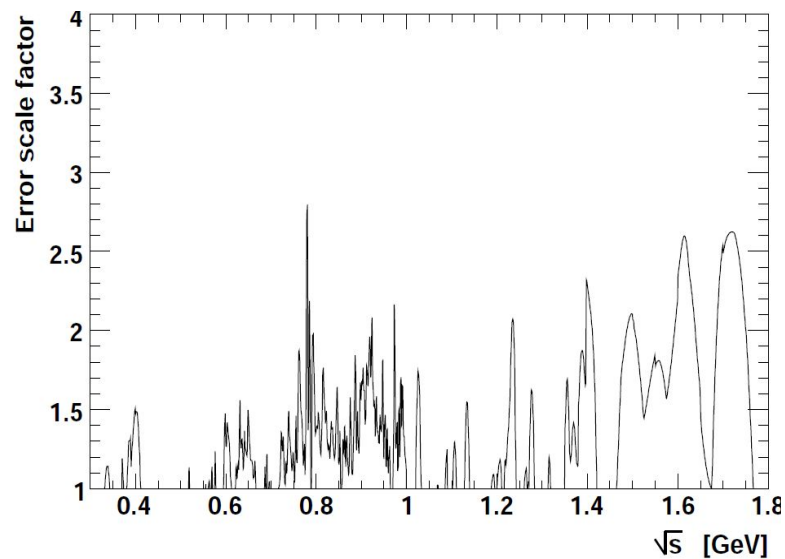
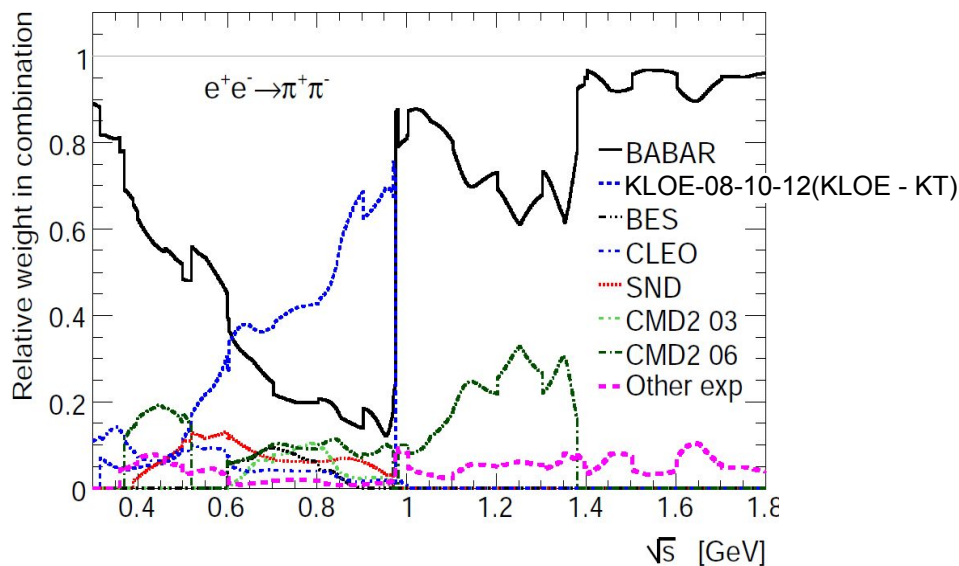
Total uncertainty: 5.9 (2003) → 2.8 (2011) → 2.6 (2017) → 2.3 (2018)

# $a_{\mu}^{\pi\pi}$ contribution [0.28; 1.8] GeV – spline-based (2018)

→ with KLOE-08-10-12 (KLOE-KT) used as input:  $506.55 \pm 2.38 [10^{-10}]$

(after uncertainty enhancement by 18% caused by the tension between inputs, taken into account through a local rescaling)

→ Compensation between uncertainty reduction for KLOE-08-10-12 (KLOE-KT), inducing a change of weights in DHMZ combination, and tension enhancement



# Uncertainties on uncertainties and on correlations

*Topic of general interest, in other fields too*

[1908.00921](#)(DHMZ), [2006.04822](#)(WP Theory Initiative)

[Back](#)

# Two different approaches for combining ( $e^+e^-$ ) data

DHMZ:

- $\chi^2$  computed locally (in each fine bin), taking into account correlations between measurements (see previous slides)
- Used to determine the weights on the measurements in the combination and their level of agreement
- Uncertainties and correlations propagated using pseudo-experiments or  $\pm 1\sigma$  shifts of each uncertainty component

KNT:

- $\chi^2$  computed globally (for full mass range)

$$\chi_I^2 = \sum_{i=1}^{N_{\text{tot}}} \sum_{j=1}^{N_{\text{tot}}} (R_i^{(m)} - \mathcal{R}_m^{i,I}) \mathbf{C}_I^{-1}(i^{(m)}, j^{(n)}) (R_j^{(n)} - \mathcal{R}_n^{j,I}) \quad \text{KNT (1802.02995)}$$

$$\chi^2 = \sum_{i=1}^{195} \sum_{j=1}^{195} (\sigma_{\pi\pi(\gamma)}^0(i) - \bar{\sigma}_{\pi\pi(\gamma)}^0(m)) \mathbf{C}^{-1}(i^{(m)}, j^{(n)}) (\sigma_{\pi\pi(\gamma)}^0(j) - \bar{\sigma}_{\pi\pi(\gamma)}^0(n)) \quad \text{KLOE-KMT (1711.03085)}$$

- relies on description of correlations on long ranges

- *One of the main sources of differences for the uncertainty on  $a_\mu$*

# Evaluation of uncertainties and correlations ( $e^+e^-$ )

	$\sigma_{\pi\pi\gamma}$	$\sigma_{\pi\pi}^0$	$F_\pi$	$\Delta^{\pi\pi}a_\mu$
Reconstruction Filter	negligible			
Background subtraction	Tab. 1		0.3%	
Trackmass	0.2%			
Pion cluster ID	negligible			
Tracking efficiency	0.3%			
Trigger efficiency	0.1%			
Acceptance	Tab. 2		0.2%	
Unfolding	Tab. 3		negligible	
L3 filter	0.1%			
$\sqrt{s}$ dependence of $H$	-	Tab. 4		0.2%
Luminosity	0.3%			
Experimental systematics	0.6%			
FSR resummation	-	0.3%		
Radiator function $H$	-	0.5%		
Vacuum Polarization	-	0.1%	-	0.1%
Theory systematics	0.6%			

→ Systematics *evaluated* in  $\sim$ wide mass ranges with sharp transitions

$M_{\pi\pi}^2$ range (GeV <sup>2</sup> )	Systematic error (%)
$0.35 \leq M_{\pi\pi}^2 < 0.39$	0.6
$0.39 \leq M_{\pi\pi}^2 < 0.43$	0.5
$0.43 \leq M_{\pi\pi}^2 < 0.45$	0.4
$0.45 \leq M_{\pi\pi}^2 < 0.49$	0.3
$0.49 \leq M_{\pi\pi}^2 < 0.51$	0.2
$0.51 \leq M_{\pi\pi}^2 < 0.64$	0.1
$0.64 \leq M_{\pi\pi}^2 < 0.95$	-

KLOE 08 (0809.3950)

KLOE 10 (1006.5313)

	$\sigma_{\pi\pi\gamma}$	$\sigma_{\pi\pi}^{\text{bare}}$	$ F_\pi ^2$	$\Delta a_\mu^{\pi\pi}$
	threshold ; $\rho$ -peak			(0.1 - 0.85 GeV <sup>2</sup> )
Background Filter	0.5% ; 0.1%			negligible
Background subtraction	3.4% ; 0.1%			0.5%
$f_0 + \rho\pi$ bkg.	6.5% ; negl.			0.4%
$\Omega$ cut	1.4% ; negl.			0.2%
Trackmass cut	3.0% ; 0.2%			0.5%
$\pi$ -e PID	0.3% ; negl.			negligible
Trigger	0.3% ; 0.2%			0.2%
Acceptance	1.9% ; 0.3%			0.5%
Unfolding	negl. ; 2.0%			negligible
Tracking	0.3%			
Software Trigger (L3)	0.1%			
Luminosity	0.3%			
Experimental syst.				1.0%
FSR treatment	-	7% ; negl.		0.8%
Radiator function $H$	-	0.5%		
Vacuum Polarization	-	Ref. 34	-	0.1%
Theory syst.				0.9%

# Evaluation of uncertainties and correlations ( $e^+e^-$ )

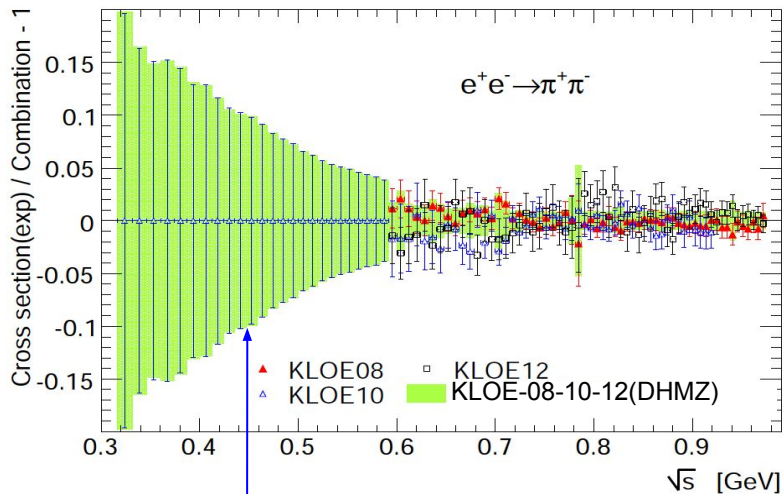
Sources	0.3-0.4	0.4-0.5	0.5-0.6	0.6-0.9	0.9-1.2	1.2-1.4	1.4-2.0	2.0-3.0
trigger/ filter	5.3	2.7	1.9	1.0	0.7	0.6	0.4	0.4
tracking	3.8	2.1	2.1	1.1	1.7	3.1	3.1	3.1
$\pi$ -ID	10.1	2.5	6.2	2.4	4.2	10.1	10.1	10.1
background	3.5	4.3	5.2	1.0	3.0	7.0	12.0	50.0
acceptance	1.6	1.6	1.0	1.0	1.6	1.6	1.6	1.6
kinematic fit ( $\chi^2$ )	0.9	0.9	0.3	0.3	0.9	0.9	0.9	0.9
correl $\mu\mu$ ID loss	3.0	2.0	3.0	1.3	2.0	3.0	10.0	10.0
$\pi\pi/\mu\mu$ non-cancel.	2.7	1.4	1.6	1.1	1.3	2.7	5.1	5.1
unfolding	1.0	2.7	2.7	1.0	1.3	1.0	1.0	1.0
ISR luminosity	3.4	3.4	3.4	3.4	3.4	3.4	3.4	3.4
sum (cross section)	13.8	8.1	10.2	5.0	6.5	13.9	19.8	52.4

BABAR (1205.2228)

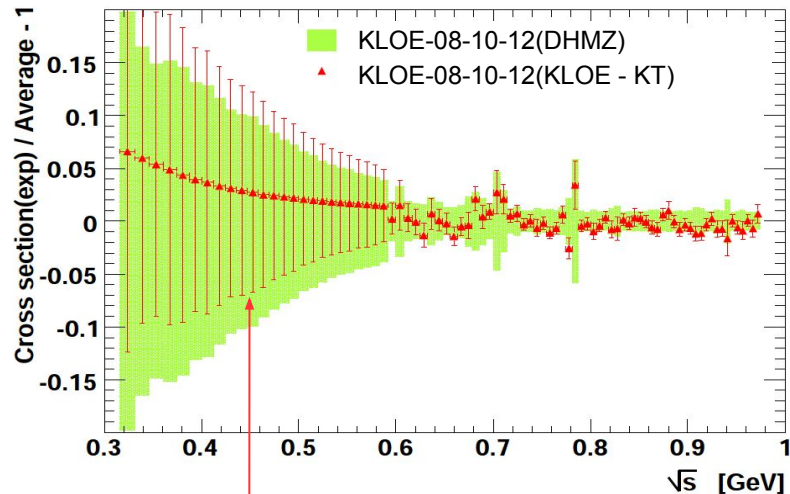
→ Systematics *evaluated* in  $\sim$ wide mass ranges with sharp transitions  
(statistics limitations when going to narrow ranges)



# Combining the 3 KLOE measurements



Local combination (DHMZ)



Information propagated between mass regions, through shifts of systematics - relying on correlations, amplitudes and shapes of systematics (KLOE-KT)

# Combining the 3 KLOE measurements - $a_{\mu}^{\pi\pi}$ contribution

KLOE08  $a_{\mu}[0.6 ; 0.9] : 368.3 \pm 3.2 [10^{-10}]$

KLOE10  $a_{\mu}[0.6 ; 0.9] : 365.6 \pm 3.3$

KLOE12  $a_{\mu}[0.6 ; 0.9] : 366.8 \pm 2.5$

→ Correlation matrix:

	08		10		12	
--	----	--	----	--	----	--

-----  
08 |      1    0.70    0.35

10 |    0.70      1    0.19

12 |    0.35    0.19      1

→ Amount of independent information provided by each measurement

→ KLOE-08-10-12(DHMZ) -  $a_{\mu}[0.6 ; 0.9] : 366.5 \pm 2.8$  (Without  $\chi^2$  rescaling:  $\pm 2.2$ )

→ Conservative treatment of uncertainties and correlations (*not perfectly known*) in weight determination

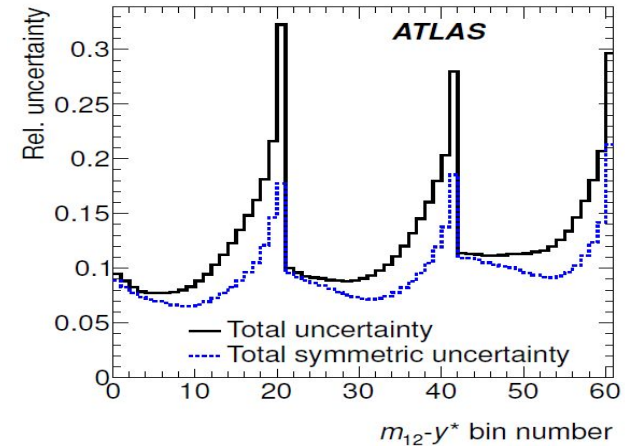
→ KLOE-08-10-12(KLOE-KT) -  $a_{\mu}[0.6 ; 0.9]\text{GeV} : 366.9 \pm 2.2$  (Includes  $\chi^2$  rescaling)

→ Assuming perfect knowledge of the correlations to minimize average uncertainty

# $\chi^2$ definitions and properties

$$\chi^2(\mathbf{d}; \mathbf{t}) = \sum_{i,j} (d_i - t_i) \cdot [C^{-1}(\mathbf{t})]_{ij} \cdot (d_j - t_j) \quad C_{ij} = C_{ij}^{stat} + \sum_k s_i^k \cdot s_j^k$$

$$\chi^2(\mathbf{d}; \mathbf{t}) = \min_{\beta_a} \left\{ \sum_{i,j} \left[ d_i - \left( 1 + \sum_a \beta_a \cdot (\epsilon_a^\pm(\beta_a))_i \right) t_i \right] \cdot [C_{su}^{-1}(\mathbf{t})]_{ij} \cdot \left[ d_j - \left( 1 + \sum_a \beta_a \cdot (\epsilon_a^\pm(\beta_a))_j \right) t_j \right] + \sum_a \beta_a^2 \right\},$$



- Two  $\chi^2$  definitions, with systematic uncertainties included in covariance matrix or treated as fitted “nuisance parameters”
- Equivalent for symmetric Gaussian uncertainties  
(1312.3524 - ATLAS)
- *Both approaches assume the knowledge of the amplitude, shape (phase-space dependence) and correlations of systematic uncertainties*

# Comparing lattice QCD and data-driven results in systematically improvable ways

[2308.04221](#) (BMW & DMZ)

Guiding ideas:

- Need *rigorous* and *realistic* treatment of uncertainties and correlations at all levels  
(Underestimated uncertainties do not bring scientific progress & can put studies on wrong path)
- Caution about significance:  
statistics-dominated measurement; prediction uncertainty limited by non-Gaussian systematic effects
- Studies for understanding differences between data-driven and Lattice QCD approaches need to follow similar standards as the  $g-2$  experiment: *double-blinding*

[Back](#)

# Lattice calculations and comparisons w.r.t. dispersive

→ Lattice: employ simulations to compute electromagnetic-current two-point function

$$C(t) = \frac{1}{3e^2} \sum_{i=1}^3 \int d^3x \langle J_i(\vec{x}, t) J_i(0) \rangle$$

$$\frac{J_\mu}{e} = \frac{2}{3} \bar{u} \gamma_\mu u - \frac{1}{3} \bar{d} \gamma_\mu d - \frac{1}{3} \bar{s} \gamma_\mu s + \frac{2}{3} \bar{c} \gamma_\mu c - \frac{1}{3} \bar{b} \gamma_\mu b + \frac{2}{3} \bar{t} \gamma_\mu t$$

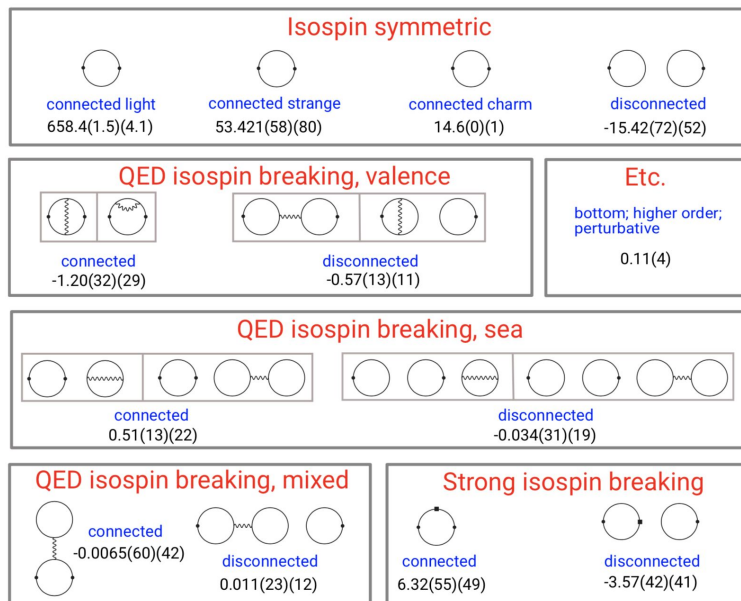
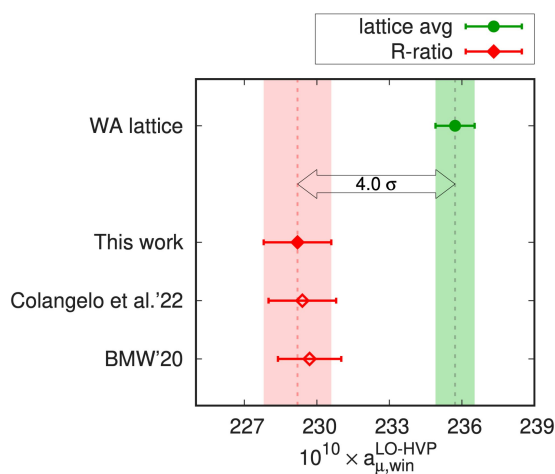
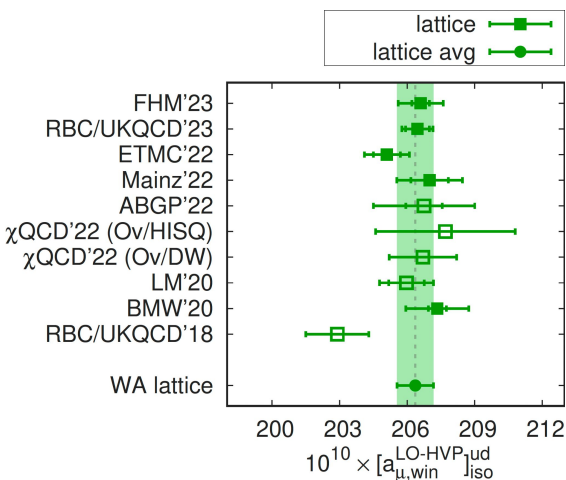
$a_\mu^{\text{LO-HVP}}$ ,  $a_{\mu, \text{win}}^{\text{LO-HVP}}$ ,  $\delta(\Delta\alpha_{\text{had}}^{(5)}) \equiv \Delta\alpha_{\text{had}}^{(5)}(-1 \rightarrow -10\text{GeV}^2), \dots$  are weighted sums of  $C(t)$  over  $t$

Based on BMW'20 (+) with preliminary  $\delta(\Delta\alpha_{\text{had}}^{(5)})$

$$C(t) = \frac{1}{24\pi^2} \int_0^\infty ds \sqrt{s} R(s) e^{-|t|\sqrt{s}}$$

→ Tensions between lattice and data-driven (DD) HVP results

$$[\Delta a_\mu^{\text{LO-HVP}}]_{\text{lat-DD}} \sim 2.1\sigma \quad [\Delta a_{\mu, \text{win}}^{\text{LO-HVP}}]_{\text{lat-DD}} \sim 4.0\sigma$$



All contributions to  $C(t)$ , with all limits taken:

$$a \rightarrow 0, L \rightarrow \infty, M_\pi \rightarrow M_\pi^\phi, \dots$$

# Lattice $\leftrightarrow$ R-ratio comparison: *requirements*

$$C(t) = \frac{1}{24\pi^2} \int_0^\infty ds \sqrt{s} R(s) e^{-|t|\sqrt{s}}$$

→ R-ratio → lattice: “straightforward” (integrate R-ratio)

→ Lattice → R-ratio: inverse Laplace transform (ill-posed problem)

*(Former) Status for lattice calculations:*

→ Very few HVP quantities computed on lattice with:

- All contributions to  $C(t)$ : flavors, various contractions, QED and SIB corrections
- All limits taken:  $a \rightarrow 0$ ,  $L \rightarrow \infty$ ,  $M_\pi \rightarrow M_\pi^\phi$ , ...

→ None with correlations among lattice HVP observables

→ None with uncertainties on these correlations (important for checking stability of inverse problem)

→ Developed statistical approach that:

- Provides useful information with limited lattice input
- Can be systematically improved with more lattice input
- Can (eventually) incorporate physical constraints
- Includes measure of agreement of lattice & R-ratio results with comparison hypothesis
- Accounts for all correlations in lattice and R-ratio observables ...
- ... including uncertainties on these

# Lattice covariances: method

→ Uncertainties and correlations critical for comparisons

→ Use extension of BMW uncertainty method with stat. resampling and syst. histogram, with flat and Akaike Information Criterion (AIC) weights

→ Applicable for observables:  $\{a_j\} = \left\{ a_\mu^{\text{LO-HVP}}, a_{\mu,\text{win}}^{\text{LO-HVP}}, \delta\left(\Delta\alpha_{\text{had}}^{(5)}\right), \dots \right\}$

$$H(A, B) = \sum_{\substack{\phi, \psi_A^{\text{flat}}, \psi_A^{\text{AIC}} \\ \psi_B^{\text{flat}}, \psi_B^{\text{AIC}}} w_A\left(\phi, \psi_A^{\text{flat}}, \psi_A^{\text{AIC}}\right) \times w_B\left(\phi, \psi_B^{\text{flat}}, \psi_B^{\text{AIC}}\right) \\ \times \mathcal{N}_2\left(A, B, \bar{A}\left(\phi, \psi_A^{\text{flat}}, \psi_A^{\text{AIC}}\right), \bar{B}\left(\phi, \psi_B^{\text{flat}}, \psi_B^{\text{AIC}}\right), C_{\text{stat}}^{(1)}\left(\phi, \psi_A^{\text{flat}}, \psi_A^{\text{AIC}}, \psi_B^{\text{flat}}, \psi_B^{\text{AIC}}\right)\right)$$

$$\text{with } w\left(\phi, \psi^{\text{flat}}, \psi^{\text{AIC}}\right) = \frac{\text{AIC}\left(\phi, \psi^{\text{flat}}, \psi^{\text{AIC}}\right)}{\sum_{\psi^{\text{AIC}}} \text{AIC}\left(\phi, \psi^{\text{flat}}, \psi^{\text{AIC}}\right)}$$

↑ correlated / independent choices

→ Build covariance matrix from quantiles of three 1D distributions  $\left( a_\mu^{\text{LO-HVP}}, a_{\mu,\text{win}}^{\text{LO-HVP}}, a_\mu^{\text{LO-HVP}} + a_{\mu,\text{win}}^{\text{LO-HVP}} \right)$

→ Separate stat. & syst. by solving (for  $\lambda = 2$ )

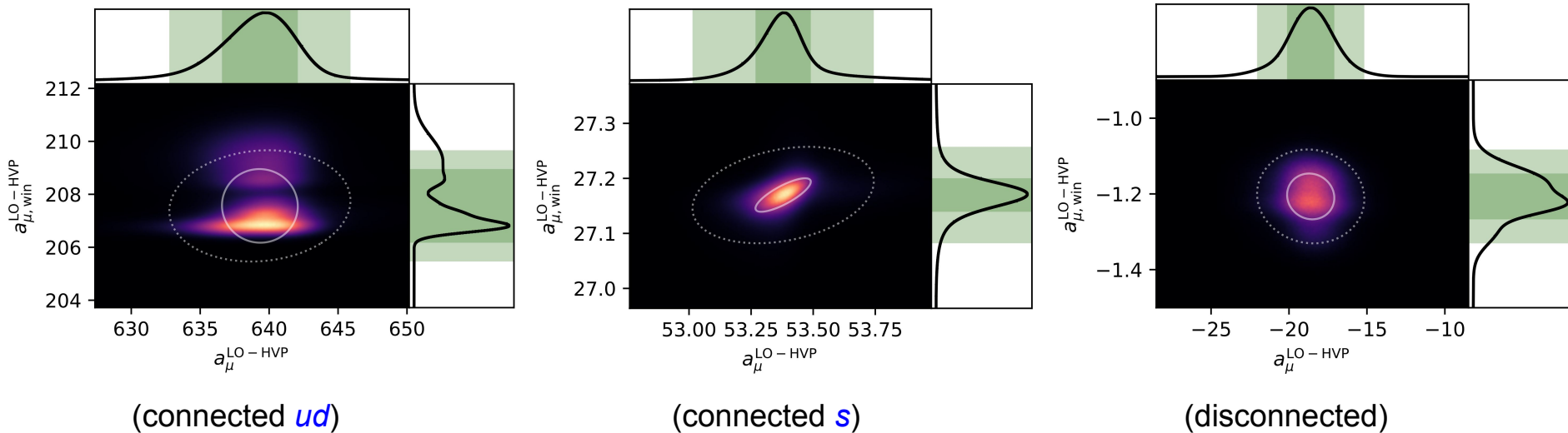
$$C_{\text{stat}} + C_{\text{syst}} = C$$

$$\lambda C_{\text{stat}} + C_{\text{syst}} = C_\lambda$$

# Lattice covariances: results

→  $\delta(\Delta\alpha_{\text{had}}^{(5)})$  largely uncorrelated with other two observables

→ Uncertainties and correlations of  $a_{\mu}^{\text{LO-HVP}}$  &  $a_{\mu,\text{win}}^{\text{LO-HVP}}$  contributions (units of  $10^{-10}$ )



→ Double peak structure due to the variation  $\alpha_S^{(n=0,3)}$  in continuum extrapolation

→ Taken into account by considering  $1\sigma$  &  $2\sigma$  quantiles



# Uncertainties on lattice covariances

- *Uncertainties on covariance matrix* could potentially compromise the inverse problem
- *Stat. error on error* estimated from bootstrap on only 48 jackknife samples (sufficient for this study)
- *Syst. error on error* from:

- For: ud, s, QED, SIB connected, and disconnected

→ Get uncertainties from 1 or  $2\sigma$  quantiles

→ 0 or 100% correlations in a → 0 uncertainties of  $T = a_{\mu}^{\text{LO-HVP}}$  and  $W = a_{\mu, \text{win}}^{\text{LO-HVP}}$ , with  $C = T - W$

$$C_{TW} = C_{TW}^{\text{other}} + \begin{bmatrix} (dW)^2 + (dC)^2 & \{0, 1\} \times (dW)^2 \\ \{0, 1\} \times (dW)^2 & (dW)^2 \end{bmatrix}_{\text{cont}}$$

- Similarly for c

→ Result (in units of  $10^{-20}$ ):

$$C_{\text{lat}}^{1\sigma, 0\%} = \begin{bmatrix} 30.13(4.88) & -0.05(0.03) \\ -0.05(0.03) & 1.95(0.47) \end{bmatrix}$$

$$C_{\text{lat}}^{2\sigma, 0\%} = \begin{bmatrix} 34.04(16.80) & 0.32(0.05) \\ 0.32(0.05) & 1.12(0.07) \end{bmatrix}$$

$$C_{\text{lat}}^{1\sigma, 100\%} = \begin{bmatrix} 30.13(4.88) & 1.56(0.03) \\ 1.56(0.03) & 1.95(0.47) \end{bmatrix}$$

$$C_{\text{lat}}^{2\sigma, 100\%} = \begin{bmatrix} 34.04(16.80) & 1.94(0.05) \\ 1.94(0.05) & 1.12(0.07) \end{bmatrix}$$

# Testing lattice

→ 1-by-1 comparison of moment integrals

Observable	lattice <sup>BMW'20</sup>	data-driven	diff.	% diff.	$\sigma$	$p$ -value [%]
$a_\mu^{\text{LO-HVP}} \times 10^{10}$	707.5(5.5)	694.0(4.0)	13.5(6.8)	1.9(1.0)	2.0	4.7
$a_{\mu,\text{win}}^{\text{LO-HVP}} \times 10^{10}$	236.7(1.4)	229.2(1.4)	7.5(2.0)	3.2(0.8)	3.8	0.01
$[\Delta_{\text{had}}^{(5)}\alpha(-10 \text{ GeV}^2) - \Delta_{\text{had}}^{(5)}\alpha(-1 \text{ GeV}^2)] \times 10^4$	48.67(0.32)	48.02(0.32)	0.65(0.45)	1.3(0.9)	1.4	15.

→ **Simultaneous** comparisons with correlations

$$\chi^2(a_j) = \sum_{j,k} [a_j^{\text{lat}} - a_j] [C_{\text{lat}}^{-1}]_{jk} [a_k^{\text{lat}} - a_k] + \sum_{j,k} [a_j^{\text{R}} - a_j] [C_{\text{R}}^{-1}]_{jk} [a_k^{\text{R}} - a_k]$$

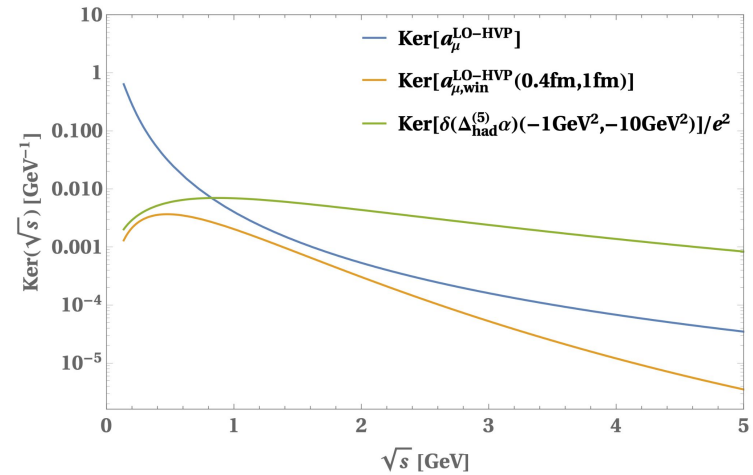
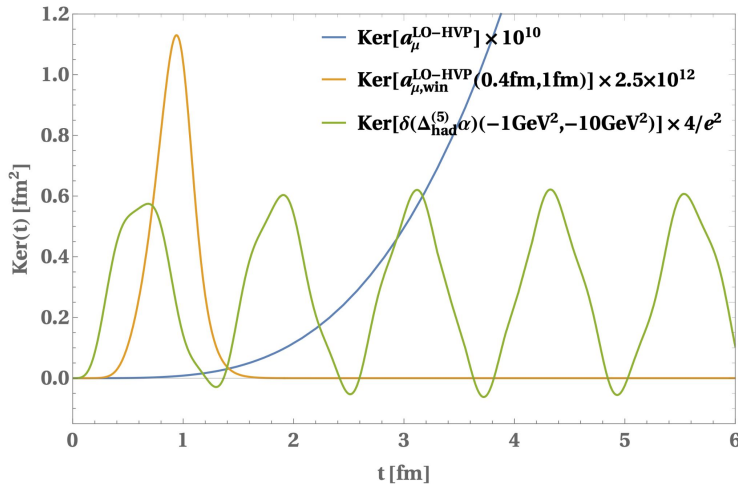
$$\chi_{\text{min}}^2 = \sum_{j,k} [a_j^{\text{lat}} - a_j^{\text{R}}] [(C_{\text{lat}} + C_{\text{R}})^{-1}]_{jk} [a_k^{\text{lat}} - a_k^{\text{R}}]$$

# observ.	$\chi^2/\text{dof}$	$p$ -value [%]
2	14.4/2 – 18.8/2	0.002 – 0.017
3	14.4/3 – 18.8/3	0.009 – 0.63

→ Some dilution compared to  $a_{\mu,\text{win}}^{\text{LO-HVP}}$  alone, but still significant tension

→ (Taking into account the shapes of integral kernels) Differences could be explained by: a  $C(t)$  that is enhanced in  $t \sim [0.4, 1.5]$  fm, also probably for  $t \gtrsim 1.5$  fm, with possible suppression for  $t \lesssim 0.4$  fm

# Consequences of direct lattice / dispersive moments comparison for $C(t)$



→ SD:ID:LD windows

- 10%:33%:57% for  $a_\mu^{\text{LO-HVP}}$
- 70%:29%:1% for  $\delta(\Delta\alpha_{\text{had}}^{(5)})$

+ Taking into account the tensions and agreements above:

→ Excess in  $C(t)$  for  $t \sim [0.4, 1.5]$  fm

→ Probably for  $t \gtrsim 1.5$  fm

→ Possible suppression for  $t \lesssim 0.4$  fm (mainly based on preliminary  $\delta(\Delta\alpha_{\text{had}}^{(5)})$ )

# Testing R-ratio: methodology

→ Chop  $a_j^R$  into contributions  $a_{j,b}^R$  from same  $\sqrt{s}$ -intervals  $I_b$  for all  $j$  :

$$a_j^R = \sum_b a_{j,b}^R$$

→ To accommodate lattice results  $a_j^{\text{lat}}$ , allow common rescaling of  $a_{j,b}^R$ , for all  $j$ , in certain  $I_b$  :

$$a_j^{\text{lat}} = \sum_b \gamma_b \cdot a_{j,b}^R$$

- Simplest interpretation: R-ratio rescaled in  $I_b$

- However, constrains shape of R-ratio modification in limited way: *physical deformation may be allowed*

→ If  $N_j \geq N_b$ , system (over-)constrained: solved here for one  $\gamma$  via weighted average and/or  $\chi^2$  minimization, while avoiding too strong assumptions about the knowledge of uncertainties and correlations

$$a_j^{\text{lat}} = \sum_{b \in A} \gamma \cdot a_{j,b}^R + \sum_{b \in B} a_{j,b}^R \quad \rightarrow \quad \gamma = \frac{a_j^{\text{lat}} - \sum_{b \in B} a_{j,b}^R}{\sum_{b \in A} a_{j,b}^R} \equiv \tilde{\gamma}_j ;$$

$$\chi^2(\gamma) = \sum_{j,k} [\gamma - \tilde{\gamma}_j] \left[ \left( C_{\text{lat}}^{\tilde{\gamma}} + C_{\text{R}}^{\tilde{\gamma}} \right)^{-1} \right]_{jk} [\gamma - \tilde{\gamma}_k]$$

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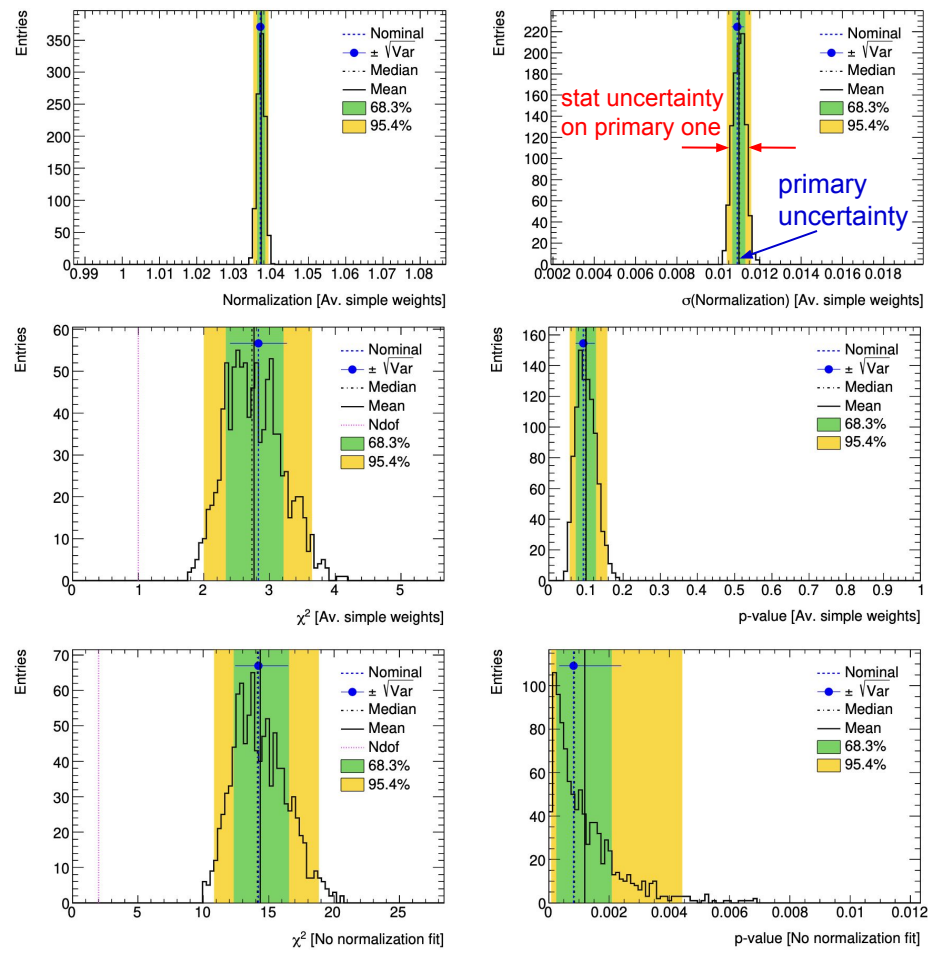

$$\begin{aligned} \chi^2(\gamma) = & \sum_{j,k} \left[ a_j^{\text{lat}} - \sum_{b \in A} \gamma \cdot a_{jb} - \sum_{b \in B} a_{jb} \right] [C_{\text{lat}}^{-1}]_{jk} \left[ a_k^{\text{lat}} - \sum_{c \in A} \gamma \cdot a_{kc} - \sum_{c \in B} a_{kc} \right] \\ & + \sum_{(jb),(kc)} \left[ a_{jb}^R - a_{jb} \right] [C_{\text{R}}^{-1}]_{(jb)(kc)} \left[ a_{kc}^R - a_{kc} \right] \end{aligned}$$

→ Somewhat different interpretation, still compatible results

# Sensitivity to the lattice statistical uncertainties on covariance matrix

- Employ 2<sup>nd</sup> order sampling (bootstraps on jackknife samples) to build distributions for the quantities of interest: re-run procedure with fluctuated lattice covariance matrix
- Quantiles of these distributions to quantitatively evaluate the impact
- Normalisation factor and its uncertainty from fit precisely determined
- Conclusions about  $\chi^2$  and  $p$ -values stable within lattice statistical uncertainties on covariance matrix

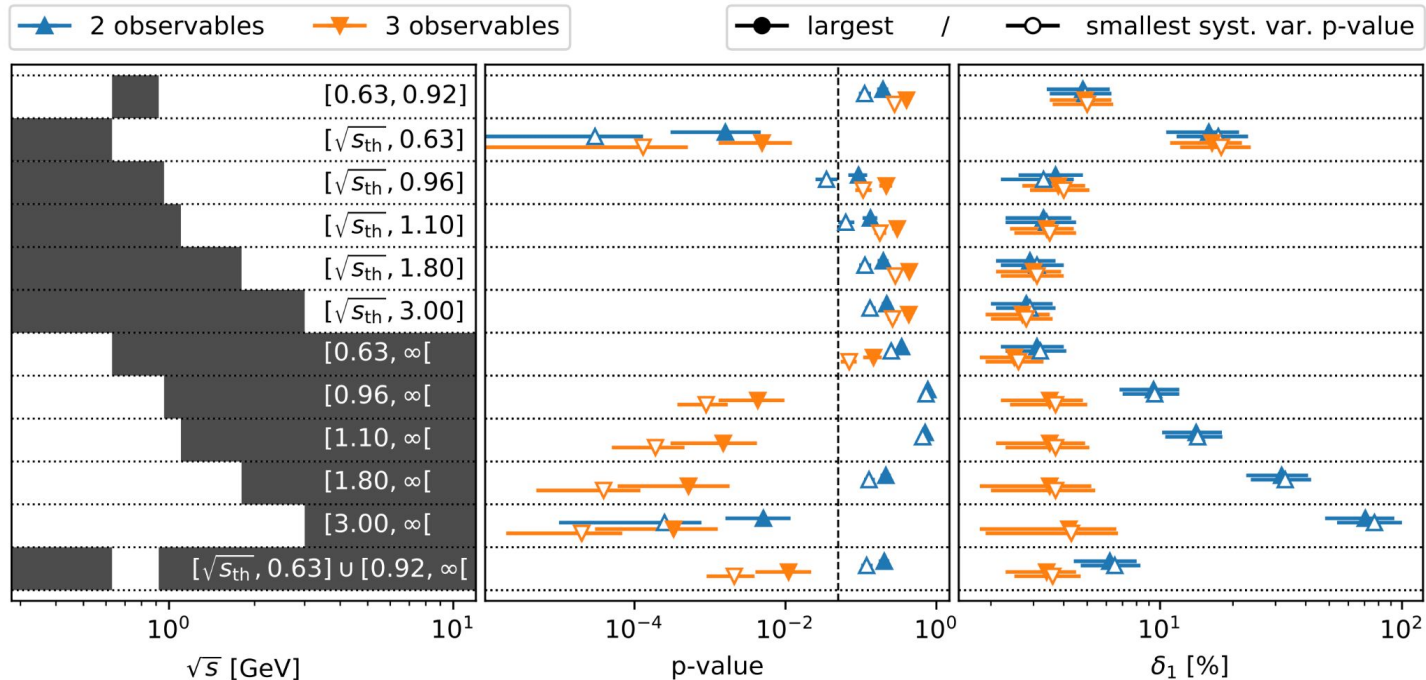
2 input moment integrals; Normalisation fit < 0.96 GeV;  
Lattice covariance matrix "0"



# Testing R-ratio: results

- Lattice → R-ratio: inverse Laplace transform (ill-posed problem)
- *New in this study*: Correlations among lattice HVP observables and uncertainties on these correlations

Consider  $a_j = a_{\mu}^{\text{LO-HVP}}, a_{\mu, \text{win}}^{\text{LO-HVP}}$  (2 observables) with  $a_j = \delta(\Delta\alpha_{\text{had}}^{(5)})$  (3 observables)



- Differences could be explained by enhancing measured R-ratio around (/any larger interval including)  $\rho$ -peak
- Outcome of the studies stable within stat. and syst. uncertainties on lattice covariance matrices
- Rescalings beyond the uncertainties of  $\text{Re}^+e^-$  → No problems for EW fits in case of 3-observable

# Testing R-ratio: summary of results

Modifications to measured R-ratio that could explain lattice results are:

→ Possible in  $\rho$ -peak interval  $[0.63, 0.92]$  GeV for 2 & 3 observables

- Requires rescaling of observables in that interval by  $\sim (5.0 \pm 1.5)\%$

→ Disfavored in interval below  $\rho$ -peak,  $[\sqrt{s}_{\text{th}}, 0.63]$  GeV

→ Possible in  $[\sqrt{s}_{\text{th}}, \sqrt{s}_{\text{max}}]$  with  $\sqrt{s}_{\text{max}} : 0.96 \rightarrow 3.0$  GeV that include  $\rho$ -peak, for 2 & 3 observables

- Rescalings  $\sim (4 \pm 1)\% \rightarrow (3 \pm 1)\%$  for  $\sqrt{s}_{\text{max}} \nearrow$

→ Possible in  $[\sqrt{s}_{\text{min}}, \infty[$  with  $\sqrt{s}_{\text{min}} : 0.63 \rightarrow 1.8$  GeV, for 2 observables

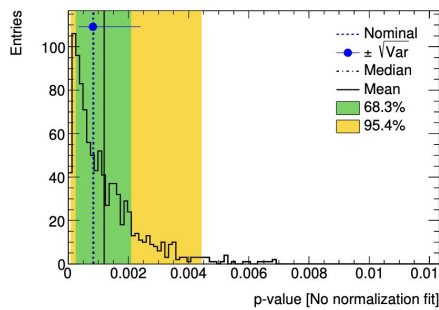
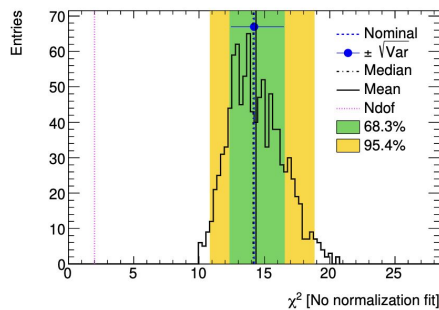
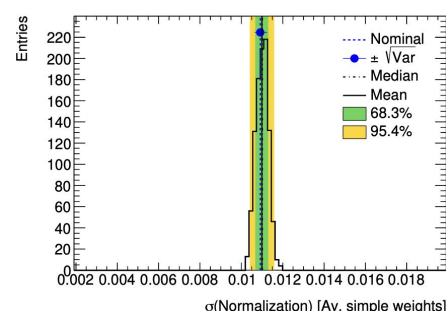
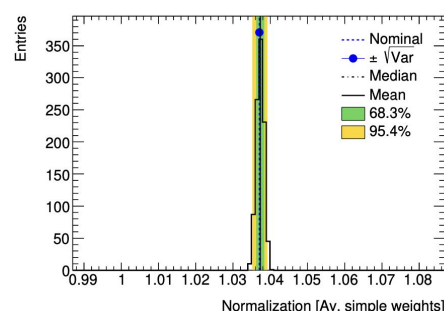
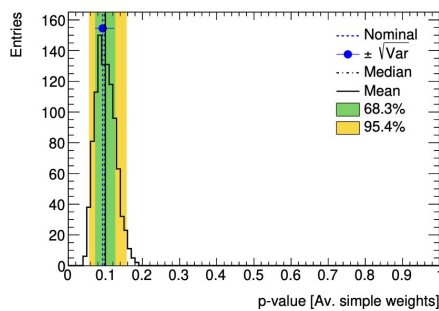
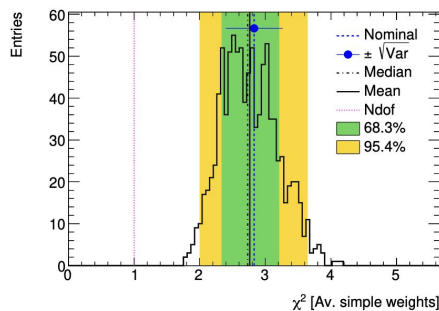
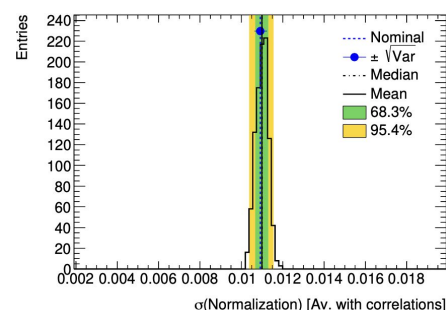
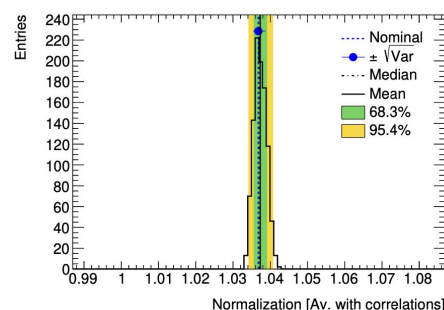
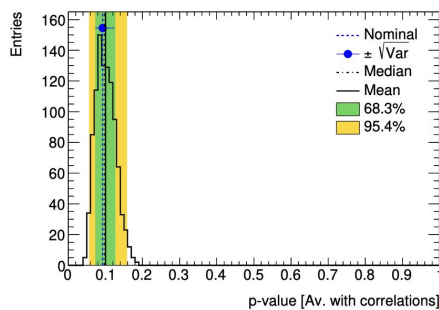
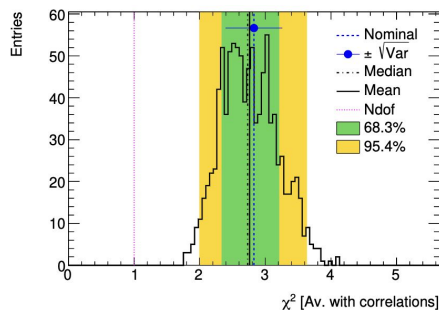
- Rescalings  $\sim (3 \pm 1)\% \rightarrow (32 \pm 9)\%$  for  $\sqrt{s}_{\text{min}} \nearrow$

→ Disfavored in  $[3.0 \text{ GeV}, \infty[$ , for 2 & 3 observables

→ Adding  $\delta(\Delta\alpha_{\text{had}}^{(5)})$  constraint eliminates the possibility of rescalings in  $[\sqrt{s}_{\text{min}}, \infty[$  with  $\sqrt{s}_{\text{min}} : 0.96 \rightarrow 3.0$  GeV that do not include  $\rho$ -peak

# Results - Normalisation < 0.96 GeV; lattice covariance matrix “0”

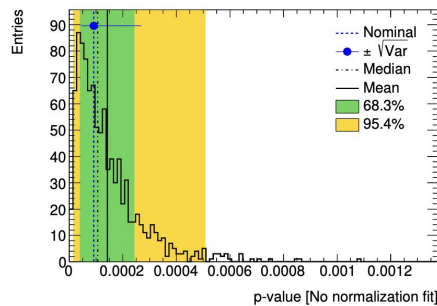
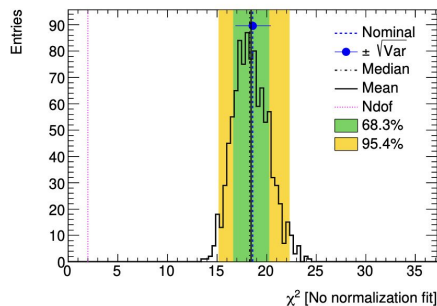
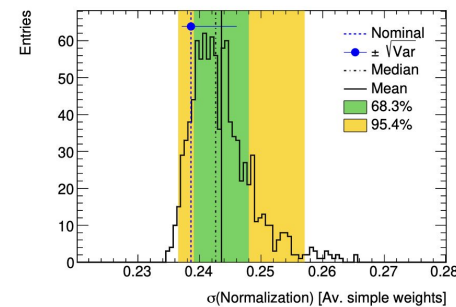
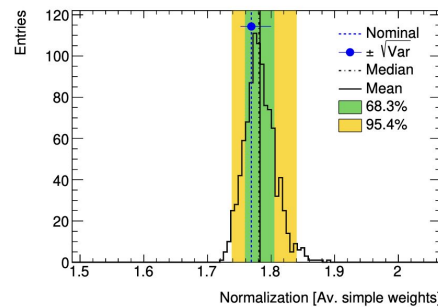
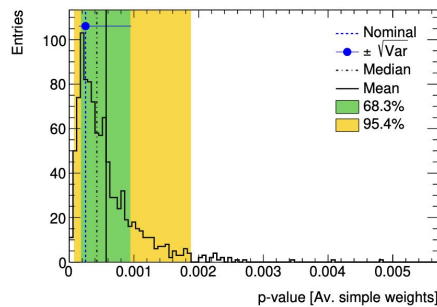
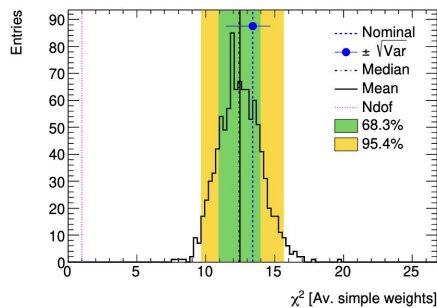
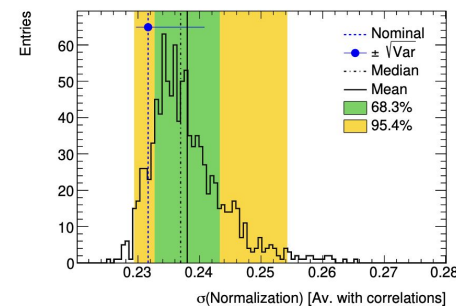
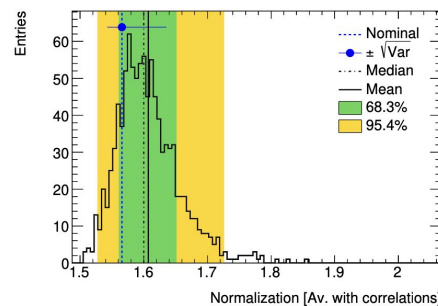
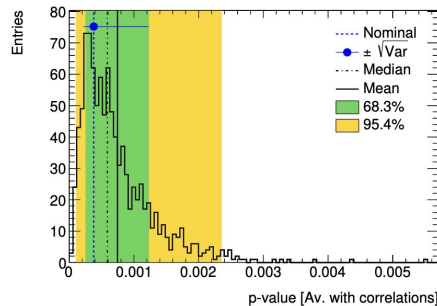
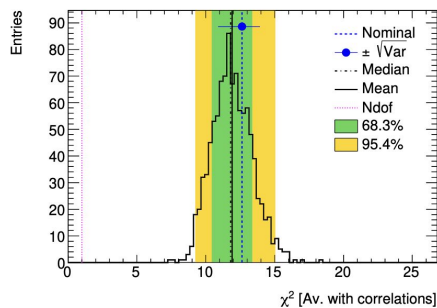
## 2 input moment integrals





# Results - Normalisation > 3 GeV; lattice covariance matrix “3”

## 2 input moment integrals



# Testing R-ratio: results

Number of observables	$I_1$ [GeV]	Lat.cov. $\delta_1 \equiv (\gamma_1 - 1)$	$\chi^2/ndof$	$p$ -value	$\delta_1 \times \Delta\alpha_{\text{had}}^{(5)}(M_Z^2)/I_1 \times 10^4$	
2	$[\sqrt{s_{\text{th}}, 0.63}]$	0	15.9(5.3) $^{+0.9}_{-0.8}\%$	10.0 $^{+2.1}_{-1.9}/1$	0.16 $^{+0.31}_{-0.13}\%$	0.80
2	$[\sqrt{s_{\text{th}}, 0.63}]$	3	17.4(5.7) $^{+0.6}_{-0.5}\%$	17.4 $^{+2.2}_{-1.9}/1$	0.003 $^{+0.010}_{-0.010}\%$	0.88
2	$[0.63, \infty[$	0	3.1(0.9) $^{+0.05}_{-0.05}\%$	0.9 $^{+0.1}_{-0.1}/1$	34.6 $^{+3.2}_{-3.2}\%$	8.49
2	$[0.63, \infty[$	3	3.2(0.9) $^{+0.02}_{-0.02}\%$	1.3 $^{+0.1}_{-0.1}/1$	25.2 $^{+2.8}_{-2.2}\%$	8.71
3	$[\sqrt{s_{\text{th}}, 0.63}]$	0	16.4(5.4) $^{+0.9}_{-0.6}\%$	10.6 $^{+2.2}_{-1.9}/2$	0.49 $^{+0.73}_{-0.31}\%$	0.83
3	$[\sqrt{s_{\text{th}}, 0.63}]$	3	17.9(5.8) $^{+0.6}_{-0.5}\%$	17.8 $^{+2.1}_{-1.9}/2$	0.013 $^{+0.038}_{-0.016}\%$	0.91
3	$[0.63, \infty[$	0	2.5(0.7) $^{+0.06}_{-0.07}\%$	3.8 $^{+0.9}_{-0.5}/2$	14.7 $^{+0.16}_{-0.40}\%$	6.68
3	$[0.63, \infty[$	3	2.6(0.7) $^{+0.04}_{-0.04}\%$	5.3 $^{+0.5}_{-0.4}/2$	7.0 $^{+1.9}_{-1.6}\%$	6.96
2	$[\sqrt{s_{\text{th}}, 0.96}]$	0	3.7(1.1) $^{+0.1}_{-0.1}\%$	2.8 $^{+0.5}_{-0.4}/1$	9.3 $^{+2.8}_{-2.3}\%$	1.32
2	$[\sqrt{s_{\text{th}}, 0.96}]$	3	3.9(1.1) $^{+0.06}_{-0.06}\%$	4.4 $^{+1.1}_{-0.5}/1$	3.5 $^{+1.4}_{-1.0}\%$	1.39
2	$[0.96, \infty[$	0	9.4(2.6) $^{+0.04}_{-0.04}\%$	0.09 $^{+0.01}_{-0.00}/1$	77.0 $^{+1.2}_{-1.1}\%$	22.59
2	$[0.96, \infty[$	3	9.5(2.5) $^{+0.02}_{-0.02}\%$	0.12 $^{+0.01}_{-0.01}/1$	72.9 $^{+1.5}_{-1.5}\%$	22.75
3	$[\sqrt{s_{\text{th}}, 0.96}]$	0	3.8(1.1) $^{+0.09}_{-0.09}\%$	3.1 $^{+0.4}_{-0.3}/2$	21.7 $^{+4.3}_{-4.2}\%$	1.36
3	$[\sqrt{s_{\text{th}}, 0.96}]$	3	4.0(1.1) $^{+0.05}_{-0.05}\%$	4.5 $^{+0.3}_{-0.4}/2$	10.7 $^{+3.2}_{-2.3}\%$	1.42
3	$[0.96, \infty[$	2	3.5(1.3) $^{+0.2}_{-0.2}\%$	10.9 $^{+2.2}_{-2.2}/2$	0.43 $^{+0.54}_{-0.30}\%$	8.35
3	$[0.96, \infty[$	3	3.7(1.3) $^{+0.1}_{-0.1}\%$	14.1 $^{+1.5}_{-1.2}/2$	0.089 $^{+0.083}_{-0.052}\%$	8.91
2	$[\sqrt{s_{\text{th}}, 1.1}]$	0	3.3(1.0) $^{+0.1}_{-0.1}\%$	2.2 $^{+0.4}_{-0.3}/1$	13.4 $^{+3.2}_{-2.9}\%$	1.40
2	$[\sqrt{s_{\text{th}}, 1.1}]$	3	3.4(1.0) $^{+0.05}_{-0.04}\%$	3.5 $^{+0.3}_{-0.4}/1$	6.3 $^{+1.9}_{-1.3}\%$	1.46
2	$[1.1, \infty[$	0	14.1(3.9) $^{+0.07}_{-0.08}\%$	0.1 $^{+0.02}_{-0.02}/1$	70.9 $^{+1.6}_{-1.6}\%$	33.01
2	$[1.1, \infty[$	3	14.3(3.8) $^{+0.04}_{-0.04}\%$	0.2 $^{+0.02}_{-0.02}/1$	65.8 $^{+1.8}_{-1.4}\%$	33.31
3	$[\sqrt{s_{\text{th}}, 1.1}]$	0	3.4(1.0) $^{+0.07}_{-0.07}\%$	2.4 $^{+0.3}_{-0.3}/2$	30.3 $^{+4.5}_{-4.5}\%$	1.44
3	$[\sqrt{s_{\text{th}}, 1.1}]$	3	3.5(1.0) $^{+0.04}_{-0.04}\%$	3.5 $^{+0.3}_{-0.3}/2$	17.8 $^{+3.8}_{-2.8}\%$	1.49
3	$[1.1, \infty[$	2	3.5(1.4) $^{+0.2}_{-0.2}\%$	13.0 $^{+2.7}_{-2.0}/2$	0.15 $^{+0.27}_{-0.12}\%$	8.14
3	$[1.1, \infty[$	3	3.7(1.4) $^{+0.1}_{-0.1}\%$	17.1 $^{+1.9}_{-1.6}/2$	0.019 $^{+0.027}_{-0.014}\%$	8.70
2	$[\sqrt{s_{\text{th}}, 1.8}]$	0	2.9(0.8) $^{+0.1}_{-0.1}\%$	1.7 $^{+0.3}_{-0.2}/1$	19.8 $^{+3.4}_{-3.2}\%$	1.63
2	$[\sqrt{s_{\text{th}}, 1.8}]$	3	3.1(0.9) $^{+0.05}_{-0.03}\%$	2.5 $^{+0.2}_{-0.3}/1$	11.3 $^{+2.4}_{-1.8}\%$	1.69
2	$[1.8, \infty[$	0	31.8(9.1) $^{+0.6}_{-0.6}\%$	1.5 $^{+0.2}_{-0.2}/1$	21.4 $^{+3.4}_{-3.4}\%$	70.17
2	$[1.8, \infty[$	3	32.9(9.2) $^{+0.4}_{-0.4}\%$	2.3 $^{+0.2}_{-0.2}/1$	12.8 $^{+2.5}_{-1.8}\%$	72.62
3	$[\sqrt{s_{\text{th}}, 1.8}]$	0	3.0(0.9) $^{+0.05}_{-0.05}\%$	1.7 $^{+0.2}_{-0.2}/2$	43.7 $^{+5.2}_{-5.2}\%$	1.65
3	$[\sqrt{s_{\text{th}}, 1.8}]$	3	3.1(0.9) $^{+0.03}_{-0.03}\%$	2.5 $^{+0.2}_{-0.3}/2$	28.6 $^{+4.5}_{-3.5}\%$	1.70
3	$[1.8, \infty[$	2	3.5(1.7) $^{+0.2}_{-0.1}\%$	15.1 $^{+3.6}_{-2.0}/2$	0.052 $^{+0.130}_{-0.046}\%$	7.79
3	$[1.8, \infty[$	3	3.7(1.7) $^{+0.08}_{-0.08}\%$	20.3 $^{+2.4}_{-2.0}/2$	0.0039 $^{+0.0081}_{-0.0034}\%$	8.18
2	$[\sqrt{s_{\text{th}}, 3.0}]$	0	2.8(0.8) $^{+0.06}_{-0.06}\%$	1.5 $^{+0.2}_{-0.2}/1$	22.0 $^{+3.4}_{-2.5}\%$	2.03
2	$[\sqrt{s_{\text{th}}, 3.0}]$	3	2.9(0.8) $^{+0.03}_{-0.03}\%$	2.3 $^{+0.2}_{-0.2}/1$	13.2 $^{+2.5}_{-1.9}\%$	2.10
2	$[3.0, \infty[$	0	70.5(22.4) $^{+0.6}_{-3.2}\%$	7.8 $^{+1.1}_{-1.4}/1$	0.51 $^{+0.66}_{-0.35}\%$	143.42
2	$[3.0, \infty[$	3	76.9(23.9) $^{+2.2}_{-2.2}\%$	13.4 $^{+1.5}_{-1.5}/1$	0.025 $^{+0.052}_{-0.024}\%$	156.38
3	$[\sqrt{s_{\text{th}}, 3.0}]$	0	2.7(0.8) $^{+0.05}_{-0.05}\%$	1.7 $^{+0.3}_{-0.2}/2$	43.1 $^{+5.8}_{-6.1}\%$	1.97
3	$[\sqrt{s_{\text{th}}, 3.0}]$	3	2.8(0.8) $^{+0.03}_{-0.03}\%$	2.7 $^{+0.3}_{-0.3}/2$	26.3 $^{+4.3}_{-3.7}\%$	2.04
3	$[3.0, \infty[$	2	4.2(2.4) $^{+0.09}_{-0.08}\%$	16.0 $^{+3.9}_{-2.6}/2$	0.033 $^{+0.084}_{-0.030}\%$	8.59
3	$[3.0, \infty[$	3	4.3(2.4) $^{+0.06}_{-0.05}\%$	21.7 $^{+2.7}_{-2.2}/2$	0.0020 $^{+0.0049}_{-0.0018}\%$	8.80
2	$[0.63, 0.92]$	0	4.8(1.4) $^{+0.1}_{-0.1}\%$	1.7 $^{+0.3}_{-0.2}/1$	19.6 $^{+3.4}_{-3.2}\%$	1.42
2	$[0.63, 0.92]$	3	4.9(1.4) $^{+0.06}_{-0.05}\%$	2.5 $^{+0.2}_{-0.3}/1$	11.2 $^{+2.4}_{-1.8}\%$	1.47
2	$[\sqrt{s_{\text{th}}, 0.63}] \cup [0.92, \infty[$	0	6.2(1.8) $^{+0.1}_{-0.1}\%$	1.6 $^{+0.2}_{-0.2}/1$	20.4 $^{+3.4}_{-3.2}\%$	15.33
2	$[\sqrt{s_{\text{th}}, 0.63}] \cup [0.92, \infty[$	3	6.5(1.8) $^{+0.07}_{-0.07}\%$	2.4 $^{+0.2}_{-0.2}/1$	11.9 $^{+2.4}_{-1.8}\%$	15.88
3	$[0.63, 0.92]$	0	4.9(1.4) $^{+0.08}_{-0.08}\%$	1.8 $^{+0.2}_{-0.2}/2$	40.2 $^{+4.0}_{-4.1}\%$	1.45
3	$[0.63, 0.92]$	3	5.0(1.4) $^{+0.05}_{-0.05}\%$	2.6 $^{+0.2}_{-0.2}/2$	27.9 $^{+4.0}_{-4.1}\%$	1.50
3	$[\sqrt{s_{\text{th}}, 0.63}] \cup [0.92, \infty[$	0	3.4(1.1) $^{+0.1}_{-0.1}\%$	9.0 $^{+1.8}_{-1.3}/2$	1.1 $^{+1.1}_{-0.7}\%$	8.30
3	$[\sqrt{s_{\text{th}}, 0.63}] \cup [0.92, \infty[$	3	3.6(1.1) $^{+0.08}_{-0.08}\%$	12.4 $^{+1.3}_{-1.1}/2$	0.21 $^{+0.18}_{-0.12}\%$	8.80

# Considering more observables in the data-driven approach

→ Enhancement of available information limited by the (anti-)correlations among the moment integrals

Moment integral	Correlation coefficients																		
$\Delta\alpha_{\text{had}}^{(5)}(-10 \text{ GeV}^2)$	1																		
$\Delta\alpha_{\text{had}}^{(5)}(-9 \text{ GeV}^2)$	0.999	1																	
$\Delta\alpha_{\text{had}}^{(5)}(-8 \text{ GeV}^2)$	0.999	0.999	1																
$\Delta\alpha_{\text{had}}^{(5)}(-7 \text{ GeV}^2)$	0.996	0.998	0.999	1															
$\Delta\alpha_{\text{had}}^{(5)}(-6 \text{ GeV}^2)$	0.993	0.995	0.998	0.999	1														
$\Delta\alpha_{\text{had}}^{(5)}(-5 \text{ GeV}^2)$	0.986	0.990	0.994	0.997	0.999	1													
$\Delta\alpha_{\text{had}}^{(5)}(-4 \text{ GeV}^2)$	0.976	0.981	0.986	0.991	0.995	0.999	1												
$\Delta\alpha_{\text{had}}^{(5)}(-3 \text{ GeV}^2)$	0.960	0.966	0.973	0.980	0.986	0.993	0.998	1											
$\Delta\alpha_{\text{had}}^{(5)}(-2 \text{ GeV}^2)$	0.931	0.939	0.948	0.957	0.967	0.977	0.987	0.996	1										
$\Delta\alpha_{\text{had}}^{(5)}(-1 \text{ GeV}^2)$	0.874	0.885	0.896	0.909	0.923	0.938	0.955	0.973	0.990	1									
$a_{\mu,\text{win}}^{\text{LO-HVP}}[0, 0.1] \text{ fm}$	0.806	0.791	0.774	0.753	0.728	0.698	0.660	0.611	0.543	0.442	1								
$a_{\mu,\text{win}}^{\text{LO-HVP}}[0.1, 0.4] \text{ fm}$	0.959	0.955	0.949	0.942	0.931	0.916	0.895	0.864	0.813	0.723	0.864	1							
$a_{\mu,\text{win}}^{\text{LO-HVP}}[0.4, 0.7] \text{ fm}$	0.876	0.887	0.899	0.912	0.926	0.940	0.954	0.966	0.972	0.958	0.428	0.786	1						
$a_{\mu,\text{win}}^{\text{LO-HVP}}[0.7, 1] \text{ fm}$	0.711	0.726	0.743	0.762	0.784	0.809	0.838	0.873	0.91	0.961	0.206	0.509	0.893	1					
$a_{\mu,\text{win}}^{\text{LO-HVP}}[1, 1.3] \text{ fm}$	0.604	0.619	0.636	0.656	0.678	0.705	0.738	0.778	0.831	0.901	0.123	0.365	0.775	0.973	1				
$a_{\mu,\text{win}}^{\text{LO-HVP}}[1.3, 1.6] \text{ fm}$	0.553	0.568	0.584	0.604	0.626	0.653	0.686	0.728	0.783	0.861	0.093	0.305	0.710	0.941	0.993	1			
$a_{\mu,\text{win}}^{\text{LO-HVP}}[1.6, 2.6] \text{ fm}$	0.508	0.522	0.537	0.556	0.577	0.604	0.636	0.677	0.733	0.814	0.074	0.260	0.647	0.891	0.963	0.987	1		
$a_{\mu,\text{win}}^{\text{LO-HVP}}[2.6, 4] \text{ fm}$	0.419	0.431	0.445	0.461	0.479	0.502	0.530	0.567	0.617	0.694	0.052	0.197	0.523	0.753	0.840	0.885	0.944	1	
$a_{\mu,\text{win}}^{\text{LO-HVP}}[4, \infty] \text{ fm}$	0.312	0.321	0.332	0.344	0.358	0.375	0.397	0.426	0.466	0.528	0.034	0.137	0.381	0.565	0.646	0.698	0.787	0.942	1

→ Employing blinding approach in BMW - DHMZ collaboration: here sharing only uncertainties and correlations for dispersive result while pending lattice-based calculations of new moments

# Considering more observables in the data-driven approach

→ Quantify available information through the distribution of eigenvalues for covariance, correlation and normalized covariance matrices (complementary information): strong correlations yield small eigenvalues

Moment integral	Total uncertainty	Covariance	Correlation	Normalized covariance
$a_{\mu, \text{win}}^{\text{LO-HVP}} [t_{\min}, t_{\max}]$				
[0, 0.1] fm	$8.18 \cdot 10^{-12}$	$3.12 \cdot 10^{-20}$	4.85	$2.04 \cdot 10^{-4}$
[0.1, 0.4] fm	$3.86 \cdot 10^{-11}$	$3.52 \cdot 10^{-21}$	1.76	$8.39 \cdot 10^{-5}$
[0.4, 0.7] fm	$6.43 \cdot 10^{-11}$	$6.38 \cdot 10^{-22}$	$3.32 \cdot 10^{-1}$	$1.45 \cdot 10^{-5}$
[0.7, 1] fm	$8.00 \cdot 10^{-11}$	$1.53 \cdot 10^{-22}$	$4.88 \cdot 10^{-2}$	$2.06 \cdot 10^{-6}$
[1, 1.3] fm	$7.90 \cdot 10^{-11}$	$8.77 \cdot 10^{-24}$	$4.27 \cdot 10^{-3}$	$1.87 \cdot 10^{-7}$
[1.3, 1.6] fm	$6.32 \cdot 10^{-11}$	$4.64 \cdot 10^{-25}$	$4.47 \cdot 10^{-4}$	$1.88 \cdot 10^{-8}$
[1.6, $\infty$ [ fm	$1.15 \cdot 10^{-10}$	$7.89 \cdot 10^{-26}$	$1.51 \cdot 10^{-5}$	$6.34 \cdot 10^{-10}$

←  $[C_R]_{ij} / (a_i^R \cdot a_j^R)$

Moment integral	Total uncertainty	Covariance	Correlation	Normalized covariance
$a_{\mu, \text{win}}^{\text{LO-HVP}} [t_{\min}, t_{\max}]$				
[0, 0.1] fm	$8.18 \cdot 10^{-12}$	$2.76 \cdot 10^{-20}$	6.07	$2.49 \cdot 10^{-4}$
[0.1, 0.4] fm	$3.86 \cdot 10^{-11}$	$3.17 \cdot 10^{-21}$	1.99	$9.28 \cdot 10^{-5}$
[0.4, 0.7] fm	$6.43 \cdot 10^{-11}$	$5.11 \cdot 10^{-22}$	$6.86 \cdot 10^{-1}$	$2.72 \cdot 10^{-5}$
[0.7, 1] fm	$8.00 \cdot 10^{-11}$	$1.33 \cdot 10^{-22}$	$2.29 \cdot 10^{-1}$	$9.71 \cdot 10^{-6}$
[1, 1.3] fm	$7.90 \cdot 10^{-11}$	$1.52 \cdot 10^{-23}$	$2.12 \cdot 10^{-2}$	$8.76 \cdot 10^{-7}$
[1.3, 1.6] fm	$6.32 \cdot 10^{-11}$	$1.20 \cdot 10^{-24}$	$2.36 \cdot 10^{-3}$	$9.89 \cdot 10^{-8}$
[1.6, 2.6] fm	$9.32 \cdot 10^{-11}$	$2.91 \cdot 10^{-25}$	$3.90 \cdot 10^{-4}$	$1.59 \cdot 10^{-8}$
[2.6, 4] fm	$2.01 \cdot 10^{-11}$	$2.15 \cdot 10^{-26}$	$3.43 \cdot 10^{-5}$	$1.41 \cdot 10^{-9}$
[4, $\infty$ [ fm	$2.64 \cdot 10^{-12}$	$1.78 \cdot 10^{-27}$	$5.83 \cdot 10^{-7}$	$2.50 \cdot 10^{-11}$

→ 2 extra moment integrals add ~1 d.o.f.

# Considering more observables in the data-driven approach

Moment integral	Total uncertainty	Covariance	Eigenvalues	
			Correlation	Normalized covariance ← $[C_R]_{ij} / (a_i^R \cdot a_j^R)$
$\Delta\alpha_{\text{had}}^{(5)}(-10 \text{ GeV}^2)$	$4.80 \cdot 10^{-5}$	$1.48 \cdot 10^{-8}$	14.74	$5.21 \cdot 10^{-4}$
$\Delta\alpha_{\text{had}}^{(5)}(-9 \text{ GeV}^2)$	$4.64 \cdot 10^{-5}$	$2.17 \cdot 10^{-10}$	3.26	$1.34 \cdot 10^{-4}$
$\Delta\alpha_{\text{had}}^{(5)}(-8 \text{ GeV}^2)$	$4.48 \cdot 10^{-5}$	$4.09 \cdot 10^{-12}$	$7.36 \cdot 10^{-1}$	$2.92 \cdot 10^{-5}$
$\Delta\alpha_{\text{had}}^{(5)}(-7 \text{ GeV}^2)$	$4.29 \cdot 10^{-5}$	$1.94 \cdot 10^{-14}$	$2.40 \cdot 10^{-1}$	$1.02 \cdot 10^{-5}$
$\Delta\alpha_{\text{had}}^{(5)}(-6 \text{ GeV}^2)$	$4.09 \cdot 10^{-5}$	$2.22 \cdot 10^{-16}$	$2.16 \cdot 10^{-2}$	$8.90 \cdot 10^{-7}$
$\Delta\alpha_{\text{had}}^{(5)}(-5 \text{ GeV}^2)$	$3.85 \cdot 10^{-5}$	$3.40 \cdot 10^{-18}$	$2.45 \cdot 10^{-3}$	$1.02 \cdot 10^{-7}$
$\Delta\alpha_{\text{had}}^{(5)}(-4 \text{ GeV}^2)$	$3.57 \cdot 10^{-5}$	$1.99 \cdot 10^{-20}$	$4.17 \cdot 10^{-4}$	$1.68 \cdot 10^{-8}$
$\Delta\alpha_{\text{had}}^{(5)}(-3 \text{ GeV}^2)$	$3.23 \cdot 10^{-5}$	$9.27 \cdot 10^{-23}$	$3.92 \cdot 10^{-5}$	$1.59 \cdot 10^{-9}$
$\Delta\alpha_{\text{had}}^{(5)}(-2 \text{ GeV}^2)$	$2.78 \cdot 10^{-5}$	$1.40 \cdot 10^{-23}$	$2.33 \cdot 10^{-6}$	$8.94 \cdot 10^{-11}$
$\Delta\alpha_{\text{had}}^{(5)}(-1 \text{ GeV}^2)$	$2.07 \cdot 10^{-5}$	$7.03 \cdot 10^{-25}$	$1.15 \cdot 10^{-7}$	$4.78 \cdot 10^{-12}$
$a_{\mu,\text{win}}^{\text{LO-HVP}} [0, 0.1] \text{ fm}$	$8.18 \cdot 10^{-12}$	$1.47 \cdot 10^{-25}$	$3.46 \cdot 10^{-10}$	$1.22 \cdot 10^{-14}$
$a_{\mu,\text{win}}^{\text{LO-HVP}} [0.1, 0.4] \text{ fm}$	$3.86 \cdot 10^{-11}$	$7.96 \cdot 10^{-28}$	$3.76 \cdot 10^{-13}$	$1.19 \cdot 10^{-17}$
$a_{\mu,\text{win}}^{\text{LO-HVP}} [0.4, 0.7] \text{ fm}$	$6.43 \cdot 10^{-11}$	$1.66 \cdot 10^{-28}$	$2.28 \cdot 10^{-13}$	$7.40 \cdot 10^{-18}$
$a_{\mu,\text{win}}^{\text{LO-HVP}} [0.7, 1] \text{ fm}$	$8.00 \cdot 10^{-11}$	$4.50 \cdot 10^{-30}$	$9.74 \cdot 10^{-14}$	$3.10 \cdot 10^{-18}$
$a_{\mu,\text{win}}^{\text{LO-HVP}} [1, 1.3] \text{ fm}$	$7.90 \cdot 10^{-11}$	$5.25 \cdot 10^{-31}$	$3.07 \cdot 10^{-14}$	$9.72 \cdot 10^{-19}$
$a_{\mu,\text{win}}^{\text{LO-HVP}} [1.3, 1.6] \text{ fm}$	$6.32 \cdot 10^{-11}$	$8.57 \cdot 10^{-32}$	$1.40 \cdot 10^{-14}$	$4.38 \cdot 10^{-19}$
$a_{\mu,\text{win}}^{\text{LO-HVP}} [1.6, 2.6] \text{ fm}$	$9.32 \cdot 10^{-11}$	$-4.56 \cdot 10^{-27}$	$-2.24 \cdot 10^{-14}$	$-7.06 \cdot 10^{-19}$
$a_{\mu,\text{win}}^{\text{LO-HVP}} [2.6, 4] \text{ fm}$	$2.01 \cdot 10^{-11}$	$-2.01 \cdot 10^{-23}$	$-3.79 \cdot 10^{-14}$	$-1.19 \cdot 10^{-18}$
$a_{\mu,\text{win}}^{\text{LO-HVP}} [4, \infty[ \text{ fm}$	$2.64 \cdot 10^{-12}$	$-1.12 \cdot 10^{-22}$	$-1.22 \cdot 10^{-13}$	$-3.83 \cdot 10^{-18}$

→ 10 extra moment integrals, but no additional independent d.o.f.

# Detailed conclusions for lattice / dispersive comparisons

→ Presented flexible method for comparing lattice QCD and data-driven HVP results

→ Find that discrepancies/agreements between lattice and data-driven results for

$$a_{\mu}^{\text{LO-HVP}}, a_{\mu, \text{win}}^{\text{LO-HVP}}, \delta\left(\Delta\alpha_{\text{had}}^{(5)}\right)$$

On lattice side, result from:

- a  $C(t)$  that is enhanced in  $t \sim [0.4, 1.5]$  fm
- also probably for  $t \gtrsim 1.5$  fm
- with possible suppression for  $t \lesssim 0.4$  fm (mainly based on preliminary  $\delta\left(\Delta\alpha_{\text{had}}^{(5)}\right)$ )

On data-driven side, could be explained by:

- enhancing measured R-ratio around  $\rho$ -peak
- or in any larger interval including  $\rho$ -peak

→ Lattice and measured R-ratio correlations of uncertainties critical for drawing such conclusions

# Detailed conclusions for lattice / dispersive comparisons

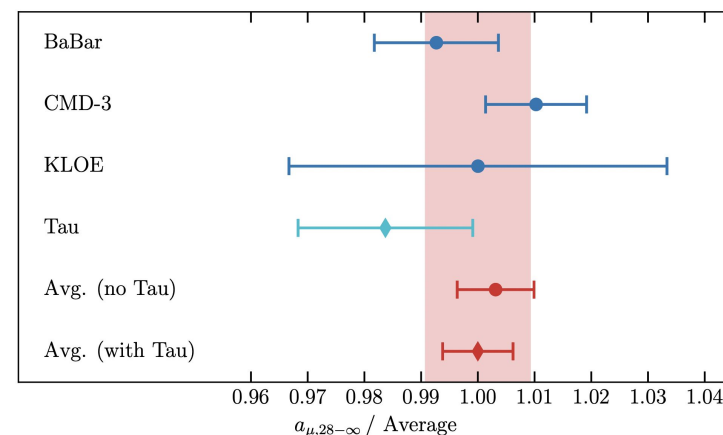
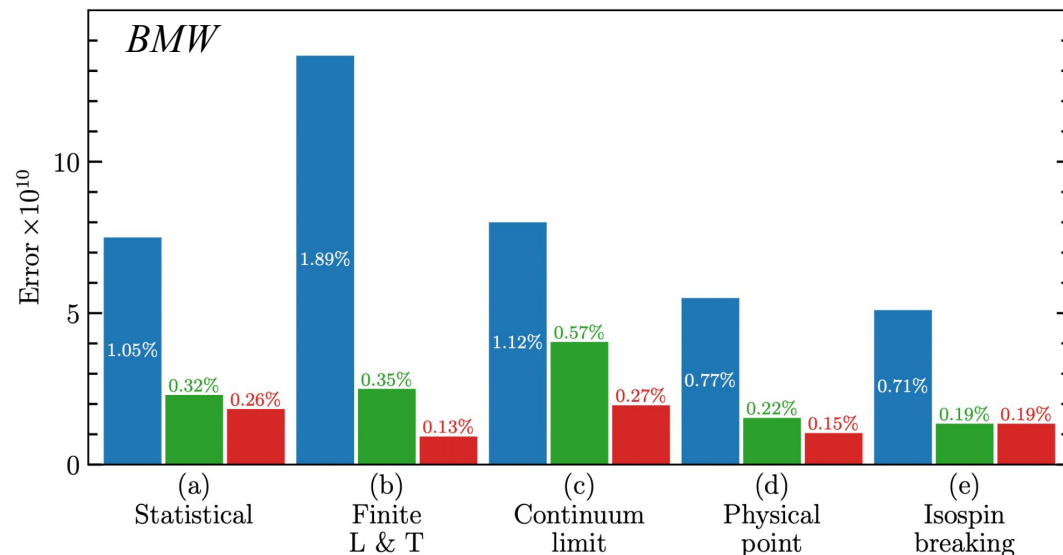
- Important to check that uncertainties on uncertainties and correlations do not spoil picture, especially for inverse problem
  - checked here for lattice stat and syst uncertainties
  - must do so for measured R-ratio uncertainties
- Also important not to share results between 2 approaches before they are final (mutual blinding)
- With more HVP observables, many generalizations possible, also including physics-driven constraints
- However, limit on independent HVP observables in data-driven and lattice approaches
- Same methods can be used to combine determinations of lattice and data-driven results for HVP observables, once differences are understood
- No problems with EW fits in case of 3-observable comparisons (not shown)

# Some references to related work on HVP

- Windows proposed in RBC/UKQCD arXiv:1801.07224
- Discussed in context of detailed comparison in Colangelo et al arXiv:2205.12963
- Consequences of rescaling of measured R-ratio studied in Crivellin et al arXiv:2003.04886, Keshavarzi et al arXiv:2006.12666, de Rafael arXiv:2006.13880, Malaescu et al arXiv:2008.08107
- Consequences of lattice  $\Delta a_\mu^{\text{LO-HVP}}$  on  $\pi^+\pi^-$  contributions to R-ratio with physical constraints in Colangelo et al arXiv:2010.07943
- Use of Backus-Gilbert method for reconstruction of smeared R-ratio from lattice C(t) in Hansen et al arXiv:1903.06476, Alexandrou et al arXiv:2212.08467
- Proposal for comparing measured R-ratio and lattice C(t) via spectral-width sumrules in Boito et al arXiv:2210.13677
- ... (many other references for reconstructing spectral functions from lattice correlators)



# Merging dispersive and lattice HVP calculations

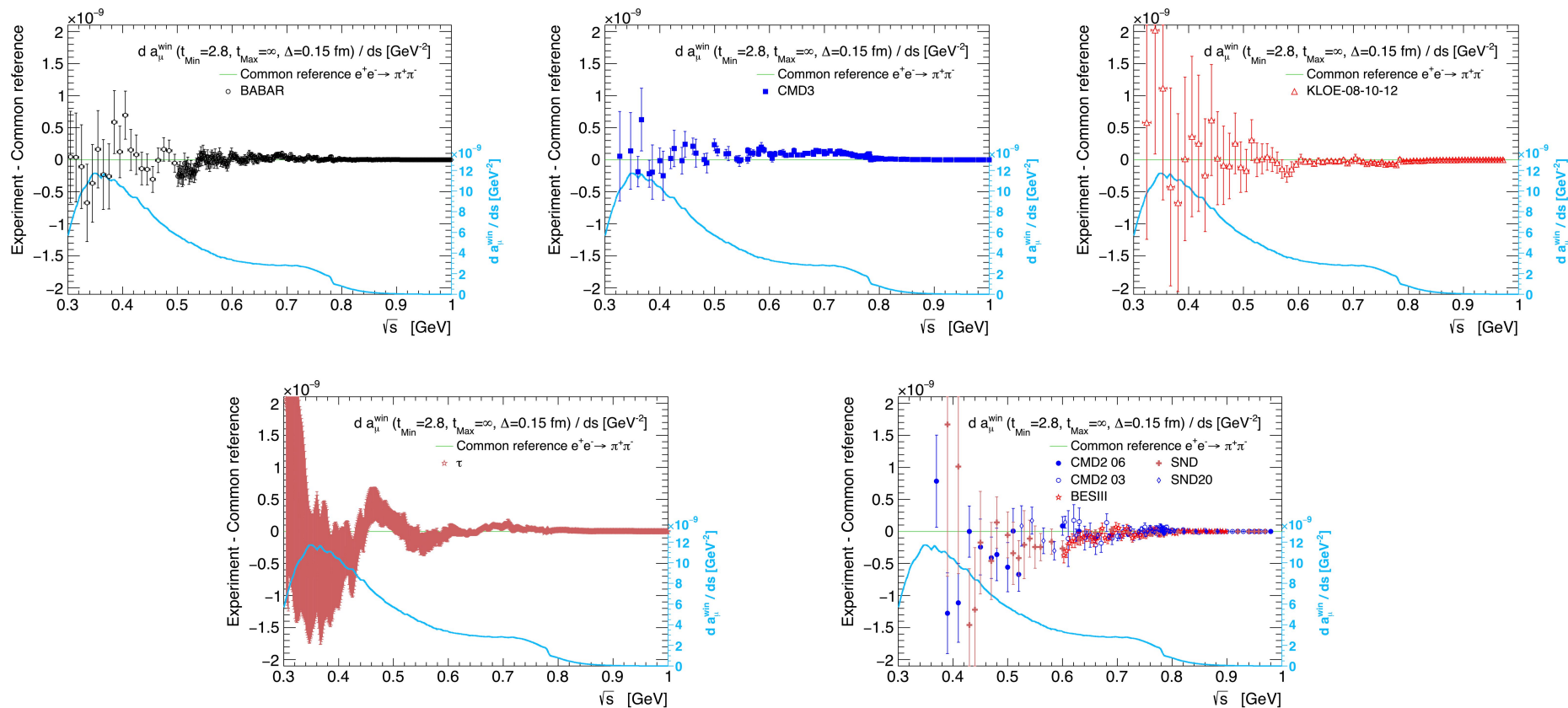


2407.10913

→ Currently most precise prediction, based on improved lattice QCD (BMW) + data-driven inputs (DMZ) at large- $t$  (input data in good agreement at low energy)

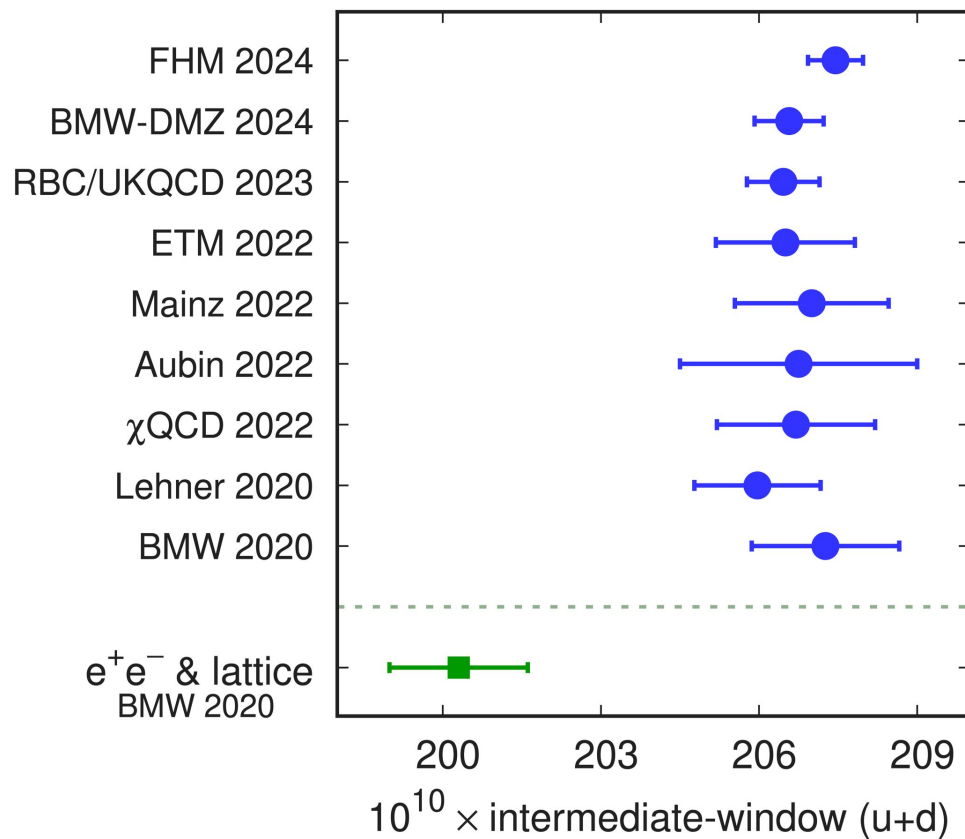
Back

# Merging dispersive and lattice HVP calculations



[2407.10913](https://arxiv.org/abs/2407.10913)

# Comparisons of lattice HVP calculations

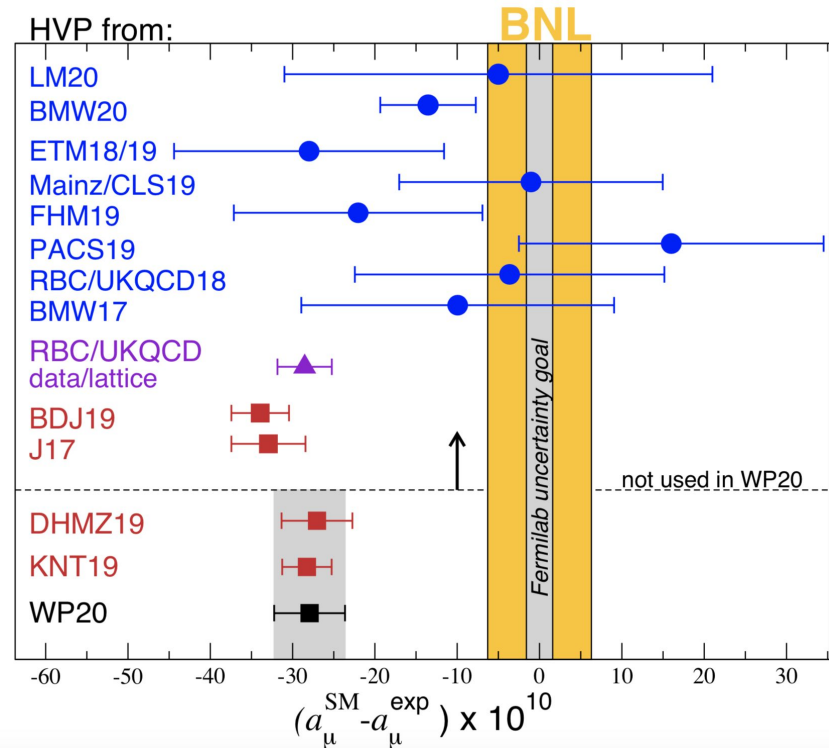


[CERN Courier March-April '25](#)

# Impact of correlations between $a_\mu$ and $\alpha_{\text{QED}}$ on the EW fit

[2008.08107](#)(BM, Matthias Schott)

See also: Crivellin et al, 2003.04886;  
Keshavarzi et al., 2006.12666 ;de Rafael,  
2006.13880; Colangelo et al, 2010.07943



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# Approaches considered for treating the $a_\mu - \alpha_{\text{QED}}$ correlations

Studied approaches probing different hypotheses concerning the possible source(s) of the  $a_\mu$  tension(s) :

(0) Scaling factor applied to the HVP contribution from some energy range of the hadronic spectrum

→ Approaches taking into account (*for the first time*) the full correlations between the uncertainties of the HVP contributions to  $a_\mu$  and  $\alpha_{\text{QED}}$ , based on input from DHMZ 19 (arXiv:1908.00921):

correlations between points/bins of a measurement in a given channel, between different measurements in the same channel, between different channels; full treatment of the BABAR-KLOE tension in the  $\pi^+\pi^-$  channel

Computation (Energy range)	$a_\mu^{\text{HVP, LO}} [10^{-10}]$	$\Delta\alpha_{\text{had}}(M_Z^2) [10^{-4}]$	$\rho$
Phenomenology (Full HVP)	$694.0 \pm 4.0$	$275.3 \pm 1.0$	44%
Phenomenology ([Th.; 1.8 GeV])	$635.5 \pm 3.9$	$55.4 \pm 0.4$	86%
Phenomenology ([Th.; 1 GeV])	$539.8 \pm 3.8$	$36.3 \pm 0.3$	99.5%
Lattice (Full HVP)BMW 20 (v1)	$712.4 \pm 4.5$	-	-

(1) Cov. matrix of  $a_\mu$  and  $\alpha_{\text{QED}}$  (Pheno) described by a nuisance parameter ( $\text{NP}_1$ ) impacting both quantities (used to shift  $a_\mu$  to some “target” value - coherent shift applied to  $\alpha_{\text{QED}}$ ) and another one ( $\text{NP}_2$ ) impacting only  $\alpha_{\text{QED}}$  (used in the EW fit)

Note: “target” values chosen in order to reach agreement with the BMW 20 prediction / Experimental  $a_\mu (\pm 1\sigma)$

Uncertainty components	$a_\mu^{\text{HVP, LO}}$	$\Delta\alpha_{\text{had}}(M_Z^2)$
$\text{NP}_1$	$\sigma(a_\mu^{\text{HVP, LO}})$	$\sigma(\Delta\alpha_{\text{had}}(M_Z^2)) \cdot \rho$
$\text{NP}_2$	0	$\sigma(\Delta\alpha_{\text{had}}(M_Z^2)) \cdot \sqrt{1 - \rho^2}$

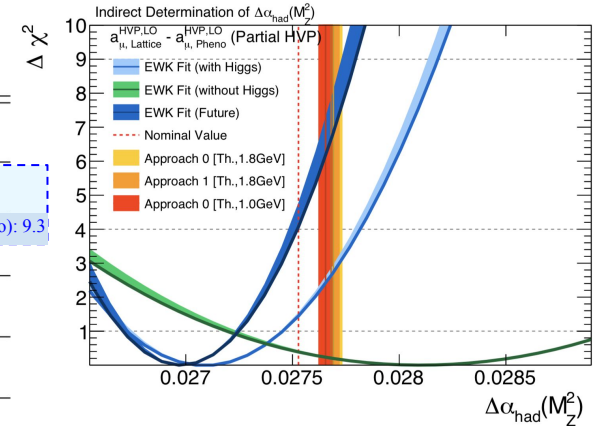
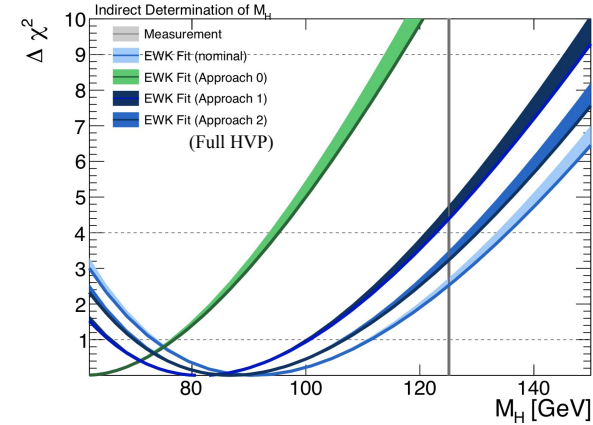
(2) Include the HVP contribution to  $a_\mu$  as extra parameter in the EW fit, constrained by the Pheno & BMW 20 values

Note: Also accounted for the coherent impact of  $\alpha_s$  on the HVP contribution and on the EW fit

# Results: comparing the Phenomenology & BMW 20 values

$a_\mu^{\text{HVP, LO}}$ shift (Energy range)	Approach 0		Approach 1		
	Scaling factor	$\Delta' \alpha_{\text{had}}(M_Z^2)$	Shift NP <sub>1</sub>	$\sigma'(\Delta \alpha_{\text{had}}(M_Z^2))$	$\Delta' \alpha_{\text{had}}(M_Z^2)$
$a_\mu^{\text{HVP, LO}}(\text{Lattice}) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ (Full HVP)	1.027	0.02826	4.6	$9.0 \cdot 10^{-5}$	0.02774
$(a_\mu^{\text{HVP, LO}}(\text{Lattice}) - 1\sigma) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ (Full HVP)	1.020	0.02808	3.5	$9.0 \cdot 10^{-5}$	0.02769
$a_\mu^{\text{HVP, LO}}(\text{Lattice}) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ ([Th.; 1.8 GeV])	1.029	0.02769	4.7	$9.5 \cdot 10^{-5}$	0.02768
$(a_\mu^{\text{HVP, LO}}(\text{Lattice}) - 1\sigma) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ ([Th.; 1.8 GeV])	1.022	0.02765	3.5	$9.5 \cdot 10^{-5}$	0.02764
$a_\mu^{\text{HVP, LO}}(\text{Lattice}) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ ([Th.; 1 GeV])	1.034	0.02765	-	-	-
$(a_\mu^{\text{HVP, LO}}(\text{Lattice}) - 1\sigma) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ ([Th.; 1 GeV])	1.026	0.02762	-	-	-

→ Large scaling factors (w.r.t. exp. uncertainties) & significant shifts of NP<sub>1</sub>



$a_\mu^{\text{HVP, LO}}$ shift (Energy range)	Nominal		Approach 0		Approach 1		Approach 2	
	$\Delta' \alpha_{\text{had}}(M_Z^2)$	$\chi^2/\text{ndf}$	$\Delta' \alpha_{\text{had}}(M_Z^2)$	$\chi^2/\text{ndf}$	$\Delta' \alpha_{\text{had}}(M_Z^2)$	$\chi^2/\text{ndf}$	$\Delta' \alpha_{\text{had}}(M_Z^2)$	$\chi^2/\text{ndf}$
	0.02753	18.6/16 (p=0.29)	-	-	-	-	0.02753	28.1/17 (p=0.04)
$a_\mu^{\text{HVP, LO}}(\text{Lattice}) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ (Full HVP)	-	-	0.02826	27.6/16 (p=0.04)	0.02774	20.3/16 (p=0.21)	-	-
$a_\mu^{\text{HVP, LO}}(\text{Lattice}) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ ([Th.; 1.8 GeV])	-	-	0.02769	19.9/16 (p=0.22)	0.02768	19.8/16 (p=0.23)	-	-
$a_\mu^{\text{HVP, LO}}(\text{Lattice}) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ ([Th.; 1.0 GeV])	-	-	0.02765	19.6/16 (p=0.24)	-	-	-	-

→ Addressing the BMW 20 - Pheno difference for  $a_\mu$  has little impact on the EW fit, except for the unrealistic scenario rescaling the full HVP contribution

Note: Similar conclusions for the comparison with the Experimental  $a_\mu$  value (see next slides)

# Scaling factors and NP shifts

$a_\mu^{\text{HVP, LO}}$ shift (Energy range)	Approach 0		Approach 1		
	Scaling factor	$\Delta' \alpha_{\text{had}}(M_Z^2)$	Shift NP <sub>1</sub>	$\sigma' (\Delta \alpha_{\text{had}}(M_Z^2))$	$\Delta' \alpha_{\text{had}}(M_Z^2)$
$a_\mu^{\text{HVP, LO}} - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ (Full HVP)	1.027	0.02826	4.6	$9.0 \cdot 10^{-5}$	0.02774
$(a_\mu^{\text{HVP, LO}} - 1\sigma) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ (Full HVP)	1.020	0.02808	3.5	$9.0 \cdot 10^{-5}$	0.02769
$a_\mu^{\text{HVP, LO}} - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ ([Th.; 1.8 GeV])	1.029	0.02769	4.7	$9.5 \cdot 10^{-5}$	0.02768
$(a_\mu^{\text{HVP, LO}} - 1\sigma) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ ([Th.; 1.8 GeV])	1.022	0.02765	3.5	$9.5 \cdot 10^{-5}$	0.02764
$a_\mu^{\text{HVP, LO}} - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ ([Th.; 1 GeV])	1.034	0.02765	-	-	-
$(a_\mu^{\text{HVP, LO}} - 1\sigma) - a_\mu^{\text{HVP, LO}}(\text{Pheno})$ ([Th.; 1 GeV])	1.026	0.02762	-	-	-
$a_\mu^{\text{Exp}} - a_\mu^{\text{SM}}(\text{Pheno})$ (Full HVP)	1.037	0.02856	6.6	$9.0 \cdot 10^{-5}$	0.02782
$(a_\mu^{\text{Exp}} - 1\sigma) - a_\mu^{\text{SM}}(\text{Pheno})$ (Full HVP)	1.028	0.02831	5.0	$9.0 \cdot 10^{-5}$	0.02775
$a_\mu^{\text{Exp}} - a_\mu^{\text{SM}}(\text{Pheno})$ ([Th.; 1.8 GeV])	1.041	0.02776	6.6	$9.5 \cdot 10^{-5}$	0.02774
$(a_\mu^{\text{Exp}} - 1\sigma) - a_\mu^{\text{SM}}(\text{Pheno})$ ([Th.; 1.8 GeV])	1.031	0.02770	5.0	$9.5 \cdot 10^{-5}$	0.02769
$a_\mu^{\text{Exp}} - a_\mu^{\text{SM}}(\text{Pheno})$ ([Th.; 1 GeV])	1.048	0.02771	-	-	-
$(a_\mu^{\text{Exp}} - 1\sigma) - a_\mu^{\text{SM}}(\text{Pheno})$ ([Th.; 1 GeV])	1.036	0.02766	-	-	-

→ Large scaling factors (w.r.t. uncertainties) & significant shifts of NP<sub>1</sub>

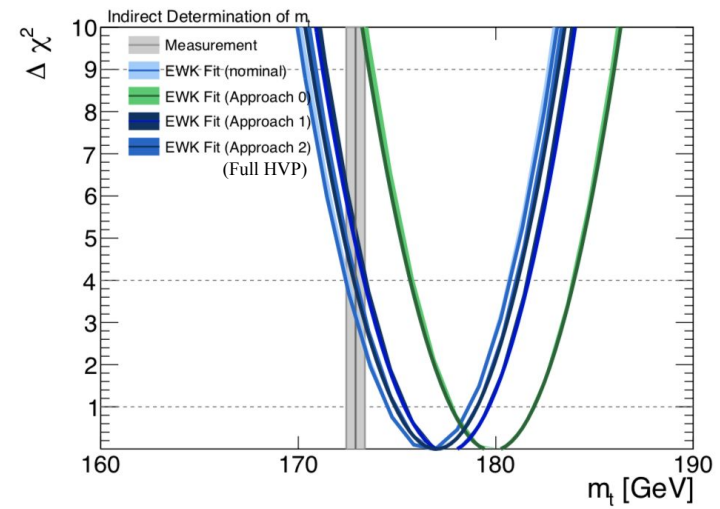
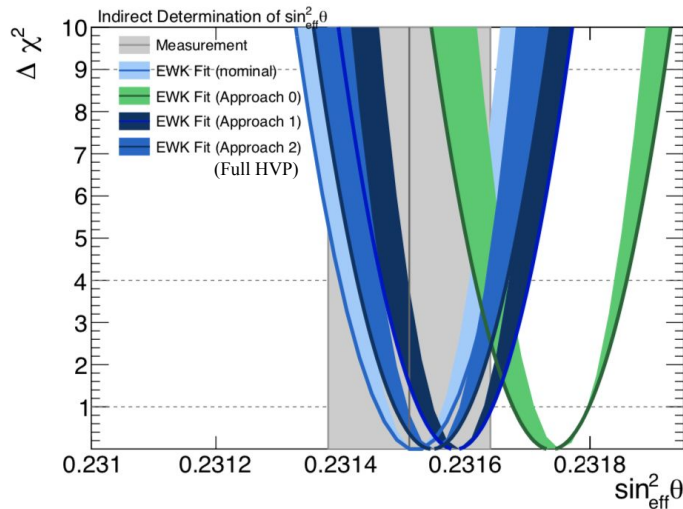
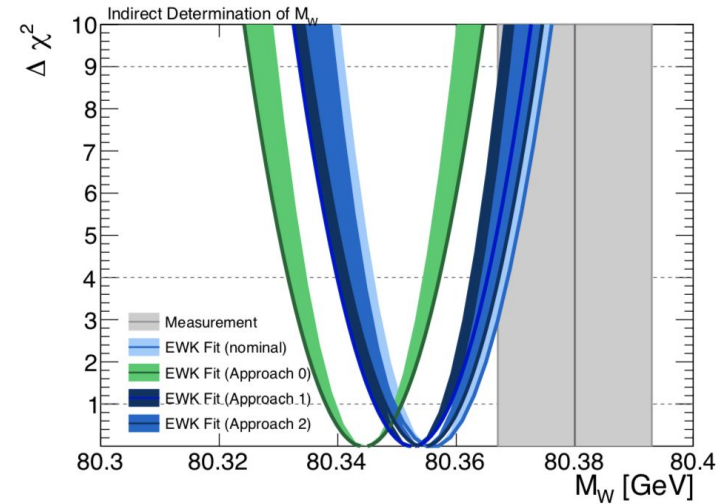
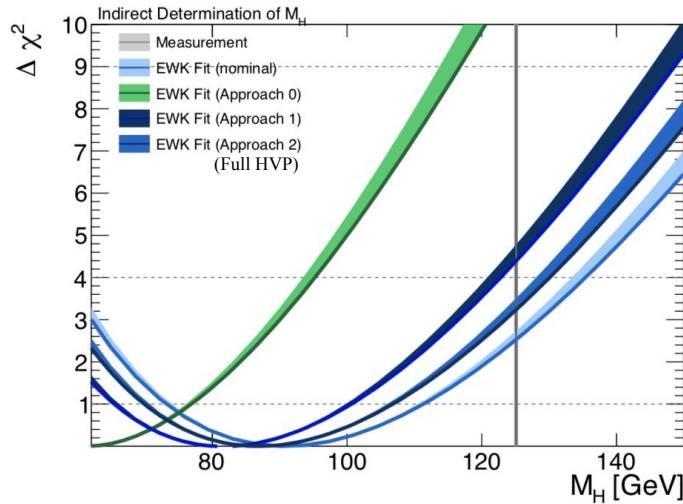
# EW fit inputs and $\chi^2$ results

LEP/LHC/Tevatron					
$M_Z$ [GeV]	$91.188 \pm 0.002$	$R_c^0$	$0.1721 \pm 0.003$	$M_H$ [GeV]	$125.09 \pm 0.15$
$\sigma_{\text{had}}^0$ [nb]	$41.54 \pm 0.037$	$R_b^0$	$0.21629 \pm 0.00066$	$M_W$ [GeV]	$80.380 \pm 0.013$
$\Gamma_Z$ [GeV]	$2.495 \pm 0.002$	$A_c$	$0.67 \pm 0.027$	$m_t$ [GeV]	$172.9 \pm 0.5$
$A_l$ (SLD)	$0.1513 \pm 0.00207$	$A_l$ (LEP)	$0.1465 \pm 0.0033$	$\sin^2 \theta_{\text{eff}}^l$	$0.2314 \pm 0.00023$
$A_{\text{FB}}^l$	$0.0171 \pm 0.001$	$m_c$ [GeV]	$1.27_{-0.11}^{+0.07}$ GeV	After HL-LHC	
$A_{\text{FB}}^c$	$0.0707 \pm 0.0035$	$m_b$ [GeV]	$4.20_{-0.07}^{+0.17}$ GeV	$M_W$ [GeV]	$80.380 \pm 0.008$
$A_{\text{FB}}^b$	$0.0992 \pm 0.0016$	$\alpha_s(M_Z)$	$0.1198 \pm 0.003$	$\sin^2 \theta_{\text{eff}}^l$	$0.2314 \pm 0.00012$
$R_l^0$	$20.767 \pm 0.025$	$\Delta\alpha_{\text{had}}^{(5)}(M_Z^2) [10^{-5}]$	$2760 \pm 9$	$m_t$ [GeV]	$172.9 \pm 0.3$

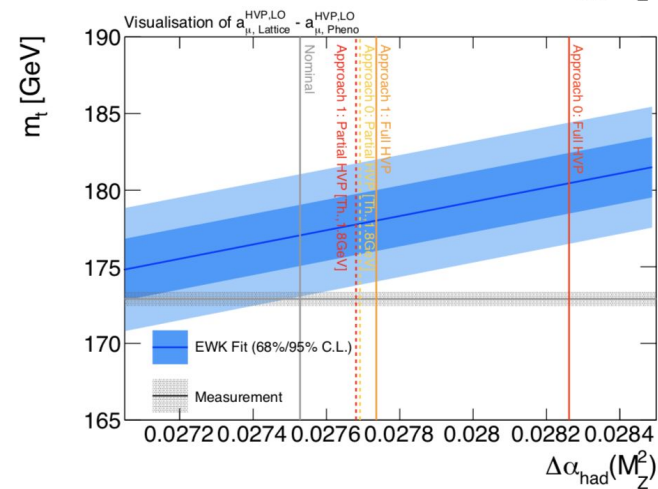
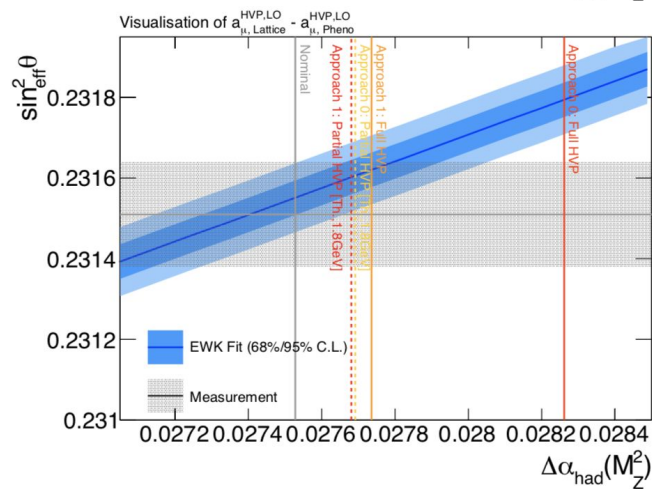
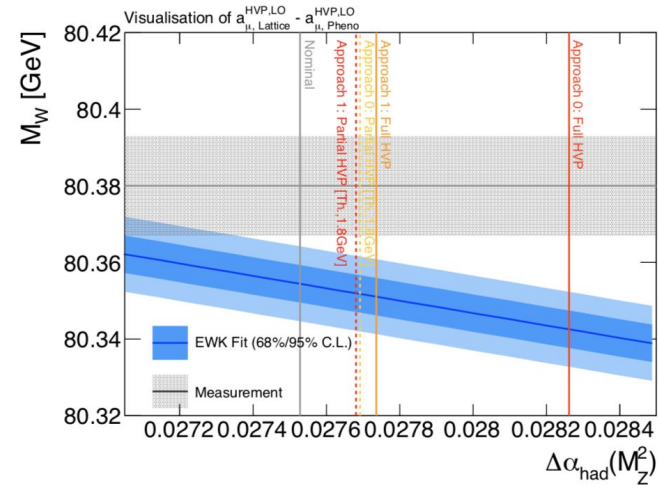
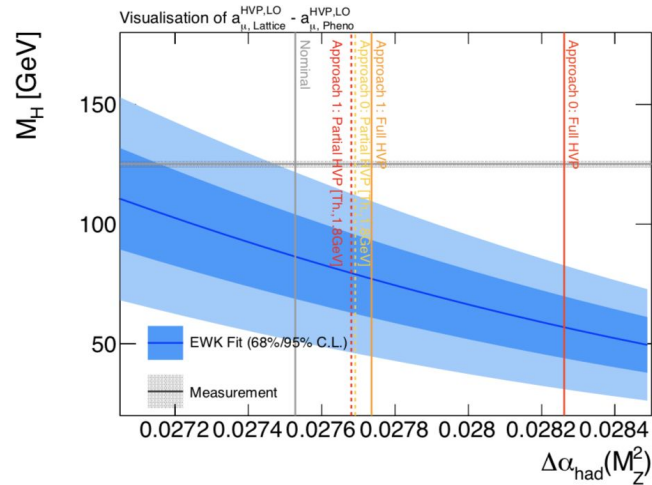
$a_\mu^{\text{HVP, LO shift}}$ (Energy range)	Nominal		Approach 0		Approach 1		Approach 2	
	$\Delta' \alpha_{\text{had}}(M_Z^2)$	$\chi^2/\text{ndf}$	$\Delta' \alpha_{\text{had}}(M_Z^2)$	$\chi^2/\text{ndf}$	$\Delta' \alpha_{\text{had}}(M_Z^2)$	$\chi^2/\text{ndf}$	$\Delta' \alpha_{\text{had}}(M_Z^2)$	$\chi^2/\text{ndf}$
	0.02753	18.6/16 (p=0.29)	-	-	-	-	0.02753	28.1/17 (p=0.04)
$a_\mu^{\text{HVP, LO (Lattice)}} - a_\mu^{\text{HVP, LO (Pheno)}}$ (Full HVP)	-	-	0.02826	27.6/16 (p=0.04)	0.02774	20.3/16 (p=0.21)	-	$\chi^2(\text{BMW20-Pheno}): 9.3$
$a_\mu^{\text{HVP, LO (Lattice)}} - a_\mu^{\text{HVP, LO (Pheno)}}$ ([Th.; 1.8 GeV])	-	-	0.02769	19.9/16 (p=0.22)	0.02768	19.8/16 (p=0.23)	-	-
$a_\mu^{\text{HVP, LO (Lattice)}} - a_\mu^{\text{HVP, LO (Pheno)}}$ ([Th.; 1.0 GeV])	-	-	0.02765	19.6/16 (p=0.24)	-	-	-	-
$a_\mu^{\text{Exp}} - a_\mu^{\text{SM (Pheno)}}$ (Full HVP)	-	-	0.02856	33.6/16 (p=0.01)	0.02782	21.2/16 (p=0.17)	-	-
$a_\mu^{\text{Exp}} - a_\mu^{\text{SM (Pheno)}}$ ([Th.; 1.8 GeV])	-	-	0.02776	20.6/16 (p=0.19)	0.02774	20.4/16 (p=0.20)	-	-
$a_\mu^{\text{Exp}} - a_\mu^{\text{SM (Pheno)}}$ ([Th.; 1.0 GeV])	-	-	0.02771	20.1/16 (p=0.22)	-	-	-	-



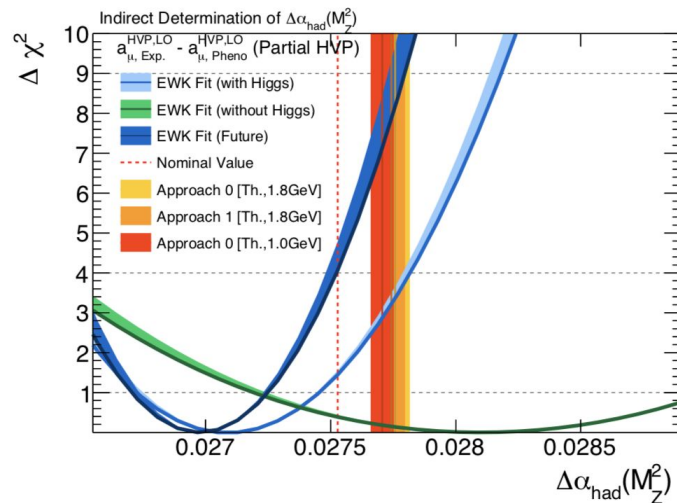
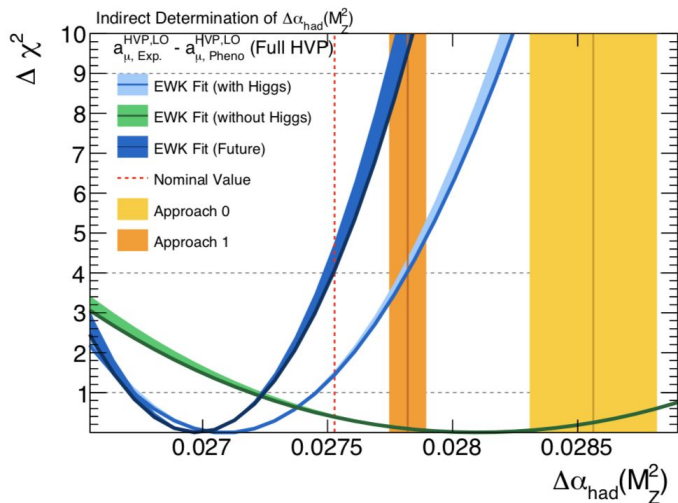
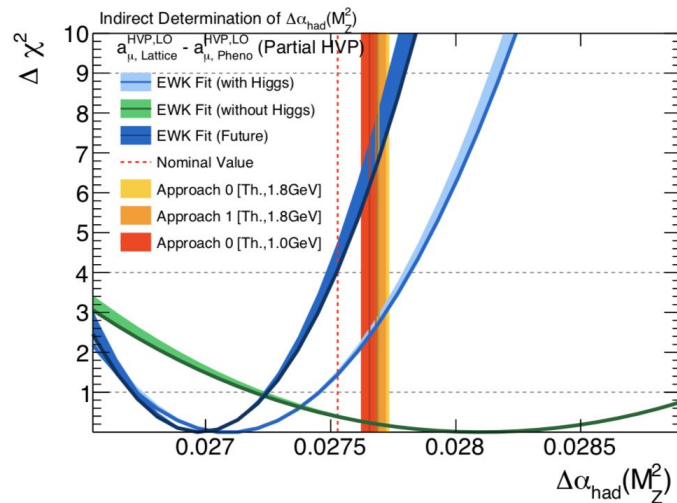
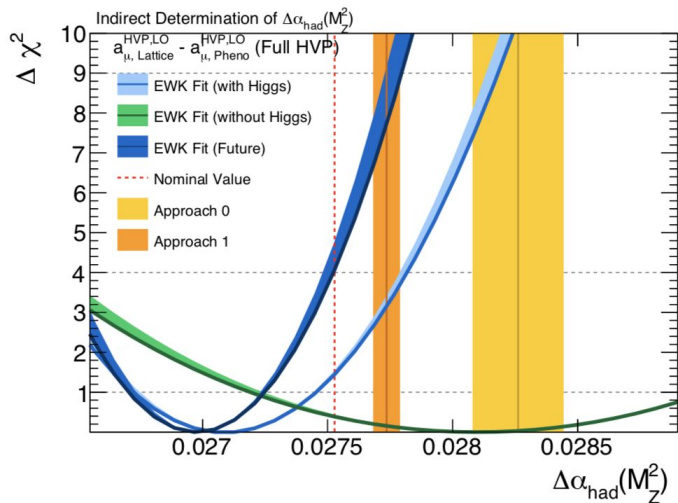
# EW fit results: $\chi^2$ scans



# EW fit results: parameter scans for varying $\Delta\alpha_{\text{had}}(M_Z^2)$



# EW fit results: indirect determination of $\Delta\alpha_{\text{had}}(M_Z^2)$

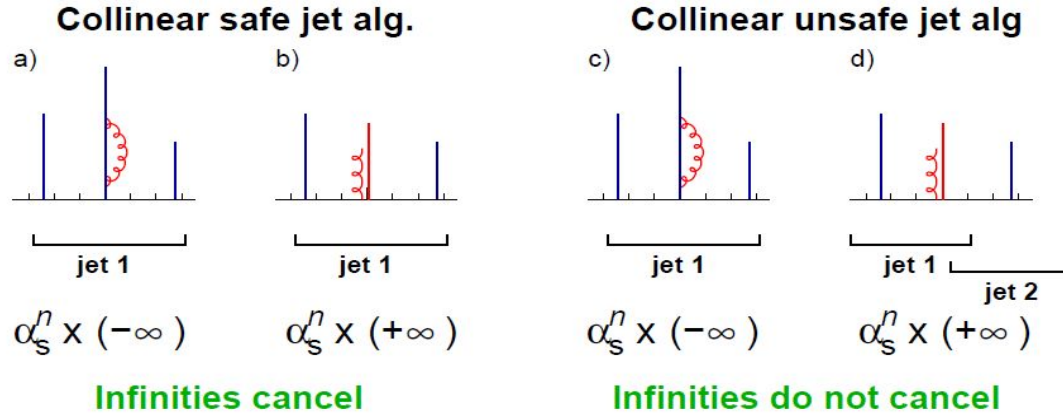


# Backup jet studies

# Jet definitions

- Main guidelines: infrared and collinear safety

(guaranteed in data, but important for the theoretical interpretation at all orders)



G. Salam  
0906.1833

- Most used algorithms:

→ Sequential recombination, using distance between objects: anti- $k_r$ ,  $k_r$ , Cambridge/Aachen

$$d_{ij} = \min(p_{ti}^{2p}, p_{tj}^{2p}) \frac{\Delta R_{ij}^2}{R^2}, \quad \Delta R_{ij}^2 = (y_i - y_j)^2 + (\phi_i - \phi_j)^2$$

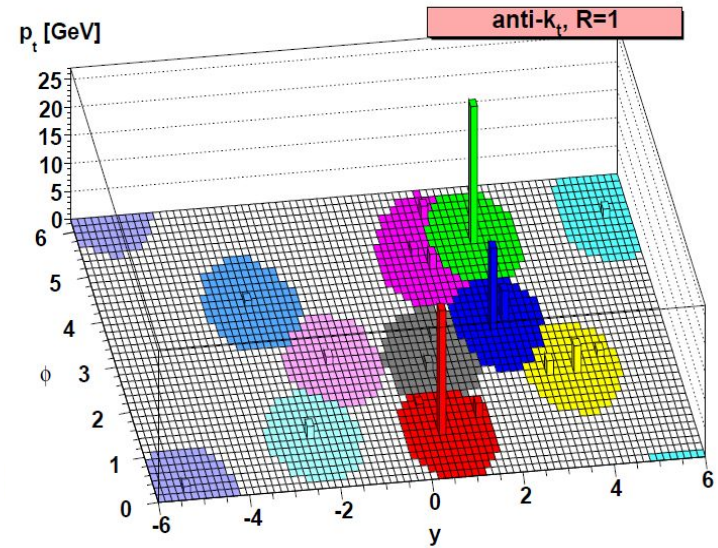
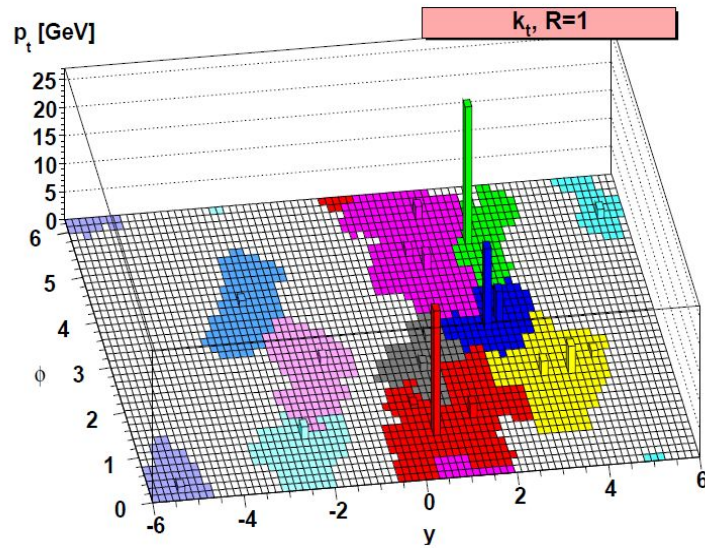
$$d_{iB} = p_{ti}^{2p},$$

→ Algorithms applicable coherently at the “truth” and reconstructed level

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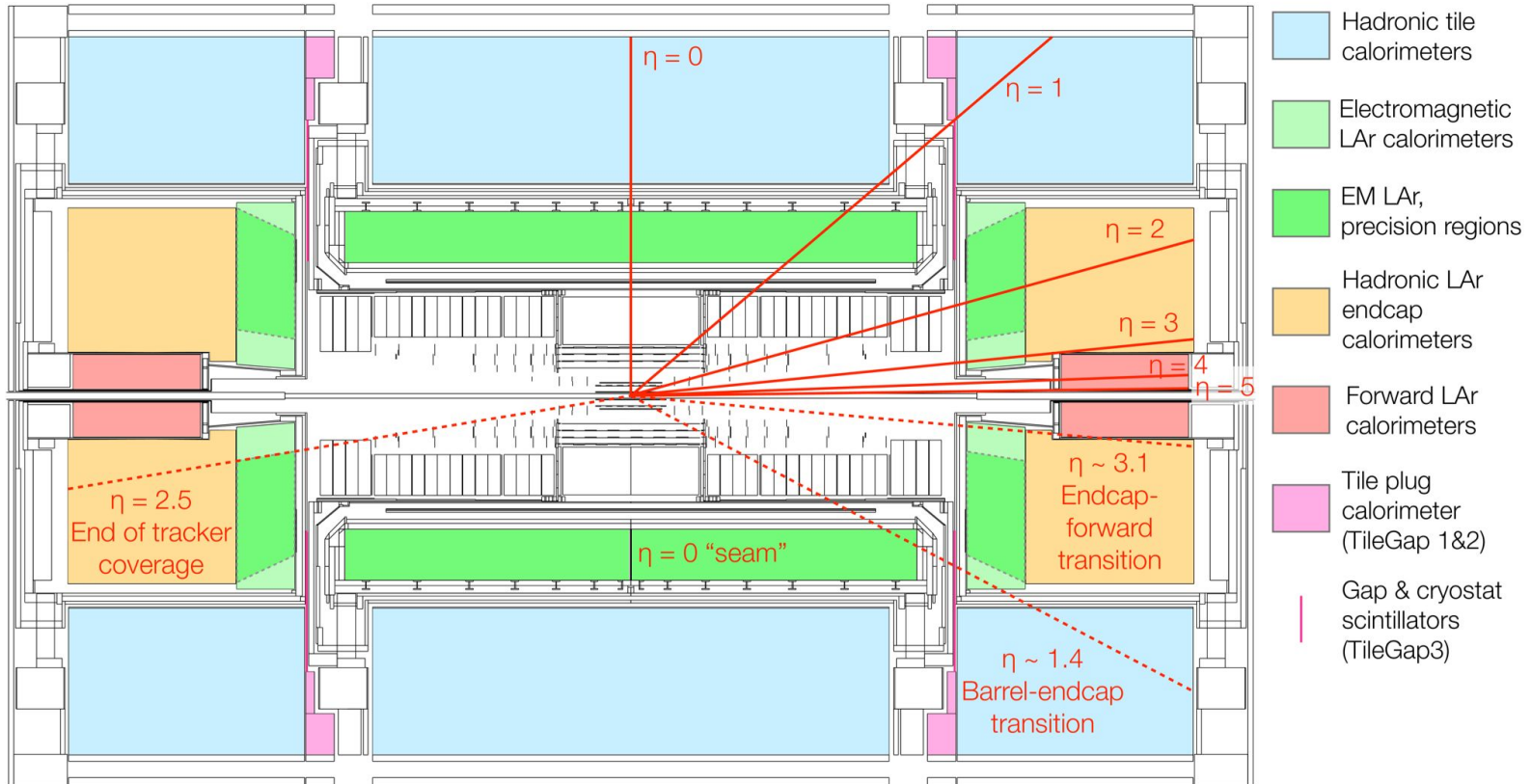
# Jet definitions

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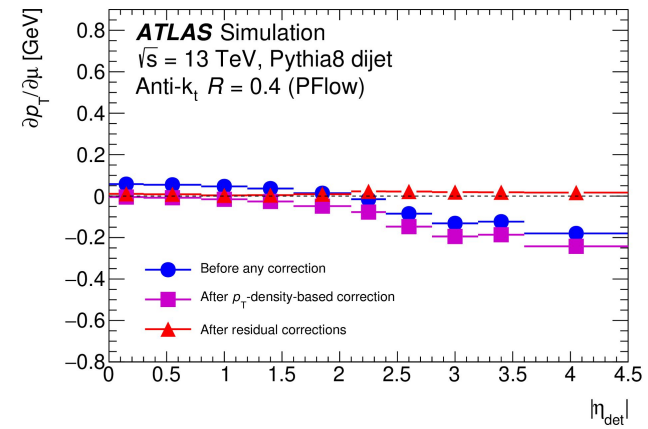
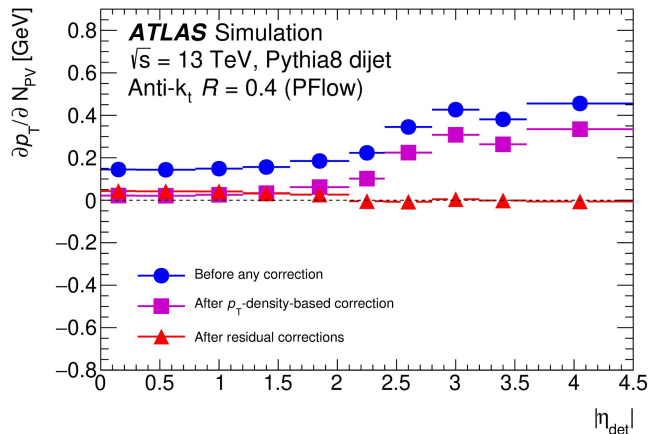
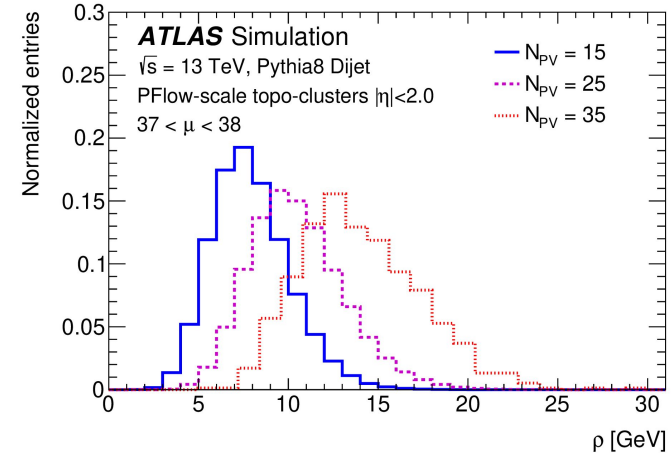
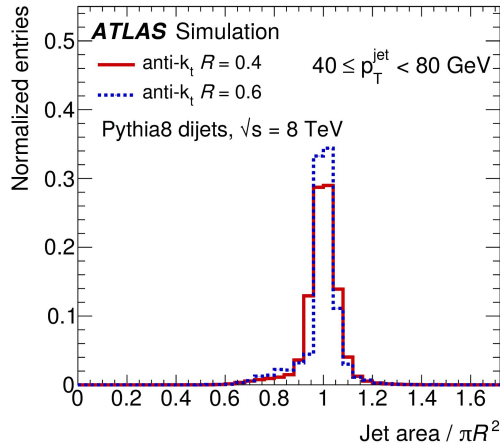
→  $\text{anti-}k_t$  jets have more circular shapes compared to  $k_t$

# ATLAS detector



# Pile-up correction(s)

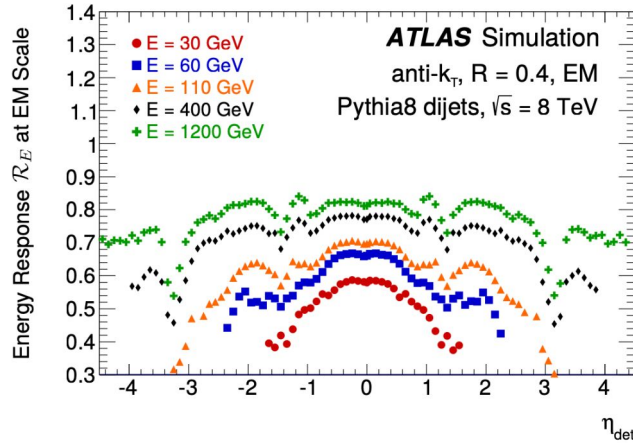
$$p_T \mapsto p_T - \rho A_T - \alpha (N_{PV} - 1) - \beta \mu$$



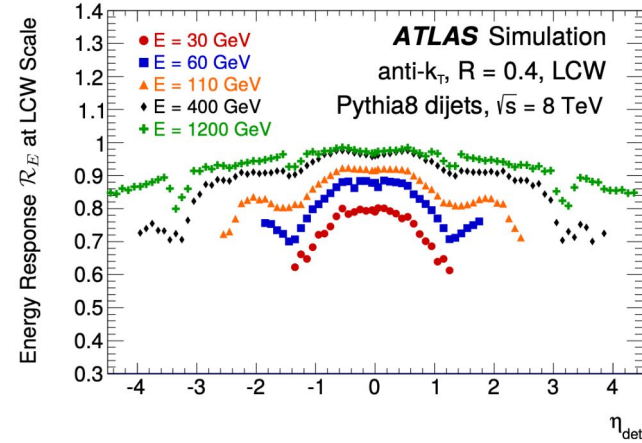
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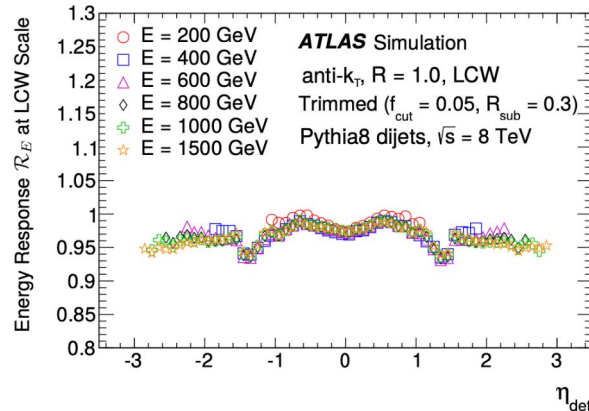
# MC-based jet energy and mass scales



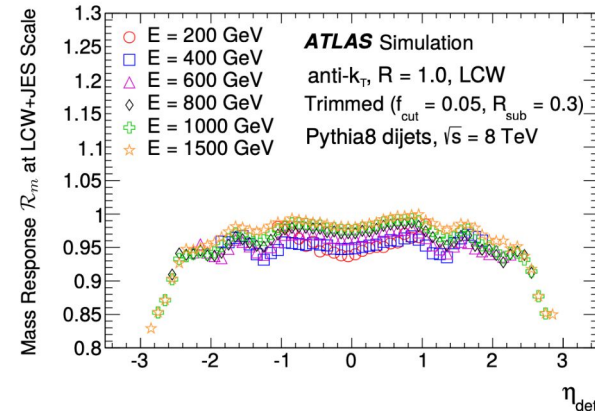
(a) Jet energy response  $\mathcal{R}_E$ , EM scale,  $R = 0.4$



(b) Jet energy response  $\mathcal{R}_E$ , LCW scale,  $R = 0.4$



(c) Jet energy response  $\mathcal{R}_E$ , LCW scale,  $R = 1.0$

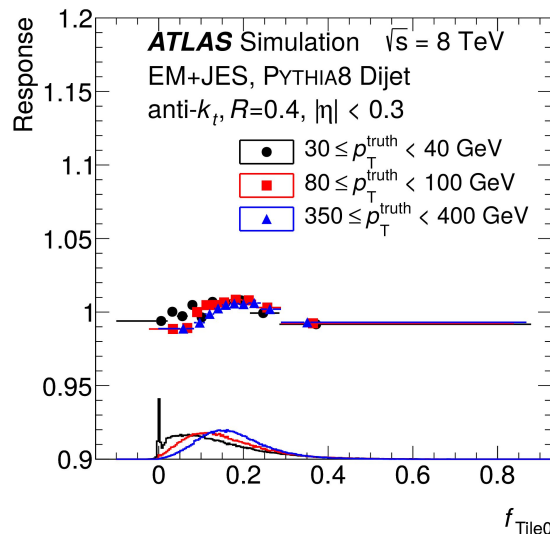
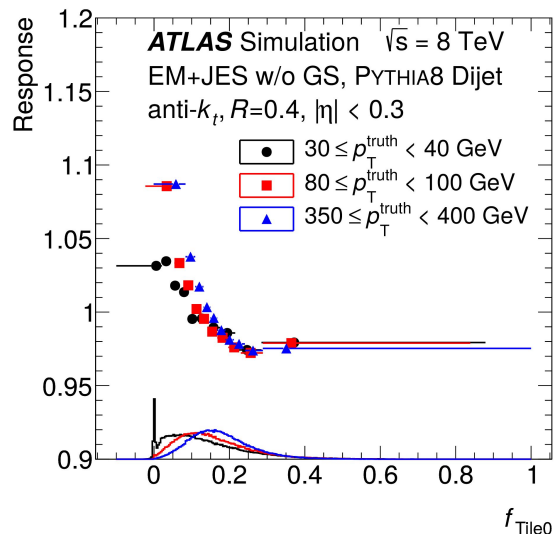


(d) Jet mass response  $\mathcal{R}_m$ , LCW scale,  $R = 1.0$

# GSC & GNNC – variables & sensitivity

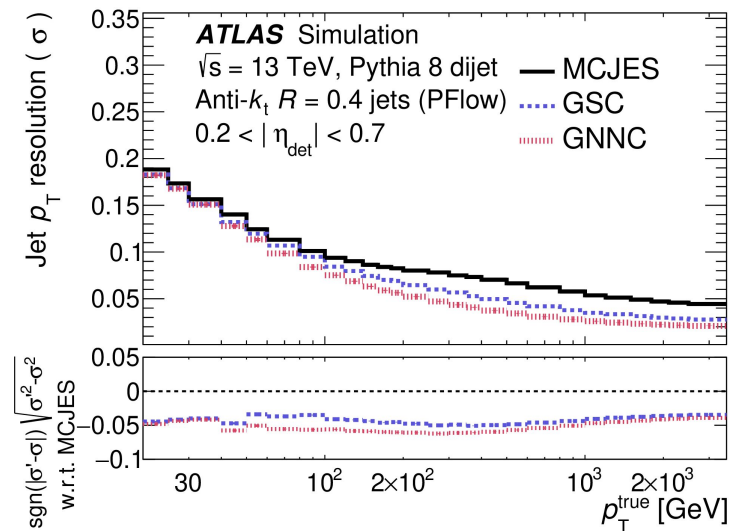
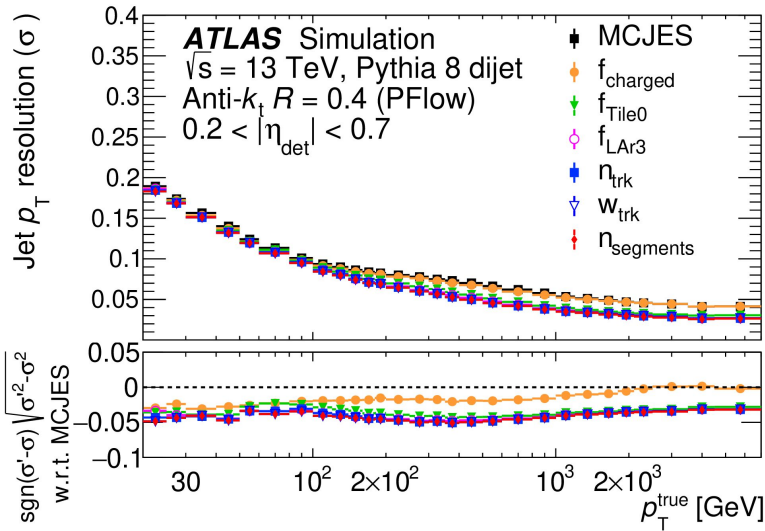
Calorimeter	$f_{\text{LAr},0-3}$	the $E_{\text{frac}}$ measured in the 0th-3rd layer of the EM LAr calorimeter
	$f_{\text{TILE0-2}}^*$	the $E_{\text{frac}}$ measured in the 0th-2nd layer of the hadronic tile calorimeter
	$f_{\text{HEC},0-3}$	the $E_{\text{frac}}$ measured in the 0th-3rd layer of the hadronic end cap calorimeter
	$f_{\text{FCAL},0-2}$	the $E_{\text{frac}}$ measured in the 0th-2nd layer of the forward calorimeter
	$N_{90\%}$	The minimum number of clusters containing 90% of the jet energy.
Jet kinematics	$p_{\text{T}}^{\text{JES}}^*$	The jet $p_{\text{T}}$ after the MCJES calibration
	$\eta_{\text{det}}$	The detector $\eta$
Tracking	$w_{\text{track}}^*$	the average $p_{\text{T}}$ -weighted transverse distance in the $\eta$ - $\phi$ plane between the jet axis and all tracks of $p_{\text{T}} > 1$ GeV ghost-associated with the jet
	$N_{\text{track}}^*$	the number of tracks with $p_{\text{T}} > 1$ GeV ghost-associated with the jet
	$f_{\text{charged}}^*$	the fraction of the jet $p_{\text{T}}$ measured from ghost-associated tracks
Muon segments	$N_{\text{segments}}^*$	the number of muon track segments ghost-associated with the jet
Pileup	$\mu$	The average number of interactions per bunch crossing
	$N_{\text{PV}}$	The number of reconstructed primary vertices

Table 1: List of variables used as input to the GNNC. Variables with a \* correspond to variables that are also used by the GSC.



→ Little / no dependence left for the response after having applied the corrections

# GSC & GNNC – performance



→ Improvement of the JER with the GSC and some further improvement using GNNC

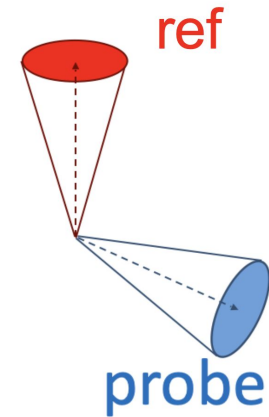
→ Work during Laura Boggia's QT, aiming to merge GNNC and the in-situ step

→ *Potential to eventually merge all the MC-based and in-situ steps*

# In-situ $\eta$ – intercalibration calibration method

→ Central reference method / matrix method (using more combinations of central-forward bins, obtaining hence a better statistical precision)

→ Deriving calibration factors in bins of  $p_T$ ,  $\eta$



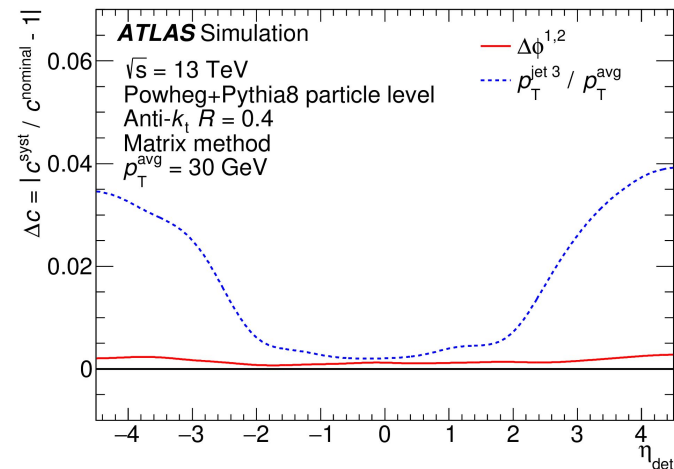
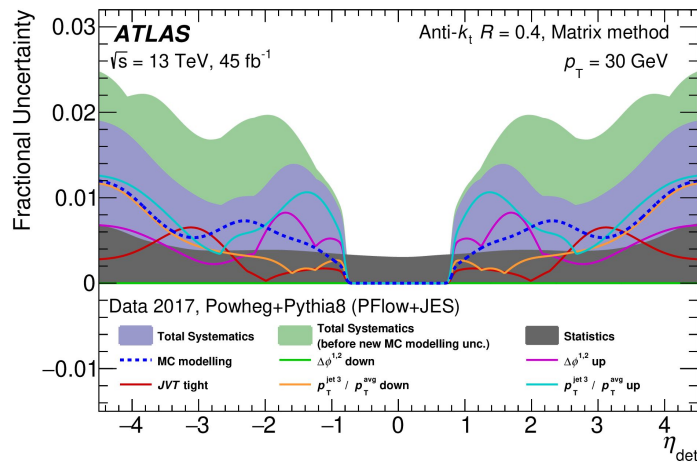
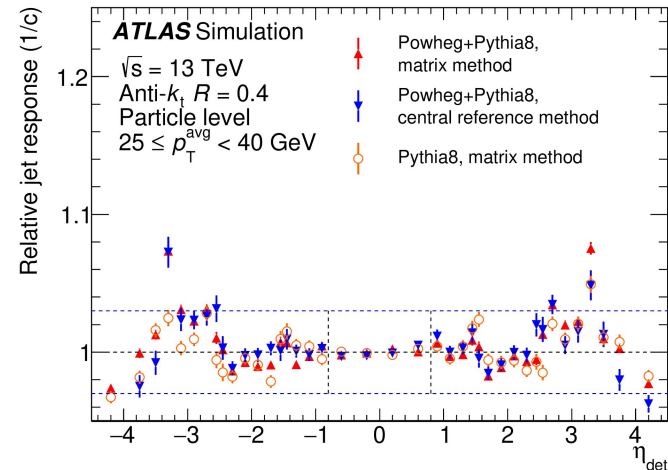
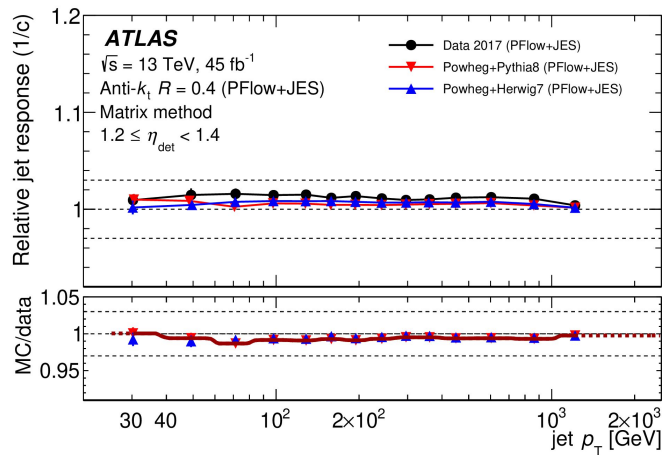
$$S(c_{1x}, \dots, c_{Nx}) = \sum_{j=2}^N \sum_{i=1}^{j-1} \left( \frac{1}{\Delta \langle \mathcal{R}_{ijx} \rangle} (c_{ix} \langle \mathcal{R}_{ijx} \rangle - c_{jx}) \right)^2 + X(c_{ix})$$

$$X(c_i) = \lambda \left( \frac{1}{N} \sum_{i=1}^N c_i - 1 \right)^2$$

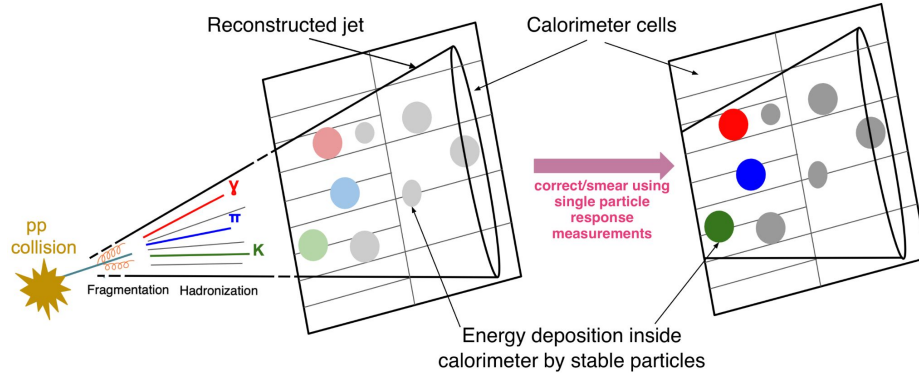
$$\sum_{i=1}^{\alpha-1} \left( \left( \frac{-\langle \mathcal{R}_{i\alpha} \rangle}{\Delta^2 \langle \mathcal{R}_{i\alpha} \rangle} + \frac{\lambda}{N^2} \right) c_i \right) + \left( \sum_{i=1}^{\alpha-1} \frac{1}{\Delta^2 \langle \mathcal{R}_{i\alpha} \rangle} + \sum_{i=\alpha+1}^N \frac{\langle \mathcal{R}_{\alpha i} \rangle^2}{\Delta^2 \langle \mathcal{R}_{\alpha i} \rangle} + \frac{\lambda}{N^2} \right) c_\alpha + \sum_{i=\alpha+1}^N \left( \left( \frac{-\langle \mathcal{R}_{\alpha i} \rangle}{\Delta^2 \langle \mathcal{R}_{\alpha i} \rangle} + \frac{\lambda}{N^2} \right) c_i \right) - \frac{\lambda}{N} = 0$$

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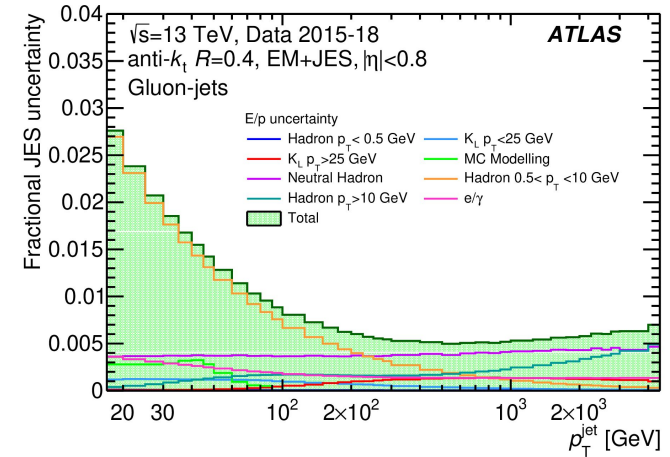
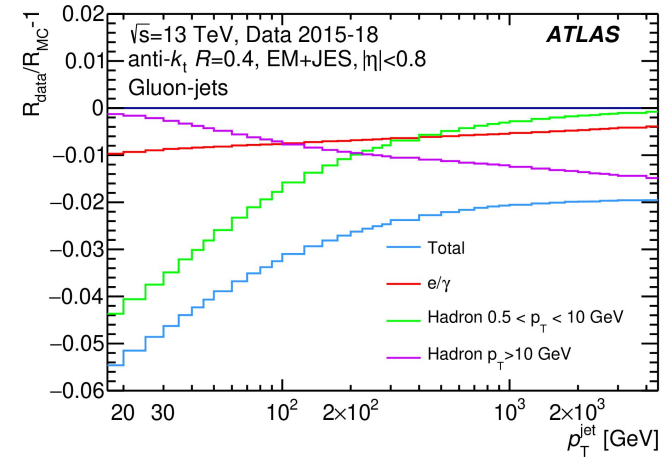
# In-situ $\eta$ – intercalibration calibration method



# E/p method



→ Important improvement of the E/p method  
(Lata Panwar's PostDoc)

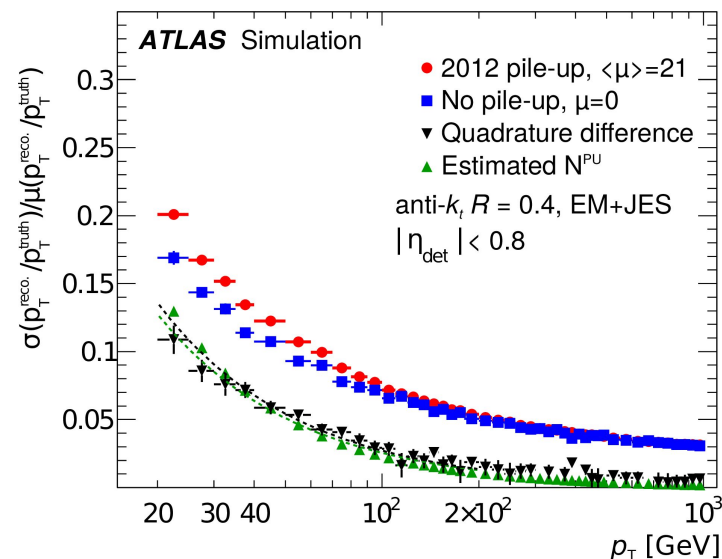


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# JER noise term

→ Constraint using random cones in zero-bias data and from distribution of soft jet momenta

→ Closure test for random cone method in MC



	EM $R = 0.4$	LCW $R = 0.4$	EM $R = 0.6$	LCW $R = 0.6$
$\langle \sigma_\rho \rangle$ ( $Z \rightarrow \mu\mu$ , data) [GeV]	1.81	3.25	1.81	3.25
$\langle \sigma_\rho \rangle$ ( $Z \rightarrow \mu\mu$ , MC) [GeV]	2.09	3.72	2.09	3.72
$\langle \sigma_\rho \rangle \sqrt{A}$ ( $Z \rightarrow \mu\mu$ , data) [GeV]	1.28	2.30	1.92	3.46
Random cone, data [GeV]	1.52	2.61	2.42	4.19
Difference [%]	16	12	21	17
$\langle \sigma_\rho \rangle \sqrt{A}$ ( $Z \rightarrow \mu\mu$ , MC) [GeV]	1.48	2.64	2.22	3.96
Random cone, MC [GeV]	1.60	2.73	2.61	4.49
Difference [%]	7.5	4.4	15	12

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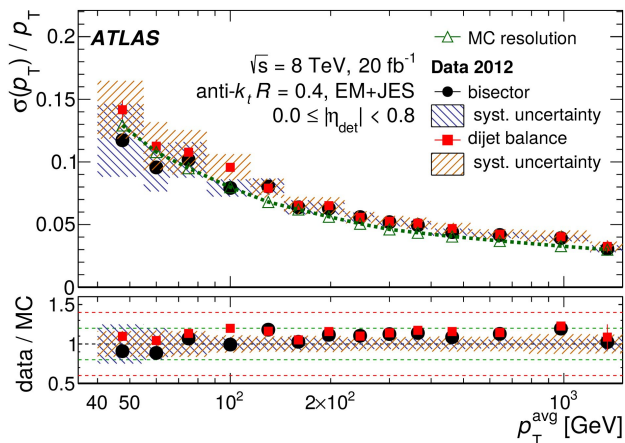
# Resolution measurement from dijet balance

## Dijet balance

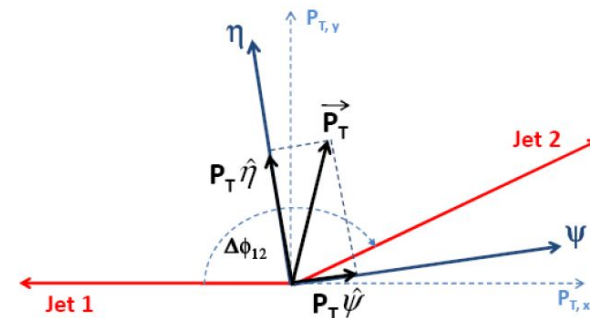
$$\mathcal{A} = \frac{p_T^{\text{probe}} - p_T^{\text{ref}}}{p_T^{\text{avg}}}$$

$$\sigma_{\mathcal{A}}^{\text{probe}} = \frac{\sigma_{p_T}^{\text{probe}} \oplus \sigma_{p_T}^{\text{ref}}}{p_T^{\text{avg}}} = \left\langle \frac{\sigma_{p_T}}{p_T} \right\rangle_{\text{probe}} \oplus \left\langle \frac{\sigma_{p_T}}{p_T} \right\rangle_{\text{ref}}$$

- MC-based subtraction of truth-level smearing
- Closure test in MC

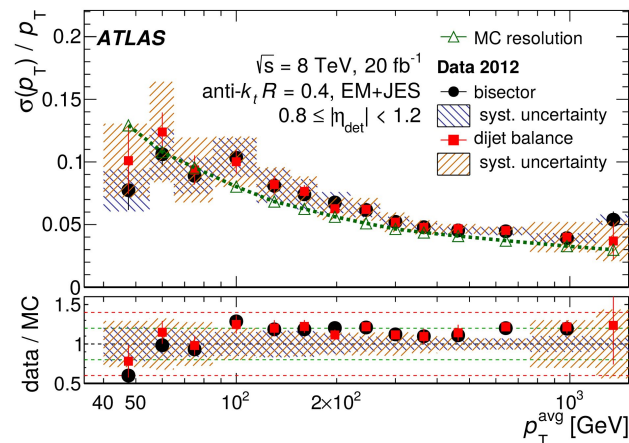


## Bisector



- Attempt to distinguish physics and detector effects

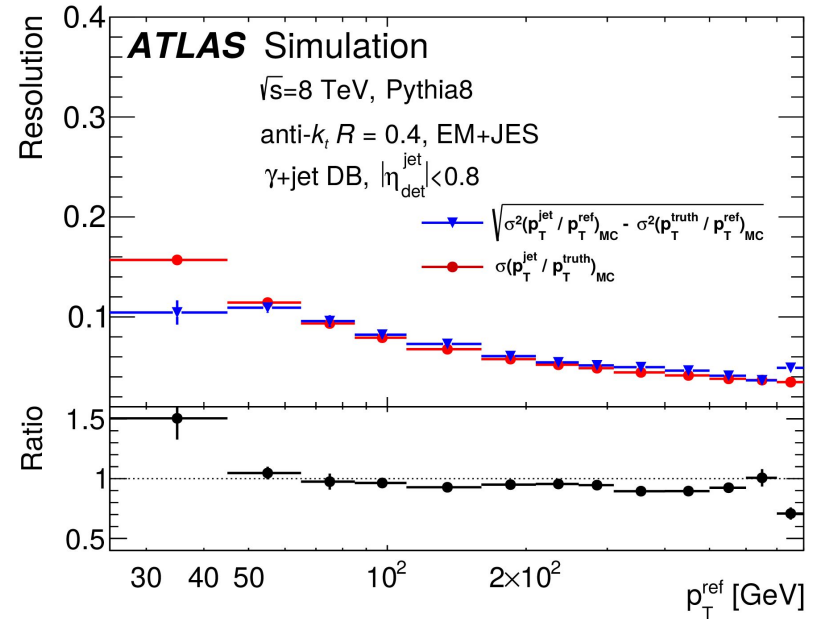
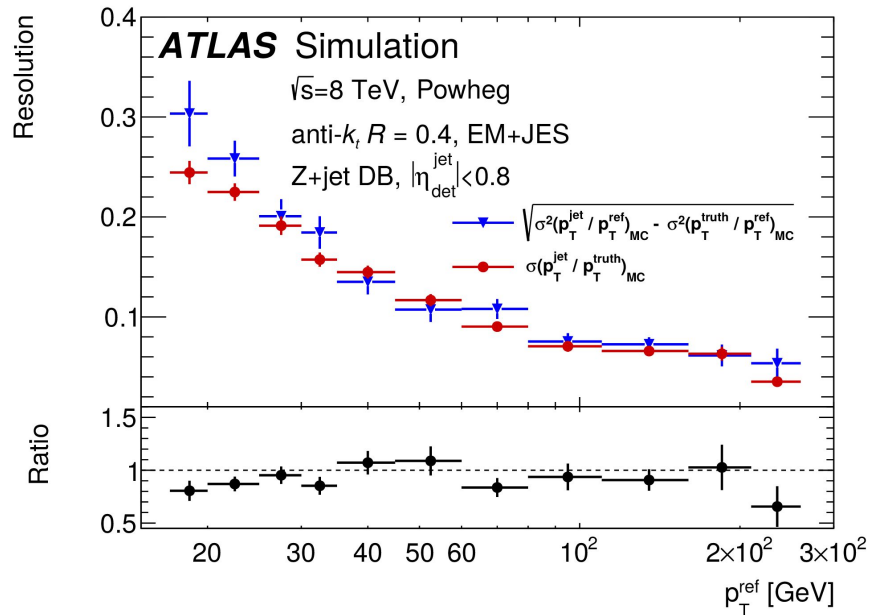
$$\frac{\sigma_{p_T}}{p_T} = \frac{\sigma_{\Psi} \Theta \sigma_{\eta}}{p_T \sqrt{2} \langle |\cos \Delta\phi_{12}| \rangle}$$



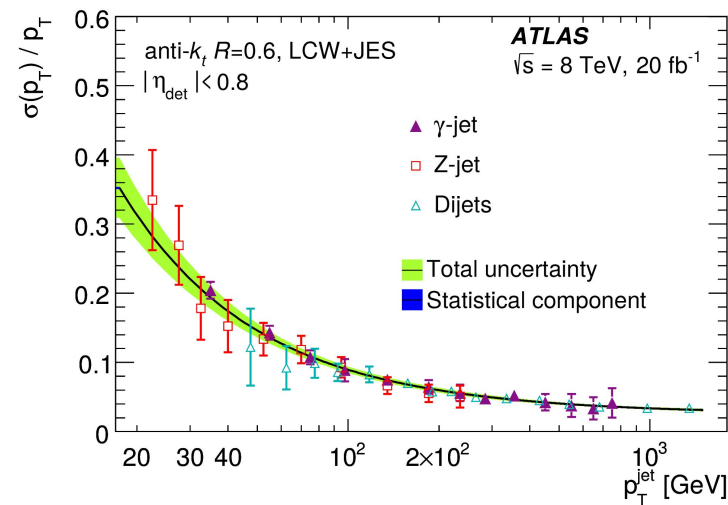
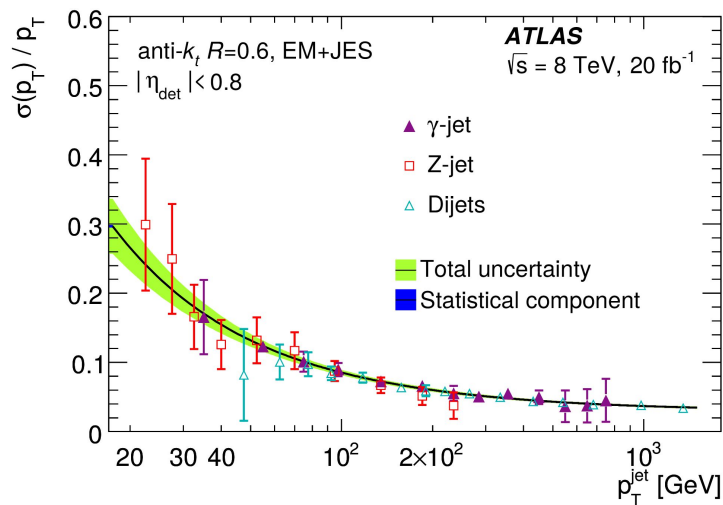
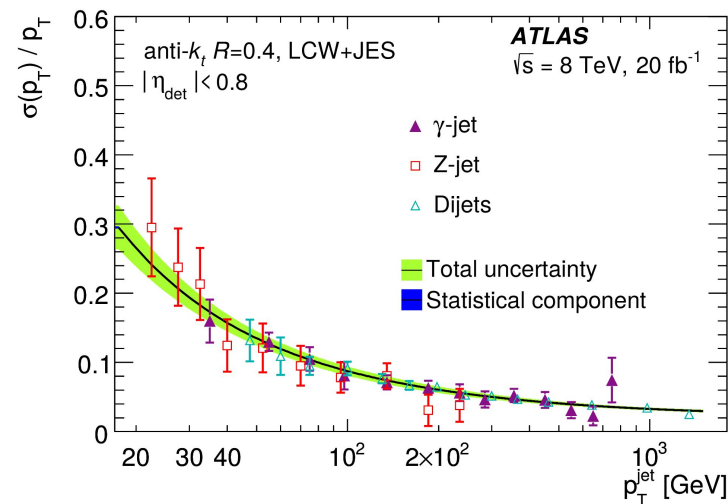
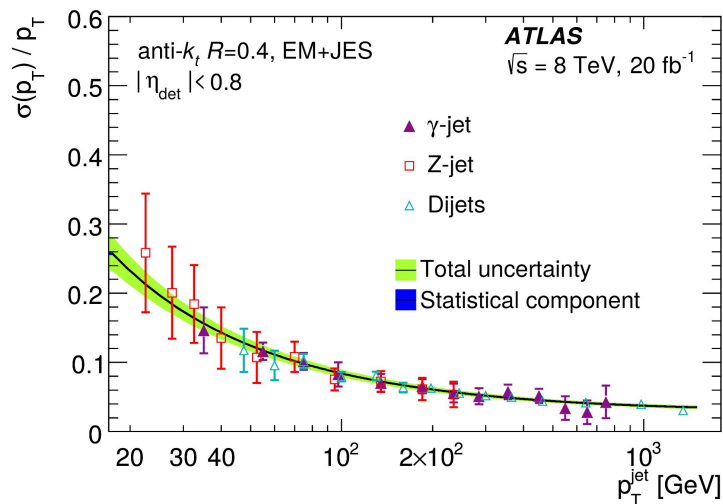


# V + jet: Evaluating the JER

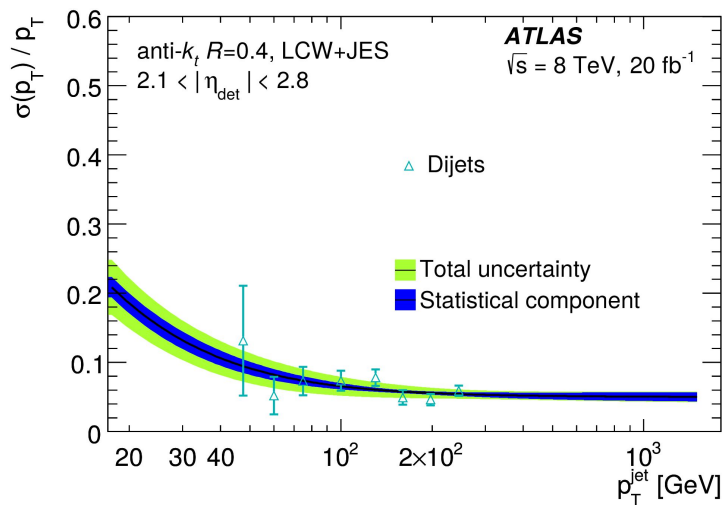
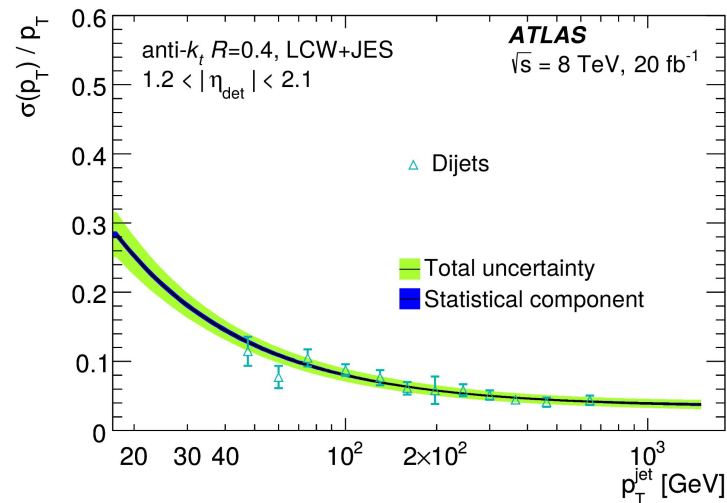
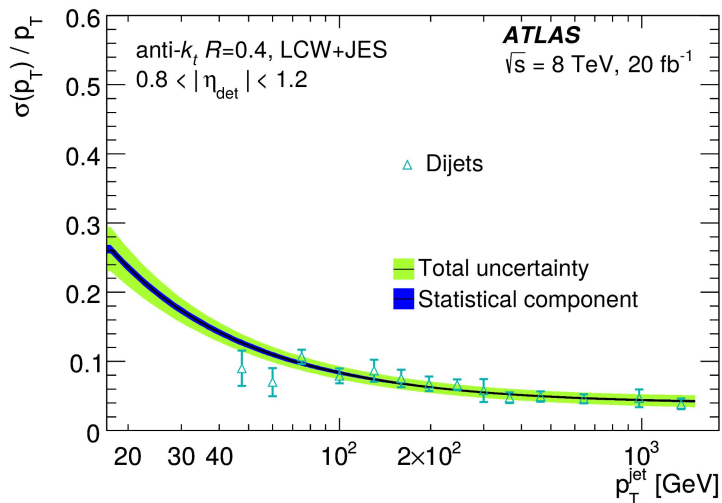
- MC-based subtraction of truth-level smearing
- Closure test in MC



# $\gamma$ -jet + Z-jet + Dijets + Zero Bias JER fit ( $\eta < 0.8$ )



# JER in-situ combination



# $\gamma$ -jet + Z-jet + Dijets + Zero Bias JER fit (EM JES, R=0.4, $\eta < 0.8$ )

- Propagation of in-situ uncertainties with toys and  $\pm 1\sigma$  shifts: good agreement between the two methods (except for two fits at large  $\eta$  - poor constraint on S & non-Gaussian distribution)
- $\chi^2$  definition does not include uncertainty on N (only uncertainties on data points are included)  
Small ( $\sim 1$ -2 units) changes in  $\chi^2$  value for  $N \pm 0.632$  GeV
- Uncertainty band not rescaled by global  $\chi^2/\text{ndof}$  (recall: local rescaling for JES)

In-situ uncertainties (N fixed) :

$$S = 0.713 \pm 0.067 \sqrt{\text{GeV}}$$

$$C = 0.030 \pm 0.003$$

$$\text{Corr}(S; C) = -0.25 \text{ (in-situ)}$$

Zero-bias:

$$N = 3.325 \pm 0.632 \text{ GeV}$$

$$N : 0.632 ; S : -0.038 ; C : 0.001$$

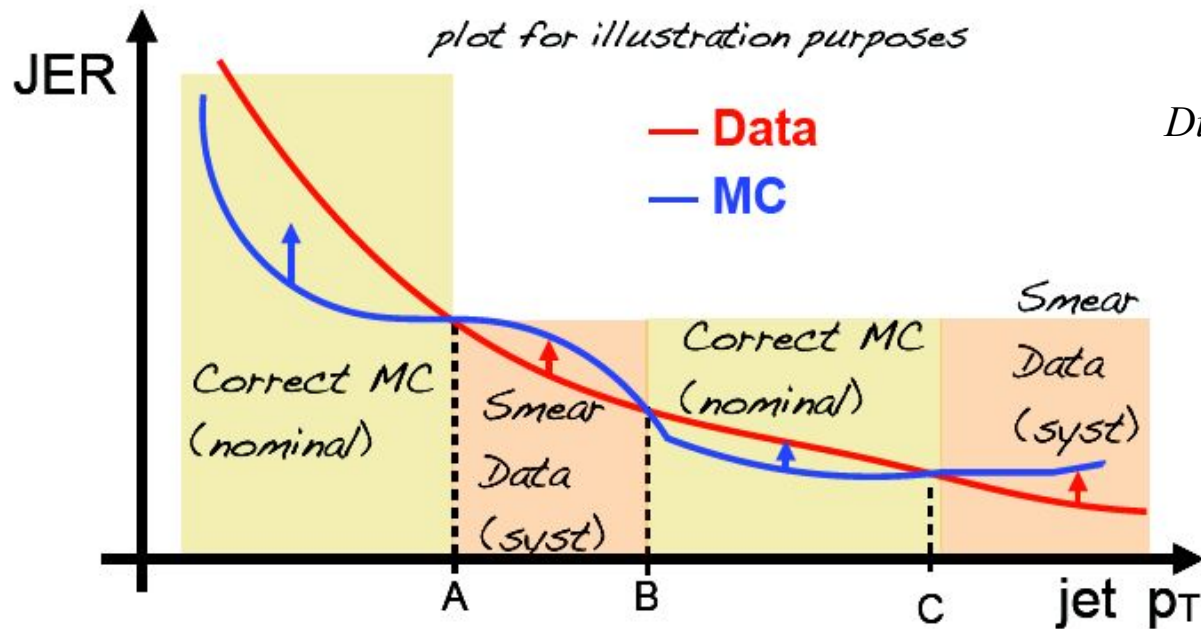
$$N : -0.632 ; S : 0.030 ; C : -0.001$$

$$\chi^2_{\text{diag}}/\text{ndof} = 8 / 35$$

$$\chi^2_{\text{corr}}/\text{ndof} = 71 / 35$$

→ Some indication of higher order contributions to JER function

# Propagation of JER uncertainties and correlations

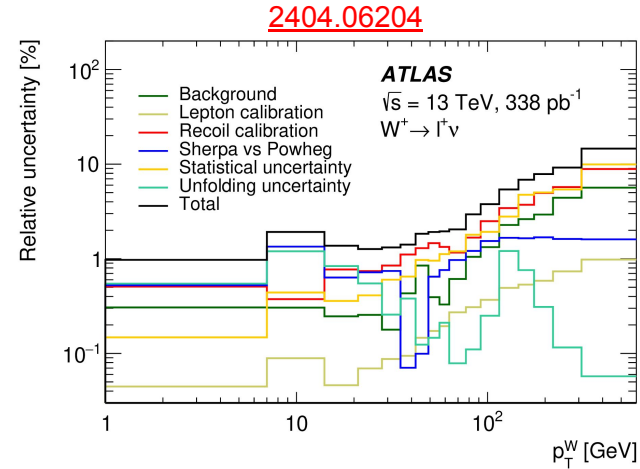
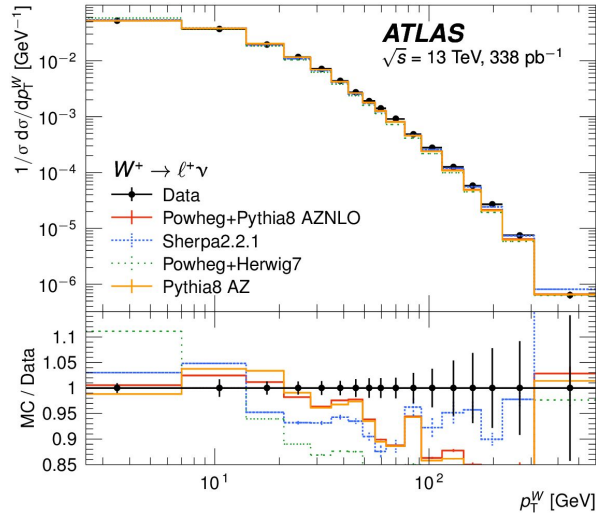


Dimitris Varouchas - PostDoc

- **Nominal value**
  - ♦ Correct MC, if  $JER_{Data} > JER_{MC}$
- **Systematic variation (for the uncertainty propagation only)**
  - ♦ Smear Data, if  $JER_{Data} < JER_{MC}$

# Some challenging unfolding examples

→  $p_T(W)$ : large resolution effects for MET reconstruction & need relatively fine binning in order to discriminate among theoretical predictions



→ Unfolding in a different context:

inverse Laplace transform to convert spacelike lattice QCD results into timelike quantities

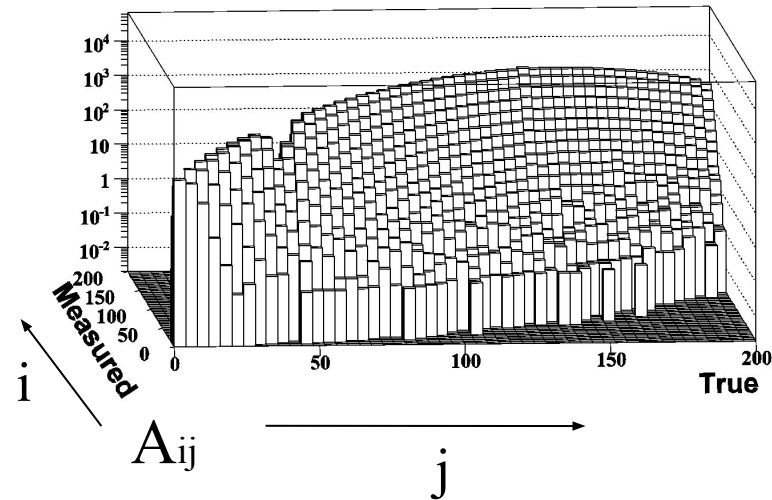
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# An Iterative, Bayes-inspired Unfolding Method

$$P_{ij} = \frac{A_{ij}}{\sum_{k=1}^{n_d} A_{kj}}$$

$$\tilde{P}_{ij} = \frac{A_{ij}}{\sum_{k=1}^{n_u} A_{ik}} ; u = \tilde{P} \cdot d$$

→ Note:  $\tilde{P}_{ij}$  depends on the shape of the truth distribution in MC



- 1<sup>st</sup> unfolding, where the original transfer matrix is used

→ 1) Transfer matrix improvement (hence of the unfolding probability matrix)

Reweight the truth MC distribution based on previous unfolding result.

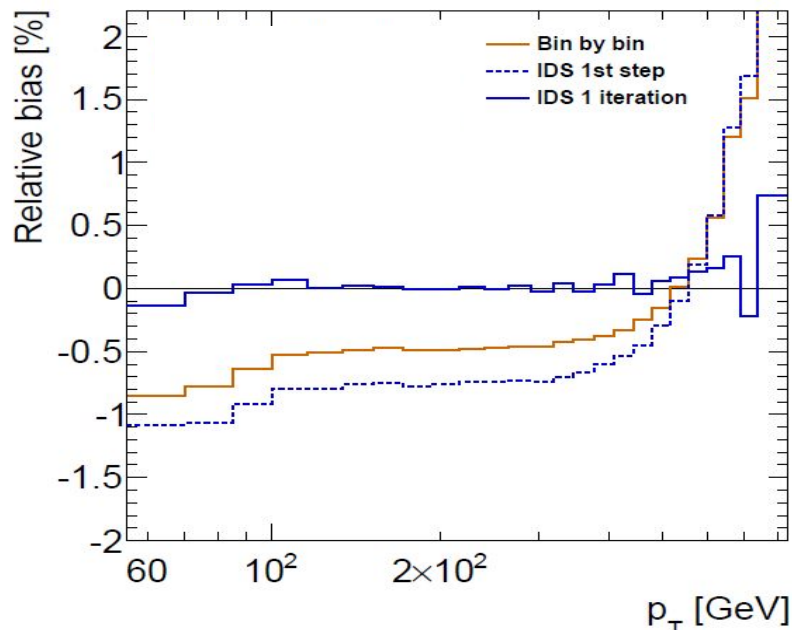
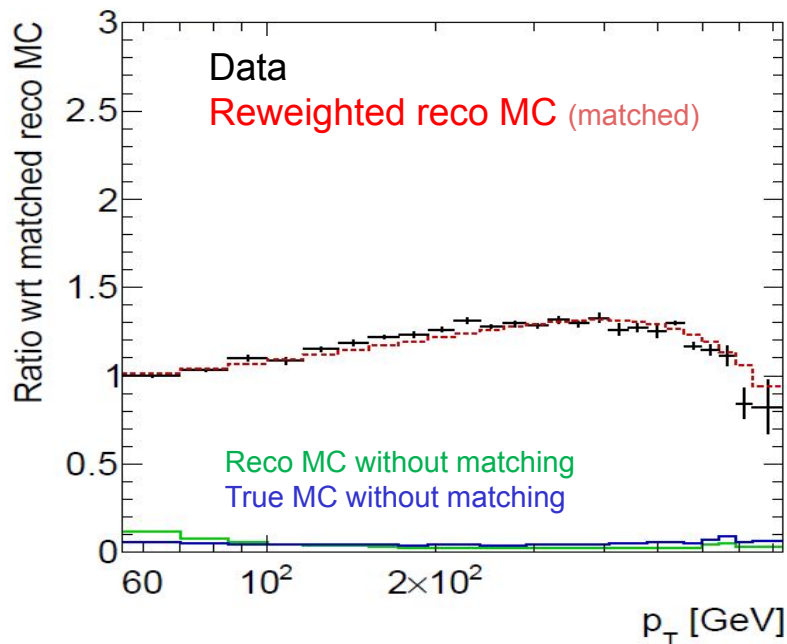
→ 2) Improved unfolding

→ Choice on number of iterations = regularization

→ Other methods exist, like e.g. dynamical local regularization in IDS (treatment of fluctuations in each bin, at each step of the procedure)

# Data-driven closure test: motivation, procedure, example

- In-situ determination of the unfolding uncertainty related to the data/MC shape difference and to the regularization :
- reweight true MC by smooth function: improved data/recoMC agreement
  - unfold the reweighted reconstructed MC
  - compare with reweighted true MC



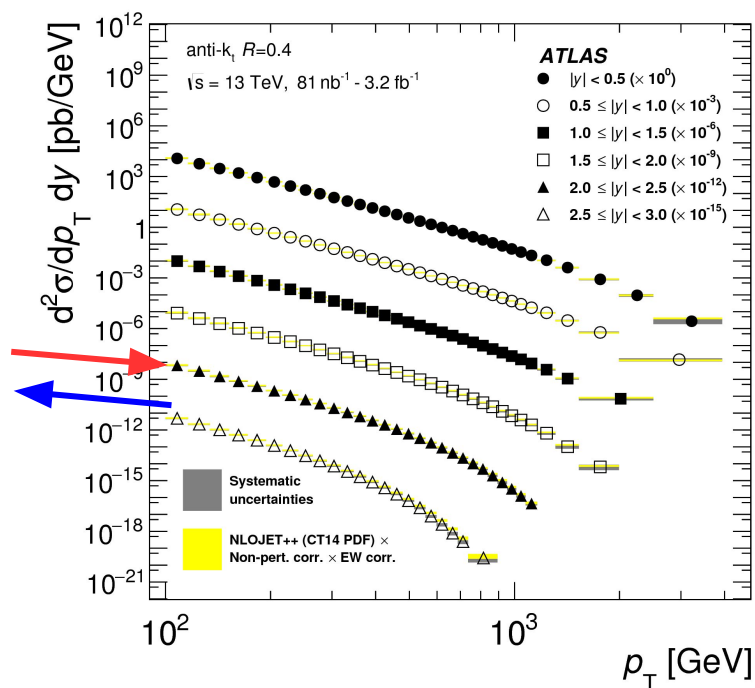
→ Applicable in cases without very different degenerated solutions (see eigenvalues of folding matrix, quality of the data/reweighted MC etc.). In other cases allows to learn about the ill-posedness of the problem

→ Method introduced in [arXiv:0907.3791](https://arxiv.org/abs/0907.3791), used in [arXiv:1112.6297](https://arxiv.org/abs/1112.6297) etc. ... [arXiv:2405.20041](https://arxiv.org/abs/2405.20041) (Omnifold 24-d) ...



# Choice of the phase-space

- Selection defining phase-space at “truth” level – as close as possible to the reconstructed-level selection:  
*minimize extrapolation to reduce model dependence*
- Include over-/under-flow bins when migrations to the region of interest are relevant  
→ These extra bins are generally not published



# Iterative methods: choice of the number of iterations

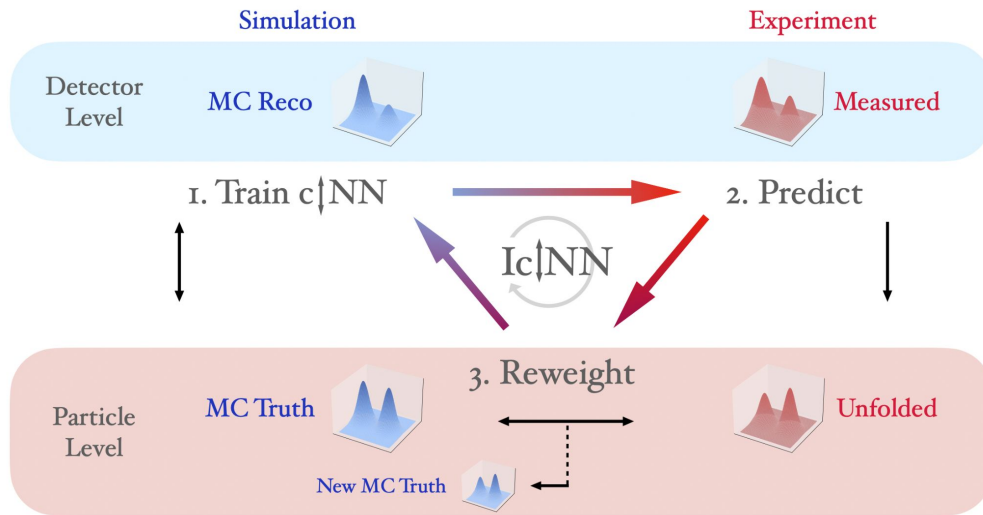
- Number of iterations = regularization parameter: optimising variance / bias
- Take into account: *systematic uncertainty related to the unfolding method (bias due to MC/data shape difference & regularization)*; impact on statistical uncertainties & correlations; constraints induced on binning choice
- Compare data and the modified reconstructed MC: see how much information is left to be propagated from the data shape to the truth MC shape
  - bin-by-bin comparison or using a  $\chi^2$  (see e.g. [arxiv:0907.3791](https://arxiv.org/abs/0907.3791), [arXiv:2404.06204](https://arxiv.org/abs/2404.06204))
- Suggestion in IBU publication: compare results from consecutive steps ([NIM A 362, 487 \(1995\)](#))
  - risk of  $\sim$ small changes between consecutive steps, while having a significant bias

# Tests of the unfolding

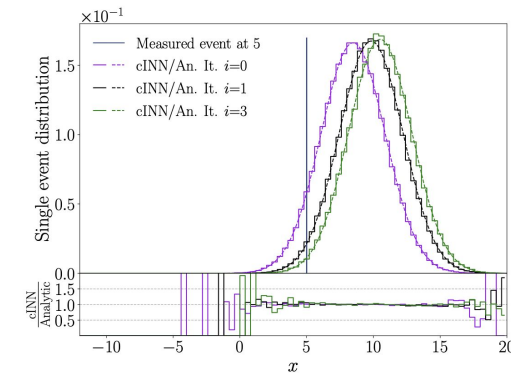
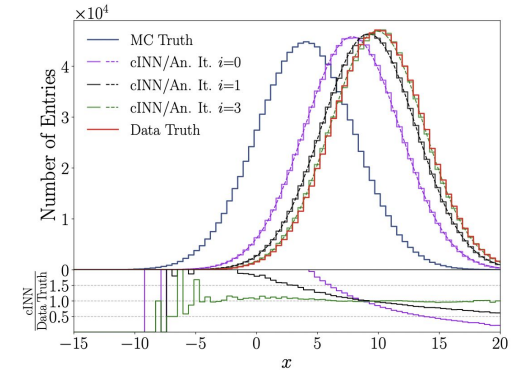
- “Technical closure test” → same MC for the transfer matrix and input distribution (pseudo-data) - expect perfect agreement between unfolding result and truth MC
- “Data-driven closure test” → allows to evaluate a systematic related to the unfolding method and the choice of regularization (see next slides)
- “Linearity test” → MC samples with various truth inputs; check linear dependence between unfolded and truth values of a quantity of interest
- “Pull test” → relevant only for unfolding methods providing an estimate of the statistical uncertainties (i.e. not from pseudo-experiments) - tests their reliability

# ML-based unfolding

- ML-based methods allow to enhance the dimensionality & obtain results event-by-event: enables computing secondary quantities [arXiv:2109.13243](https://arxiv.org/abs/2109.13243)
- IcINN: iteratively improve (reweight) MC simulation; publish unfolded distributions for each data event



[arXiv:2212.08674](https://arxiv.org/abs/2212.08674)

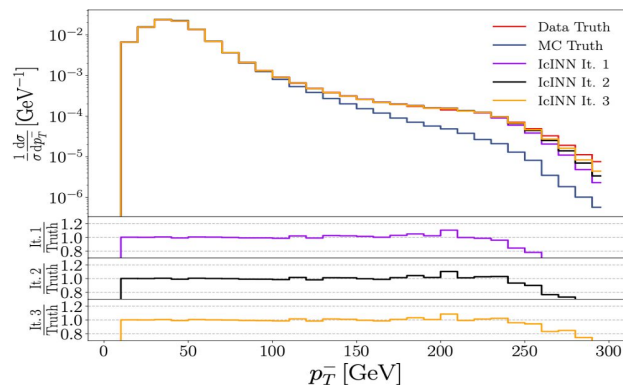
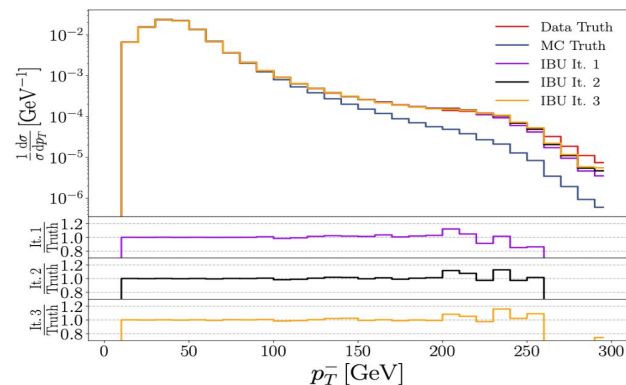


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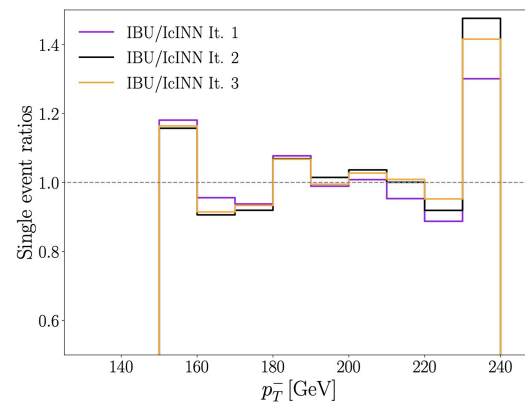
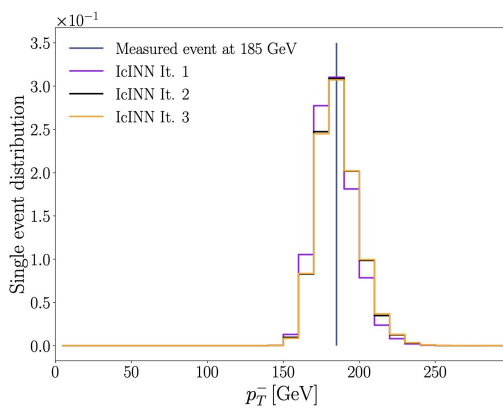
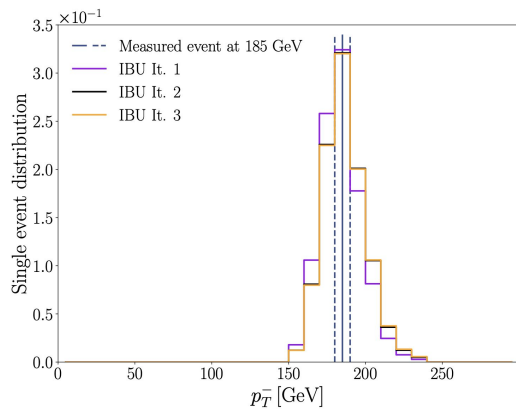
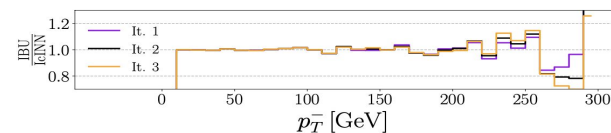
# Comparison of Transfer Matrix- / ML-based unfolding

→ Comparison typically performed for the full unfolded distributions

→ *New*: even-by-event comparison [arXiv:2310.17037](https://arxiv.org/abs/2310.17037)



$$pp \rightarrow Z\gamma\gamma, Z \rightarrow \mu^-\mu^+$$

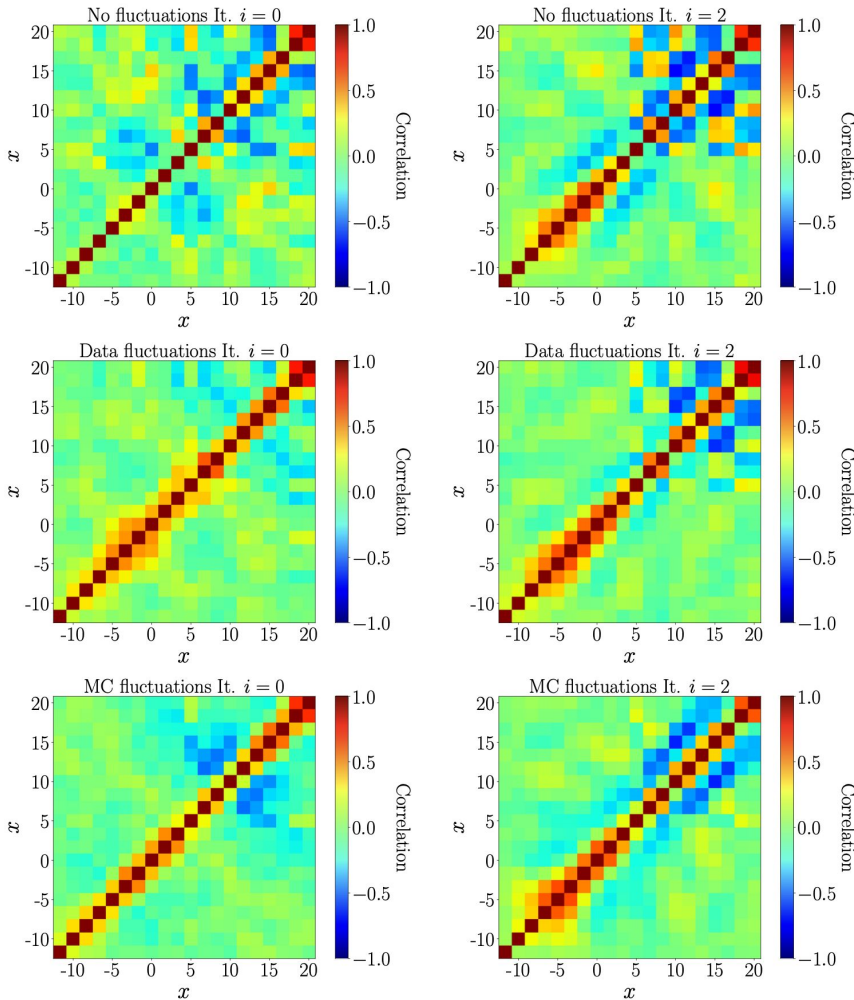


# Comparison of Transfer Matrix- / ML-based unfolding

→ Bootstrap method implemented for IcINN, for small number of observables (GPU challenge for training):

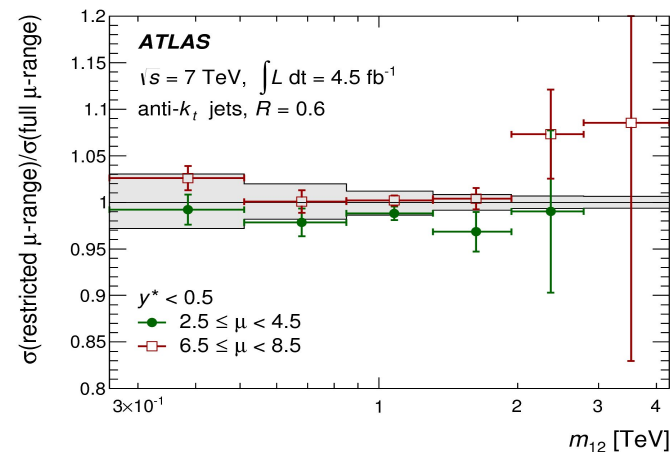
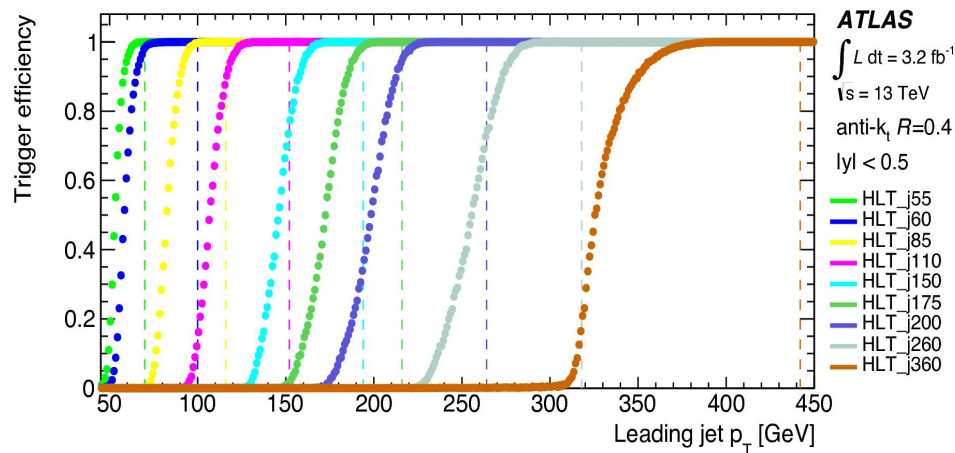
*Evaluate statistical uncertainties and correlations*

[arXiv:2212.08674](https://arxiv.org/abs/2212.08674)



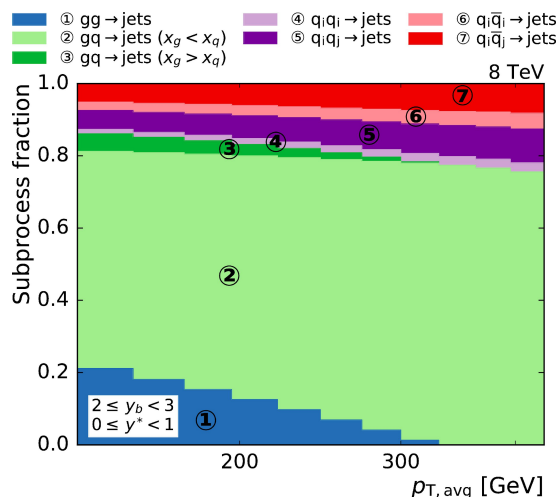
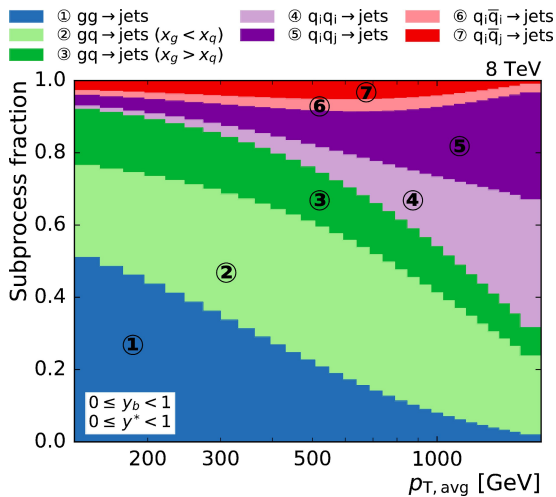
# Trigger and pile-up

- Trigger prescales and pile-up treatment take into account variations in data-taking conditions
- Jet trigger efficiencies determined in-situ using unbiased samples
- Each trigger used in the region where it is fully efficient

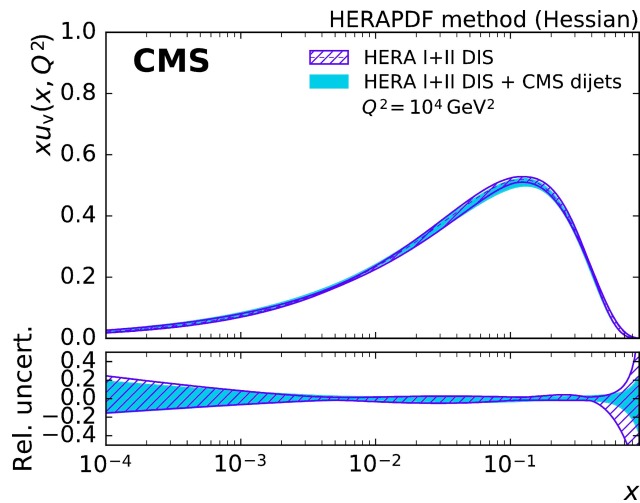
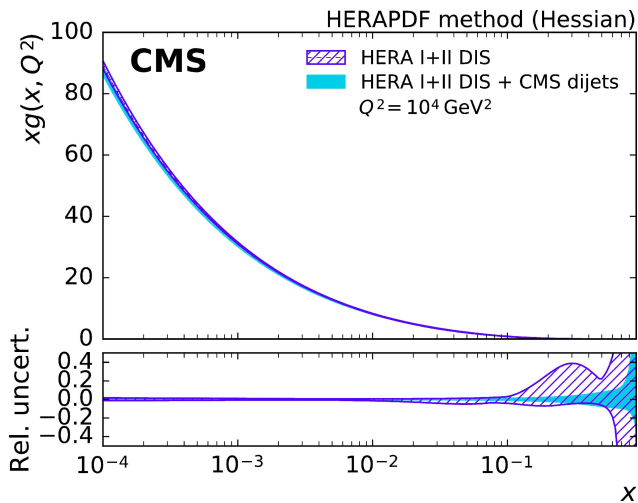


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# Dijet 3D measurement - CMS



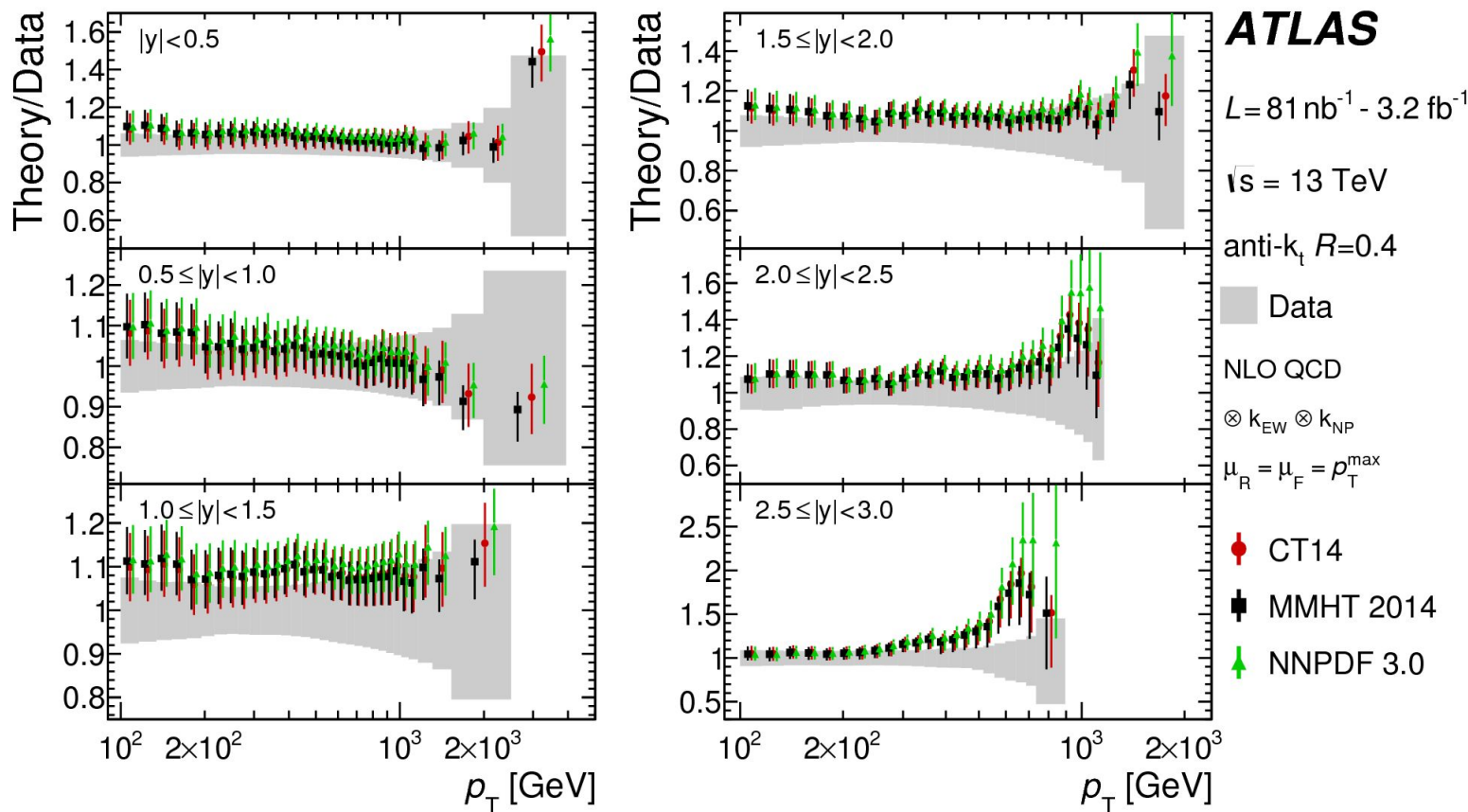
→ Contributions of various processes in different phase-space regions: sensitivity to PDFs





# Inclusive jet cross sections: theory/data

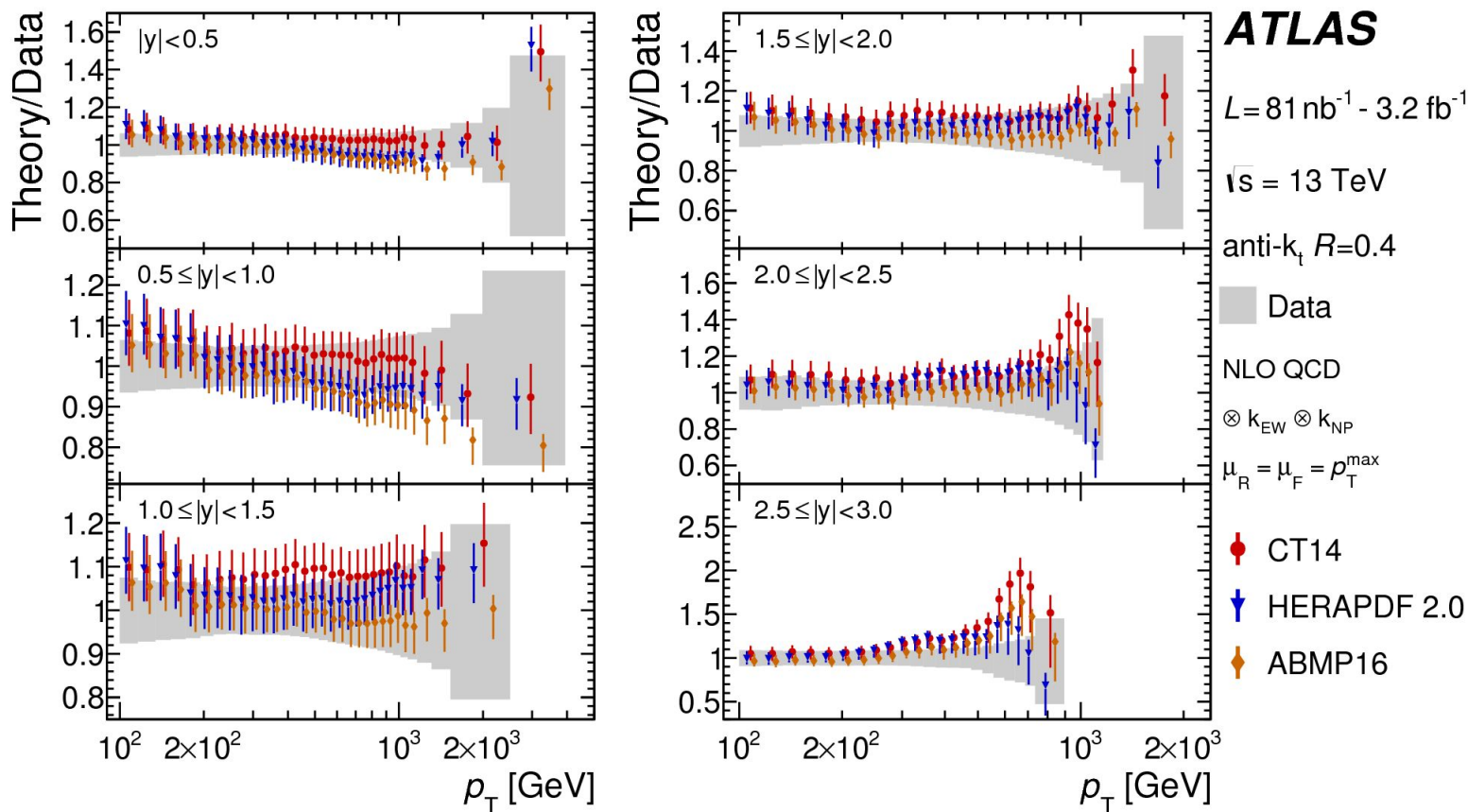
→ Good data/theory agreement within uncertainties observed for most PDF sets:  
CT14, MMHT 2014, NNPDF 3.0, HERAPDF 2.0, ABMP16



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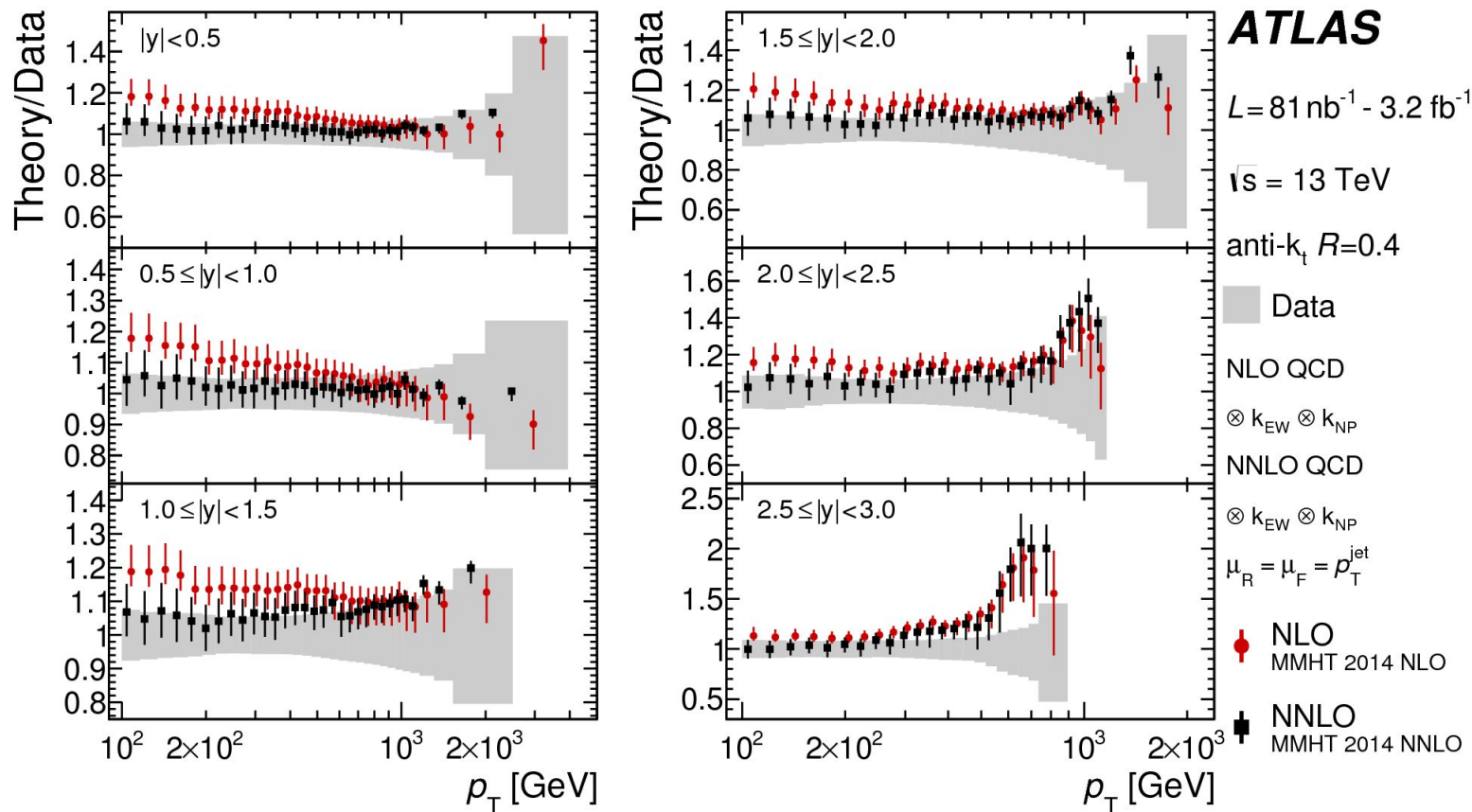
# Inclusive jet cross sections: theory/data

→ Good data/theory agreement within uncertainties observed for most PDF sets:  
 CT14, MMHT 2014, NNPDF 3.0, HERAPDF 2.0, ABMP16



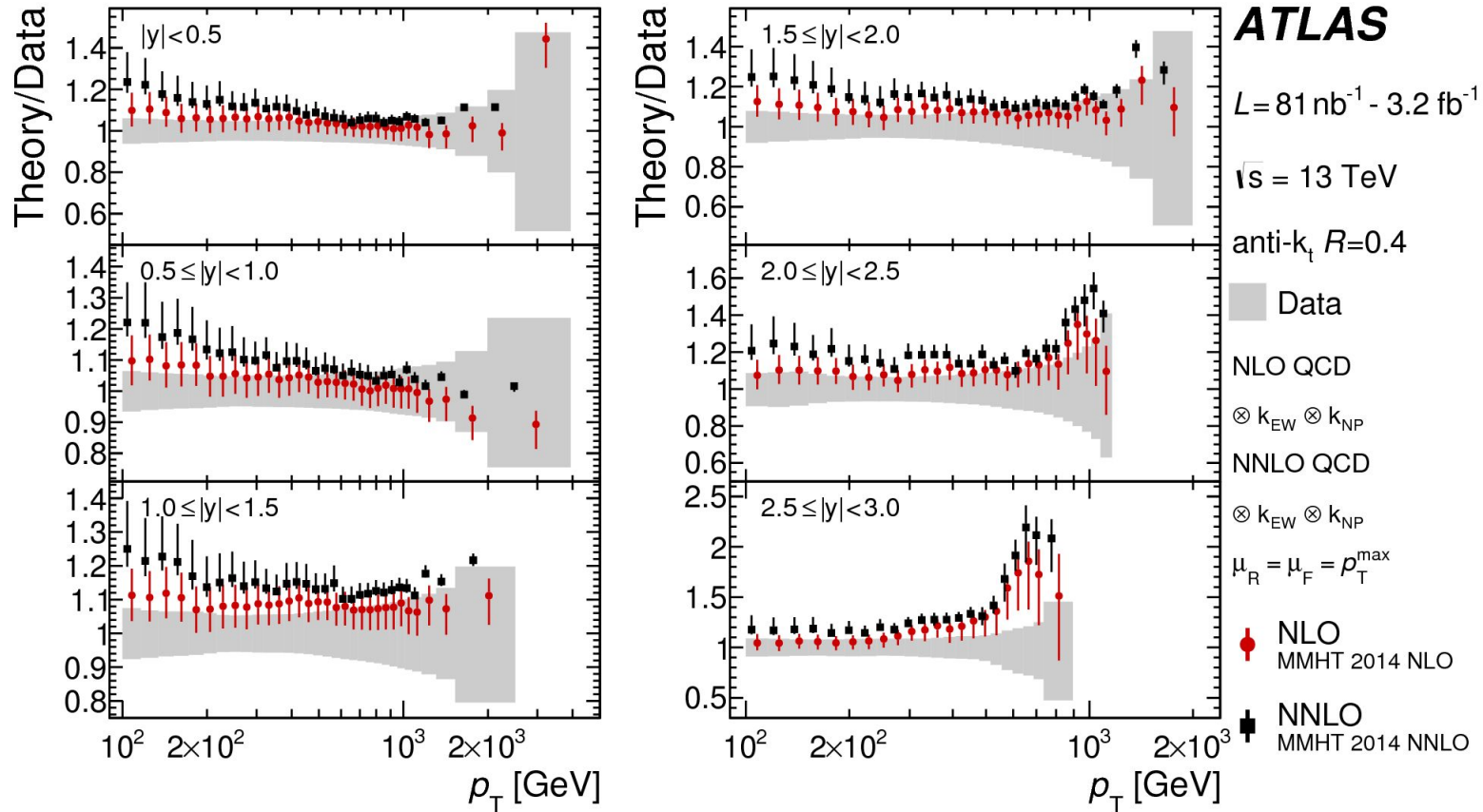
# Inclusive jet cross sections: NLO/NNLO

→ Better data/theory agreement for NNLO, when using the  $p_T^{\text{jet}}$  scale choice



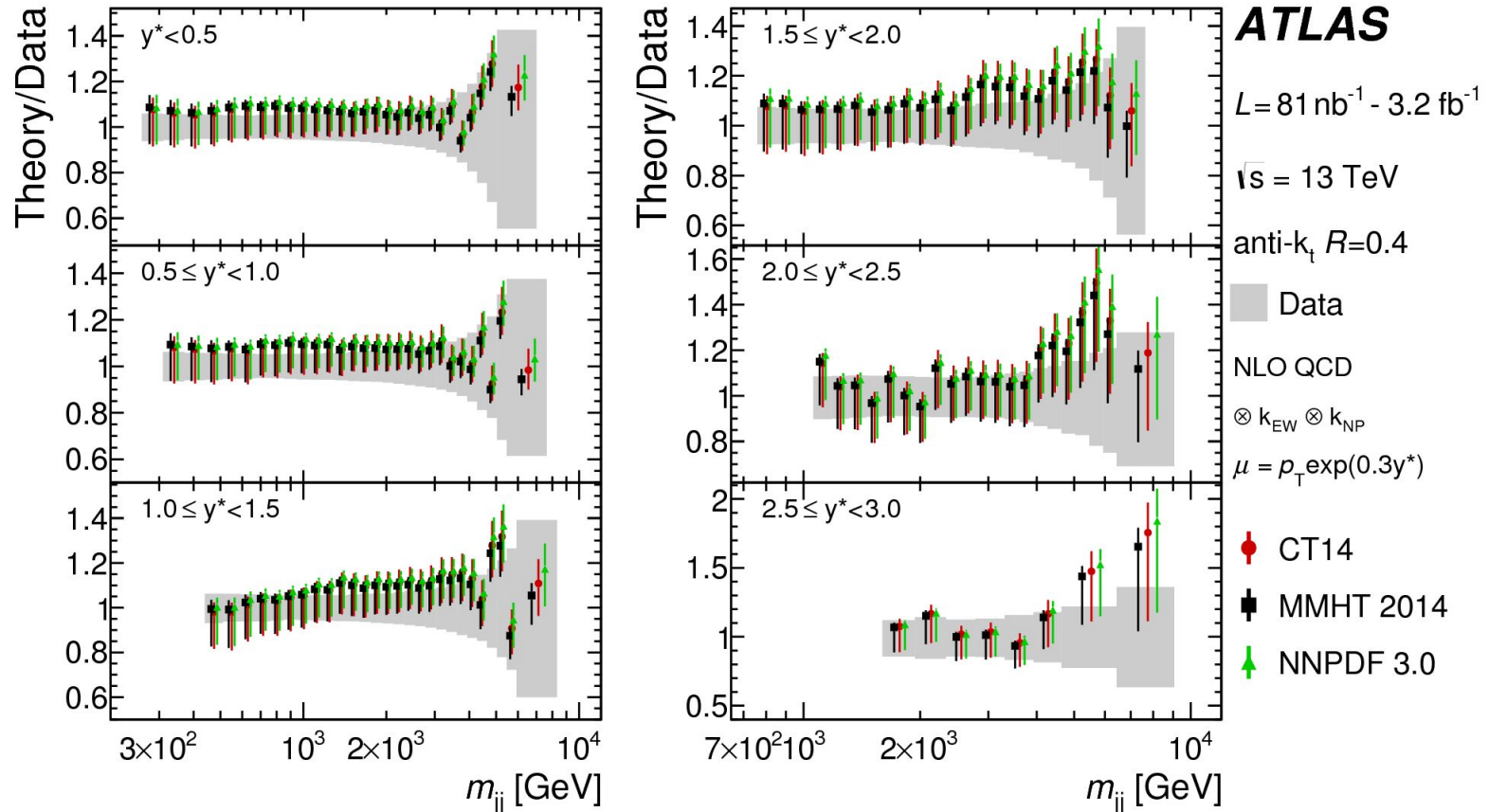
# Inclusive jet cross sections: NLO/NNLO

→ Better data/theory agreement for NLO, when using the  $p_T^{\max}$  scale choice



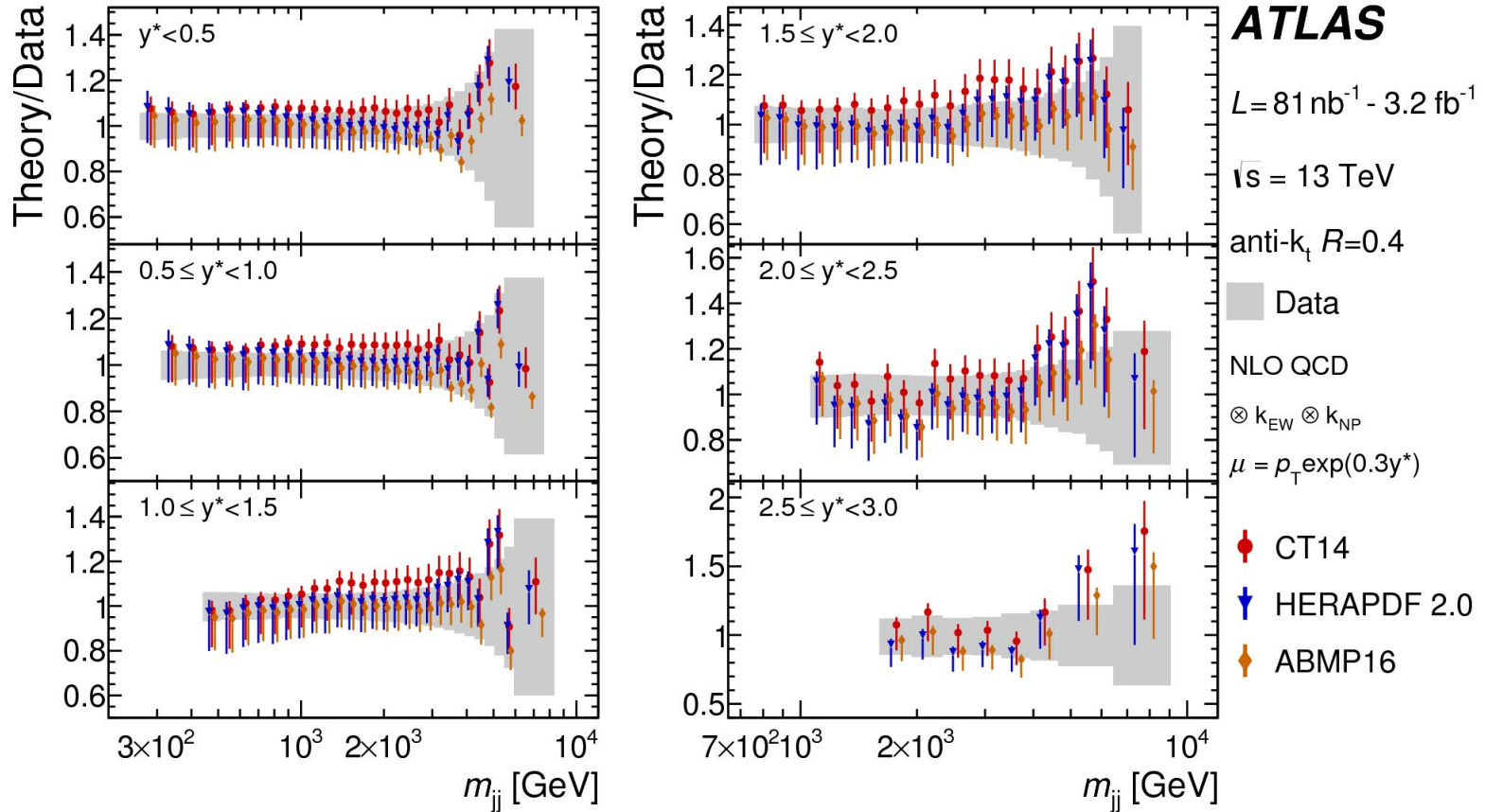
# Dijet cross sections: theory/data

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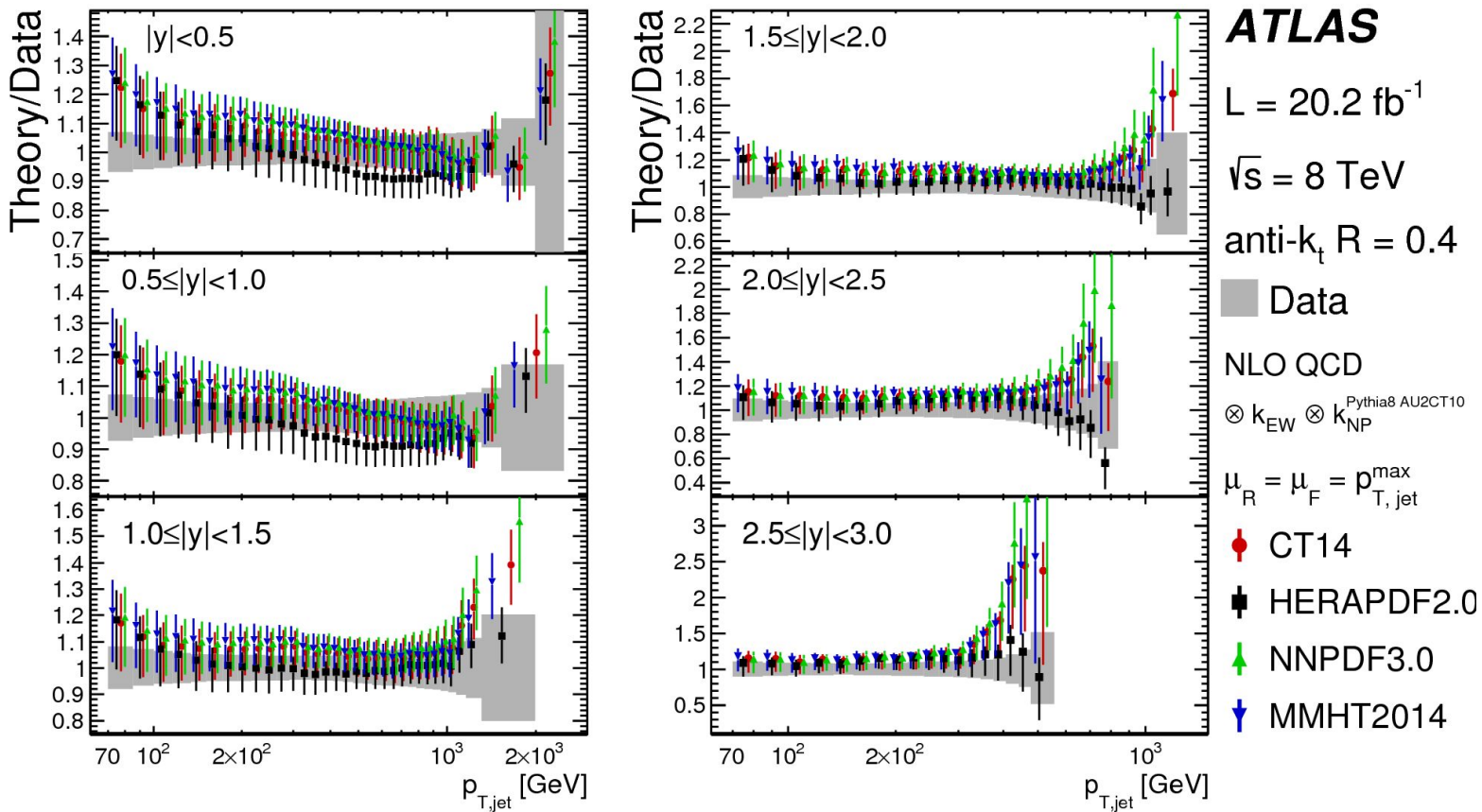
# Dijet cross sections: theory/data

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CT14, MMHT 2014, NNPDF 3.0, HERAPDF 2.0, ABMP16

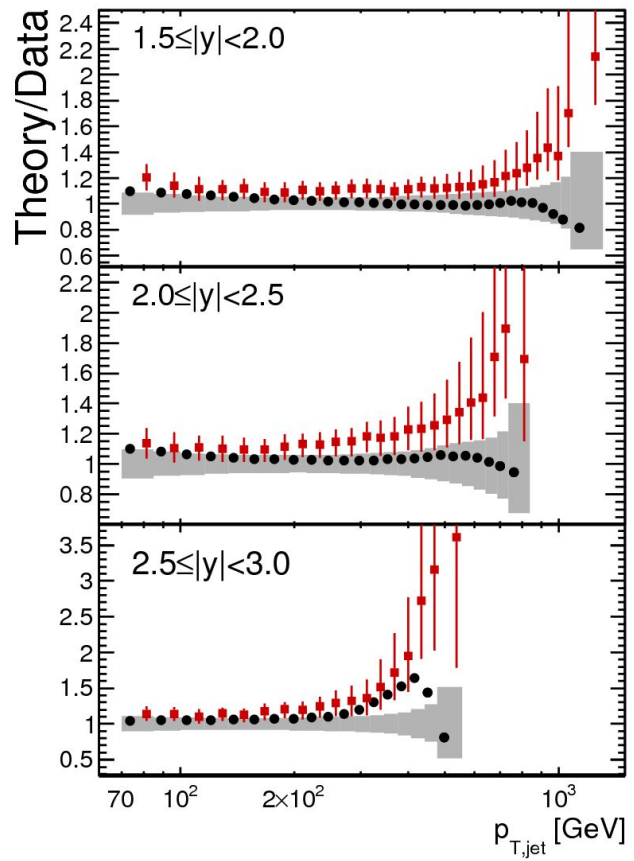
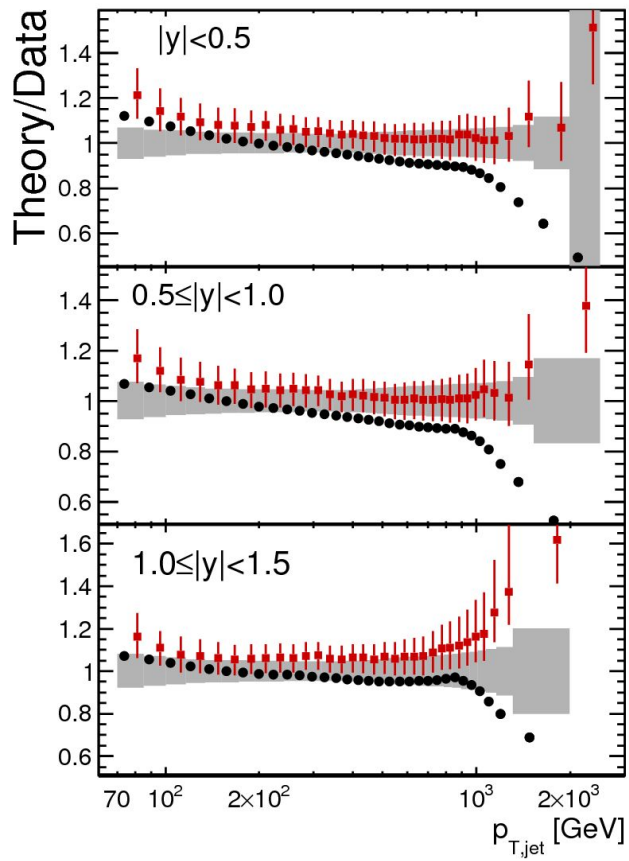


# Inclusive jet cross sections at $\sqrt{s}=8$ TeV: Theory/Data

→ Good data/theory agreement within uncertainties observed for most PDF sets



# Inclusive jet cross sections at $\sqrt{s}=8$ TeV: Theory/Data



**ATLAS**

$L = 20.2 \text{ fb}^{-1}$

$\sqrt{s} = 8 \text{ TeV}$

anti- $k_t$   $R = 0.4$

■ Data

■ NLO QCD (CT10)

⊗  $k_{EW}$  ⊗  $k_{NP}^{Pythia8 AU2CT10}$

$\mu_R = \mu_F = p_{T,jet}^{max}$

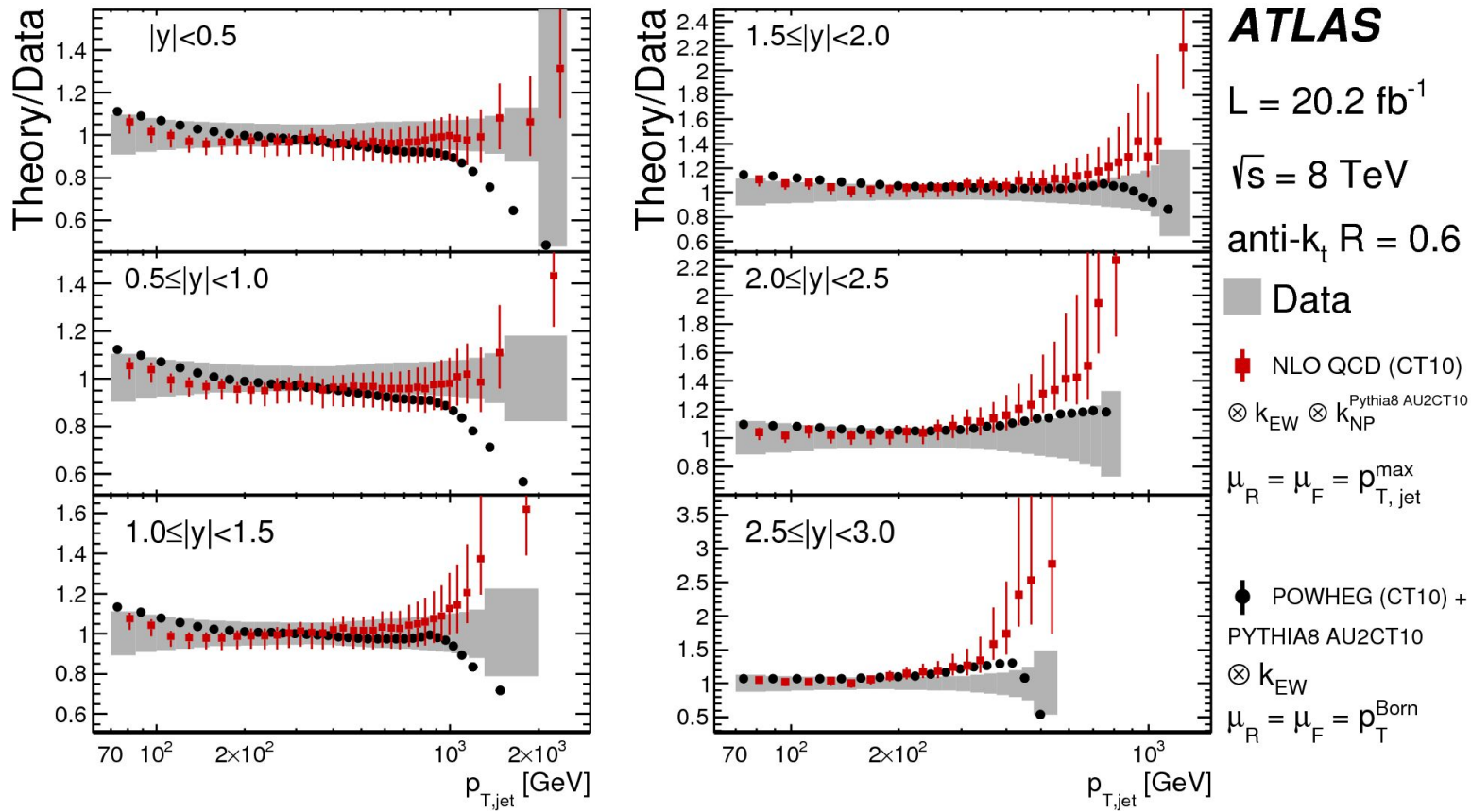
● POWHEG (CT10) + PYTHIA8 AU2CT10

⊗  $k_{EW}$

$\mu_R = \mu_F = p_T^{Born}$



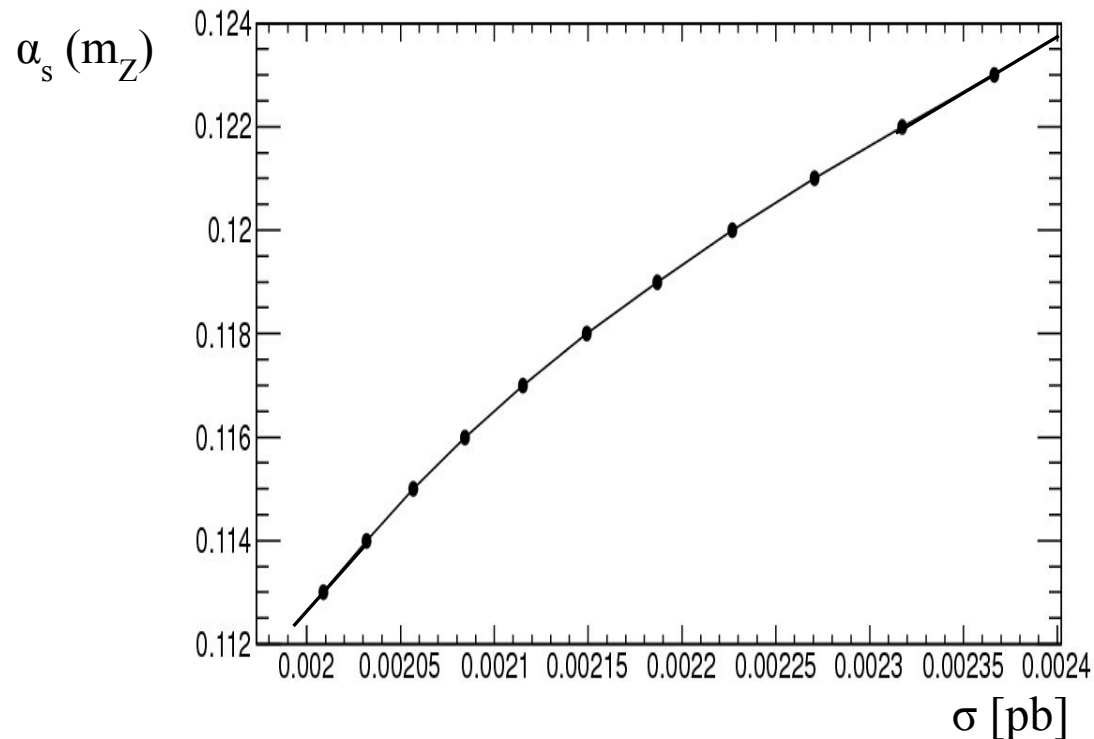
# Inclusive jet cross sections at $\sqrt{s}=8$ TeV: Theory/Data



# $\alpha_s$ determination in a given ( $p_T ; |y|$ ) bin

→ Theory prediction:  $\sigma(\alpha_s)$  using NLOJET++(CT10) with NP corrections

→ Computed for the  $\alpha_s$  values that were used in the PDF fits: interpolation between these values and extrapolation outside the covered range

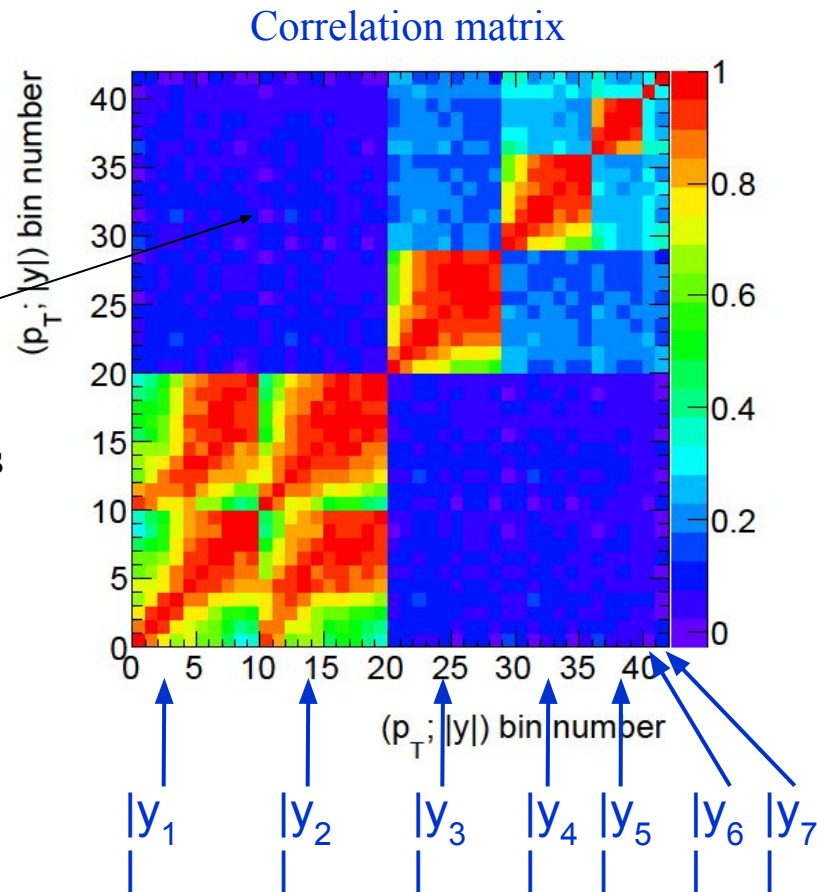


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# Covariance & correlation matrices - $\alpha_s$ - all $|y|$ bins

→ Using toys: both statistical and systematic uncertainties are fluctuated, taking into account correlations;  $\alpha_s(\sigma_{\text{toy}})$  determined for each toy

Small correlations of systematic uncertainties between central and forward regions



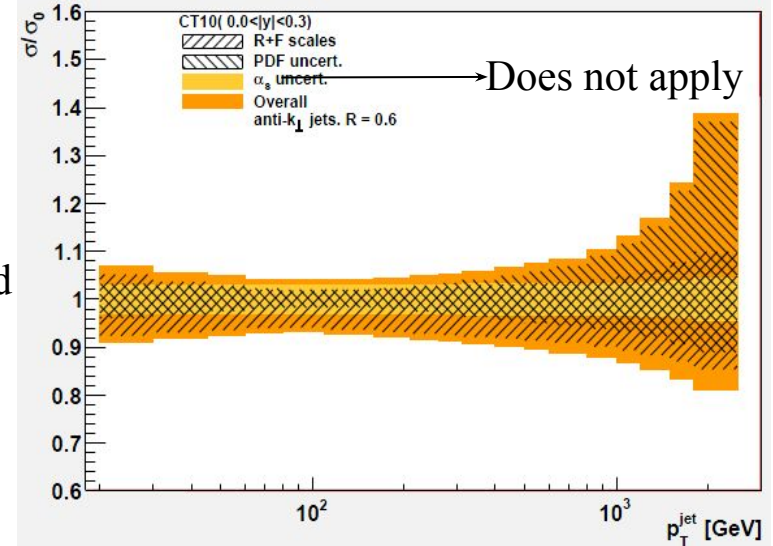
# Theoretical uncertainties & full result for $\alpha_s$ from inclusive jets

## 1) Scale uncertainty:

tested  $(\mu_r; \mu_f) \in \{(1;1), (0.5;1), (2;1), (1;0.5), (1;2), (0.5;0.5), (2;2)\}$

## 2) PDF(CT10) uncertainties propagated using 26 (positive and negative) nuisance parameters

## 3) Uncertainty due to the PDF choice



$$\alpha_s = 0.1151 \pm 0.0001(\text{stat}) \pm 0.0047(\text{exp syst}) \pm 0.0060(\text{jet size}) \pm 0.0014(\text{p}_T \text{ range})$$

$$+ 0.0044 \text{ (scale)} \pm 0.0010(\text{PDF e. v.}) + 0.0022 \text{ (PDF choice)} \pm +0.0009 \text{ (NP corr.)}$$

$$- 0.0011 \text{ (scale)} \pm 0.0010(\text{PDF e. v.}) - 0.0015 \text{ (PDF choice)} \pm -0.0034 \text{ (NP corr.)}$$

→ Good agreement with world average

→ Largest systematic uncertainty from choice of the jet size (R)

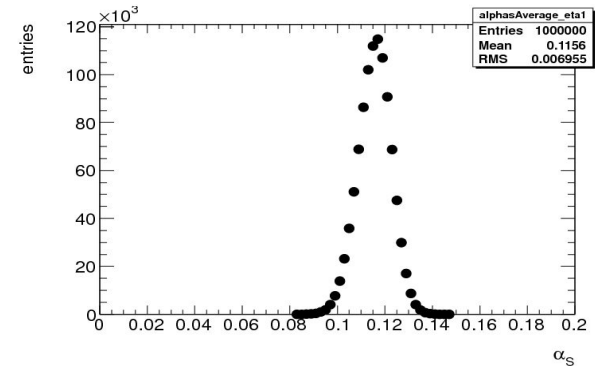
→ *Scale choice effectively different for jets with R=0.4 / R=0.6*

*Is this well motivated for a given type of events for which we just consider various observables?  
 (i.e. jet cross-sections with various jet sizes)*

# 1) Simple average - central $|y|$ bin

→ In each toy make simple average (weight= $1/N_{\text{bins}}$ ) of the  $\alpha_s$  values in 10 bins (45–600GeV) where the  $\alpha_s$  scan in PDF fits “covers” well the distribution:  
discard few low precision bins at low and high  $p_T$

→  $\alpha_s = 0.1156 + 0.0067 - 0.0072$  (exp.)



# 2) Weighted average - central $|y|$ bin

→ In each toy make weighted average (weight= $1/\sigma_i^2$ ) of the  $\alpha_s$  values in 10 bins (45–600GeV)

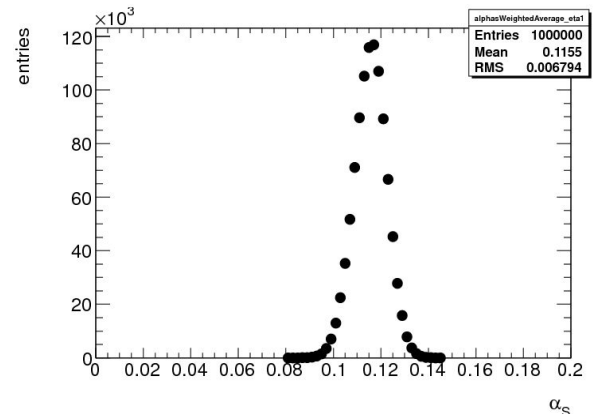
→  $\alpha_s = 0.1155 + 0.0067 - 0.0069$  (exp.)

→ Small uncertainty reduction /simple average

→  $\chi^2_{\text{diag}}/\text{dof} = 0.049$  (9 dof)

Correlations important to evaluate fit quality

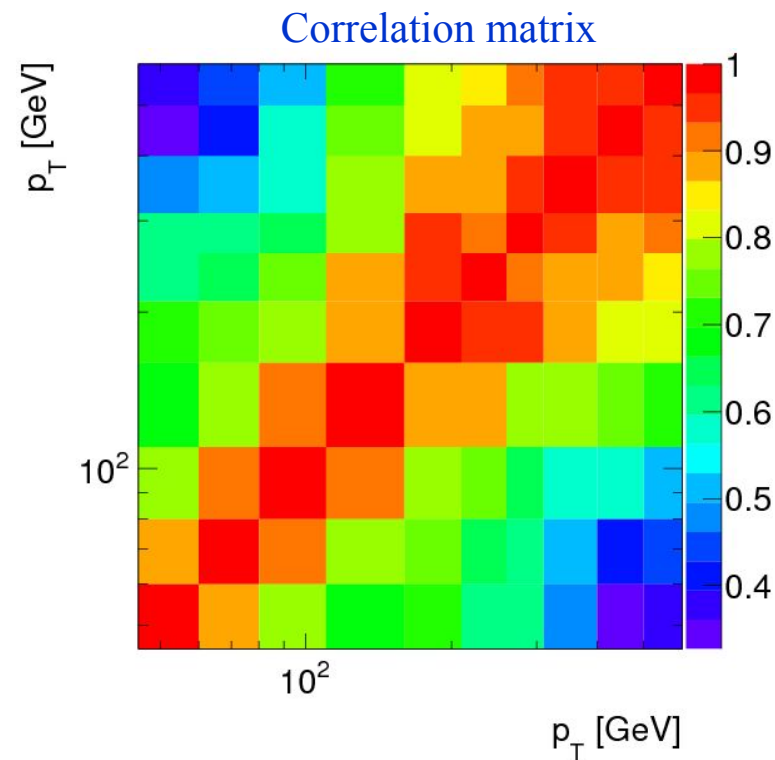
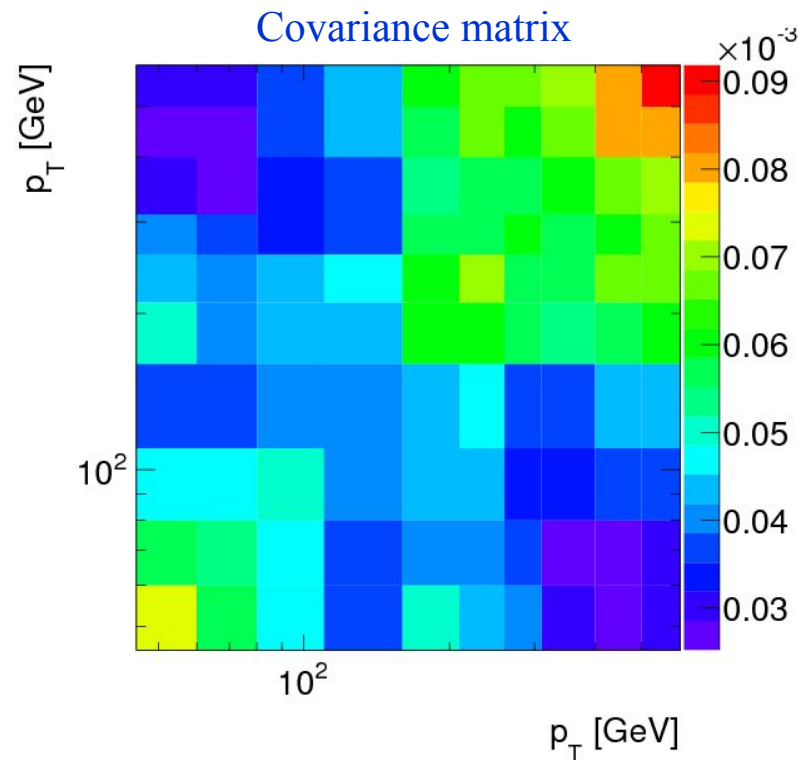
→ Weights  $\in [0;1]$



### 3) Average using $\chi^2$ minimization - central $|y|$ bin

$$\chi^2 = (\alpha_s - \alpha_s^{\text{av}})C^{-1}(\alpha_s - \alpha_s^{\text{av}})^T$$

→ Covariance and correlation matrices computed after symmetrization of  $\alpha_s$  uncertainties  
(compensation of asymmetries in data uncertainties and non-linear effects in  $\sigma(\alpha_s)$ )



### 3) Average using $\chi^2$ minimization - central $|y|$ bin

→ In each toy compute  $\alpha_s$  average by minimizing

$$\chi^2 = (\alpha_s - \alpha_s^{\text{av}})C^{-1}(\alpha_s - \alpha_s^{\text{av}})^T$$

→  $\alpha_s = 0.1160 \pm 0.0051$  (exp.)

→  $\chi^2_{\text{correl}}/\text{dof} = 0.39$  (9 dof)

→ Problem in the weights:

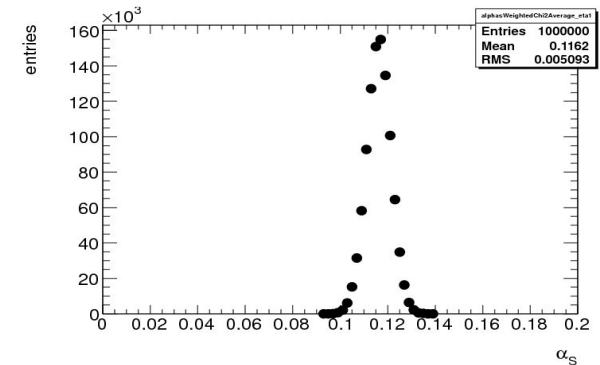
{0.019; 0.646; -0.785; 1.576; -1.262; -0.157; 0.576; 0.991; -0.116; -0.488}

“Standard” problem of  $\chi^2$  in presence of strong correlations

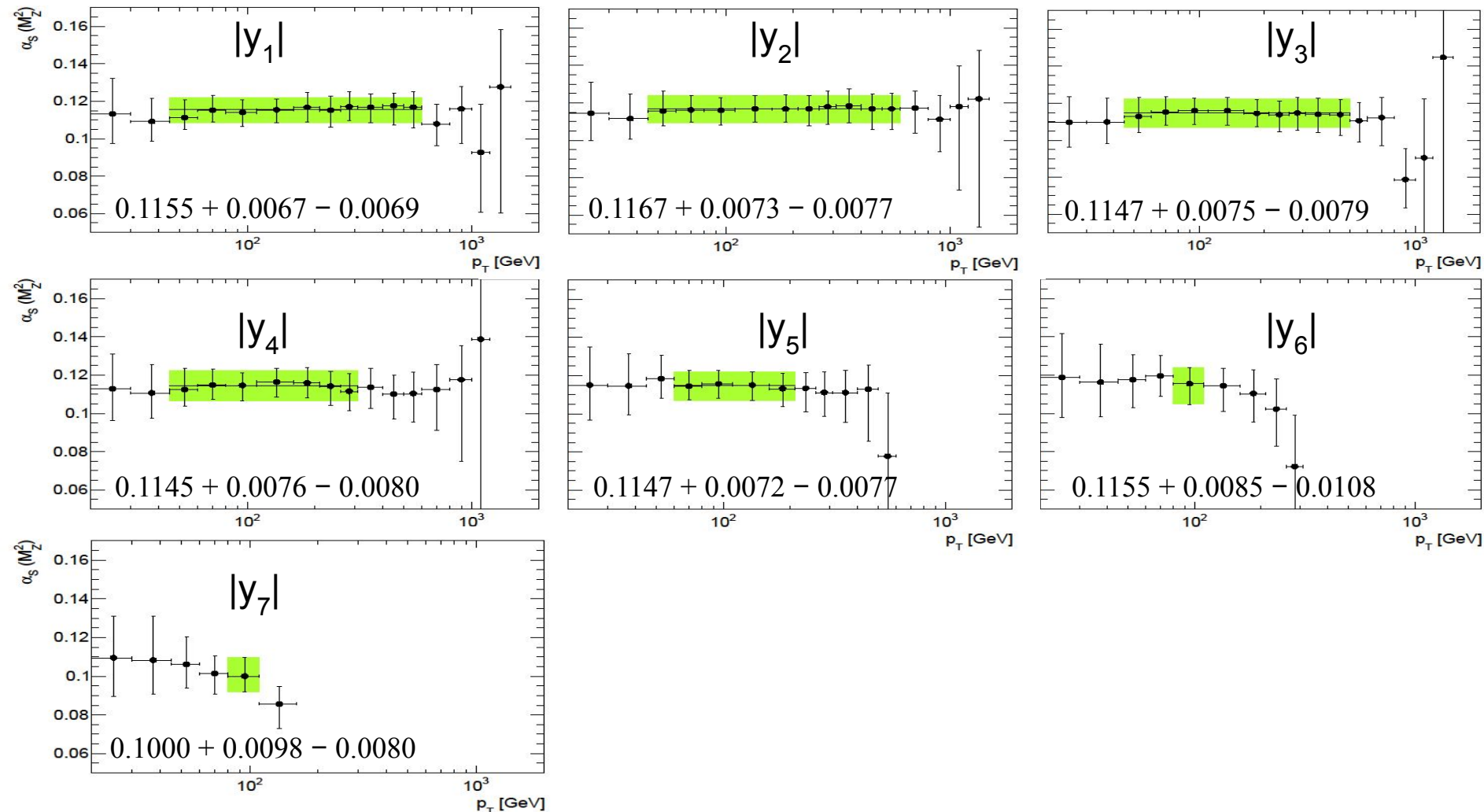
→ Important reduction of the uncertainty/weighted av.

$\chi^2$  minimization  $\sim$  uncertainty minimization, but **fully relies on the correlations of uncertainties** for determining the nominal value and the final (smallest) uncertainty: **ignores uncertainties on correlations !** (which do exist for any data set !)

→ The weighted average uses correlations only in the uncertainty propagation

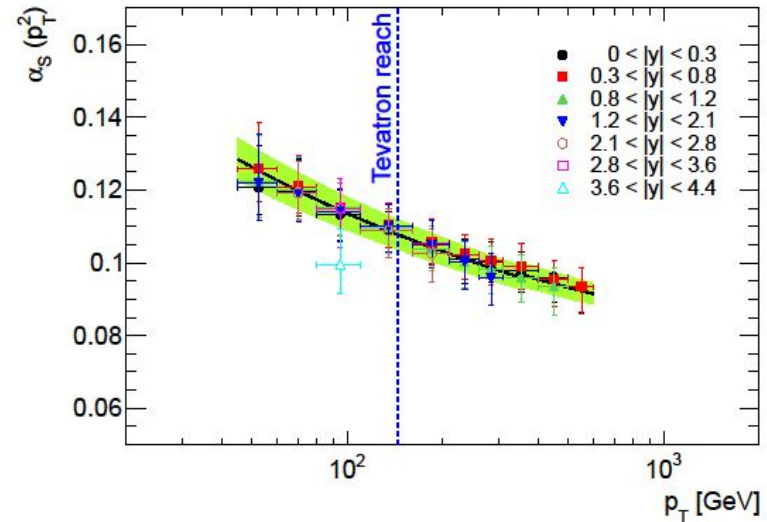
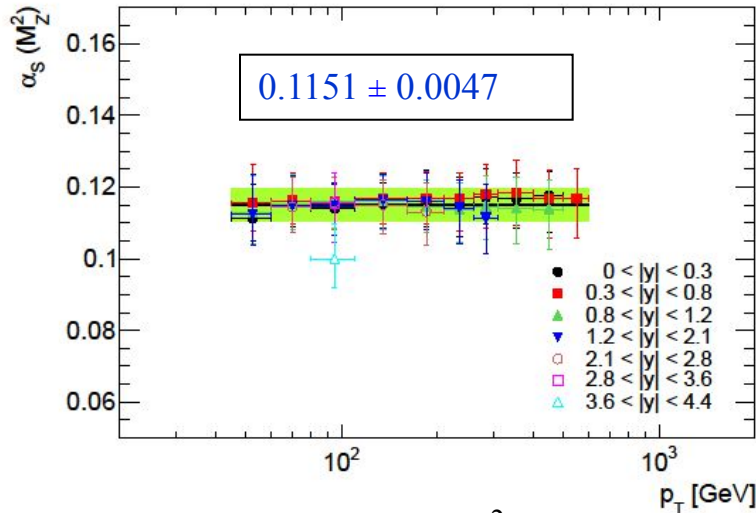


# $\alpha_s$ from inclusive jets: weighted averages - all $|y|$ bins





# $\alpha_s$ from inclusive jets: weighted average & running



→ Weighted average ( $\sim 1/\sigma_i^2$ ) for nominal result:

- avoid biases in the weights due to uncertainties on correlations
- full propagation of correlations for the evaluation of uncertainties

→ Reduction of the uncertainty for the average when including all  $|y|$  bins:

small correlations of the systematic uncertainties between different  $|y|$  bins

→  $\chi^2_{\text{correl}}/\text{dof} = 0.54$  (41 dof)

→ Test of RGE for  $p_T \in [45;600]\text{GeV}$

→ Shift of central value & error reduction minimizing  $\chi^2_{\text{correl}}$ :  $(0.1165 \pm 0.0033)$ ;

negative weights present here too

# TEEC – $\alpha_s$ scale dependence / choice

$\langle Q \rangle$ (GeV)	TEEC	$\alpha_s(Q^2)$ value (NNPDF 3.0)
412	$0.0966 \pm 0.0014$ (exp.)	$^{+0.0054}_{-0.0015}$ (scale) $\pm 0.0009$ (PDF) $\pm 0.0001$ (NP)
437	$0.0964 \pm 0.0012$ (exp.)	$^{+0.0048}_{-0.0011}$ (scale) $\pm 0.0009$ (PDF) $\pm 0.0002$ (NP)
472	$0.0955 \pm 0.0011$ (exp.)	$^{+0.0051}_{-0.0015}$ (scale) $\pm 0.0009$ (PDF) $\pm 0.0001$ (NP)
522	$0.0936 \pm 0.0011$ (exp.)	$^{+0.0043}_{-0.0010}$ (scale) $\pm 0.0010$ (PDF) $\pm 0.0001$ (NP)
604	$0.0933 \pm 0.0011$ (exp.)	$^{+0.0050}_{-0.0014}$ (scale) $\pm 0.0011$ (PDF) $\pm 0.0003$ (NP)
810	$0.0907 \pm 0.0013$ (exp.)	$^{+0.0049}_{-0.0020}$ (scale) $\pm 0.0011$ (PDF) $\pm 0.0002$ (NP)

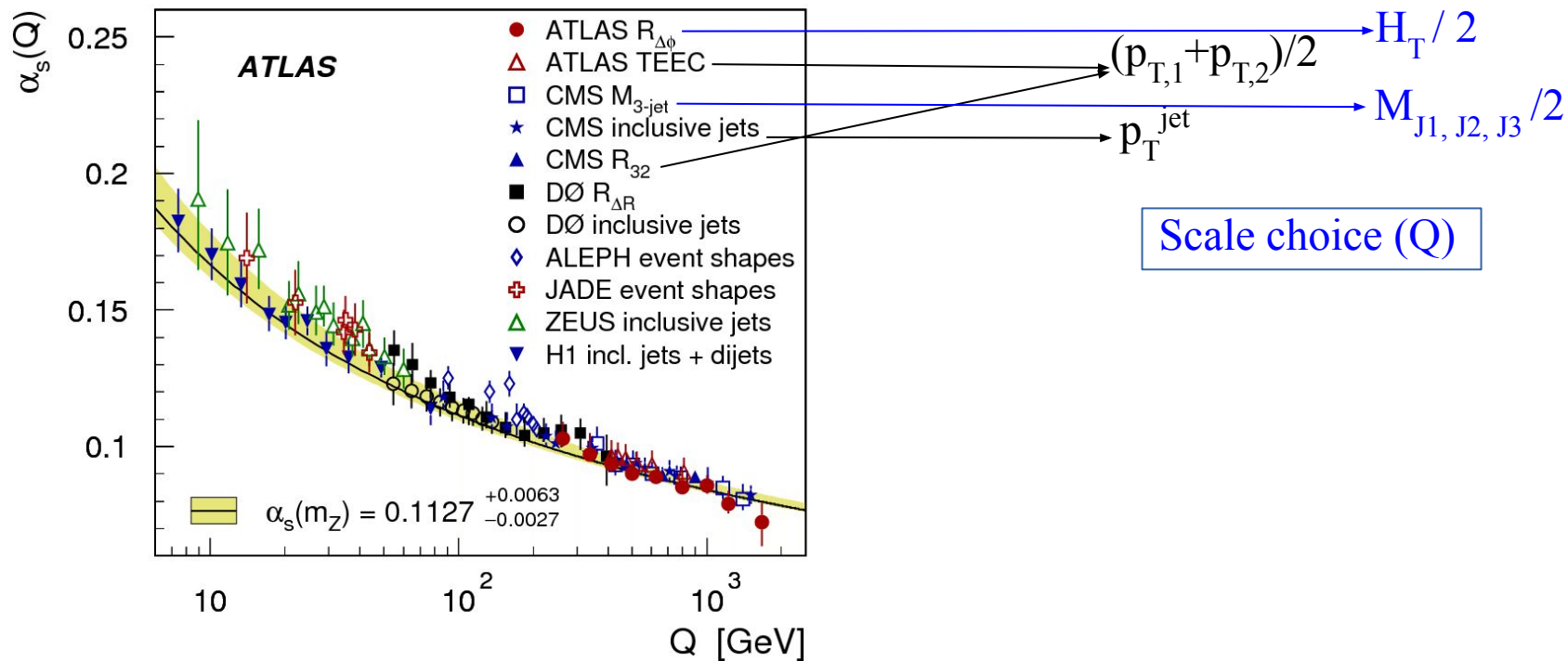
$\langle p_{T3} \rangle$ (GeV)	$\alpha_s(\langle p_{T3} \rangle)$ value (TEEC, NNPDF 3.0)
169	$0.1072 \pm 0.0017$ (exp.) $^{+0.0067}_{-0.0019}$ (scale) $\pm 0.0011$ (PDF) $\pm 0.0001$ (NP)
174	$0.1074 \pm 0.0014$ (exp.) $^{+0.0060}_{-0.0014}$ (scale) $\pm 0.0012$ (PDF) $\pm 0.0002$ (NP)
179	$0.1068 \pm 0.0014$ (exp.) $^{+0.0064}_{-0.0019}$ (scale) $\pm 0.0012$ (PDF) $\pm 0.0001$ (NP)
186	$0.1052 \pm 0.0014$ (exp.) $^{+0.0054}_{-0.0013}$ (scale) $\pm 0.0013$ (PDF) $\pm 0.0001$ (NP)
197	$0.1060 \pm 0.0014$ (exp.) $^{+0.0065}_{-0.0018}$ (scale) $\pm 0.0014$ (PDF) $\pm 0.0004$ (NP)
215	$0.1052 \pm 0.0018$ (exp.) $^{+0.0066}_{-0.0027}$ (scale) $\pm 0.0015$ (PDF) $\pm 0.0003$ (NP)

→ For observables like  $R_{3/2}$ ,  $N_{3/2}$ ,  $R_{\Delta\phi}$  and (A)TEEC, sensitivity to  $\alpha_s$  originates from probability of emission of extra radiation (3<sup>rd</sup> jet etc.)

→ Effect acknowledged by evolving  $\alpha_s$  to  $\langle p_{T3} \rangle$  (significantly lower than  $\langle H_{T2} \rangle$ )

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# Thoughts on RGE tests through jet measurements



→ Can one really claim tests of RGE at scales from event-level observables ???

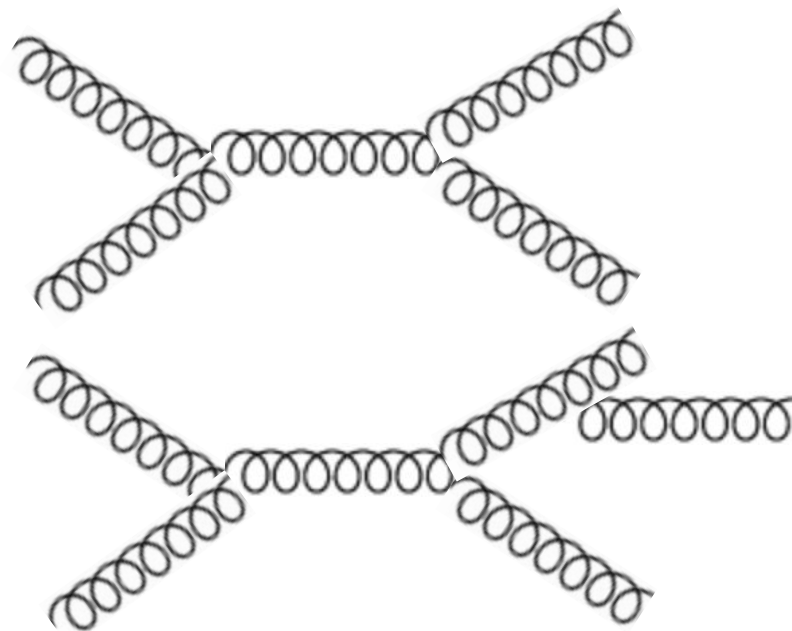
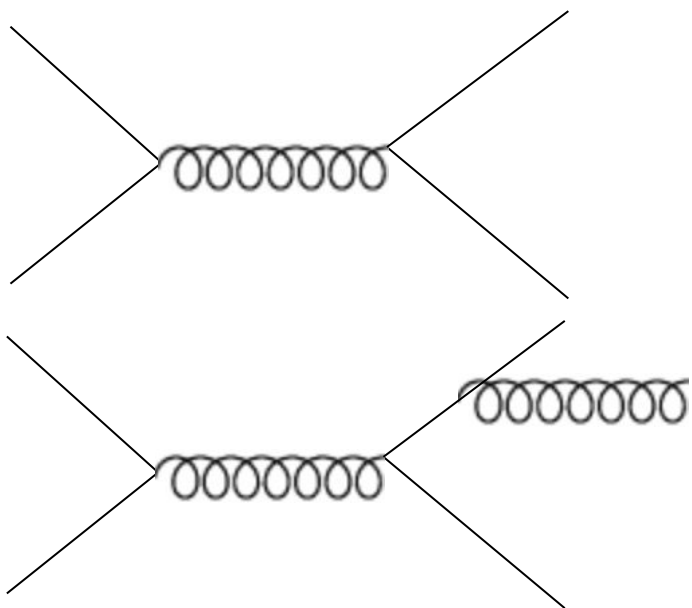
e.g.  $p_T^{\text{lead. jet}}(R_{3/2})$ ,  $p_T^{(\text{all jets})}(N_{3/2})$ ,  $(p_{T,1} + p_{T,2})/2$ ,  $H_T/2$ ,  $M_{J_1, J_2, J_3}/2$  (large even for low  $p_{T,1-3}$ )

→ “Traditional criteria” of minimizing uncertainties/k-factors is not relevant here

→ Relevant scale for RGE test using  $R_{3/2}$ ,  $N_{3/2}$ ,  $R_{\Delta\phi}$  and (A)TEEC related to  $p_{T,3}$  (low)

Need consistency between scale for theory calculation and RGE test claim; MiNLO procedure may provide a way forward.

# Thoughts on PDF sensitivity in $\alpha_s$ evaluations from jet Xsec ratios



→ *PDF uncertainties non-negligible* (typically between total experimental and NLO scale uncertainty) *for cross-section ratio measurements & (A)TEEC:*

- probability of extra radiation (which makes these observables non-trivial) *sensitive to the type of partons in the initial state*
- both  $\alpha_s$  & PDF sensitivities of the observables are reduced when taking ratios and they are both relevant for the  $\alpha_s$  evaluation

# Result quantitative comparisons for “all” PDFs

- Individual  $|y|$  bins, wide  $p_T$  ranges: p-values generally  $> 4\%$  ( $\sim 1\%$  or lower for  $R=0.6$ ,  $0.5 < |y| < 1$  at 8 TeV,  $1.5 < |y| < 2$  at 13 TeV), decreasing when considering wider phase-space regions
- Full  $|y|$  range, wide  $p_T$  ranges: p-values  $\ll 10^{-3}$   
( $p_T > 100$  GeV)  $\chi^2/\text{ndf}$ :  $\sim 313\text{-}385/159$  (8 TeV);  $384\text{-}475/177$  (13 TeV)
- Data/theory tension also seen initially by CMS in arXiv:1410.6765 when using the original data, uncertainties and correlations from arXiv:1212.6660
- CMS noticed that “Changing the correlation in the JES uncertainty from 0% to 100% produces a steep rise in  $\chi^2/\text{ndf}$ ” and modified the correlation model
- Good data/theory agreement on full phase-space for ATLAS dijets (13 TeV)
- Full  $|y|$  range, narrow  $p_T$  ranges: good data/theory agreement for  $70 < p_T < 100$  GeV;  
p-values are often below  $10^{-3}$  for the other narrow  $p_T$  ranges
- Pairs of  $|y|$  bins (consecutive / central-forward), narrow  $p_T$  ranges at  $>100$  GeV:  
Good data/theory agreement  $\rightarrow$  source of low p-values not in a single  $|y|$  bin, nor due to some possible central/forward tension
- Little sensitivity to choice of non-perturbative correction and to scale choice

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# Role of uncertainty correlations

- Correlations of uncertainties between various phase-space regions have a **key role in  $\chi^2$  evaluation** (e.g. ignoring correlations yields a very small  $\chi^2/\text{ndf}$ )
- Experimental uncertainties (examples for ATLAS measurements):
  - JES in-situ statistical uncertainties: correlations well known (e.g. > 240 components for calibration using dijet balance reduce  $\chi^2$  by more than 200 units )
  - JES Flavour Response, JES MJB Fragmentation, JES Pile-up Rho Topology:  
“2-point systematics” from comparison of various MC generators – unknown correlations
- Theoretical uncertainties:
  - $\alpha_s$ , PDFs: correlations (generally) well known
  - Scale variations, alternative scale choice, non-perturbative corrections:  
“2-point systematics” – unknown correlations
- *Good understanding of the sources of systematic uncertainties required in order to evaluate uncertainties on correlations: performed detailed tests using realistic alternative correlation scenarios*

# Testing realistic alternative correlation assumptions

→ 18 options for splitting the systematics with unknown correlations in  
 2 or 3 sub-components with smooth  $p_T$  and/or  $|y|$  dependence

Splitting option	Sub-component(s) definition(s), completed by complementary
1	$L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot \text{uncertainty}$
2	$L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot 0.5 \cdot \text{uncertainty}$
3	$L(p_T[\text{TeV}], 0.1, 2.5) \cdot \text{uncertainty}$
4	$L(p_T[\text{TeV}], 0.1, 2.5) \cdot 0.5 \cdot \text{uncertainty}$
5	$L((\ln(p_T[\text{TeV}]))^2, (\ln(0.1))^2, (\ln(2.5))^2) \cdot \text{uncertainty}$
6	$L((\ln(p_T[\text{TeV}]))^2, (\ln(0.1))^2, (\ln(2.5))^2) \cdot 0.5 \cdot \text{uncertainty}$
7	$L( y , 0, 3) \cdot \text{uncertainty}$
8	$L( y , 0, 3) \cdot 0.5 \cdot \text{uncertainty}$
9	$L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot L( y , 0, 3) \cdot \text{uncertainty}$
10	$L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot \sqrt{1 - L( y , 0, 3)^2} \cdot \text{uncertainty}$
11	$L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot L( y , 0, 3) \cdot 0.5 \cdot \text{uncertainty}$
12	$L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot \sqrt{1 - L( y , 0, 3)^2} \cdot 0.5 \cdot \text{uncertainty}$
13	$L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot \sqrt{1 - L( y , 0, 1.5)^2} \cdot \text{uncertainty}$ $L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot L( y , 1.5, 3) \cdot \text{uncertainty}$
14	$L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot \sqrt{1 - L( y , 0, 1)^2} \cdot \text{uncertainty}$ $L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot L( y , 1, 3) \cdot \text{uncertainty}$
15	$L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot \sqrt{1 - L( y , 0, 2)^2} \cdot \text{uncertainty}$ $L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5)) \cdot L( y , 2, 3) \cdot \text{uncertainty}$
16	$\sqrt{1 - L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5))^2} \cdot \sqrt{1 - L( y , 0, 1.5)^2} \cdot \text{uncertainty}$ $\sqrt{1 - L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5))^2} \cdot L( y , 1.5, 3) \cdot \text{uncertainty}$
17	$\sqrt{1 - L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5))^2} \cdot \sqrt{1 - L( y , 0, 1)^2} \cdot \text{uncertainty}$ $\sqrt{1 - L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5))^2} \cdot L( y , 1, 3) \cdot \text{uncertainty}$
18	$\sqrt{1 - L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5))^2} \cdot \sqrt{1 - L( y , 0, 2)^2} \cdot \text{uncertainty}$ $\sqrt{1 - L(\ln(p_T[\text{TeV}]), \ln(0.1), \ln(2.5))^2} \cdot L( y , 2, 3) \cdot \text{uncertainty}$

→ Tested for *experimental and theoretical systematic uncertainties*

*Changed theorists' view on how to interpret our measurements*

→ One component added to the ones listed for each option in the table, to keep total uncertainty unchanged

$$L(x, \min, \max) = (x - \min) / (\max - \min)$$

# Testing realistic alternative correlation assumptions

→ Splitting the *theory systematic uncertainties* with unknown correlations in 6 sub-components with smooth  $p_T$  and  $|y|$  dependence

$$f_1(p_T, y) = C(p_T, y) \cdot c_1 / \log(M(y)/p_T),$$

$$f_2(p_T, y) = C(p_T, y) \cdot c_2 \cdot y^2 / \log(M(y)/p_T),$$

$$f_3(p_T, y) = C(p_T, y) \cdot c_3,$$

$$f_4(p_T, y) = C(p_T, y) \cdot c_4 \cdot y^2,$$

$$f_5(p_T, y) = C(p_T, y) \cdot c_5 \cdot \log(15p_T/M(y)),$$

$$f_6(p_T, y) = C(p_T, y) \cdot c_6 \cdot y^2 \cdot \log(15p_T/M(y))$$

$$M(y) = \sqrt{s} \cdot \exp(-y)$$

Based on:

Phys. Rev. D81 (2010) 035018

arXiv:0907.5052 [hep-ph]

→ 3 options for various values of the coefficients ( $c_1 - c_6$ )



# Quantitative comparison between data and NLO theory prediction

8 TeV – ATLAS inclusive jets (arXiv:1706.03192)

Rapidity ranges	$P_{\text{obs}}$			
	CT14	MMHT2014	NNPDF3.0	HERAPDF2.0
Anti- $k_t$ jets $R = 0.4$				
$ y  < 0.5$	44%	28%	25%	16%
$0.5 \leq  y  < 1.0$	43%	29%	18%	18%
$1.0 \leq  y  < 1.5$	44%	47%	46%	69%
$1.5 \leq  y  < 2.0$	3.7%	4.6%	7.7%	7.0%
$2.0 \leq  y  < 2.5$	92%	89%	89%	35%
$2.5 \leq  y  < 3.0$	4.5%	6.2%	16%	9.6%
Anti- $k_t$ jets $R = 0.6$				
$ y  < 0.5$	6.7%	4.9%	4.6%	1.1%
$0.5 \leq  y  < 1.0$	1.3%	0.7%	0.4%	0.2%
$1.0 \leq  y  < 1.5$	30%	33%	47%	67%
$1.5 \leq  y  < 2.0$	12%	16%	15%	3.1%
$2.0 \leq  y  < 2.5$	94%	94%	91%	38%
$2.5 \leq  y  < 3.0$	13%	15%	20%	8.6%

→ Generally good agreement for individual  $|y|$  bins

Splitting options for $R = 0.4$	CT14	NNPDF3.0
JES Flavour Response Opt 7	268/159	257/159
JES MJB Fragmentation Opt 17		
JES Pile-up Rho topology Opt 18		
Scale variations Opt 17		
Alternative scale choice Opt 7		
Non-perturbative corrections Opt 7		
JES Flavour Response Opt 7		
JES MJB Fragmentation Opt 17		
JES Pile-up Rho topology Opt 18		
Scale variations Opt 20		
Alternative scale choice Opt 17		
Non-perturbative corrections Opt 7		

$\chi^2/\text{ndf}$	$P_{\text{T}}^{\text{jet,max}}$		$P_{\text{T}}^{\text{jet}}$	
	$R = 0.4$	$R = 0.6$	$R = 0.4$	$R = 0.6$
$p_{\text{T}} > 70 \text{ GeV}$				
CT14	349/171	398/171	340/171	392/171
HERAPDF2.0	415/171	424/171	405/171	418/171
NNPDF3.0	351/171	393/171	350/171	393/171
MMHT2014	356/171	400/171	354/171	399/171
$p_{\text{T}} > 100 \text{ GeV}$				
CT14	321/159	360/159	313/159	356/159
HERAPDF2.0	385/159	374/159	377/159	370/159
NNPDF3.0	333/159	356/159	331/159	356/159
MMHT2014	335/159	364/159	333/159	362/159
$100 < p_{\text{T}} < 900 \text{ GeV}$				
CT14	272/134	306/134	262/134	301/134
HERAPDF2.0	350/134	331/134	340/134	326/134
NNPDF3.0	289/134	300/134	285/134	299/134
MMHT2014	292/134	311/134	284/134	308/134
$100 < p_{\text{T}} < 400 \text{ GeV}$				
CT14	128/72	149/72	118/72	145/72
HERAPDF2.0	148/72	175/72	141/72	170/72
NNPDF3.0	119/72	141/72	115/72	139/72
MMHT2014	132/72	143/72	122/72	140/72

→ Tension when including all  $|y|$  bins

# Quantitative comparison between data and NLO theory prediction

13 TeV – ATLAS inclusive jets and dijets (arXiv:1711.02692)

Rapidity ranges	$P_{\text{obs}}$				
	CT14	MMHT 2014	NNPDF 3.0	HERAPDF 2.0	ABMP16
$p_{\text{T}}^{\text{max}}$					
$ y  < 0.5$	67%	65%	62%	31%	50%
$0.5 \leq  y  < 1.0$	5.8%	6.3%	6.0%	3.0%	2.0%
$1.0 \leq  y  < 1.5$	65%	61%	67%	50%	55%
$1.5 \leq  y  < 2.0$	0.7%	0.8%	0.8%	0.1%	0.4%
$2.0 \leq  y  < 2.5$	2.3%	2.3%	2.8%	0.7%	1.5%
$2.5 \leq  y  < 3.0$	62%	71%	69%	25%	55%
$p_{\text{T}}^{\text{jet}}$					
$ y  < 0.5$	69%	67%	66%	30%	46%
$0.5 \leq  y  < 1.0$	7.4%	8.9%	8.6%	3.4%	2.0%
$1.0 \leq  y  < 1.5$	69%	62%	68%	45%	54%
$1.5 \leq  y  < 2.0$	1.3%	1.6%	1.4%	0.1%	0.5%
$2.0 \leq  y  < 2.5$	8.7%	6.6%	7.4%	1.0%	3.6%
$2.5 \leq  y  < 3.0$	65%	72%	72%	28%	59%

→ Generally good agreement for inclusive jets for individual  $|y|$  bins

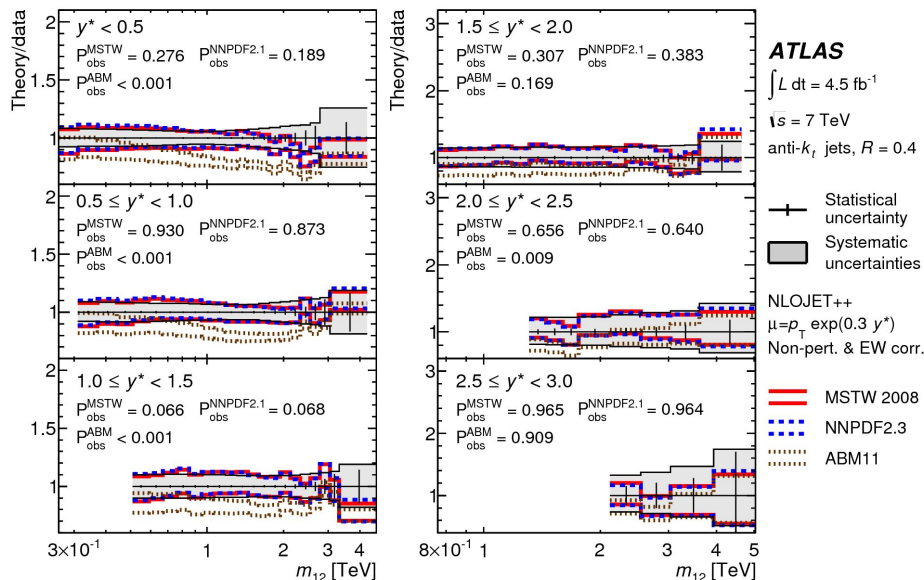
$\chi^2/\text{dof}$ all $ y $ bins	CT14	MMHT 2014	NNPDF 3.0	HERAPDF 2.0	ABMP16
$p_{\text{T}}^{\text{max}}$	419/177	431/177	404/177	432/177	475/177
$p_{\text{T}}^{\text{jet}}$	399/177	405/177	384/177	428/177	455/177

→ Tension when including all  $|y|$  bins for inclusive jets

$y^*$ ranges	$P_{\text{obs}}$				
	CT14	MMHT 2014	NNPDF 3.0	HERAPDF 2.0	ABMP16
$y^* < 0.5$	79%	59%	50%	71%	71%
$0.5 \leq y^* < 1.0$	27%	23%	19%	32%	31%
$1.0 \leq y^* < 1.5$	66%	55%	48%	66%	69%
$1.5 \leq y^* < 2.0$	26%	26%	28%	9.9%	25%
$2.0 \leq y^* < 2.5$	41%	34%	29%	3.6%	20%
$2.5 \leq y^* < 3.0$	45%	46%	40%	25%	38%
all $y^*$ bins	9.4%	6.5%	11%	0.1%	5.1%

→ Good data/theory agreement for dijets

# PDF comparisons for dijets at 7 TeV

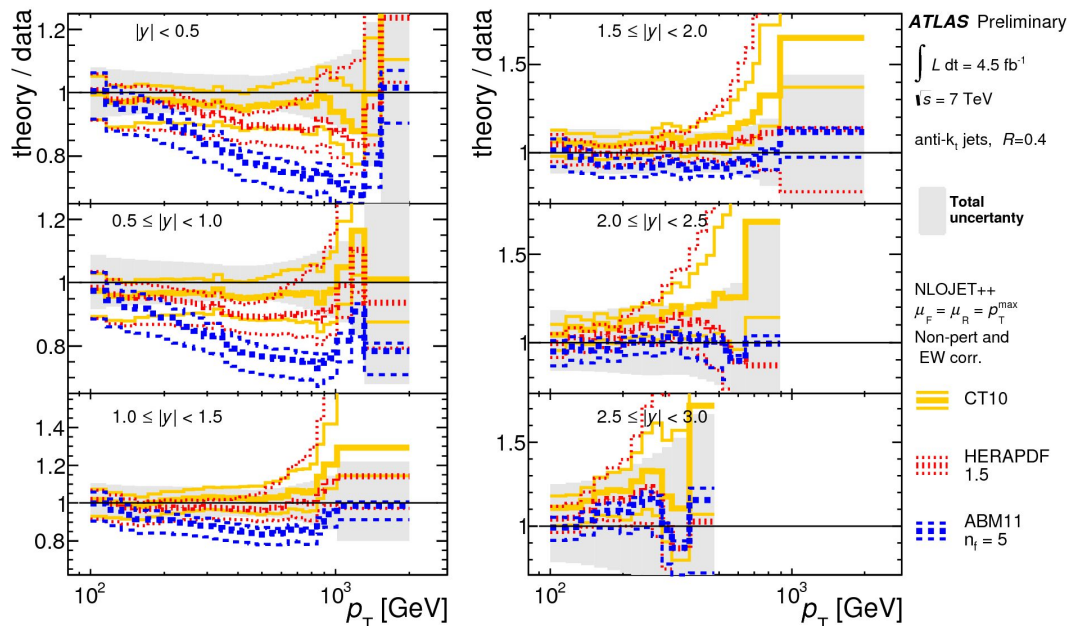


PDF set	$y^*$ ranges	mass range (full/high)	$P_{\text{obs}}$	
			$R = 0.4$	$R = 0.6$
CT10	$y^* < 0.5$	high	0.742	0.785
	$y^* < 1.5$	high	0.080	0.066
	$y^* < 1.5$	full	0.324	0.168
HERAPDF1.5	$y^* < 0.5$	high	0.688	0.504
	$y^* < 1.5$	high	0.025	0.007
	$y^* < 1.5$	full	0.137	0.025
MSTW 2008	$y^* < 0.5$	high	0.328	0.533
	$y^* < 1.5$	high	0.167	0.183
	$y^* < 1.5$	full	0.470	0.352
NNPDF2.1	$y^* < 0.5$	high	0.405	0.568
	$y^* < 1.5$	high	0.151	0.125
	$y^* < 1.5$	full	0.431	0.242
ABM11	$y^* < 0.5$	high	0.024	$< 10^{-3}$
	$y^* < 1.5$	high	$< 10^{-3}$	$< 10^{-3}$
	$y^* < 1.5$	full	$< 10^{-3}$	$< 10^{-3}$

→ **Sensitivity to PDFs:** level of agreement strongly depends on the PDF set and phase-space region

→ **Valuable experimental inputs to constrain proton PDFs:** Published information on **cross-sections & uncertainties, with their correlations and asymmetries**

# PDF comparisons for inclusive jets at 7 TeV



$y$ ranges	$P_{\text{obs}}$ (ATLAS Preliminary)					
	NLO PDF set:	CT10	MSTW2008	NNPDF2.1	HERAPDF1.5	ABM11
$ y  < 0.5$		84%	61%	72%	56%	< 0.1%
$0.5 \leq  y  < 1.0$		91%	93%	89%	49%	< 0.1%
$1.0 \leq  y  < 1.5$		89%	88%	85%	93%	2.7%
$1.5 \leq  y  < 2.0$		93%	88%	91%	75%	55%
$2.0 \leq  y  < 2.5$		86%	82%	85%	26%	57%
$2.5 \leq  y  < 3.0$		95%	94%	97%	82%	85%

→  $P_{\text{obs}}$  strongly depends on the PDF set and phase-space region

# Quantitative data/theory comparison & limit setting

→ Define global test statistic (over a set of bins), at the unfolded level, using full information on correlations and uncertainty distributions

-Likelihood: can accommodate non-Gaussian distributions

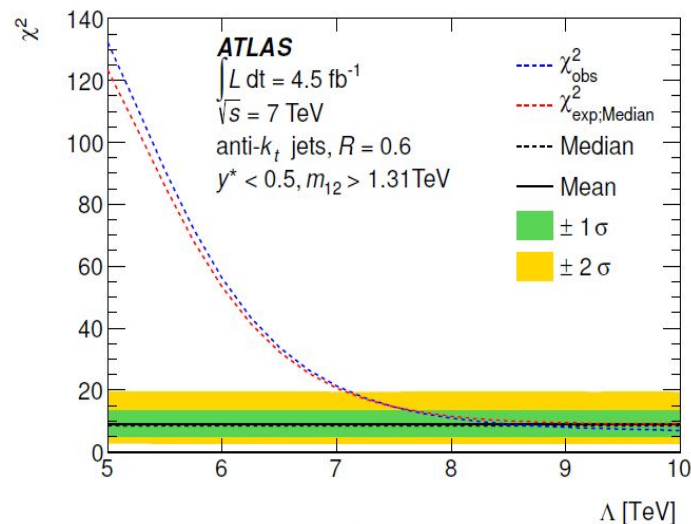
- $\chi^2$ (with correlations) standard definitions require symmetrization of uncertainties but can be generalized

→ Perform scan of parameters of theory model (ex:  $\Lambda$  for CI) and apply test statistic to compare with unfolded data: Obtain “observed”  $\chi^2(\Lambda)$

→ Generate toys with full set of uncertainties (transposed to SM+NP theory): comparison between nominal / fluctuated SM+NP : allows to reconstruct the distribution of the test statistic (important when using a  $\chi^2$  with error symmetrization).

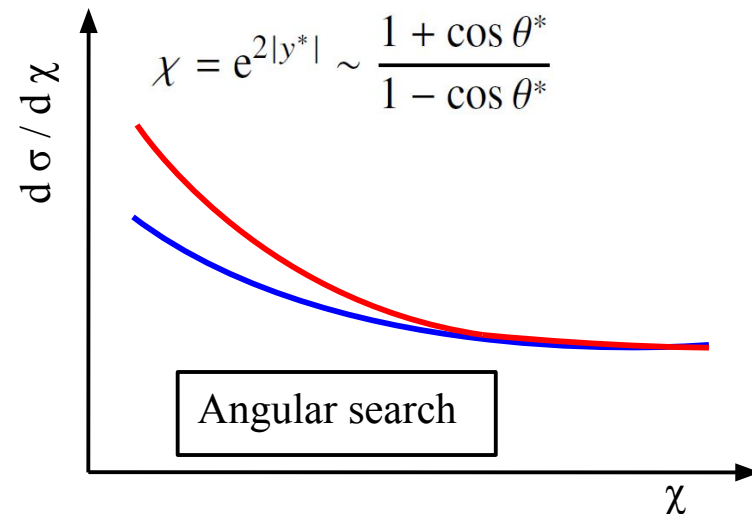
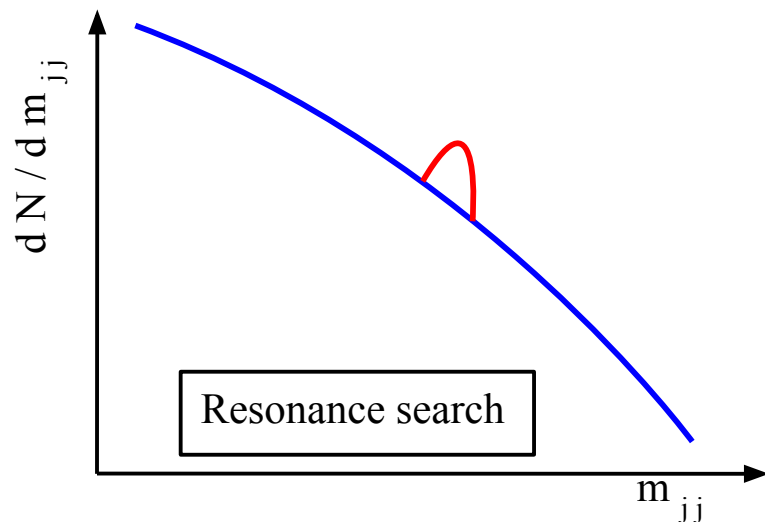
→ p-value to test data/QCD agreement

→ CLs =  $p_{s+b}/(1-p_b)$  for setting limits on New Physics



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# Resonance and angular searches



→ Sensitive to narrow resonances (QBH,  $q^*$  etc.)

→ Smooth QCD background described by smooth (fitted) function

→ Sensitive to resonant and non-resonant signals (CI, QBH etc.)

→ ~Flat QCD background described by MC, normalized to data

Tests of the folding-based approach:

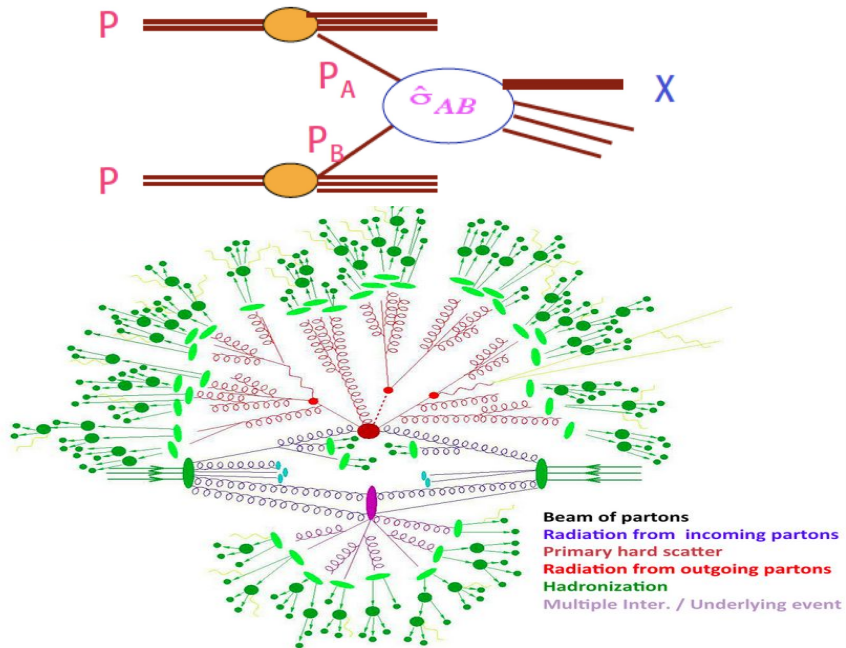
→ Detailed comparisons of transfer matrices for various signal models

→ Studied statistical uncertainties and correlations induced by transfer matrix – negligible impact

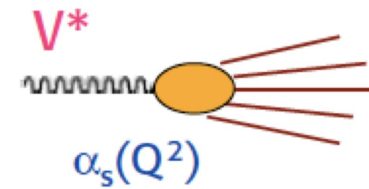
→ Studied possibility to use geometrical matching when building “A”

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# Comparison of LHC / FCCee “environments”



Pile-up



@ FCCee:

→ Short distance interaction of virtual bosons with quarks

→ No PDFs

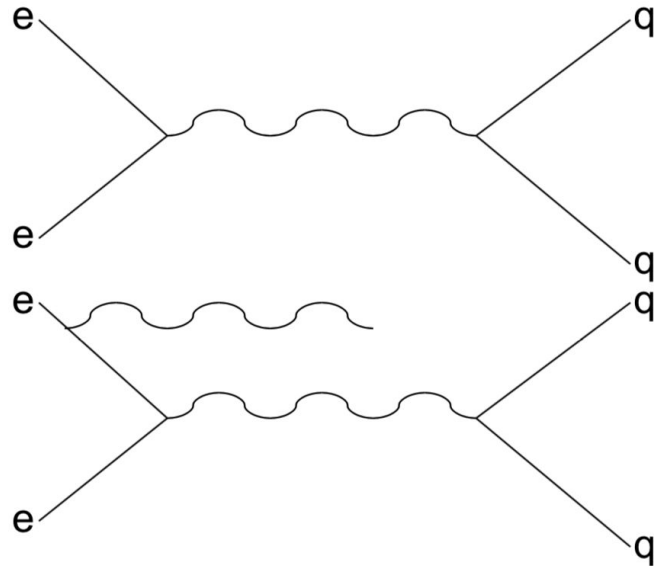
→ No underlying event & MPI

→ No pile-up

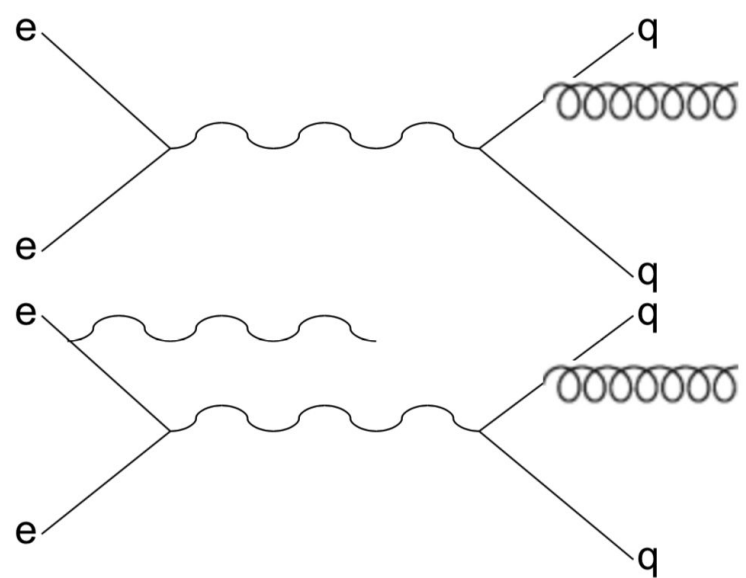
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# $\alpha_s$ evaluation from *(ISR) jet production*

N<sup>3</sup>LO + NNLL (arXiv:1902.08158)



N<sup>2</sup>LO + NLLA (arXiv:1205.3714)



→ Sensitivity to  $\alpha_s$  e.g. from 3/2 jet ratios (OR jet rates w.r.t. total hadronic Xsec)

→ High luminosity allows to select large samples of events with collinear / large angle ISR photons: allows to scan  $\sqrt{s}$ ' with the same detector and collider conditions – important for RGE test

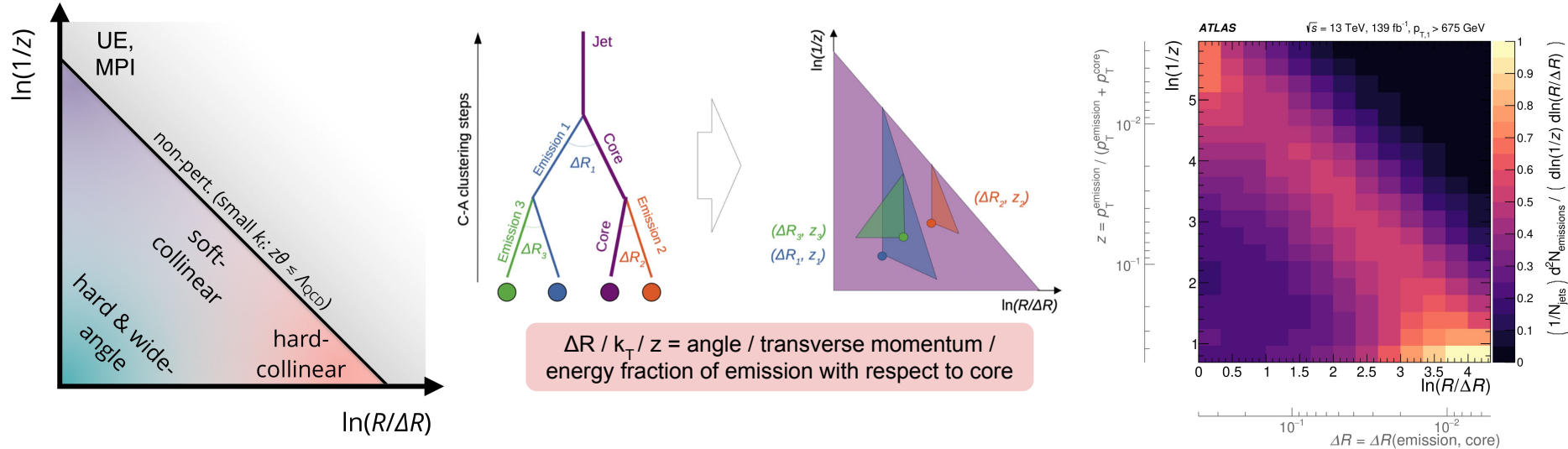
→ Need to study jet and photon energy calibration and resolution, acceptance and reconstruction efficiency etc. in view of *optimizing the detector design*

Should be able to target  $\delta\alpha_s / \alpha_s < 1\%$



# Opportunities for jet substructure studies

→ Numerous algorithms/methods developed for studying into detail the jet substructure in the LHC environment:  
 Important for understanding QCD effects inside jets, jet tagging (e.g. q/g, boosted top, H→bb), New Physics searches



[2004.03540](#)

→ The Lund Jet Plane allows for an effective way to distinguish (non)perturbative effects

→ Huge potential for doing precision studies of jet substructure in the clean FCC-ee environment

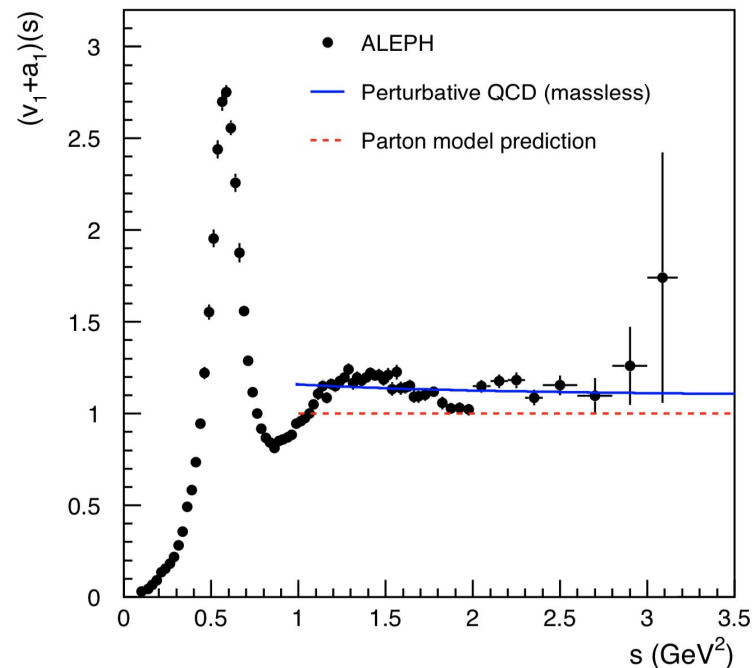
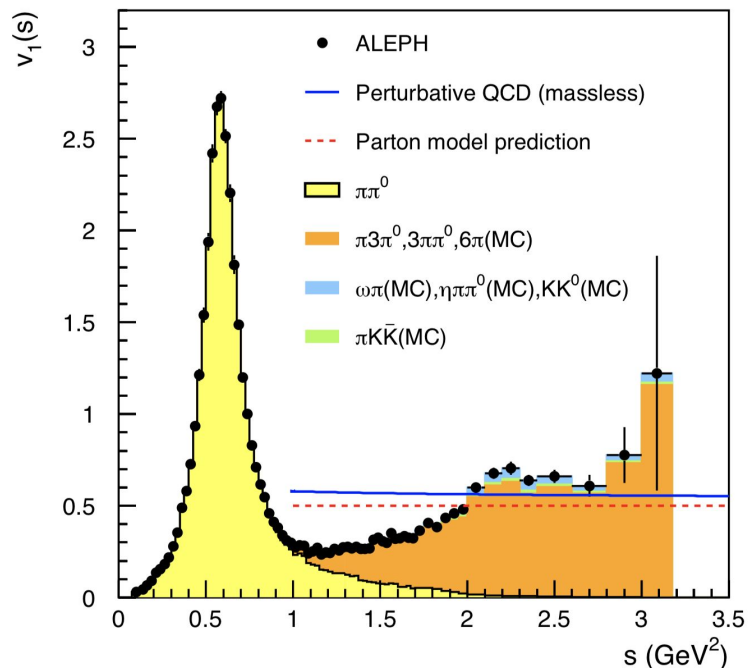
→ Need to *perform detector optimization* in terms of granularity, energy resolution, (tracking/calorimeter) acceptance

# $\alpha_S$ evaluation from *hadronic $\tau$ decays*

→  $\tau$  hadronic spectral functions (SFs) from ALEPH, unfolded of detector effects

→ Broadly used for  $(g-2)_\mu$  predictions and other QCD studies

1312.1501



$$v_1 / a_1 \left[ \tau^- \rightarrow V^- / A^- \bar{\nu}_\tau \right] \propto \frac{\text{BR} \left[ \tau^- \rightarrow V^- / A^- \bar{\nu}_\tau \right]}{\text{BR} \left[ \tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau \right]}$$

branching fractions

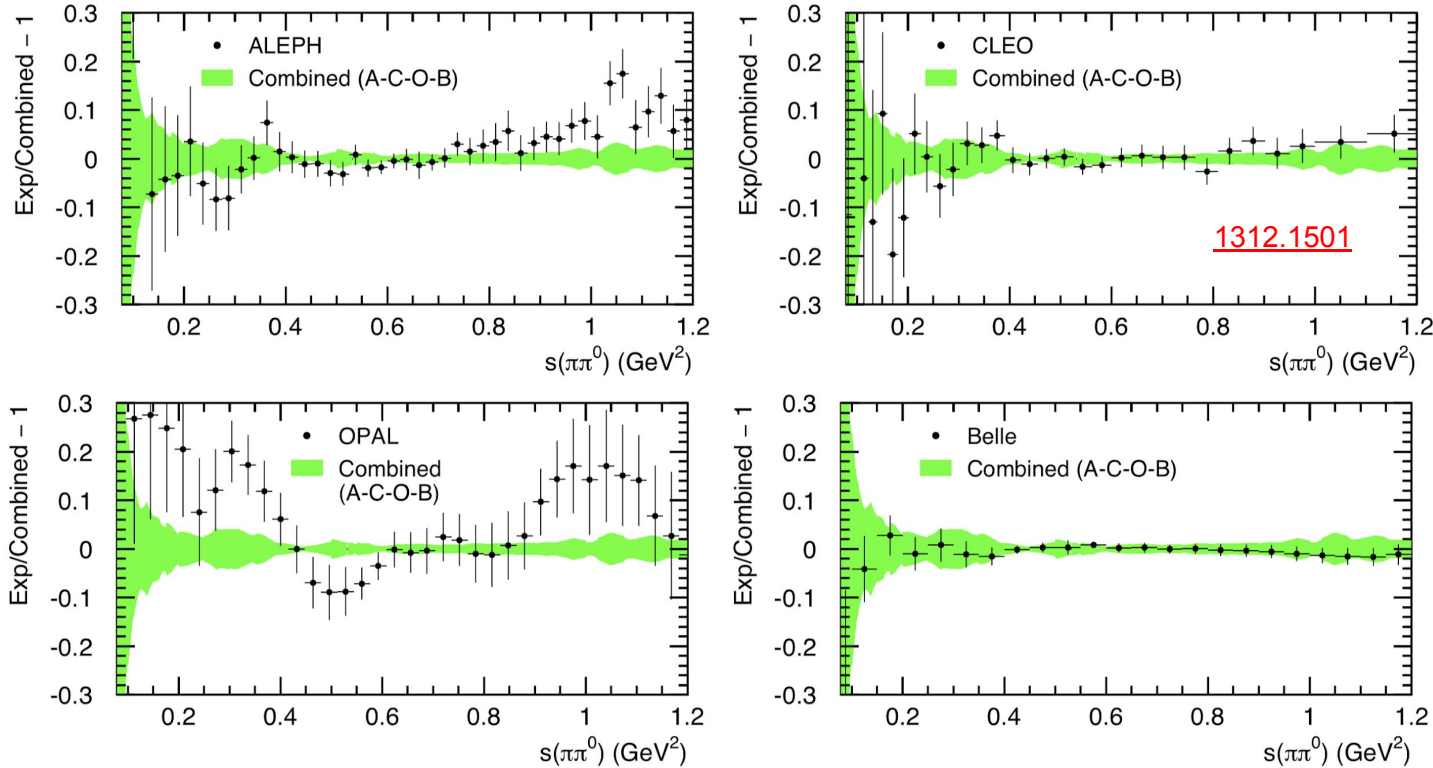
$$\frac{1}{N_{V/A}} \frac{dN_{V/A}}{ds} \frac{m_\tau^2}{(1-s/m_\tau^2)^2 (1+s/m_\tau^2)}$$

mass spectrum

kinematic factor

# $\alpha_s$ evaluation from *hadronic $\tau$ decays*

→  $\tau$  hadronic spectral functions ( $\pi\pi^0$  channel) from various experiments



→ Normalisation: branching fractions best determined by ALEPH (large boost, high granularity)

→ Shape best determined by Belle (high statistics); improvements @ Belle II

→ What precision can one achieve at FCCee? Need to study acceptance, reconstruction efficiency, resolution etc. in view of *optimizing the detector design* for SFs measurements

# $\alpha_S$ evaluation from *hadronic $\tau$ decays*

$$R_\tau = \frac{\Gamma(\tau^- \rightarrow \text{hadrons } \nu_\tau)}{\Gamma(\tau^- \rightarrow e^- \nu_e \nu_\tau)}$$

EW correction:  
 $\approx 0.0010$  (neglected)

Perturbative quark-mass terms:

$$\sum_{i,j=u,d,s} C_{U,ij}(s_0) \frac{m_i(s_0) \cdot m_j(s_0)}{s_0}$$

$$R_{\tau,U}(s_0) \propto |V_{CKM}|^2 S_{EW} \left( 1 + \delta^{(0)} + \delta'_{EW} + \delta_U^{(2,m_q)} + \sum_{D=4,6,\dots} \delta_U^{(D)} \right)$$

Perturbative contribution

$$-2\pi i \oint_{|S|=s_0} \frac{ds}{s_0} \left( 1 - 2\frac{s}{s_0} + 2\left(\frac{s}{s_0}\right)^3 - \left(\frac{s}{s_0}\right)^4 \right) \cdot D(s)$$

where:  $D(s) = \frac{1}{4\pi^2} \sum_{n=0}^{\infty} \tilde{K}_n(\xi) a_s^n(-\xi s)$

Adler function to avoid unphysical subtractions:

$$D(s) \equiv -s \frac{d\Pi(s)}{ds}$$

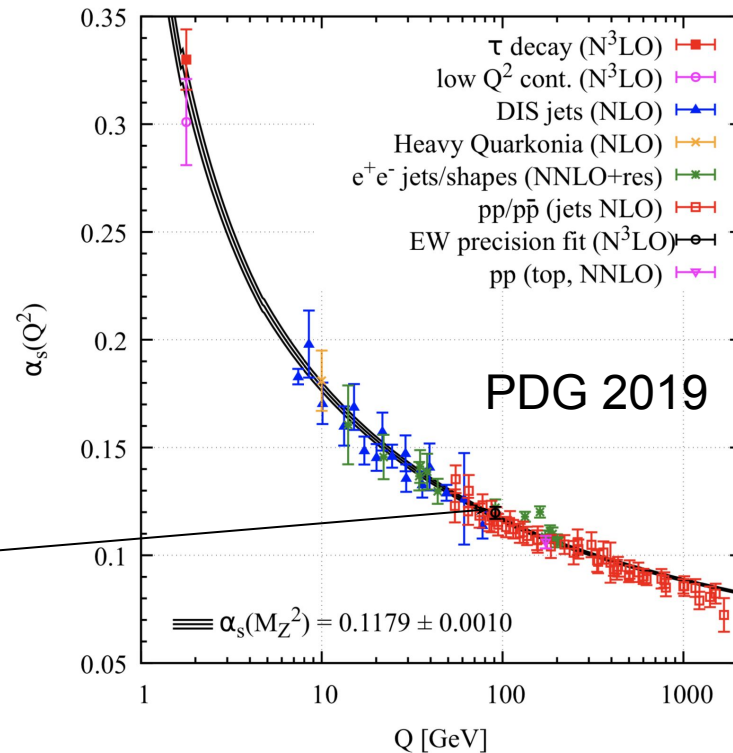
Nonperturbative contribution

$$\delta_U^{(D)} = (-1)^D \sum_{\dim O=D} C_U(\mu) \frac{\langle O(\mu) \rangle_U}{s_0^{D/2}}$$

- Theoretical prediction available at N<sup>3</sup>LO: need for even higher precision at the time of FCC-ee to reduce dominant uncertainty from perturbative series (CIPT/FOPT), to benefit from the statistical precision ( $\delta\alpha_S / \alpha_S \ll 1\%$ )
- More precise SFs will allow to better pin down non-perturbative corrections and probe the structure of the QCD vacuum (condensates)

# $\alpha_s$ evaluation from *hadronic Z decays*

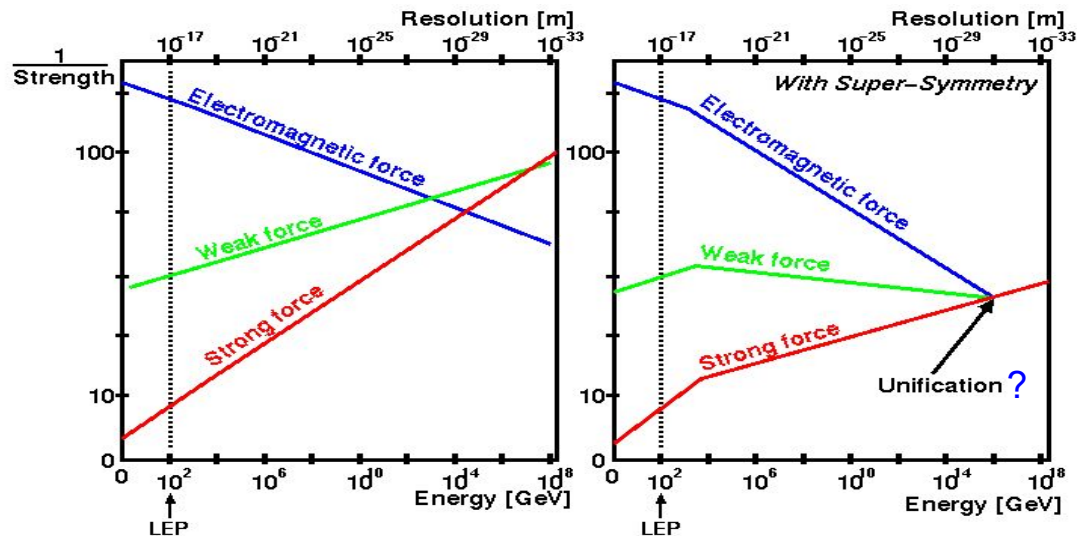
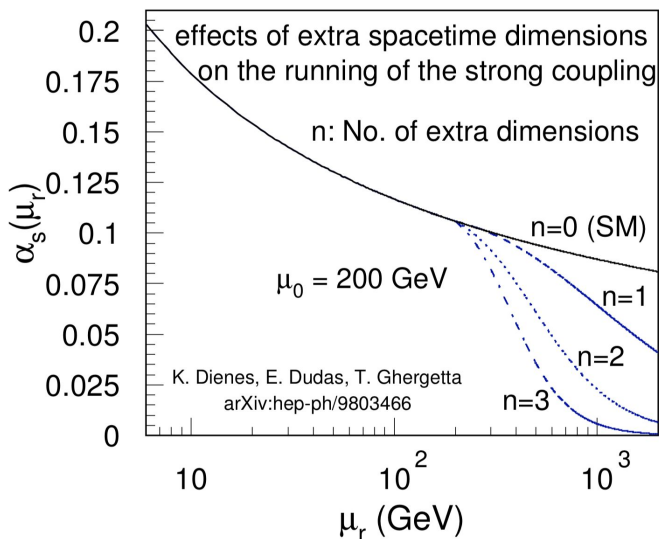
- Theoretical prediction available at N<sup>3</sup>LO
- Better convergence of the perturbative series and less non-perturbative corrections compared to precise determinations at lower scales (e.g. from  $\tau$  decays)



- Used for “reference value”:  
determinations at other energies  
evolved at the  $m_Z$  scale and then  
compared to *test the RGE from QCD*

- Need to study acceptance and reconstruction efficiency etc. in view of *optimizing detector design*

# Ultimate goal: *test RGE & unification of couplings*



- A deviation from the SM prediction for the RGE can be an indication of New Physics
- Are the coupling constants unified at the Planck scale?
- *Need to evaluate the strong coupling at multiple scales, with high precision*

