

# SEARCHING FOR NEW PHYSICS USING MUONIC TRANSITIONS

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based on Haxton, McElvain, Menzo, Rule, JZ, 2406.13818;  
Fox, Hostert, Menzo, Pospelov, JZ, 2407.03450; 2306.15631;  
Bigaran, Fox, Gouttenoire, Harnik, Krnjaic, Menzo, JZ, 2503.07722

Moriond EW, Mar 27 2025

# MUONS AND NEW PHYSICS

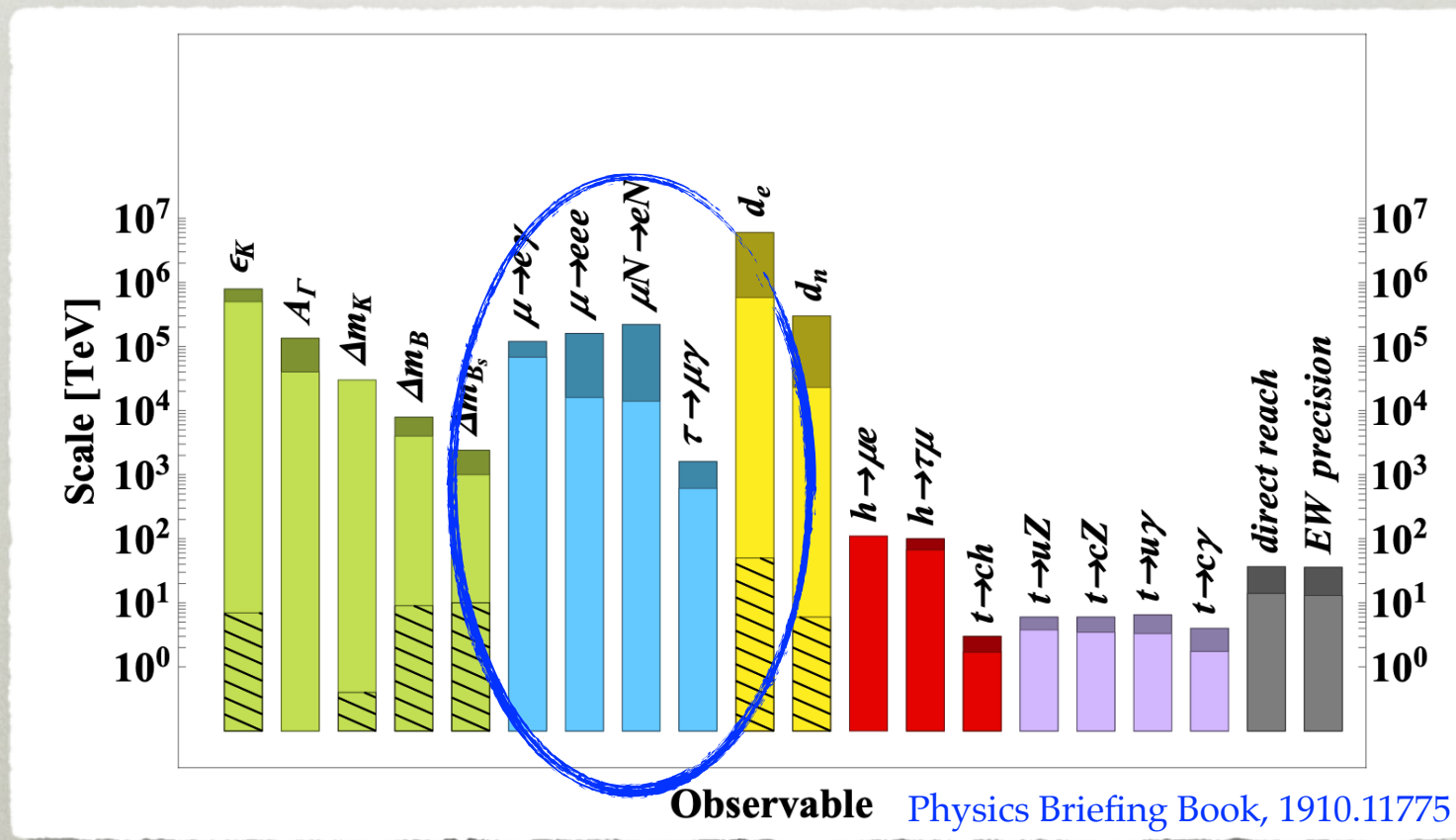
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- muon the lightest unstable particle in the SM
- relatively easy to produce  $\Rightarrow$  large samples available
- can use muons to search for
  - heavy new physics
  - light new physics

# SEARCHING FOR HEAVY NEW PHYSICS

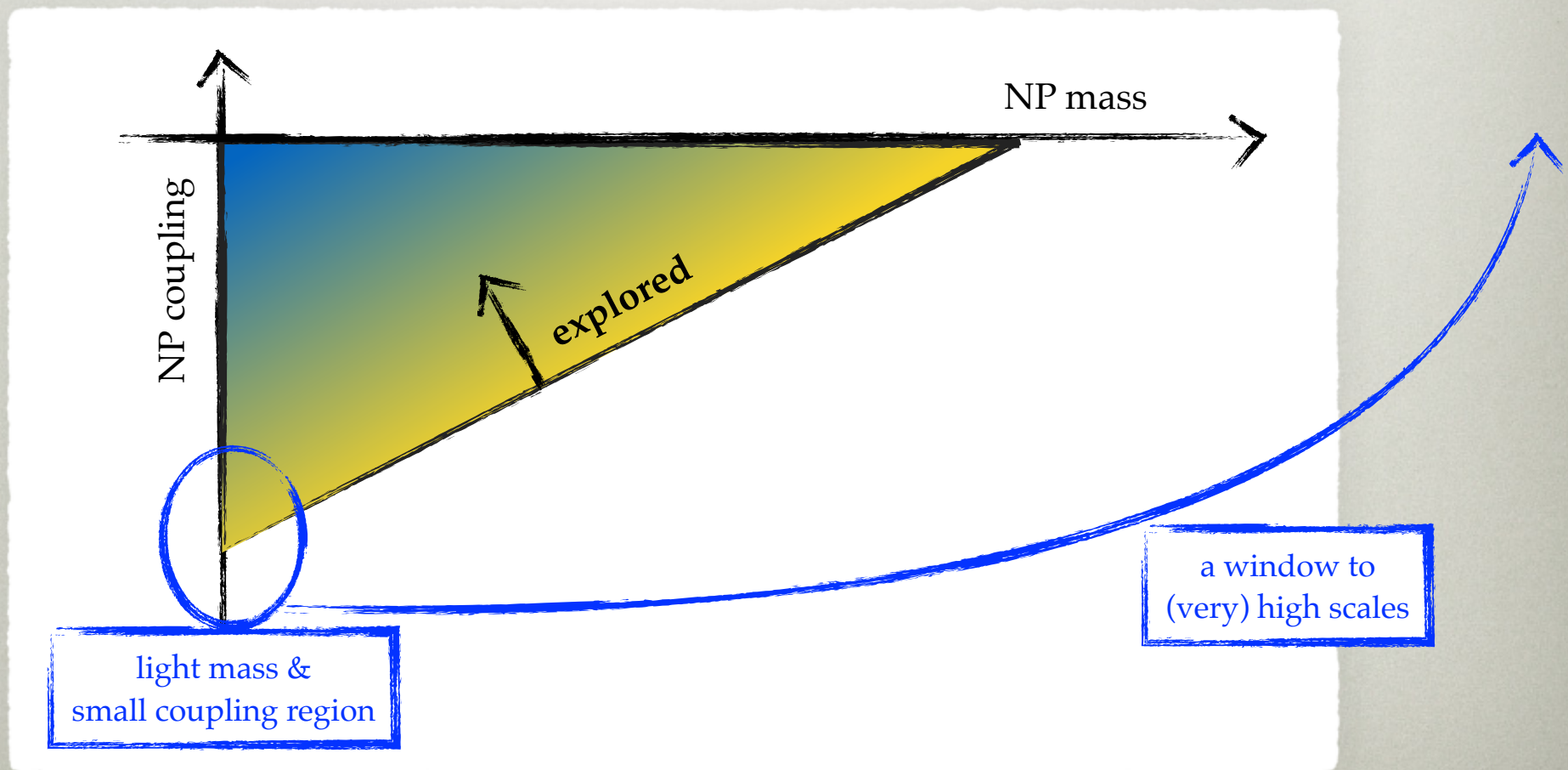
see Ana Teixeira's talk for an in depth overview

- high effective scales probed



# SEARCHING FOR LIGHT NEW PHYSICS

- light particles: a window to high UV dynamics
  - the gain comes from very small SM muon decay width



# THE REST OF THE TALK

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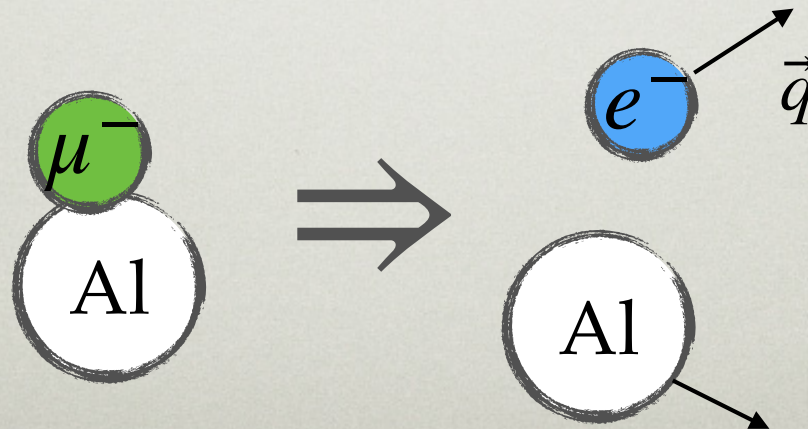
- heavy new physics
  - EFT based  $\mu \rightarrow e$  predictions
- searching for light NP using muons
  - flavor violating QCD axion in  $\mu \rightarrow ea$
  - time dependent searches in  $\mu \rightarrow e\phi$
  - $\mu \rightarrow 5e$
  - ....

EFT BASED  $\mu \rightarrow e$   
PREDICTIONS

# $\mu \rightarrow e$ KINEMATICS

Haxton, McElvain, Menzo, Rule, JZ, 2406.13818

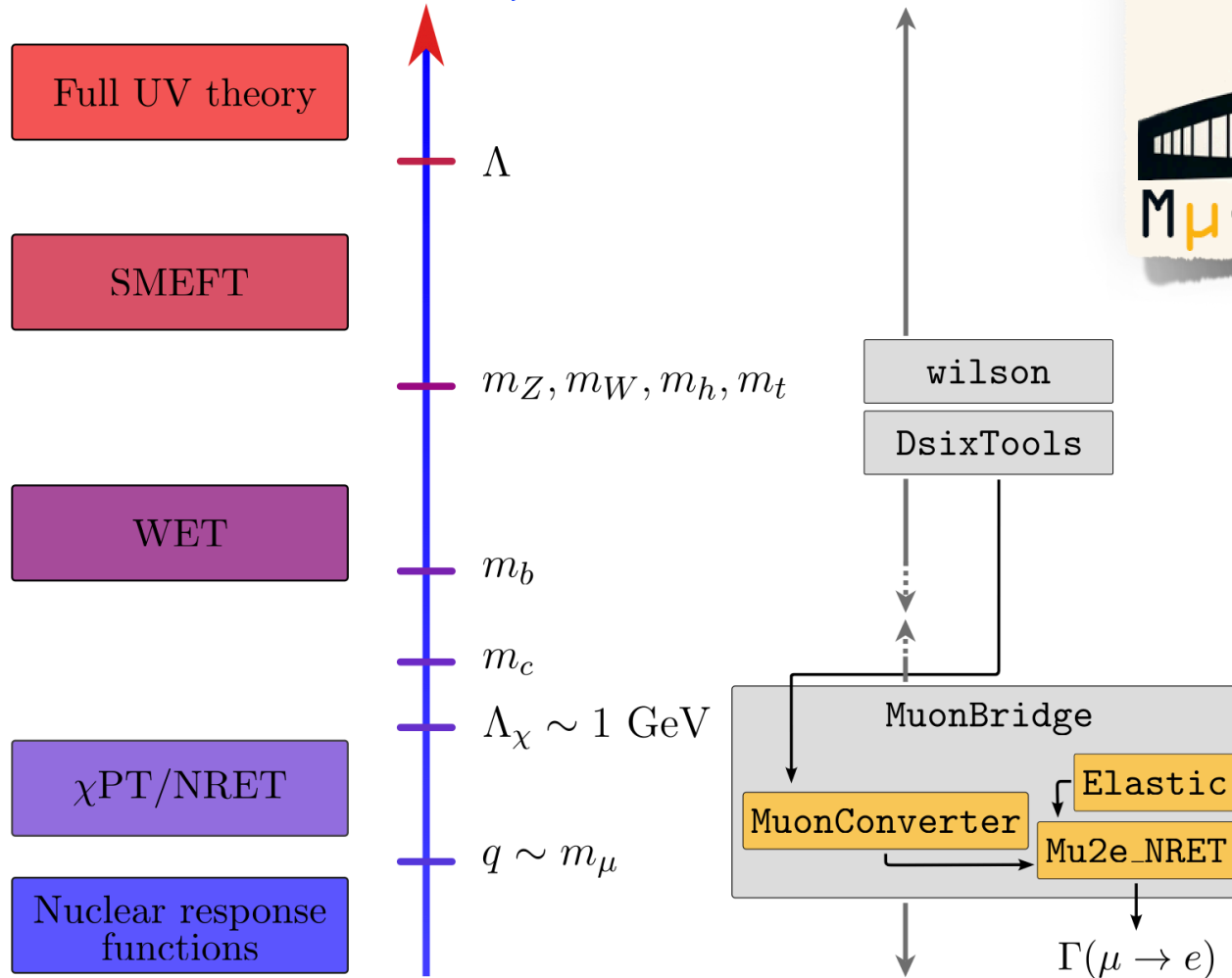
- initial state:  $\mu^-$  in 1s orbital
- final state: relativistic  $e^-$  with three momentum
  - $E_{\mu}^{\text{bind}} \ll m_{\mu}$  (for  $^{27}\text{Al}$   $E_{\mu}^{\text{bind}} \simeq 0.463$  MeV)  
 $\Rightarrow |\vec{q}| \sim \mathcal{O}(100 \text{ MeV})$



# TOWER OF EFTS

## ⇒ MUONBRIDGE CODE

Haxton, McElvain, Menzo, Rule, JZ, 2406.13818





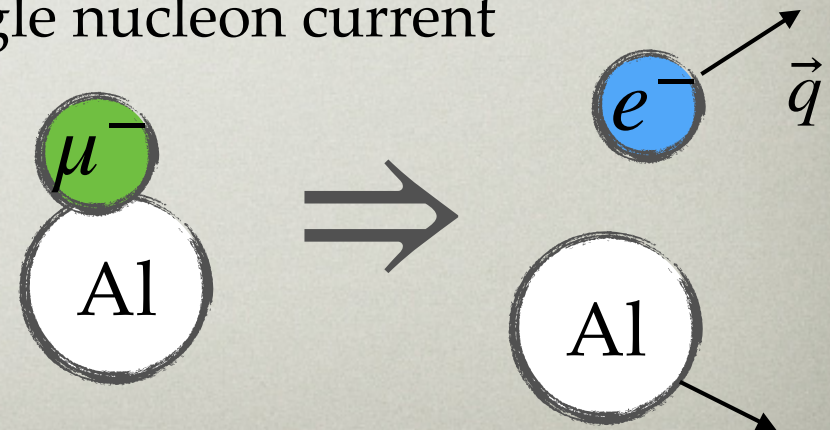
# NON-RELATIV. EXPANSION

- a hierarchy of small parameters

$$y \equiv \left(\frac{qb}{2}\right)^2 > |\vec{v}_N| > |\vec{v}_\mu| > |\vec{v}_T|$$

$b \sim$  nuclear size      $\vec{v}_N = (\vec{k}_1 + \vec{k}_2)/2$  average nucleon velocity     bound muon velocity     velocity of outgoing target nucleus

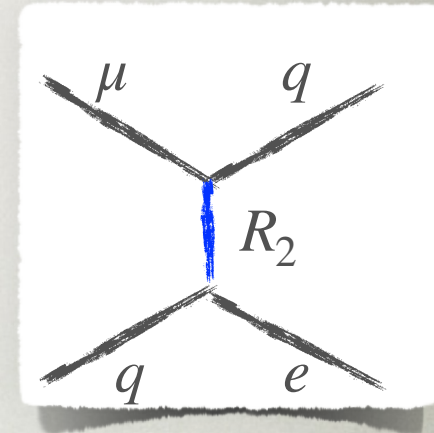
- $y \sim 0.2 - 0.5 \Rightarrow$  nuclear scales are being probed
- Chiral EFT: interactions with single nucleon current dominate
- can expand in  $v_N$  and  $v_\mu$ 
  - we keep  $\mathcal{O}(v_N)$ ,  $\mathcal{O}(v_\mu)$  terms



# LEPTOQUARK EXAMPLE

- scalar leptoquark  $R_2$  in the  $(3, 2, 7/6)$  of the SM gauge group

$$\mathcal{L} \supset y_{2ij}^{RL} \bar{u}_R^i R_2 L_L^j + y_{2ij}^{LR} \bar{e}_R^i R_2^* Q_L^j + \text{h.c.},$$

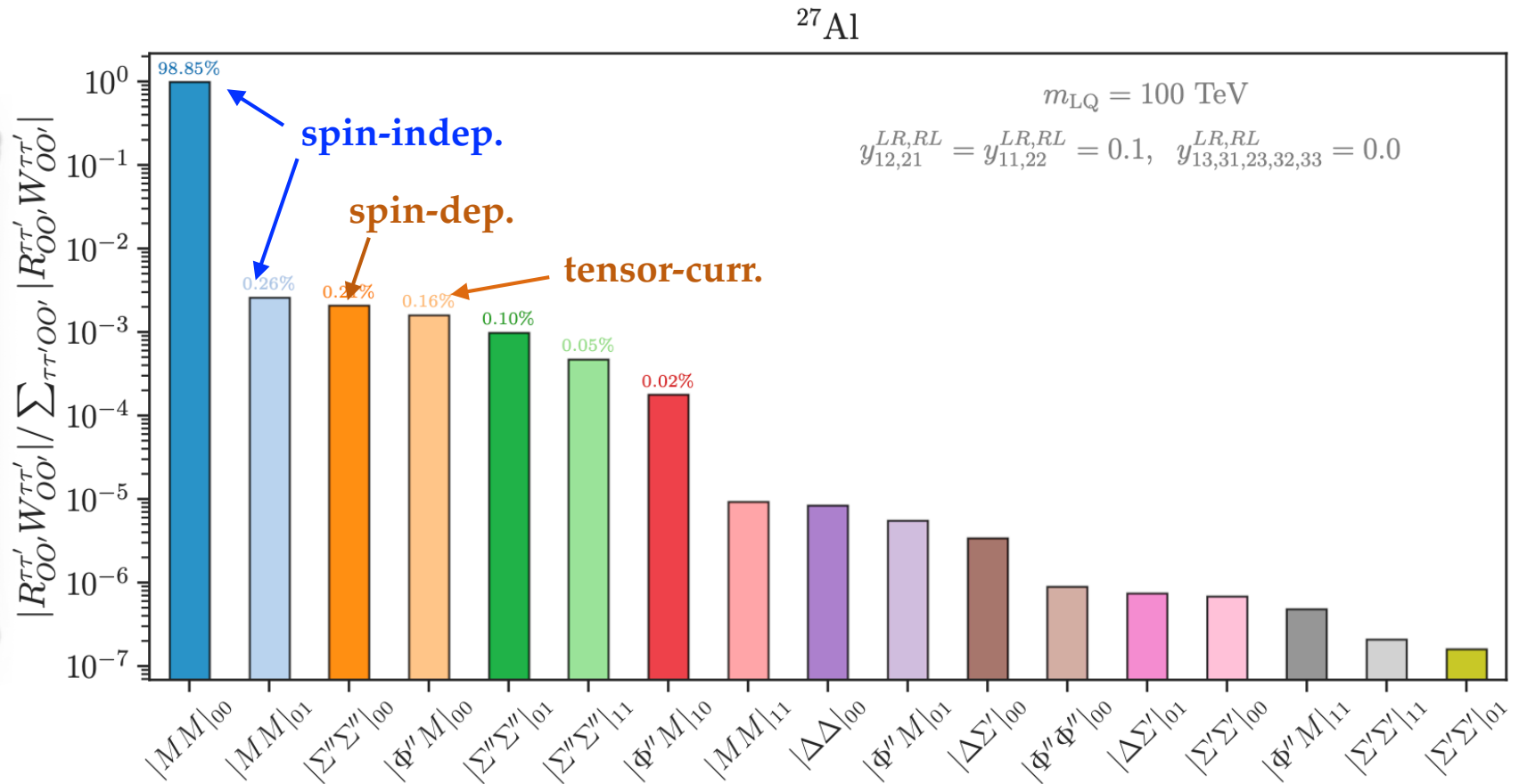


- integrating out  $R_2 \Rightarrow$  generates all 10 dim 6 ops in WET
  - including tensor currents
  - these have coherently enhanced contri. at subleading powers in  $\nu_N, \nu_\mu \Rightarrow$  kept in MuonBridge

# DIFFERENT CONTRIBS.

- typical point in the parameter space is dominated by spin independent contrib.

relative contribution.  
to the  $\mu \rightarrow e$  rate



# SEARCHING FOR LIGHT NEW PHYSICS

# SEARCHING FOR LIGHT NEW PHYSICS

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- can we use large datasets of stopped muons for light NP searches?
  - answer experiment dependent
- three examples
  - QCD axion in  $\mu \rightarrow ea$
  - time dependent  $\mu \rightarrow e\phi$
  - $\mu \rightarrow 5e$  at Mu3e

# QCD AXION

# AXION LIKE MODELS

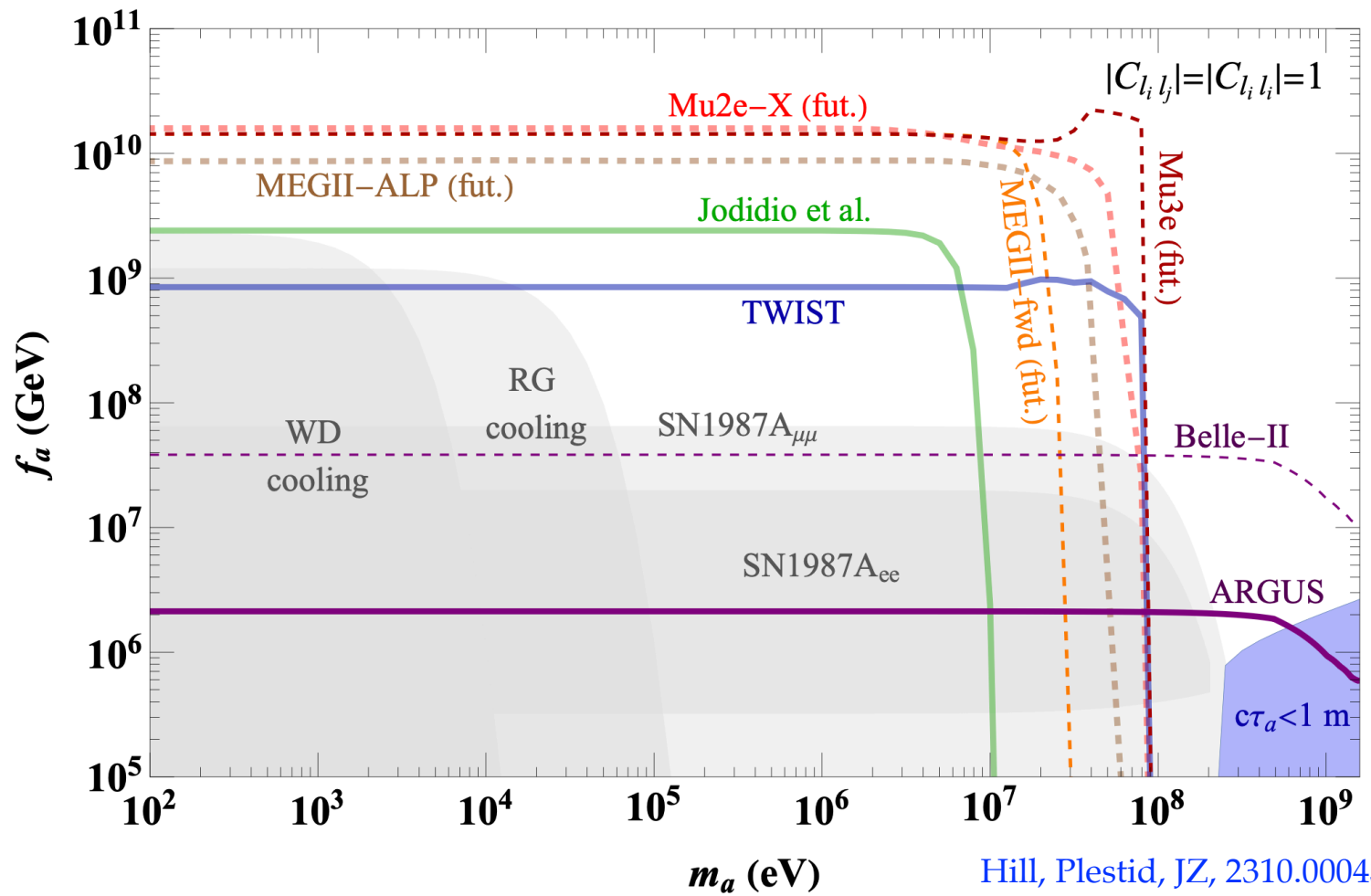
- any spontaneously broken global symmetry  $\Rightarrow$  (p)NGB
  - if "light enough" can be DM
- in general couplings to gluons, photons, SM fermions

$$\mathcal{L}_{\text{eff}} = \frac{\alpha_s}{8\pi} \frac{a}{f_a} G\tilde{G} + \frac{E}{N} \frac{\alpha_{\text{em}}}{8\pi} \frac{a}{f_a} F\tilde{F} + \frac{\partial_\mu a}{2f_a} \bar{f}_i \gamma^\mu (C_{f_i f_j}^V + C_{f_i f_j}^A \gamma_5) f_j$$

$$F_{f_i f_j}^{V,A} \equiv \frac{2f_a}{C_{f_i f_j}^{V,A}}$$

$$F_{l_i l_j} = \frac{2f_a}{\sqrt{|C_{l_i l_j}^V|^2 + |C_{l_i l_j}^A|^2}}$$

- in general ALPs will have flavor violating couplings
  - here focus on enhanced couplings to leptons



$(\gamma_5) f_j$

$$\frac{2f_a}{\sqrt{|C_{l_i l_j}|^2 + |C_{l_i l_j}^A|^2}}$$

- in general ALPs will have flavor violating couplings
  - here focus on enhanced couplings to leptons



TIME DEPENDENT

$$\mu \rightarrow e\phi$$

# NON-ABELIAN PNOB

Bigaran, Fox, Gouttenoire, Harnik, Krnjaic, Menzo, JZ, 2503.07722

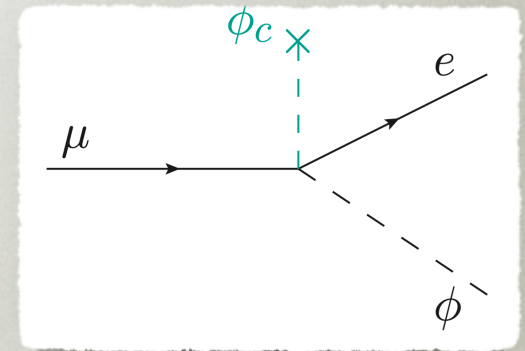
- if DM a non-Abelian pNOB the interactions with the SM of the form

$$\mathcal{L}_{\text{int}} \supset \frac{i\phi_c(\partial_\mu\phi)}{2f^2} \bar{\ell}_i \gamma^\mu (C_{ij}^V + C_{ij}^A \gamma_5) \ell_j,$$

- example in the SM:  $\pi^\pm$  interacting with leptons via photon exchange
- classical  $\phi$  background induces time dependent  $\mu \rightarrow e\phi$  decays

$$\phi_c(t) = \phi_0 \cos(m_\phi t + \delta)$$

$$\phi_0 = \frac{\sqrt{2\rho_\phi}}{m_\phi}$$



- time dep. searches can be more sensitive  
 $\Leftrightarrow$  for systematics dominated searches
- time dep. a smoking gun signal of DM
- can do the same type of search at MEG-II, Mu2e-X, COMET-X
  - can do time dep.  $\tau \rightarrow \ell\phi$  search at Belle II

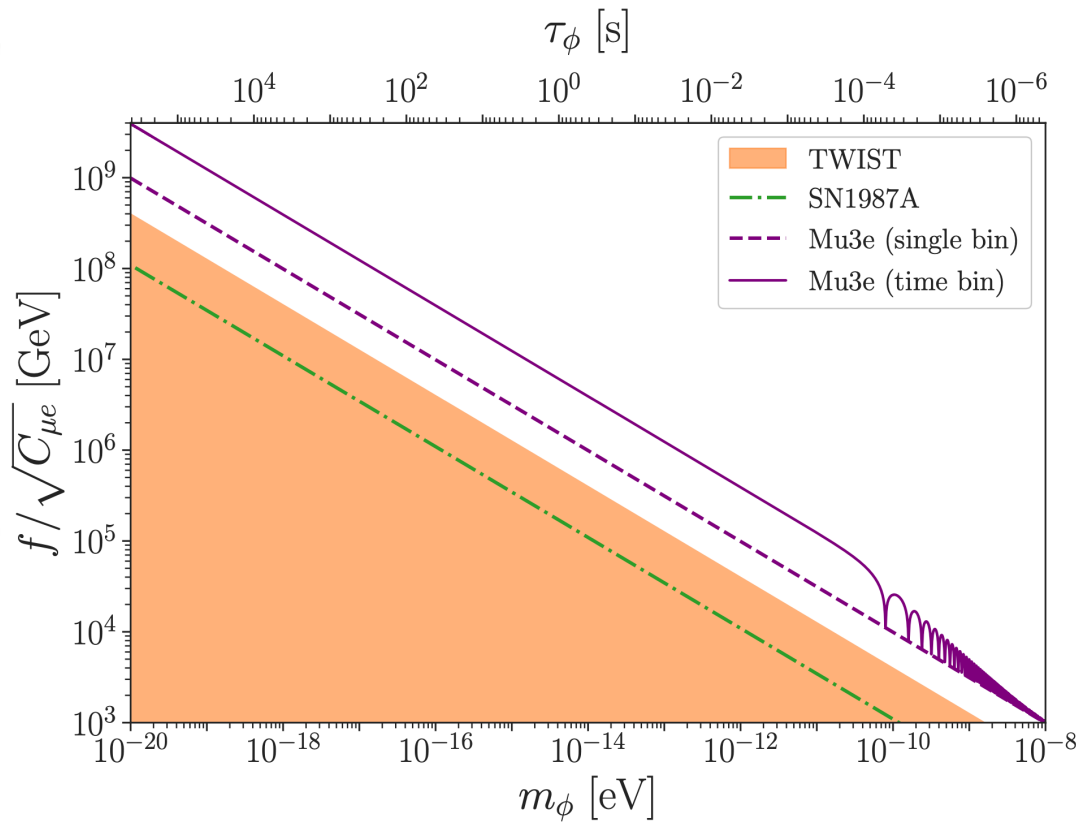
# AN PNGB

an, Fox, Gouttenoire, Harnik, Krnjaic, Menzo, JZ, 2503.07722

interactions with the SM of the form

$$(V_{ij} + C_{ij}^A \gamma_5) \ell_j,$$

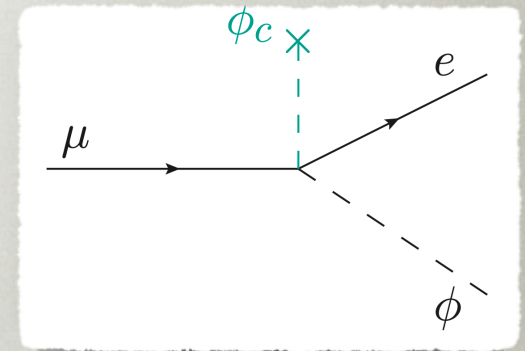
coupling with leptons via photon exchange



$$\phi_c(t) = \phi_0 \cos(m_\phi t + \delta)$$

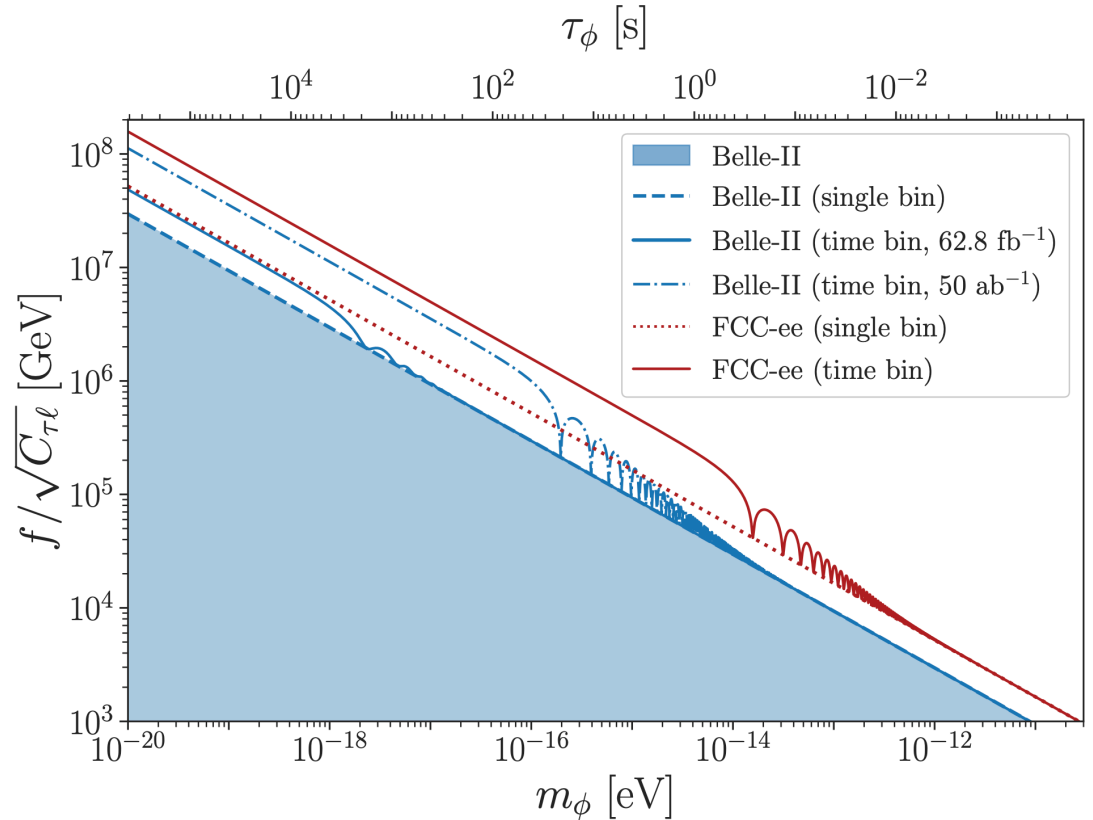
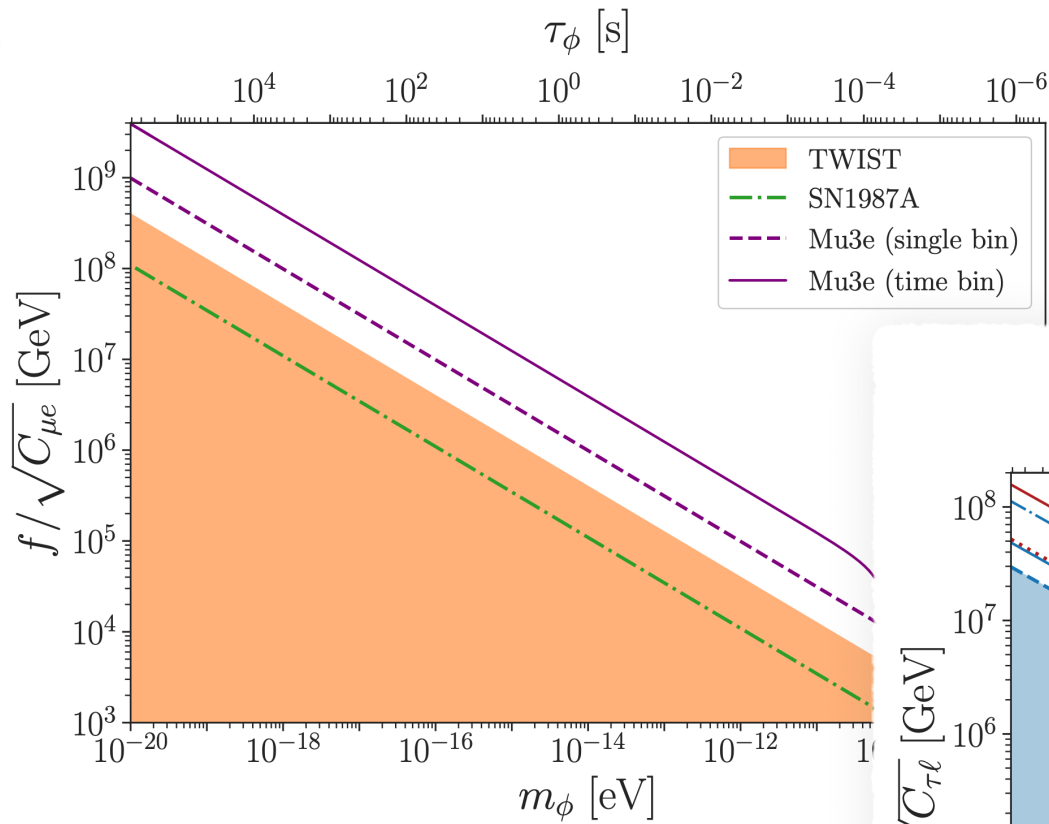
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in, Fox, Gouttenoire, Harnik, Krnjaic, Menzo, JZ, 2503.07722



$$\phi_c(t) = \phi_0 \cos(m_\phi t + \delta)$$

- time dep. searches can be  $\Leftrightarrow$  for systematics dominated
- time dep. a smoking gun signal of DIM
- can do the same type of search at MEG-II, Mu2e-X, COMET-X
  - can do time dep.  $\tau \rightarrow \ell \phi$  search at Belle II

$$\mu \rightarrow 5e$$

# A SIMPLE DARK SECTOR MODEL

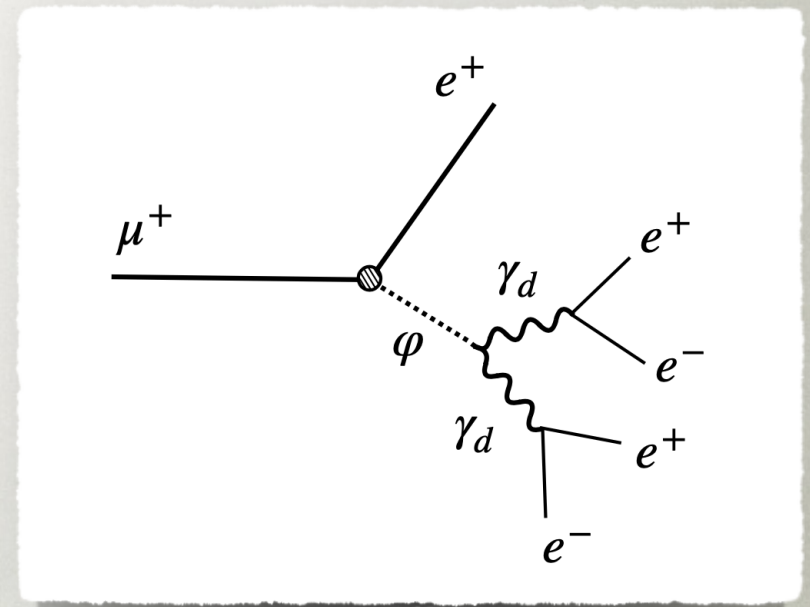
Hostert, Menzo, Pospelov, JZ, 2306.15631

- higgsed dark abelian gauge group  $U(1)_d$

- dark photon  $\gamma_d$
- light dark Higgs  $h_d$

- coupling to SM

- kinetic mixing
- flavor violating  
high dim op., e.g.,  $\frac{C_{ij}}{\Lambda} \bar{l}_i e_j H \phi$



- very conservative reach at Mu3e:  $\text{Br}(\mu \rightarrow 5e) \sim 10^{-12}$

- accepting all accidental bckgs.

- $\Rightarrow$  reach on effective UV scale,  $\Lambda \sim 10^{15}$  GeV

# OTHER LIGHT NEW PHYSICS SIGNATURES

- muon can only decay to  $e, \gamma$ , inv  $\Rightarrow$  a "finite" numb. of differ. observable signatures

$e \setminus \gamma$	0	1	2	3	4
1		$e\gamma$	$e2\gamma$	$e3\gamma$	$e4\gamma$
3	$3e$	$3e\gamma$	$3e2\gamma$	...	
5	$5e$	$5e\gamma$	...		
7	$7e$	...			
...					

MEG-II      MEG, 2005.00339  
 Mu3e      2306.15631  
 Greljo, Palavrić, Tunja, JZ, work in progress

- other options
  - displaced vertices in  $\mu \rightarrow 3e$       Knapen, Opferkuch, Redigolo, Tammaro, 2410.13941
  - $\mu \rightarrow 3ea$       Knapen, Langhoff, Opferkuch, Redigolo, 2311.17915
  - BNV annihilation  $\mu^- p \rightarrow$  dark sect.      Fox, Hostert, Menzo, Pospelov, JZ, 2407.03450
  - ....

# CONCLUSIONS

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- EFT approach well suited for predicting the  $\mu \rightarrow e$  conversion rates
  - results available in the form of a public code **MuonBridge**
- rare muon decays can be use to search for light NP
  - QCD axion, time dep. npNGB DM,  $\mu \rightarrow 5e, \dots$



# BACKUP SLIDES

# LIGHT NEW PHYSICS $\Rightarrow$ PROBE OF HIGH SCALES

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- rare decays into a light state,  $X$ , e.g.,  $\mu \rightarrow e\phi$ ,
  - exquisite probes of UV physics
- parametric gains compared to probing NP through dim-6 ops
  - the reason is that the SM decay widths are power suppressed  $\Gamma_\ell \propto m_\ell^5/m_W^4$
- if  $\mu \rightarrow e\phi$  through dim 5 op. suppressed by  $1/f_\phi \Rightarrow Br(\mu \rightarrow e\phi) \propto (m_W^2/f_\phi m_\mu)^2$

# WET

Haxton, McElvain, Menzo, Rule, JZ, 2406.13818

- 10 dimension 6 ops

$$Q_{1,q}^{(6)} = (\bar{e}\gamma_\alpha\mu)(\bar{q}\gamma^\alpha q),$$

$$Q_{3,q}^{(6)} = (\bar{e}\gamma_\alpha\mu)(\bar{q}\gamma^\alpha\gamma_5q),$$

$$Q_{5,q}^{(6)} = (\bar{e}\mu)(\bar{q}q),$$

$$Q_{7,q}^{(6)} = (\bar{e}\mu)(\bar{q}i\gamma_5q),$$

$$Q_{9,q}^{(6)} = (\bar{e}\sigma^{\alpha\beta}\mu)(\bar{q}\sigma_{\alpha\beta}q),$$

$$Q_{2,q}^{(6)} = (\bar{e}\gamma_\alpha\gamma_5\mu)(\bar{q}\gamma^\alpha q),$$

$$Q_{4,q}^{(6)} = (\bar{e}\gamma_\alpha\gamma_5\mu)(\bar{q}\gamma^\alpha\gamma_5q).$$

$$Q_{6,q}^{(6)} = (\bar{e}i\gamma_5\mu)(\bar{q}q),$$

$$Q_{8,q}^{(6)} = (\bar{e}i\gamma_5\mu)(\bar{q}i\gamma_5q),$$

$$Q_{10,q}^{(6)} = (\bar{e}i\sigma^{\alpha\beta}\gamma_5\mu)(\bar{q}\sigma_{\alpha\beta}q).$$

- additional 16 operators at dimension 7
- related to other WET bases used in the literature by linear transf.
- note: tensor currents appear already at dimension 6

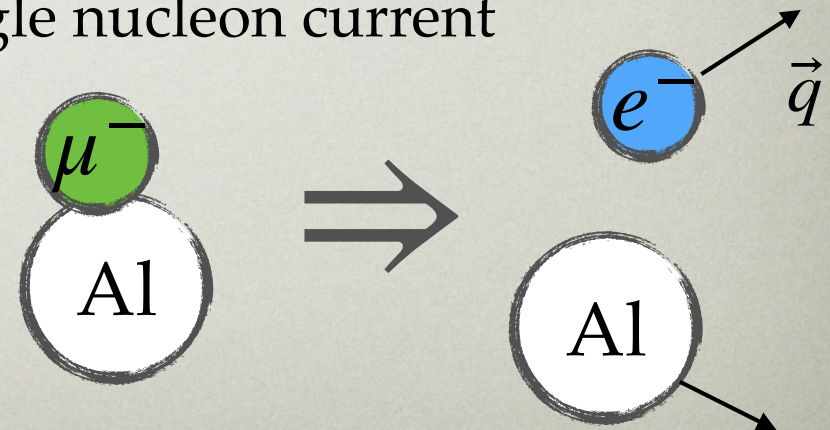
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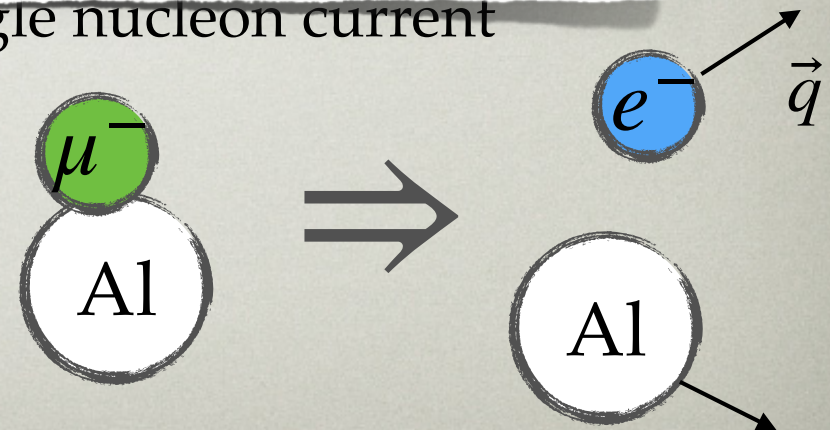
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# NON-RELATIVISTIC LIMIT

- all
 

$\mathcal{O}_1 = 1_L 1_N,$	$\mathcal{O}'_2 = 1_L i\hat{q} \cdot \vec{v}_N,$
$\mathcal{O}_3 = 1_L i\hat{q} \cdot [\vec{v}_N \times \vec{\sigma}_N],$	$\mathcal{O}_4 = \vec{\sigma}_L \cdot \vec{\sigma}_N,$
$\mathcal{O}_5 = \vec{\sigma}_L \cdot (i\hat{q} \times \vec{v}_N),$	$\mathcal{O}_6 = i\hat{q} \cdot \vec{\sigma}_L i\hat{q} \cdot \vec{\sigma}_N,$
$\mathcal{O}_7 = 1_L \vec{v}_N \cdot \vec{\sigma}_N,$	$\mathcal{O}_8 = \vec{\sigma}_L \cdot \vec{v}_N,$
$\mathcal{O}_9 = \vec{\sigma}_L \cdot (i\hat{q} \times \vec{\sigma}_N),$	$\mathcal{O}_{10} = 1_L i\hat{q} \cdot \vec{\sigma}_N,$
$\mathcal{O}_{11} = i\hat{q} \cdot \vec{\sigma}_L 1_N,$	$\mathcal{O}_{12} = \vec{\sigma}_L \cdot [\vec{v}_N \times \vec{\sigma}_N],$
$\mathcal{O}'_{13} = \vec{\sigma}_L \cdot (i\hat{q} \times [\vec{v}_N \times \vec{\sigma}_N]),$	$\mathcal{O}_{14} = i\hat{q} \cdot \vec{\sigma}_L \vec{v}_N \cdot \vec{\sigma}_N,$
$\mathcal{O}_{15} = i\hat{q} \cdot \vec{\sigma}_L i\hat{q} \cdot [\vec{v}_N \times \vec{\sigma}_N],$	$\mathcal{O}'_{16} = i\hat{q} \cdot \vec{\sigma}_L i\hat{q} \cdot \vec{v}_N.$
- y
- Chiral EFT: interactions with single nucleon current dominate
- can expand in  $v_N$  and  $v_\mu$ 
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# MATCHING ONTO NRET

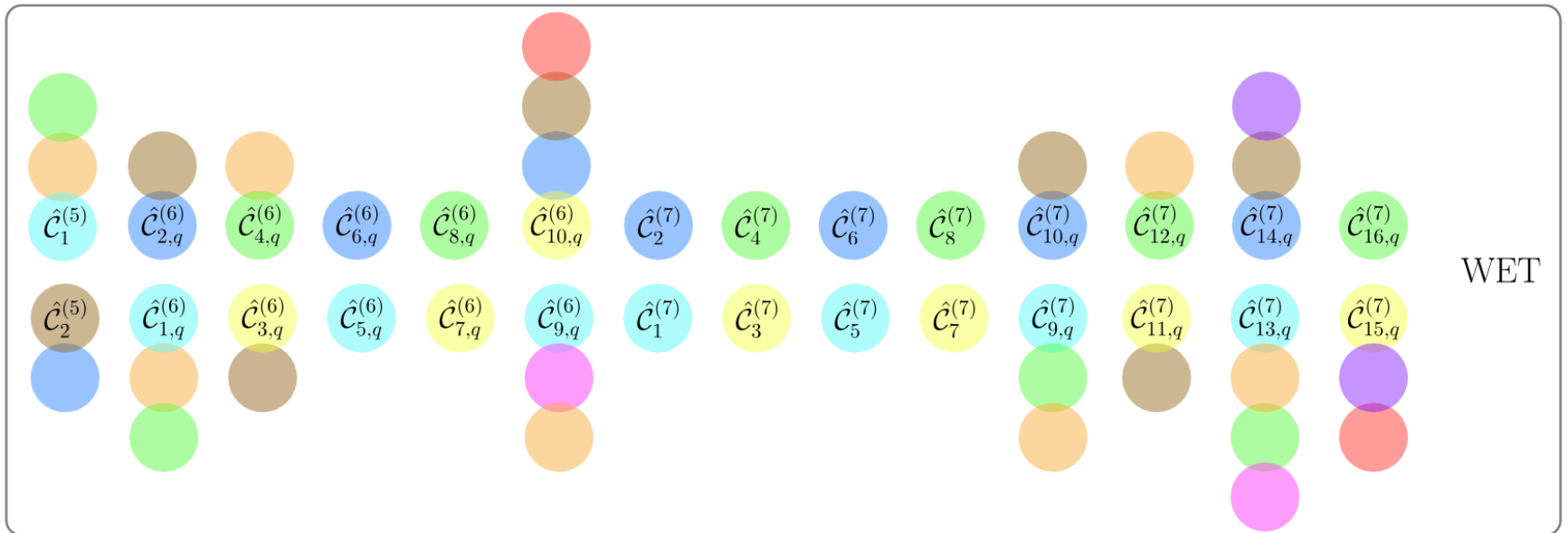
- NRET effective Lagrangian

$$\mathcal{L}_{\text{eff}}^{\text{NRET}} = \sum_{N=n,p} \sum_{i=1}^{16} c_i^N \mathcal{O}_i^N + \dots,$$

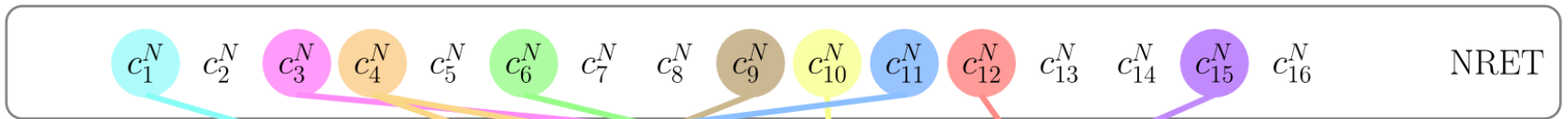
- low energy coefficients  $c_i^N$  functions of  $\vec{q}_{\text{eff}}^2$ 
  - for  $\mu \rightarrow e$  this is a constant
  - their values from nonperturbative matching of WET to NRET
    - follow from nucleon matrix elements  $\langle N | \mathcal{O}_i | N \rangle$
- for  $\mu \rightarrow e$  transition rate prediction need nuclear physics  $|\langle A | \mathcal{O}_i | A \rangle|^2$ 
  - nuclear response functions  $W_i$

$$W_M \sim \mathcal{O}(A^2) \gg \left\{ W_{\Sigma'}, W_{\Sigma''}, \frac{q_{\text{eff}}^2}{m_N^2} W_{\tilde{\Phi}''} \right\} \gg \left\{ \frac{q_{\text{eff}}^2}{m_N^2} W_{\Delta}, \frac{q_{\text{eff}}^2}{m_N^2} W_{\tilde{\Phi}'} \right\}$$

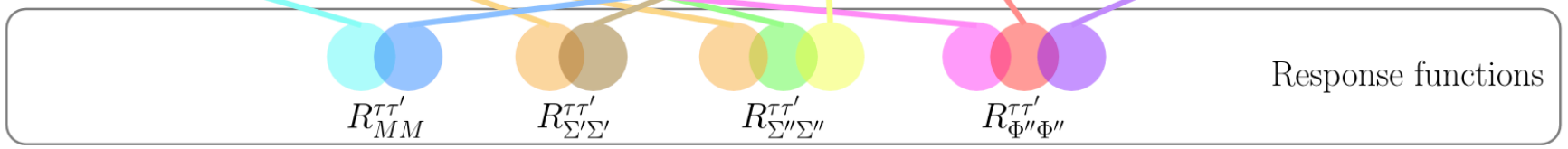
- $\mu \rightarrow e$  rate:  $\Gamma(\mu \rightarrow e) \propto \sum_i R_i W_i$



WET



NRET



Response functions

• for  $\mu \rightarrow e$  transition rate prediction need nuclear physics  $|\langle N | \mathcal{O}_i | N \rangle|$

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- $\mu \rightarrow e$  rate:  $\Gamma(\mu \rightarrow e) \propto \sum_i R_i W_i$

# MUONBRIDGE

- the code/repository **MuonBridge** consists of three modules

- **MuonConverter**: matches WET to NRET

- can interface with RG running codes

- **Mu2e\_NRET**: calculates the  $\mu \rightarrow e$  rate

$$B(\mu^- \rightarrow e^-) = \frac{\Gamma[\mu^- + (A, Z) \rightarrow e^- + (A, Z)]}{\Gamma[\mu^- + (A, Z) \rightarrow \nu_\mu + (A, Z - 1)]},$$

- particle physics input from **MuonConverter**, i.e., WET Wilson coeffs.  $C_i$
- **Elastic**: a database of shell model density matrices for calculating nuclear form factors
- comes in both **Python** and **Mathematica** versions





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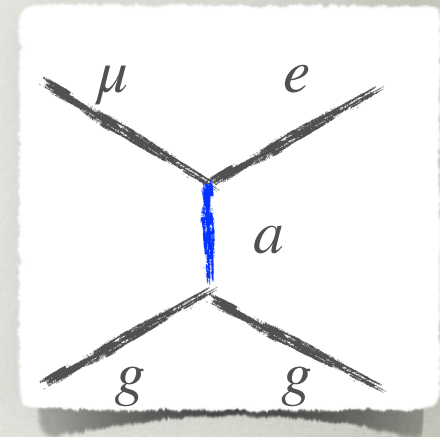
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- $\mu \rightarrow e$  rate:  $\Gamma(\mu \rightarrow e) \propto \sum_i R_i W_i$

# LIGHT ALP

- the same formalism trivially extends to light mediators
- example light ALP coupling to  $\mu e$  and gluons
- strictly speaking WET no longer an appropriate EFT
  - but trivial fix, allow Wilson coeffs to be  $q^2$  dependent,  $C_i \propto 1/(m_a^2 - q^2)$
  - since  $a$  only weakly couples to gluons: corrections to QCD can be neglected, i.e., just an external probe
  - in  $\mu \rightarrow e$  the  $q$  is fixed, so  $C_i$  are even constants



# MATCHING ONTO NRET

- NRET effective Lagrangian

$$\mathcal{L}_{\text{eff}}^{\text{NRET}} = \sum_{N=n,p} \sum_{i=1}^{16} c_i^N \mathcal{O}_i^N + \dots,$$

- low energy coefficients are functions of  $\vec{q}_{\text{eff}}^2$  in general
  - for  $\mu \rightarrow e$  this is a constant
- their values from nonperturbative matching of WET to NRET
  - follow from nucleon matrix elements
  - for instance

$$c_1^N = \sum_q \frac{1}{m_q} \hat{C}_{5,q}^{(6)} F_S^{q/N} + \hat{C}_1^{(7)} F_G^N + \hat{C}_5^{(7)} F_\gamma^N,$$
$$+ \sum_q \hat{C}_{1,q}^{(6)} F_1^{q/N} + m_+ \sum_q \hat{C}_{9,q}^{(7)} F_1^{q/N} - \sum_q \frac{q^2}{2m_N m_q} \hat{C}_{13,q}^{(7)} (F_{T,1}^{q/N} - 4F_{T,2}^{q/N})$$

• NR

$$\langle N' | \bar{q} \gamma^\mu q | N \rangle = \bar{u}'_N \left[ F_1^{q/N}(q^2) \gamma^\mu - \frac{i}{2m_N} F_2^{q/N}(q^2) \sigma^{\mu\nu} q_\nu \right] u_N,$$

$$\langle N' | \bar{q} \gamma^\mu \gamma_5 q | N \rangle = \bar{u}'_N \left[ F_A^{q/N}(q^2) \gamma^\mu \gamma_5 - \frac{1}{2m_N} F_{P'}^{q/N}(q^2) \gamma_5 q^\mu \right] u_N,$$

$$\langle N' | m_q \bar{q} q | N \rangle = F_S^{q/N}(q^2) \bar{u}'_N u_N,$$

$$\langle N' | m_q \bar{q} i \gamma_5 q | N \rangle = F_P^{q/N}(q^2) \bar{u}'_N i \gamma_5 u_N,$$

$$\langle N' | \frac{\alpha_s}{12\pi} G^{a\mu\nu} G_{\mu\nu}^a | N \rangle = F_G^N(q^2) \bar{u}'_N u_N,$$

$$\langle N' | \frac{\alpha_s}{8\pi} G^{a\mu\nu} \tilde{G}_{\mu\nu}^a | N \rangle = F_{\tilde{G}}^N(q^2) \bar{u}'_N i \gamma_5 u_N,$$

$$\langle N' | m_q \bar{q} \sigma^{\mu\nu} q | N \rangle = \bar{u}'_N \left[ F_{T,0}^{q/N}(q^2) \sigma^{\mu\nu} - \frac{i}{2m_N} \gamma^{[\mu} q^{\nu]} F_{T,1}^{q/N}(q^2) - \frac{i}{m_N^2} q^{[\mu} k_{12}^{\nu]} F_{T,2}^{q/N}(q^2) \right] u_N,$$

• the

$$\langle N' | \frac{\alpha}{12\pi} F^{\mu\nu} F_{\mu\nu} | N \rangle = F_\gamma^N(q^2) \bar{u}'_N u_N,$$

$$\langle N' | \frac{\alpha}{8\pi} F^{\mu\nu} \tilde{F}_{\mu\nu} | N \rangle = F_{\tilde{\gamma}}^N(q^2) \bar{u}'_N i \gamma_5 u_N.$$

• for instance

$$c_1^N = \sum_q \frac{1}{m_q} \hat{C}_{5,q}^{(6)} F_S^{q/N} + \hat{C}_1^{(7)} F_G^N + \hat{C}_5^{(7)} F_\gamma^N,$$

$$+ \sum_q \hat{C}_{1,q}^{(6)} F_1^{q/N} + m_+ \sum_q \hat{C}_{9,q}^{(7)} F_1^{q/N} - \sum_q \frac{q^2}{2m_N m_q} \hat{C}_{13,q}^{(7)} \left( F_{T,1}^{q/N} - 4F_{T,2}^{q/N} \right)$$

# CONVERSION RATE

- matrix elements squared of NRET operators leads to nuclear response functions  $W_i$

$$\Gamma(\mu \rightarrow e) = \frac{G_F^2}{\pi} \frac{q_{\text{eff}}^2}{1 + \frac{q}{M_T}} |\phi_{1s}^{Z_{\text{eff}}}(\vec{0})|^2 \sum_{\tau=0,1} \sum_{\tau'=0,1} \left\{ \begin{aligned} & \left[ \tilde{R}_{MM}^{\tau\tau'} W_{MM}^{\tau\tau'}(q_{\text{eff}}) + \tilde{R}_{\Sigma''\Sigma''}^{\tau\tau'} W_{\Sigma''\Sigma''}^{\tau\tau'}(q_{\text{eff}}) + \tilde{R}_{\Sigma'\Sigma'}^{\tau\tau'} W_{\Sigma'\Sigma'}^{\tau\tau'}(q_{\text{eff}}) \right] \\ & + \frac{q_{\text{eff}}^2}{m_N^2} \left[ \tilde{R}_{\Phi''\Phi''}^{\tau\tau'} W_{\Phi''\Phi''}^{\tau\tau'}(q_{\text{eff}}) + \tilde{R}_{\tilde{\Phi}'\tilde{\Phi}'}^{\tau\tau'} W_{\tilde{\Phi}'\tilde{\Phi}'}^{\tau\tau'}(q_{\text{eff}}) + \tilde{R}_{\Delta\Delta}^{\tau\tau'} W_{\Delta\Delta}^{\tau\tau'}(q_{\text{eff}}) \right] \\ & - \frac{2q_{\text{eff}}}{m_N} \left[ \tilde{R}_{\Phi''M}^{\tau\tau'} W_{\Phi''M}^{\tau\tau'}(q_{\text{eff}}) + \tilde{R}_{\Delta\Sigma'}^{\tau\tau'} W_{\Delta\Sigma'}^{\tau\tau'}(q_{\text{eff}}) \right] \end{aligned} \right\}$$

$$\tilde{R}_{MM}^{\tau\tau'} = \tilde{c}_1^\tau \tilde{c}_1^{\tau'*} + \tilde{c}_{11}^\tau \tilde{c}_{11}^{\tau'*},$$

$$\tilde{R}_{\Sigma''\Sigma''}^{\tau\tau'} = (\tilde{c}_4^\tau - \tilde{c}_6^\tau)(\tilde{c}_4^{\tau'*} - \tilde{c}_6^{\tau'*}) + \tilde{c}_{10}^\tau \tilde{c}_{10}^{\tau'*}.$$

$$\tilde{R}_{\Sigma'\Sigma'}^{\tau\tau'} = \tilde{c}_4^\tau \tilde{c}_4^{\tau'*} + \tilde{c}_9^\tau \tilde{c}_9^{\tau'*}$$

$$c_i \equiv \tilde{c}_i/v^2 = \sqrt{2}G_F \tilde{c}_i,$$

- particle physics is in products of Wilson coefficients  $R_i$ 
  - for instance the terms not vanishing in  $v_N \rightarrow 0$  limit