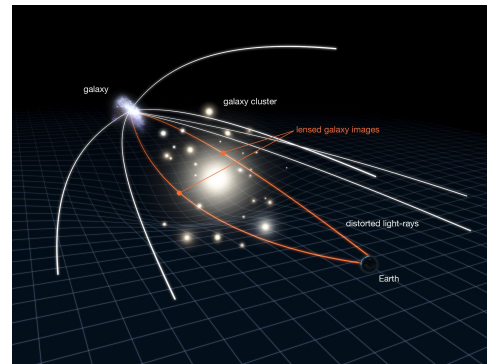


# Galaxy clusters in LSST working group

- The goal of the working group is to obtain constraints on  $\Omega_m$  and  $\sigma_8$  from the  $\lambda$  - CDM
- To use galaxy clusters as a probe, we want to analyze the number of clusters in the sky and the lensing signal from source galaxies behind the cluster
- For this analysis we need:
  - Data (simulations for the moment)
  - Organize the data into catalogs with the needed information
  - Make the theoretical predictions (counts, weak lensing, covariance)
  - Implement the likelihood between data and prediction
  - Run an MCMC



# Data

Given an area, redshift and richness region of the sky:

- The number  $\mathbf{N}$  of clusters in a region of the sky
- The shear signal from source galaxies behind the cluster

$$\begin{aligned}\gamma_+(\vartheta) &:= \oint \frac{d\varphi}{2\pi} \gamma_+(\vartheta, \varphi) = \bar{\kappa}(\vartheta) - \kappa(\vartheta) \equiv \frac{\Delta\Sigma(\vartheta)}{\Sigma_{\text{cr}}} \\ \gamma_\times(\vartheta) &:= \oint \frac{d\varphi}{2\pi} \gamma_\times(\vartheta, \varphi) = 0,\end{aligned}$$

# Theory Prediction

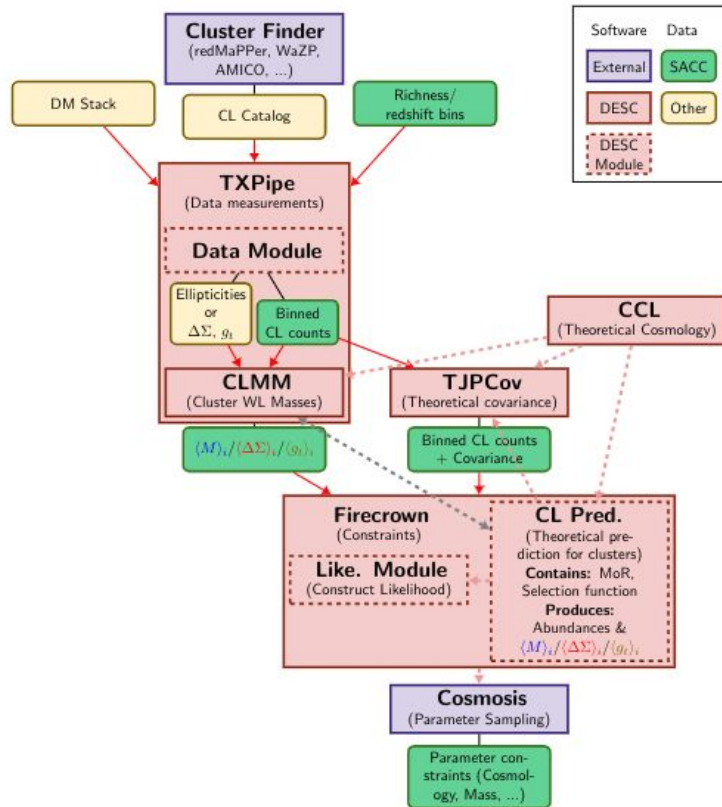
- Counts prediction (binning in z-richness)

$$N_\alpha^\beta = \int_{z_\alpha}^{z_{\alpha+1}} dz \frac{dV}{dz d\Omega} \int_{\ln \lambda_\beta}^{\ln \lambda_{\beta+1}} d \ln \lambda \int_0^\infty dM \frac{dn}{dM} P(\ln \lambda | M, z)$$

- Shear profile prediction

$$\begin{aligned}\Delta\Sigma_{\lambda_\beta}(R) &\equiv \frac{1}{N_{\Delta\Sigma}(R; \lambda_{\beta, \min}, \lambda_{\beta, \max})} \\ &\times \int_{z_{\min}}^{z_{\max}} dz \int_{M_{\min}}^{M_{\max}} dM \frac{\chi^2(z)}{H(z)} \\ &\times w_l(z; R) \frac{dn}{dM} S(M | \lambda_{\beta, \min}, \lambda_{\beta, \max}) \Delta\Sigma(R; M, z)\end{aligned}$$

# The Desc Cluster Pipeline





# Firecrown

**Python package that offers the DESC framework for implementing likelihood**

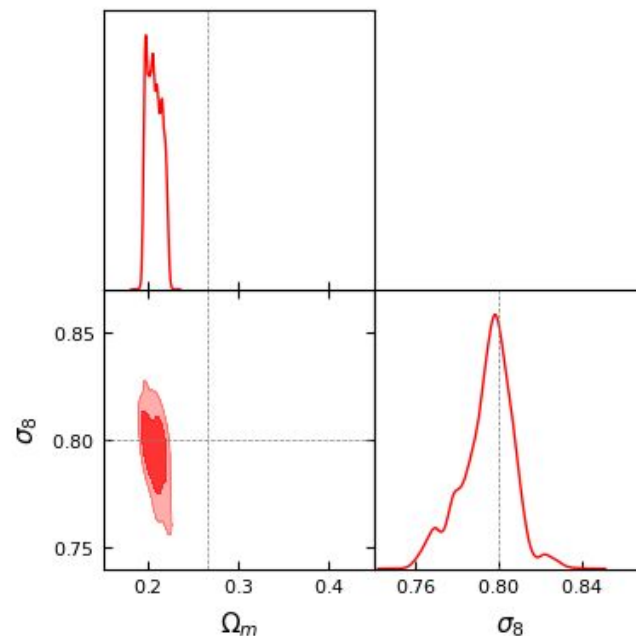
<b>Key properties</b>	<b>Relevant cluster codes</b>
<ul style="list-style-type: none"><li>• Dedicated likelihood implementations for 3x2 point, weak lensing, cluster counts, super nova, etc.</li><li>• Connection with cosmosis (MCMC) and NumCosmo.</li><li>• Data Reading</li></ul>	<b>Cluster theory model</b>
	Cluster counts prediction implementation <ul style="list-style-type: none"><li>• true mass/true redshift</li><li>• Mass proxy / redshift proxy</li></ul>
	<b>Cluster Number Counts statics</b>
	Likelihood construction, data reading/writing.

# Requirements to add new Likelihood to Firecrown

- Add theory predictions for new effect
- Add code to read and organize new data from the SACC format
- Create a recipe for theoretical prediction
- Create new statistics object that will call the recipe

# ✦ Shear Profile implementation

- I have been working on the theoretical prediction and an statistics object implementation in Firecrown to perform the shear profile analysis
- First test on mock data
  - Wrong evaluation of  $\Omega_m$  probably due Cov matrix or integration method



## Next goals

- Find source of error with ideal data
- Test code on Redmapper catalog run on CosmoDC2
- Add more systematics (misscentering, triaxiality, purity, completeness)
- Start working on TxPipe
- Make sure the systematics used throughout the pipeline are consistent