



UNIVERSITÀ
DEGLI STUDI
FIRENZE



Emergent Scalar Field Dynamics on Curved Spacetime in Group Field Theory

Roukaya Dekhil

In collaboration with Federico Greco, Stefano Liberati
and Daniele Oriti

Bridging high and low energies in search of quantum gravity - 2025
Cost Action CA23130 First Annual Conference

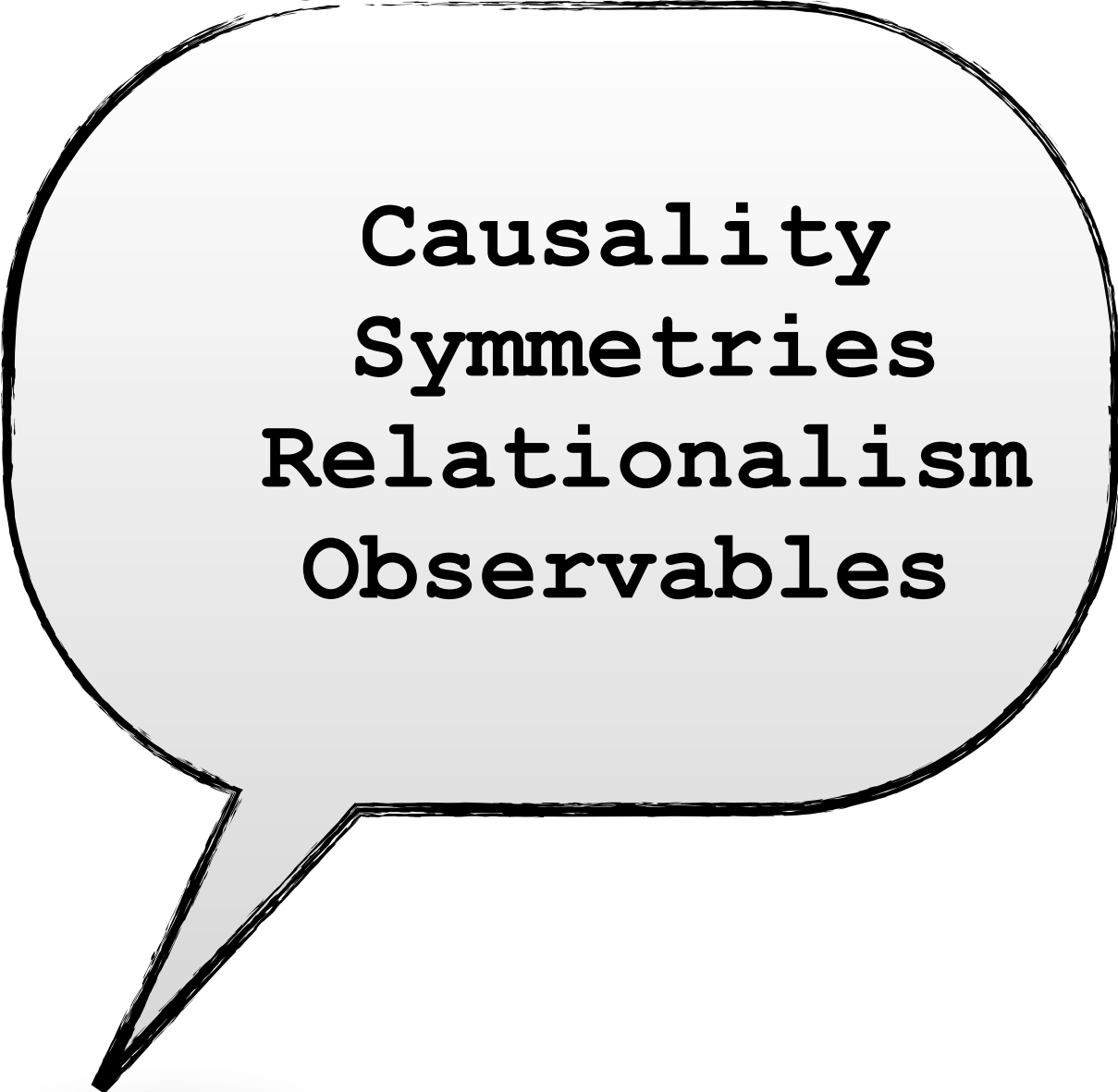
10th of July 2025, Paris

Outline

Outline

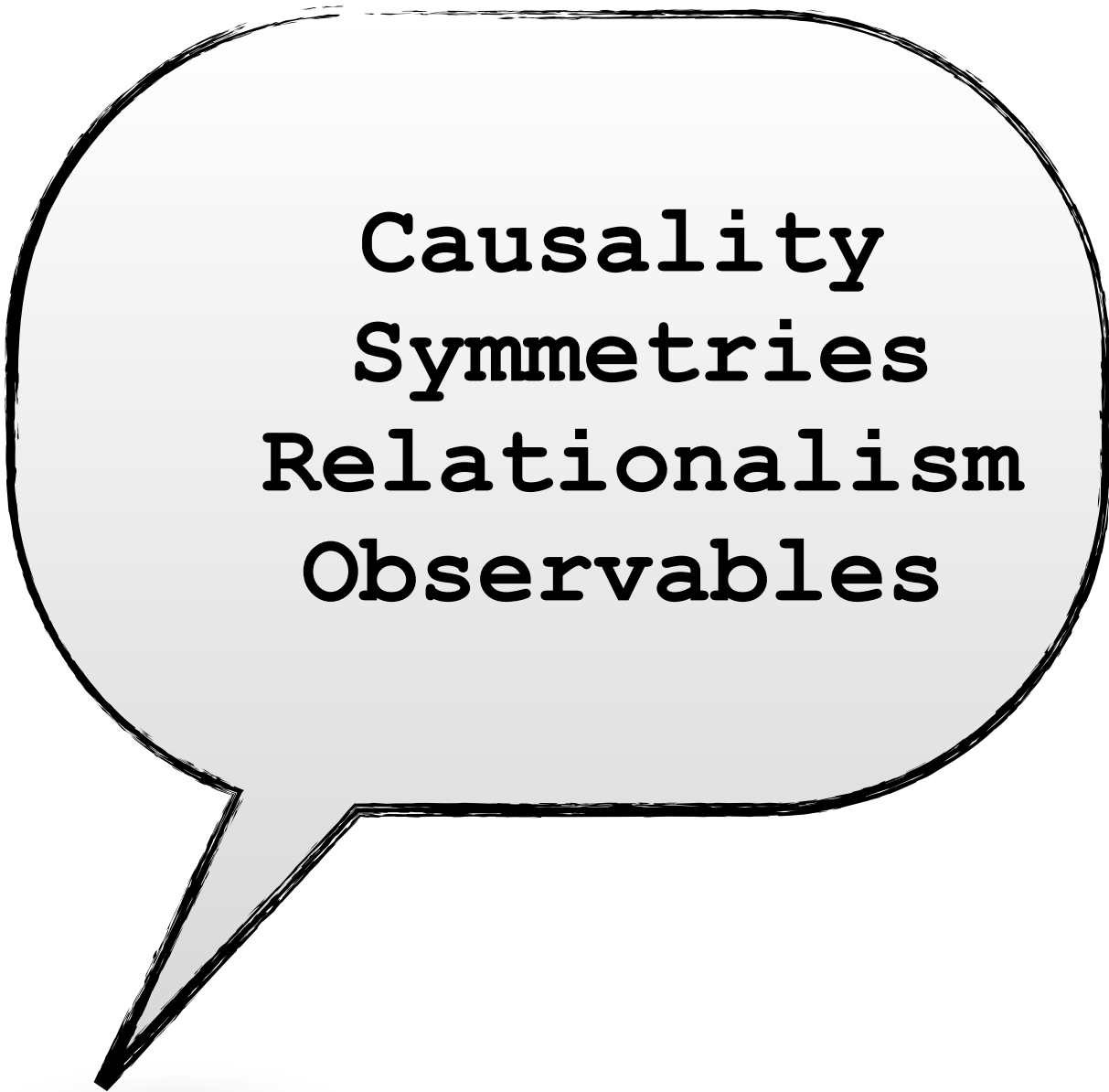
- 1.The program of Quantum Gravity
- 2.Group field theory cosmology
- 3.Emergent scalar field dynamics on curved spacetime from GFT:
background contribution
- 4.Emergent field theory on curved spacetime from GFT: scalar
perturbations

1. The program of (background independent) Quantum Gravity



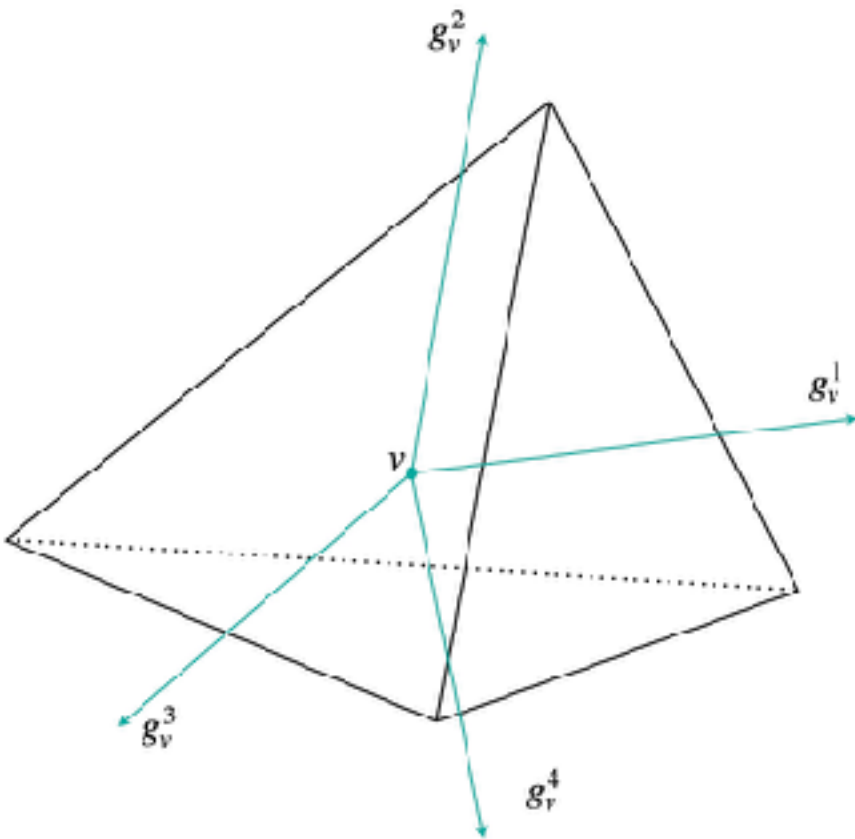
Causality
Symmetries
Relationalism
Observables

1. The program of (background independent) Quantum Gravity



Pre-geometric entities

Atoms of spacetime

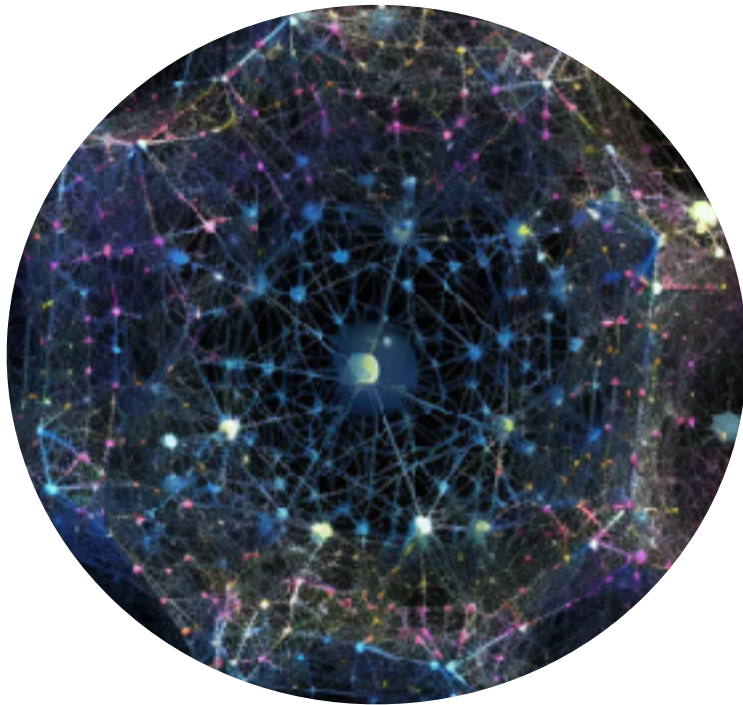


1. The program of (background independent) Quantum Gravity

Causality
Symmetries
Relationalism
Observables

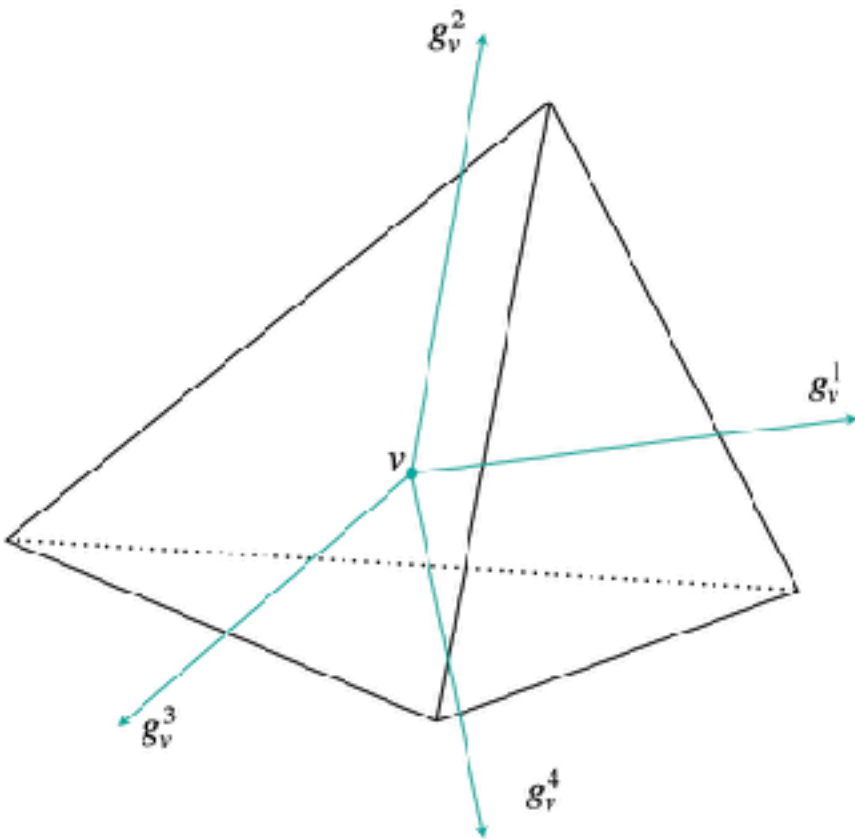
Quantum dynamics

Dynamical evolution and action principle



Pre-geometric entities

Atoms of spacetime



1. The program of (background independent) Quantum Gravity

Causality
Symmetries
Relationalism
Observables

Emergence of continuum
and classical spacetime

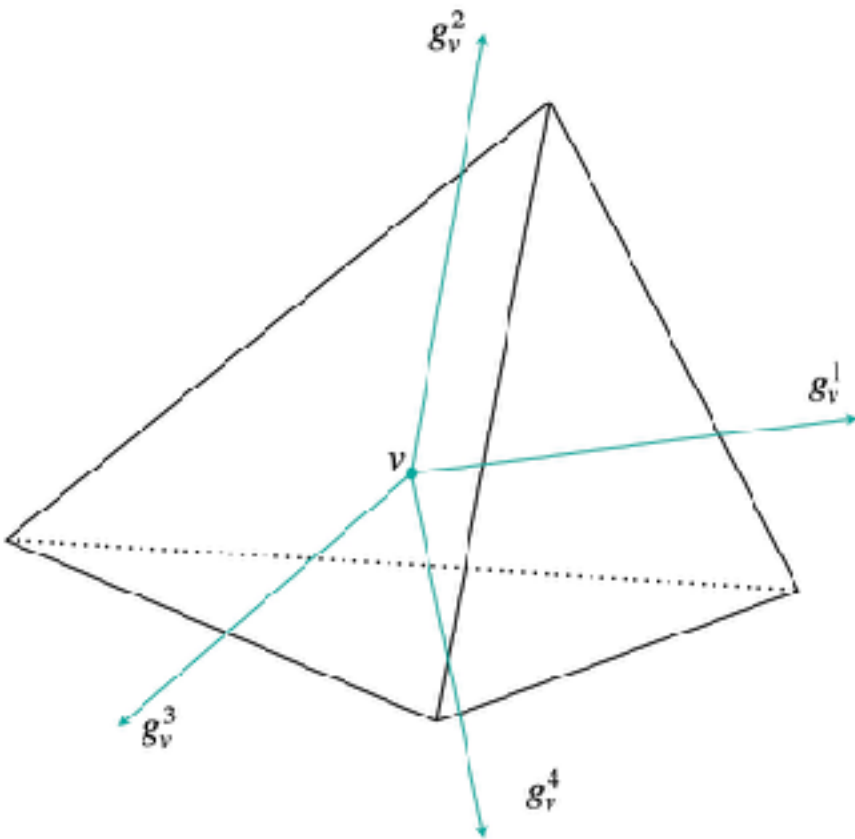
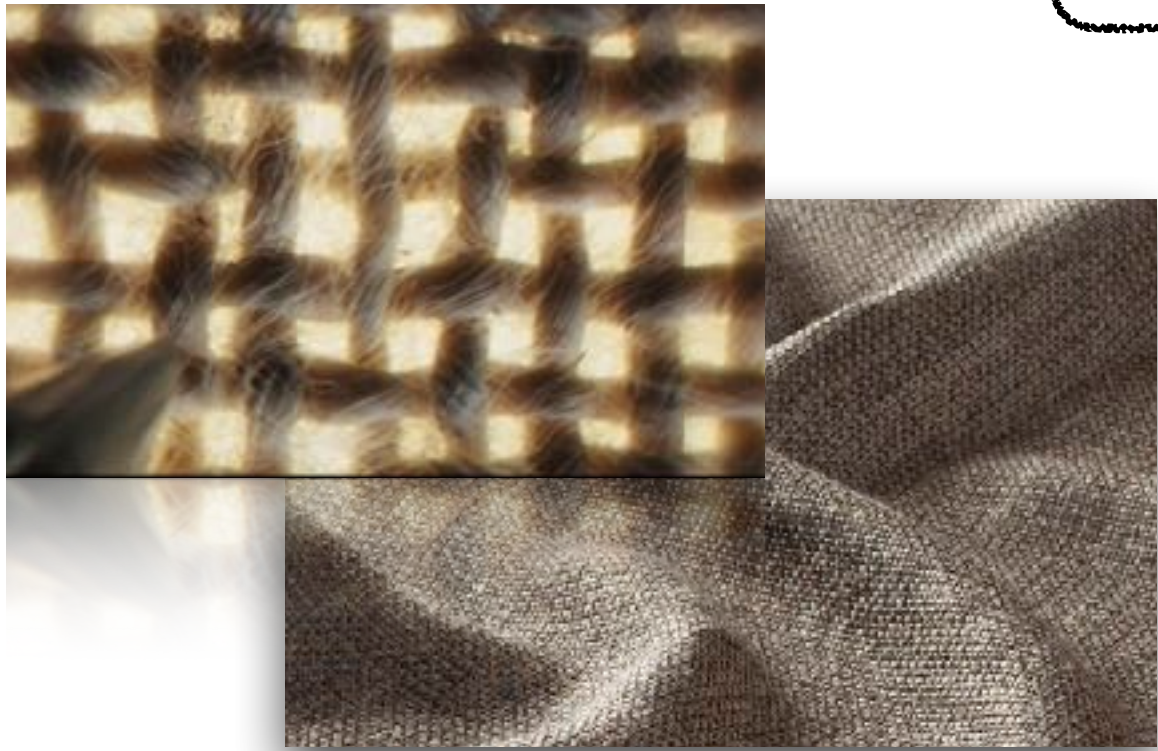
*Appropriate definition of continuum limit
and approximations to recover GR*

Quantum dynamics

*Dynamical evolution and action
principle*

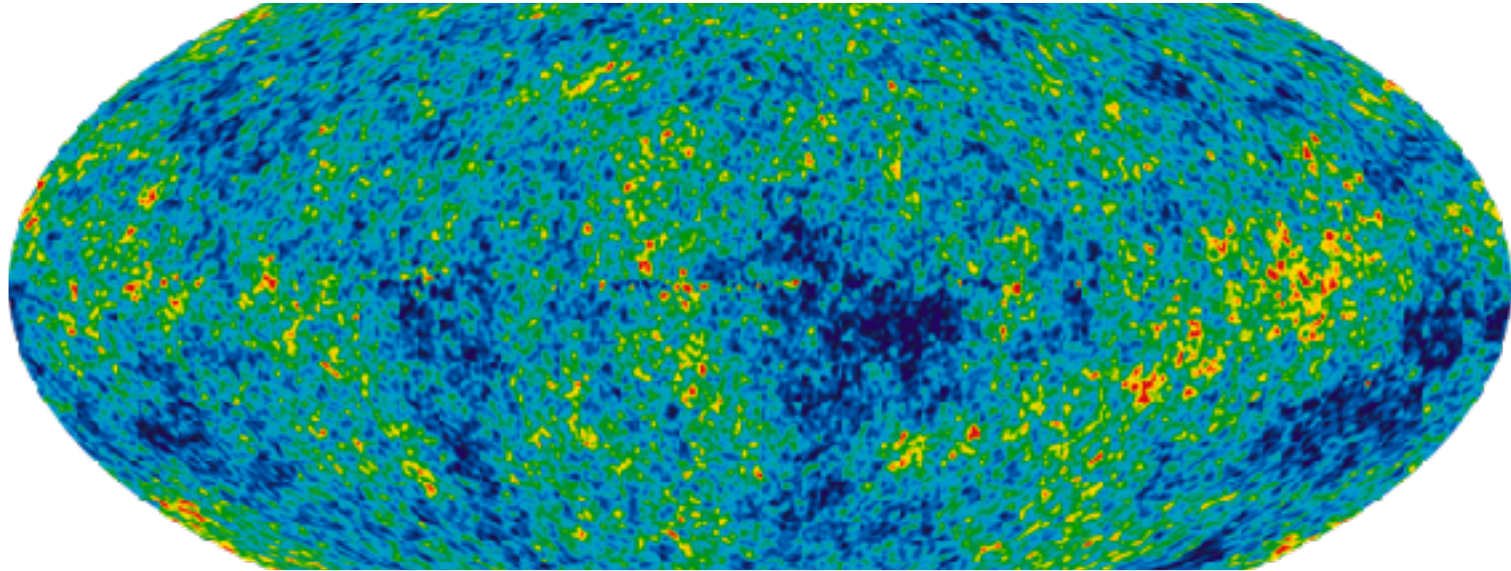
Pre-geometric entities

Atoms of spacetime



1. The program of (background independent) Quantum Gravity

Contact with observations
and providing answer to
open questions

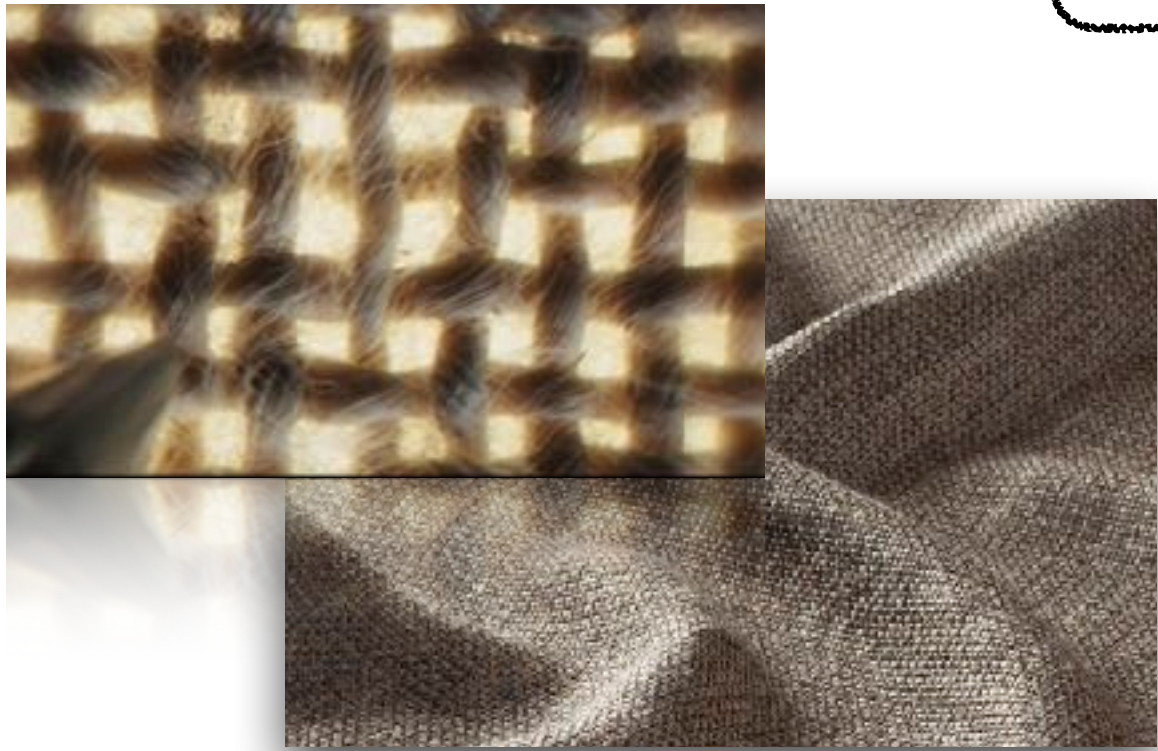


Causality
Symmetries
Relationalism
Observables

*Generating a dictionary between the
QG language
and the observational one*

Emergence of continuum
and classical spacetime

*Appropriate definition of continuum limit
and approximations to recover GR*



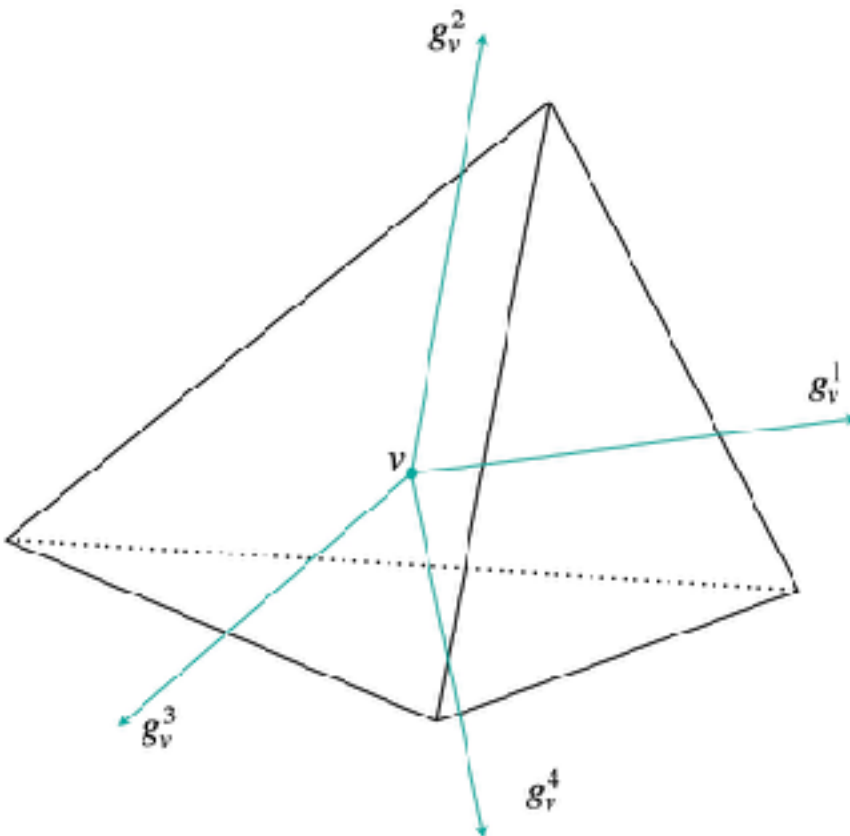
Quantum dynamics

*Dynamical evolution and action
principle*



Pre-geometric entities

Atoms of spacetime



2. Group field theories

Quantum field theory



Group field theories

2. Group field theories

Quantum field theory

Field theories *on* spacetime



Group field theories

2. Group field theories

Quantum field theory

Field theories *on* spacetime

Functions of the spacetime
coordinates



Group field theories

2. Group field theories

Quantum field theory

Field theories *on* spacetime

Functions of the spacetime
coordinates



Group field theories

Field theories *of* spacetime
atoms

2. Group field theories

Quantum field theory

Field theories *on* spacetime

Functions of the spacetime
coordinates

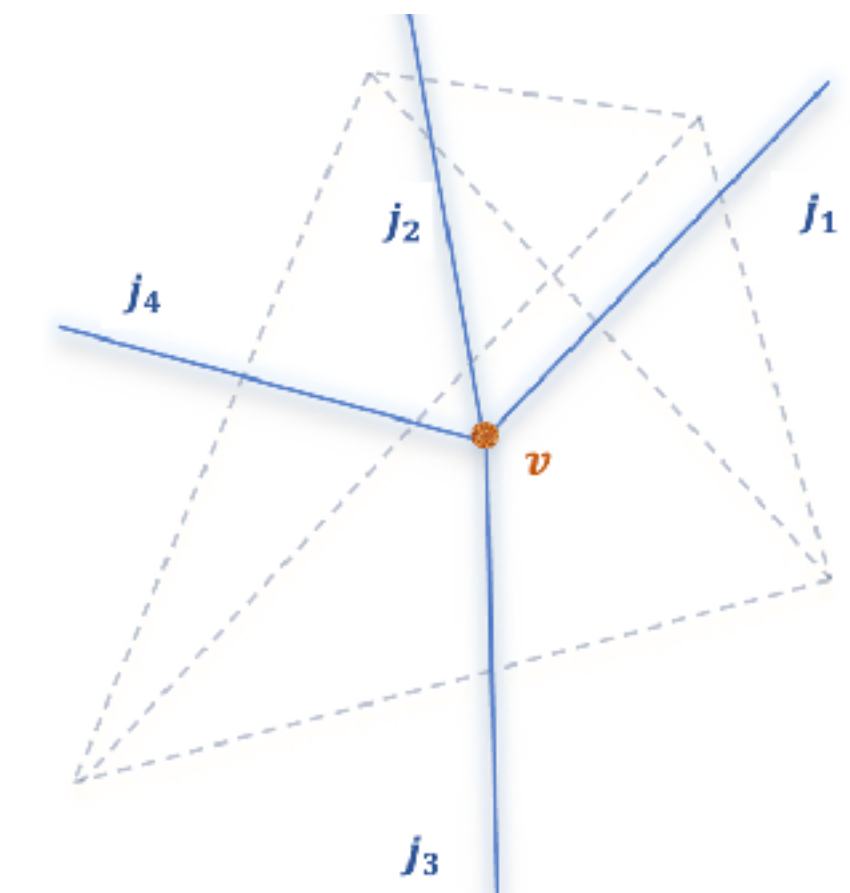


Group field theories

Field theories *of* spacetime
atoms

Functions of group elements that
give rise to spacetime

$$\begin{aligned}\varphi : G^4 &\rightarrow \mathbb{C} \\ (g_1, \dots, g_4) &\mapsto \varphi(g_1, \dots, g_4) \equiv \varphi(g_I) .\end{aligned}$$



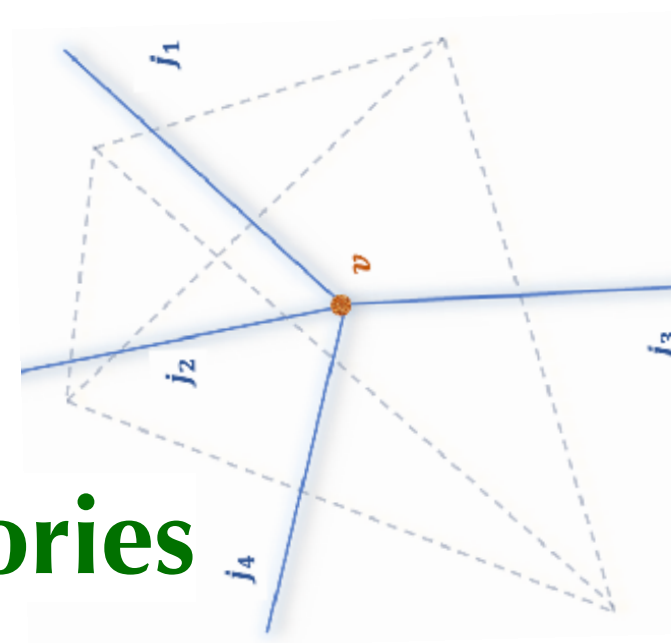
2. Group field theories

Quantum field theory

Field theories *on* spacetime

Functions of the spacetime coordinates

Action: interactions between particles on spacetime



Group field theories

Field theories *of* spacetime atoms

Functions of group elements that give rise to spacetime

$$\varphi : G^4 \rightarrow \mathbb{C}$$

$$(g_1, \dots, g_4) \mapsto \varphi(g_1, \dots, g_4) \equiv \varphi(g_I).$$

Action (non-local interactions): generates quantum geometries

$$\begin{aligned} S_{\text{GFT}}[\varphi^*, \varphi] &= \int [dg_I]^2 \varphi^*(g_I) \mathcal{K}(g_I, g'_I) \varphi(g'_I) \\ &+ \sum_i \frac{\lambda_i}{D_i} \int [dg_I]^{D_i} \varphi^*(g_{I_1}) \dots \mathcal{U}_i(g_{I_1}, \dots, g_{I_{D_i}}) \dots \varphi(g_{I_{D_i}}) \\ &+ \text{c.c.}, \end{aligned}$$

2. Group field theory cosmology: *do they have all the ingredients?*



2. Group field theory cosmology: *do they have all the ingredients?*



Pre-geometric entities

Atoms of spacetime

Causality: Group
 $SL(2, \mathbb{C})$, $SU(2)$, ...

2. Group field theory cosmology: *do they have all the ingredients?*



Quantum dynamics

Dynamical evolution and action principle

Relational evolution with respect to scalar matter fields (clock and rods)



Pre-geometric entities

Atoms of spacetime

Causality: Group $SL(2,\mathbb{C})$, $SU(2)$, ...

2. Group field theory cosmology: *do they have all the ingredients?*



Emergence of continuum
and classical spacetime

*Appropriate definition of continuum limit
and approximations to recover GR*

Effective: mean-field
approach

GFT condensate

Quantum dynamics

*Dynamical evolution and action
principle*

*Relational evolution with respect to scalar
matter fields (clock and rods)*



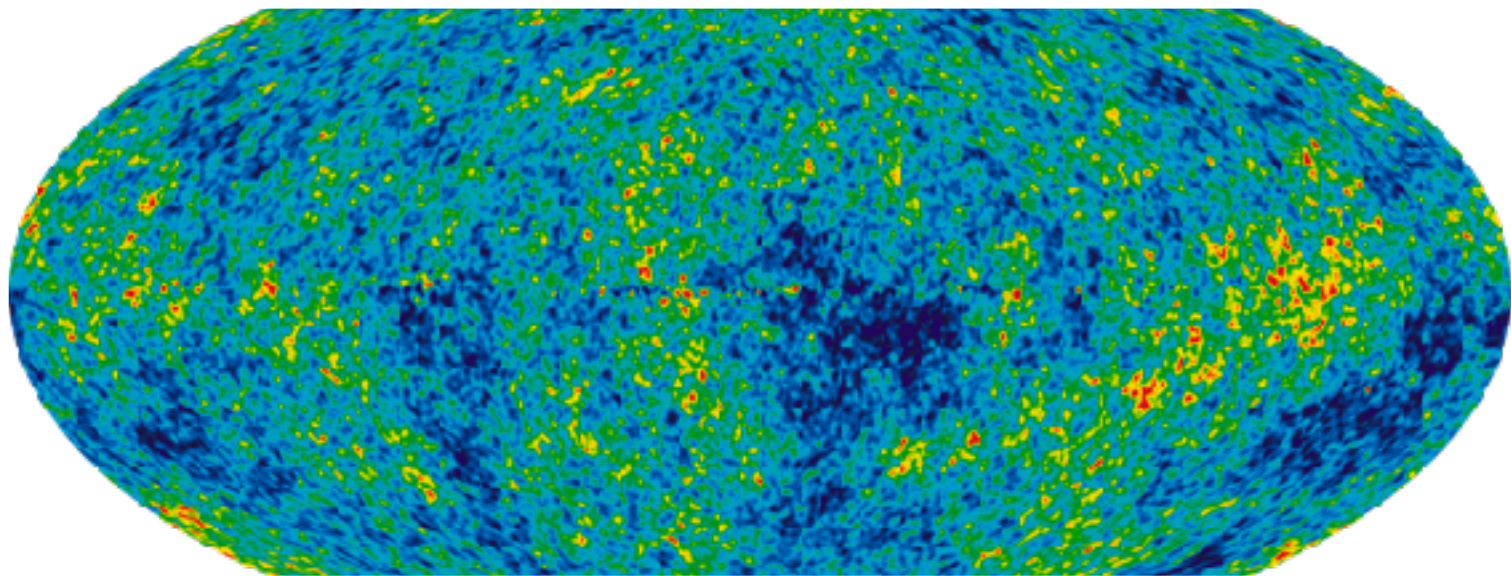
Pre-geometric entities

Atoms of spacetime

Causality: Group
 $SL(2,\mathbb{C})$, $SU(2)$, ...

2. Group field theory cosmology: *do they have all the ingredients?*

Contact with observations
and providing answers to
open questions



- Coupling to scalar field:
matter content of the
universe
- Scalar perturbations

*Generating a dictionary between the
QG language
and the observational one in
cosmology*

Emergence of continuum
and classical spacetime

Effective: mean-field
approach

GFT condensate

*Appropriate definition of continuum limit
and approximations to recover GR*

Quantum dynamics

*Relational evolution with respect to scalar
matter fields (clock and rods)*



*Dynamical evolution and action
principle*

Pre-geometric entities

Causality: Group
 $SL(2,\mathbb{C})$, $SU(2)$, ..

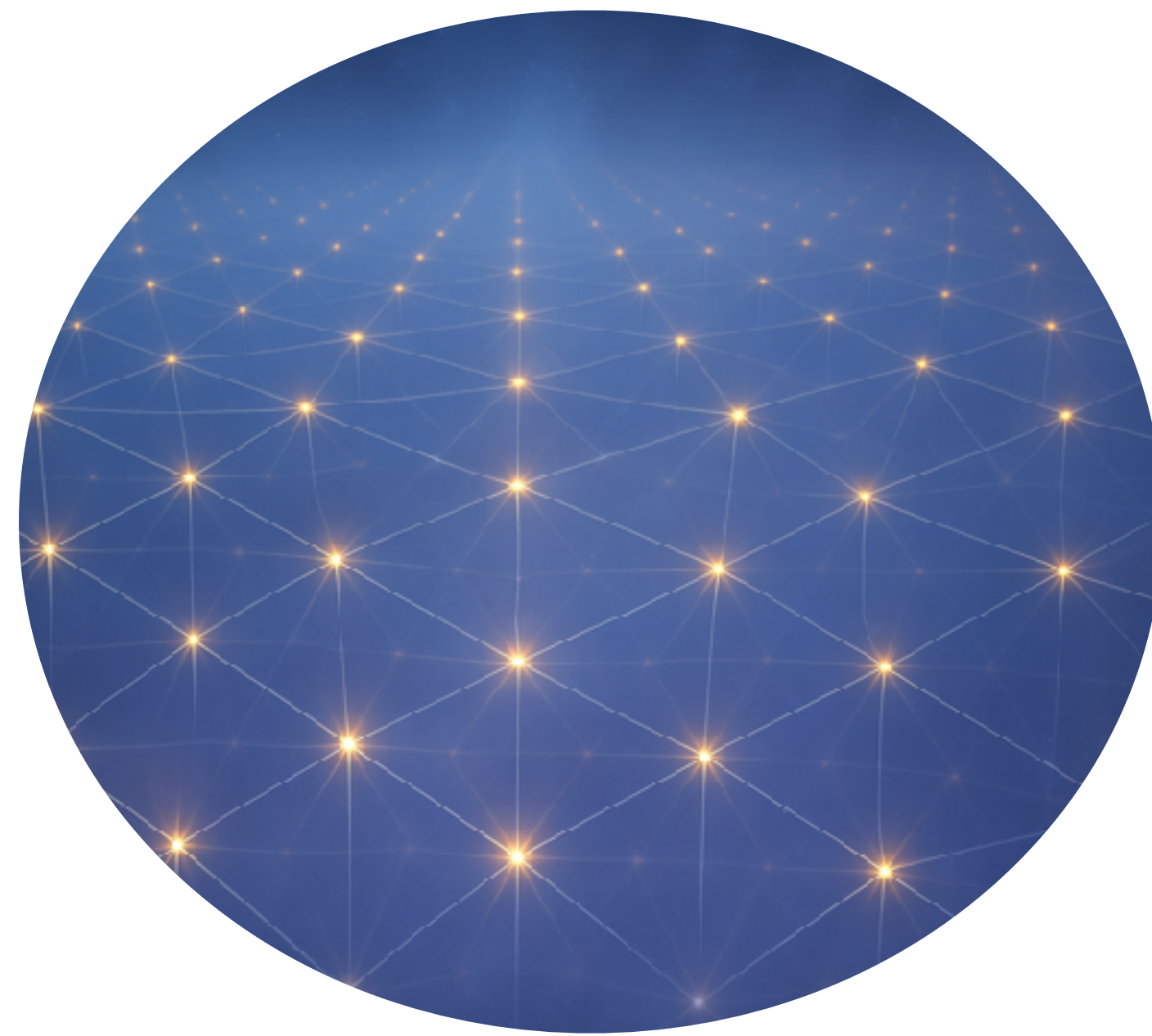
Atoms of spacetime



Ingredients of the GFT condensate (kinematics)

Classical set up

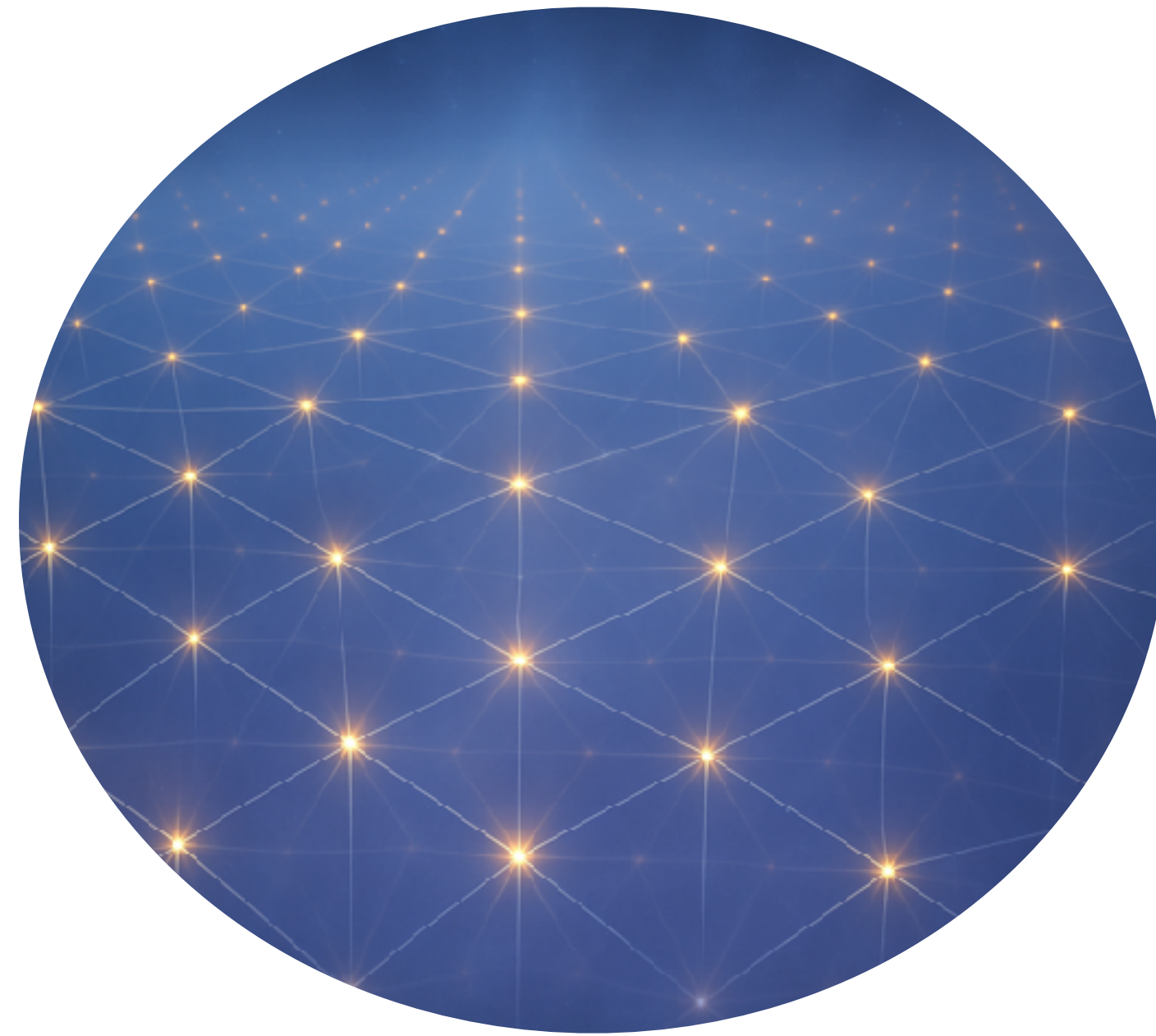
GFT condensate setting



Ingredients of the GFT condensate (kinematics)

Classical set up

Relational GR for
homogeneous isotropic
universe: FLRW line element



GFT condensate setting

Ingredients of the GFT condensate (kinematics)

Classical set up

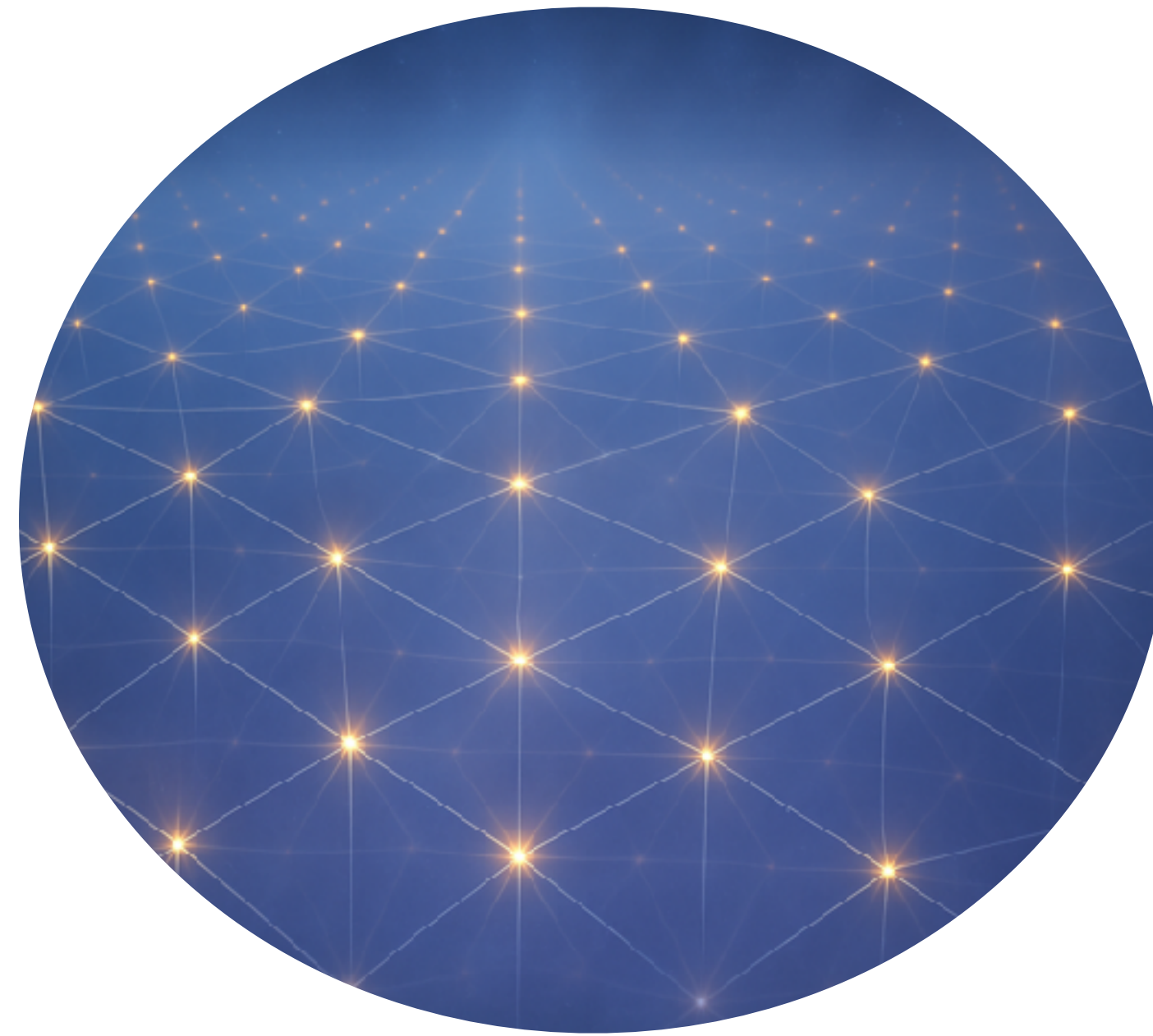
Relational GR for
homogeneous isotropic
universe: FLRW line element

Five **massless scalar fields**
minimally coupled to GR

Classical action for five
scalar fields

$$S = \frac{1}{2} \int d^4x \sqrt{-g} M_{\mu\nu}^{(\lambda)} g^{ab} \partial_a \chi^\mu \partial_b \chi^\nu$$
$$- \frac{\alpha_\phi}{2} \int d^4x \sqrt{-g} g^{ab} \partial_a \phi \partial_b \phi$$

GFT condensate setting



Ingredients of the GFT condensate (kinematics)

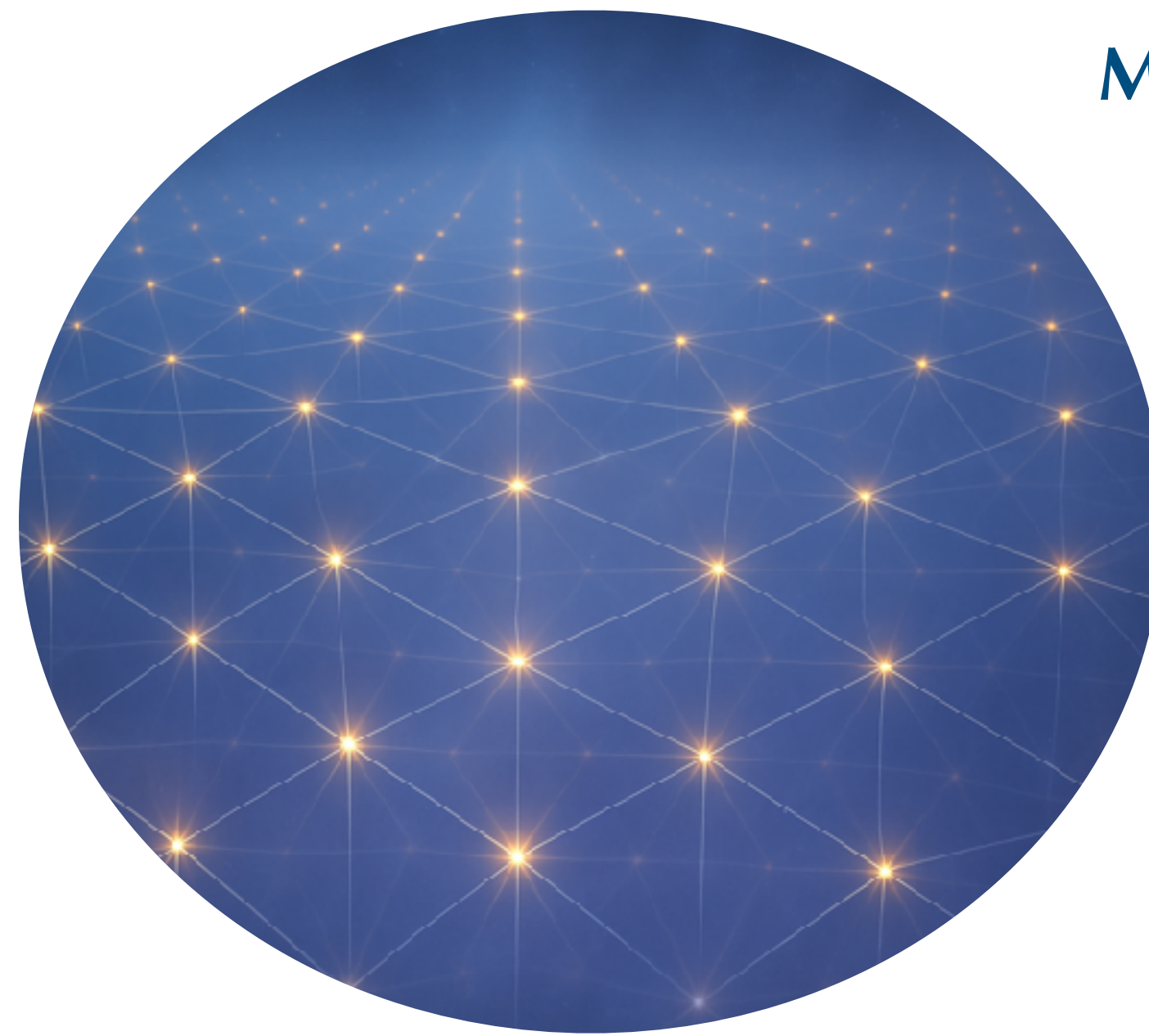
Classical set up

Relational GR for
homogeneous isotropic
universe: FLRW line element

Five **massless scalar fields**
minimally coupled to GR

Classical action for five
scalar fields

$$S = \frac{1}{2} \int d^4x \sqrt{-g} M_{\mu\nu}^{(\lambda)} g^{ab} \partial_a \chi^\mu \partial_b \chi^\nu \\ - \frac{\alpha_\phi}{2} \int d^4x \sqrt{-g} g^{ab} \partial_a \phi \partial_b \phi$$



GFT condensate setting

Matter fields

{ 4 scalar fields: clock and rods
 χ^μ
1 scalar field: matter content
of the universe ϕ



Ingredients of the GFT condensate (kinematics)

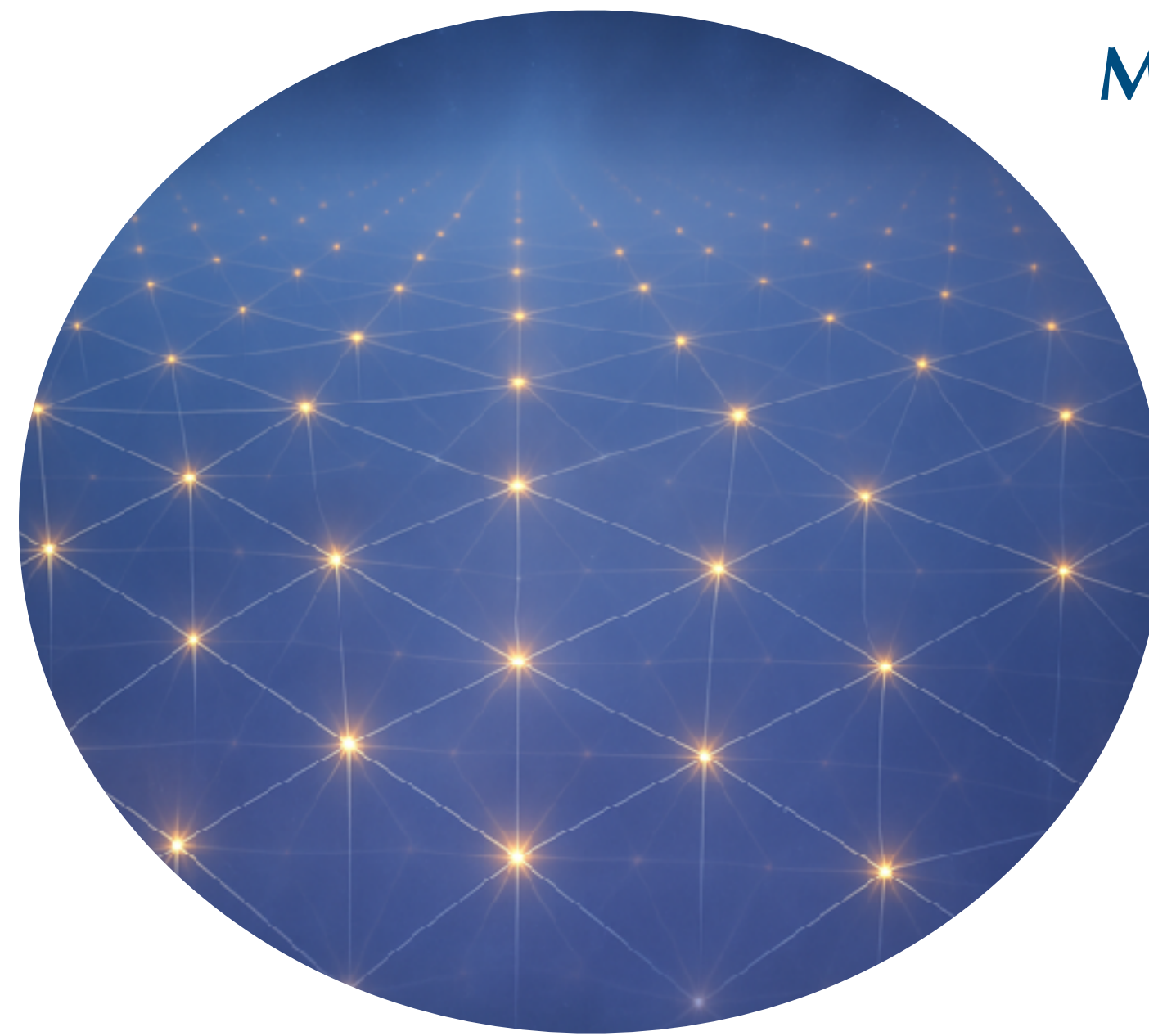
Classical set up

Relational GR for
homogeneous isotropic
universe: FLRW line element

Five **massless scalar fields**
minimally coupled to GR

Classical action for five
scalar fields

$$S = \frac{1}{2} \int d^4x \sqrt{-g} M_{\mu\nu}^{(\lambda)} g^{ab} \partial_a \chi^\mu \partial_b \chi^\nu \\ - \frac{\alpha_\phi}{2} \int d^4x \sqrt{-g} g^{ab} \partial_a \phi \partial_b \phi$$



GFT condensate setting

Matter fields

{ 4 scalar fields: clock and rods
 χ^μ
1 scalar field: matter content
of the universe ϕ

Isotropy

Dependence on a single spin j



Ingredients of the GFT condensate (kinematics)

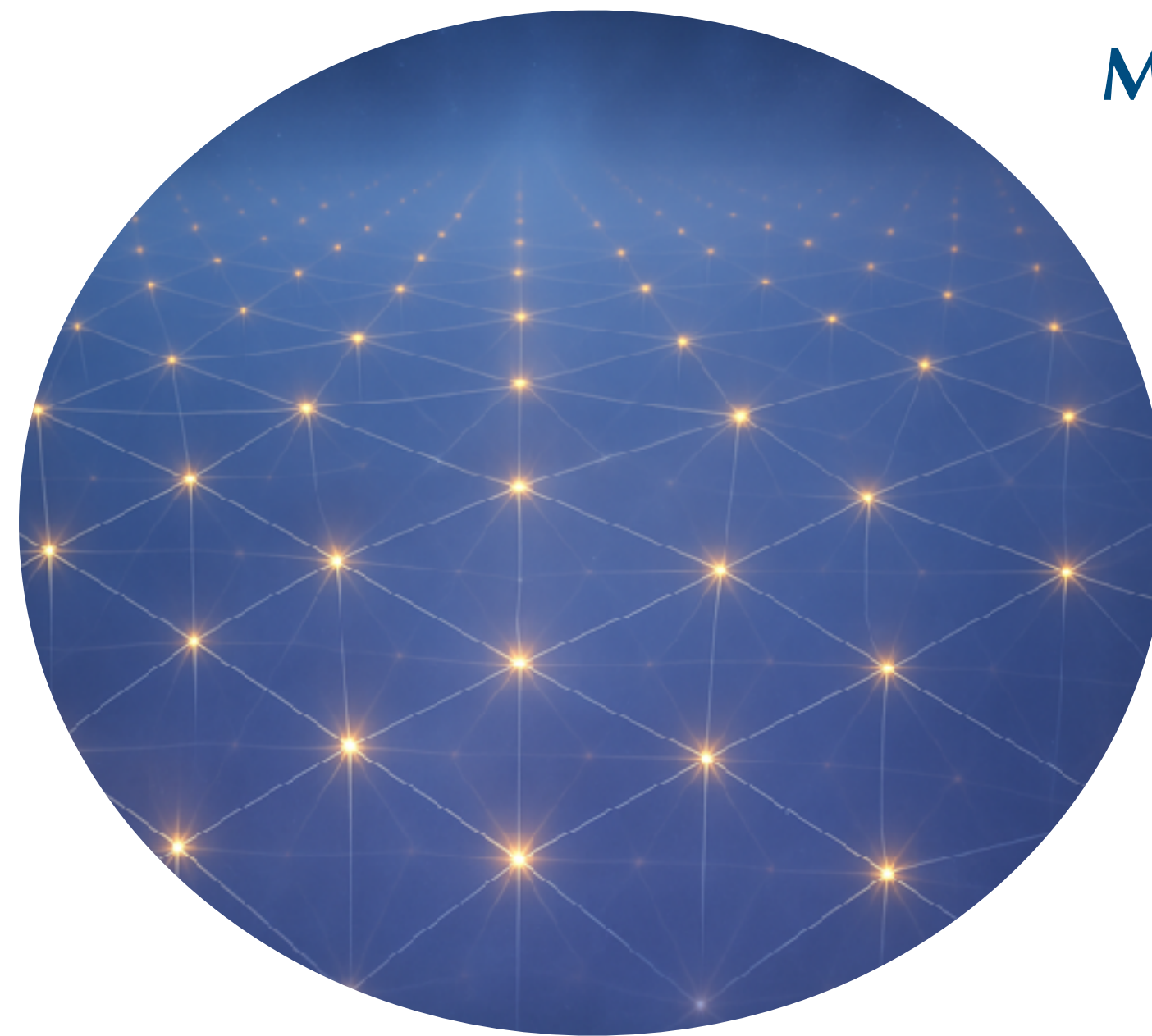
Classical set up

Relational GR for
homogeneous isotropic
universe: FLRW line element

Five **massless scalar fields**
minimally coupled to GR

Classical action for five
scalar fields

$$S = \frac{1}{2} \int d^4x \sqrt{-g} M_{\mu\nu}^{(\lambda)} g^{ab} \partial_a \chi^\mu \partial_b \chi^\nu \\ - \frac{\alpha_\phi}{2} \int d^4x \sqrt{-g} g^{ab} \partial_a \phi \partial_b \phi$$



GFT condensate setting

Matter fields { 4 scalar fields: clock and rods
 χ^μ
1 scalar field: matter content
of the universe ϕ

Isotropy Dependence on a single spin j

$\sigma(g_I, \chi^\mu, \phi)$: Wavefunction on minisuperspace

The domain of the condensate wavefunction $\simeq$
minisuperspace of homogeneous geometries



Ingredients of the GFT condensate (kinematics)

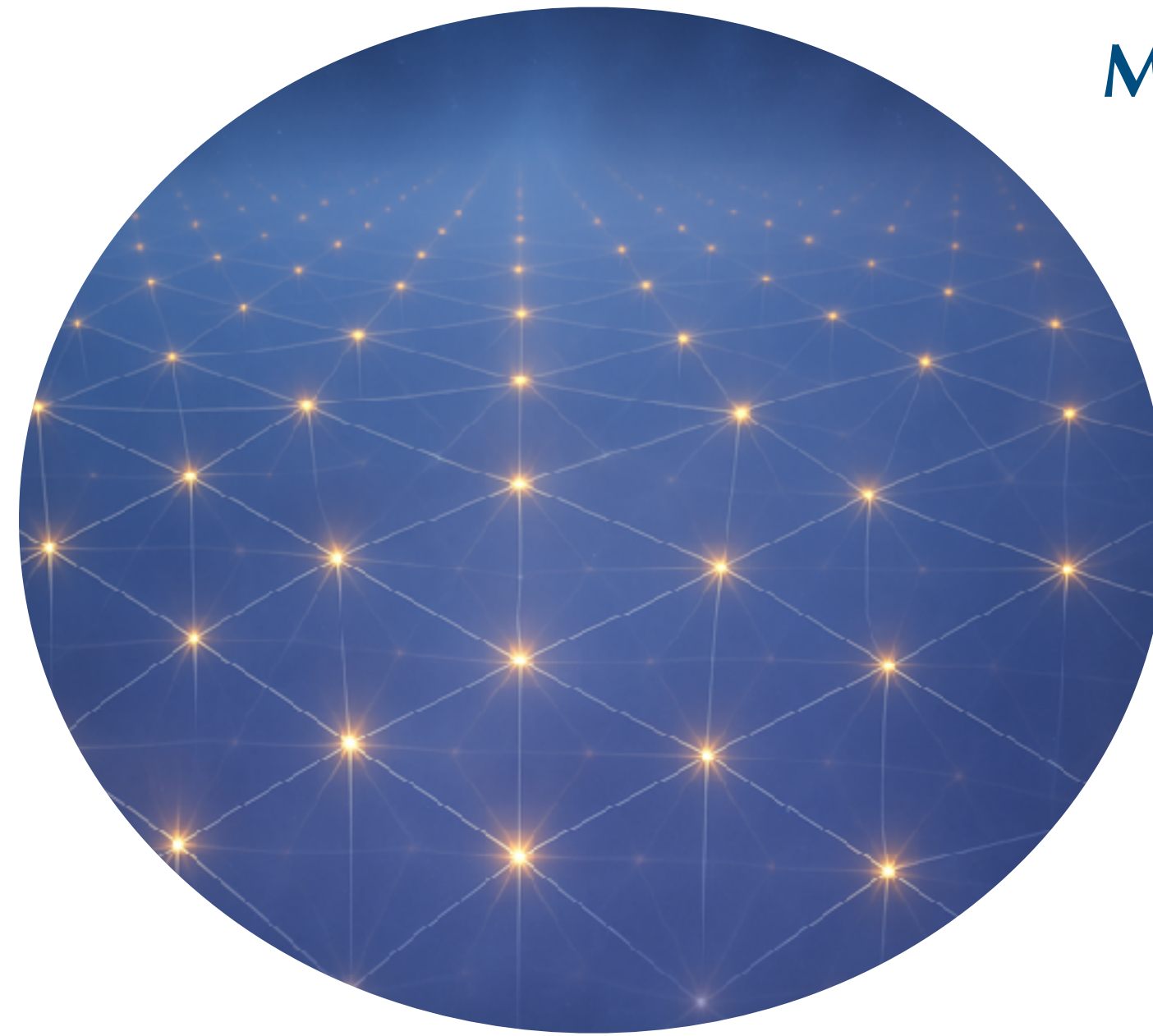
Classical set up

Relational GR for
homogeneous isotropic
universe: FLRW line element

Five **massless scalar fields**
minimally coupled to GR

Classical action for five
scalar fields

$$S = \frac{1}{2} \int d^4x \sqrt{-g} M_{\mu\nu}^{(\lambda)} g^{ab} \partial_a \chi^\mu \partial_b \chi^\nu - \frac{\alpha_\phi}{2} \int d^4x \sqrt{-g} g^{ab} \partial_a \phi \partial_b \phi$$



GFT condensate setting



Matter fields

4 scalar fields: clock and rods
 χ^μ
1 scalar field: matter content
of the universe ϕ

Isotropy

Dependence on a single spin j

$\sigma(g_I, \chi^\mu, \phi)$: Wavefunction on minisuperspace

The domain of the condensate wavefunction $\simeq$
minisuperspace of homogeneous geometries

Factorisation

$$\sigma_{\epsilon, \delta, \pi_0, \pi_x; \chi^\mu}(g_I, \chi^\mu, \phi) = \eta_\epsilon(\chi^0 - x^0; \pi_0) \eta_\delta(|\chi - \mathbf{x}|; \pi_x) \tilde{\sigma}(g_I, \chi^\mu, \phi)$$



Effective dynamics of GFT condensates (dynamics)

Effective dynamics of GFT condensates (dynamics)

- Extract effective hydrodynamic equations: **averaged** form of the equations of motion

$$\left\langle \sigma_{\epsilon^\mu}; x^\mu, \pi_\mu \left| \frac{\delta S_{\text{GFT}} [\hat{\varphi}, \hat{\varphi}^\dagger]}{\delta \hat{\varphi}^\dagger (g_I, x^\mu)} \right| \sigma_{\epsilon^\mu}; x^\mu, \pi_\mu \right\rangle = 0.$$

Effective dynamics of GFT condensates (dynamics)

- Extract effective hydrodynamic equations: **averaged** form of the equations of motion

$$\left\langle \sigma_{\epsilon^\mu; x^\mu, \pi_\mu} \left| \frac{\delta S_{\text{GFT}} [\hat{\varphi}, \hat{\varphi}^\dagger]}{\delta \hat{\varphi}^\dagger (g_I, x^\mu)} \right| \sigma_{\epsilon^\mu; x^\mu, \pi_\mu} \right\rangle = 0. \quad \partial_0^2 \tilde{\sigma}_j (x, \pi_\phi) - i\gamma \partial_0 \tilde{\sigma}_j (x, \pi_\phi) - E_j (\pi_\phi) \tilde{\sigma}_j (x, \pi_\phi) + \alpha^2 \nabla^2 \tilde{\sigma}_j (x, \pi_\phi) = 0$$
$$\tilde{\sigma}_j (x, \pi_\phi) = \rho (x, \pi_\phi) e^{i\theta_j(x, \pi_\phi)}$$

Effective dynamics of GFT condensates (dynamics)

- Extract effective hydrodynamic equations: **averaged** form of the equations of motion

$$\left\langle \sigma_{\epsilon^\mu; x^\mu, \pi_\mu} \left| \frac{\delta S_{\text{GFT}} [\hat{\phi}, \hat{\phi}^\dagger]}{\delta \hat{\phi}^\dagger (g_I, x^\mu)} \right| \sigma_{\epsilon^\mu; x^\mu, \pi_\mu} \right\rangle = 0. \quad \partial_0^2 \tilde{\sigma}_j (x, \pi_\phi) - i\gamma \partial_0 \tilde{\sigma}_j (x, \pi_\phi) - E_j (\pi_\phi) \tilde{\sigma}_j (x, \pi_\phi) + \alpha^2 \nabla^2 \tilde{\sigma}_j (x, \pi_\phi) = 0$$

$$\tilde{\sigma}_j (x, \pi_\phi) = \rho (x, \pi_\phi) e^{i\theta_j(x, \pi_\phi)}$$

- Physical entities of interest correspond to **hydrodynamic averages**: expectation values of basic operators

$$\langle \hat{\Phi} \rangle = \phi_0 = N_0 \partial_{\pi_\phi} \theta_0$$

$$\langle \hat{V} \rangle = V_j \rho_0^2 = a^3$$

Effective dynamics of GFT condensates (dynamics)

- Extract effective hydrodynamic equations: **averaged** form of the equations of motion

$$\left\langle \sigma_{\epsilon^\mu; x^\mu, \pi_\mu} \left| \frac{\delta S_{\text{GFT}} [\hat{\phi}, \hat{\phi}^\dagger]}{\delta \hat{\phi}^\dagger (g_I, x^\mu)} \right| \sigma_{\epsilon^\mu; x^\mu, \pi_\mu} \right\rangle = 0. \quad \partial_0^2 \tilde{\sigma}_j (x, \pi_\phi) - i\gamma \partial_0 \tilde{\sigma}_j (x, \pi_\phi) - E_j (\pi_\phi) \tilde{\sigma}_j (x, \pi_\phi) + \alpha^2 \nabla^2 \tilde{\sigma}_j (x, \pi_\phi) = 0$$

$$\tilde{\sigma}_j (x, \pi_\phi) = \rho (x, \pi_\phi) e^{i\theta_j(x, \pi_\phi)}$$

- Physical entities of interest correspond to **hydrodynamic averages**: expectation values of basic operators

$$\langle \hat{\Phi} \rangle = \phi_0 = N_0 \partial_{\pi_\phi} \theta_0$$

$$\langle \hat{V} \rangle = V_j \rho_0^2 = a^3$$

Perturbative framework with respect to the spatial gradient (slightly inhomogeneous relational quantities)

$$\rho_j = \rho_0 + \delta\rho, \quad \theta_j = \theta_0 + \delta\theta.$$

Effective dynamics of GFT condensates (dynamics)

- Extract effective hydrodynamic equations: **averaged** form of the equations of motion

$$\left\langle \sigma_{\epsilon^\mu}; x^\mu, \pi_\mu \left| \frac{\delta S_{\text{GFT}} [\hat{\phi}, \hat{\phi}^\dagger]}{\delta \hat{\phi}^\dagger (g_I, x^\mu)} \right| \sigma_{\epsilon^\mu}; x^\mu, \pi_\mu \right\rangle = 0. \quad \partial_0^2 \tilde{\sigma}_j (x, \pi_\phi) - i\gamma \partial_0 \tilde{\sigma}_j (x, \pi_\phi) - E_j (\pi_\phi) \tilde{\sigma}_j (x, \pi_\phi) + \alpha^2 \nabla^2 \tilde{\sigma}_j (x, \pi_\phi) = 0$$

$$\tilde{\sigma}_j (x, \pi_\phi) = \rho (x, \pi_\phi) e^{i\theta_j(x, \pi_\phi)}$$

- Physical entities of interest correspond to **hydrodynamic averages**: expectation values of basic operators

$$\langle \hat{\Phi} \rangle = \phi_0 = N_0 \partial_{\pi_\phi} \theta_0$$

$$\langle \hat{V} \rangle = V_j \rho_0^2 = a^3$$

Perturbative framework with respect to the spatial gradient (slightly inhomogeneous relational quantities)

$$\rho_j = \rho_0 + \delta\rho, \quad \theta_j = \theta_0 + \delta\theta.$$



Dynamics of the
background quantities



Dynamics of the
perturbed quantities

3. Emergent scalar field theory on curved spacetime from GFT

$$a^3 = \langle \hat{V} \rangle = V_j \rho_0^2$$

Background level

Evolution equation for the background
GFT phase

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2} \right) = 0$$

- Function of the clock and the scalar field momentum
- Minisuperspace

3. Emergent scalar field theory on curved spacetime from GFT

$$a^3 = \langle \hat{V} \rangle = V_j \rho_0^2$$

Background level

Evolution equation for the background
GFT phase

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2} \right) = 0$$

- Function of the clock and the scalar field momentum
- Minisuperspace

Expectation value of the scalar field
operator

$$\langle \hat{\Phi} \rangle = \phi_0 = N_0 \partial_{\pi_\phi} \theta_0$$

3. Emergent scalar field theory on curved spacetime from GFT

$$a^3 = \langle \hat{V} \rangle = V_j \rho_0^2$$

Background level

Evolution equation for the background
GFT phase

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2} \right) = 0$$

- Function of the clock and the scalar field momentum
- Minisuperspace

Expectation value of the scalar field
operator

$$\langle \hat{\Phi} \rangle = \phi_0 = N_0 \partial_{\pi_\phi} \theta_0$$

- Study the dynamics of the background homogeneous scalar field (necessary for the scalar perturbation treatment)

3. Emergent scalar field theory on curved spacetime from GFT

$$a^3 = \langle \hat{V} \rangle = V_j \rho_0^2$$

Background level

Evolution equation for the background
GFT phase

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2} \right) = 0$$

- Function of the clock and the scalar field momentum
 - Minisuperspace

Expectation value of the scalar field
operator

$$\langle \hat{\Phi} \rangle = \phi_0 = N_0 \partial_{\pi_\phi} \theta_0$$

- Study the dynamics of the background homogeneous scalar field (necessary for the scalar perturbation treatment)
- **Large volume regime (large GFT densities): this can be matched to GR results**

3. Emergent scalar field theory on curved spacetime from GFT

$$a^3 = \langle \hat{V} \rangle = V_j \rho_0^2$$

Background level

Evolution equation for the background
GFT phase

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2} \right) = 0$$

- Function of the clock and the scalar field momentum
 - Minisuperspace

Expectation value of the scalar field
operator

$$\langle \hat{\Phi} \rangle = \phi_0 = N_0 \partial_{\pi_\phi} \theta_0$$

- Study the dynamics of the background homogeneous scalar field (necessary for the scalar perturbation treatment)
- **Large volume regime (large GFT densities): this can be matched to GR results**

For the small volume regime: we have a quantum bounce replacing the classical singularity

3. Emergent scalar field theory on curved spacetime from GFT

Near the
bounce

Background level

Evolution equation for the background
GFT phase

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2} \right) = 0$$

- Function of the clock and the scalar field momentum

**Not observable: adequate
mathematical formulation!**

Expectation value of the scalar field
operator

$$\langle \hat{\Phi} \rangle = \phi_0 = N_0 \partial_{\pi_\phi} \theta_0$$

- The scalar field is homogeneous

3. Emergent scalar field theory on curved spacetime from GFT

Near the
bounce

Background level

Evolution equation for the background
GFT phase

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2} \right) = 0$$

- Function of the clock and the scalar field momentum

**Not observable: adequate
mathematical formulation!**

Expectation value of the scalar field
operator

$$\langle \hat{\Phi} \rangle = \phi_0 = N_0 \partial_{\pi_\phi} \theta_0$$

- The scalar field is homogeneous

**How do we access the dynamics of the
scalar field near the bounce ?**

Effective metric?

3. Emergent scalar field theory on curved spacetime from GFT

Near the
bounce

Background level

Evolution equation for the background
GFT phase

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2} \right) = 0$$

- Function of the clock and the scalar field momentum

**Not observable: adequate
mathematical formulation!**

Expectation value of the scalar field
operator

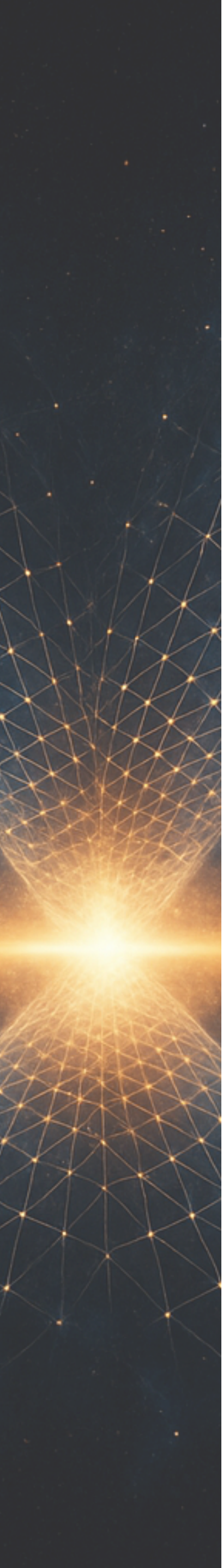
$$\langle \hat{\Phi} \rangle = \phi_0 = N_0 \partial_{\pi_\phi} \theta_0$$

- The scalar field is homogeneous

**How do we access the dynamics of the
scalar field near the bounce ?**

Effective metric?

**Closed equation of motion of the scalar
field with **no** condensate parametrisation**



Matter field dynamics near the bounce

Background level

Matter field dynamics near the bounce

Background level

1 Inversion relation: $\theta_0 = \left(a^3 f(\pi_\phi, x^0)\right)^{-1} \phi_0 + \frac{\gamma x^0}{2}$

2 Fundamental dynamics of the GFT phase:

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2}\right) = 0$$

Matter field dynamics near the bounce

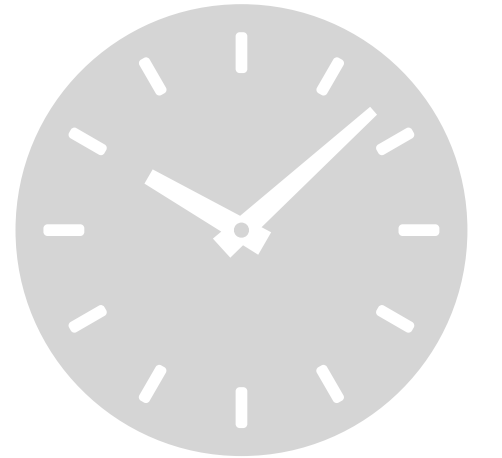
1 Inversion relation: $\theta_0 = \left(a^3 f(\pi_\phi, x^0)\right)^{-1} \phi_0 + \frac{\gamma x^0}{2}$

2 Fundamental dynamics of the GFT phase:

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2} \right) = 0$$

Background level

3 Clock is perfect

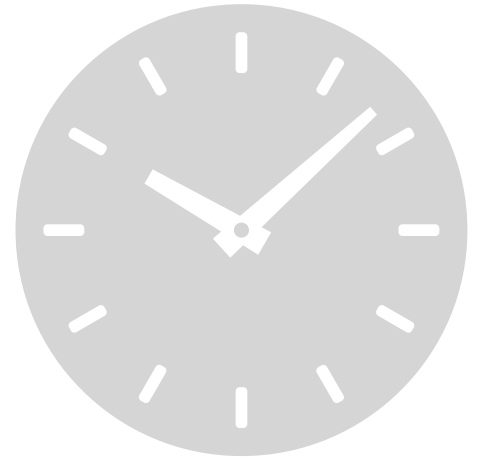


4 Assume an effective d'Alembert operator:

$$\square = \frac{1}{\sqrt{g}} \partial_\mu \left(\sqrt{g} g^{\mu\nu} \partial_\nu \phi_0 \right)$$

Matter field dynamics near the bounce

Background level



1 Inversion relation: $\theta_0 = \left(a^3 f(\pi_\phi, x^0)\right)^{-1} \phi_0 + \frac{\gamma x^0}{2}$

3 Clock is perfect

2 Fundamental dynamics of the GFT phase:

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2}\right) = 0$$

4 Assume an effective d'Alembert operator:

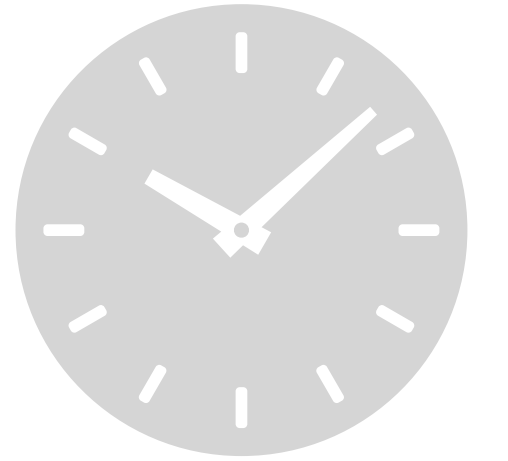
$$\square = \frac{1}{\sqrt{g}} \partial_\mu \left(\sqrt{g} g^{\mu\nu} \partial_\nu \phi_0 \right)$$

$\Rightarrow$ Closed evolution equation for the homogeneous scalar field

$$\ddot{\phi} - \left(3H + 2\frac{\dot{f}}{f}\right) \dot{\phi} + \left(3H\frac{\dot{f}}{f} - 3\dot{H} + 2\frac{\dot{f}^2}{f^2} - \frac{\ddot{f}}{f}\right) \phi = 0.$$

Matter field dynamics near the bounce

Background level



1 Inversion relation: $\theta_0 = \left(a^3 f(\pi_\phi, x^0)\right)^{-1} \phi_0 + \frac{\gamma x^0}{2}$

3 Clock is perfect

2 Fundamental dynamics of the GFT phase:

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2}\right) = 0$$

$$\square = \frac{1}{\sqrt{g}} \partial_\mu \left(\sqrt{g} g^{\mu\nu} \partial_\nu \phi_0 \right)$$

4 Assume an effective d'Alembert operator:

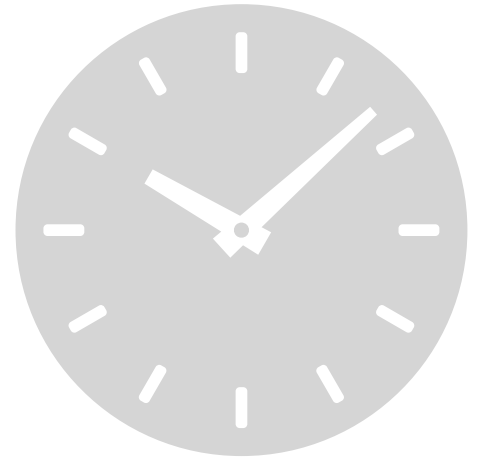
$\Rightarrow$ Closed evolution equation for the homogeneous scalar field

$$\ddot{\phi} - \left(3H + 2\frac{\dot{f}}{f}\right) \dot{\phi} + \left(3H\frac{\dot{f}}{f} - 3\dot{H} + 2\frac{\dot{f}^2}{f^2} - \frac{\ddot{f}}{f}\right) \phi = 0.$$

Read off the metric $g_{00} = Ca^{12}f^4$, $g_{ij} = a^2\delta_{ij}$

Matter field dynamics near the bounce

Background level



1 Inversion relation: $\theta_0 = \left(a^3 f(\pi_\phi, x^0)\right)^{-1} \phi_0 + \frac{\gamma x^0}{2}$

3 Clock is perfect

2 Fundamental dynamics of the GFT phase:

$$\ddot{\theta}_0 + \frac{3\dot{a}}{a} \left(\dot{\theta}_0 - \frac{\gamma}{2}\right) = 0$$

4 Assume an effective d'Alembert operator:

$$\square = \frac{1}{\sqrt{g}} \partial_\mu \left(\sqrt{g} g^{\mu\nu} \partial_\nu \phi_0 \right)$$

⇒ Closed evolution equation for the homogeneous scalar field

$$\ddot{\phi} - \left(3H + 2\frac{\dot{f}}{f}\right) \dot{\phi} + \left(3H\frac{\dot{f}}{f} - 3\dot{H} + 2\frac{\dot{f}^2}{f^2} - \frac{\ddot{f}}{f}\right) \phi = 0.$$

Read off the metric $g_{00} = Ca^{12}f^4$, $g_{ij} = a^2\delta_{ij}$

And additional terms ...

Modification with respect to the initial assumptions/classical system

What do we make of this?

Near the
bounce

$$\ddot{\phi}_0 - \left(3H + 2\frac{\dot{f}}{f} \right) \dot{\phi}_0 + g_{00} (R + \xi m^2) \phi_0 = 0.$$

□ **Non-minimal coupling:** standard field theory on curved spacetime tells us that such coupling to gravity occurs due to the natural presence of strong gravitational interaction between the geometrical degrees of freedom and those of matter.

What do we make of this?

Near the
bounce

$$\ddot{\phi}_0 - \left(3H + 2\frac{\dot{f}}{f} \right) \dot{\phi}_0 + g_{00} (R + \xi m^2) \phi_0 = 0.$$

❑ **Non-minimal coupling:** standard field theory on curved spacetime tells us that such coupling to gravity occurs due to the natural presence of strong gravitational interaction between the geometrical degrees of freedom and those of matter.

❑ **Emergent mass or an effective potential:**

- Our model naturally incorporates inflation driving the expansion of the universe.
- Chameleon mechanism: in the presence of other matter fields these scalars can acquire an effective mass parameter that is environmentally dependent.
-

What do we make of this?

Near the
bounce

$$\ddot{\phi}_0 - \left(3H + 2\frac{\dot{f}}{f} \right) \dot{\phi}_0 + g_{00} (R + \xi m^2) \phi_0 = 0.$$

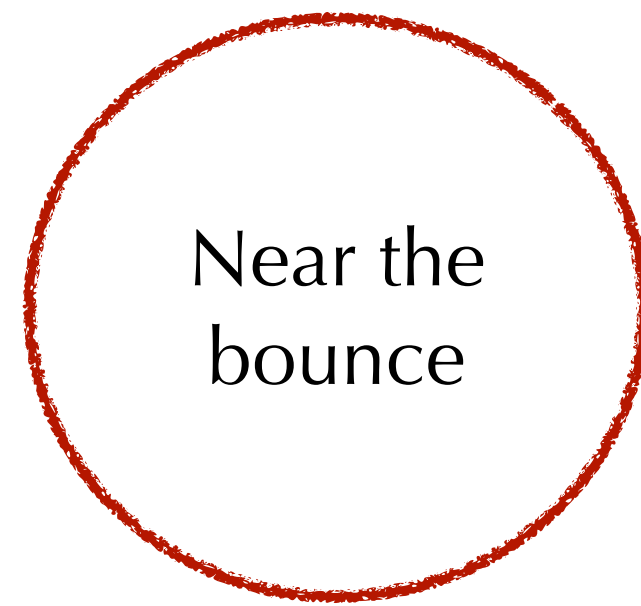
❑ **Non-minimal coupling:** standard field theory on curved spacetime tells us that such coupling to gravity occurs due to the natural presence of strong gravitational interaction between the geometrical degrees of freedom and those of matter.

❑ **Emergent mass or an effective potential:**

- Our model naturally incorporates inflation driving the expansion of the universe.
- Chameleon mechanism: in the presence of other matter fields these scalars can acquire an effective mass parameter that is environmentally dependent.
-

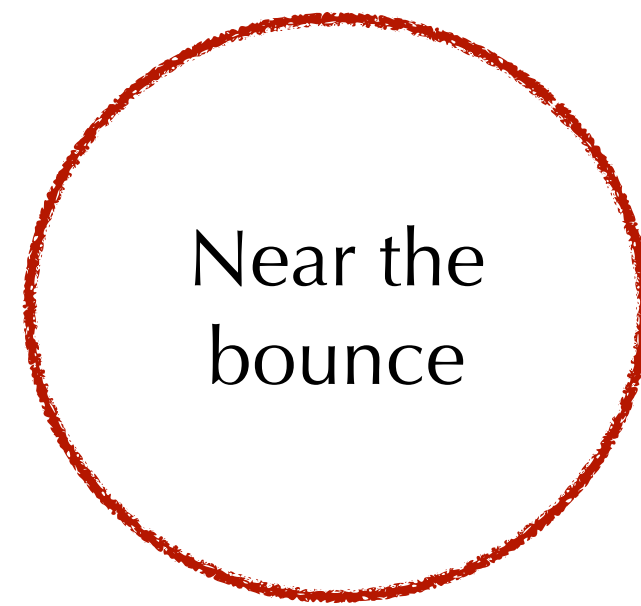
Modification with respect to classical system (**massless minimally coupled scalar field**)

4. Scalar perturbations: Prescription to extract a field theory on curved spacetime



The ingredients at hand:

4. Scalar perturbations: Prescription to extract a field theory on curved spacetime

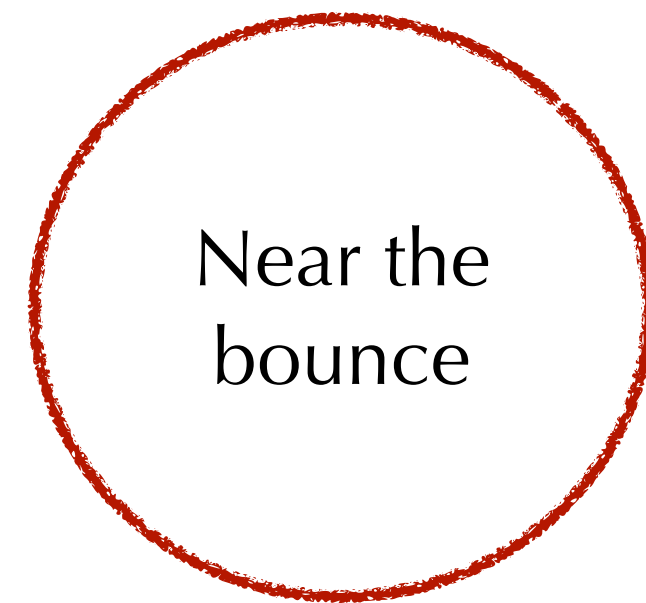


Background level

The ingredients at hand:

- ☒ Dynamics of the homogenous scalar field
- ☒ Background effective metric
- ☒ No condensate parametrisation entering the formulation

4. Scalar perturbations: Prescription to extract a field theory on curved spacetime



Background level

Perturbed level:

$$\rho_j = \rho_0 + \delta\rho, \quad \theta_j = \theta_0 + \delta\theta.$$

The ingredients at hand:

- ☒ Dynamics of the homogenous scalar field
- ☒ Background effective metric
- ☒ No condensate parametrisation entering the formulation

- ☐ Expression for the perturbed scalar field
- ☐ Dynamics of the perturbed GFT density and phase

4. Scalar perturbations: Prescription to extract a field theory on curved spacetime

Near the bounce

Background level

Perturbed level:

$$\rho_j = \rho_0 + \delta\rho, \quad \theta_j = \theta_0 + \delta\theta.$$

The ingredients at hand:

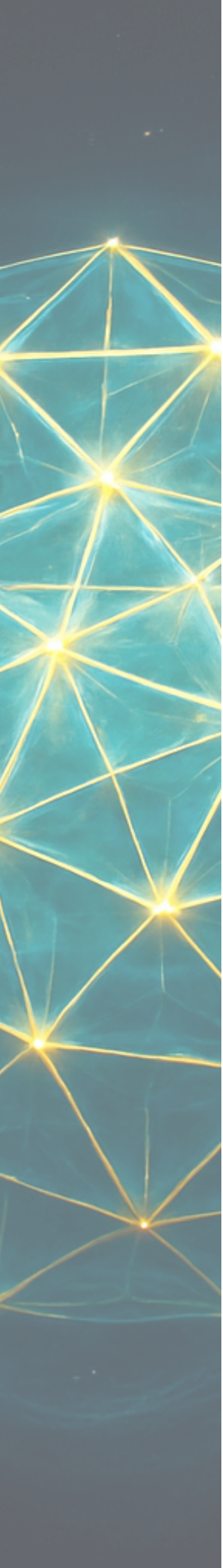
- ☒ Dynamics of the homogenous scalar field
- ☒ Background effective metric
- ☒ No condensate parametrisation entering the formulation

- ☐ Expression for the perturbed scalar field
- ☐ Dynamics of the perturbed GFT density and phase

Goal: Express the effective evolution equation of the perturbed scalar field as a field theory on curved spacetime

We need to get rid of the condensate parametrisation !

4. Scalar perturbations: Prescription to extract a FT on curved spacetime



4. Scalar perturbations: Prescription to extract a FT on curved spacetime

1. Express the coupled dynamics of the perturbed GFT phase and density in terms of only one of them

$$\delta\rho = \left(\tilde{\square} - \eta_j \right)^{-1} \mathcal{D}[\delta\theta]. \qquad \tilde{\square} \delta\theta + 2\dot{\delta\theta} \frac{\dot{\rho}_0}{\rho_0} + \mathcal{L} \left[\left(\tilde{\square} - \eta \right)^{-1} \mathcal{D}[\delta\theta] \right] = \alpha_r \nabla^2 \delta\theta.$$

4. Scalar perturbations: Prescription to extract a FT on curved spacetime

1. Express the coupled dynamics of the perturbed GFT phase and density in terms of only one of them

$$\delta\rho = \left(\tilde{\square} - \eta_j \right)^{-1} \mathcal{D}[\delta\theta]. \quad \tilde{\square} \delta\theta + 2\dot{\delta\theta} \frac{\dot{\rho}_0}{\rho_0} + \mathcal{L} \left[\left(\tilde{\square} - \eta \right)^{-1} \mathcal{D}[\delta\theta] \right] = \alpha_r \nabla^2 \delta\theta.$$

2. Find the solution to the above differential equation + derive an inversion relation

$$\delta\theta \equiv \left(\Psi(x^0, \pi_\phi) + \Phi(x, \pi_\phi) \right)^{-1} \delta\phi \quad \delta\phi = \delta\langle \hat{\Phi} \rangle_\sigma = \left[\frac{\delta N}{N_0} \phi_0 + N_0 \partial_{\pi_\phi} \delta\theta \right]_{\pi_\phi = \tilde{\pi}_\phi}$$

4. Scalar perturbations: Prescription to extract a FT on curved spacetime

1. Express the coupled dynamics of the perturbed GFT phase and density in terms of only one of them

$$\delta\rho = \left(\tilde{\square} - \eta_j \right)^{-1} \mathcal{D}[\delta\theta]. \quad \tilde{\square} \delta\theta + 2\dot{\delta\theta} \frac{\dot{\rho}_0}{\rho_0} + \mathcal{L} \left[\left(\tilde{\square} - \eta \right)^{-1} \mathcal{D}[\delta\theta] \right] = \alpha_r \nabla^2 \delta\theta.$$

2. Find the solution to the above differential equation + derive an inversion relation

$$\delta\theta \equiv \left(\Psi(x^0, \pi_\phi) + \Phi(x, \pi_\phi) \right)^{-1} \delta\phi \quad \delta\phi = \delta\langle \hat{\Phi} \rangle_\sigma = \left[\frac{\delta N}{N_0} \phi_0 + N_0 \partial_{\pi_\phi} \delta\theta \right]_{\pi_\phi = \tilde{\pi}_\phi}$$

➡ Using the inversion relation, extract the effective dynamics of the inhomogeneous scalar field:

$$\delta\ddot{\phi} - \left(2 \frac{\dot{\Psi} + \dot{\Phi}}{\Psi + \Phi} - \frac{\lambda_2}{\lambda_1} \right) \delta\dot{\phi} - \left(\frac{\ddot{\Psi} + \ddot{\Phi}}{\Psi + \Phi} - 2 \frac{(\dot{\Psi} + \dot{\Phi})^2}{(\Psi + \Phi)^2} + \frac{\lambda_2}{\lambda_1} \frac{\dot{\Psi} + \dot{\Phi}}{\Psi + \Phi} - \frac{\lambda_3}{\lambda_1} \right) \delta\phi = \frac{-\alpha_r}{\lambda_1} \nabla^2 \delta\phi.$$

4. Scalar perturbations: Prescription to extract a FT on curved spacetime

1. Express the coupled dynamics of the perturbed GFT phase and density in terms of only one of them

$$\delta\rho = \left(\tilde{\square} - \eta_j \right)^{-1} \mathcal{D}[\delta\theta]. \quad \tilde{\square} \delta\theta + 2\dot{\delta\theta} \frac{\dot{\rho}_0}{\rho_0} + \mathcal{L} \left[\left(\tilde{\square} - \eta \right)^{-1} \mathcal{D}[\delta\theta] \right] = \alpha_r \nabla^2 \delta\theta.$$

2. Find the solution to the above differential equation + derive an inversion relation

$$\delta\theta \equiv \left(\Psi(x^0, \pi_\phi) + \Phi(x, \pi_\phi) \right)^{-1} \delta\phi \quad \delta\phi = \delta\langle \hat{\Phi} \rangle_\sigma = \left[\frac{\delta N}{N_0} \phi_0 + N_0 \partial_{\pi_\phi} \delta\theta \right]_{\pi_\phi = \tilde{\pi}_\phi}$$

➡ Using the inversion relation, extract the effective dynamics of the inhomogeneous scalar field:

$$\delta\ddot{\phi} - \left(2 \frac{\dot{\Psi} + \dot{\Phi}}{\Psi + \Phi} - \frac{\lambda_2}{\lambda_1} \right) \delta\dot{\phi} - \left(\frac{\ddot{\Psi} + \ddot{\Phi}}{\Psi + \Phi} - 2 \frac{(\dot{\Psi} + \dot{\Phi})^2}{(\Psi + \Phi)^2} + \frac{\lambda_2}{\lambda_1} \frac{\dot{\Psi} + \dot{\Phi}}{\Psi + \Phi} - \frac{\lambda_3}{\lambda_1} \right) \delta\phi = \frac{-\alpha_r}{\lambda_1} \nabla^2 \delta\phi.$$

● Modified dispersion relation

● Extract the Bardeen Potentials from the effective dynamics

Indications of possible phenomenology

What did we achieve?

- ☑ Starting from a full quantum gravity setting: we extracted an effective scalar field theory on a curved spacetime.
- ☑ At early times, this resulted in the emergence of possibly:
 - Non-minimal coupling to gravity at early times —> Modified theory of gravity
 - Mass/matter or potential term that can be studied in the inflationary scenario or Chameleon mechanism
- ☑ The ground now is set for phenomenology and contact with observations!

Open questions and outlook

- ❑ Identify modified gravity theories that are produced from the effective dynamics outlined above .
- ❑ Classify such theories according to the QG parameters.
- ❑ **Phenomenological investigations:** Study the Inflationary potential, Mass term (chameleon mech.), the modified dispersion relation (dissipation effects) ..
- ❑ **Canonical quantization of the scalar field** : study all phenomena that a standard QFT theory produces: Correlation functions, CMB spectrum —> observational cosmology.

Thank you for your attention!