

IRN Terascale @ IP2I Lyon 13-15
November 2024



First-order phase transitions from a broken horizontal $SU(2)$ gauge symmetry

Anna Chrysostomou^(LPTHE), Alan Cornell^(UJ), Aldo Deandrea^(IP2I), Luc Darmé^(IP2I), Thibault Demartini^(IP Paris)



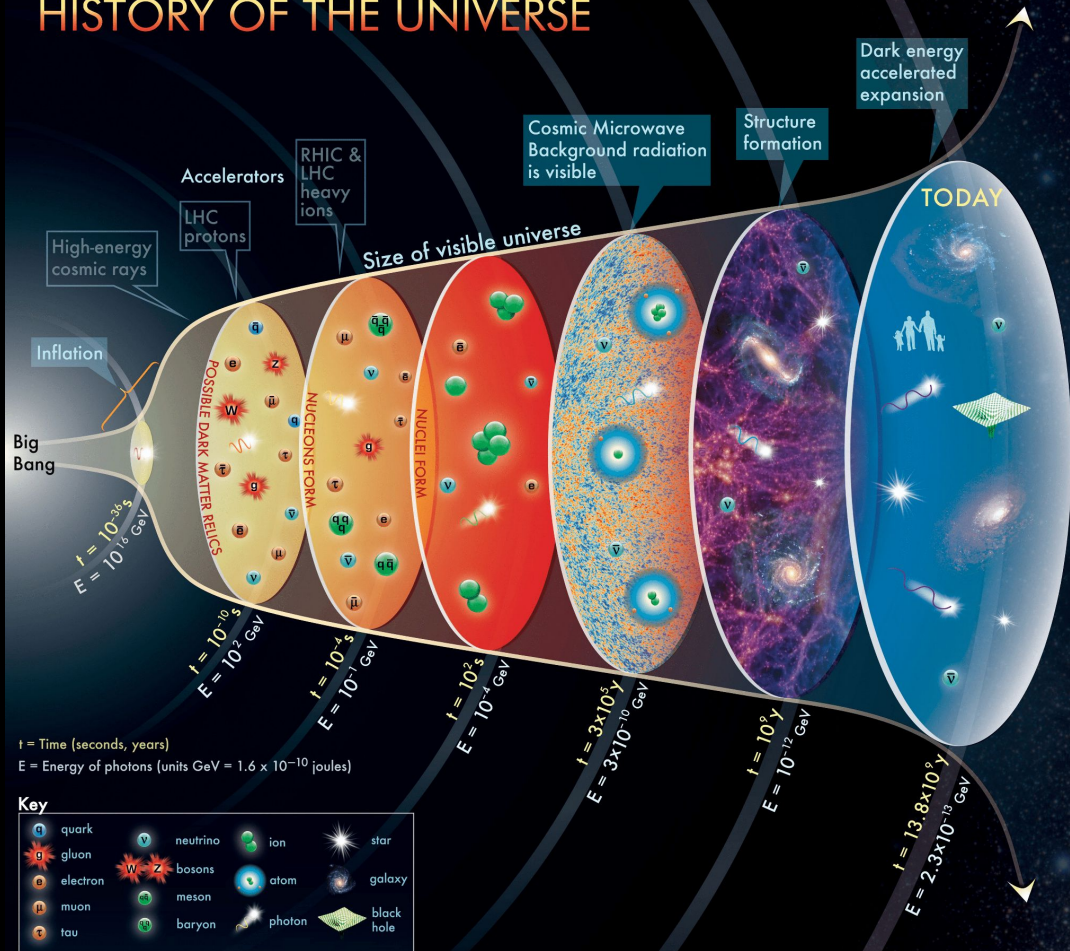
Overview

- Particle physics in the GW era: prospects & procedures
 - Understanding the first order phase transition
- (Breaking) a new horizontal $SU(2)$ flavour symmetry
- Building the finite-temperature effective potential
 - “Parwani”/Truncated Full Dressing
 - Dimensional Reduction
- Is there a first-order phase transition?

What do gravitational waves have to offer particle physicists?



HISTORY OF THE UNIVERSE

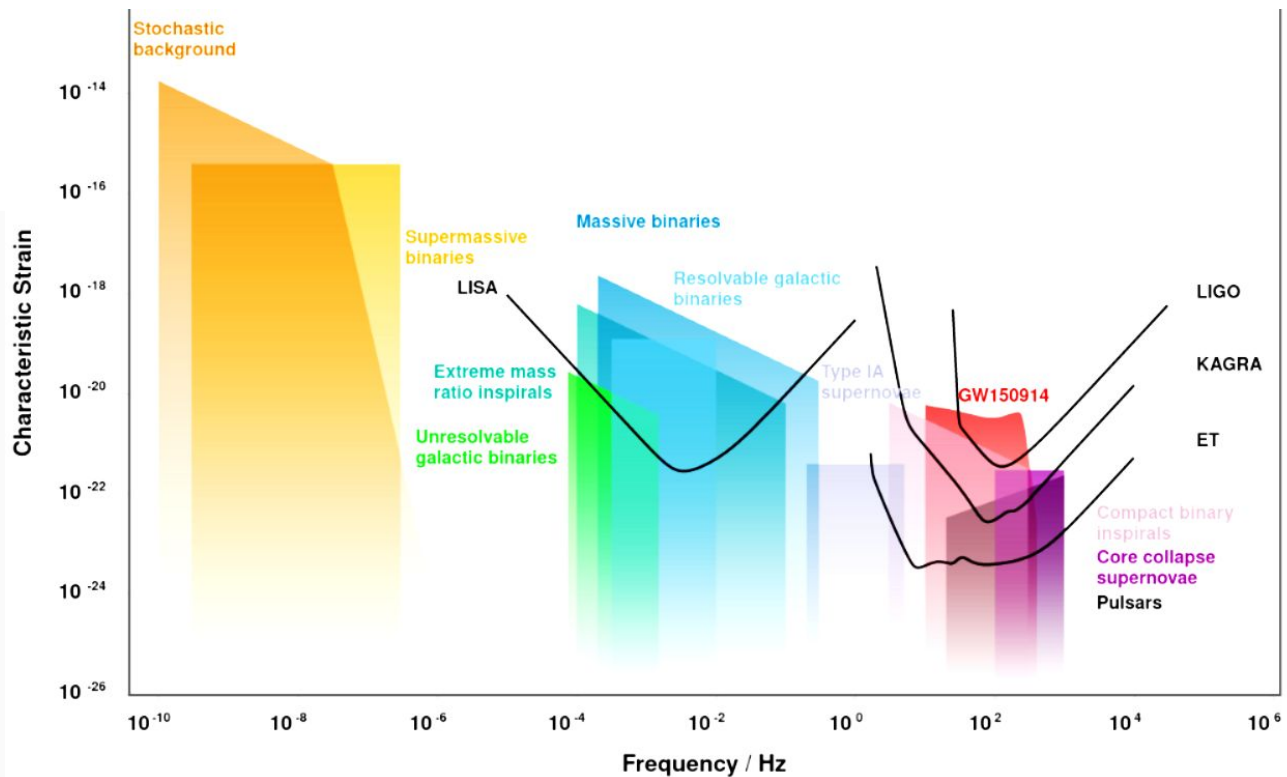


The concept for the above figure originated in a 1986 paper by Michael Turner.

The allure of gravitational waves

- *explore TeV-scale physics*
- *indirect constraints from gravitational wave experiments (complement collider experiments)*
- *early universe phase transitions \Rightarrow insights into fundamental physics e.g. symmetry breaking*
- *Phase transitions: QCD (~ 100 MeV), EW (~ 100 GeV)*
 - *Baryogenesis + baryon asymmetry, **EWSB FOPT** \leftrightarrow **BSM***
- *Inflation ($\sim 10^{13}$ TeV)*
- *Exotic: cosmic strings, primordial black holes, Planck scale*
- *Plus tests of GR : > 2 polarisation states, modified dispersion relation, sub- or super-luminal propagation, etc.*

GW experiments



Algorithm



Build $V_{\text{eff}}(\mu, T)$

Determine field content, dof, etc.
 Potential: zero-T + finite-T
 Find degenerate $\min\{V_{\text{eff}}(\mu, T)\}$

Is the PT first order ?

$$\frac{\phi_c}{T_c} \geq 1$$

Compute PT parameters

Compute Euclidean / 3d action
 Extract phase transition parameters:
 > PT strength α
 > Inverse of PT duration β/H_*
 > Bubble wall speed v_w

GW spectrum

Compute energy density of GWs as
 a function of frequency, based on
 PT parameters

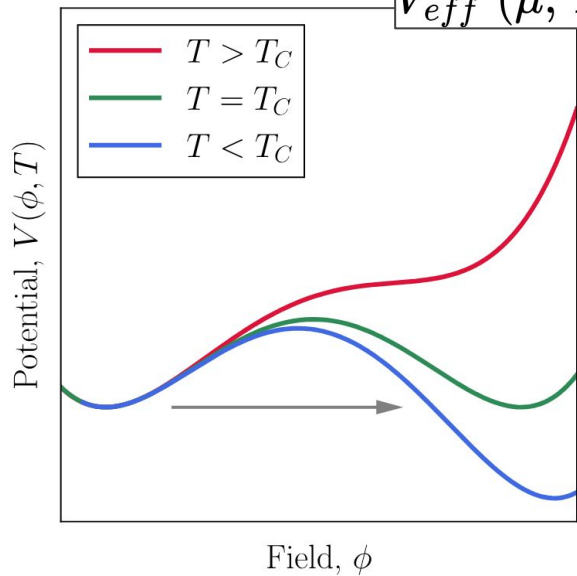
$$h^2 \Omega_{\text{GW}}(f; H_*, \alpha, \beta, v_w)$$

Sensitivity of detectors

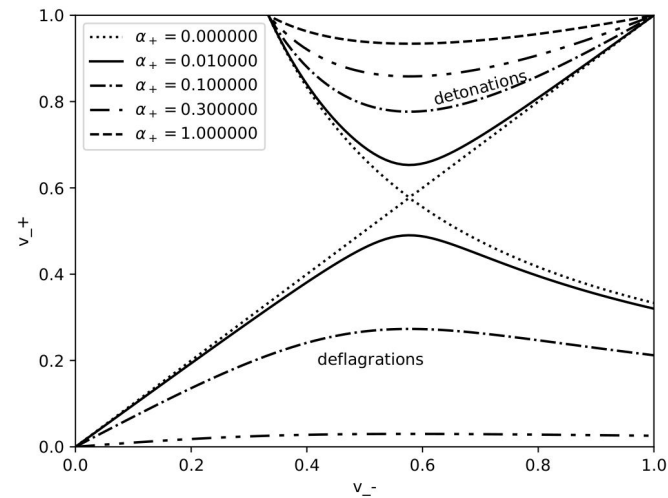
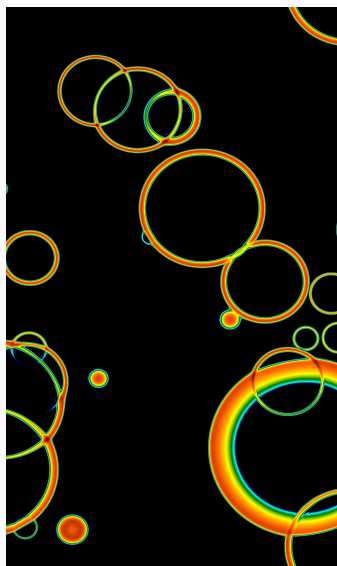
Compare GW power spectrum
 against detector sensitivity curve

Algorithm

$$V_{eff}(\mu, T) = V_{tree} + V_{CW} + V_T$$

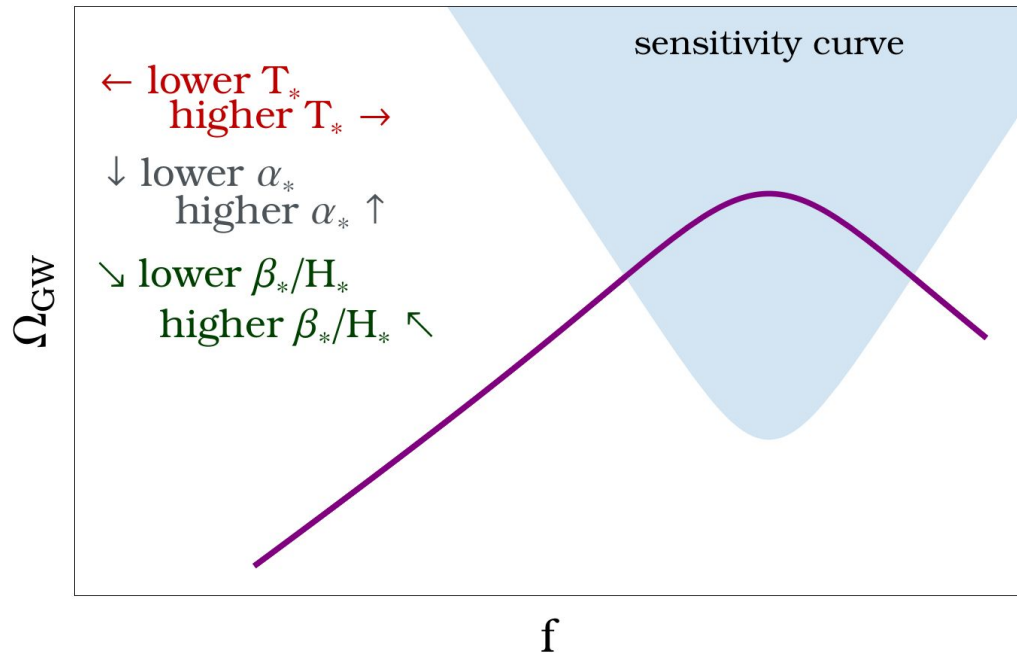


$$\frac{\phi_c}{T_c} \geq 1$$

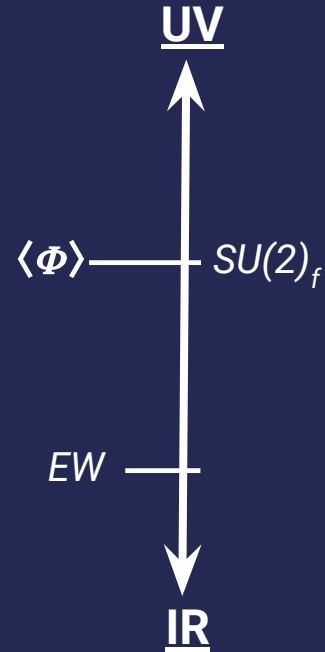


GW spectrum

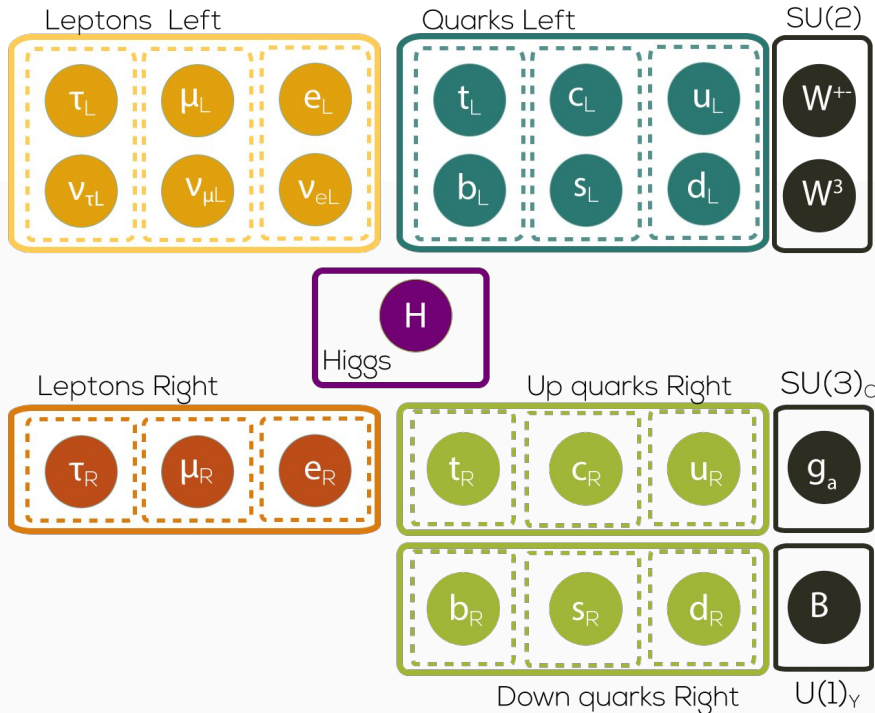
Piecewise function: broken power law joined at f_{peak}



Breaking a new « horizontal gauge symmetry » in the flavour sector



The « flavour-transfer » mechanism



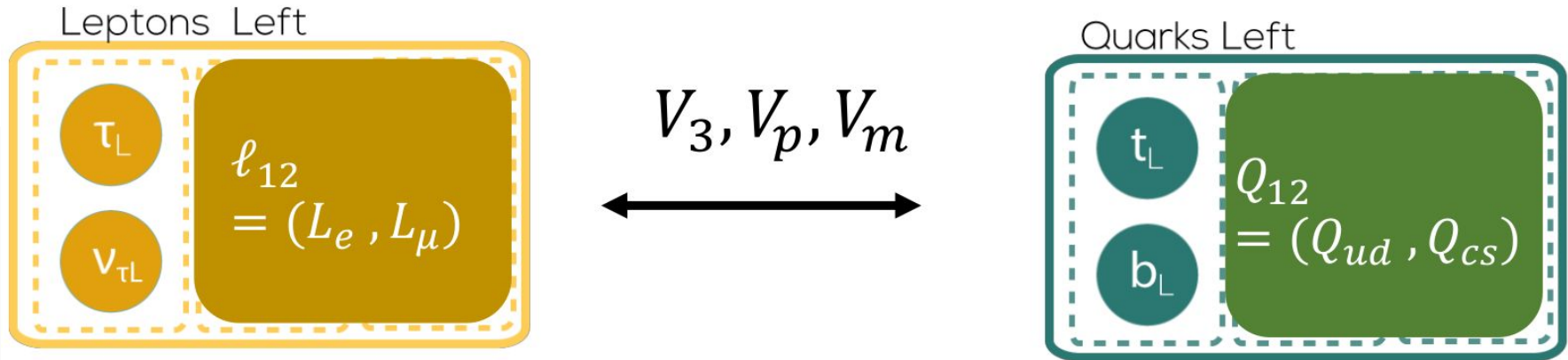
+ $SU(2)_f$

3 new « W-like » gauge bosons carrying a « flavour-charge »

$$M_{V_1}^2 = M_{V_2}^2 = M_{V_3}^2 = \frac{g_f}{2} \sum_i v_\phi^2$$

+ *rotation matrices to mass basis: V_{uL}, V_{dL}, \dots*

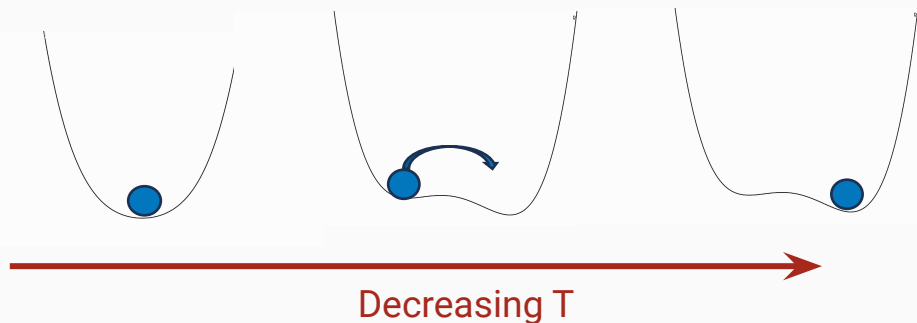
The « flavour-transfer » mechanism



*rather than break flavour, the new gauge bosons transfer flavour
from one fermionic sector to another*

True and false vacua

- To break the flavour gauge symmetries we need the appearance of a VEV for the new scalars
 - This occurs in the early universe at temperatures close to the VEV
- Flavour constraints point towards 100 TeV scale for the complete flavourful theory



~ 100 TeV
 $SU(2)_f$ breaking
by new scalar Φ

~ 0.2 TeV
EW breaking

$SU(2)_f$ and
 $SU(2)_W \times U(1)_Y$
symmetric theory

flavour bosons

$SU(2)_W \times U(1)_Y$
symmetric theory

EW bosons

$U(1)_{em}$ symmetric
theory

T

*Building the finite-temperature
effective potential: Truncated Full
Dressing vs
Dimensional Reduction*

$$\mathbf{TFD}: V_{\text{eff}}(\mu, T) \rightarrow V_{\text{eff}}(\mu + \pi T, T)$$

$$\mathbf{DR}: V_{4\text{deff}}(\mu, T) \rightarrow V_{3\text{deff}}(\mu_3, T)$$

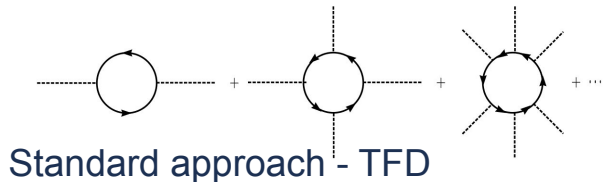
Thermal corrections : TFD vs DR

Quiros 1999,
Curtin 2006

How to compute the effective thermal potential ?

- Describe the correlation functions a QFT in a thermal bath, Greens functions can be computed **by compactifying time along the imaginary direction**
- Stability of the vacuum be estimated from this quantity (equivalent to free energy in thermodynamics)

Stay in 4D, every loop comes with an infinite sum from the modes in along the imaginary time direction



Integrate out the modes from the compactified dimension and match the 4D theory to a 3D theory

Dimensional Reduction approach (EFT-like)

More modern approach, partially automatised through DRalgo

Effective potential: Truncated Full Dressing (Parwani)

$$V_{\text{tree}}(\phi) = -\frac{1}{2}\mu_\phi^2\phi^2 + \frac{1}{4}\lambda_\phi\phi^4 + \frac{1}{2}\mu_s^2|s|^2 + \frac{1}{4}\lambda_s|s|^4 + \frac{1}{2}\lambda_{\phi s}\phi^2|s|^2$$

Field	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$SU(2)_f$	DoF
χ	1	2	1/2	1	3
Φ	1	1	0	2	1
S	3	1	2/3	2	$3 \times 2 \times 2 = 12$
V	1	1	0	3	$3 \times 3 = 9$

$$V(\phi, T) = V_{\text{tree}}(\phi) + V_{\text{CW}}(\phi) + V_T(\phi, T)$$

$$V_{\text{CW}}(\phi) = \sum_{i=\phi,\chi,f,s} \pm \frac{n_i}{64\pi^2} m_i^4 \left[\log \left\{ \frac{m_i^2}{\mu^2} \right\} - C_i \right]$$

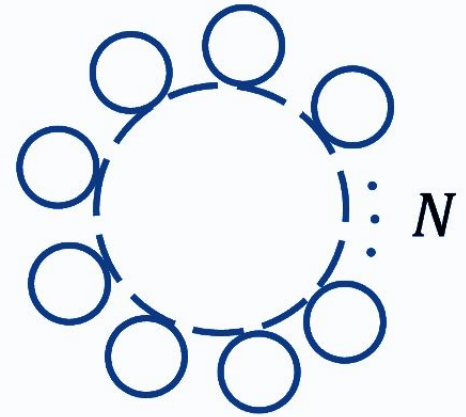
Gives the usual log-like Coleman Weinberg terms

$$V_T(\phi, T) = \sum_i \frac{n_i T^4}{2\pi^2} J_B \left(\frac{m_i^2}{T^2} \right) + \sum_i \frac{n_i T^4}{2\pi^2} J_F \left(\frac{m_i^2}{T^2} \right)$$

$$J_{B/F}(a) = \pm \int_0^\infty dy y^2 \log \left[1 \mp e^{-\sqrt{y^2+a}} \right]$$

Effective potential: Truncated Full Dressing (Parwani)

$$m^2(\phi) = m_{\text{tree}}^2(\phi) + \Pi(\phi, T) \xrightarrow{\sim g^2 T^2}$$



$$V_{\text{CW}}(\phi) = \sum_{i=\phi,\chi,f,s} \pm \frac{n_i}{64\pi^2} m_i^4 \left[\log \left\{ \frac{m_i^2}{\mu^2} \right\} - C_i \right]$$

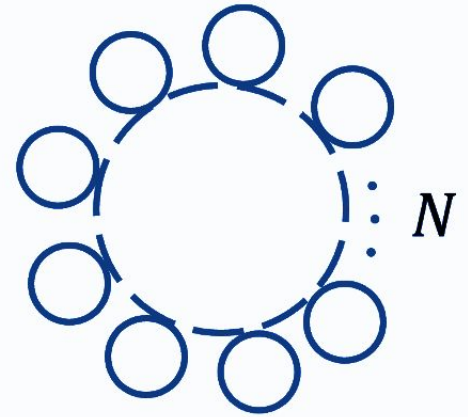
$$V_T(\phi, T) = \sum_i \frac{n_i T^4}{2\pi^2} J_B \left(\frac{m_i^2}{T^2} \right) + \sum_i \frac{n_i T^4}{2\pi^2} J_F \left(\frac{m_i^2}{T^2} \right)$$

$$\pi_\phi = \pi_\chi = \frac{\lambda_\phi}{2} + \frac{9}{48} g_f^2$$

$$\pi_f^L = \frac{3}{2} g_f^2$$

Effective potential: Truncated Full Dressing (Parwani)

$$m^2(\phi) = m_{\text{tree}}^2(\phi) + \Pi(\phi, T) \sim g^2 T^2$$



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Effective potential: Truncated Full Dressing (Parwani)

Multiple sources of theoretical uncertainty :

- Nonperturbativity (IR modes at high T) [Linde 1980]
- Inconsistencies (non-negligible $\text{Im}\{V\}$) [Weinberg & Wu 1987; Weinberg 1992]
- higher-order perturbative corrections [Arnold & Espinosa 1992]
- gauge dependence [Laine 1994]
- renormalisation scale dependence [Farakos *et al.* 1994]

See Croon *et al.* JHEP 04 (2021) 055

3D EFT approach - mitigating errors

At zero temperature, the one-loop effective potential is renormalisation group invariant

$$\frac{d}{d \log \mu} (V_{\text{tree}} + V_{1\text{-loop}}) = 0.$$

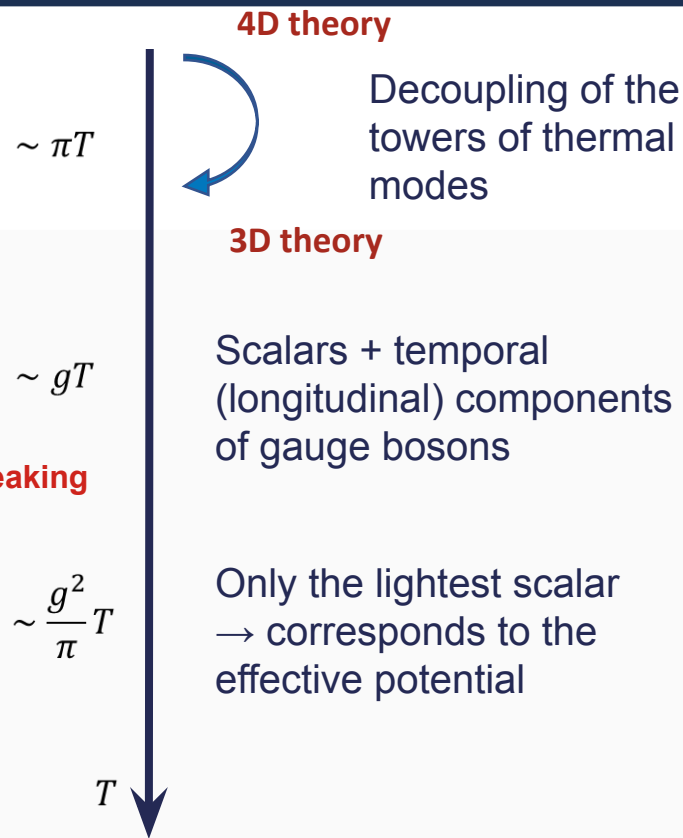
But, at high temperatures this fails, even at leading order

$$\frac{d}{d \log \mu} (V_{\text{tree}} + V_{1\text{-loop}}^{\text{thermal}}) \neq 0.$$

Symmetry breaking

The problem can be traced to the scale hierarchy $\pi T \gg m$, and to

$$\frac{d}{d \log \mu} \left(\frac{1}{2} \Pi_T \phi^2 \right).$$



3D EFT approach - mitigating errors

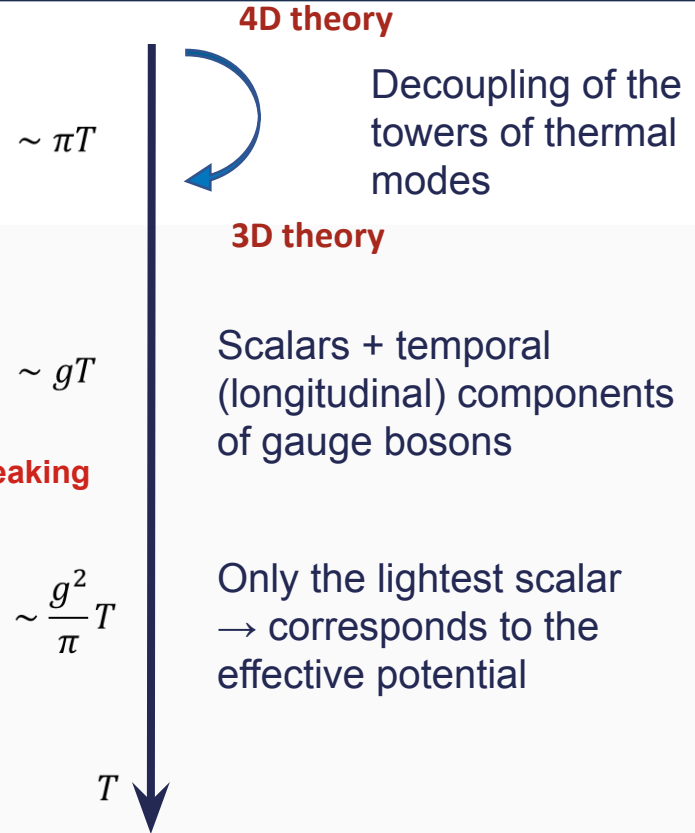
Step-by-step approach to decouple all thermal DoF

1. RGE from μ_{ini} to μ_{hard}
2. Match 4d to 3d at « hard scale »
 $\mu_{\text{hard}} \sim \pi T$ (thermal mass of fermions + transverse gauge bosons)
3. Run gT in the 3d theory
4. Decouple remaining bosonic modes, except scalar field ϕ triggering the PT

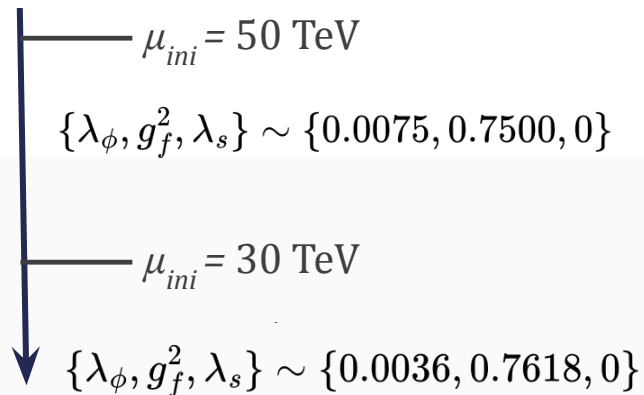
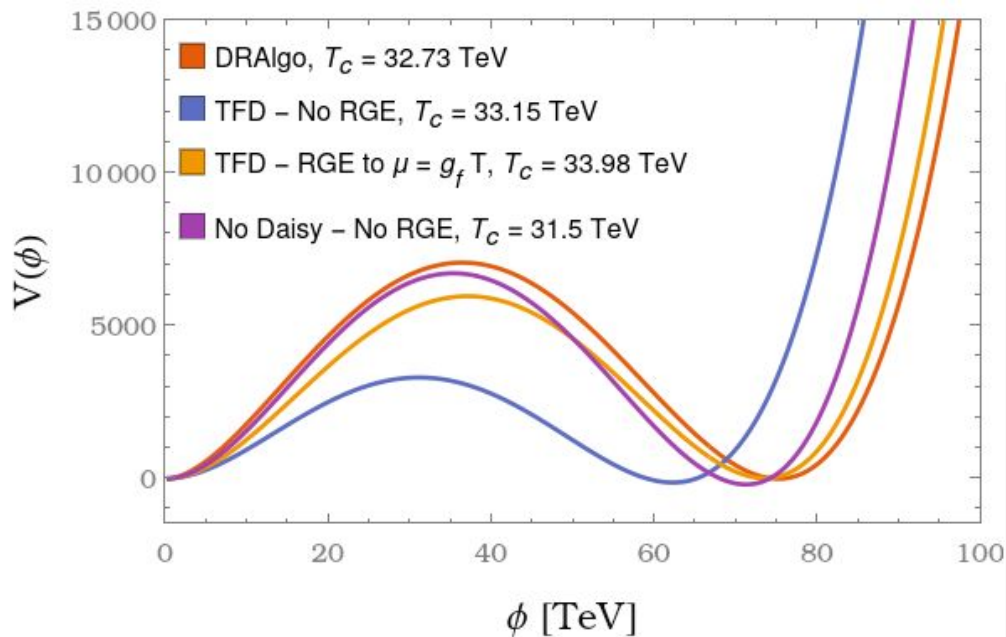
Implement using DRalgo

Up to NNLO matching in some cases !

Symmetry breaking

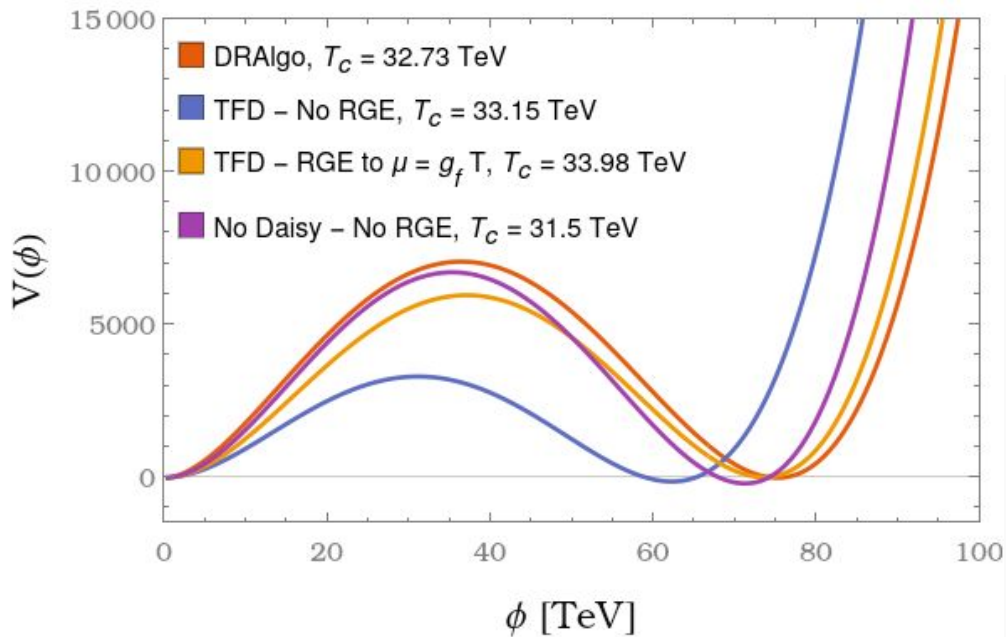


Compare against DRalgo (4d)



$$V_{4d} = V_{3d} \times T$$

Compare against DRalgo (4d)



Preliminary results

■ $\frac{\phi_c}{T_c} \sim 2.30$

■ $\frac{\phi_c}{T_c} \sim 1.88$

■ $\frac{\phi_c}{T_c} \sim 2.16$

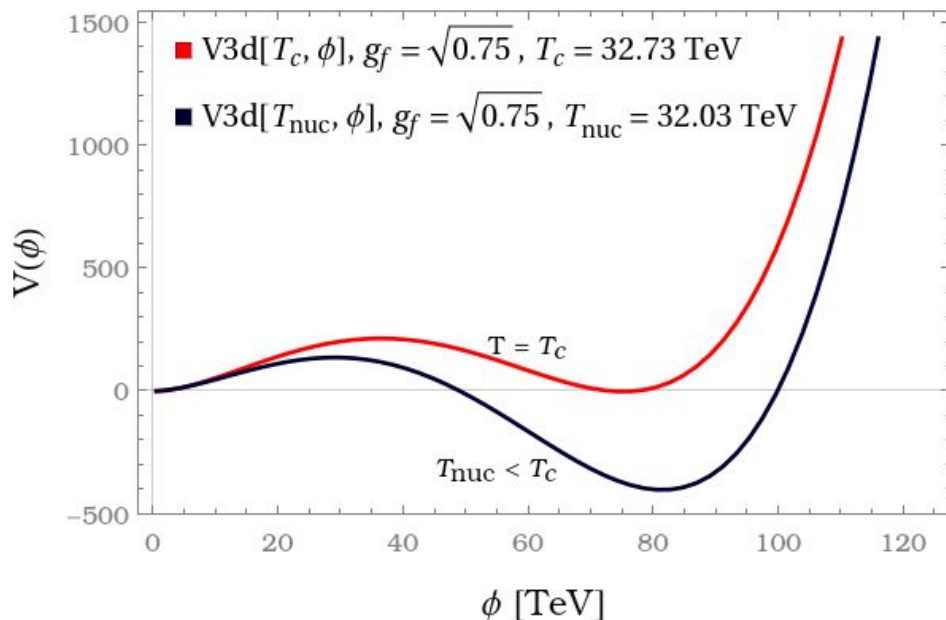
■ $\frac{\phi_c}{T_c} \sim 2.26$

*Consistently first order,
of similar strength*

Phase transition parameters using DRalgo (3d)

Preliminary results

For $\lambda_\phi = 0.0075$, $M_\phi = 10\sqrt{2}$ TeV, $\mu_\phi = -M_\phi^2/2$



$$\frac{\beta_*}{H_*} \approx 17553.10$$

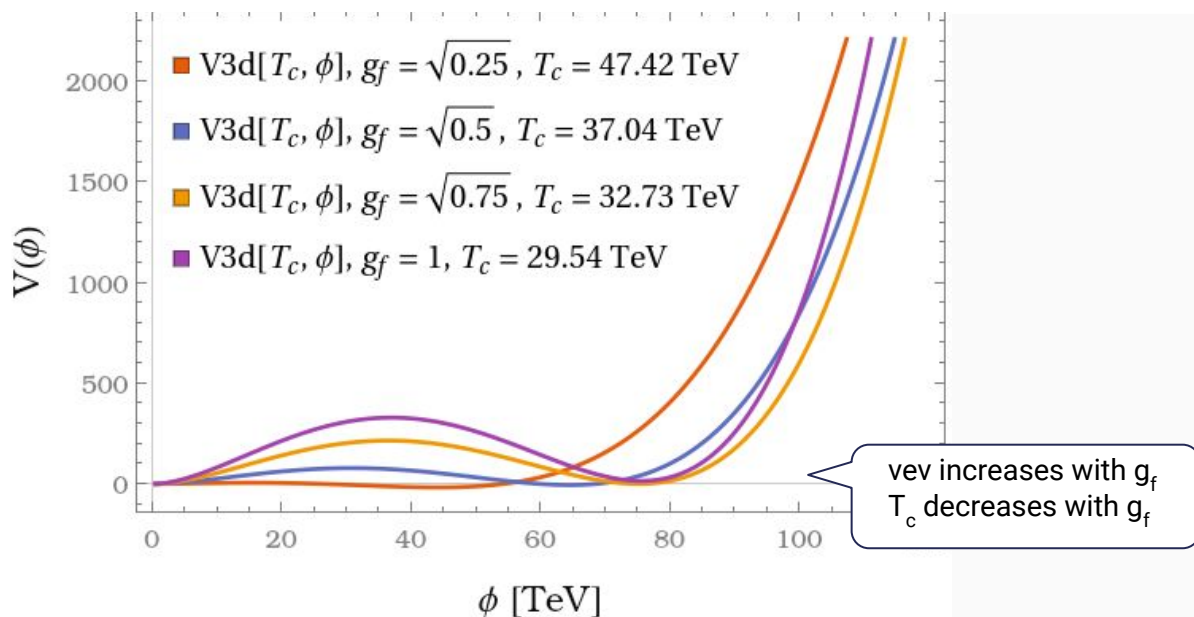
$$\alpha \approx 0.0132$$

$$v_w \sim 1$$

Phase transition parameters using DRalgo (3d)

Preliminary results

For $\lambda_\phi = 0.0075$, $M_\phi = 10\sqrt{2}$ TeV, $\mu_\phi = -M_\phi^2/2$



■ $\frac{\phi_c}{T_c} \sim 0.93$

■ $\frac{\phi_c}{T_c} \sim 1.74$

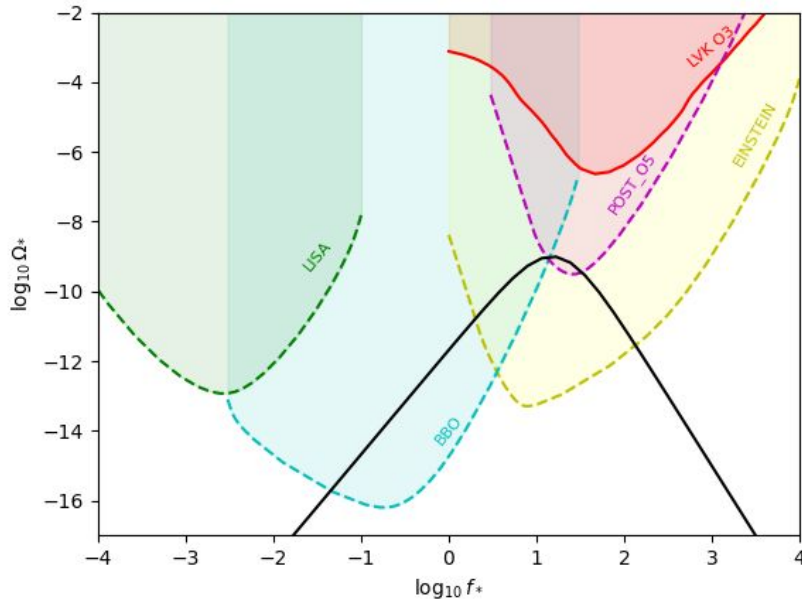
■ $\frac{\phi_c}{T_c} \sim 2.30$

■ $\frac{\phi_c}{T_c} \sim 2.56$

No FOPT if g_f is too small !

Expected GW spectrum

Preliminary results



LIGO prospects : not great

Probable causes :

- very small phase transition strength*
- large inverse-duration (i.e. short PT)*

Need to understand fully to resolve

Einstein & future detectors still viable

Conclusions

- Most models of flavour relies on broken symmetries to create the observed patterns in the SM-Higgs Yukawa couplings
- For flavour gauge symmetries, this means introducing new Higgs-like scalars, that can undergo first order phase transitions in the early universe
- Cooler phase transition for heavier flavour bosons
- Ongoing work: to finalise the effective potential based on two different approaches
 - Still discrepancies to be ironed out / understood
- The temperature range corresponding to actual flavour constraints matches the realm of LIGO/Einstein telescope range (if the PT can be made strongly-enough first order)
 - Remain: hydrodynamics simulation to improve GW spectrum predictions for our $SU(2)_f$ model

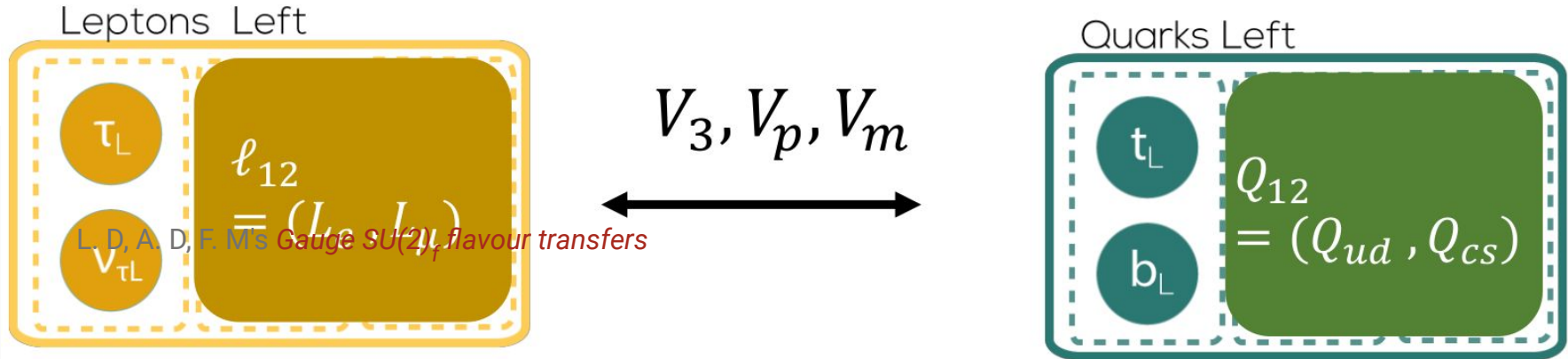
Thanks!

Any questions?

*Or write to us via
chrysostomou@lpthe.jussieu.fr*



The « flavour-transfer » mechanism



rather than break flavour, the new gauge bosons transfer flavour
from one fermionic sector to another

$$V_3 = \begin{pmatrix} 1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 0 \end{pmatrix} \quad V_p = \begin{pmatrix} 0 & 1 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \quad V_m = \begin{pmatrix} 0 & 0 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \quad \text{The corresponding generators in flavour space}$$

Temperature corrections

$$V_T(\phi, T) = \sum_i \frac{n_i T^4}{2\pi^2} J_B \left(\frac{m_i^2}{T^2} \right) + \sum_i \frac{n_i T^4}{2\pi^2} J_F \left(\frac{m_i^2}{T^2} \right)$$

$$J_{B/F}(a) = \pm \int_0^\infty dy y^2 \log \left[1 \mp e^{-\sqrt{y^2+a}} \right]$$

$$J_{B,F}^{low}(a) \approx -\sqrt{\frac{\pi}{2}} a^{3/4} e^{-\sqrt{a}} \left(1 + \frac{15}{8} a^{-1/2} + \frac{105}{128} a^{-1} \right)$$

$$J_B^{high}(a) \approx -\frac{\pi^4}{45} + \frac{\pi^2}{12} a - \frac{\pi}{6} a^{3/2} - \frac{a^2}{32} (\log(a) - c_B)$$

$$J_B(a) \approx e^{-\left(\frac{a}{6.3}\right)^4} J_B^{high}(a) + \left(1 - e^{-\left(\frac{a}{6.3}\right)^4} \right) J_B^{low}(a)$$

Uncertainties

$\Delta\Omega_{\text{GW}}/\Omega_{\text{GW}}$	4d approach	3d approach
RG scale dependence	$\mathcal{O}(10^2 - 10^3)$	$\mathcal{O}(10^0 - 10^1)$
Gauge dependence	$\mathcal{O}(10^1)$	$\mathcal{O}(10^{-3})$
High- T approximation	$\mathcal{O}(10^{-1} - 10^0)$	$\mathcal{O}(10^0 - 10^2)$
Higher loop orders	unknown	$\mathcal{O}(10^0 - 10^1)$
Nucleation corrections	unknown	$\mathcal{O}(10^{-1} - 10^0)$
Nonperturbative corrections	unknown	unknown

Sources of theoretical uncertainty and relative importance quantified by the parameter $\Delta\Omega_{\text{GW}}/\Omega_{\text{GW}}$ over the range $M = \{580 - 700\}$ GeV in the SMEFT. Although we do not have reliable estimates for the uncertainties of the 4d approach due to higher loop orders and nucleation corrections, they are expected to be much larger than the corresponding uncertainties of the 3d approach

Power counting

To illustrate next-to-leading order dimensional reduction, we consider a schematic model with scalar mass parameter μ^2 , scalar quartic coupling λ , and gauge coupling g . Given the power counting $\mu^2 \sim g^2 T^2$, $\lambda \sim g^2$, the matching of the mass parameter is

$$\begin{aligned} \bar{\mu}_3^2 = & \underbrace{\mu^2}_{\mathcal{O}(g^2)} + \underbrace{\#g^2 T^2}_{\mathcal{O}(g^2)} + \underbrace{\#g^2 \mu^2}_{\mathcal{O}(g^4)} + \underbrace{\#g^4 T^2}_{\mathcal{O}(g^4)} + \mathcal{O}(g^6) \\ & + \underbrace{\#g^2 m_D}_{\mathcal{O}(g^3)} + \underbrace{\#g^4}_{\mathcal{O}(g^4)} + \mathcal{O}(g^5), \end{aligned} \tag{1.3}$$

where the first line (with even powers of g) results from the first step, and the second line (with odd power of g) from second step of the dimensional reduction. In practice, *full* $\mathcal{O}(g^4)$ contributions are included. Going to higher orders, requires a three-loop computation for both steps of the dimensional reduction. The situation is similar for the coupling:

Power counting

$$\begin{aligned}\bar{\lambda}_3 = & \boxed{\begin{array}{c} \text{tree-level} \\ T\lambda \\ \mathcal{O}(g^2) \end{array}} + \boxed{\begin{array}{c} \text{1-loop} \\ \#g^4 \\ \mathcal{O}(g^4) \end{array}} + \mathcal{O}(g^6) \\ & + \boxed{\begin{array}{c} \text{1-loop} \\ \# \frac{g^4}{m_D} \\ \mathcal{O}(g^3) \end{array}} + \boxed{\begin{array}{c} \text{2-loop} \\ \# \frac{g^6}{m_D^2} \\ \mathcal{O}(g^4) \end{array}} + \mathcal{O}(g^5). \end{aligned} \tag{1.4}$$

Power counting

$$V_{\text{eff}}^{3\text{d}} = \underbrace{V_{\text{tree}}^{3\text{d}}}_{\mathcal{O}(g^2)} + \underbrace{V_{1\text{-loop}}^{3\text{d}}}_{\mathcal{O}(g^3)} + \underbrace{V_{2\text{-loop}}^{3\text{d}}}_{\mathcal{O}(g^4)} + \mathcal{O}(g^5) . \quad (1.5)$$