

# Super-Kamiokande experiment

Takuya Tashiro (ICRR, **University of Tokyo**)  
on behalf of Super-Kamiokande collaboration

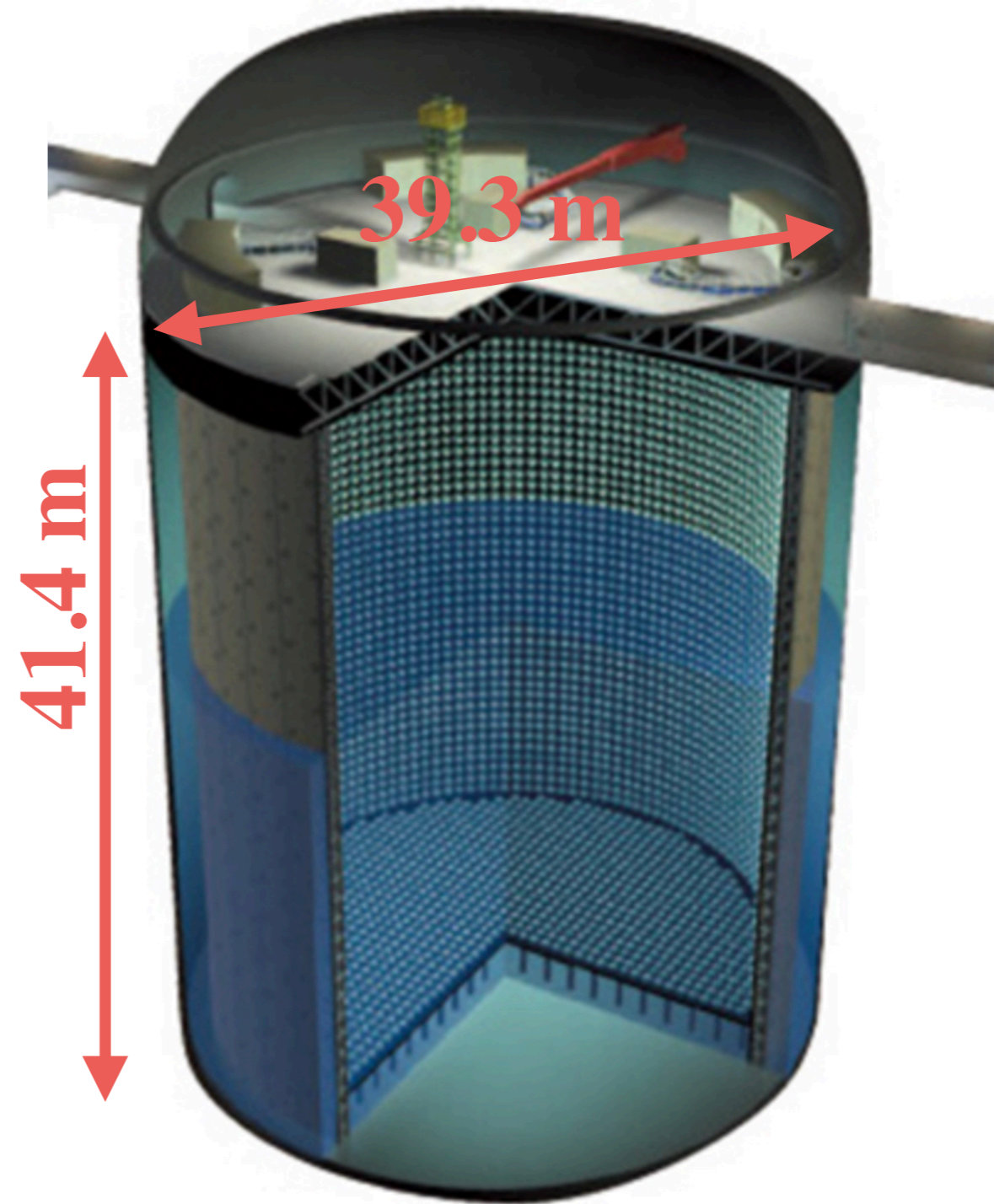
March 2025  
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# Super-Kamiokande(SK) experiment

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- ▶ SK is the world's largest underground water Cherenkov detector with a cylindrical tank of 41.4 m height and 39.3m diameter.
- ▶ Since the start of operation in 1996, SK has been taking data to explore the neutrino properties, neutrino sources, and nucleon decay.
- ▶ Discovery of neutrino oscillation by SK lead to the nobel prize in 2015.  
**Now SK aims at more precise understanding of the neutrino oscillation.**



# SK detector concept

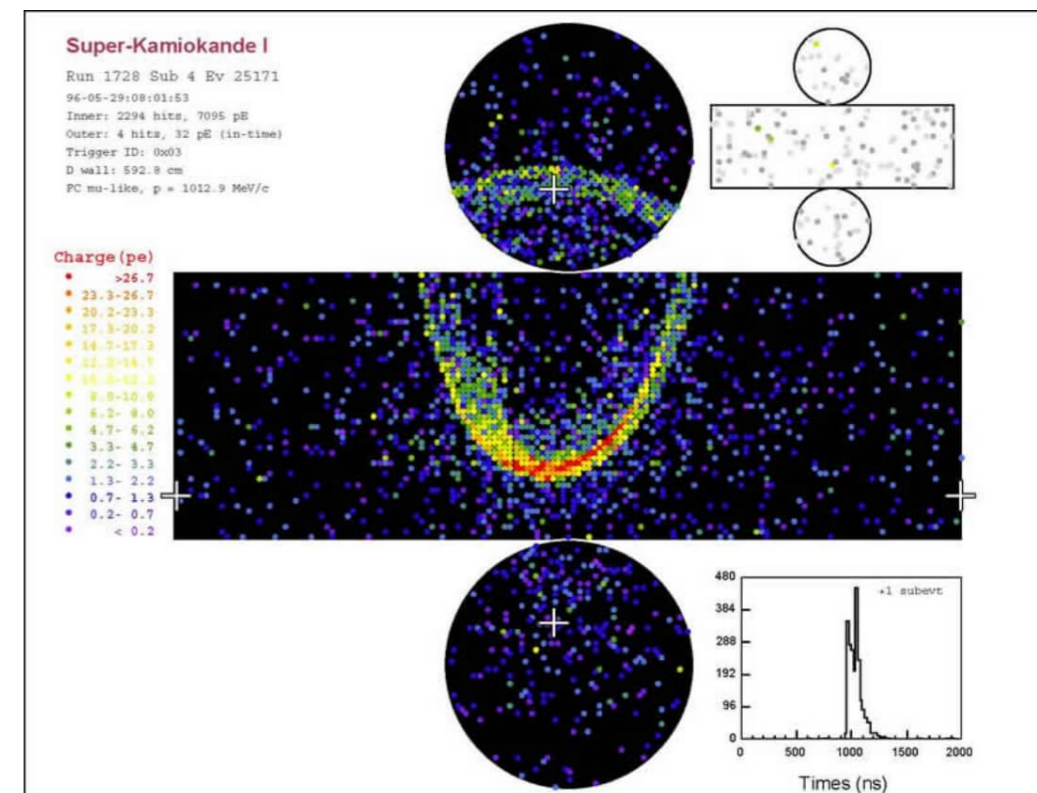
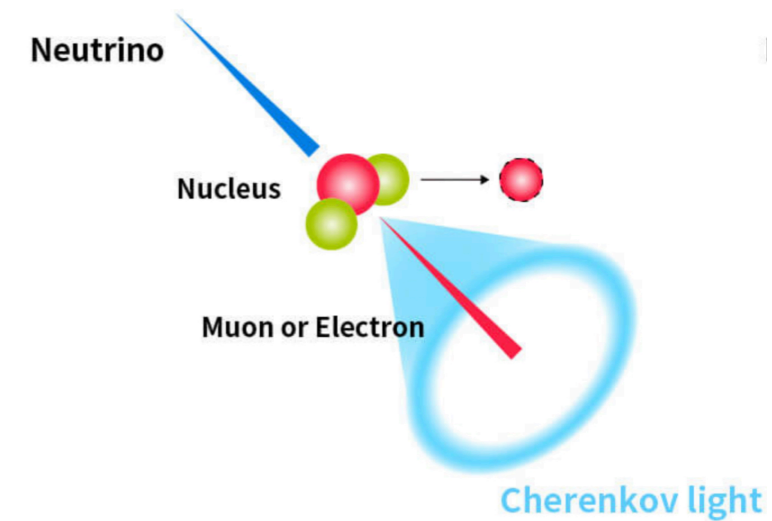
▶ The SK tank is filled with 50 kton of ultra-pure water (1996~2020) or gadolinium-loaded water (2020~)

▶ In water, a ring of Cherenkov light is emitted when a charged particle travels faster than the speed of light in water.

▶ The Cherenkov light is detected by ~11,000 PMTs attached to the tank wall and reconstructed as a ring.

▶ The particle type, momentum, direction, vertex position e.t.c. can be extracted from the amount of light, shape and direction of the ring, and photon detection time of each PMT.

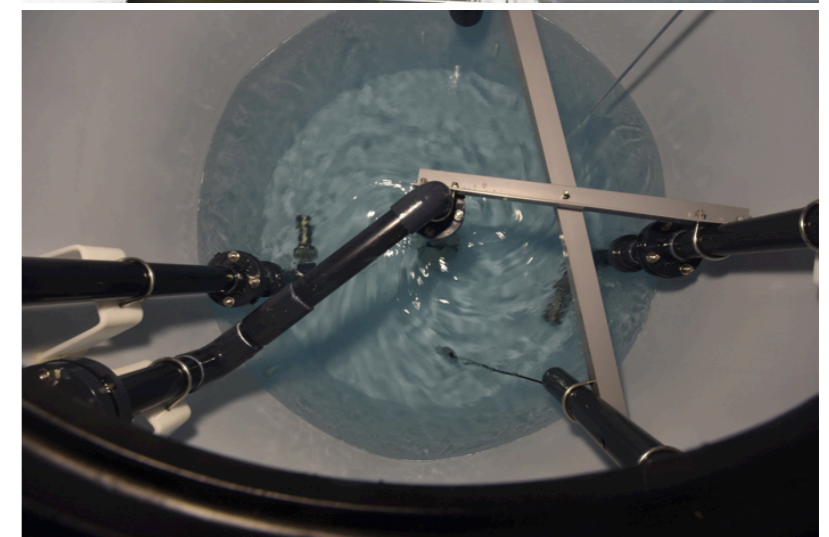
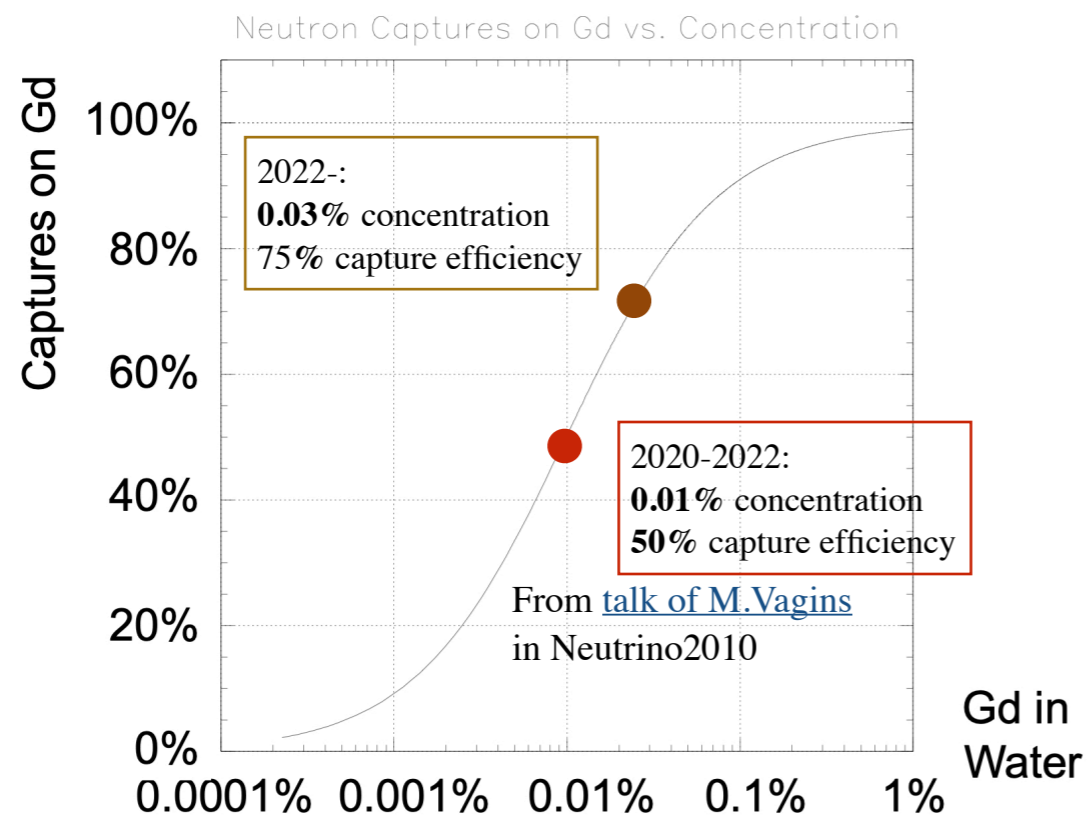
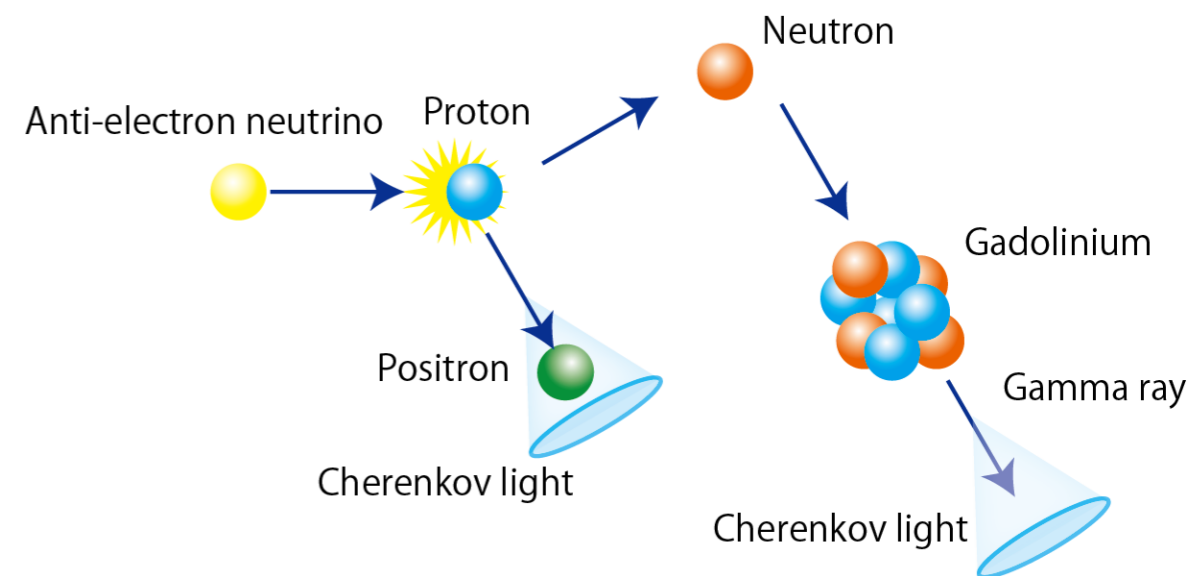
▶ Details can be found in [https://doi.org/10.1142/9789819801107\\_0008](https://doi.org/10.1142/9789819801107_0008)





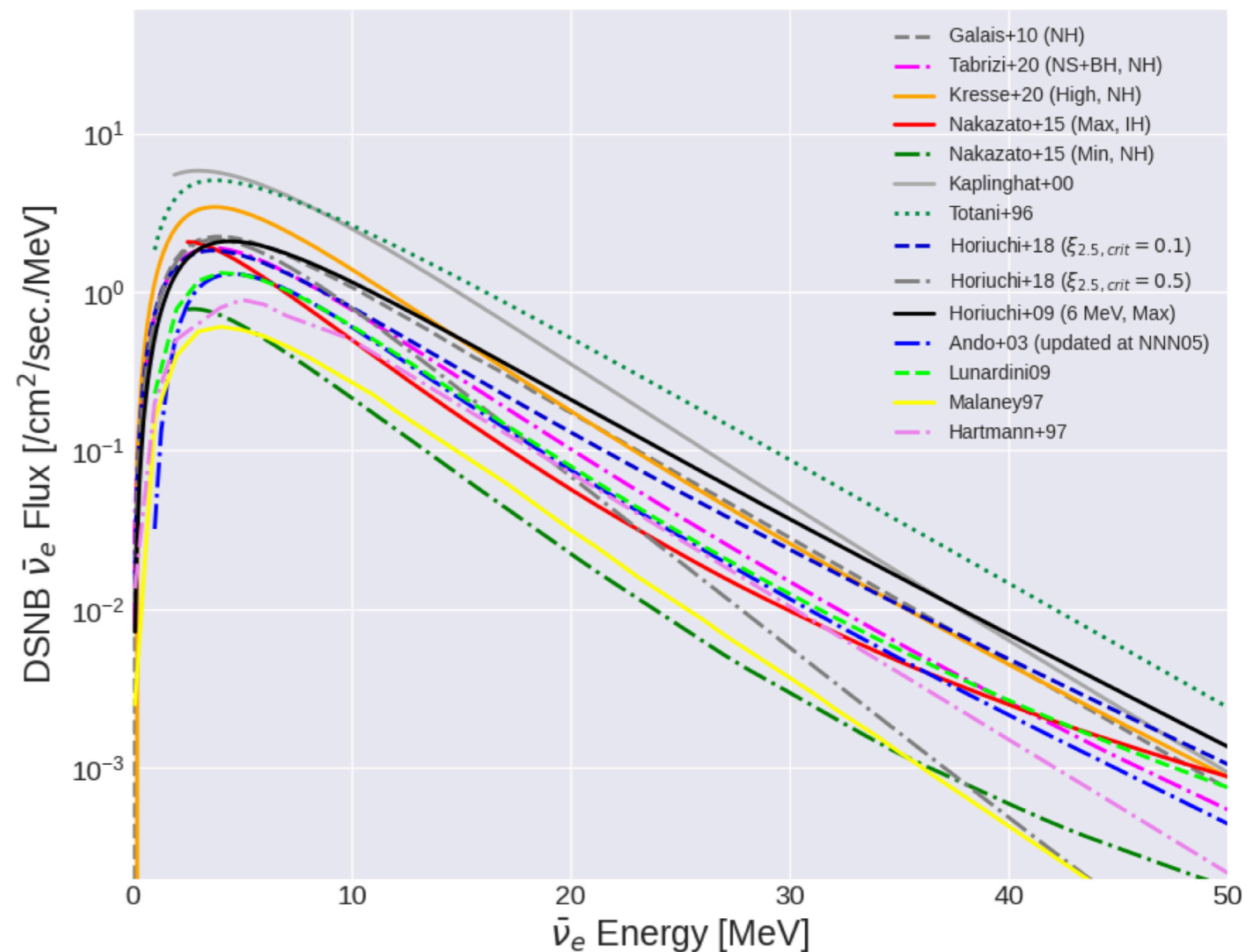
# Recent news: Gd-loading

- ▶ Since 2020, SK is operated with Gd-loaded water for neutron-tagging.
  - ▶ Neutron is captured by Gd and emit gamma-rays with the total energy of  $\sim 8$  MeV, which make EM showers emitting Cherenkov light.
- ▶ Gd loading campaigns were held in 2020 and 2022 to make the Gd concentration 0.01% and 0.03%, respectively.



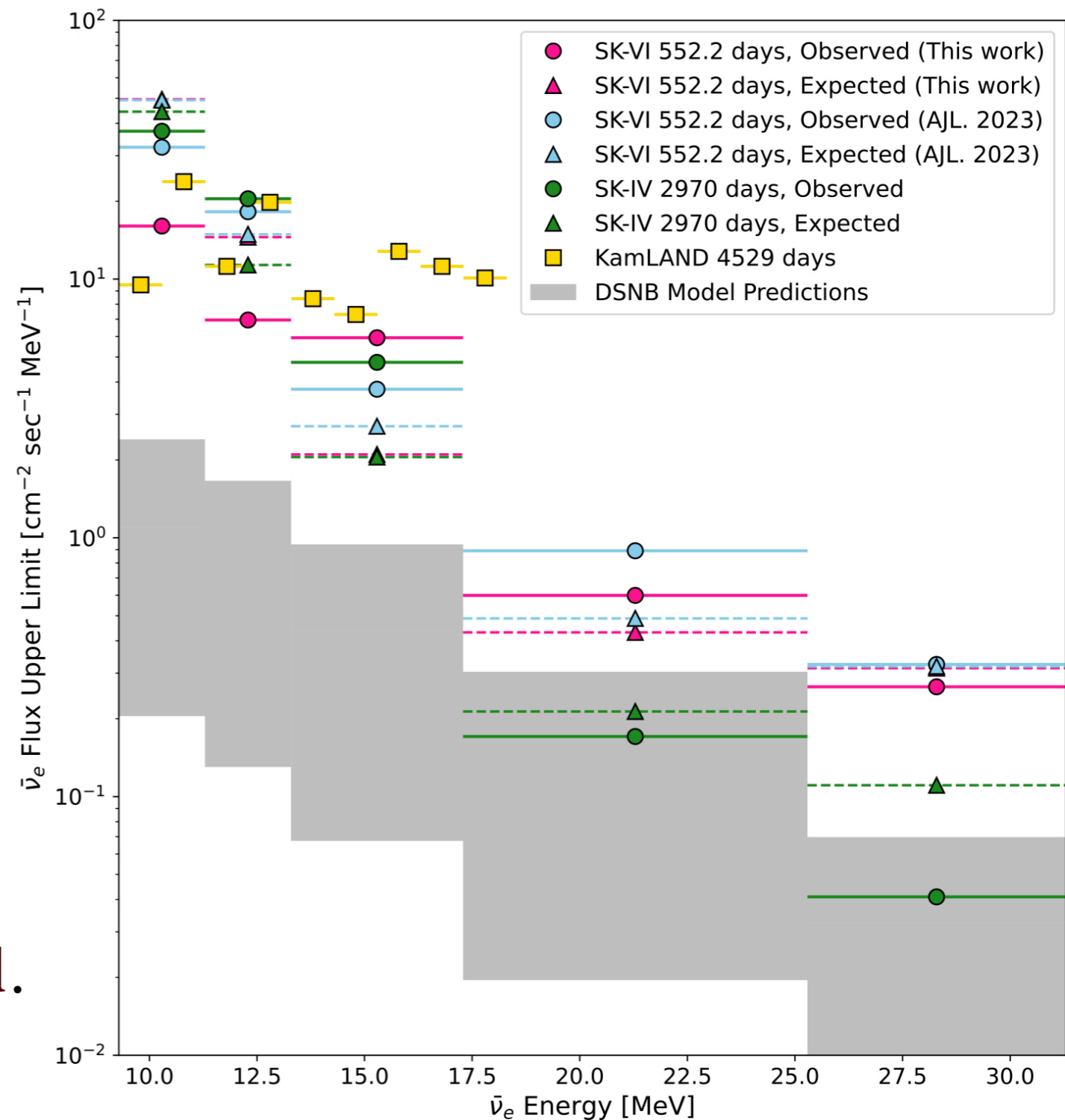


- ▶ By the analysis of SN1987a, Kamiokande revealed that neutrinos are emitted by supernovae. (Nobel prize in 2002)
- ▶ This indicates that the Universe is filled by the neutrinos emitted from past supernovae, known as diffuse supernova neutrino background(DSNB).
- ▶ The evidence of DSNB is not observed yet. Discovery one of the goals in SK.
- ▶ In the DSNB energy range ( $O(1)\sim O(10)$  MeV), the  $p + \bar{\nu}_e \rightarrow n + e^+$  process has the largest cross section Neutron tagging has important role and the sensitivity can be enhanced by Gd.





- ▶ Now the search is performed on both pure water phase of 2970 days and Gd phase of 552 days (only the 0.01% concentration period).
- ▶ The evidence of the DSNB is not observed yet, but **the sensitivity is reaching to the region predicted by theories.**
- ▶ **Further sensitivity is expected by combining more data with more Gd concentration period.** Look forward to the future results.





# Neutrino oscillation

- ▶ Neutrino oscillation probability can be determined by the MNS parameters and mass squared differences.

## MNS matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & e^{-i\delta_{CP}} \sin \theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{CP}} \sin \theta_{13} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

- ▶  $\theta_{12}, \theta_{23}, \theta_{13}$  are already experimentally determined.

$$\theta_{12} = 33.62^{+0.78}_{-0.76} \quad \theta_{23} = 47.2^{+1.9}_{-3.9} \quad \theta_{13} = 8.54^{+0.15}_{-0.15}$$

- ▶  $\delta_{CP}=0$  is not experimentally rejected by  $5\sigma$ . (T2K rejected by  $3\sigma$ ).

- ▶ Mass squared differences are also determined.

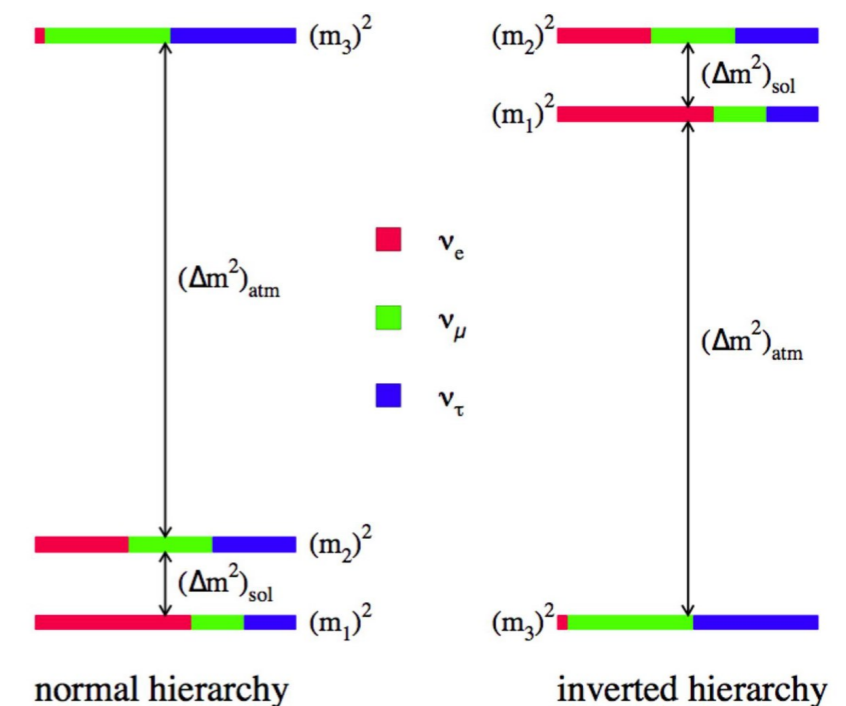
$$\Delta m_{12}^2 = 0.759 \times 10^{-4} eV^2$$

$$\Delta m_{13}^2 \sim \Delta m_{32}^2 = 23.2 \times 10^{-4} eV^2$$

- ▶ **Mass ordering is not known yet.**

$$m_3 \gg m_1, m_2 \text{ (Normal ordering)}$$

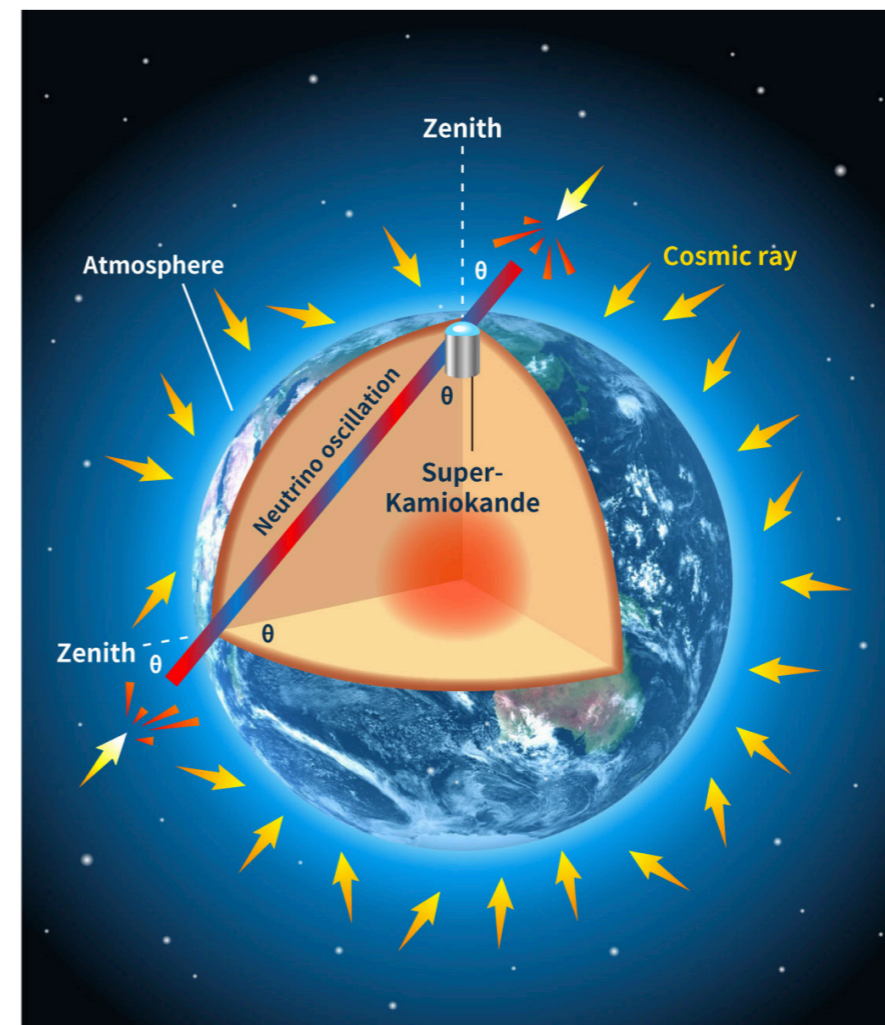
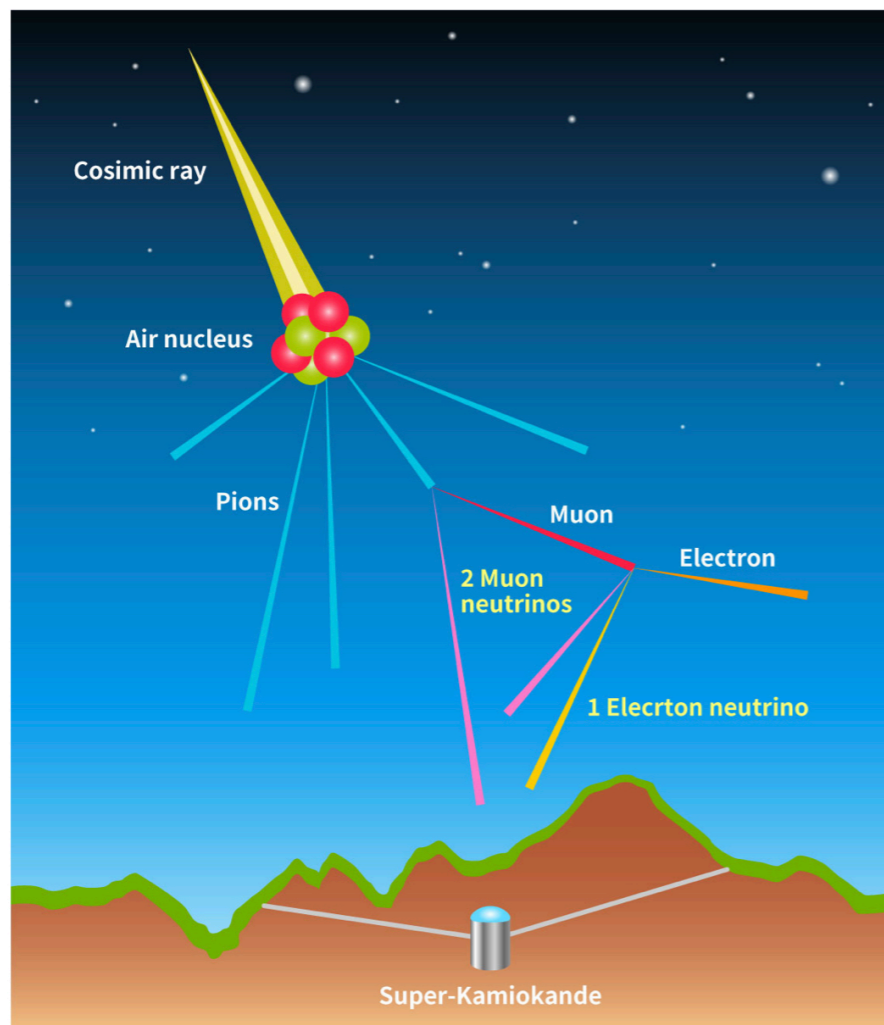
$$m_1, m_2 \gg m_3 \text{ (Inverted ordering)}$$





# Atmospheric neutrino

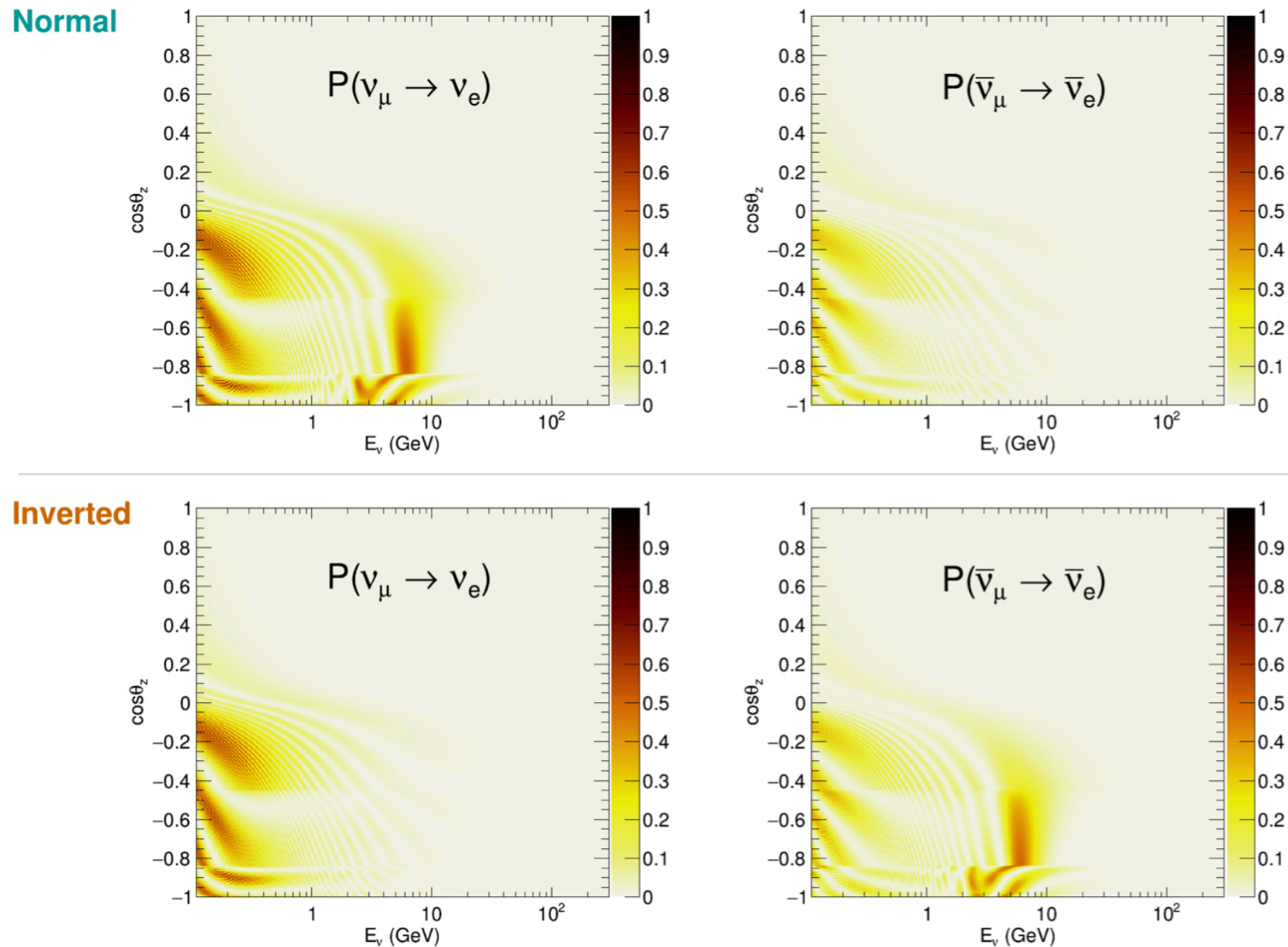
- ▶ **Atmospheric neutrino is a powerful tool to explore the neutrino oscillation.**
- ▶ By the interaction between primary cosmic ray and the atmosphere,  $\nu_\mu(\bar{\nu}_\mu)$  and  $\nu_e(\bar{\nu}_e)$  are produced with the amount ratio of 2:1.





# Latest result of atmospheric neutrino

- ▶ As a result of neutrino oscillation, the flavor ratio of the atmospheric neutrino depends on the zenith angle and energy.
- ▶ The oscillation pattern also depends on the mass ordering.



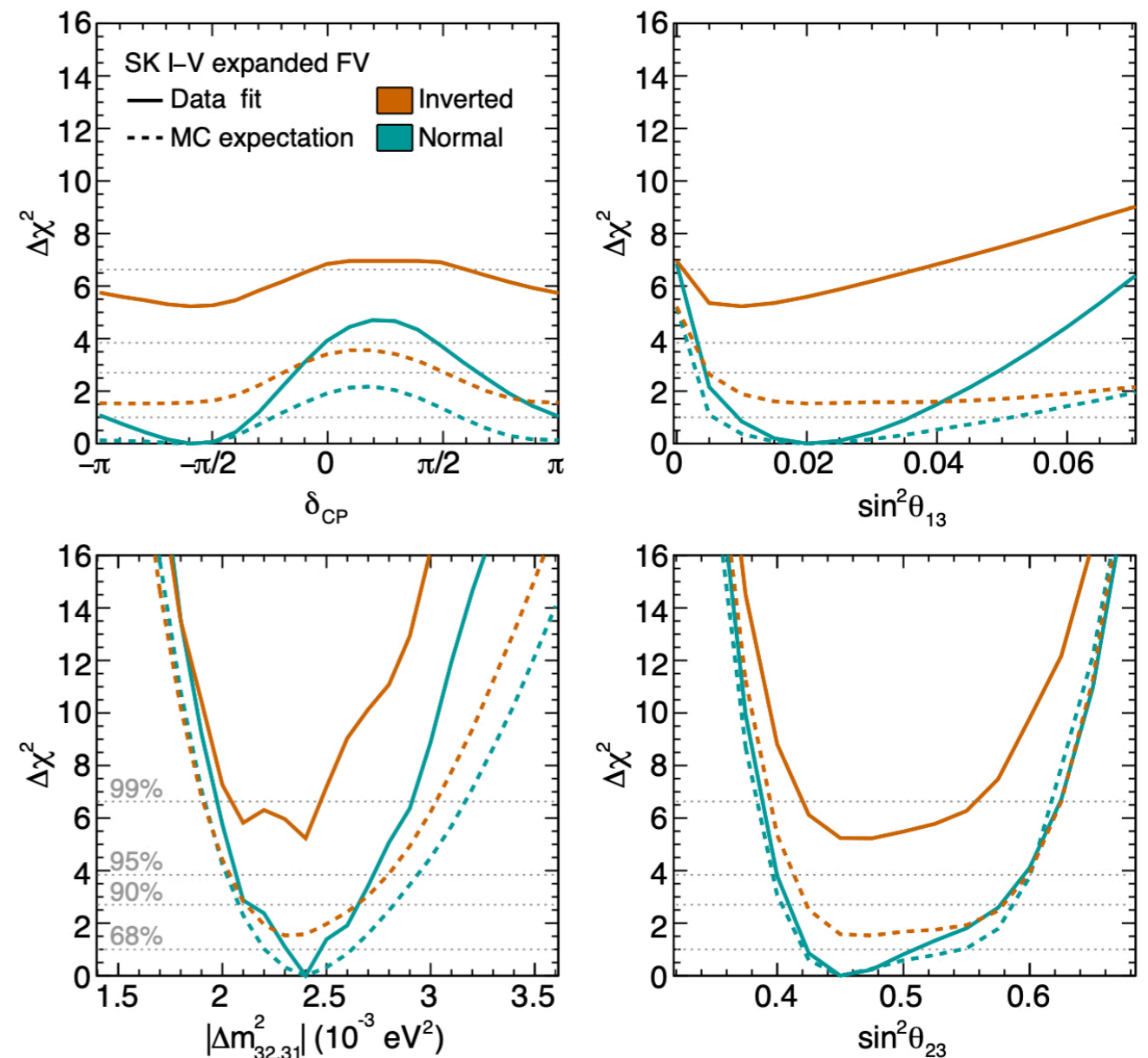


# Latest result of atmospheric neutrino

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► The latest analysis was performed with the data of complete pure water phase, between 1996 and 2020, corresponding to the exposure of 484.2 kton-year (more than 10 times larger than the exposure of the neutrino oscillation discovery analysis!).

► Detailed analysis of the  $\nu_e, \bar{\nu}_e, \nu_\mu(\bar{\nu}_\mu)$  components depending on the zenith angle and energy set strong constraints on the oscillation parameters.



From [Phys. Rev. D 109, 072014](#)



# Latest result of atmospheric neutrino

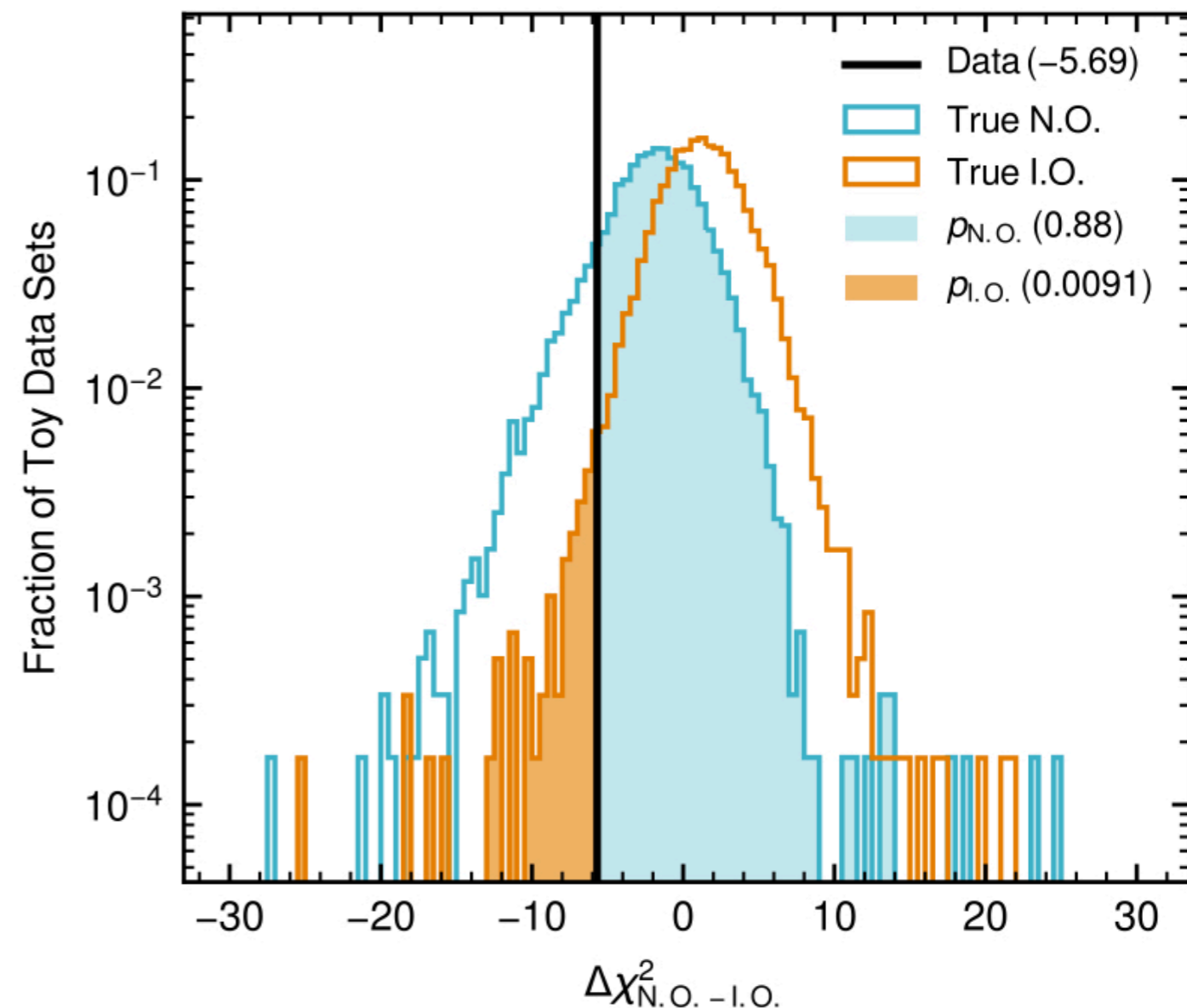
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► Detailed analysis of the  $\nu_e, \bar{\nu}_e, \nu_\mu(\bar{\nu}_\mu)$  components depending on the zenith angle and energy set strong constraints on the oscillation parameters.

► **It prefers the normal ordering at 92.3 % confidence level**

From [Phys. Rev. D 109, 072014](#)



# Solar neutrino

- ▶ As well as the atmospheric neutrinos, neutrinos coming from the Sun (solar neutrino) is an important physics target in SK.
- ▶ It is more challenging to detect the solar neutrinos in SK than the atmospheric neutrinos because the energy range is much lower.
  - ▶ Solar neutrino is  $O(1)$  MeV while the atmospheric neutrino is typically  $O(10)$  MeV  $\sim$  GeV.
  - ▶ With the improved electronics, SK can record the events with the threshold of  $\sim 3.5$  MeV in electron kinetic energy since 2008.

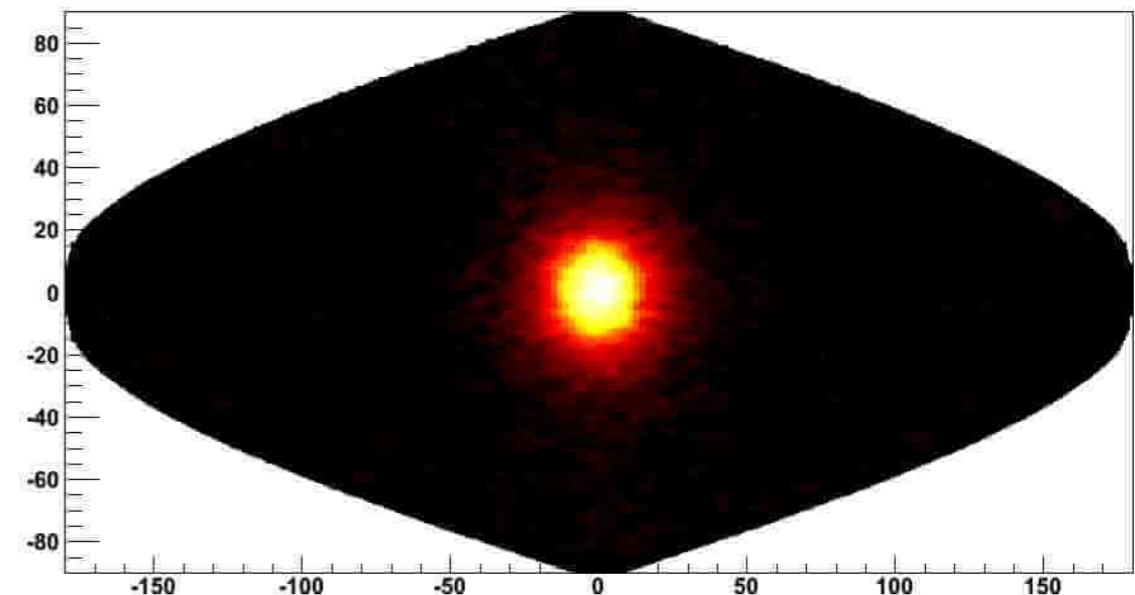
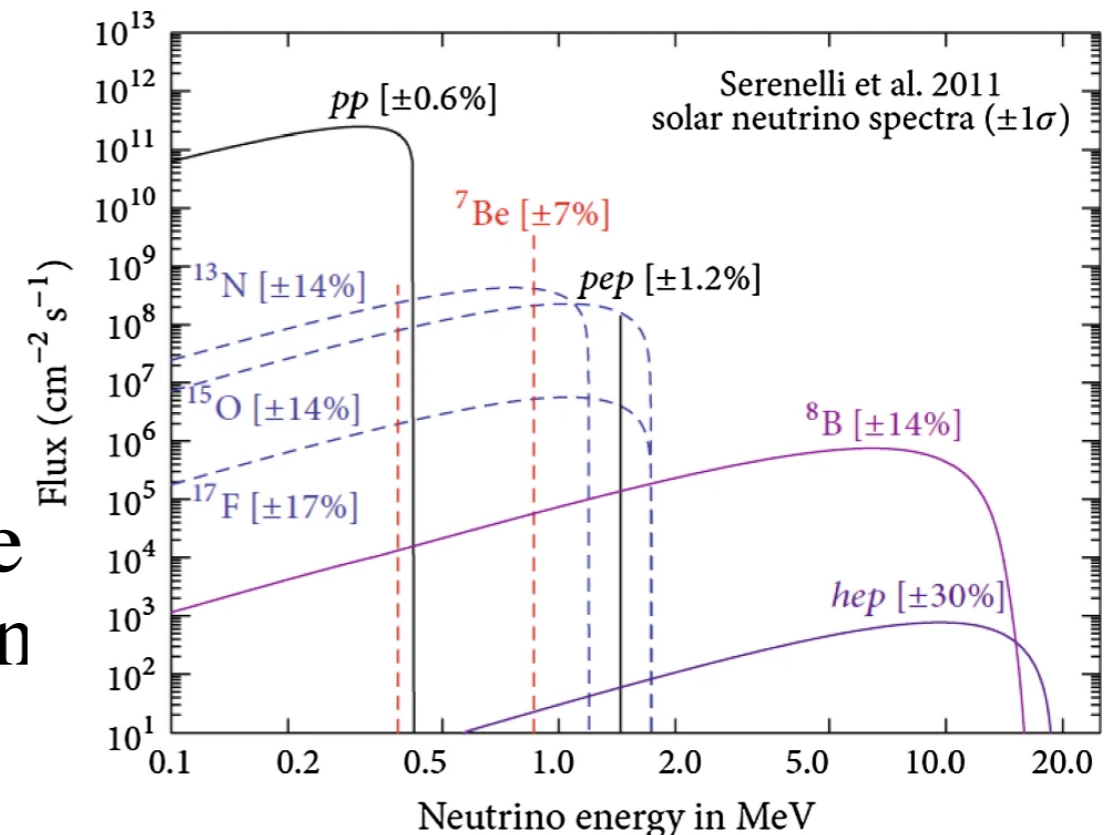


Image of the sun seen by neutrinos.  
From the [website of ICRR](#)



From [Eur. Phys. J. C 79, 298 \(2019\)](#)



# Solar neutrino oscillation

- ▶ In the solar neutrino energy range, only the  $\nu_e$  can interact with the charged current. So **the survival probability of  $\nu_e$  can be extracted from the recorded events.**
- ▶ As the  $\theta_{13}$  and  $\theta_{13}$  have small impact in the  $\nu_e$  survival probability,  $\theta_{12}$  and  $\Delta m_{21}^2$  can be constrained by the solar neutrino analysis.

- ▶ By the analysis of 5805 days data, SK set the constraint of

$$\sin^2 \theta_{12} = 0.306 \pm 0.013$$

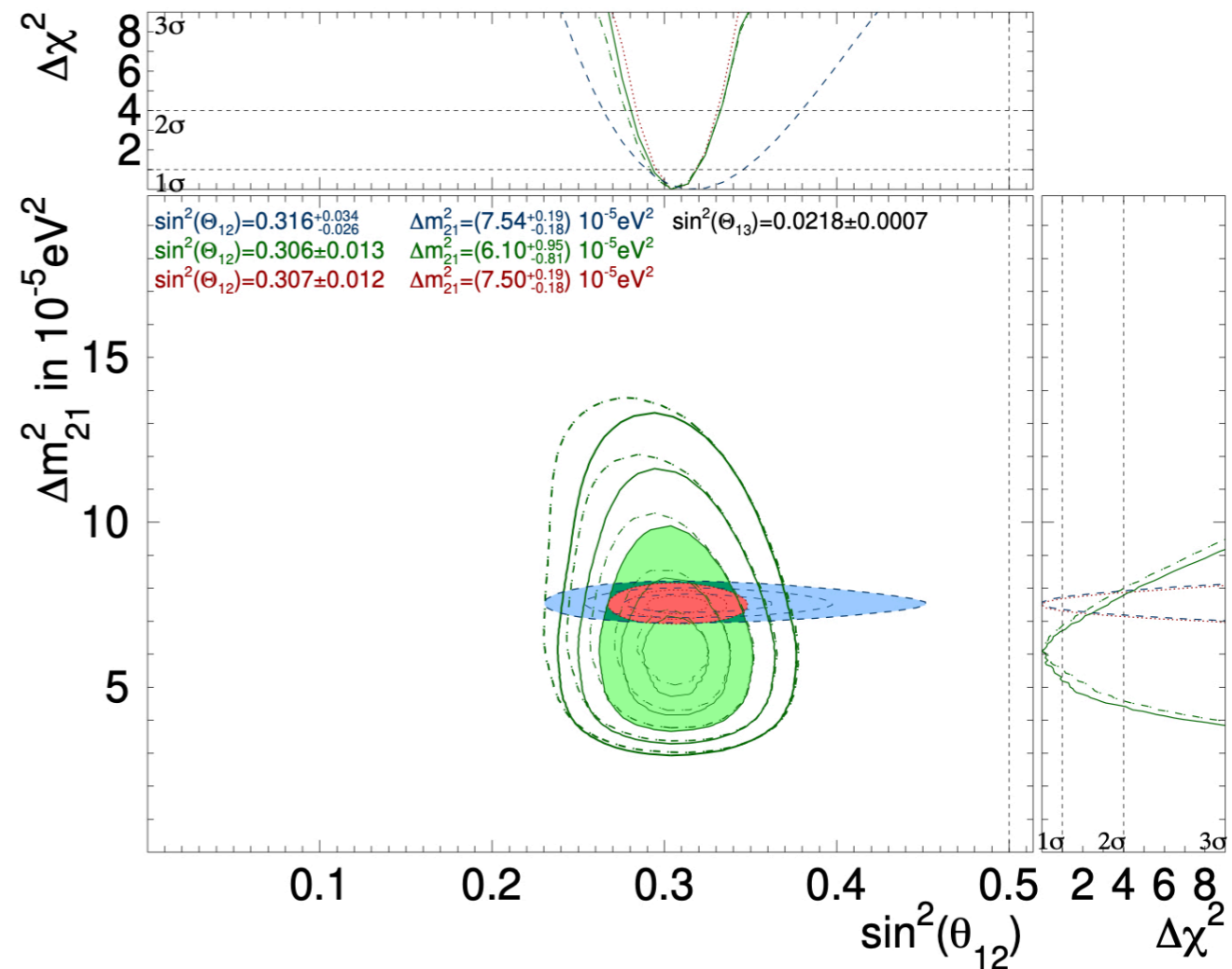
$$\Delta m_{21}^2 = (6.10^{+0.95}_{-0.81}) \times 10^{-5} \text{ eV}^2$$

- ▶ Further constraint is achieved by combination with KAMLAND,

$$\sin^2 \theta_{12} = 0.307 \pm 0.012$$

$$\Delta m_{21}^2 = (7.50^{+0.19}_{-0.18}) \times 10^{-5} \text{ eV}^2$$

- ▶  $1.5\sigma$  tension was observed in  $\Delta m_{21}^2$  between SK and KAMLAND.

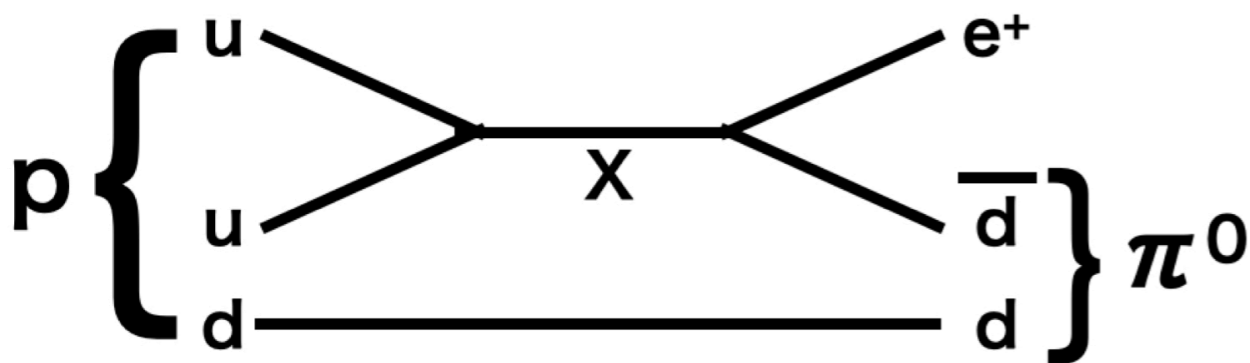


From [Phys. Rev. D 109, 092001 \(2024\)](#)

# Proton decay

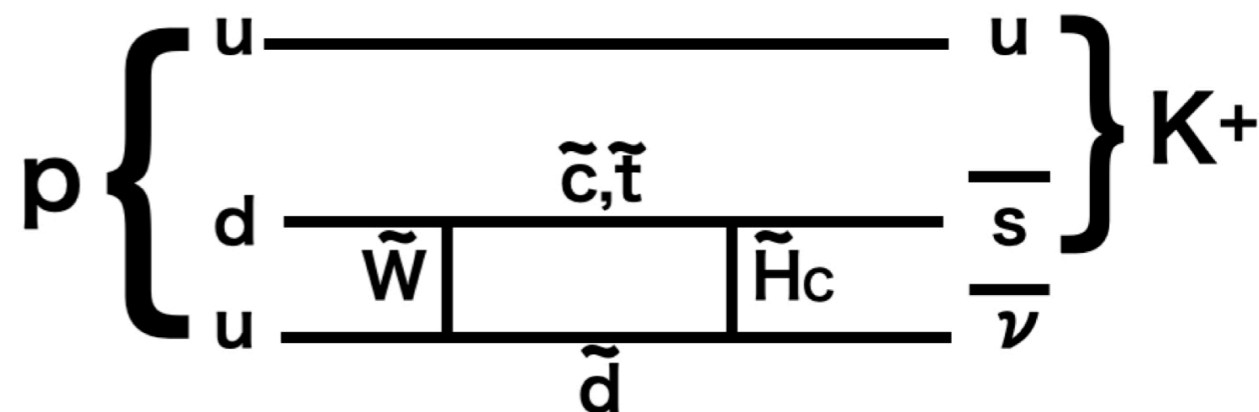
- ▶ Proton decay (or neutron decay) is also one of the important physics target in SK.
- ▶ Proton decay is prohibited in SM, but it is predicted by GUT models. **Observation of proton decay can be a strong evidence of GUT.**
- ▶ Two decay modes are intensively searched because they are predicted in leading models, and other possible modes are also widely searched.

$$p \rightarrow e^+ \pi^0$$



- ▶ Dominant in non-SUSY models
- ▶ Current limit:  
 $\tau/\Gamma > 2.4 \times 10^{34}$  years.

$$p \rightarrow \bar{\nu} K^+$$



- ▶ Dominant in SUSY models
- ▶ Current limit:  
 $\tau/\Gamma > 8.2 \times 10^{33}$  years.



# Proton decay

- ▶ Other possible decay modes including neutron decay and dinucleon decay are also searched in SK.
- ▶ In most of the modes, SK sets the best limit in the world.

Decay mode	SK detector	Exposure [kt-years]	Lifetime limit [years]	Reference	Comments
$p \rightarrow e^+\pi^0$	I-IV	450	$2.4 \times 10^{34}$	[20]	Expanded FV
$p \rightarrow \mu^+\pi^0$	I-IV	450	$1.6 \times 10^{34}$	[20]	Expanded FV
$p \rightarrow \nu\pi^+$	I-III	173	$3.9 \times 10^{32}$	[21]	
$n \rightarrow \nu\pi^0$	I-III	173	$1.1 \times 10^{33}$	[21]	
$p \rightarrow e^+\eta$	I-IV	373	$1.4 \times 10^{34}$	In prep.	
$p \rightarrow \mu^+\eta$	I-IV	373	$7.3 \times 10^{33}$	In prep.	
$p \rightarrow e^+\rho^0$	I-IV	316	$7.2 \times 10^{32}$	[22]	A part of SK IV data
$p \rightarrow \mu^+\rho^0$	I-IV	316	$5.7 \times 10^{32}$	[22]	A part of SK IV data
$p \rightarrow e^+\omega$	I-IV	316	$1.6 \times 10^{33}$	[22]	A part of SK IV data
$p \rightarrow \mu^+\omega$	I-IV	316	$2.8 \times 10^{33}$	[22]	A part of SK IV data
$n \rightarrow e^+\pi^-$	I-IV	316	$5.3 \times 10^{33}$	[22]	A part of SK IV data
$n \rightarrow \mu^+\pi^-$	I-IV	316	$3.5 \times 10^{33}$	[22]	A part of SK IV data
$n \rightarrow e^+\rho^-$	I-IV	316	$3.0 \times 10^{31}$	[22]	A part of SK IV data
$n \rightarrow \mu^+\rho^-$	I-IV	316	$6.0 \times 10^{31}$	[22]	A part of SK IV data

$p \rightarrow e^+\pi^0\pi^0$	I-V	401	$7.2 \times 10^{33}$	In prep.	
$p \rightarrow \mu^+\pi^0\pi^0$	I-V	401	$4.5 \times 10^{33}$	In prep.	
$p \rightarrow e^+e^+e^-$	I-IV	373	$3.4 \times 10^{34}$	[23]	
$p \rightarrow \mu^+e^+e^-$	I-IV	373	$2.3 \times 10^{34}$	[23]	
$p \rightarrow \mu^-e^+e^+$	I-IV	373	$1.9 \times 10^{34}$	[23]	
$p \rightarrow e^+\mu^+\mu^-$	I-IV	373	$9.2 \times 10^{33}$	[23]	
$p \rightarrow e^-\mu^+\mu^+$	I-IV	373	$1.1 \times 10^{34}$	[23]	
$p \rightarrow \mu^+\mu^+\mu^-$	I-IV	373	$1.0 \times 10^{34}$	[23]	
$p \rightarrow e^+\nu\nu$	I-IV	273	$1.7 \times 10^{32}$	[24]	A part of SK IV data
$p \rightarrow \mu^+\nu\nu$	I-IV	273	$2.2 \times 10^{32}$	[24]	A part of SK IV data
$p \rightarrow e^+\chi$	I-IV	273	$7.9 \times 10^{32}$	[25]	A part of SK IV data
$p \rightarrow \mu^+\chi$	I-IV	273	$4.1 \times 10^{32}$	[25]	A part of SK IV data
$n \rightarrow \nu\gamma$	I-IV	273	$5.5 \times 10^{32}$	[25]	A part of SK IV data
$p \rightarrow \nu K^+$	I-IV	365	$8.2 \times 10^{33}$	[26]	A part of SK IV data
$p \rightarrow \mu^+K^0$	I-IV	373	$3.6 \times 10^{33}$	[27]	
$p \rightarrow e^+K^0$	I	92	$1.3 \times 10^{33}$	[28]	
$p \rightarrow \mu^+K^0$	I	92	$1.0 \times 10^{33}$	[28]	

From the [presentation by S.Mine in NNN23](#)

$pp \rightarrow e^+e^+$	I-IV	373	$4.2 \times 10^{33}$	[29]	
$nn \rightarrow e^+e^-$	I-IV	373	$4.2 \times 10^{33}$	[29]	
$nn \rightarrow \gamma\gamma$	I-IV	373	$4.1 \times 10^{33}$	[29]	
$pp \rightarrow e^+\mu^+$	I-IV	373	$4.4 \times 10^{33}$	[29]	
$nn \rightarrow e^+\mu^-$	I-IV	373	$4.4 \times 10^{33}$	[29]	
$nn \rightarrow e^-\mu^+$	I-IV	373	$4.4 \times 10^{33}$	[29]	
$pp \rightarrow \mu^+\mu^+$	I-IV	373	$4.4 \times 10^{33}$	[29]	
$nn \rightarrow \mu^+\mu^-$	I-IV	373	$4.4 \times 10^{33}$	[29]	
$pp \rightarrow \pi^+\pi^+$	I-IV	282	$7.2 \times 10^{31}$	[30]	A part of SK IV data
$pn \rightarrow \pi^+\pi^0$	I-IV	282	$1.7 \times 10^{32}$	[30]	A part of SK IV data
$nn \rightarrow \pi^0\pi^0$	I-IV	282	$4.0 \times 10^{32}$	[30]	A part of SK IV data
$np \rightarrow e^+\nu$	I-IV	273	$2.6 \times 10^{32}$	[25]	A part of SK IV data
$np \rightarrow \mu^+\nu$	I-IV	273	$2.2 \times 10^{32}$	[25]	A part of SK IV data
$np \rightarrow \tau^+\nu$	I-IV	273	$2.9 \times 10^{31}$	[25]	A part of SK IV data
$pp \rightarrow K^+K^+$	I	92	$1.7 \times 10^{32}$	[31]	
$n \rightarrow \bar{n}$	I-IV	373	$3.6 \times 10^{32}$	[32]	

# Recent attempts

- ▶ SK tries to realize more sensitive analysis by many attempts.

- ▶ Enlarged fiducial volume

- ▶ In the convention analyses, the fiducial volume is defined as the volume  $>2\text{m}$  from the wall.
- ▶ Enlarged fiducial volume is defined as  $>1\text{m}$ , Corresponding to  **$\sim 20\%$  increase of the exposure.**
- ▶ The latest  $p \rightarrow e^+ \pi^0$  and  $p \rightarrow \mu^+ \pi^0$  analysis was performed using the enlarged volume

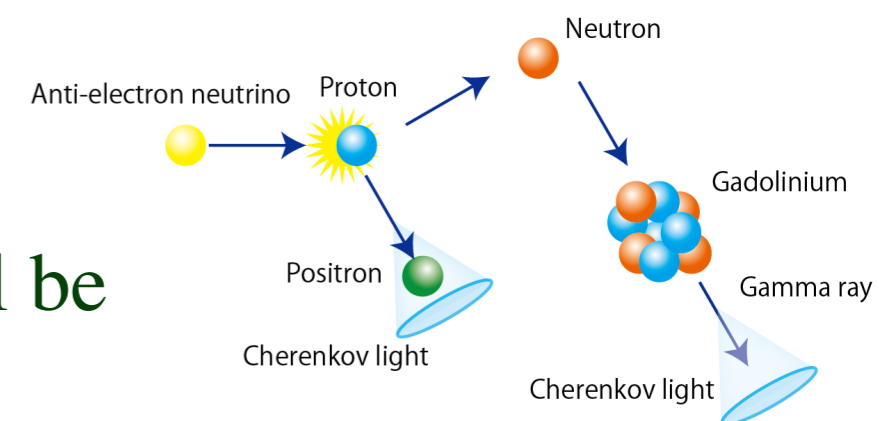
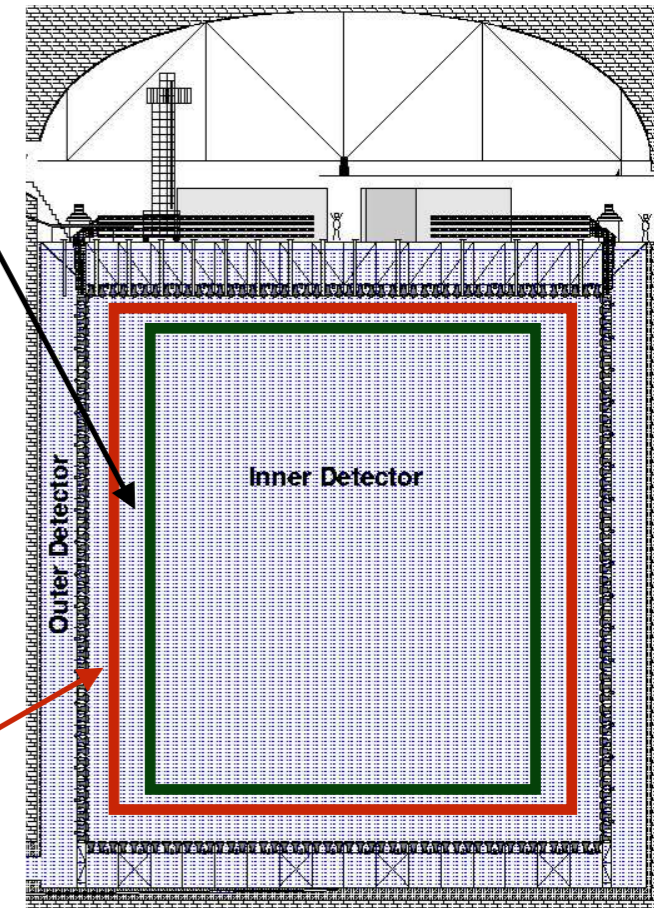
- ▶ [PhysRevD.102.112011](https://arxiv.org/abs/1102.1120)

- ▶ Neutron tagging

- ▶ Neutron tagging is applied to minimize the background from the atmospheric neutrinos.
- ▶ Neutron tagging with pure water is adopted in the data of 2008~.
- ▶ Analyses with more sensitive tagging by Gd will be published in near future.

Conventional fiducial volume

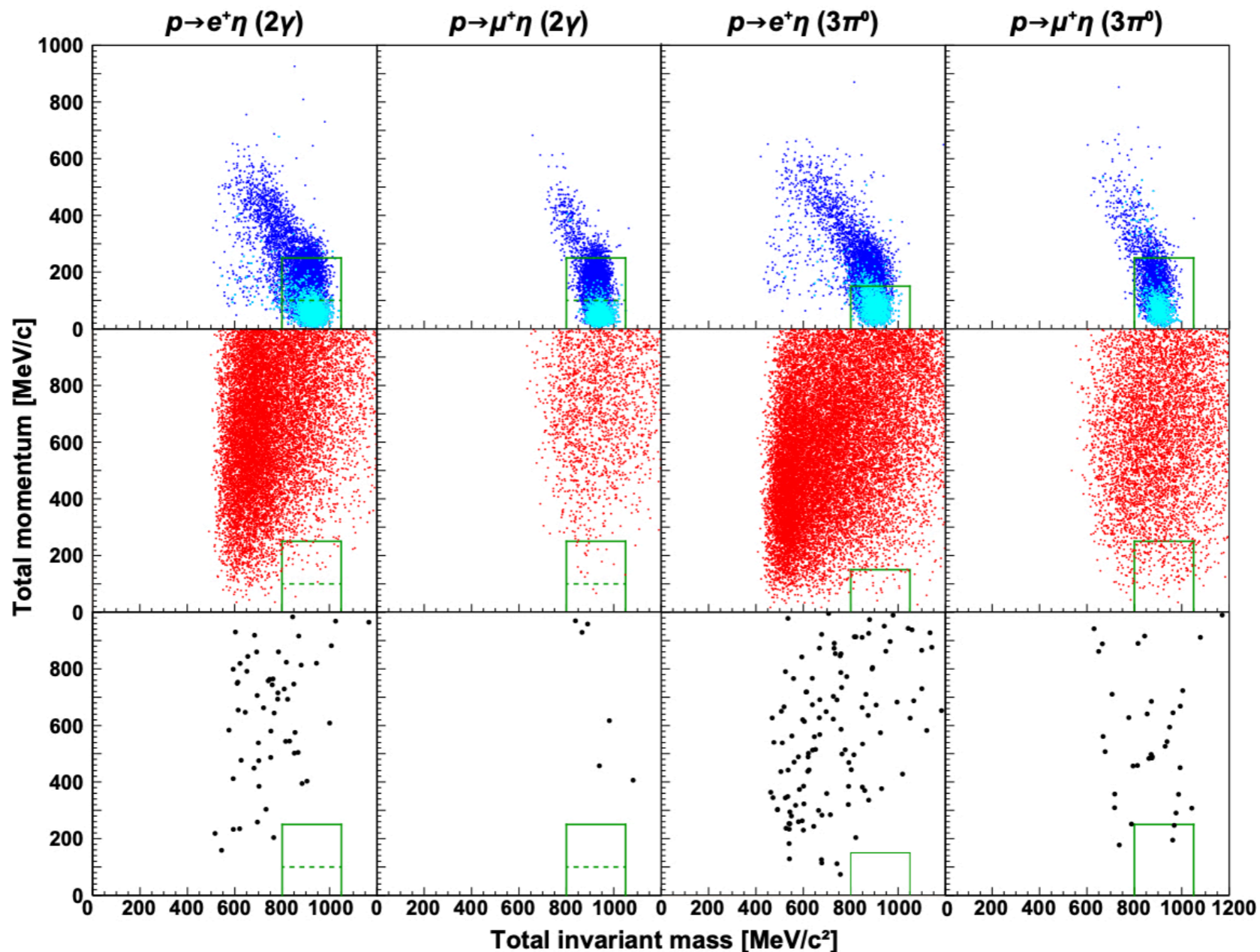
Enlarged fiducial volume





# Latest published analysis

- ▶ The search for  $p \rightarrow e^+ \eta, p \rightarrow \mu^+ \eta$  using the data of 1996-2018 was published in [Phys. Rev. D 110, 112011 \(2024\)](#).



- ▶ No significant excess was observed and the lower lifetime limit was set:  
 $\tau/\Gamma > 1.4 \times 10^{34}$  years ( $p \rightarrow e^+ \eta$ ),  $7.3 \times 10^{33}$  years ( $p \rightarrow \mu^+ \eta$ )

# Summary

- ▶ Since the start of the operation in 1996, SK has been exploring important cosmological and elementary particle physics, including neutrino and proton decay.
- ▶ In 2020, SK was upgraded by adopting the Gd-loaded water to enhance the neutron tagging performance. Now the neutron capture efficiency of 75% is achieved by the Gd concentration of 0.03%.
- ▶ In both neutrino and proton decay fields, SK has been updating many analysis channels.
- ▶ In near future, new physics results with new neutron-tagging techniques by Gd will be presented.  
**Keep Watching SK.**



# Back up