

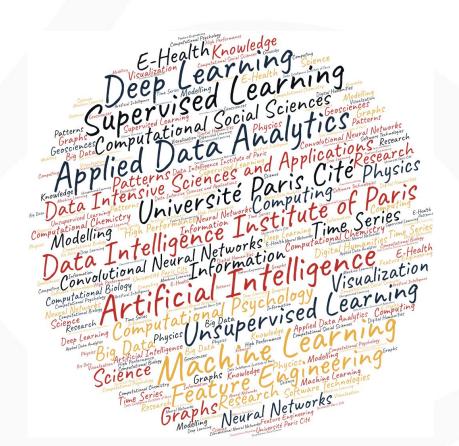


Foundation-like models for Astroparticle Physics Experiments

Signal Identification via Machine Learning Techniques

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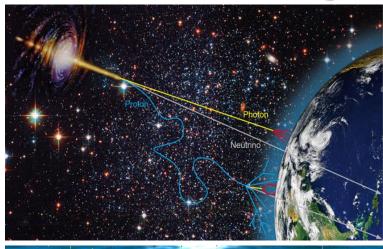
#### **Current Challenge in Astroparticle Physics Analyses:**

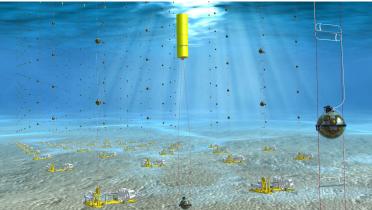
- Huge amount of Data is recorded
- Data is dominated by various types of noise
- Blows up processing time

Opportunity to accelerate analyses with Machine Learning

#### **Project**:

- Develop a Machine Learning pipeline for astroparticle physics experiments.
- Case study: KM3NeT



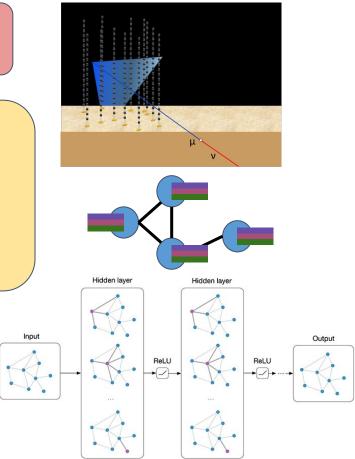




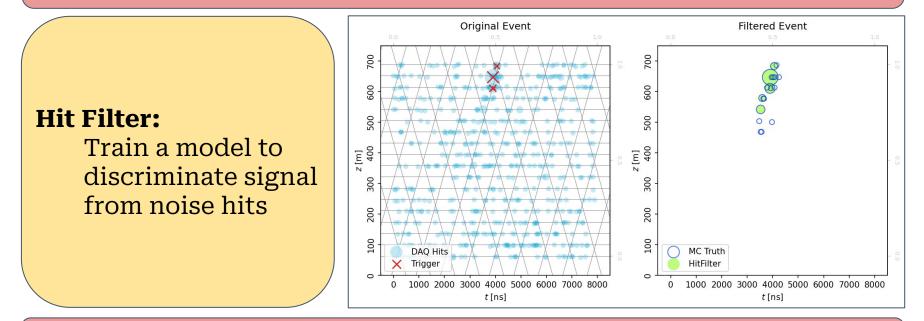
## Data can be best described by a Graph

- Nodes with features (e.g. *x*, *y*, *z*, *t*, *A*, ...)
- Edges defined via *k* nearest neighbors in Minkowski Spacetime (Captures spatial and temporal patterns)

## $\Rightarrow$ Use Graph Neural Networks

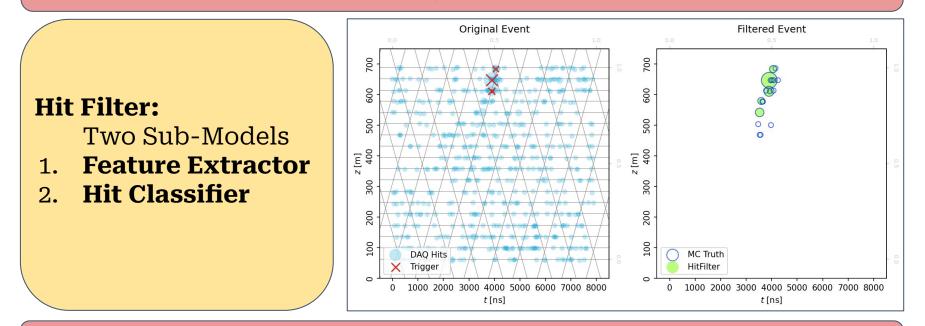


# **Challenge in KM3NeT:** Snapshot events are dominated by noise (K40)

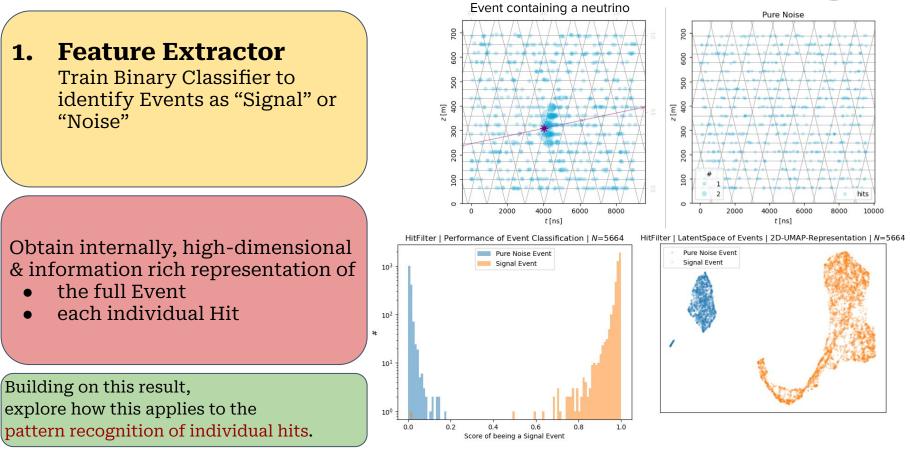


Continue analysis with cleaned event

# **Challenge in KM3NeT:** Snapshot events are dominated by noise (K40)

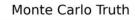


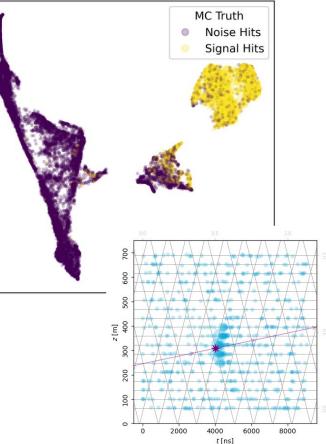
Continue analysis with cleaned event



#### Hit Filter: Latent Space of Hits





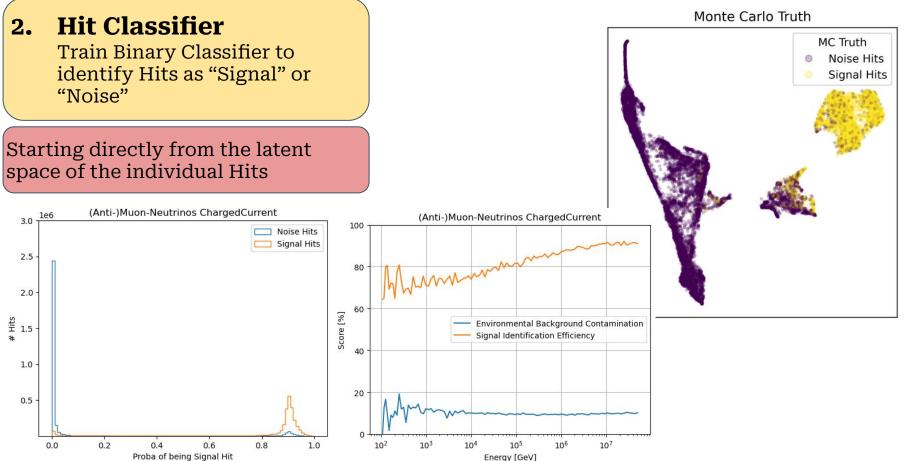


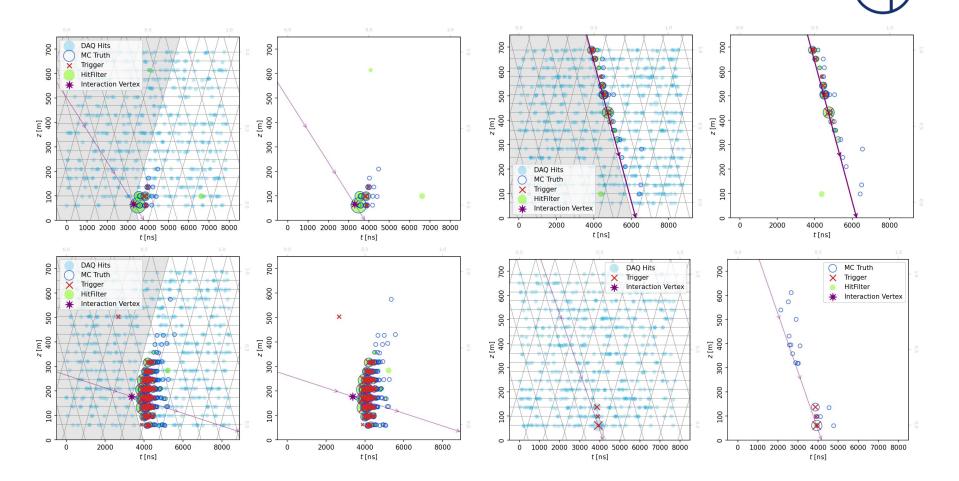
The latent space encodes both the representation of individual signals from the devices and the aggregated information within an event.

The remarkable classification accuracy of events containing neutrinos compared to pure noise events aligns closely with the strong separation achieved for individual signals in the devices.

The next step involves the identification the neutrino signals, which can be accomplished using a clustering approach or a straightforward supervised classifier.





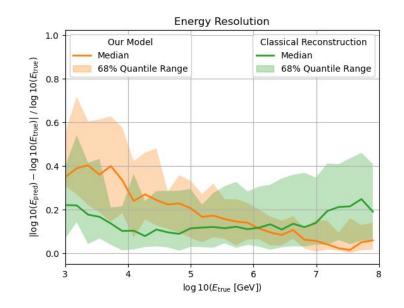


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#### **Pipeline Components**

- 1. Environmental Noise Filter Signal vs. noise discrimination
- 2. Neutrino Event Selection Atmospheric muon rejection
- Parameter Reconstruction
  Energy (Ε), direction (θ), and azimuthal angle (φ)





# Thank you for your attention!