



## **Oscillations v sur faisceaux : d' OPERA aux expériences de 2<sup>nde</sup> génération**

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Séminaire CPPM, 28 Avril 2008







**1.** Le projet CNGS1/OPERA **1.1 Concepts & potentiels 1.2 Le détecteur OPERA 1.3 Data taking and analysis** 2. Les premiers résultats du run de physique 3. Vers T2K et au-delà



Conclusions



## **1. The CNGS1/OPERA project 1.1 Concepts & physics potential**

CONNISIONE CHURPUBBLICI 425 ENATO







### The original motivations



From Kamiokande to SuperK: Atmospheric neutrinos anomaly interpretable in terms of  $v_{\mu} \rightarrow v_{\tau}$  oscillations



CHOOZ (reactor experiment) no  $v_e \rightarrow v_x$  oscillations  $\Rightarrow \Theta_{13} < 11^\circ$ 





12h

### **CERN Neutrino to Gran Sasso (CNGS)**

- Search for  $v_{\tau}$  appearance in the CNGS  $v_{\mu}$  beam
- Validation of the  $\nu_{\mu} \rightarrow \nu_{\tau}$  hypothesis
- in the "atmospheric" sector
- Secondary oscillations  $v_{\mu} \rightarrow v_{e}$  search







#### The CNGS beam

- CERN  $v_{\mu}$  beam optimized to study the  $v_{\tau}$  appearance by  $\tau$  detection
- L=730km, mean  $E_{\nu\mu}$ =17GeV
- $(v_e + v_e)/v_{\mu} = 0.68\% v_{\tau}$  prompt negligible





#### The CNGS beam







- 2 contradictory requirements:
- $\succ$  Low X-section  $\Rightarrow$  high mass
- High granularity
  - signal identification
  - background rejection

#### Target :1350 tons, 5 years running

- 26 000 neutrino interactions
  - ~150 v, interactions
  - $\cdot \sim 15$  identified
  - < 1 background event</p>

OPER/



**Hadronic background** 

**Total per channel** 

#### $\tau$ search : backgrounds



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.077

.181

.173

.095

.229

.181

.172

.764



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#### $\tau$ search : sensitivity

τ decay channels	Signal÷∆ <i>m</i> ²	Background				
	2.5 x 10 <sup>-3</sup> (eV <sup>2</sup> )	3.0 x 10 <sup>-3</sup> (eV <sup>2</sup> )				
$ au^{-}  ightarrow \mu^{-}$	2.9	4.2	0.17			
$\tau \rightarrow e^{-}$	3.5	5.0	0.18			
$ au^{-}  ightarrow h^{-}$	3.1	4.4	0.23			
$ au^{-}  ightarrow 3h$	0.9	1.3	0.18			
ALL	10.4	15.0	0.76			
	GS data taking t/year ) target mass 10					



## $\tau$ search : observation probability vs $\Delta m^2$





## $v_{\mu} \rightarrow v_{e}$ search : physics potential

- Secondary oscillations  $v_{\mu} \rightarrow v_{e}$  search
- Good e id. in the emulsions (NN analysis)
- Simultaneous fit on  $E_{e}$ , missing  $p_T$  and  $E_{vis}$  distributions



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## **1. The CNGS1/OPERA project 1.2 The OPERA detector**

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#### **The OPERA collaboration**







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#### **Nuclear emulsions technology**

- •"Emulsion Cloud Chamber" (DONUT)
- 56 Pb + 57 emulsions sheets
- Emulsions : spatial resolution
- Compact and modular structure
- CS doublet for event prediction



- 2 emulsion layers (44 μm thick) poured on a 200 μm plastic base
  - v interaction vertex search
- kink topology reconstruction
- MCS momenta measurement
  dE/dx e/π separation at low E
- e id. and  $e/\gamma$  E measurement











#### **Bricks elements & production**

Lead production (JL Goslar, Germany)

Lead storage (LNGS) Brick Assembly Machine (BAM tunnel, LNGS)

Piling and pressing section



Emulsion storage (LNGS)

Emulsion refreshin (Tono mine, Japan

> CS production (LNGS)

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### **Bricks finding & manipulation**

•With bricks one builds walls...

brick wall under filling





Brick Manipulating System (BMS)



Ph.

### **Bricks production status**



brick production march 2008



#### **The OPERA detector**

2nd SM

#### **1st SuperModule**



Glass RPC, ~100m2

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- ~2ktons Fe
- Dipolar magnet 1.55T
- T.T. : trigger, neutrino interactions localization (brick finding), kinematics Spectrometer :  $\mu$  id., charge and momentum measurement <sup>19</sup> Bricks :  $\nu$  vertex id., decay kink search, kinematics, e/ $\pi$  sep., e/ $\gamma$  E meas.



### **Target Traker modules components**

Goal : trigger and brick finding

- 62 XY planes
- $\checkmark$  32256 scintillator strips 7m  $\times$  2.5cm  $\times$  1cm (AMCRYS-H, Kharkov)
- WLS Fibres 1mm diameter (Kuraray)
- MaPMT 64 channels (Hamamatsu) :  $16 \times 62 = 992$  PMTs
- f Dedicated autotriggerable F/E electronics (min. threshold = 1/5 p.e.)



## Extruded plastic scintillator strip







#### TT active readout

#### <u>64 # MAPMT</u>

(Hamamatsu)

#### $3x3 \text{ cm}^2$







#### **Front-End Chip: final** *"slow" for charge measurement*



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- Technology: AMS BiCMOS 0.8 μm
- **Chip area** :  $10 \text{ mm}^2$
- □ Package : QFP100
- Power consumption : 185 mW

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#### channel to channel gain adjustment (1:4 range)



#### The OPERA detector construction









- Today : all electronic detectors commissioned and running since '06 except Veto (march '07) and HPT 2<sup>nd</sup> SM (1<sup>st</sup> half '07)
- General DAQ and GPS clock distribution running
- On-line software and DB schemes commissioned



## 2. Data taking and analysis



On-line electronic detectors



Offline emulsion analysis





#### **DAQ general features**

> The distributed DAQ is based on Ethernet. Each sensor (1200) is seen as a node in a Gigabit standard network. The basic "element" of the system is a daughter board ('mezzanine") embedding FPGA, FIFO,  $\mu$ −processor (AXIS)













R/O: scintillator + WLS fibres + MaPMT

The client/server protocol used relies on the CORBA standard implemented in C++ with interfaces into postgreSQL and Oracle database.

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This software is completely object oriented and uses the Interface Description Language (IDL) to describe the distributed objects independently of the <sup>24</sup> programming language. InterORB protocols guarantee interoperability.

## **Clock distribution & network architecture**

Node card i

Each individual node runs a local 100MHz clock generated via a common 20MHz clock send from a precise and stable oscillator. The oscillator is plugged onto a dedicated PCI board which locks the clock signal on the GPS and encode specific commands (propagation delay meas., reset, reboot etc).

Optical fiber

« Slave » clock

MLVDS

in Hall C (receives the GPS signal from the Outside antenna though a 8km optical fiber)

SM1





SM2





### **Event selection & brick finding**

#### **Data processing:**

data are continuously extracted from the DAQ data base

- all events are processed through the OPREC reconstruction package
- In time events are selected and stored on a dedicated file

#### **Brick finding:**

all in time events are scanned visually and sorted by categories:

- > in target event CC or NC like (according to the muon id)
- magnet interaction
- rock muon

in target events are analyzed through the brick finding package

- > each step of the reconstruction for the muon is checked manually
- iterative process including;
  - neural network for the wall finding
  - \* hadronic shower parametrization
  - \* probability estimate for the vertex location in each brick

\* exact configuration and alignment of the target at the time of the event (BMM data base)

> at the end the list of bricks is provided for each event

#### **Brick extraction:**

Listed bricks are extracted by the BMS

Extraction operations still performed in parallel with insertion

## **CC-like event processing : OPERA "baby"**

#### « Opera baby » P<sub>u</sub> : 7.87 GeV/c



## CC-like event processing : OPERA "baby"

-670 Z (cm)





700

11451 p.e. tot 1579 p.e. muon



Wall 5 Tray 24 Cell 6 prob= 0.9 Wall 6 Tray 24 Cell 6 prob= 0.09



-730

-720

-710



<del>R</del>h

#### **Processing : development lab at LNGS**





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### **Off-line emulsion scanning**

Two different systems (both operational) are running for scanning systems : "hard"-coded oriented (Japan) or "soft"-coded oriented (Europe). Both systems are ready and working at more than 20cm<sup>2</sup>/hour (10 required).



Synchronization of

objective lens and stage

Customized commercial products Software algorithms

CMOS

camera

500 fps

#### Dedicated hardware Hard coded algorithms

Constant speed stage



#### **Scanning station overview**

Emulsion scanning is performed in a fully automatic way.

About 40 microscopes are operational in the various OPERA scanning laboratories.





#### Bern emulsion scanning lab.

Predictions from electronic detectors are followed back inside the brick until tracks stop. Then a full scanning around neutrino interaction vertex is performed and the event topology and kinematics reconstructed<sup>3</sup>.



### **Off-line emulsion scanning**

- Tracks reconstruction by analysis of ~16 "slices" per emulsion
- Identification of "base"-tracks (both emulsion sheets)
- Scan back of tracks in successive emulsions of a brick



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Zoom

UnZoom



# 2. First results from physics run

#### 3 runs : august & october '06, october<sup>34</sup>'07





#### August '06 run

- 8.5 days of real beam operation (121 hours within 2 weeks of operation)
- Used for electronic detectors, DAQ, GPS commissioning and tests of CNGS-OPERA information exchange (gateway DB, early warning signal...)
- Beam side informations :



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#### **Timing issues**

- Events selection from GPS time information : T<sub>OPERA</sub> -(T<sub>SPS</sub>+TOF)<T<sub>GATE</sub>
- The events time distribution is peaked around the 2 extraction peaks times within negligible CR background (o(10-4) in ~ms windows)



-2500

2500

50.00

7500

 $\Delta t (unit = 10ns)$ 

1000 × 10

~T<sub>COINCIDENCE</sub> (typ. 200ns)

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2h



#### August '06 run

- Neutrino induced interactions : 319 detected altogether (in agreement with the predictions)
- 3/4 muons coming from the rock, 1/4 neutrino interactions in the detector

#### CC event originated from material in front of the detector







CC event originated in the detector's target material



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#### **CNGS vs cosmics events**



**Cosmic rays induced events** with a typical down-going topology

> Beam events: ~horizontal tracks





#### **October '06 run**





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#### October '06 run



- Extrapolation from ED predictions to bricks validated for beam events.
- Under evaluation for cosmics.
- Requires additional run to tune at least the brick finding procedure

- Short run due to CNGS leak problem.
- 25 neutrino events collected.
- Ikbricks in the target (BAM pilot run)
- 1 brick crossing muon identified.
- Brick extracted from the wall.



Details of the CS scanning showing the reconstructed grains.



### **Cosmics analysis**

- TOF analysis to disentangle up/down-going particles
- 10ns timestamp accuracy
- Cuts on the event topology (track length in the detector)



- 1/(β=v/c) distributions
- -1 : down-going particle / +1 : up-going particle
- Orange: on-time events (beam)
- ~250d data taking
- Data/full MC comparison under investigation





### '07 physics run

- Short run due to CNGS services electronics irradiation problems.
- 38 neutrino events collected in the OPERA target (1<sup>st</sup> SM) : 29 CC + 9 NC
- Bricks extractions + CS scanning for prediction confirmation + Scanning



## **Event 178969961 :** $v_{\mu}$ **CC interaction**





OPERA

1.Manusa

5 prongs associated to the neutrino interaction

 $\langle IP \rangle = 9 \ \mu m$ 

Electromagnetic shower pointing to the vertex  $(\gamma$  conversion)

## Event 179673325 : QE-like topology







The visual inspection allow the observation of nuclear fragments and the classification of the event as DIS







TOP VIEW (Horizontal







#### Perspectives on the CNGS : '08 physics run

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Start: May 29<sup>th</sup> End : Nov. 10<sup>th</sup> 147 days for the CNGS (200)

3\*10<sup>19</sup> p.o.t (4.5\*10<sup>19</sup>)

3470
2610
780
23
110
1.3





## 3. Vers T2K et au-delà





#### Future (2010–2020): $\theta_{13}$ , CP violation, sgn( $\Delta m^2$ )

- Proof for  $v_{\tau}$  appearance in the atmospheric sector (OPERA)
- Precise measurement of  $\Delta m_{23}^2$  and  $\theta_{23} \rightarrow is \theta_{23}$  exactly  $\pi/4$ ? (MINOS, OPERA, T2K)
- Improve the accuracy on the solar parameters

• Measure  $\theta_{13}$  (why  $\theta_{12}$  and  $\theta_{23}$  angles are large and  $\theta_{13}$  seems very small or null ?)

•Determine if the mass hierarchy the same as for charged leptons (sign of  $\Delta m_{23}^2$ ) throug MSW matter effects



•Absolute mass values (cosmology, double beta)

•Show CP violating effects (measure the phase  $\delta$  in the mixing matrix)



#### Future (2010–2020): branching point

#### $sin^{2}2\theta_{13} > 0.01$

- Conventional super  $v_{\mu}$  beam works
- Known beam technology
- Background highly nontrivial
- $v_{\epsilon}$  beam contamination not negligible but tolerable
- Upgraded sensitivity with scenarios at the second maximum (T2KK) or high energy beam (Nova)

#### $\sin^2 2\theta_{13} < 0.01$

- beta beam or
- neutrino factory required



**Branching point :** things change at  $\sin^2 2\theta_{13} \sim 0.01$  : the priority is to set the scale of the effect  $\Rightarrow$  T2K



#### The 1<sup>st</sup> Super-Beam: T2K from Tokai to SK

- Low energy neutrino beam (~0.6 GeV) from Tokai (JPARC multipurpose proton accelerator) to Super-Kamiokande by 2009.
- $\cdot$  0.75 MW from a 50 GeV proton synchrotron (neutrino beam two orders of magnitude more intense than in K2K)
- foreseen upgrades for CP violation search: 4 MW power and 1000 kton Hyper-K (2020?)



#### Physics objectives ...

- T2K will be the 1<sup>st</sup> 'next generation' physics programme
- Goal: precision measurements on the oscillation parameters
- Goal: discovery of (still) unknown phenomena (CP violation, neutrino intrinsic nature, neutrino mass hierarchy etc)





#### T2K : off-axis technology, near detector(s)



## COT.

### LAr technology in T2K

## Intermediate detector @ 2km



 $E_v$  spectrum @ 2km

~  $E_v$  spectrum @ SK w/o oscillation

→ Uncertainties from Far/Near ratio is smaller than ND@280m.

- Possible Detector configuration
  - Liquid Ar TPC
  - Water Cherenkov
    - Same target & v reconstruction algorithm as SK
  - Muon Range Detector



Facilities for 2km is to be requested in Japan after the commissioning of J-PARC facilities.





Detection of primary ionization in LArgon: 1 m.i.p ~ 20000 electrons on 3 mm

High resolution calorimetric measurement of e.m. and hadronic showers

PMs detecting UV scintillation ligth in Argon used to provide the t=0 signal of the event to measure the drift space
Each wire is sampled with a 2.5 MHz flash ADC to measure the charge distribution along the drift (time)

> 2D event reconstruction with 1 mm encode recolution.



### LAr technology in T2K

Tracking + Particle identification + Calorimetry 100 Ton mass easily feasible





8	FY2001	FY2002	FY2003	FY2004	FY2005	FY2006	FY2007	FY2008	
		Bldg	lg. construction			Beam			
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			Bldg. construction				Beam		
50 GeV			Equip. construction						
				Bldg. co	construction Be		Beam		
Materials + Life				test					
				Bldg.	constructi		Beam		
Nuclear-Particle					Equip	tion	test		
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## Conclusions

The OPERA experiment has completed almost entirely the construction of all electronic detectors and faces the last (large) effort of brick production and insertion.

The data collection from electronic detectors and from the scanning systems have been validated.

Ready for the full oscillation physics in '08.

Future neutrino physics will be oriented after the T2K campaigns (phase-I/-II). Looking forward for the ultimate facilities (neutrino factories)





## The end



#### Scanning rehearsal with bricks in NuMI beam : PEANUT





#### anning rehearsal with bricks in NuMI beam : PEANUT



**12**h



#### **Scanning in Japan**

Track Recognition by Track Selector





## **Vertex finding**

#### Scan-back strategy tested on 8 GeV/c pions to produce interactions

- Scanback: measure tracks on downstream plate and follow back to interaction point (realistic test of performance in OPERA)
- TotalScan around track disappearance points to confirm the interaction







*Inter-calibration tracks (cosmic rays, plate to plate alignment)* 



#### **CERN-LNGS UTC clocks intercalibration**

For the neutrino spill syncronization both CERN and LNGS have a double unit (including a spare) UTC clock system, but from different manifacturers. The CERN system was calibrated by the Swiss metrology institute METAS.

CERN and LNGS systems have comparable performance (<100 ns) and their single units are in both cases based on a GPS system + Rb clock.

One of the CERN UTC units was installed and running for one month in Gran Sasso in order to check for relative offsets and time stability of the two systems. Action was taken also to measure all the delays in the LNGS time distribution chain







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