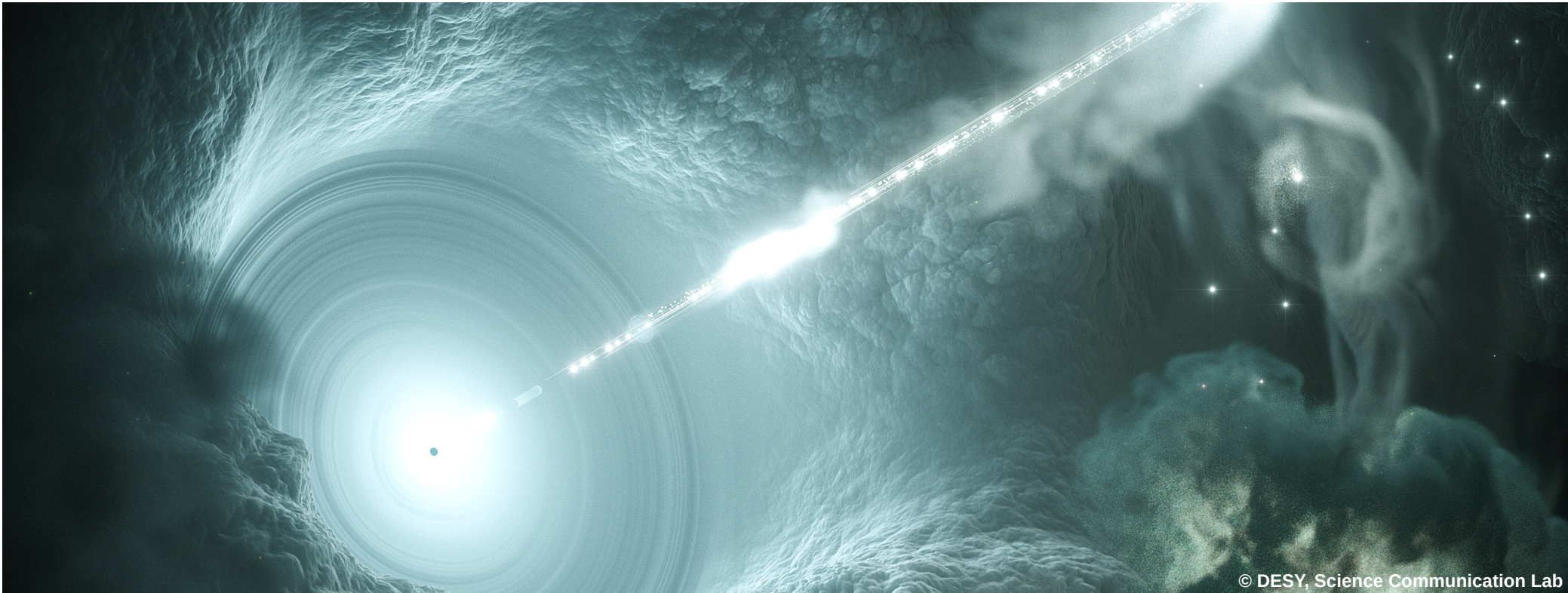




Facilitating multi-messenger modelling using Gammapy

Lea Heckmann, Matteo Cerruti, Bruno Khélifi, Justin Albinet



Active galactic nuclei (AGNs) & blazars

- Most luminous persistent sources in the Universe
- Supermassive black holes hosted at the centers of galaxies



Credit: <http://www.astro.princeton.edu/~lilew/>

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- Most luminous persistent sources in the Universe
- Supermassive black holes hosted at the centers of galaxies
- High-energy physics laboratories
- Potential multi-messenger sources (neutrinos and cosmic rays)
- But there are many open questions:
 - What is radiating?
 - How are particles accelerated?
 - In which environment is this taking place?

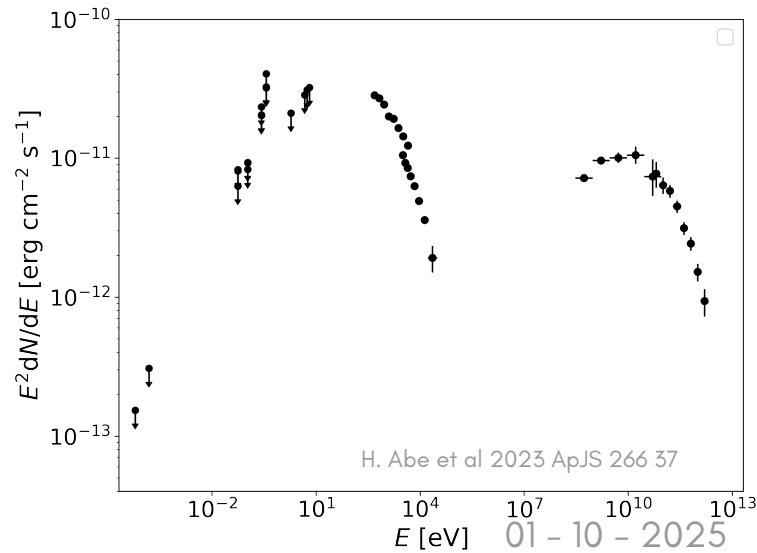


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Current AGN modelling

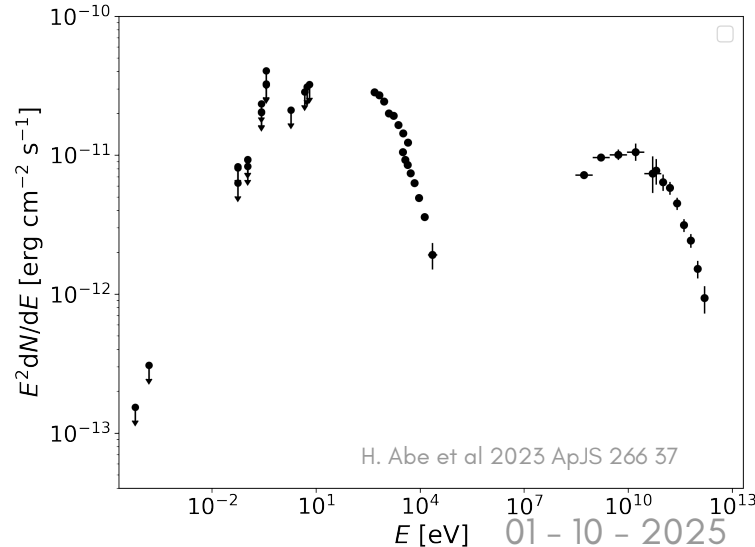
Current AGN modelling

- Combining MWL data on flux point level:



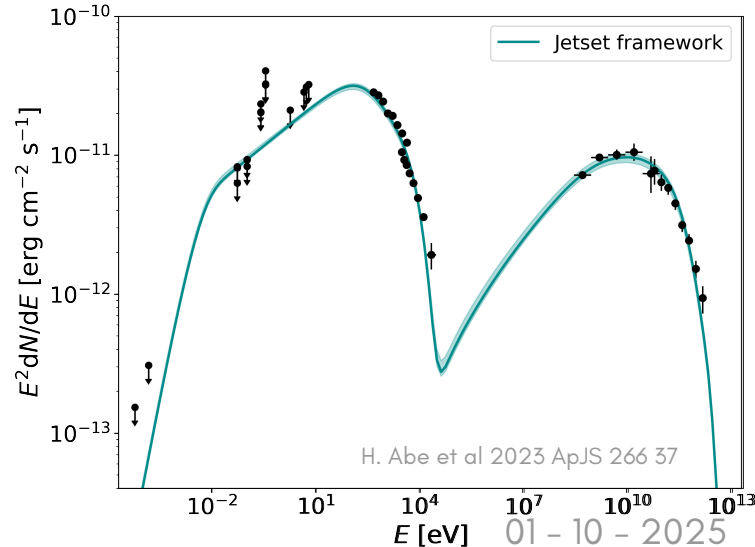
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- Combining MWL data on flux point level:
 - We construct flux points for single instruments separately
 - using different frameworks for different instruments
 - assuming simplistic models (e.g. power-law) – single instruments cannot constrain physical models (e.g. SSC)



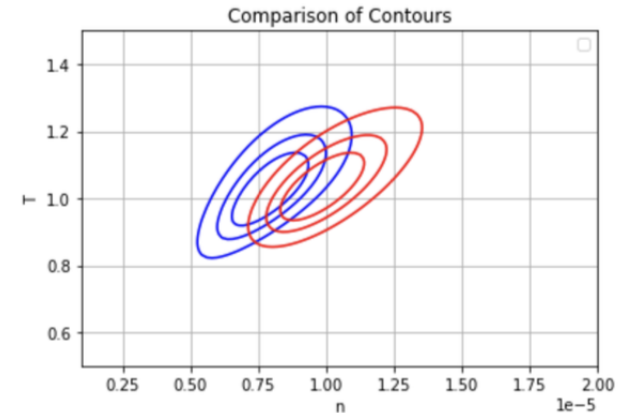
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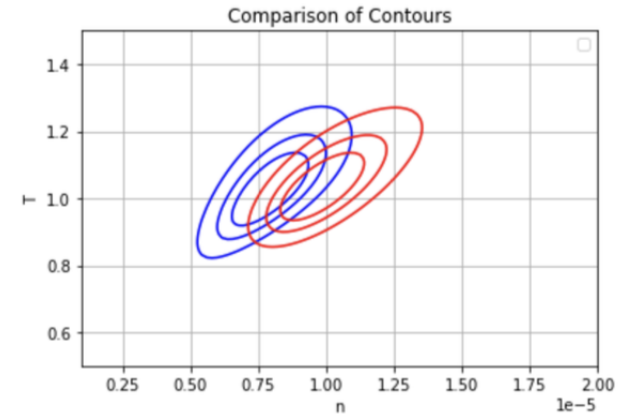
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- **Long-term goal: Enable event-based multi-messenger fits in Gammapy**
 - We start with blazars and JetSet as an example here



A new approach

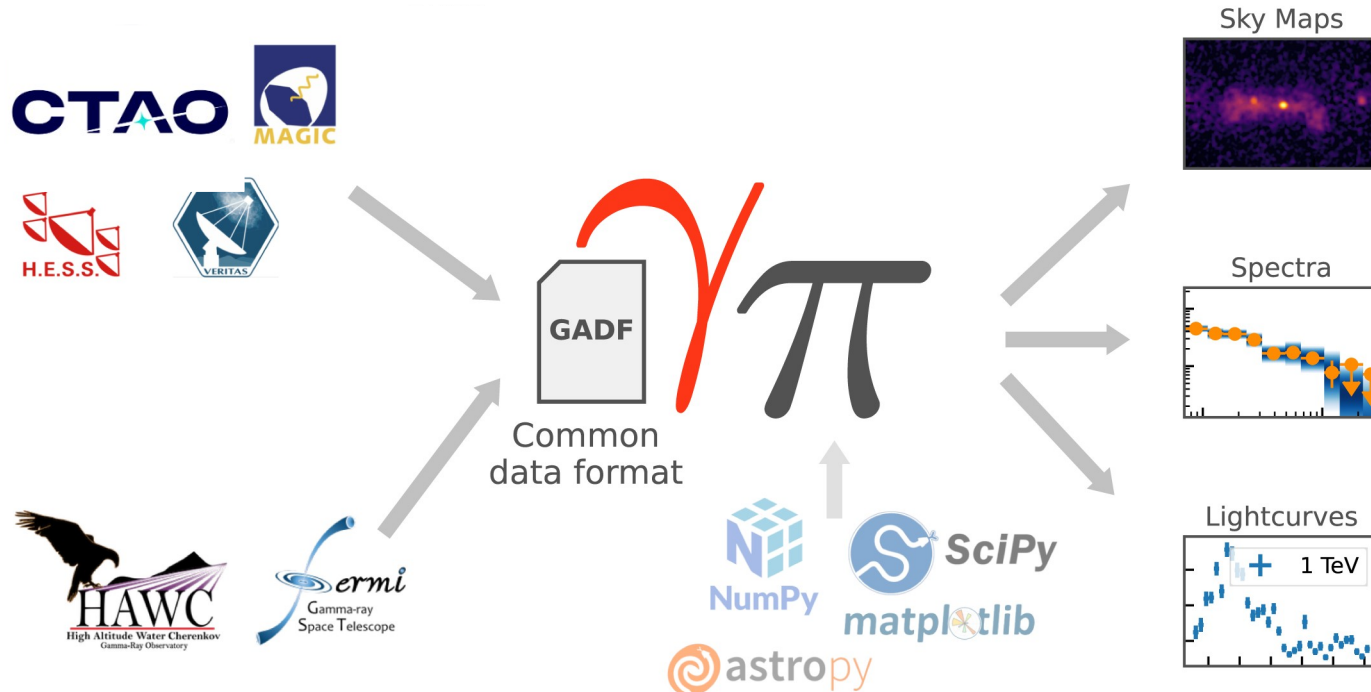
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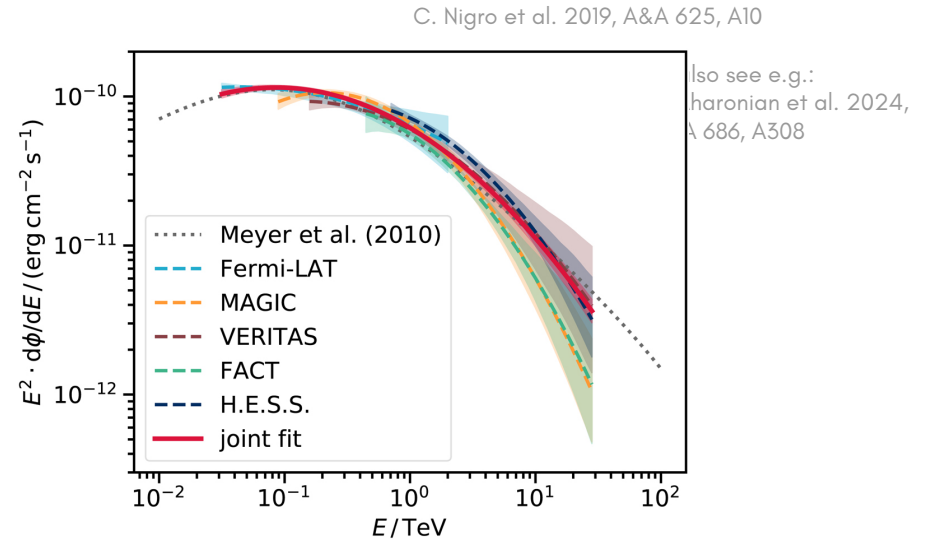
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Gammapy & MWL data

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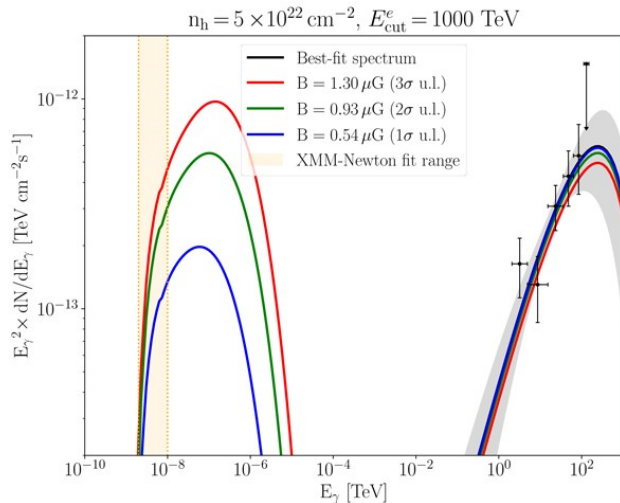


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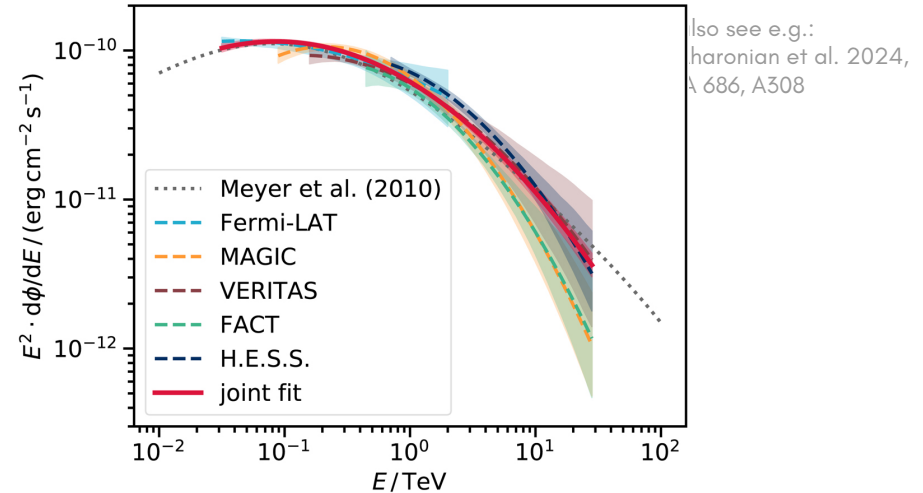
- Extension to X-rays:

<https://github.com/registerrier/gammagpy-ogip-spectra>



L. Giunti et al. 2022, A&A 667, A130

C. Nigro et al. 2019, A&A 625, A10

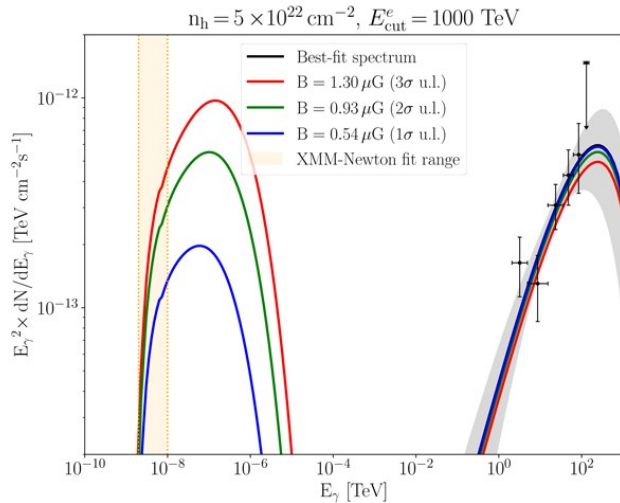


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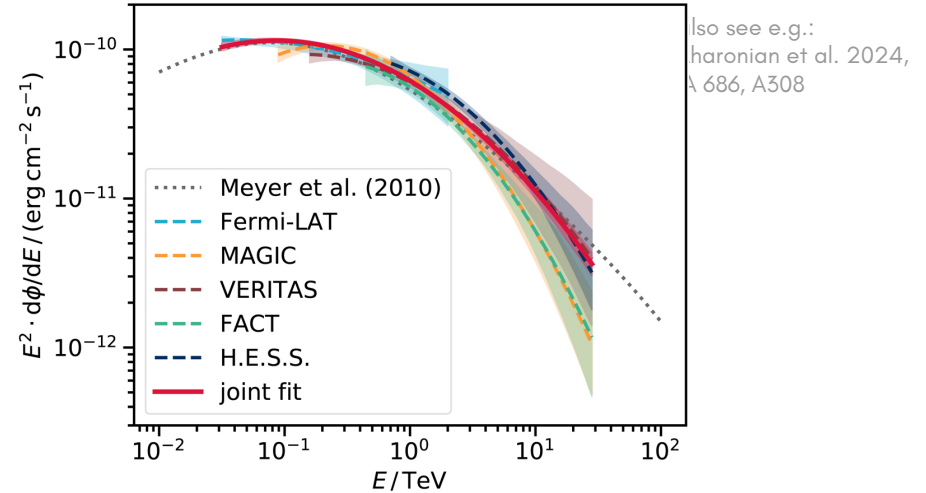
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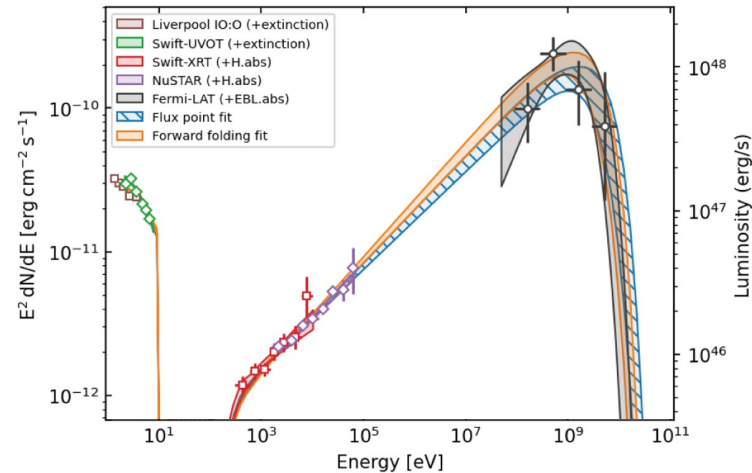
L. Giunti et al. 2022, A&A 667, A130



Also see e.g.:
Aharonian et al. 2024,
A&A, 686, A308

- Extension to UV/ optical:

https://github.com/mireianievas/gammapy_mwl_workflow



M. Nieves Rosillo et al. 2025,
A&A, 693, A287

New AGN modelling approach with Gammapy

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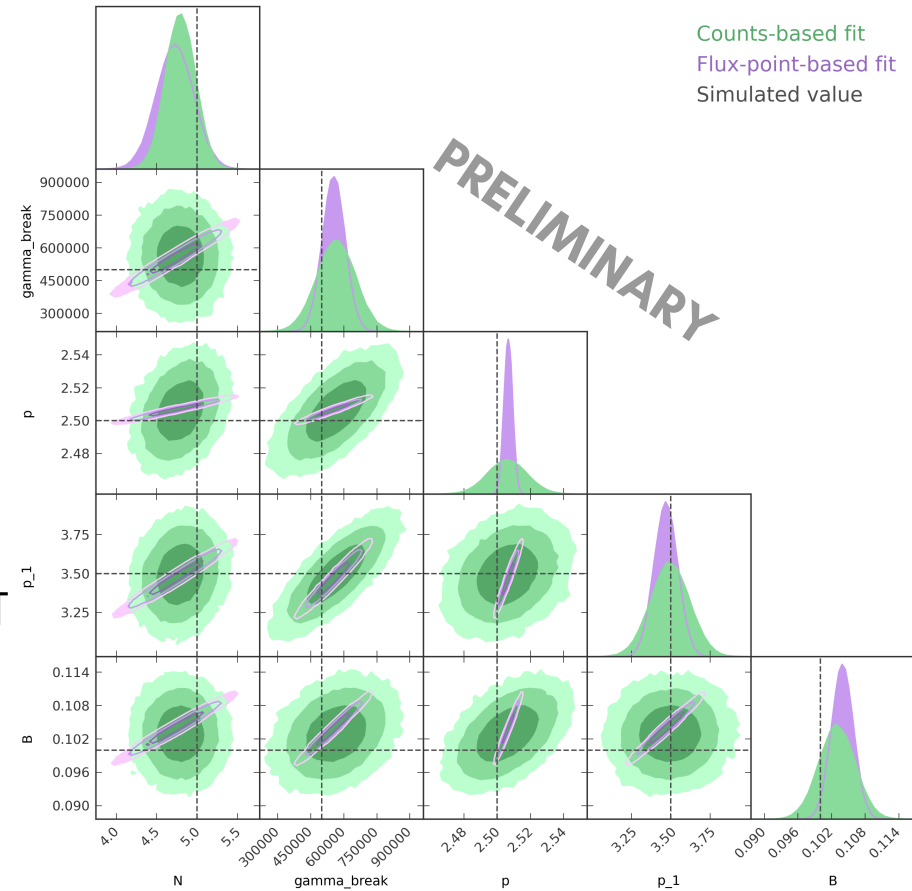
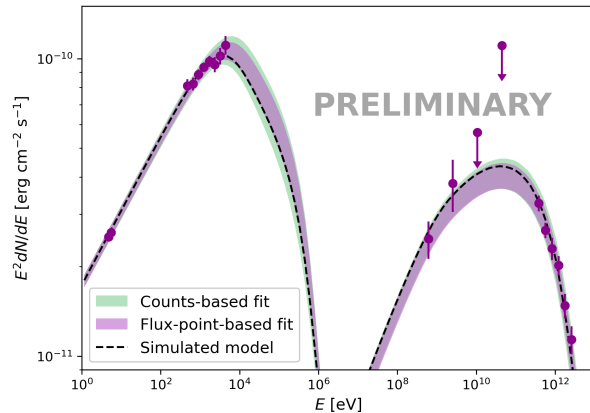
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- IRFs from UV to gamma-rays within Gammapy
 - Real data from Swift: UVOT + XRT, Fermi-LAT (UV, X-rays and gamma-rays)
 - Simulations from CTAO (Very-high-energy gamma-rays)
- Simulation of MWL data based on these IRFs using Gammapy:
- Physical model:
 - Leptonic Self-Synchrotron (SSC) model with EBL absorption
 - JetSeT ([2020ascl.soft09001T](#)) with its Gammapy plugin
- + Systematics for different bands
- + Absorption processes via the sherpa ([10.5281/zenodo.593753.](#)) module for xspec models

Gammapy & SSC

- We produce 2 MWL Gammapy datasets:
 - 1) Event data (DL3) set with event lists and IRFs
 - 2) Flux points data set:
 - Produced for each waveband
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- We fit both datasets with a SSC model using JetSet



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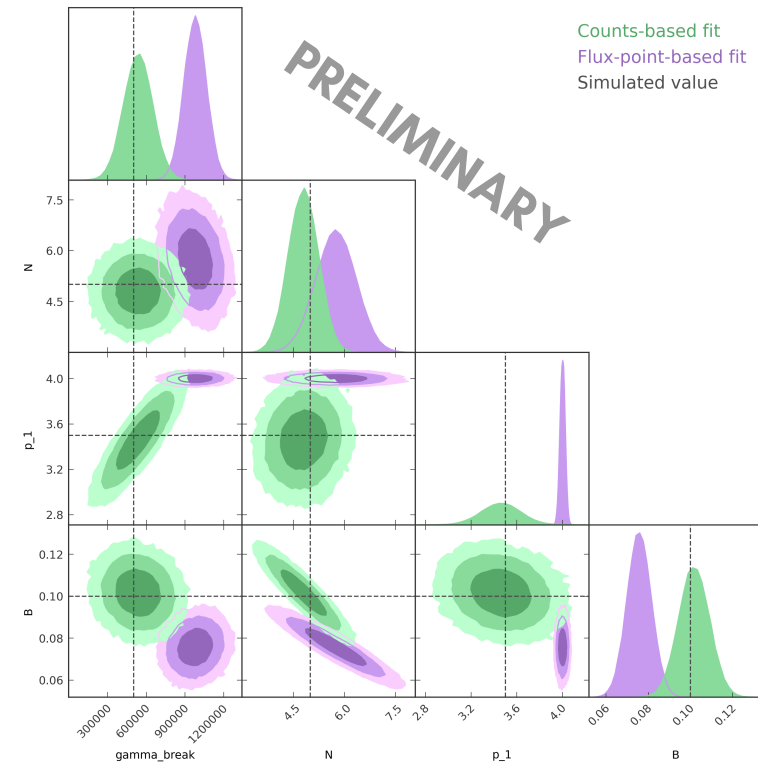
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- Using grid models and interpolation
 - Grid for E and parameters
 - SciPy RegularGridInterpolator
→ Works well from a physical point, but limited by computation time and memory usage



Facilitating multi-messenger modelling using Gammapy

- **Gammapy** enables us to **fit physical models directly on event basis**
 - Reducing biases and assumptions
 - Taking into account absorption, systematics or e.g. also ULs naturally
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- For **slow models (e.g. hadronic ones) – multimessenger fitting**
 - We are currently investigating **grid models + interpolation** methods
 - **Neural networks** and other machine learning methods might be a promising alternative

**Thank you
for your attention!**

**Get in touch!
We're open to different use-cases**