

# A Critical Look at the Higgs

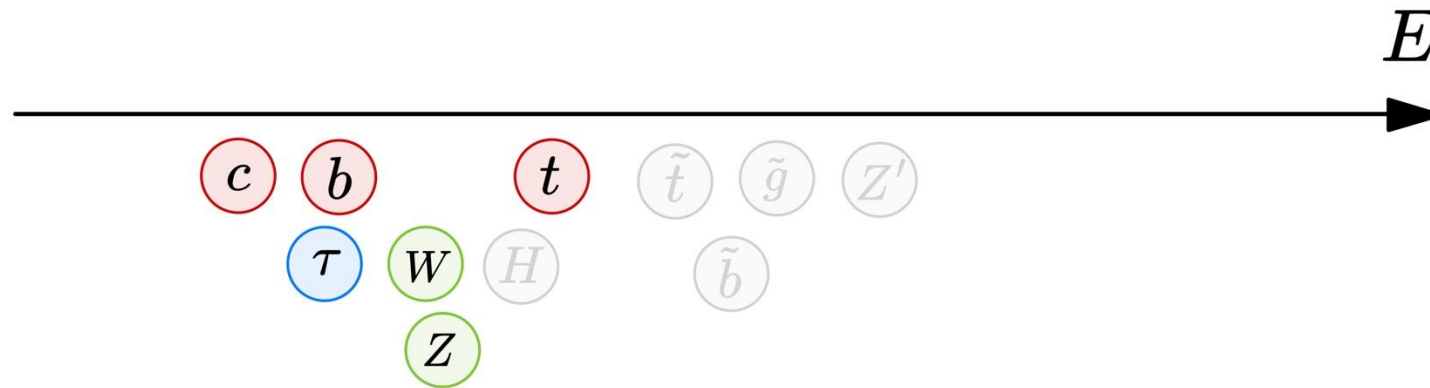
Based on 2412.03542

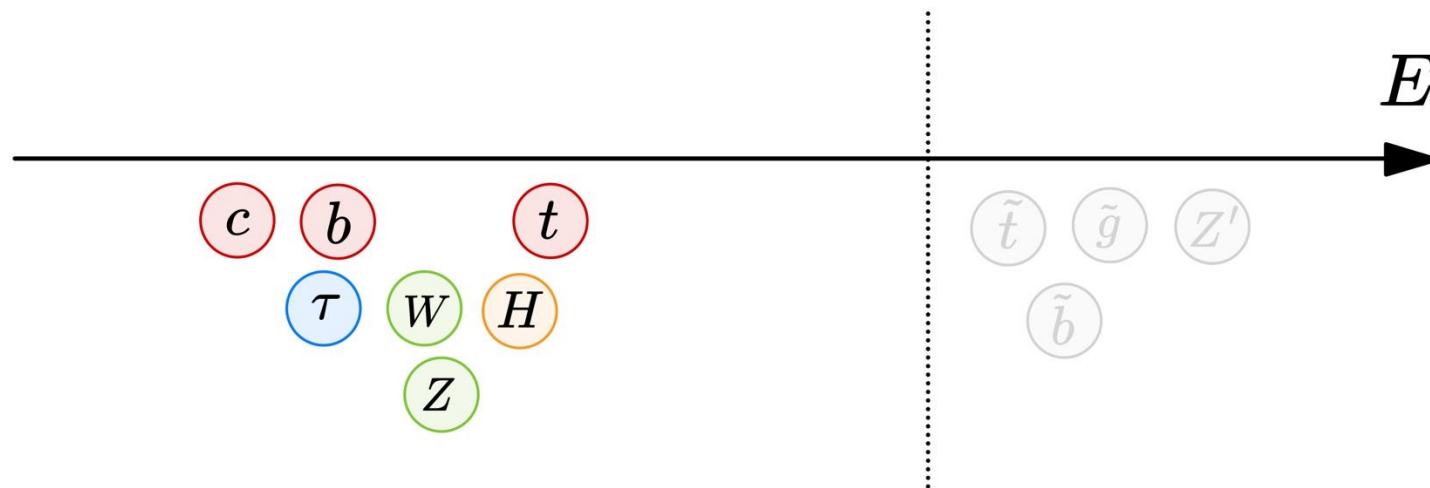
Maximilian Detering

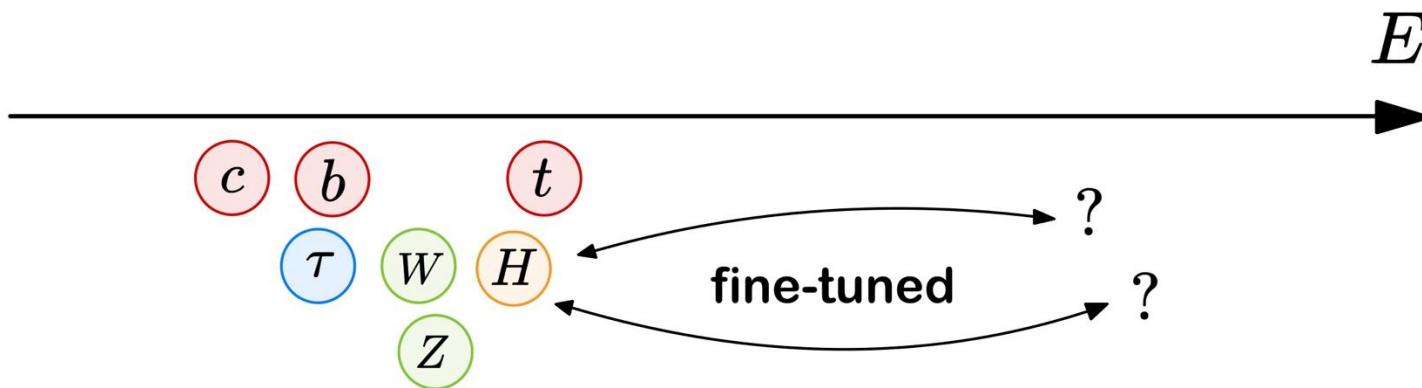
*New Approaches to Naturalness*

Institut de Physique des 2 Infinis de Lyon, France

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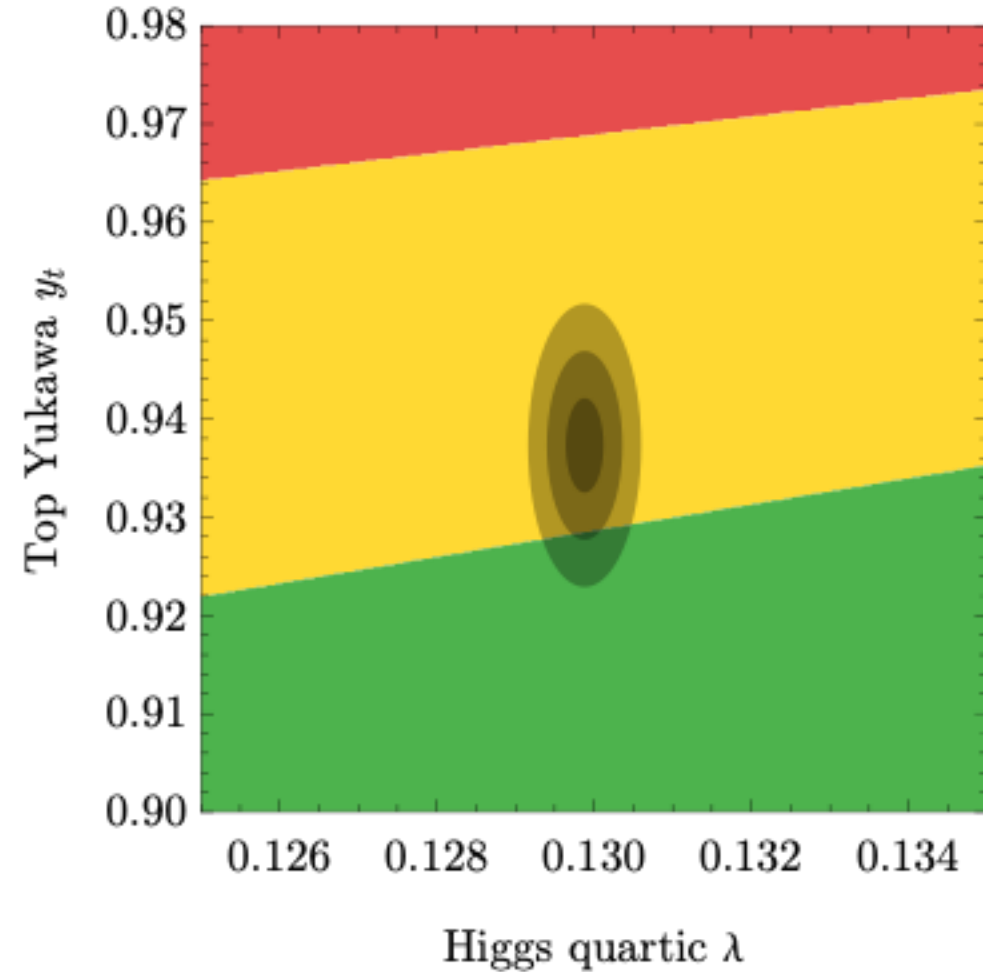
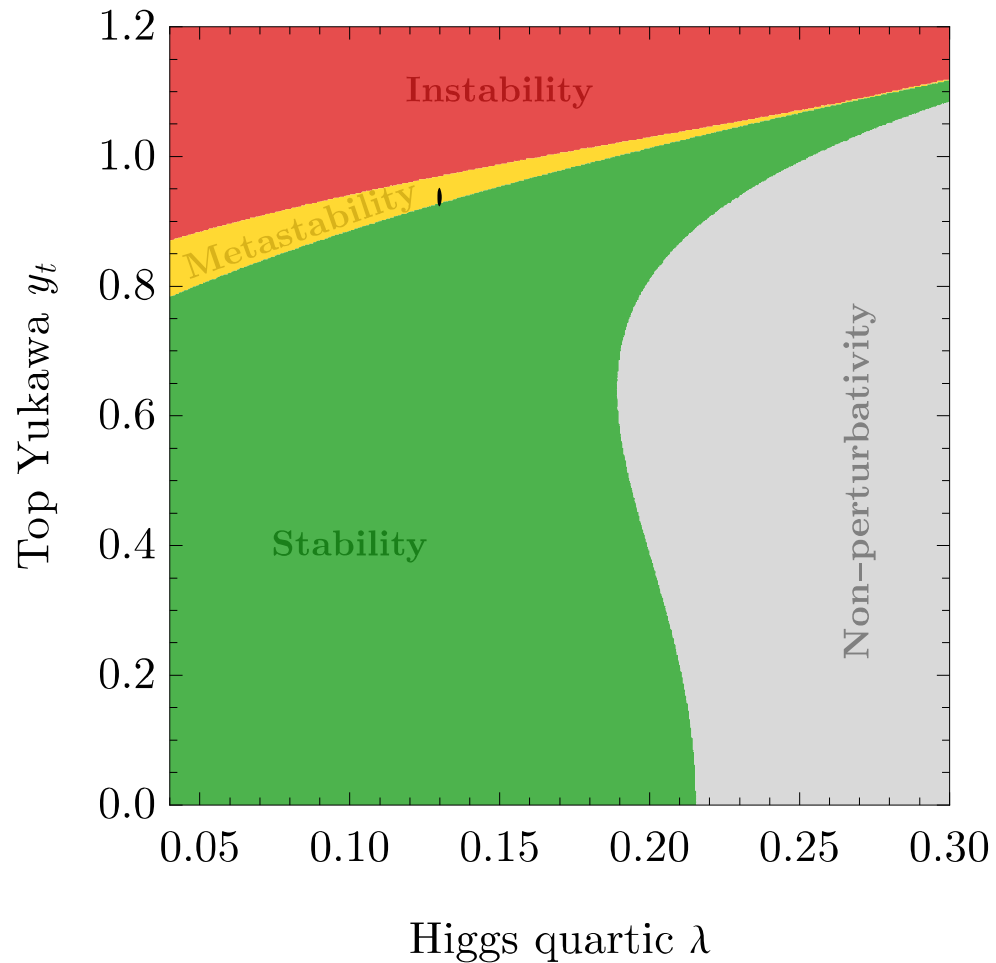
*Higgs mass not a  
fundamental parameter but  
special value explained as  
critical point in parameter  
space*

# Criticality

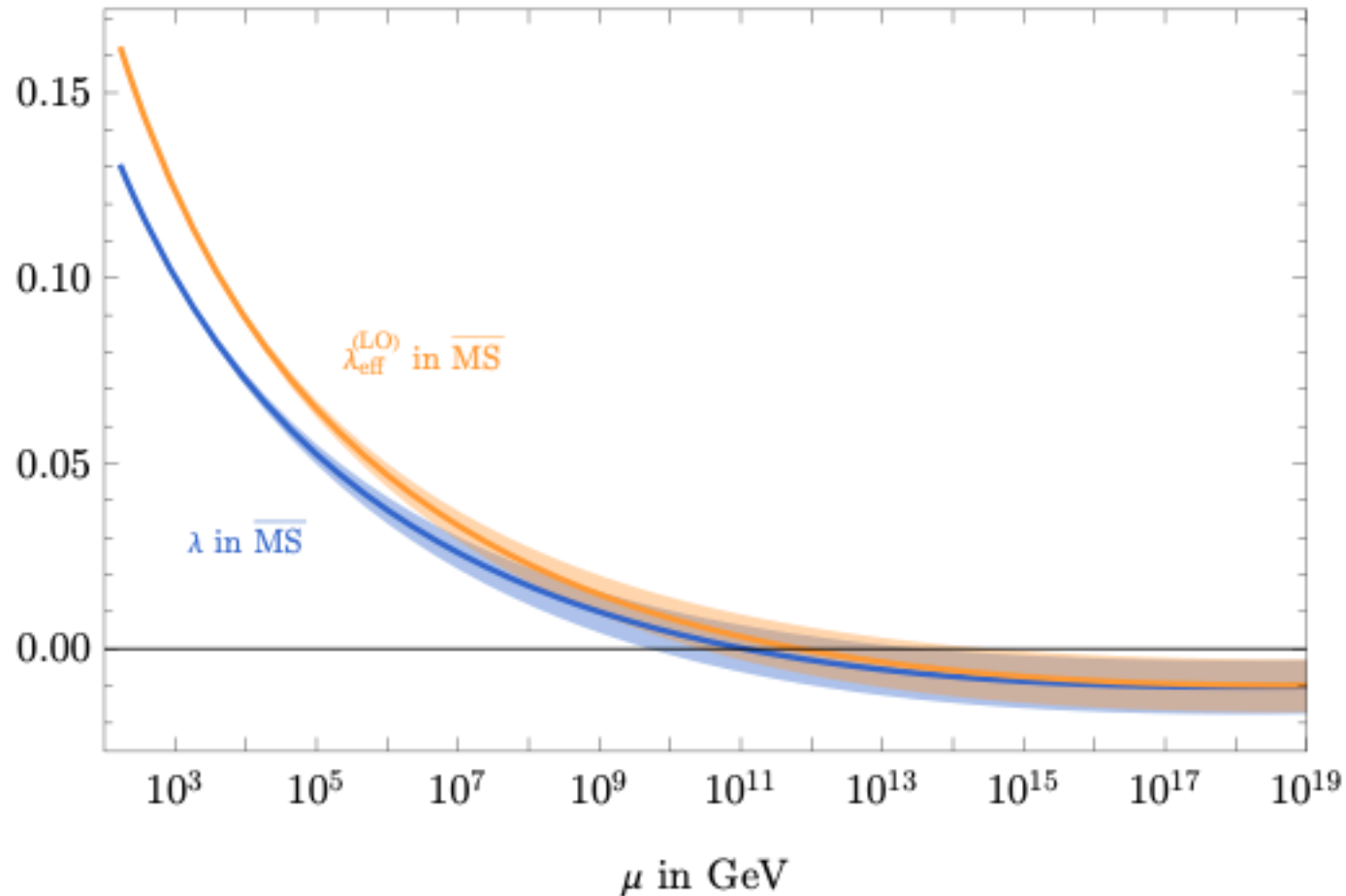
- **Classical (thermal)** phase transitions with varying thermodynamical parameters
- **Quantum** Phase Transition through change of external parameters
- **Critical values** of the parameters mark the transition
- Unconventional in particle physics context but **common in dynamical systems**

# Phases of the Standard Model

see also Buttazzo et al. '13, Degrassi et al. '12



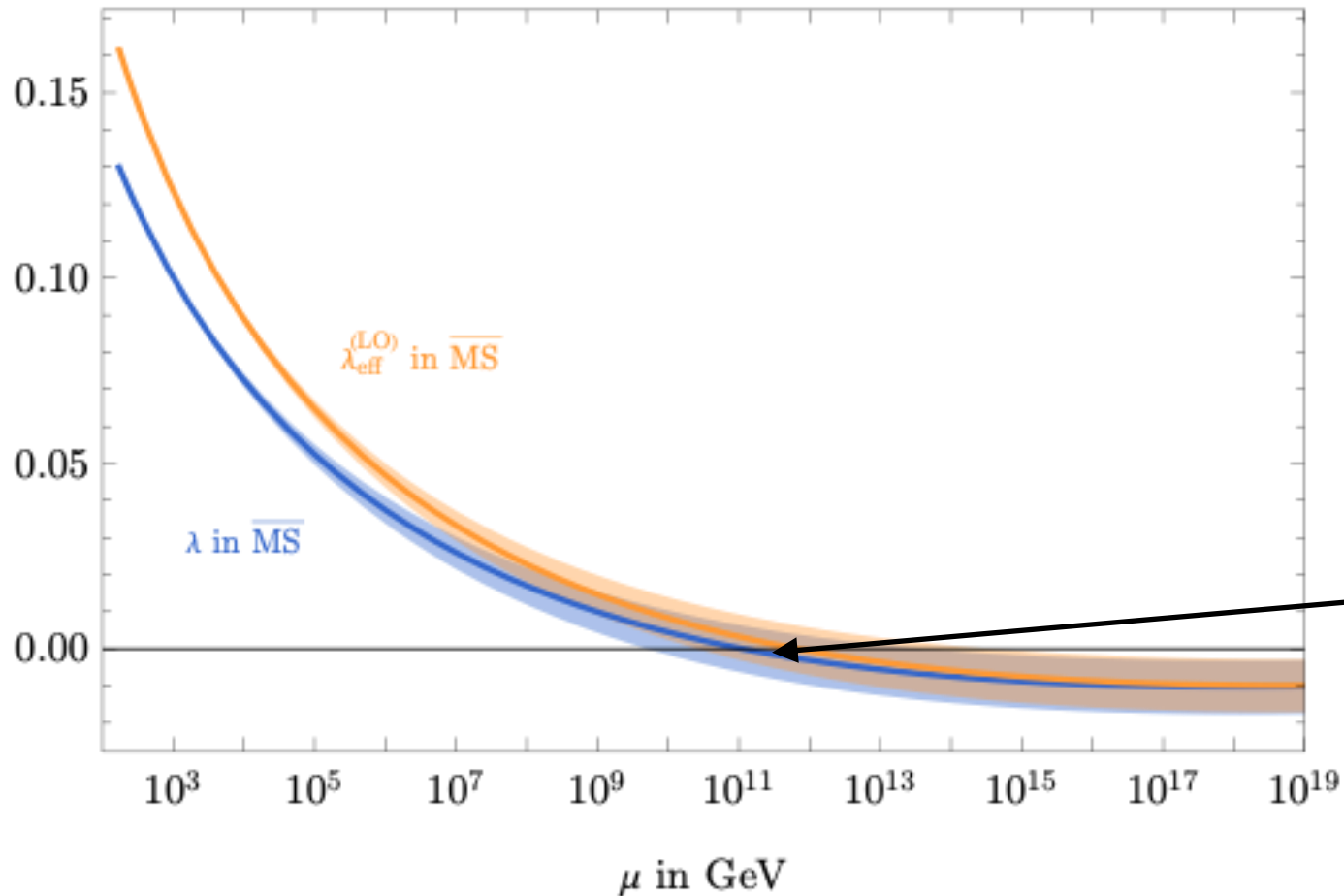
# Electroweak Vacuum Metastability



$$\beta_{\lambda}^{(1)} \sim \lambda(\lambda + y_t^2 + \text{gauge terms}) - y_t^4 + \text{gauge terms}$$



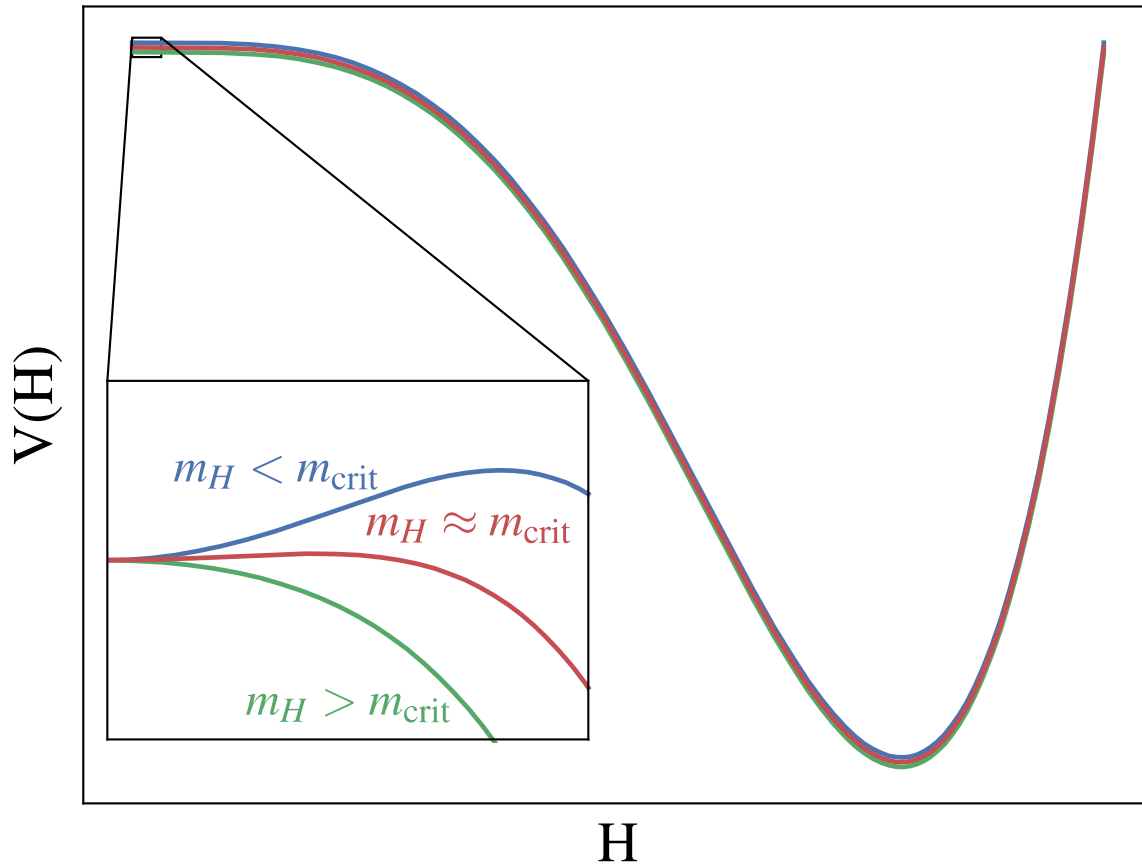
# Electroweak Vacuum Metastability



$$\beta_{\lambda}^{(1)} \sim \lambda(\lambda + y_t^2 + \text{gauge terms}) - y_t^4 + \text{gauge terms}$$

$$\lambda(\mu_I) = 0$$

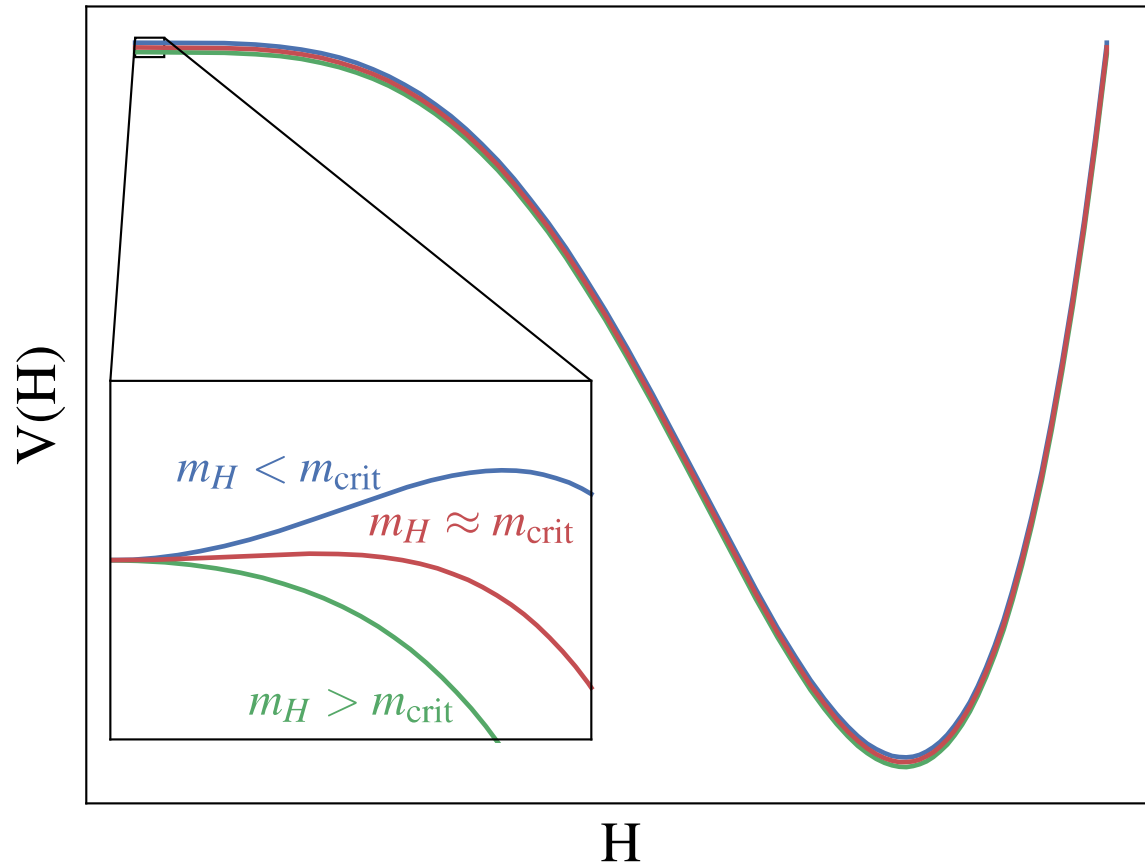
# Higgs Criticality



- Instability scale implies critical value for Higgs mass parameter
- IR + UV vacuum for  $m_H < m_{\text{crit}}$
- Near-critical point  $m \approx m_{\text{crit}}$
- Only UV vacuum for  $m_H > m_{\text{crit}}$
- Metastability bound

$$m_H^2 < m_{\text{crit}}^2 = -\frac{1}{2} e^{-\frac{3}{2}} \beta_\lambda(\mu_I) \mu_I^2$$

# Higgs Criticality



But why should we expect to live near a critical point?

# Self-organised criticality

- Criticality acts as an attractor Bak, Tang, Wiesenfeld '87
- Localisation near critical point during inflation McCullough, Giudice, You '21
- Background field varies and coupled to some operator  $\mathcal{O}$  whose expectation value changes as  $\phi$  passes through some critical value  $\phi_c$ 
$$V = (\phi - \phi_c)\mathcal{O}$$
- If  $\langle \mathcal{O} \rangle$  changes across  $\phi_c$ , stochastic evolution could localise  $\phi$  near  $\phi_c$
- Vacuum transition dynamics in the string landscape Khoury, Parrikar '19
- See also Thomas' Talk

# Vacuum Selection

- Small Higgs mass may just be the result of vacuum selection
- Given that we find ourselves in the IR vacuum, we expect an upper bound on the Higgs mass
- Does not explain *why* we are in the phase with an IR vacuum
  - selection could be realised through dynamical mechanisms mentioned
- General expectation: Higgs mass comparable to metastability bound
- But Higgs mass in the SM with substantial hierarchy to metastability bound
- Prediction: New physics leading to a saturated metastability bound

see also MD, Enguita, Gavela, Steingasser, You '25

# Axion-Higgs Criticality

- **Motivated** BSM candidate (dark matter, pNGB of global symmetries, prevalent in string theories)

- Axion-like particles are **naturally light**

- Consider simple potential

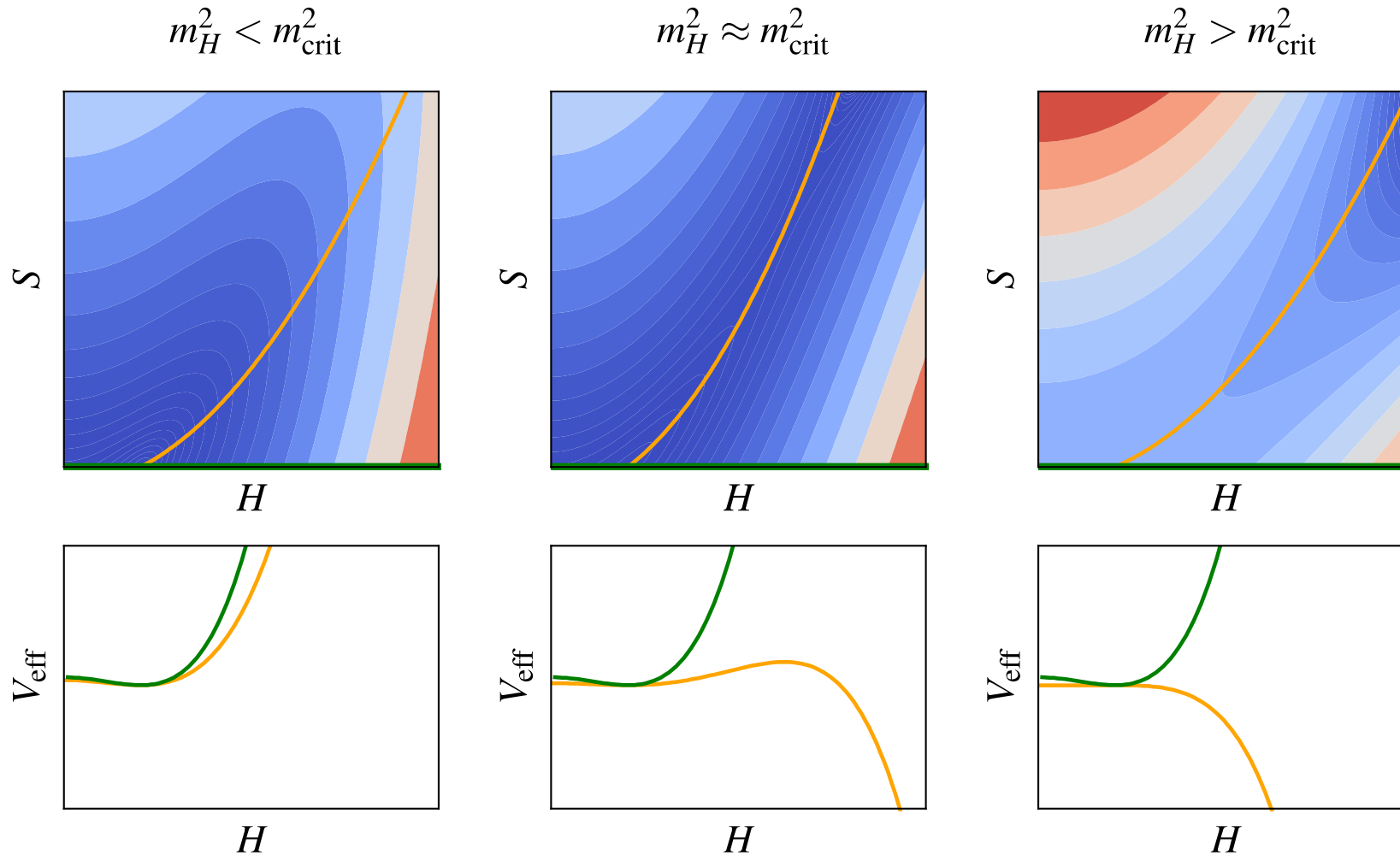
$$V_S = m_S^2 f^2 \left( 1 - \cos \left( \frac{S}{f} \right) \right) - A f H^2 \cos \left( \frac{S}{f} - \delta \right)$$

- Large decay constant

$$V_S = \frac{m_S^2}{2} S^2 - A \sin \delta S H^2$$

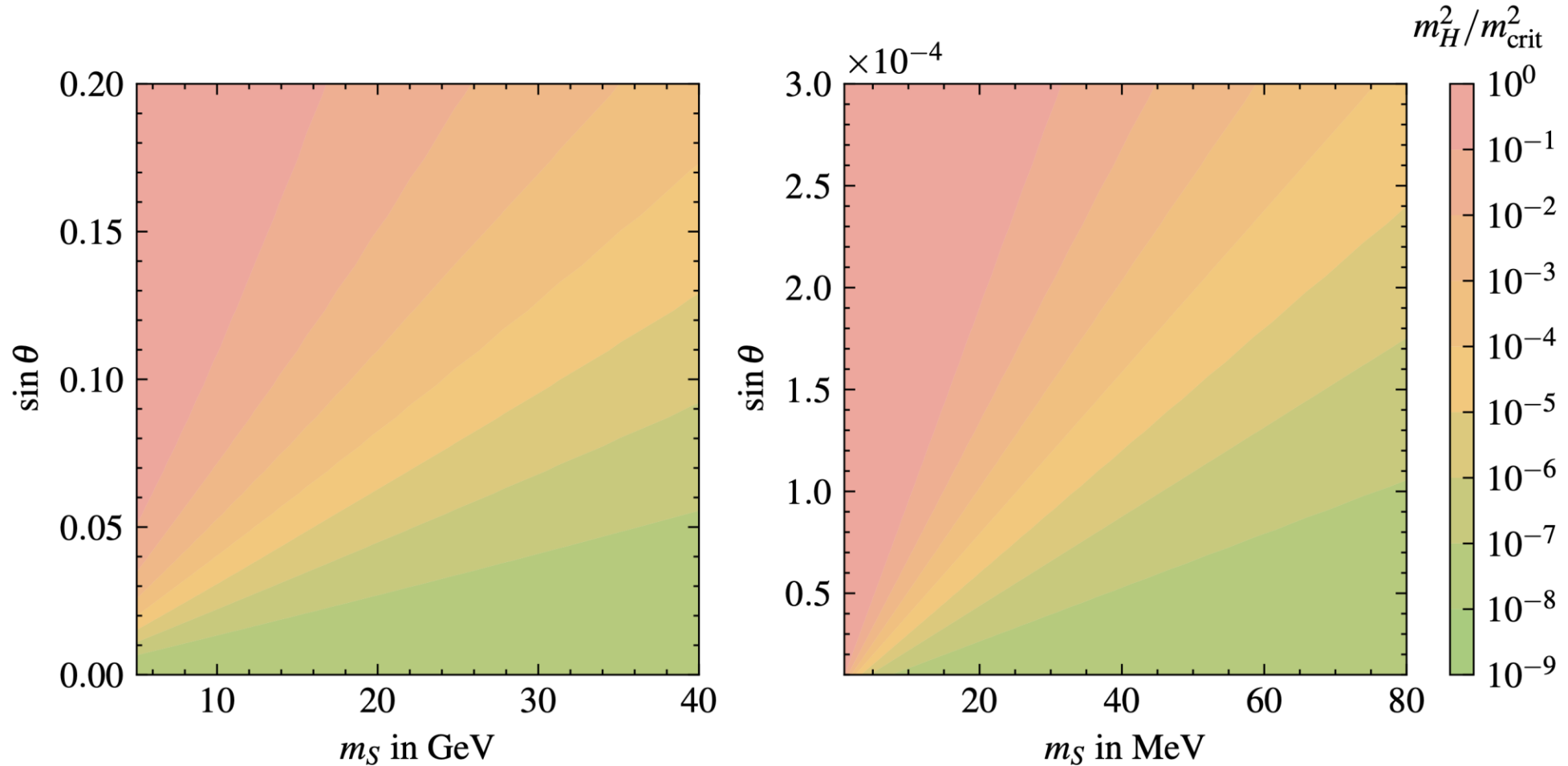
- Criticality through **mixing with Higgs** and **destabilising effect** on scalar potential

# Axion-Higgs Criticality



# Axion-Higgs Criticality

MD, You '24



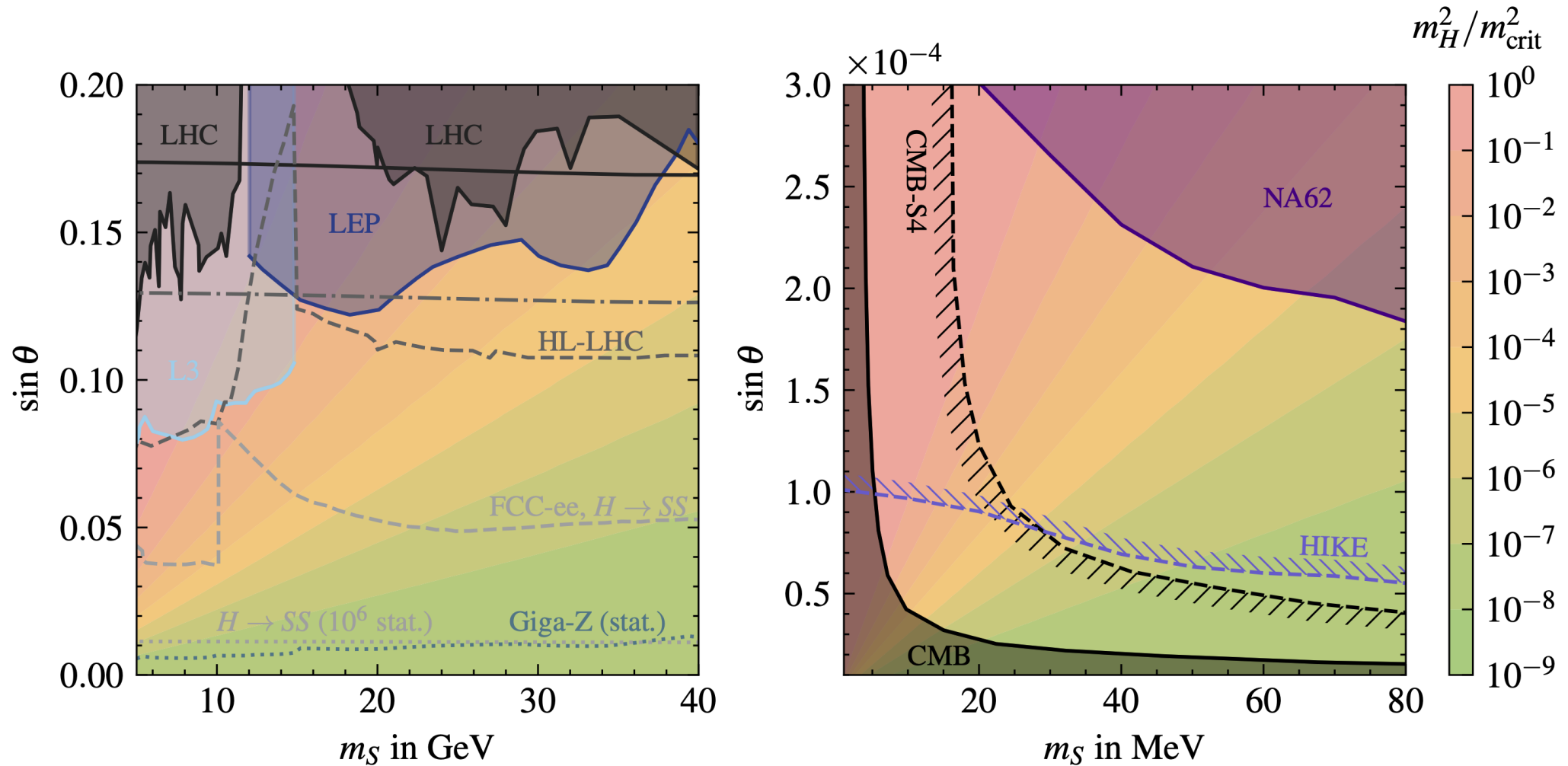


# Experimental Signatures

- Scalar direct production (SZZ vertex)
- Higgs decays
- Exotic Meson decays
- Effective number of relativistic degrees of freedom

# Axion-Higgs Criticality

MD, You '24



# Concluding Remarks

- Fine-tuning as result of **near-criticality**
- If Higgs mass is explained through cosmological criticality, then new physics coupled to the Higgs is expected
- **Criticality as new paradigm** for model building beyond naturalness
- Parameter space for Axion-Higgs Criticality **comprehensively probable** by future experiments

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## Upcoming Work:

- Dark Matter connection to Higgs criticality
- Critical Split SUSY