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Low-energy Test of Quantum Gravity via Angular Momentum Entanglement

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arXiv:2409.01364v3

Bridging high and low energies in search of quantum gravity, Paris, 9 July 2025

Supported by: ERC Synergy Grant HyperQ (No.856432), EU Project QuMicro (No.01046911) & DFG via QuantEra project LemaQume (No. 500314265)

Let me share a secret you might not know...

We don't have a universally accepted theory of quantum gravity!



So far, it is still impossible to consistently and/or uniquely quantize general relativity or to gravitize quantum mechanics

Planck length



Planck energy

10^{-35} m

10^{28} eV

10^{-18} m

10^{12} eV

10^{-15} m

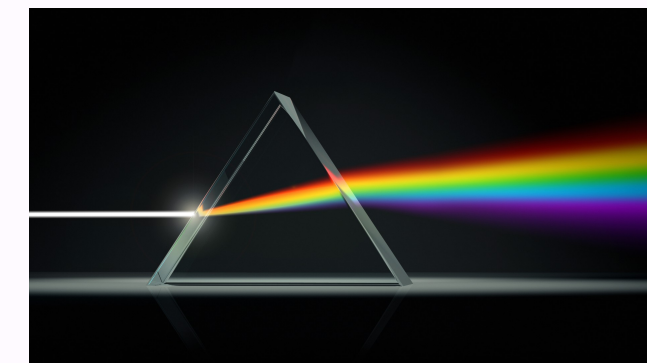
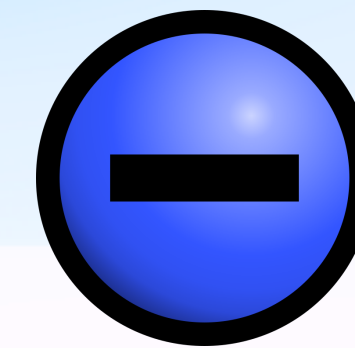
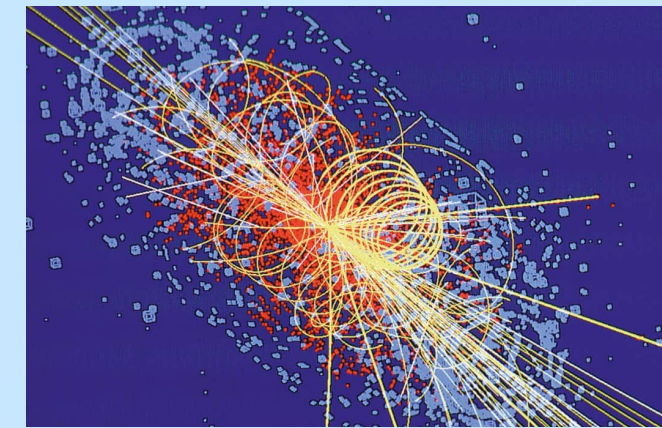
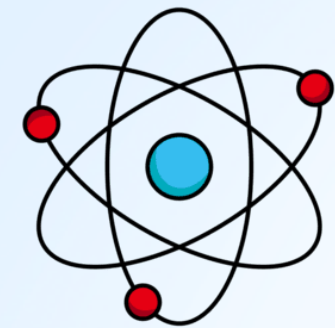
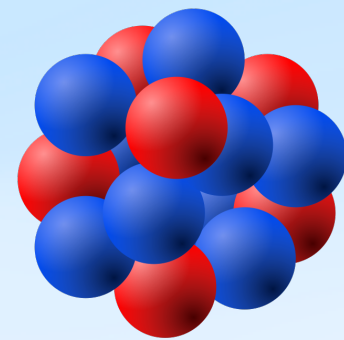
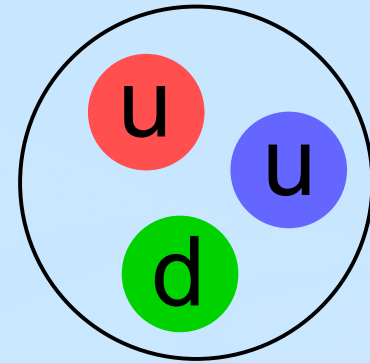
10^6 eV

10^{-10} m

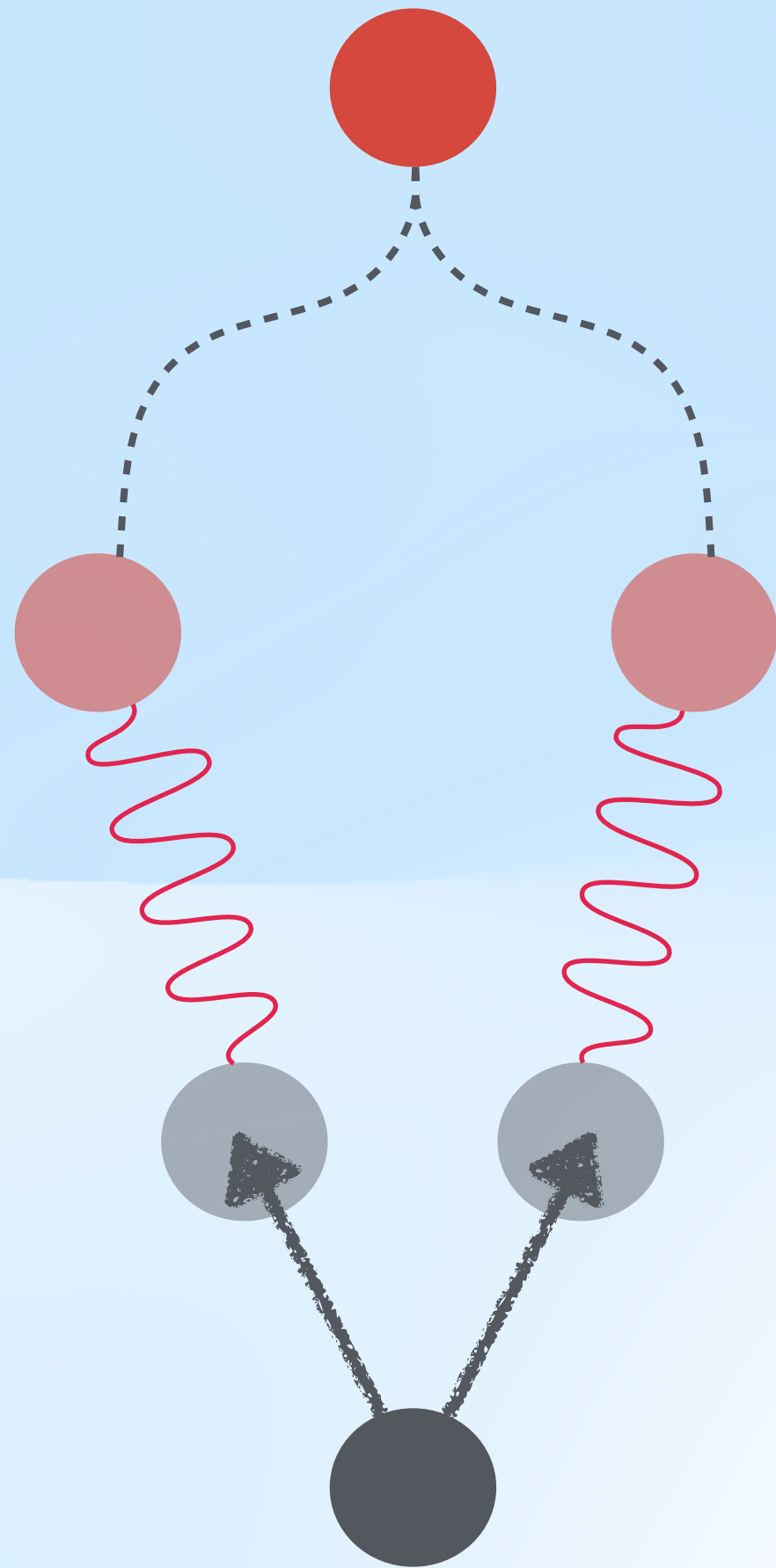
1 eV

1 m

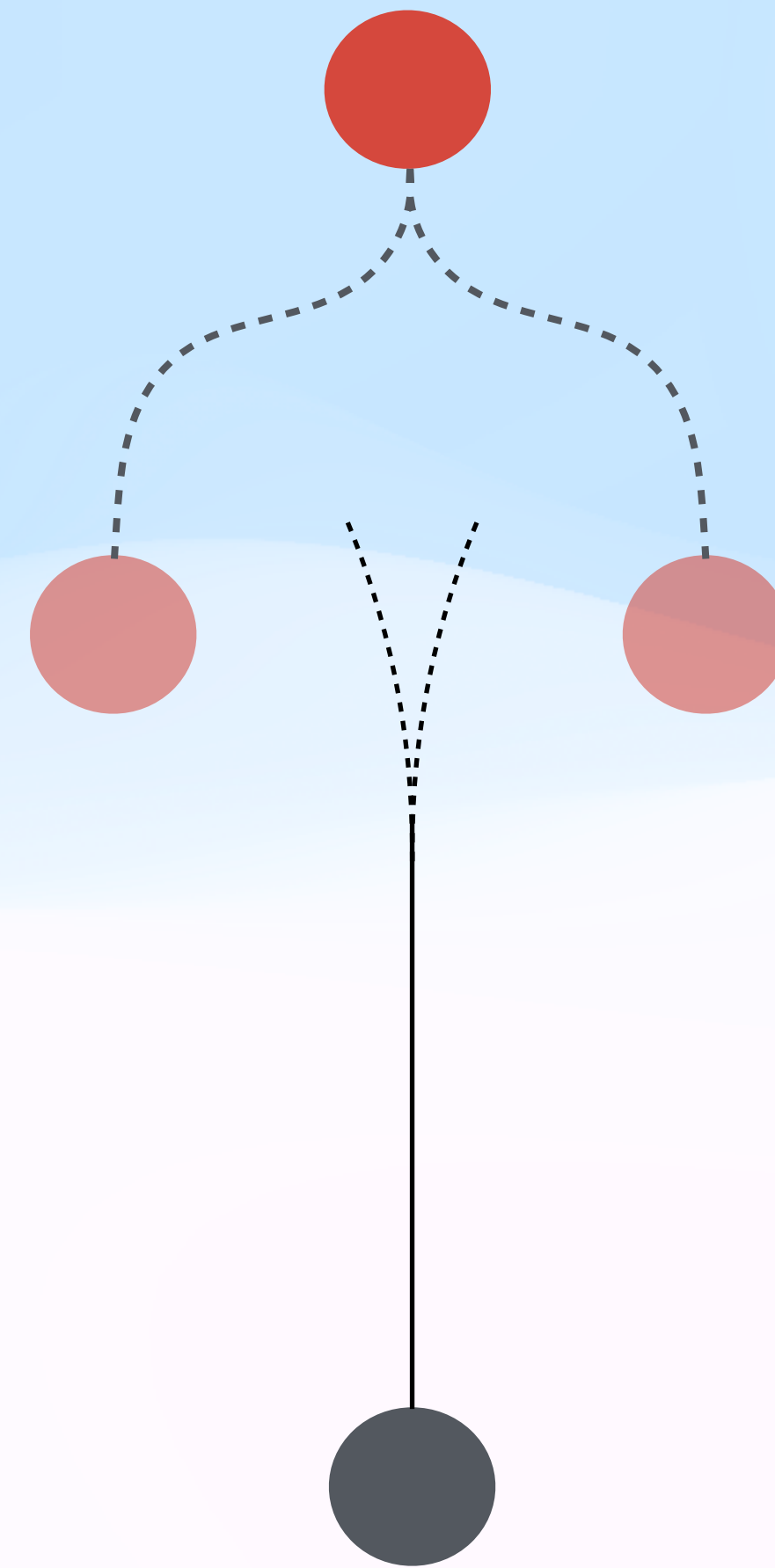
10^{-2} eV



There might be hope to witness nonclassical aspects of gravity

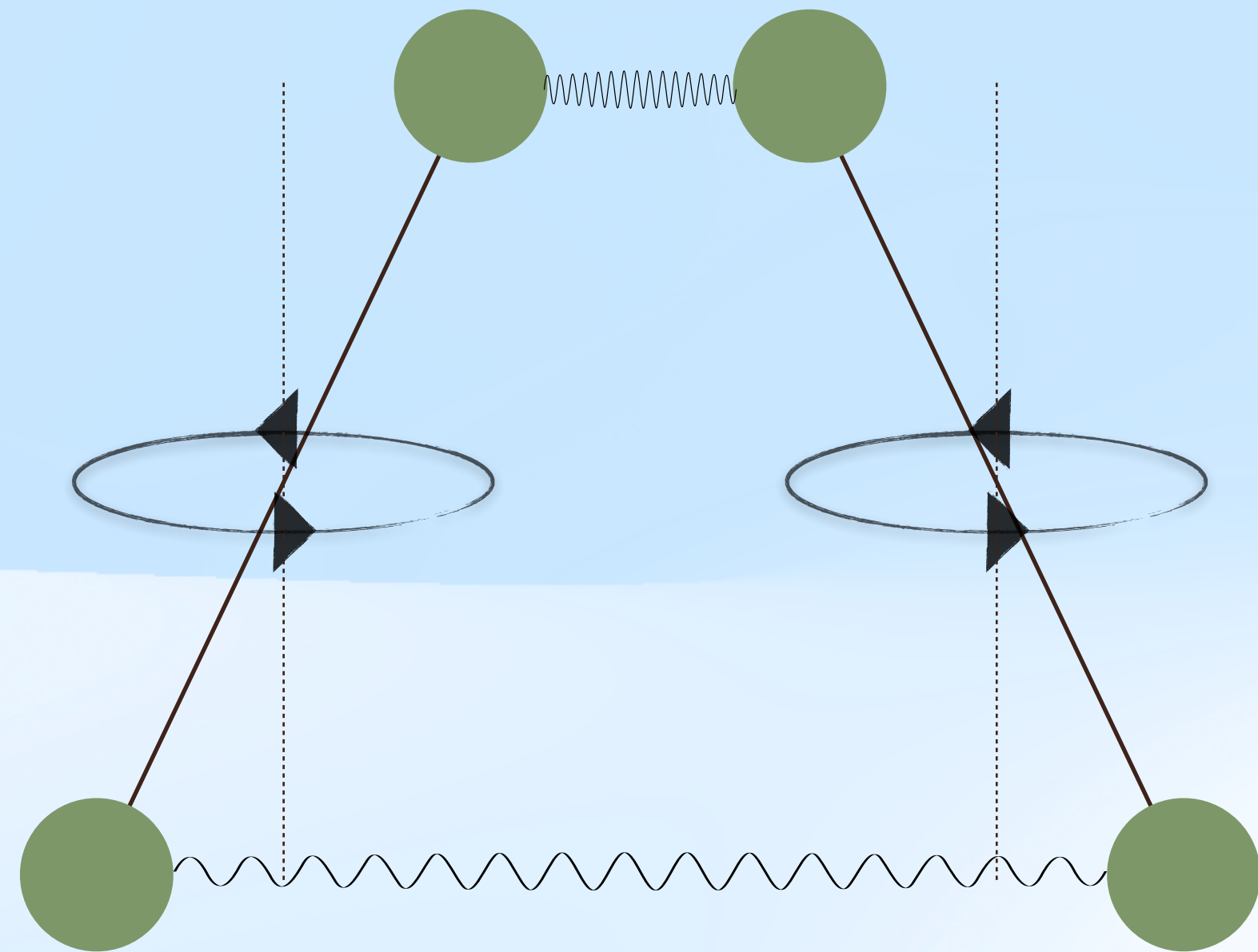


Feynman (1957)

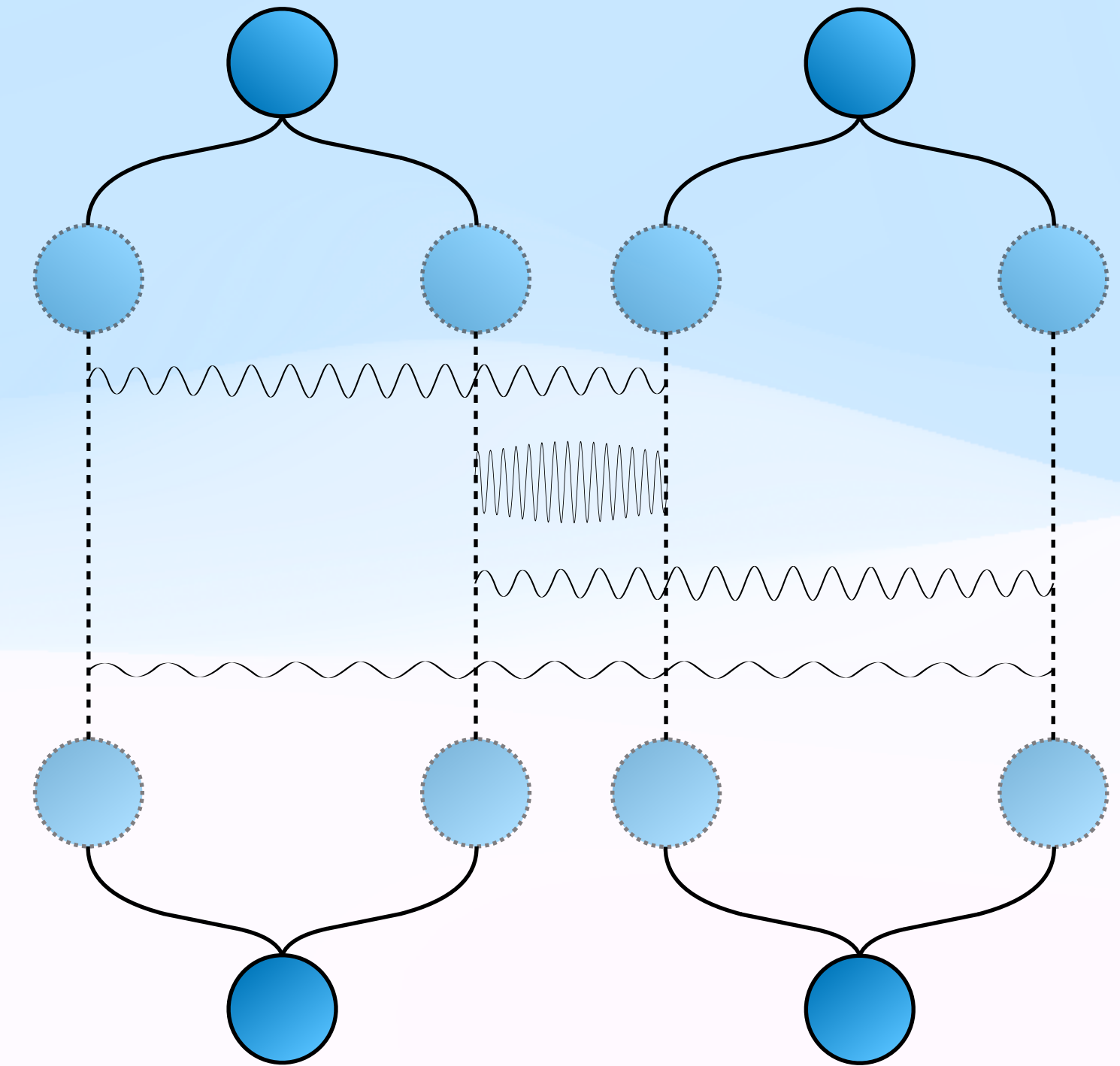


Lindner & Peres (2005)

More recent proposals revolve around the LOCC theorem



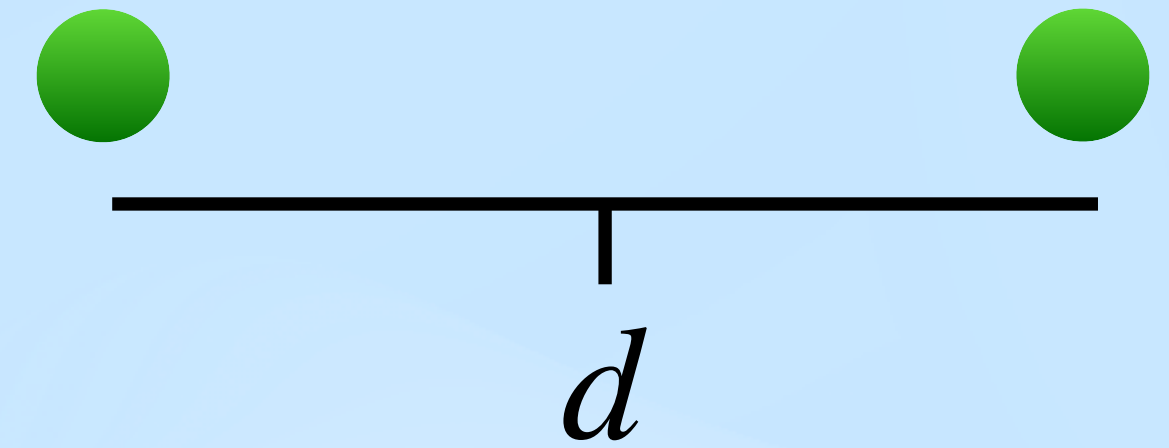
Kafri & Taylor (2013), Lami et al. (2024)



Bose et al. (2017), Marletto and Vedral (2017)

What do all these approaches have in common?

→ Good ol' Newton



In the two limits

$$\frac{d}{c} \ll t$$
$$|\mathbf{v}| \ll c$$

(near-field)

(low velocity)

$$V(\mathbf{r}) \approx \frac{Gm_A m_B}{|\mathbf{r}|}$$

Christodoulou et al. (2023)

→ Entanglement in spatial d.o.f.s

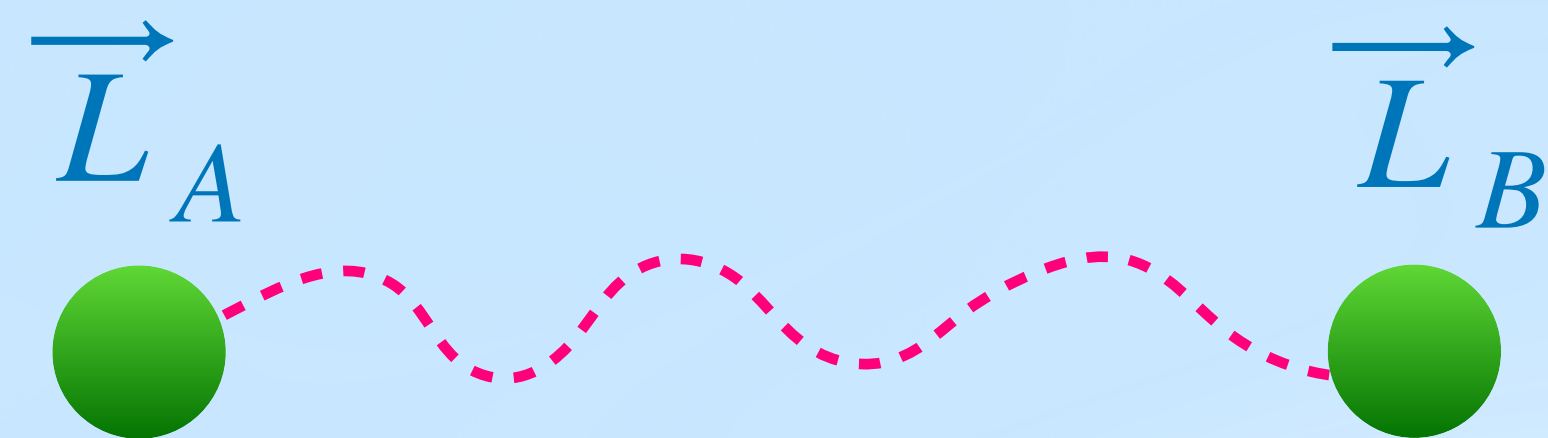
PROs

- Low Hilbert space dimensionality (only two states available)
- Easy measurement scheme (at least in principle)

CONs

- Difficulty in the realization of the initial state
- No direct general relativistic effect into play

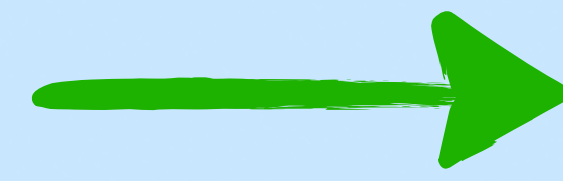
We propose an alternative scheme with Post-Newtonian corrections



$\sim$ Newtonian + Post-Newtonian

$$-\frac{G}{c^2 r^3} \left[\vec{L}_A \cdot \vec{L}_B - 3 \left(\vec{L}_A \cdot \vec{e}_r \right) \left(\vec{L}_B \cdot \vec{e}_r \right) \right]$$

Fast rotating SiO₂ microspheres
 $\omega \approx 10^7$ Hz, $R = 50\mu\text{m}$

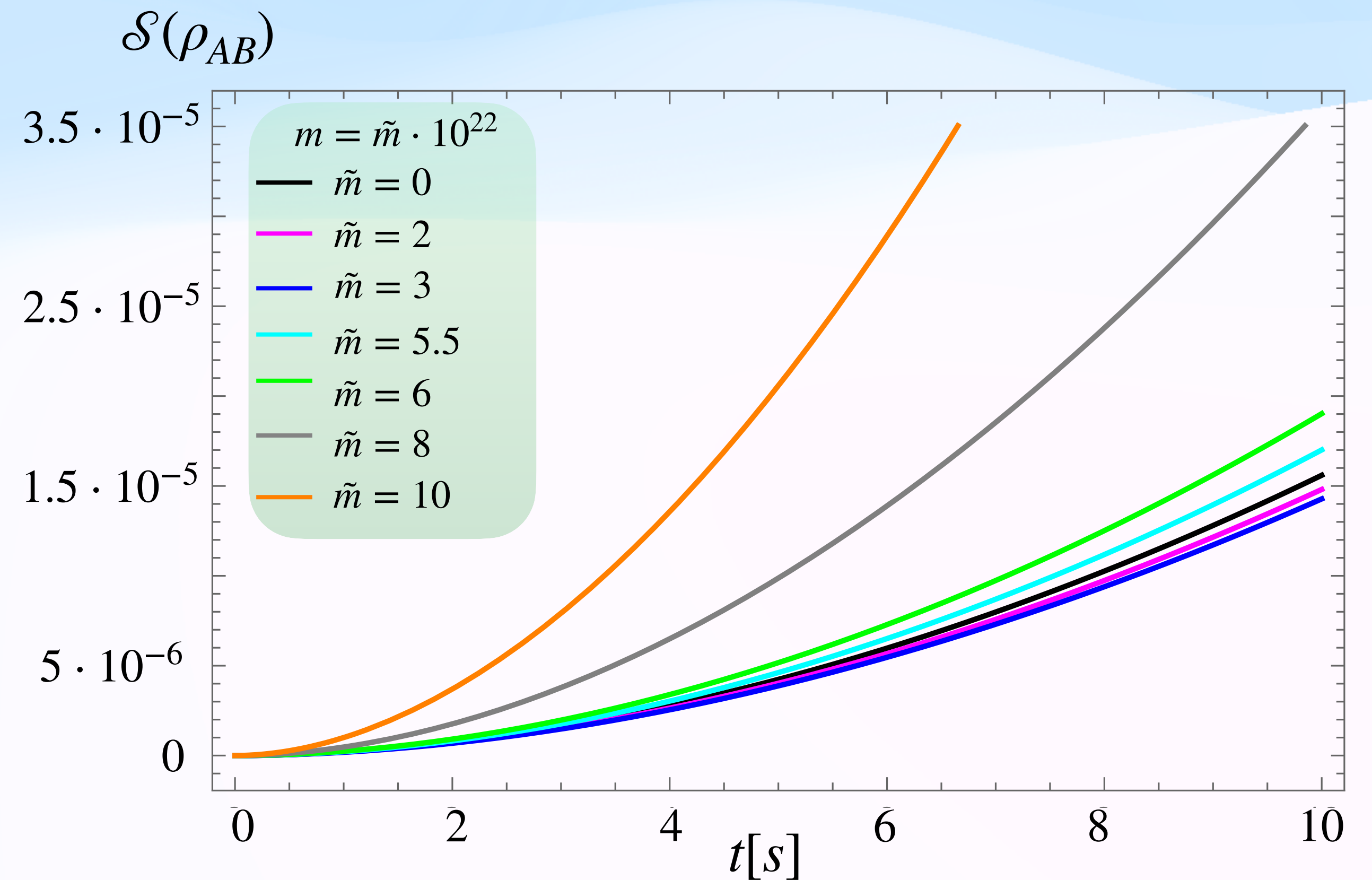


$$l = \frac{L}{\hbar} \approx 10^{23}$$

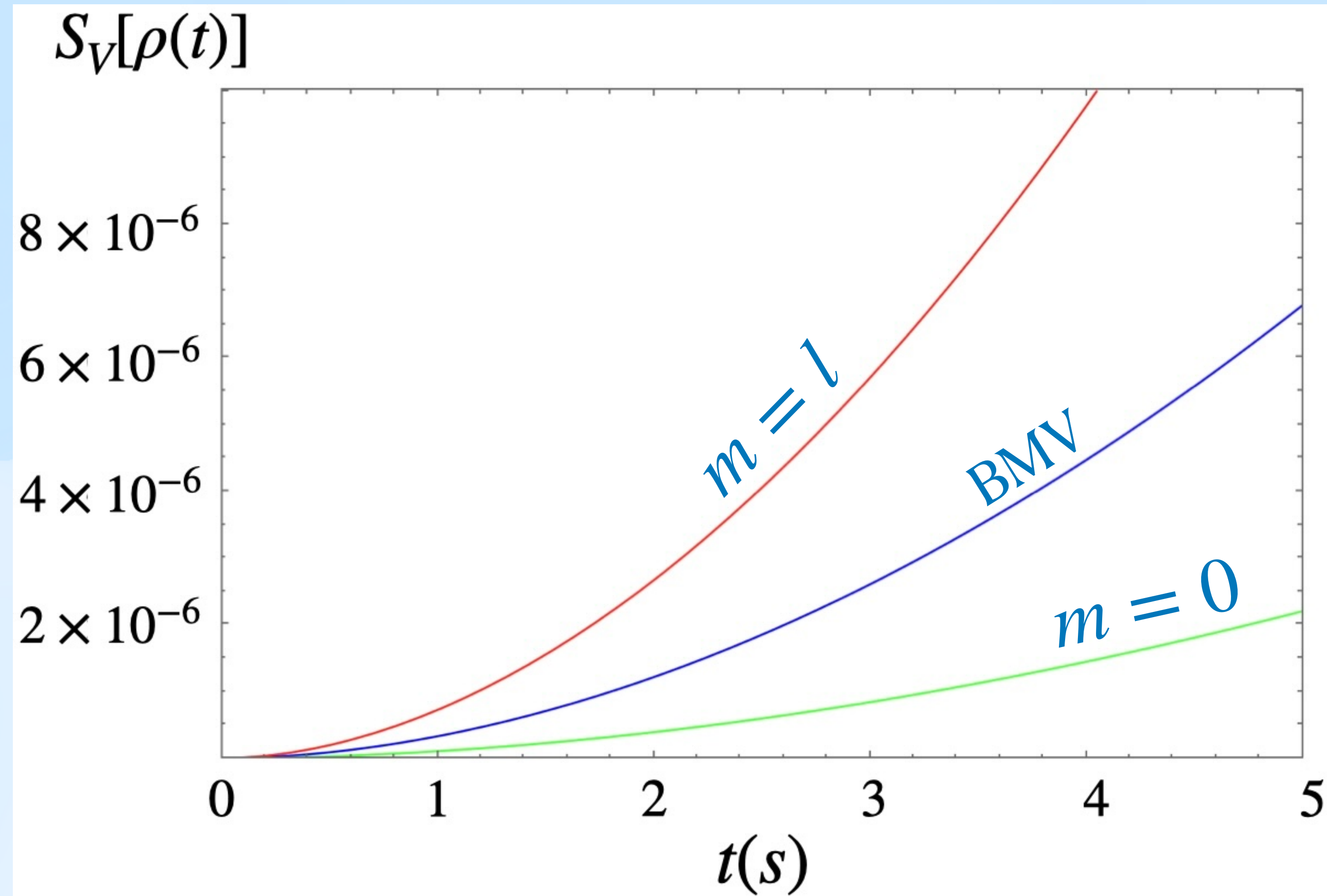
$$|\psi_{AB}(0)\rangle = (|l, m\rangle_A + |l, -m\rangle_A) \otimes (|l, m\rangle_B + |l, -m\rangle_B)$$



$$\hat{H}_I = -\frac{G\hbar^2}{2c^2 r^3} \left(\hat{L}_+^{(A)} \hat{L}_-^{(B)} + \hat{L}_-^{(A)} \hat{L}_+^{(B)} - 4\hat{L}_z^{(A)} \hat{L}_z^{(B)} \right)$$



Isn't the effect too small?



But what are the differences with other QG tests?

PROs

- NO initial superpositions required
- Most of the usual decoherence channels do not apply here

CONs

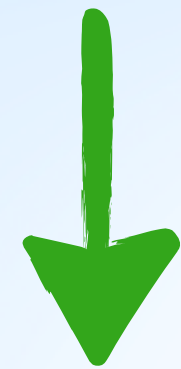
- Difficult detection of angular momentum quanta
- Other decoherence channels might be troublesome

A protocol to measure angular momentum

A fast-spinning object tends to spontaneously magnetize in the direction of the rotation axis [Barnett (1915)]

$$\hat{H}_{c.o.m.} = \frac{\hat{p}^2}{2m} + \frac{|\chi_V| V \hat{B}^2}{2\mu_0} + \frac{|\chi_V| V}{2\gamma\mu_0 I} \left(\hat{\vec{B}} \cdot \hat{\vec{L}} + \hat{\vec{L}} \cdot \hat{\vec{B}} \right),$$

With a small magnetic gradient along z



$$\Delta^2 \hat{L}_z = \left(\frac{\gamma I B_0}{2 \sin^2 \left(\frac{\Omega t}{2} \right)} \right)^2 (\Delta^2 \hat{z}(t) + \Delta^2 \hat{z}(0))$$

Angular momentum can be inferred from center of mass position!

Sources of decoherence

Magnetic dipole-dipole interaction

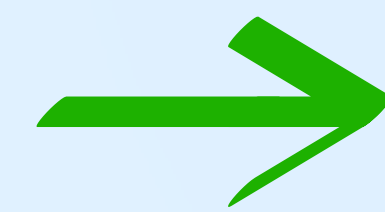
Looks just like our gravitational Hamiltonian!

$$H_{dip} = \frac{\mu_0 \gamma^2 \hbar^2}{4\pi r^3} \left[\vec{S}_A \cdot \vec{S}_B - 3 \left(\vec{S}_A \cdot \vec{e}_r \right) \left(\vec{S}_B \cdot \vec{e}_r \right) \right]$$

$\approx 10^9$ nuclear spins in a SiO_2 sphere of radius $R = 50\mu\text{m}$

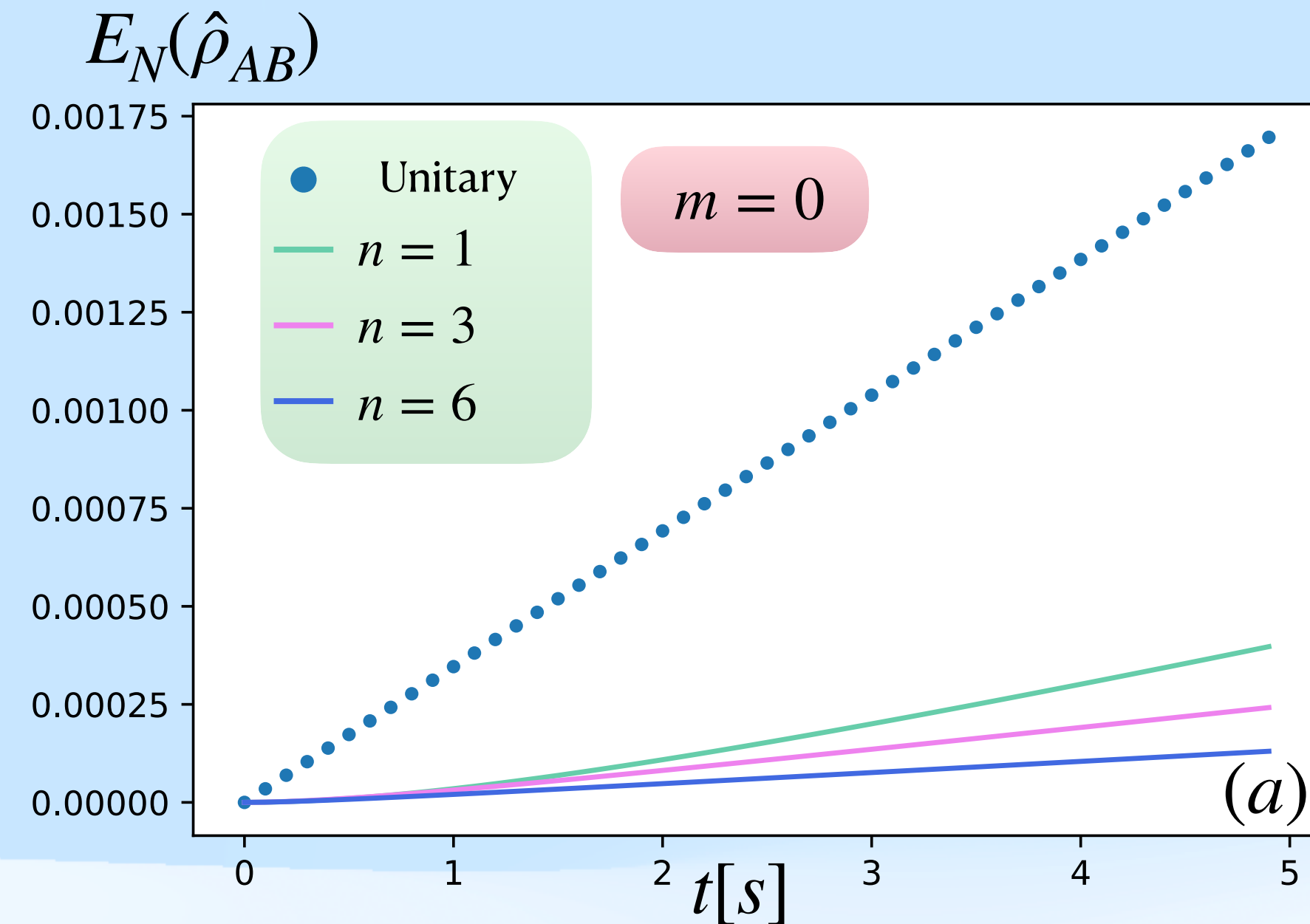
$$V_M = \frac{\mu_0 \gamma^2 \hbar^2}{4\pi r^3} \times 10^{18} \approx \cdot 10^{-28} \times p^2 \text{ J}$$

$$V_G = \frac{G \hbar^2 l^2}{c^2 r^3} \approx 1.2 \cdot 10^{-38} \text{ J}$$



$$\frac{V_M}{V_G} \approx 10^{10} p^2$$

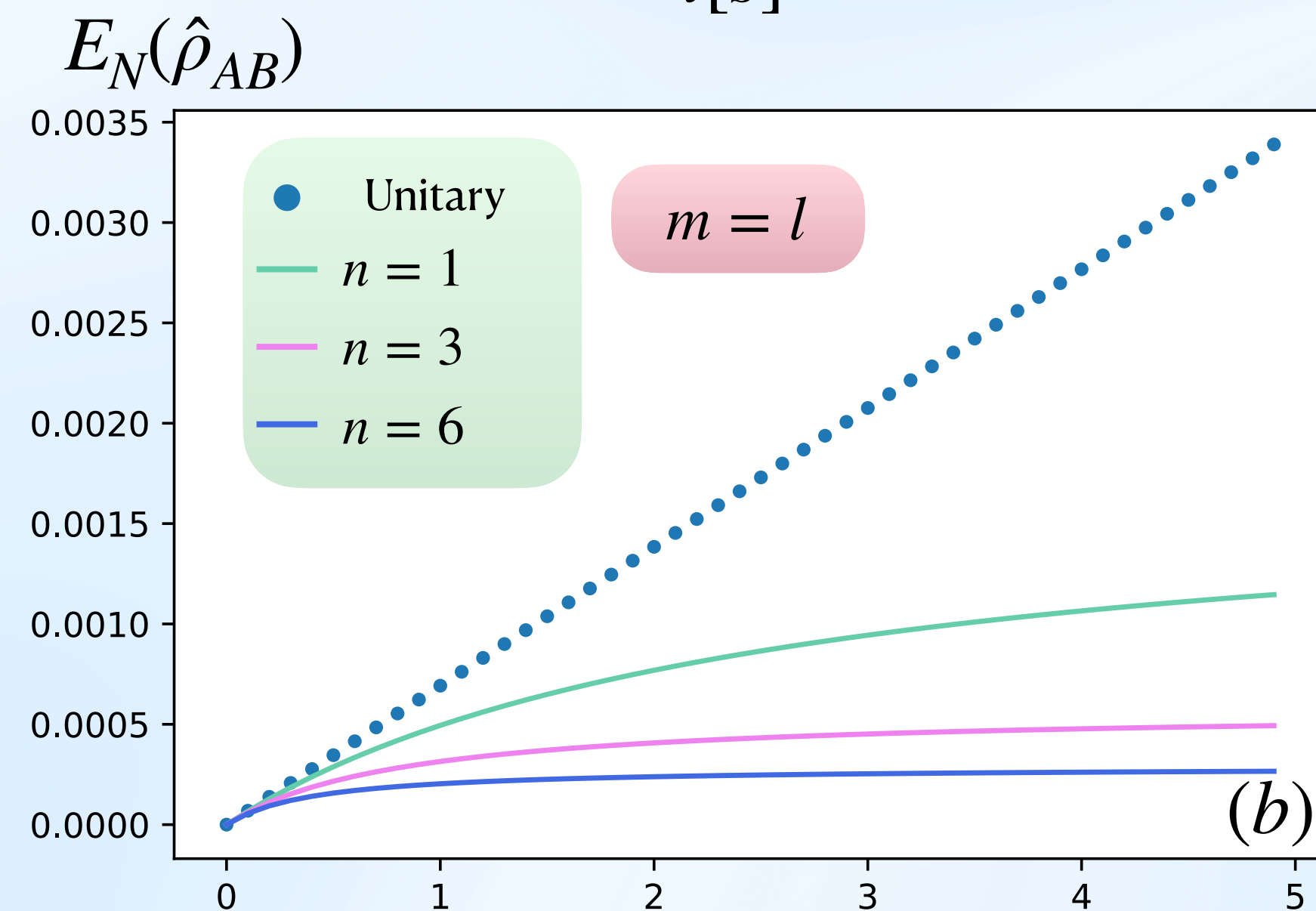
Decoherence due to one collision



The event can transfer up to n quanta to the m number

$$T = 0.1 \text{ K and } P = 10^{-16} \text{ Pa}$$

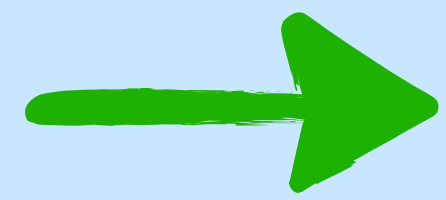
Dramatic drop of entanglement!



Probability of collision
must be $\ll 1$

Sources of decoherence

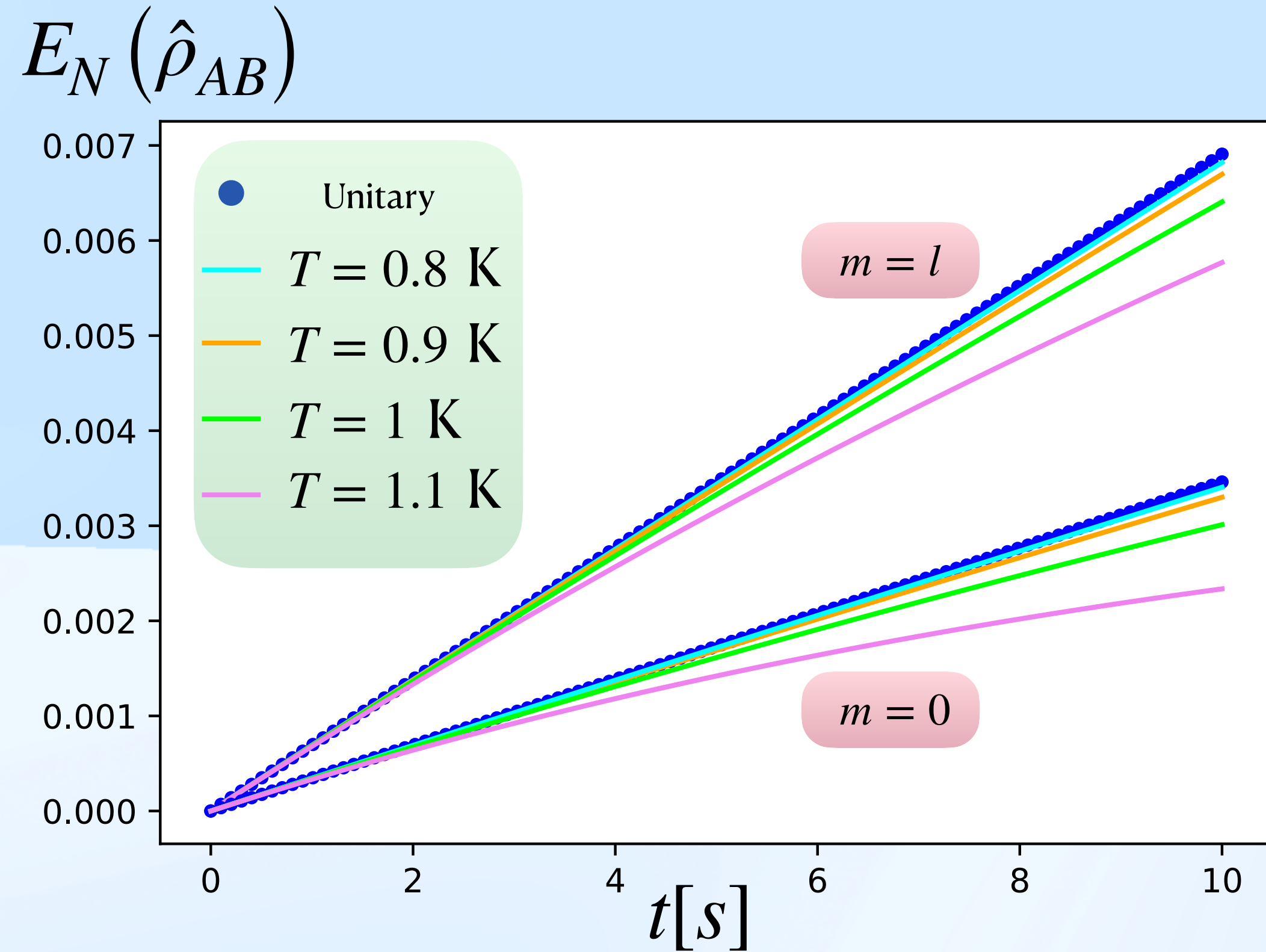
- Heating due to laser



If we start rotating the spheres at $T_0 = 1\text{K}$, with a laser of 300nm , this leads to $T_f = 1.13\text{K}$ to achieve $\omega = 10^7\text{Hz}$

Is this temperature detrimental for the entanglement?

Decoherence due to thermal radiation



$$\begin{aligned} \frac{d}{dt}\hat{\rho}_S(t) = & -\frac{i}{\hbar} [\hat{H}_I^{AB}, \hat{\rho}_S(t)] \\ & + \sum_{l \geq 0} \frac{\Delta_l^3 \hbar^2}{6c^3 I^3 \epsilon_0} (1 + N(\Delta_l)) \left[\hat{\vec{A}}^{AB}(\Delta_l) \cdot \hat{\rho}_S(t) \hat{\vec{A}}^{AB\dagger}(\Delta_l) - \frac{1}{2} \left\{ \hat{\vec{A}}^{AB\dagger}(\Delta_l) \cdot \hat{\vec{A}}^{AB}(\Delta_l), \hat{\rho}_S(t) \right\} \right] \\ & + \sum_{l \geq 0} \frac{\Delta_l^3 \hbar^2}{6c^3 I^3 \epsilon_0} N(\Delta_l) \left[\hat{\vec{A}}^{AB\dagger}(\Delta_l) \cdot \hat{\rho}_S(t) \hat{\vec{A}}^{AB}(\Delta_l) - \frac{1}{2} \left\{ \hat{\vec{A}}^{AB}(\Delta_l) \cdot \hat{\vec{A}}^{AB\dagger}(\Delta_l), \hat{\rho}_S(t) \right\} \right], \end{aligned}$$

With $\Delta_l = 2(l + 1)$

$$\hat{A}_1^{AB} = \sum_{i=A,B} \sum_{m=-l}^l c_m^{(1)} |l, m\rangle_i \langle l+1, m+1|_i$$

$$\hat{A}_2^{AB} = \sum_{i=A,B} \sum_{m=-l}^l c_m^{(2)} |l, m\rangle_i \langle l+1, m-1|_i$$

$$\hat{A}_3^{AB} = \sum_{i=A,B} \sum_{m=-l}^l c_m^{(3)} |l, m\rangle_i \langle l+1, m|_i$$

Take-home message

- Many proposals for low-energy signatures of quantum gravity
- Angular momentum entanglement can be a promising alternative
- Effort in seeking alternative tests to the ones already available is required

Thanks for the attention!



Check the preprint here!