

# Disclosing signatures of dark matter with astroparticle messengers

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**Silvia Manconi**

Sorbonne University & LPTHE

December 5, 2024

FRIF Day 2024

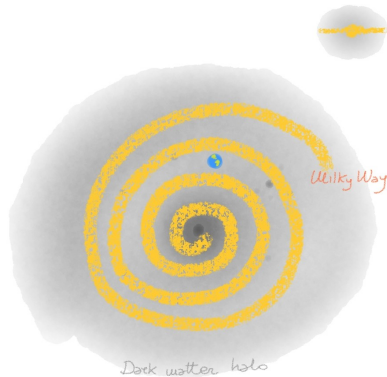
**Silvia Manconi**, Chaire de Professor Junior, started 1<sup>st</sup> December 2024

*Astroparticle physics and theory models of dark matter*

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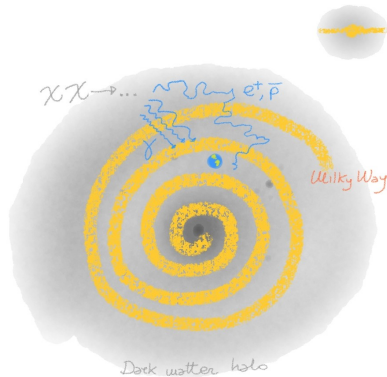
- Nature of **Dark Matter**: indirect searches
- **Astroparticles** origin: sources, propagation



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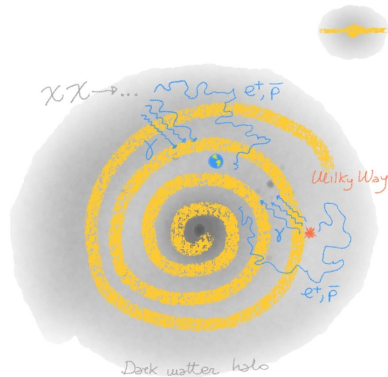
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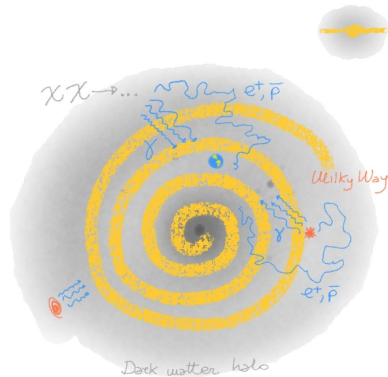
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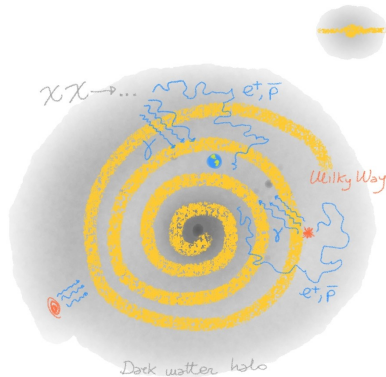
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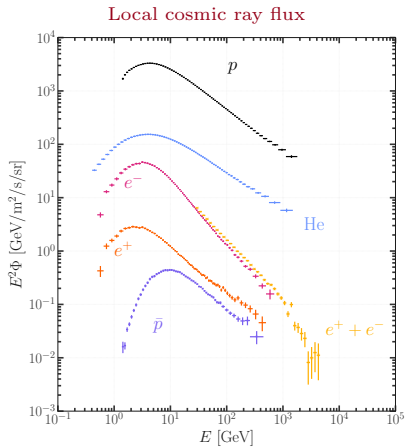
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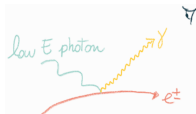
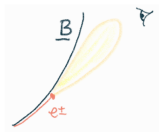
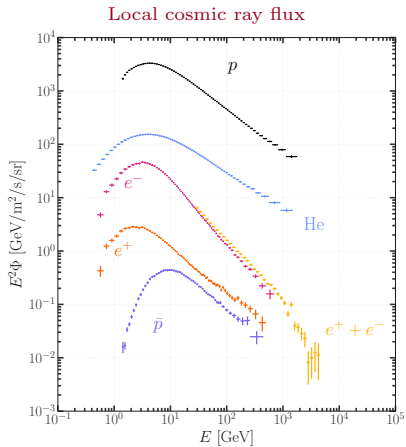
Phenomenology of high energy astroparticles (GeV-TeV)  
in the Milky Way and beyond  
from exotic and astrophysical sources

# Astroparticle messengers: precision era

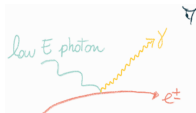
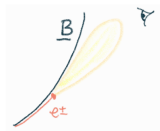
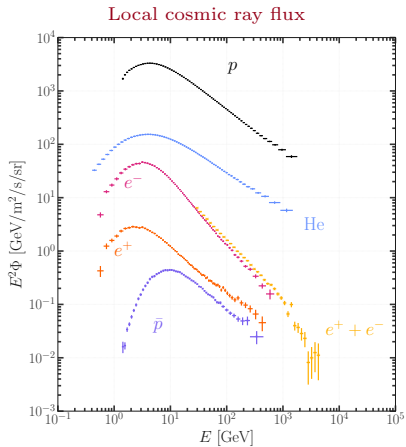




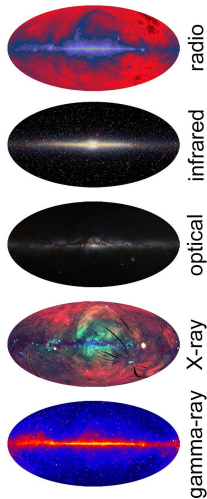
# Astroparticle messengers: precision era



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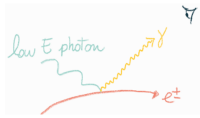
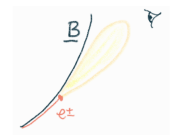
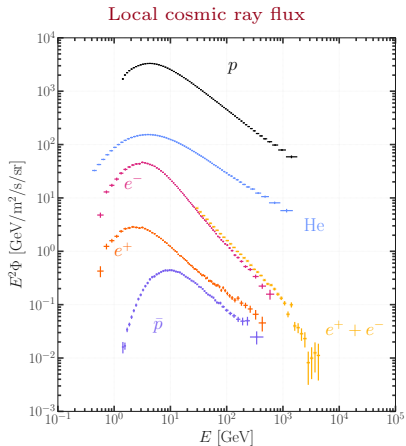


## Multiwavelength photons

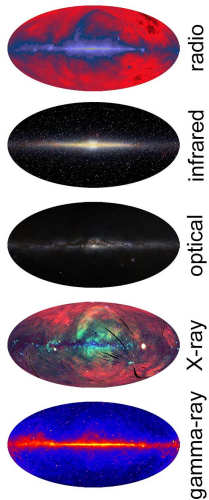


[NASA]

# Astroparticle messengers: precision era



## Multiwavelength photons



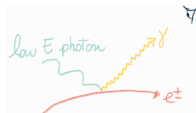
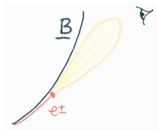
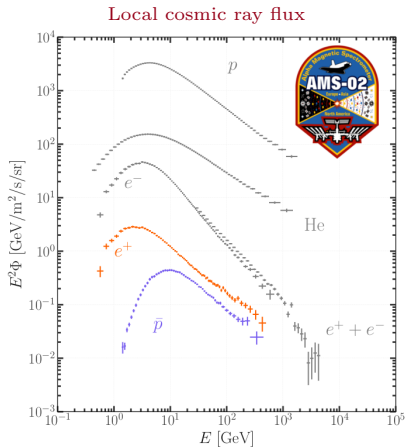
[NASA]



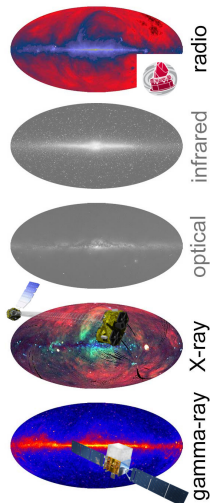
“Ultralight” DM   “Light” DM   WIMP   Composite DM   Primordial black holes

[1904.07915]

# Astroparticle messengers: precision era



## Multiwavelength photons



[NASA]



“Ultralight” DM “Light” DM WIMP Composite DM Primordial black holes

[1904.07915]

**Undergrad** Cagliari/Turin (IT) + CERN

*Theoretical physics of fundamental interactions*

**PhD** Turin + SLAC (US) [defended Jan'19]

*F. Donato, N. Fornengo*

**Postdoc** TTK RWTH Aachen (DE) [Oct'19-Nov'22]

*M. Kraemer, F. Kahlhoefer, J. Lesgourges*

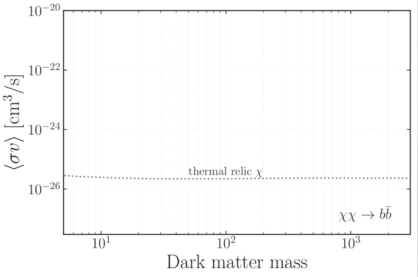
**Postdoc** LAPTh Annecy (FR) [Dec'22-Nov'24]

*F. Calore, P. Serpico, Y. Genolini*

*Marie Skłodowska-Curie fellowship: Signatures of cosmic rays and new fundamental particles in the Very high energy sky (VerSi)*



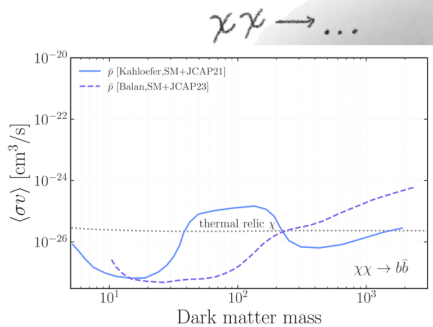
# I search for dark matter...



$\chi\chi \rightarrow \dots$



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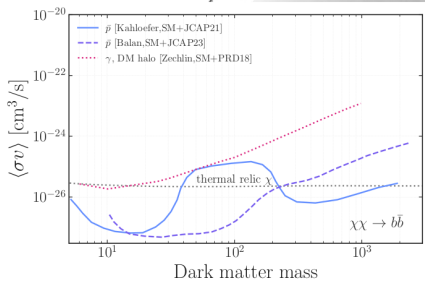


## Cosmic antimatter

[Kahlofer, SM+JCAP'21]

[Balan, SM+JCAP'23]

# I search for dark matter...



Cosmic  
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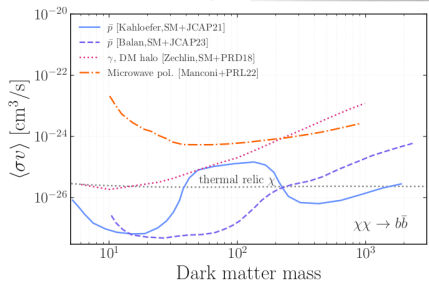
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Gamma rays  
at high latitudes

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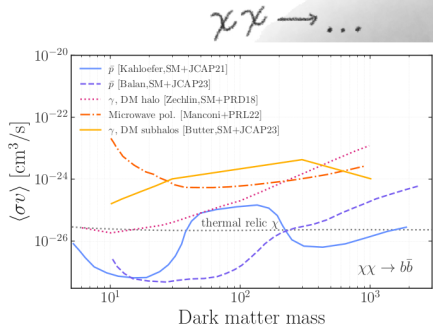
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Microwave  
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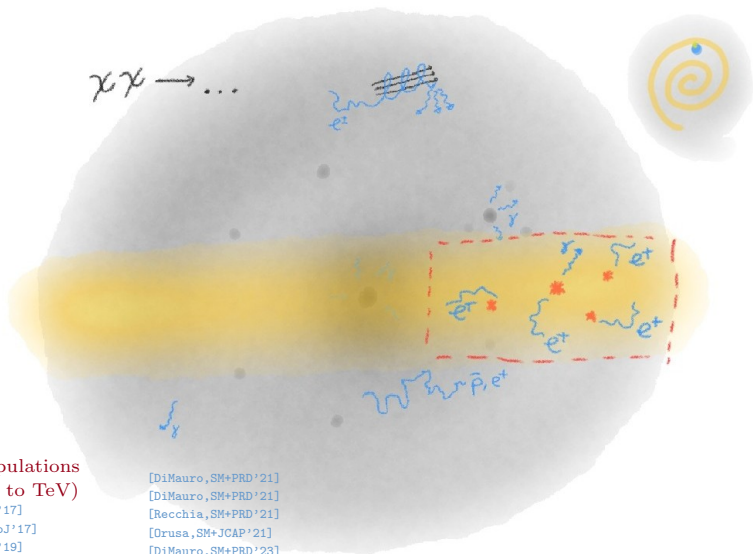
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Gamma rays  
from subhalos &  
machine  
learning

[Butter, SM+JCAP'23]



**Stellar populations**  
( $e^\pm$ , radio to TeV)

[Manconi+JCAP'17]

[DiMauro, SM+ApJ'17]

[Manconi+JCAP'19]

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[DiMauro, SM+PRD'20]

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[Recchia, SM+PRD'21]

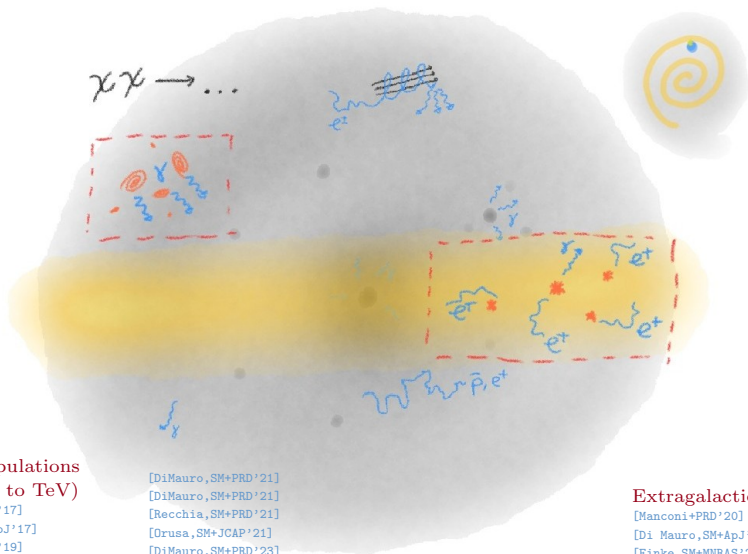
[Orusa, SM+JCAP'21]

[DiMauro, SM+PRD'23]

[Manconi+A&A'24]

[Orusa, SM+24]

# ...on top of astrophysical backgrounds



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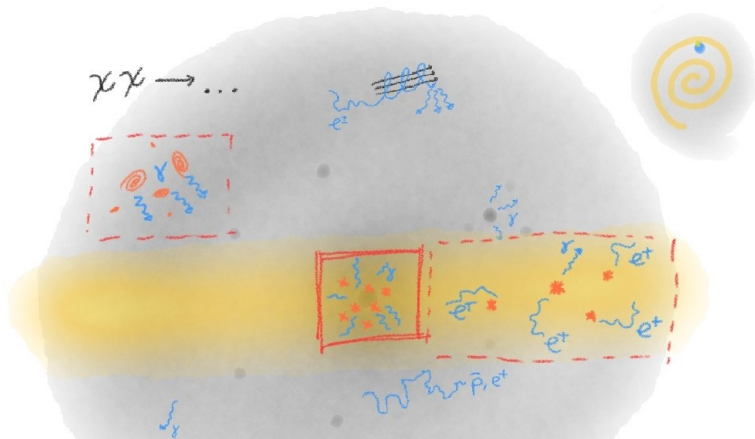
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## Extragalactic

[Manconi+PRD'20]  
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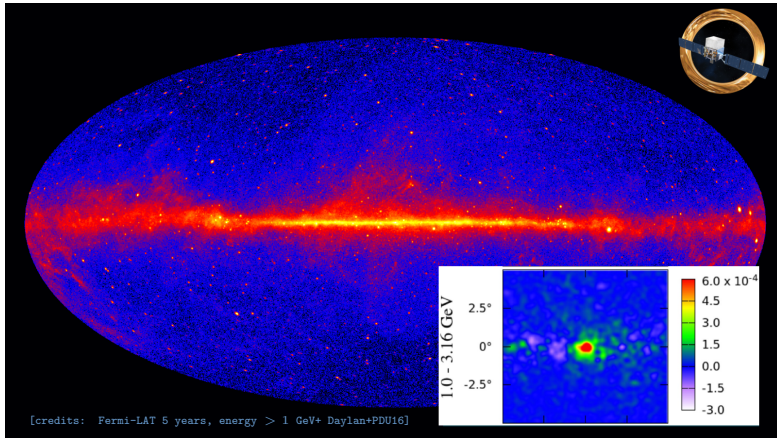
## Galactic center

[Calore, SM+PRL'21]  
[Manconi+PRD'24]

## Extragalactic

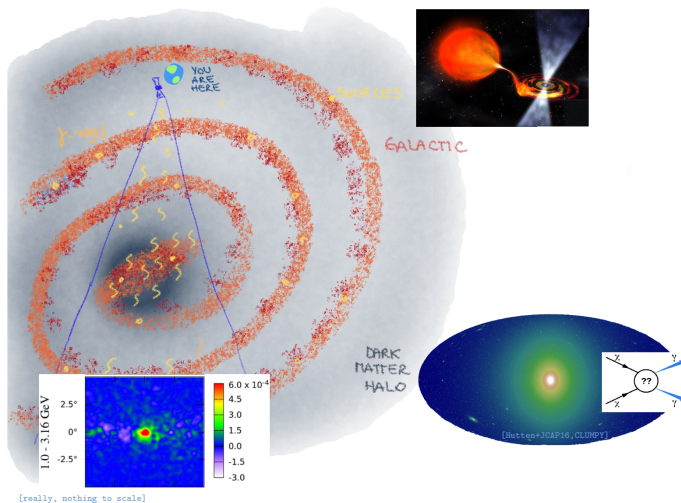
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# The Galactic Center Excess



Statistically significant excess in Fermi-LAT data  
with respect to diffuse emission and resolved astrophysical sources  
few % of 2-20 GeV inner Galaxy flux [hot topic... N+1 papers]

# Dark matter or stellar populations?



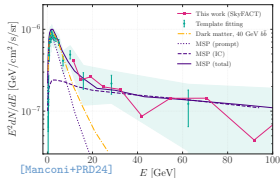
New physics phenomena:

- + **Dark Matter:** annihilation in Galactic halo
- + **Stellar:** population of new, faint sources

# Debated, likely stellar

*Significant community effort & debate: developing new methods, tackling bias from background diffuse emission mis-modeling*

## Energy spectrum



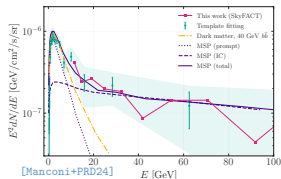
- High energy tail  
 $E > 10$  GeV  
[Linden+PRD'16, Manconi+PRD'24]
- Naturally explained by inverse Compton of  $e^\pm$  in stellar population
- Dark matter: more tuning



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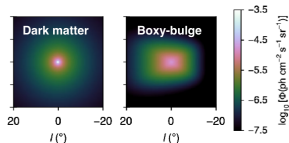
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## Morphology

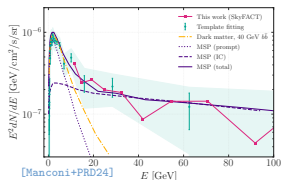


- Longitudinal asymmetry at  $\sim 10$  deg
- Stellar-like: independent evidence  
[Bartels+NA'18, Macias+NA'18, JCAP'19, Calore, SM+PRL'21, PRD'24, Song+MNRAS24]

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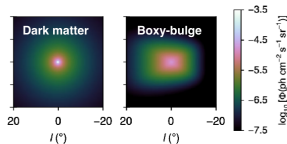
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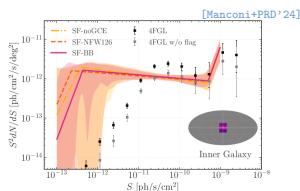
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## Statistics of photon counts

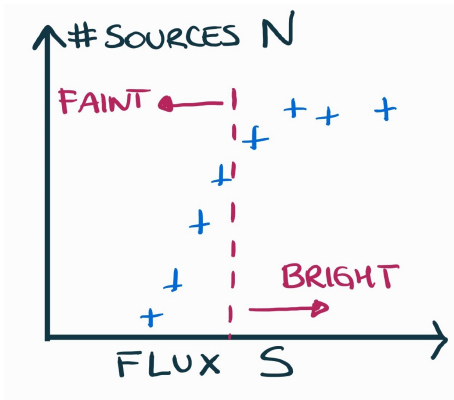


- Unresolved point sources: independent evidences  
[Buschmann+PRD'20, Calore, SM+PRL'21, List+PRD'21, Manconi+PRD'24]
- Source number depends on luminosity function [Dinsmore+JCAP'22]

**Corroborating an (at least) partial stellar origin of the excess**

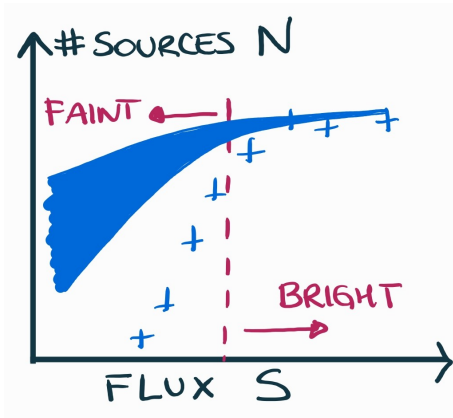
# Dissecting the Galactic center GeV excess

Photon count statistics: advanced methods to interpret Fermi-LAT data  
extracting crucial observables



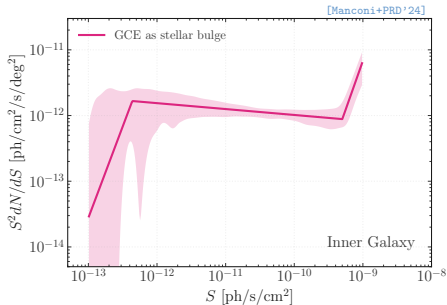
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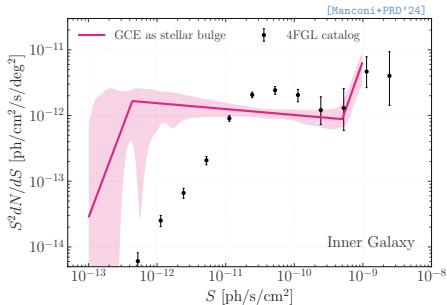
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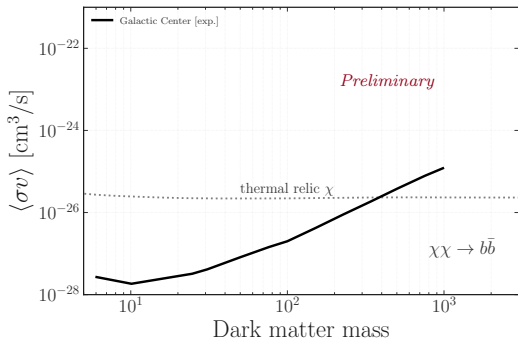
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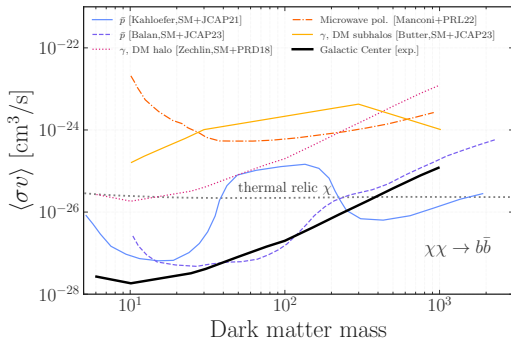


- Robustly addressing diffuse mismodeling, we revealed flux and spatial distribution of faint sources in the inner Galaxy [Calore, SM+PRL'21]
- Highest energy: energy spectrum, morphology & photon counts support (partial) stellar origin [Manconi+PRD'24]

Constraints to GeV dark matter: potentially strongest in GeV mass range

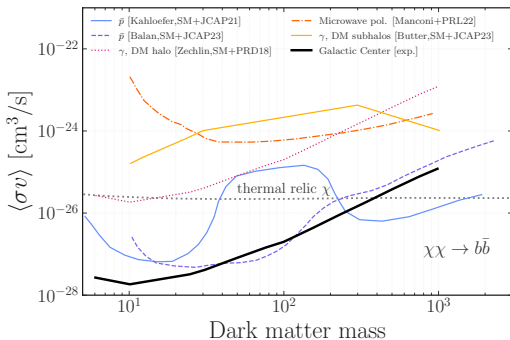


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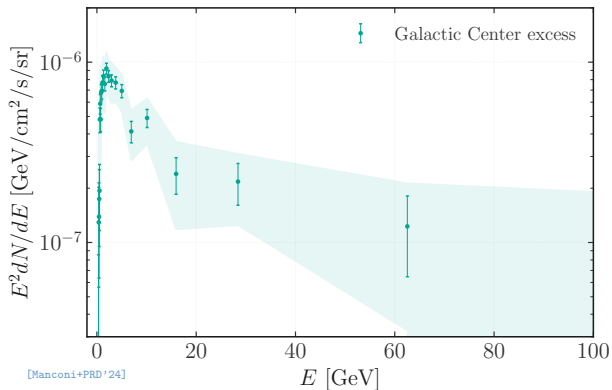
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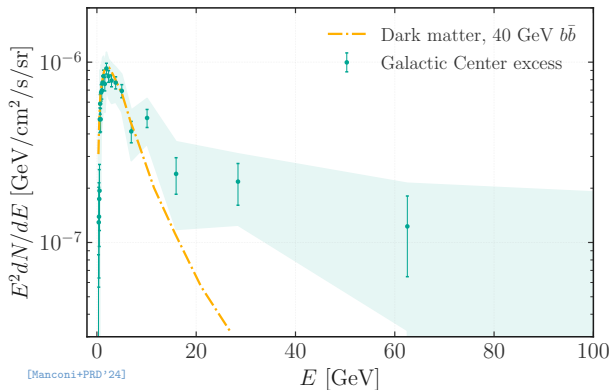
Next:

- Extend to nearby Galaxy (dwarf spheroidal, M31, LMC)
- Exploit X-ray and TeV current and forthcoming telescopes (eRosita, ART-XC, LHAASO, Cherenkov Telescope Array (CTA),...)

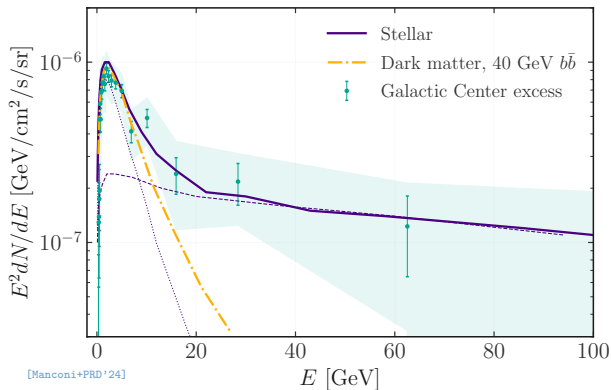
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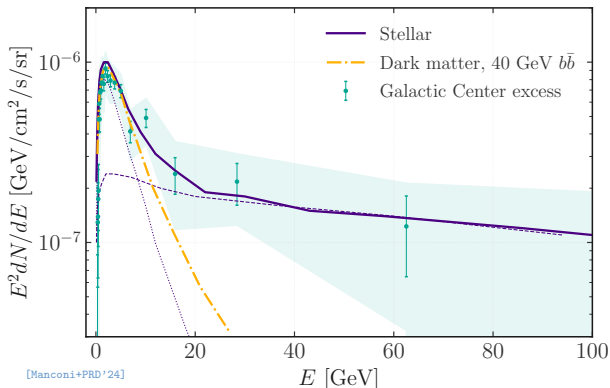


## *Nature of $> 10$ GeV gamma ray emission at the Galactic Center*



1. **Stellar population models:** particle ( $e^\pm$ ,  $p?$ ), multiwavelength
2. **Consequences for dark matter** in the Milky Way and in nearby galaxies

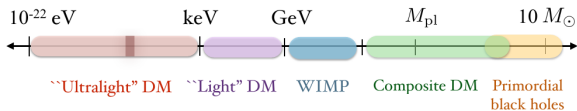
## *Nature of $> 10$ GeV gamma ray emission at the Galactic Center*



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using current and upcoming X-ray and TeV telescopes  
(eRosita, HAWC, Cherenkov Telescope Array (CTA),...)

## Probing broad range of dark matter theories



### Axion-like particles

CTA: signatures in TeV  $\gamma$ -ray spectra

Machine learning signal processing

[Manconi, Nippel, Kraemer, in prep]

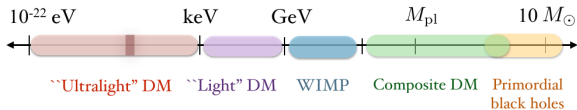
### Global fits of theories Beyond SM

legacy of GeV astroparticles

GAMBIT, including astro uncertainties

e.g. [Kahlofer, SM+21, Balan, SM+23]

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[Manconi, Nippel, Kraemer, in prep]

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e.g. [Kahlofer, SM+21, Balan, SM+23]



Antimatter asymmetry; X-rays, MeV to probe sub-GeV dark matter

## Phenomenology of dark matter & astroparticles with multi-messenger, multi-wavelength and machine learning methods



**Teaching:** Modélisation Numérique en Physique (L2, Jan-May 2025)  
*open to collaborate to future courses in astroparticles*

**Supervision:** [Postdoc opening](#) (deadline 16/12)+ [M2-PhD](#) with M.Cirelli  
*open to co-supervise L&M students*



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