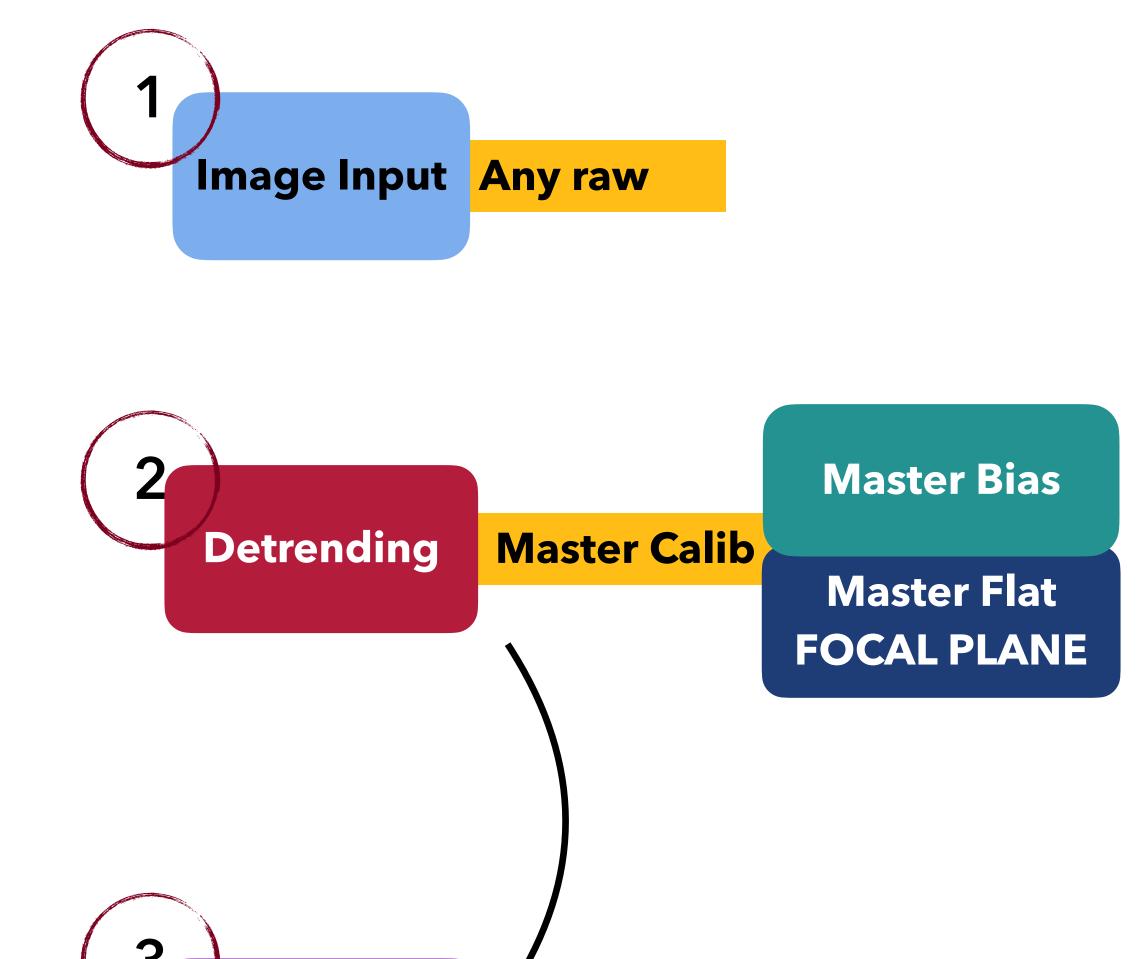
# ztfin2p3 - Pipeline status Marie Aubert @ ZTF France - 07/11/24

How does it work?
What does it do?

- Image I/O and handling (e.g WCS) : ztfimg
- Calibration and everything else: ztfin2p3
- 3 Main steps from Raw to Aperture catalogues:



**Aperture** 

Catalogues on the fly for

*U*bercal

→ No sciimg ztfin2p3 @ CC

1 Image Input Any raw

How does it work? What does it do?

- Image Input and handling (e.g WCS): ztfimg
- Image Input is the basis  $\rightarrow$  Applies to any raw opened in the pipeline :

#### **Image Input:**

ztfimg.RawCCD ztfimg.RawQuadrant

- 1. Open
- 2. Pocket correction (date > 20191022)
- 3. Corr for NL from linearity tables / rcid (date dependent)
- 3. Corr for overscan:
- Only last five columns used for model.
- Model is cubic polynomial applied to sigma-clipped overscans

How does it work?
What does it do?

- Image I/O and handling (e.g WCS) : ztfimg
- Calibration ztfin2p3

1 Image Input Any raw



Image I/O

- 1. Use last 10 bias / day
- 2. Pixel-level sigma-clipped (3) median

**Master Bias** 

- → Store ( with Header → package versions)
- → Image are store in fits files / CCD.

ztfin2p3.pipe.newpipe
 class BiasPipe

How does it work? What does it do?

- Image I/O and handling (e.g WCS) : ztfimg
- Calibration ztfin2p3

1 Image Input Any raw



Master Calib

#### **Master Flat**

**Image I/O** 

- 1. Pixel-level sigma-clipped (3) median / LED
- 2. Normalize flat / LED / CDD with CCD Median
- . Average CCD & norm to get Flat / Filter
- 4. Re-open everything to estimate median level at Focal Plane level
  - → Store in Header
  - → Image are store as CCD normalized in fits files.

ztfin2p3.pipe.newpipe
 class FlatPipe

**Master Bias** 

How does it work?
What does it do?

- Image I/O and handling (e.g WCS) : ztfimg
- Calibration ztfin2p3

1 Image Input Any raw



**Master Calib** 

#### Image I/O

**Master Bias** 

Master Flat FOCAL PLANE

#### Detrending

- 1. Get the nearest calib files  $\rightarrow$  CCD, filter, date matching.
- 2. Open raw file with ztfimg  $\bigcirc$
- 3. Correct for bias ( $I_{\text{raw}} I_{\text{master bias}}$ )
- 4. Correct for flat at focal plane level  $I_{\text{sci}} = \frac{I_{\text{bias corr}}}{I_{\text{Flat}}} \times \frac{N_{\text{fp}}}{N_{\text{ccd}}}$ 
  - 5. Separate  $I_{\text{sci}}$  into four quadrant  $\rightarrow$  IPAC like sci format.
  - 6. Get associated headers & go from there

ztfin2p3.science
build\_science\_image
build\_science\_header

#0utput : ztfimg.ScienceQuadrant x4

How does it work?
What does it do?

- Image I/O and handling (e.g WCS) : ztfimg
- Calibration ztfin2p3

Image Input Any raw **Master Bias** Detrending **Master Calib Master Flat Master Calib FOCAL PLANE** ztfin2p3.aperture **Aperture** 

get\_aperture\_photometry

Image I/O

**Detrending** 

- **Aperture**
- Match Image to source catalogue to each quad : **GAIA DR3 Ben ver.**
- 2. R.A, Dec to ra, dec with pec-vel
- 3.  $\rightarrow$  (x, y) quadrant coord.
- 4. sep package aperture
- 5. Format and store to parquet

# Status of the pipeline

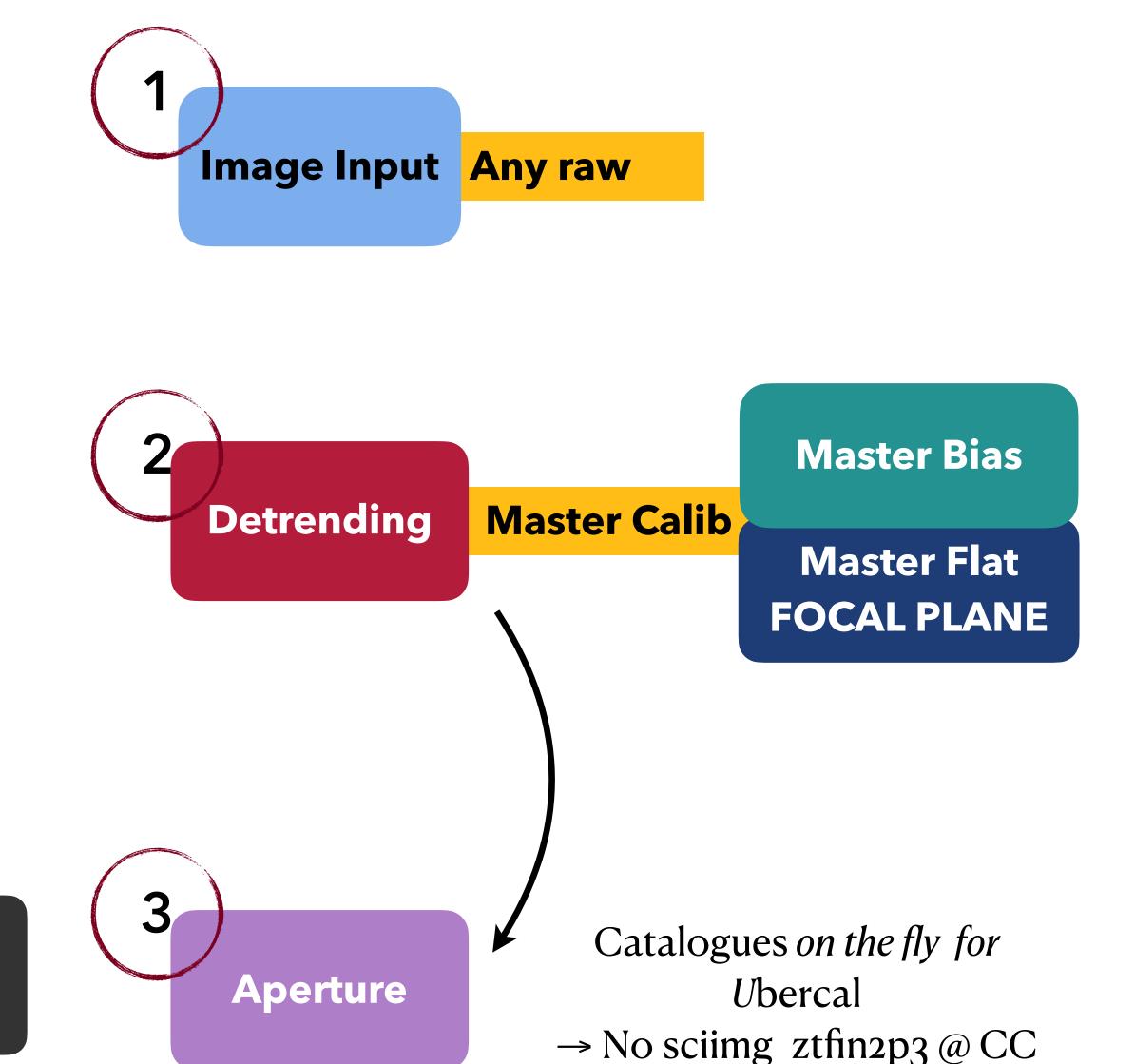
→ Ready to run (until the next issue)

Implementation aspects are coded.

Simon did a lot of refactoring & developed command line environments.

Simple command line commands: (please don't run them unless discussed)

\$ ztfin2p3 run calib [options]
\$ ztfin2p3 run d2a[options]



Interacts with CC Slurm → Create relevant batch job izy pizy.

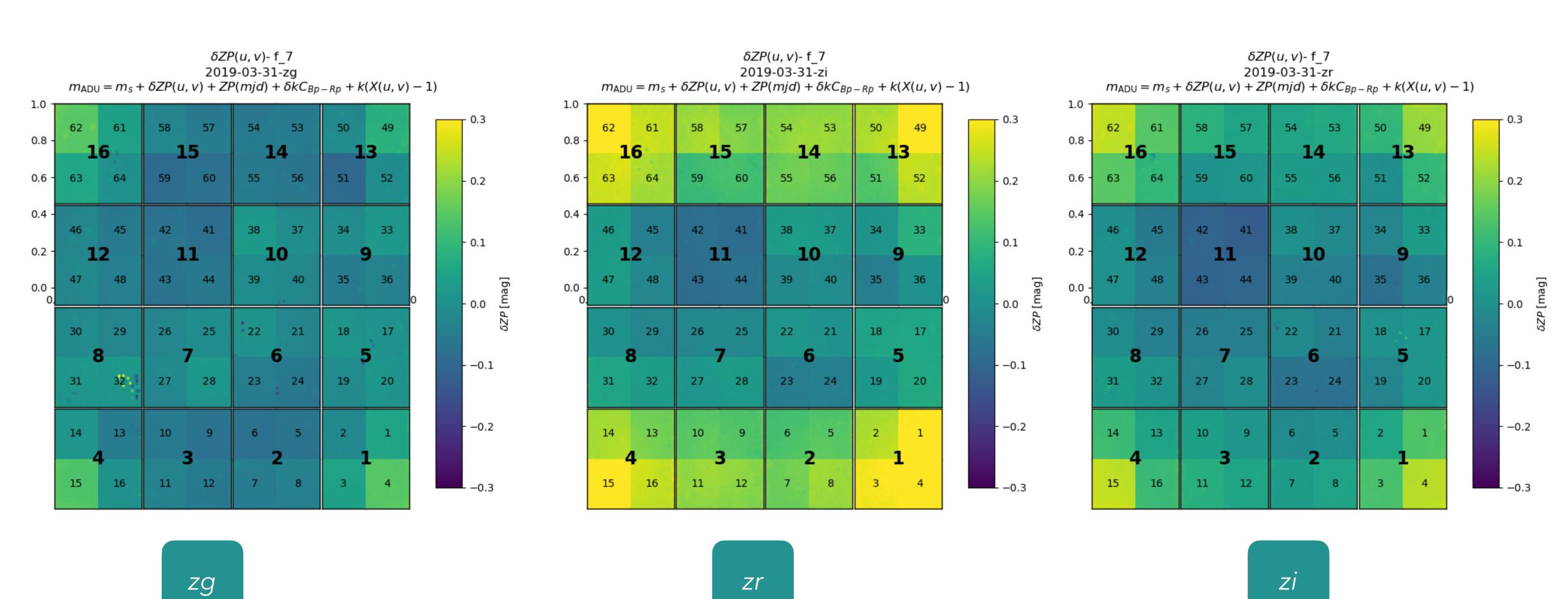
## Pipeline validation—as it is

In progress

- Pipeline will be tested both ways:
  - Starflats fits to aperture catalogues
  - Check with Uberflats if overlap possible (for now toy sample for Ben)
- Ran: 2019-03-31 in the three bands.

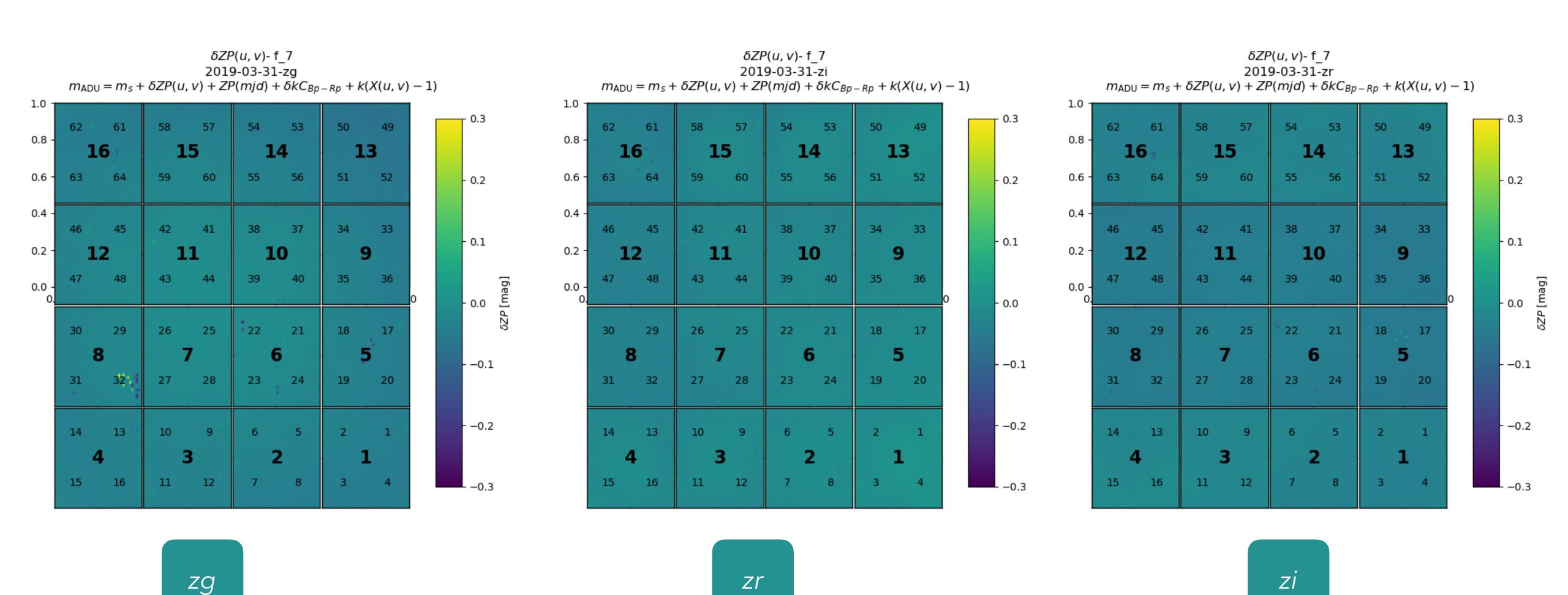
### Starflats 2019-03-31

### Focal plane residuals with CCD normalization



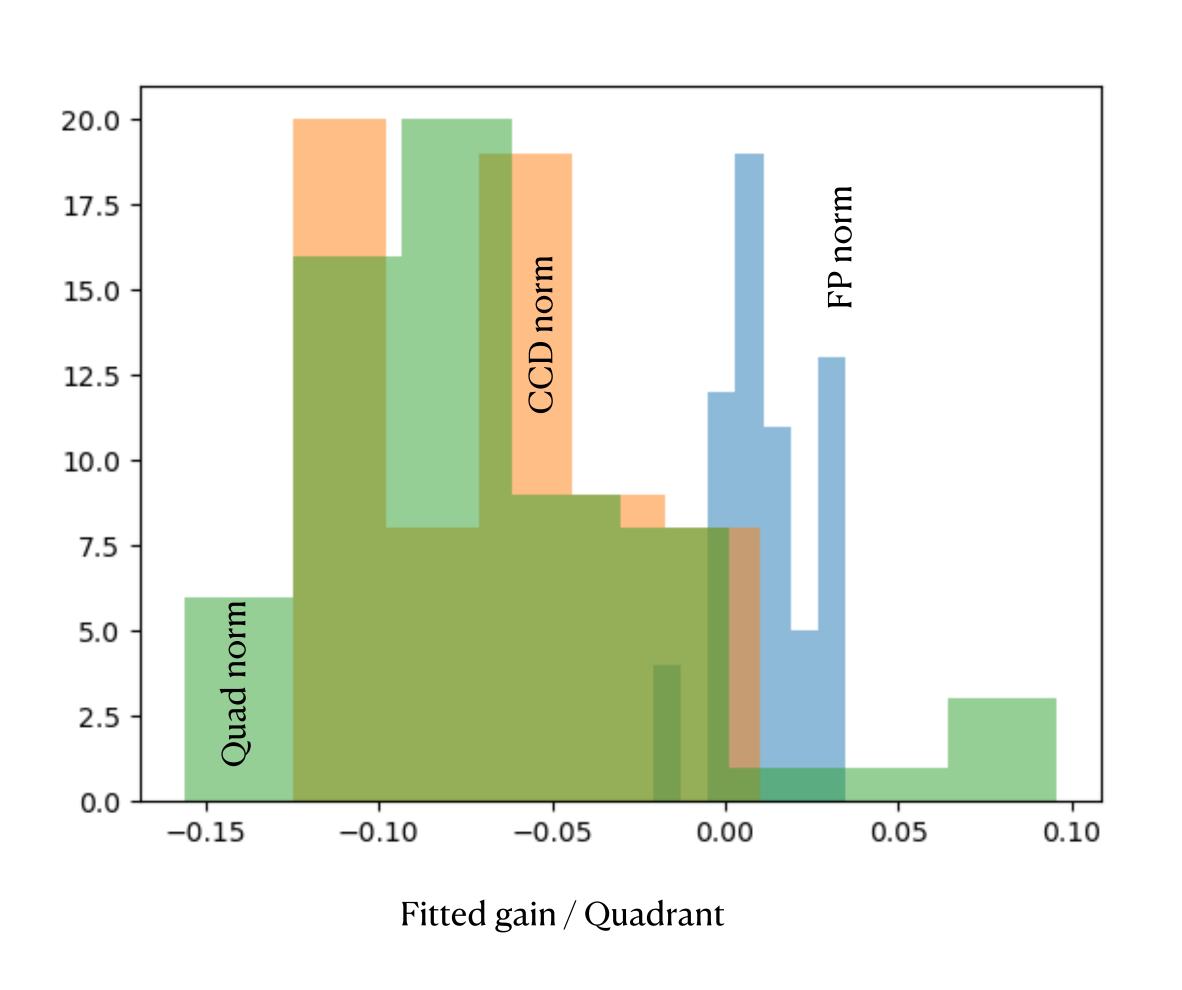
# Pipeline validation — as it is

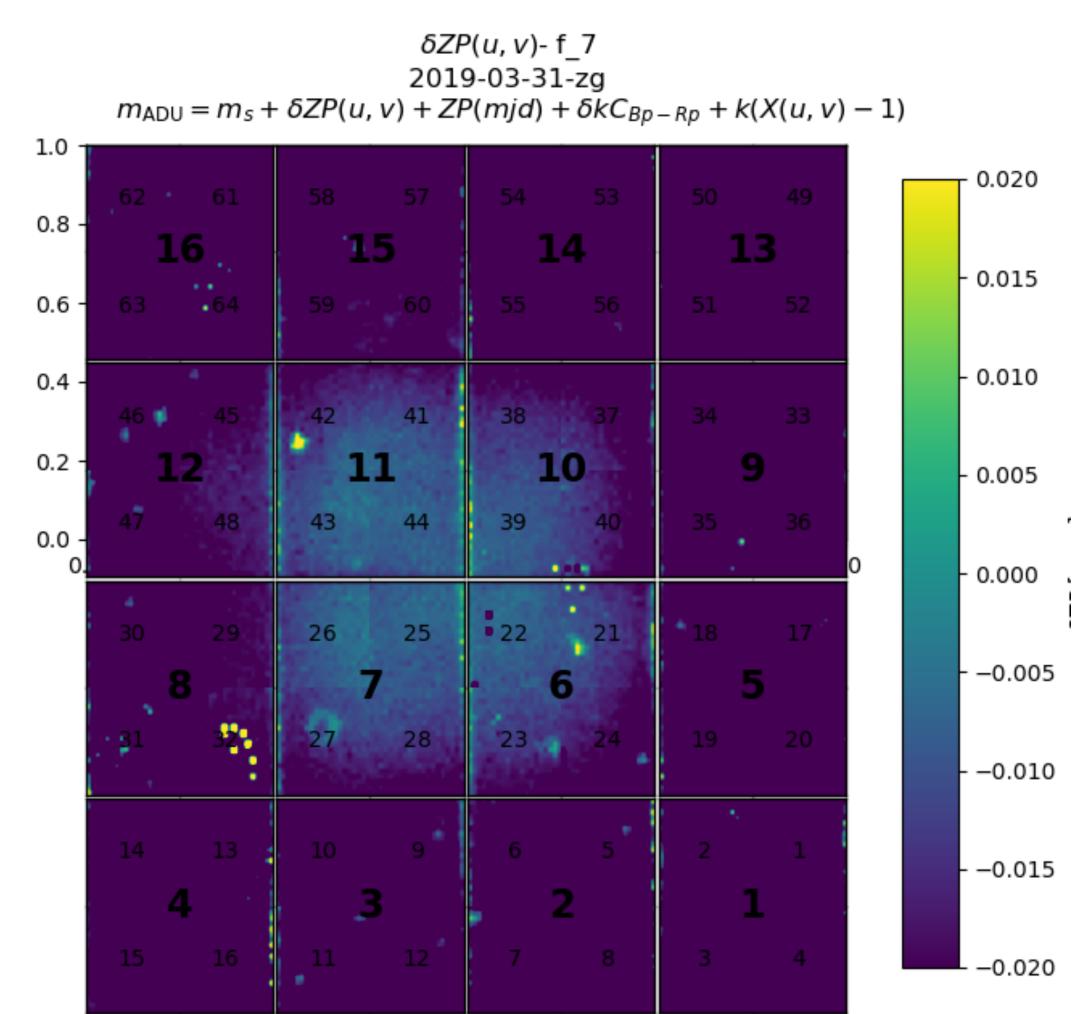
### Focal plane residuals with Focal Plane normalization



# Pipeline validation — as it is

Focal plane residuals with Focal Plane normalization





## Pipeline validation — as it is

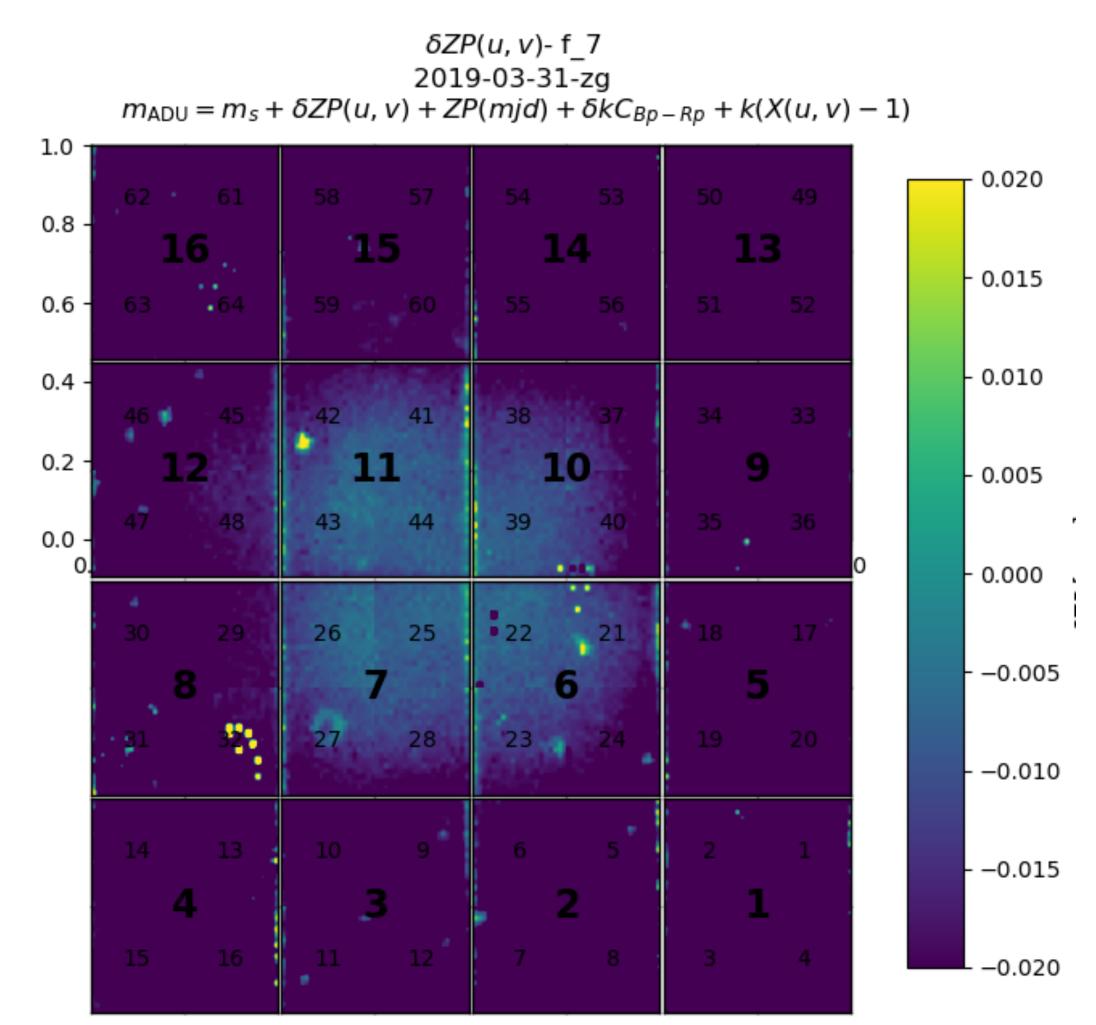
### Focal plane residuals with Focal Plane normalization

### Focal plane residuals:

- Uniform but some residual vignetting
  - → Off centering of the flat fields?
  - → Some flatness issue
- Some residuals gain jumps as well.

Checks to be done with other starflats. (bands and date -> persistent behaviour?)

Need to validate that with pocket as well.



# Starflats & Pipeline validation:

→ Halted because no much room @ CCIN2P3

# Now addressing concerns:

- ★ What's done : → See previously
- \* Missing aspects in pipeline:
  - → Fringe is TBD, Pocket needs dual validation in 2021 and 2023 (after NL changes), BF @ Nicolas?
  - → Validation thingies → finish star flat
- \* Stored in pipeline:
  - → Aperture catalogue, some images for SM (needs to discuss the strategy in detail).
- \* Room @ CC:
  - → 1.2PB as temporary fix but we are at 1.09PB. Cleaning will be done
- \* Strategy (on-disk, on the fly) long-term data management:?
- \* Bookkeeping: Not per say.

# Conclusion & what needs doing by Dec. 2024

#### Personnel vision on what needs to be done.

- Check that downloaded data is OK: (in progress)
  - → No corrupted raw files & mask files.
  - → All necessary masks are there.
  - → No missing header.

2019 special:
Before removing check
that nothing will be
lost

- Remove data (soon): (probably Sebastien / Mickaël)
  - → List files **NOT** to be deleted.
- Personal to do list:
  - → Finish creating data for starflats (with Pocket validation as well)
  - → SM needs to check how they run their pipeline.
  - → Check stability of flats for some Ubercal period hints.
  - → fringez implementation.

# Pipeline vs Needs of subsequent analyses?

What's your opinion on the matter?