

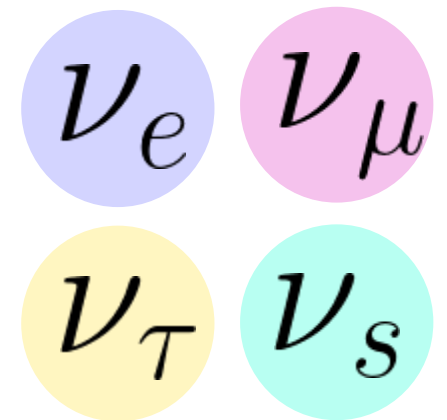
Boosting the production of sterile neutrino dark matter with self-interactions

Maria Dias

In collaboration with:
Stefan Vogl
Recontres de Moriond 2025

Sterile neutrinos as dark matter

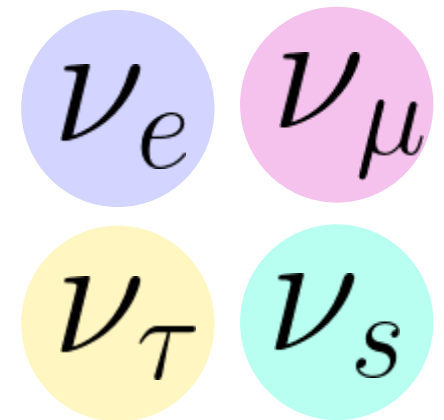
What are sterile neutrinos?



Sterile neutrinos as dark matter

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- Singlets under the SM gauge group

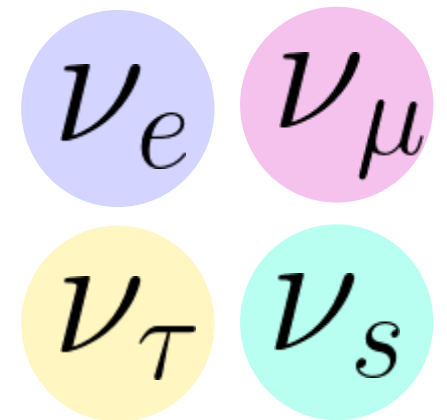


Sterile neutrinos as dark matter

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- Only interact with the SM through mass mixing

$$|\nu_4\rangle = \cos(\theta)|\nu_s\rangle + \sin(\theta)|\nu_a\rangle$$



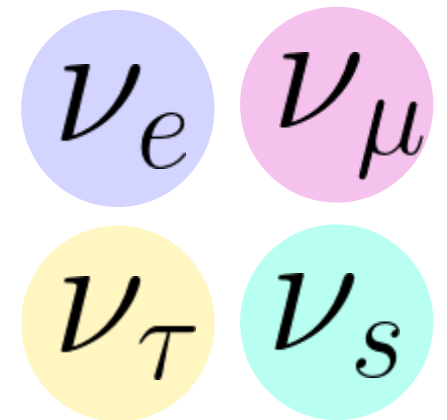
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- Well motivated BSM candidates



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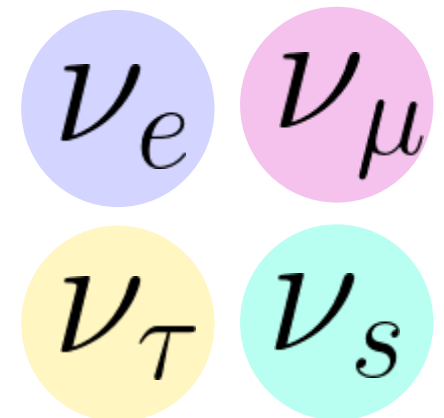
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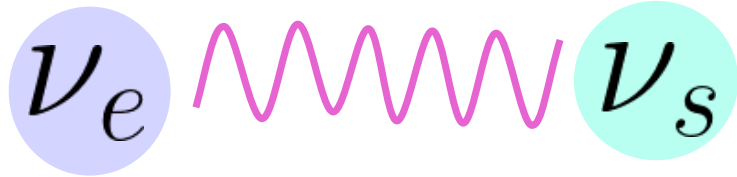
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- Well motivated BSM candidates

↳ keV sterile neutrinos
are good DM
candidates



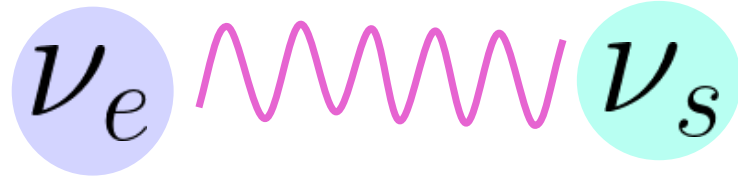
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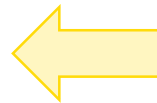
The Dodelson-Widrow (DW) scenario

- SN are produced in the early universe through oscillations

Sterile neutrinos as dark matter

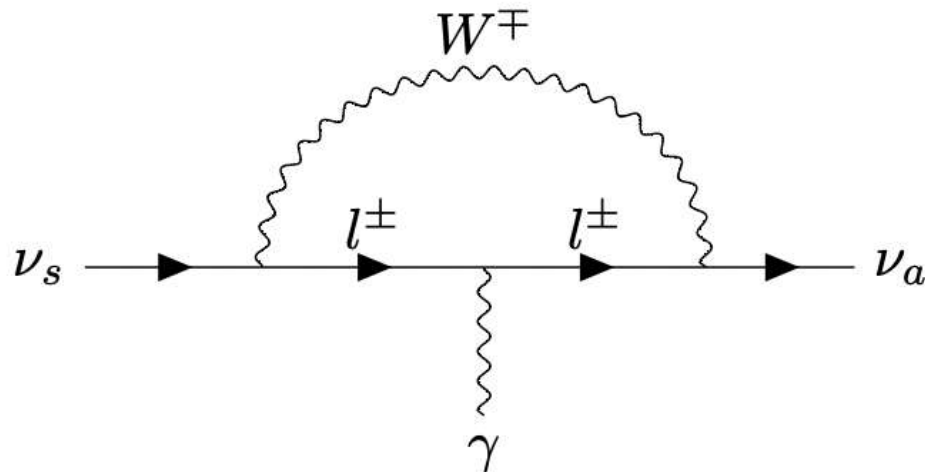


Can be searched for by current and future X-ray telescopes

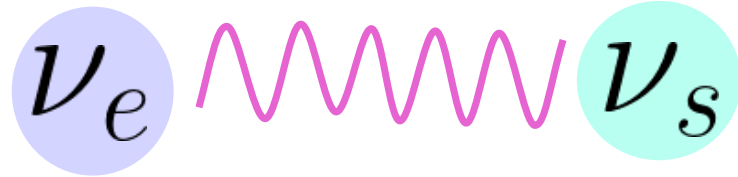


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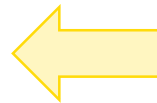
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Sterile neutrinos as dark matter



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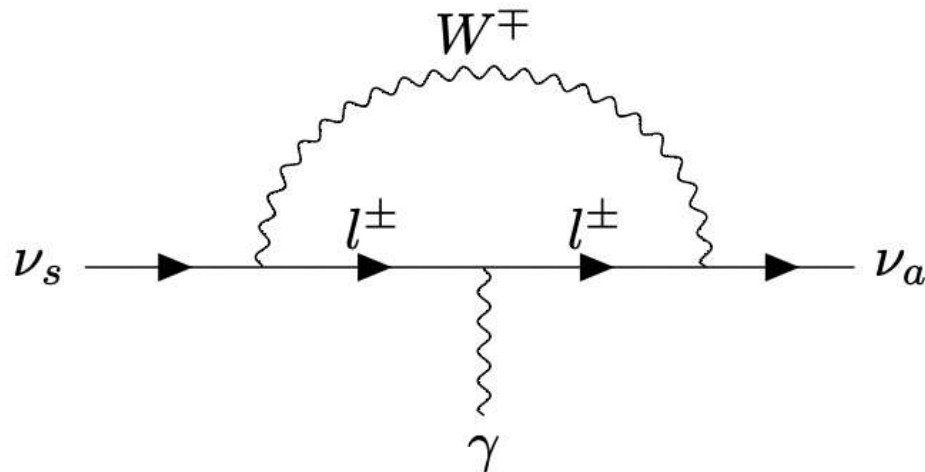


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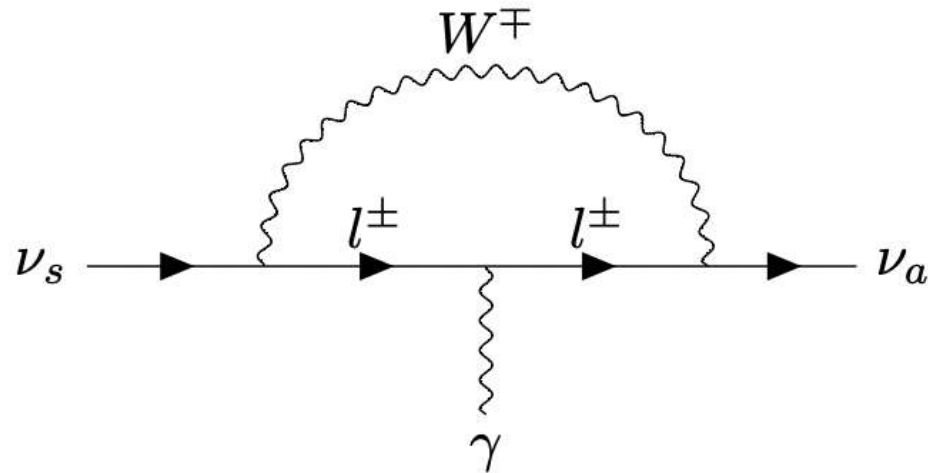
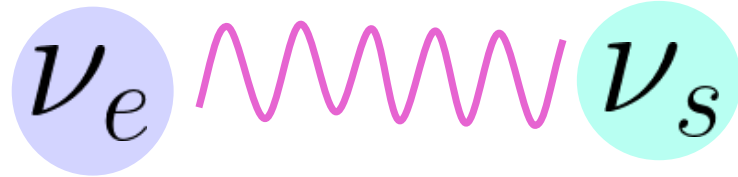
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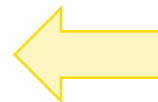
DW is mostly excluded



Sterile neutrinos as dark matter



$$\mathcal{L}_{\text{int}} = y \bar{\nu}_s \nu_s \phi.$$



The Dodelson-Widrow (DW) scenario

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DW is mostly excluded

How to extend DW?



Self-interactions

Self-interacting sterile neutrinos



The new interaction modifies the DW scenario in two key ways:

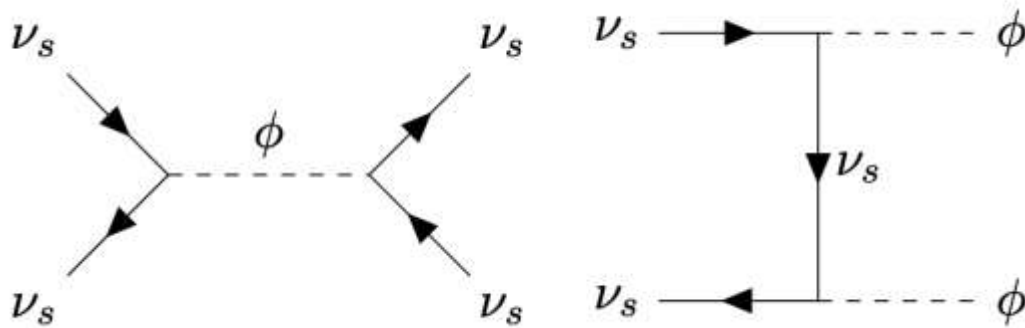
Self-interacting sterile neutrinos



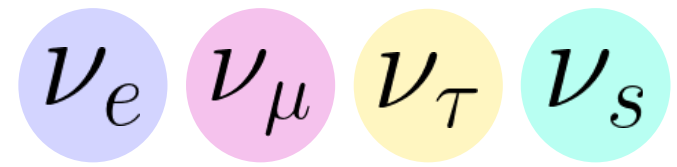
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1. New scattering rate

$$\Gamma_t = \Gamma_a + \Gamma_s$$



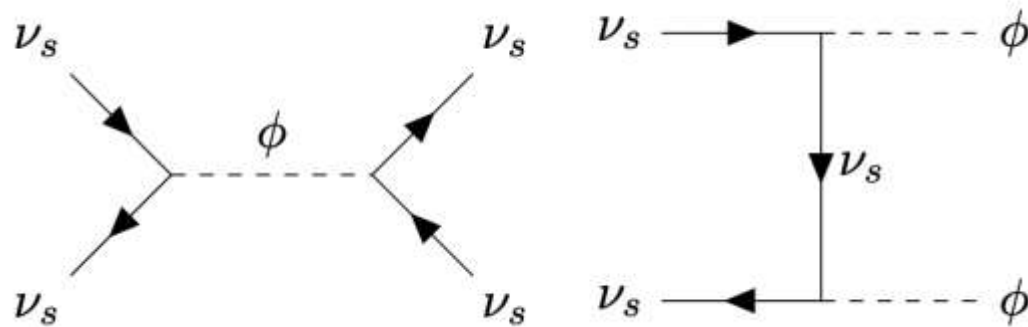
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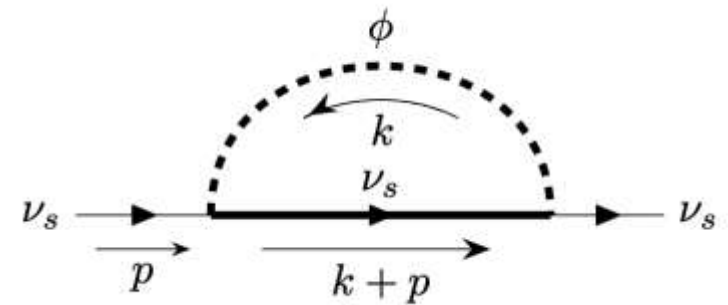
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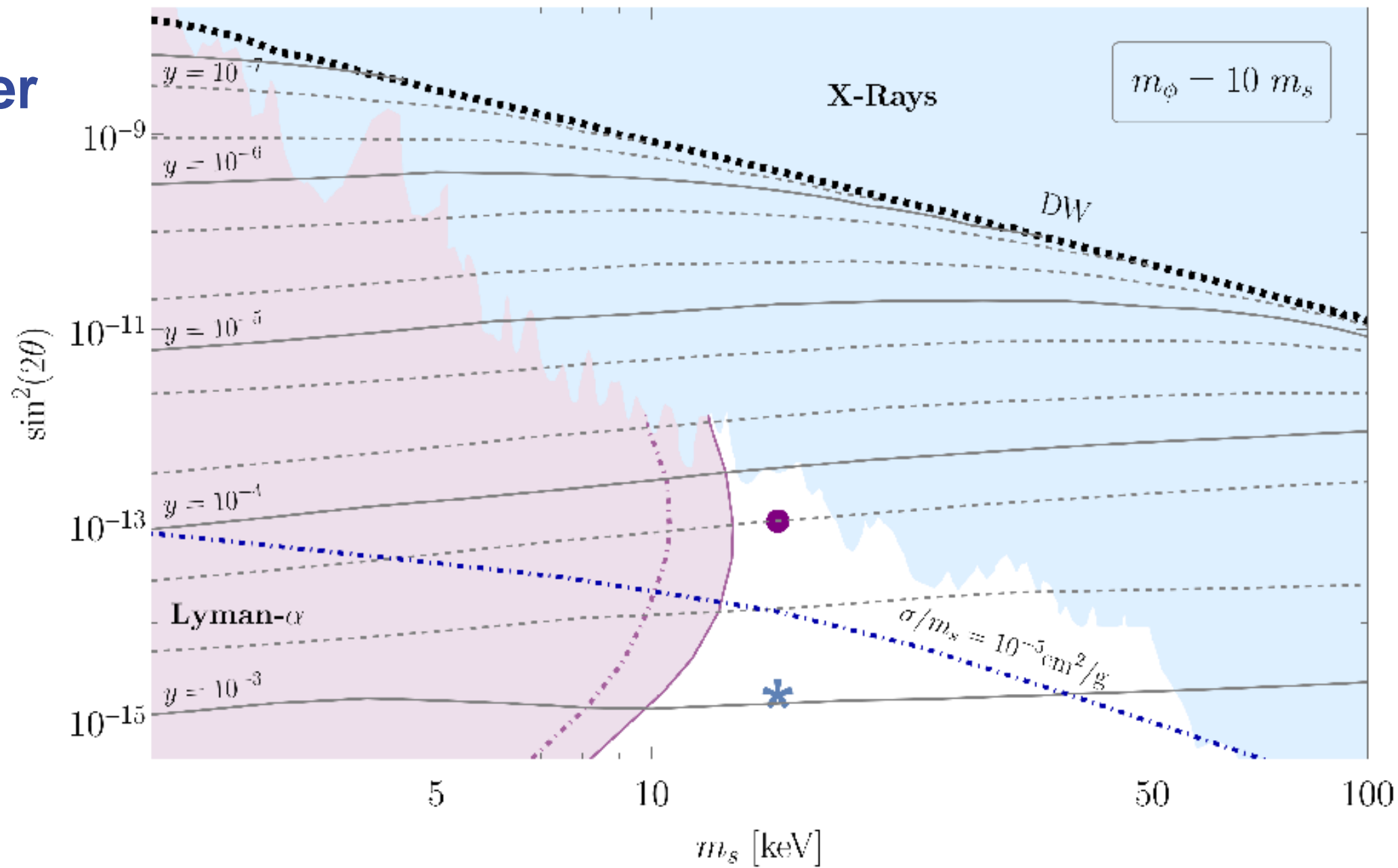


2. Correction to the neutrino's self energy

$$V_{\text{eff}} = V_a - V_s$$

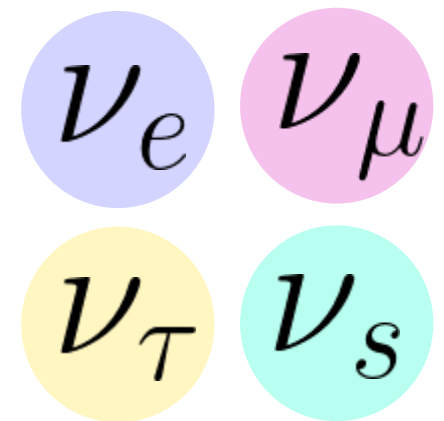


The parameter space



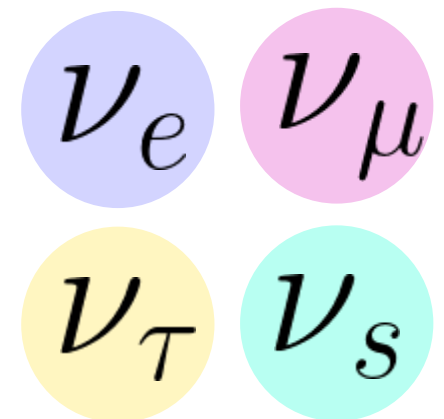
Conclusions

- Sterile neutrinos are attractive BSM candidates



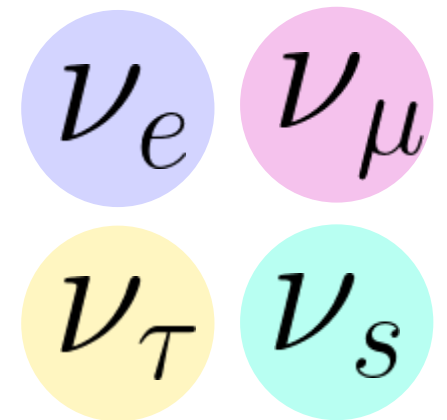
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Conclusions

- Sterile neutrinos are attractive BSM candidates
- The canonical DW mechanism is mostly excluded
- Self-interactions among the sterile neutrinos can open new portions of parameter space



Thank you for your attention

