

# Hyper-Kamiokande status

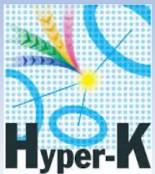


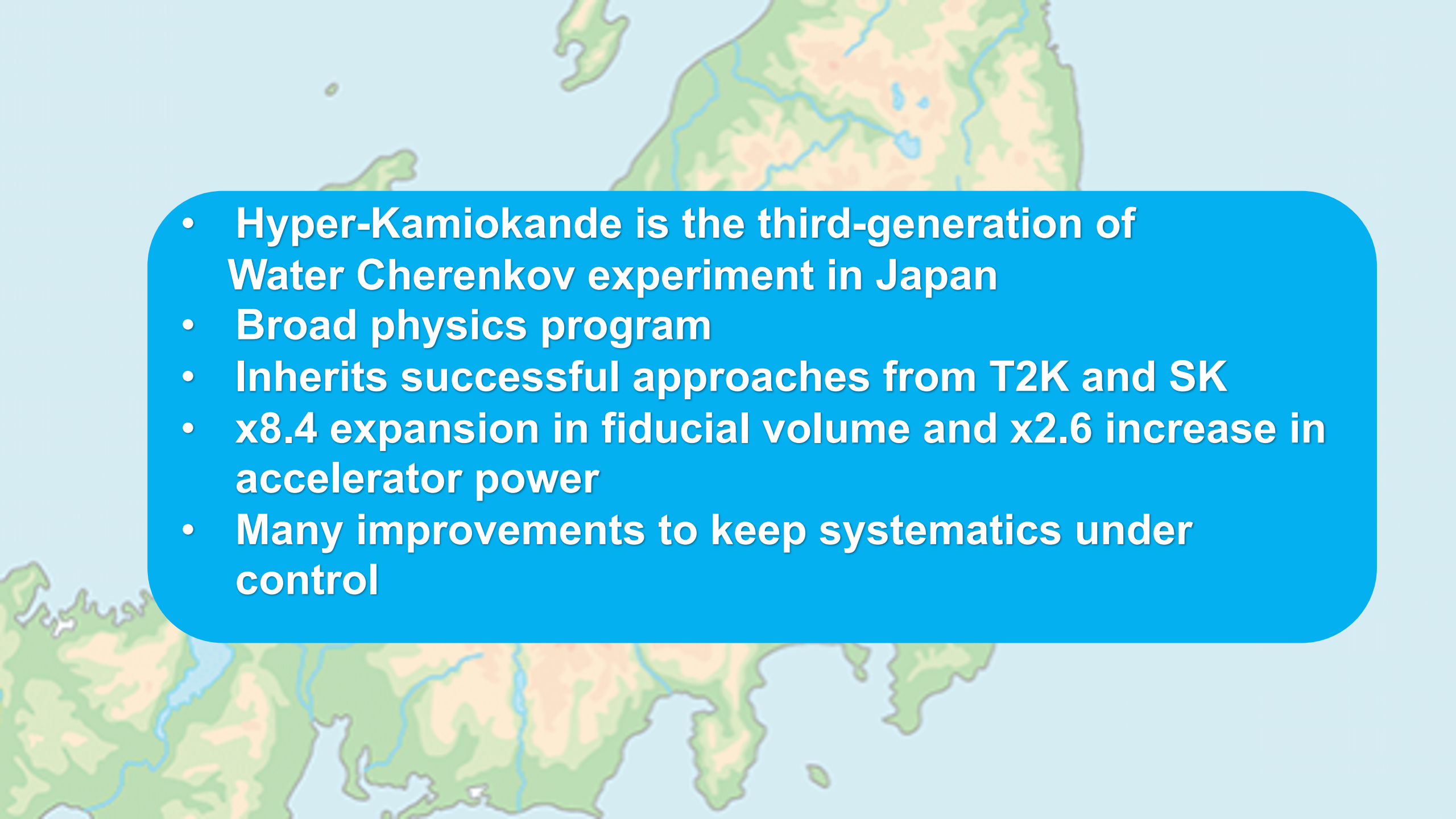
**Denis Carabadjac**

dcarabadjac@llr.in2p3.fr

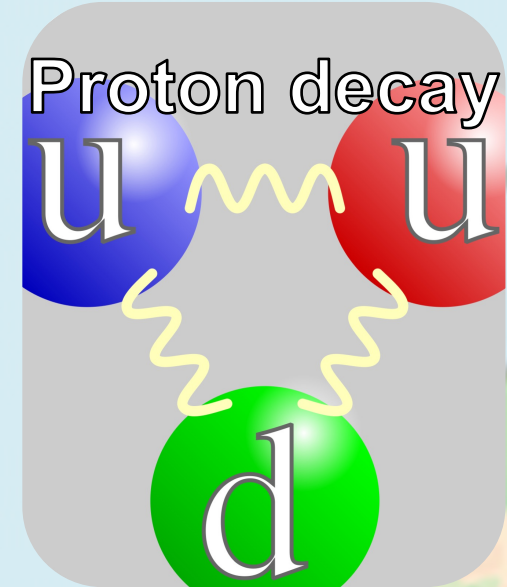
**IRN meeting**

10.10.2024



- 
- A topographic map of Japan is visible in the background, showing the main islands and surrounding waters. The map uses green and brown colors to represent elevation and terrain.
- **Hyper-Kamiokande is the third-generation of Water Cherenkov experiment in Japan**
  - **Broad physics program**
  - **Inherits successful approaches from T2K and SK**
  - **x8.4 expansion in fiducial volume and x2.6 increase in accelerator power**
  - **Many improvements to keep systematics under control**

# Hyper-K has an extensive physics program



- GUT probe



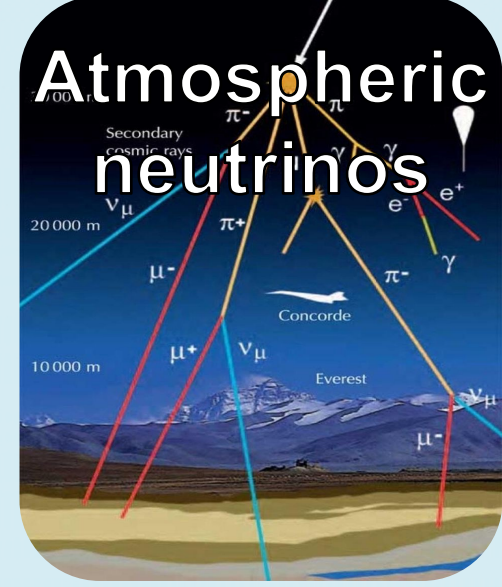
- MSW effect
- Hep neutrino search
- Day-night asymmetry



- SN neutrino detection
- Dark matter indirect detection



- Precise measurements of neutrino oscillations
  - Searching for CP-violation → matter-antimatter asymmetry
  - Neutrino mass ordering



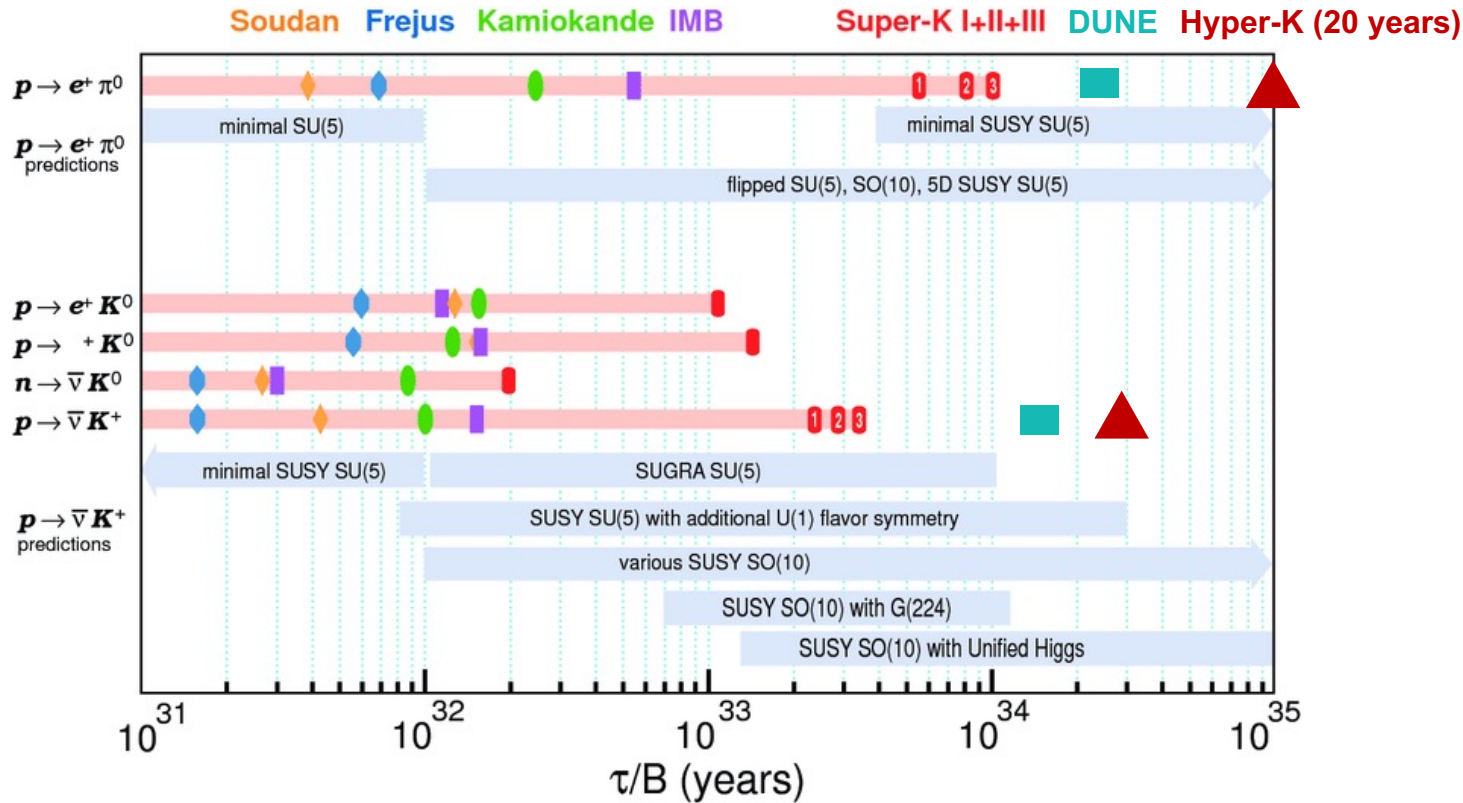
Physics potential:

# Proton decay search

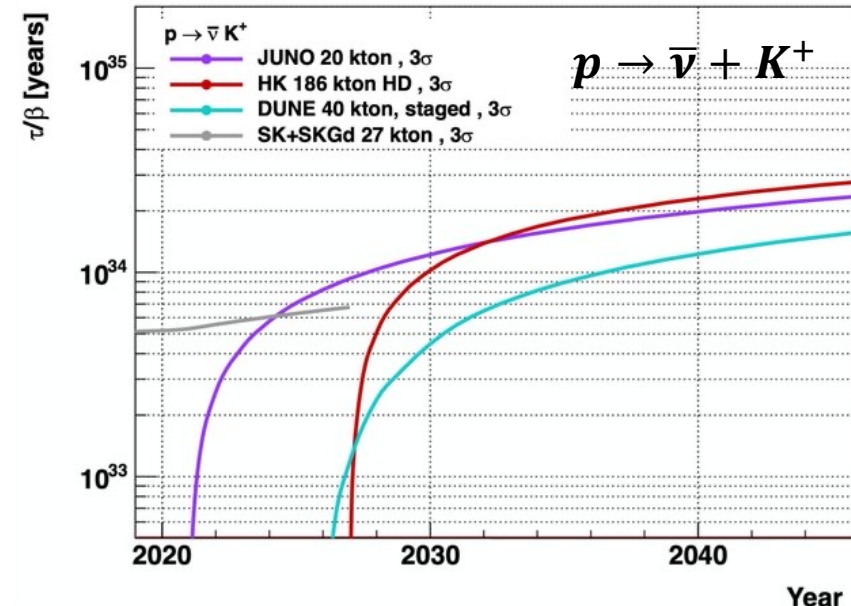
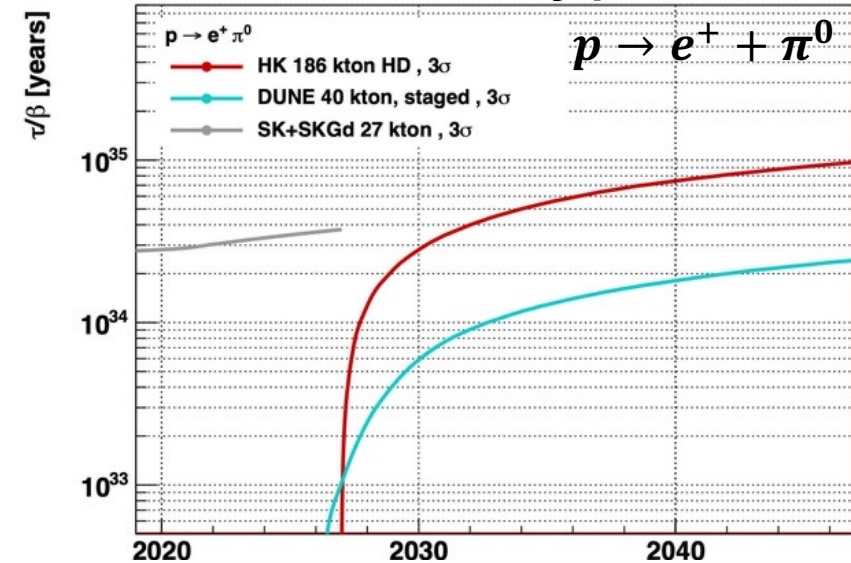
# Proton decay search



- Nucleon decay - direct probe for GUT
- Hyper-K provides world leading sensitivity for proton decay due to large mass and free protons in target material (H<sub>2</sub>O)
- 3σ discovery potential - half-life of 10<sup>35</sup> years for  $p \rightarrow e^+ + \pi^0$  and 3×10<sup>34</sup> years for  $p \rightarrow \bar{\nu} + K^+$  after 20 years



## 3σ discovery potential

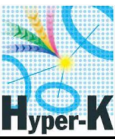


**Physics potential:**

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# **Astrophysics neutrino**

# Supernovae neutrino

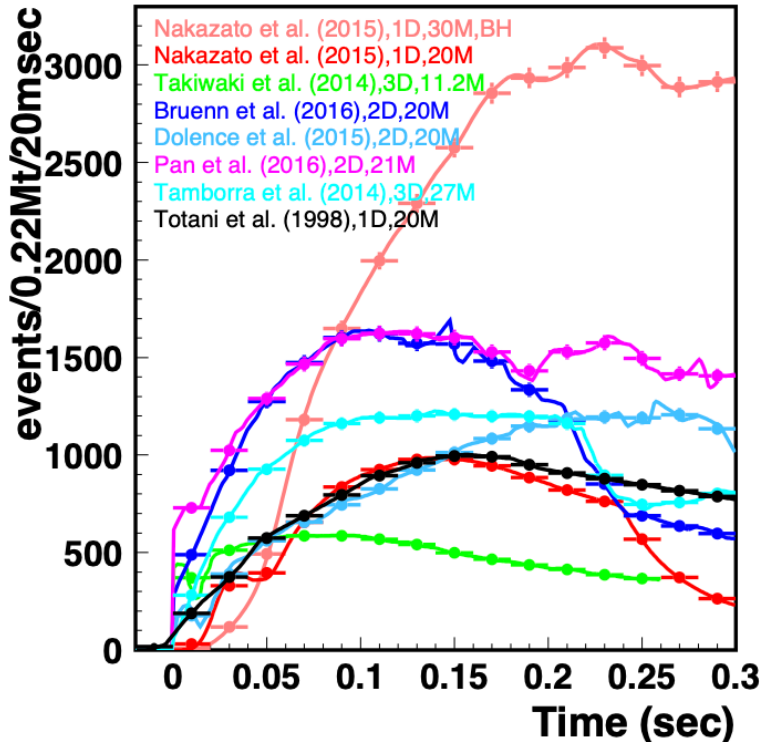


## Direct SN $\nu$

Tens of thousand events (~from centre of Milky Way)  
+ precise timing information



- accurate supernovae models tests
- mass ordering measurement
- absolute mass of neutrinos sensitivity (0.5 to 1.3 eV)



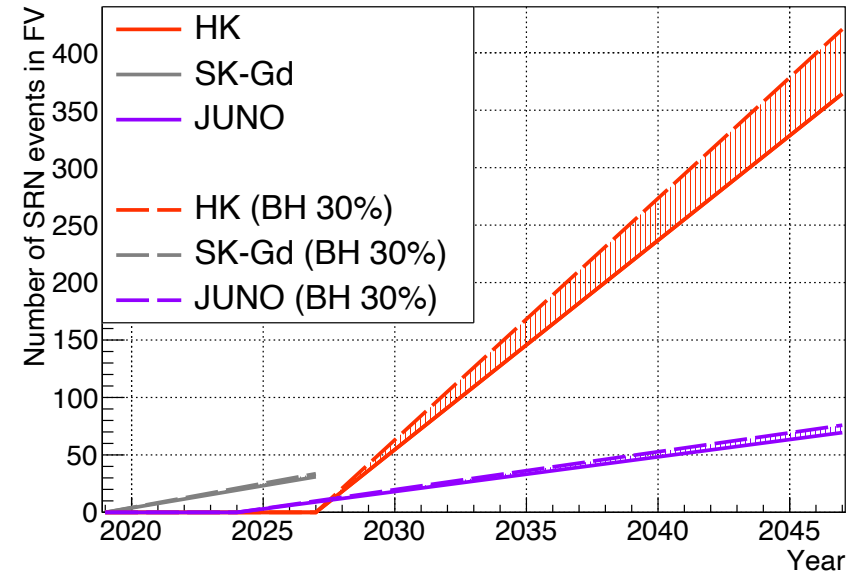
Interaction detection	Integrated number of events	
	NO	IO
$\bar{\nu}_e + p$	57836	74852
$\nu + e^-$	3615	3580

10 kpc supernova using the Livermore model

## DSNB

- 4.2 (5.7)  $\sigma$  sign. after 10 (20) years\*

Total number without event selection efficiency



\* Assumed the flux prediction from model described here [arXiv:astro-ph/0202450](https://arxiv.org/abs/2002.02450) and  $E_\nu \in [16; 30]$  MeV after spallation bkg reduction

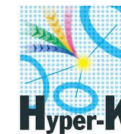
Physics potential:

**Beam and atmosph.  
neutrino**



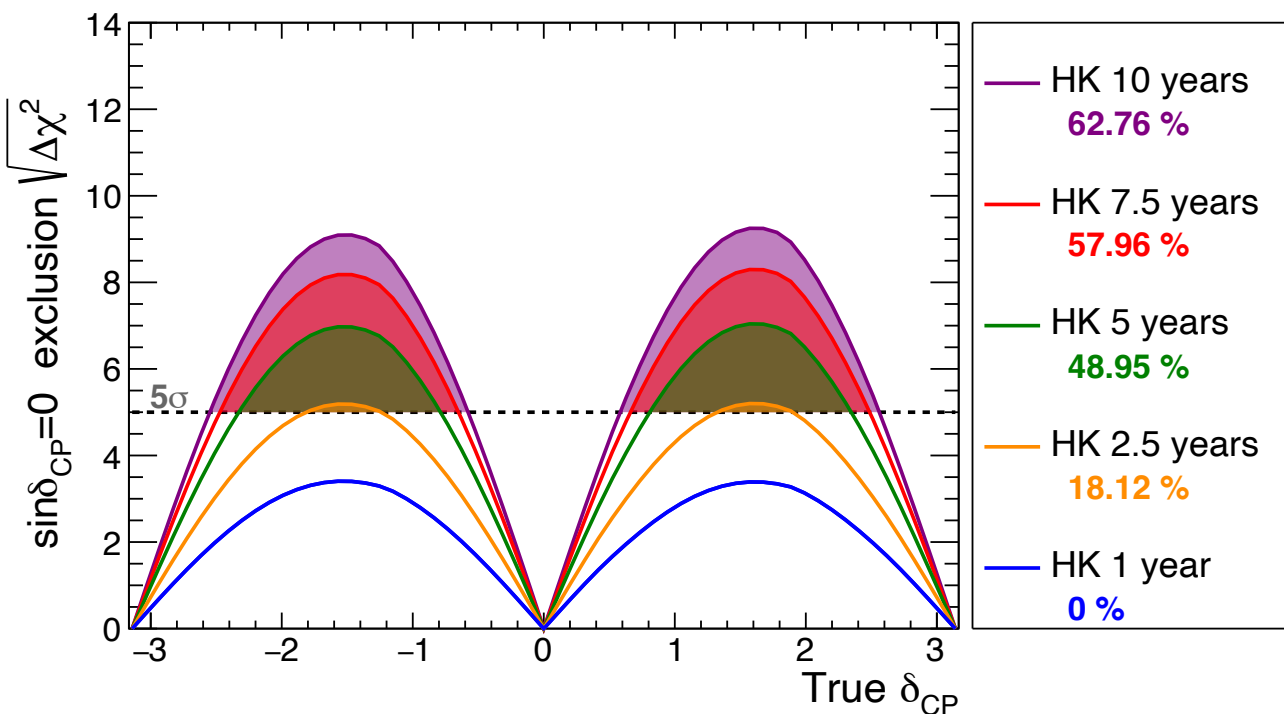
- New extensive LBL sensitivity studies have been updated last year
  - All analysis have been performed by **French** PhD students: Claire Dalmazzone (LPNHE) and myself (LLR, IRFU)
  - The paper on LBL sensitivity studies is under preparation on behalf of the HK collaboration
- Based on T2K analysis (Neutrino2020 version):
    - $\text{Hyper - K MC} = \text{Super - K MC} \times \text{Reweights (flux, detector, POT)}$
    - Event selections, fitting strategy and statistical treatment adopted from T2K
    - Systematics parametrisation adopted from T2K
- Three systematics models considered:
    - **Statistics only** - no systematics
    - **T2K 2020 syst.** - T2K constraints
    - **Improved syst.** – expected improved constraints from IWCD, upgraded ND280 measurements and high statistics
  - MO assumed to be known

# Long-baseline program: CP-violation



## Long-baseline experiment provides unique possibility for precise CP-V measurement

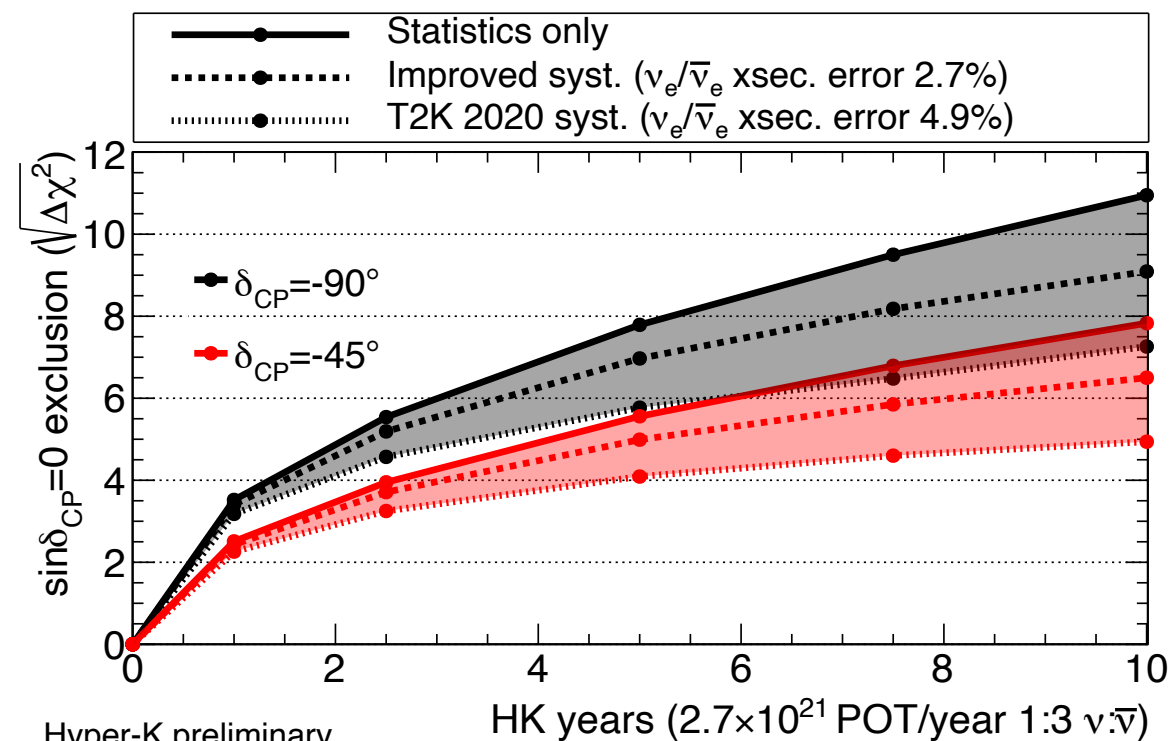
- Allows direct CP-V search having well-controlled  $\nu$  and  $\bar{\nu}$  beams
- **5 $\sigma$**  sensitivity to CP-V discovery for **62% of true  $\delta_{CP}$**  values in 10 years (improved syst.)
- After **2.5 years 5 $\sigma$  CP-V discovery** for maximal CP-V



Hyper-K preliminary

True normal ordering (known), Improved systematics

$$\sin^2\theta_{13}=0.0218\pm 0.0007, \sin^2\theta_{23}=0.528, \Delta m_{32}^2=2.509\times 10^{-3}\text{eV}^2/c^4$$



Hyper-K preliminary

True normal ordering (known)

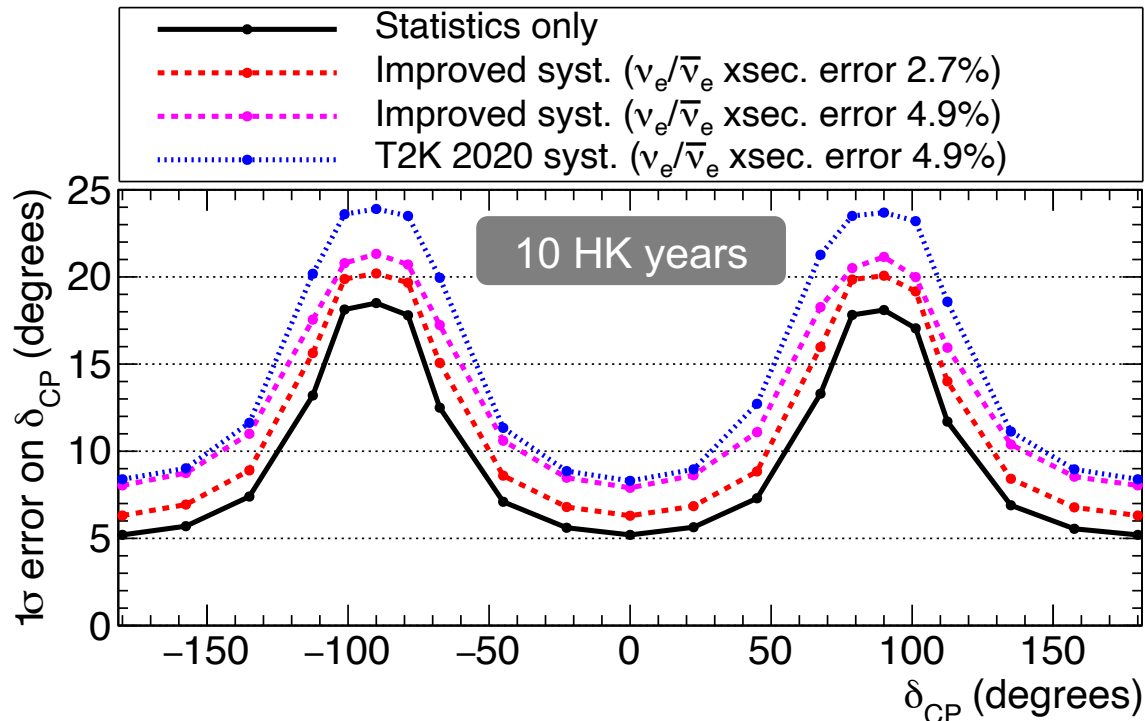
$$\sin^2\theta_{13}=0.0218\pm 0.0007, \sin^2\theta_{23}=0.528, \Delta m_{32}^2=2.509\times 10^{-3}\text{eV}^2/c^4$$

HK years ( $2.7\times 10^{21}$  POT/year 1:3  $\nu:\bar{\nu}$ )

# Long-baseline program: $\delta_{CP}$ precision

## Long-baseline experiment provides unique possibility for precise $\delta_{CP}$ measurement

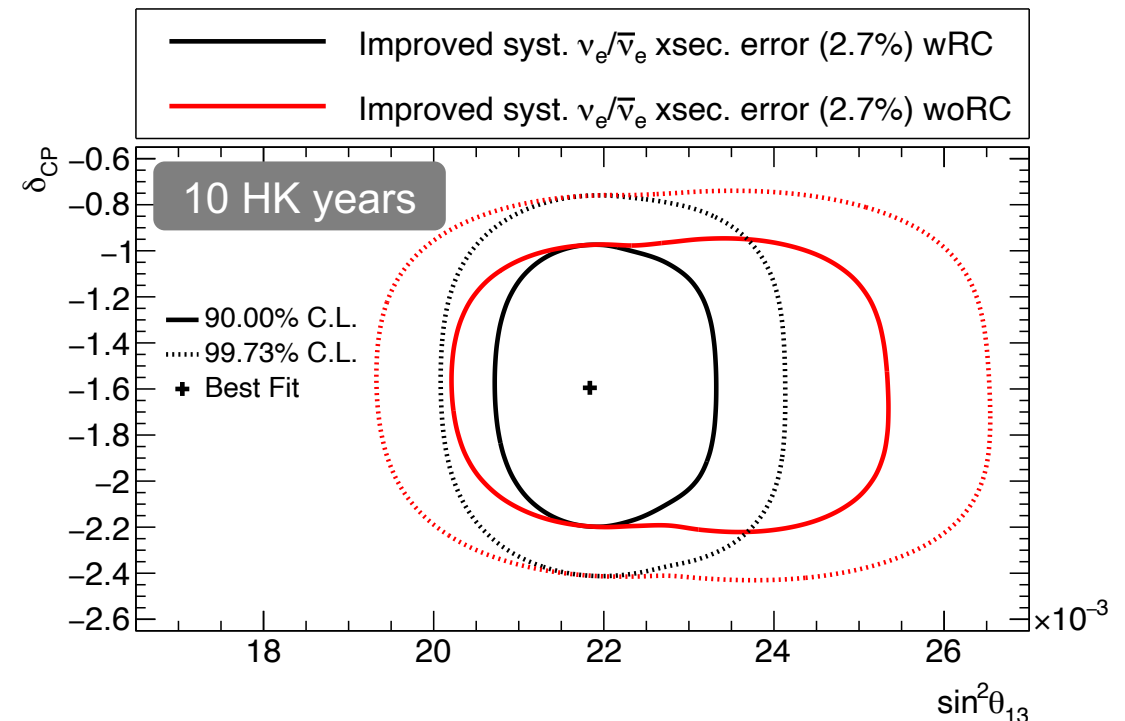
- $5^\circ - 22^\circ$  for  $1\sigma$  precision  $\delta_{CP}$  measurement
- Depends on true value of  $\delta_{CP}$  and highly on  $\sigma(\nu_e)/\sigma(\bar{\nu}_e)$
- Small impact from reactor constraint ( $1:3 \nu:\bar{\nu}$  allows to split the degeneracy with  $\sin^2 \theta_{13}$ )



Hyper-K preliminary

True normal ordering (known), HK 10 Years ( $2.7 \times 10^{22}$  POT  $1:3 \nu:\bar{\nu}$ )

$\sin^2 \theta_{13} = 0.0218 \pm 0.0007$ ,  $\sin^2 \theta_{23} = 0.528$ ,  $\Delta m_{32}^2 = 2.509 \times 10^{-3} \text{eV}^2/c^4$

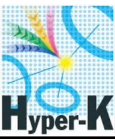


Hyper-K preliminary

True normal ordering (known), 10 years ( $2.7 \times 10^{22}$  POT  $1:3 \nu:\bar{\nu}$ )

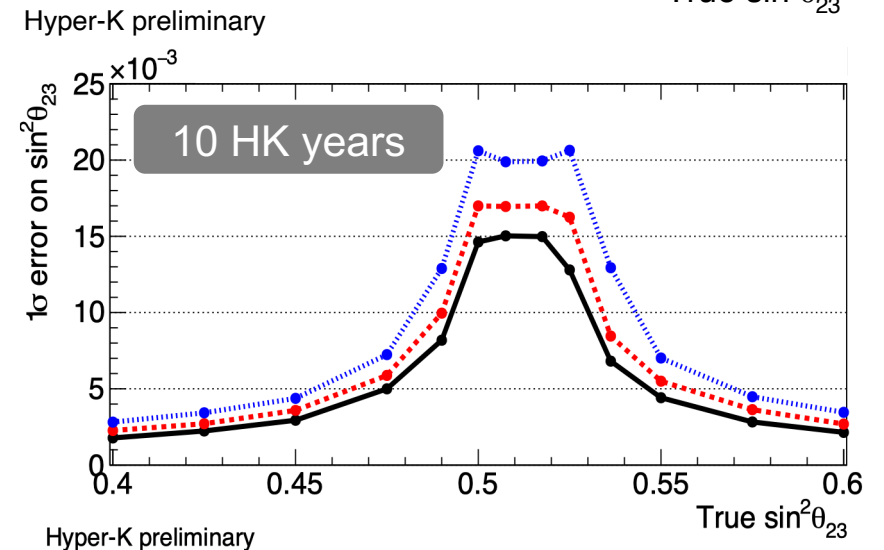
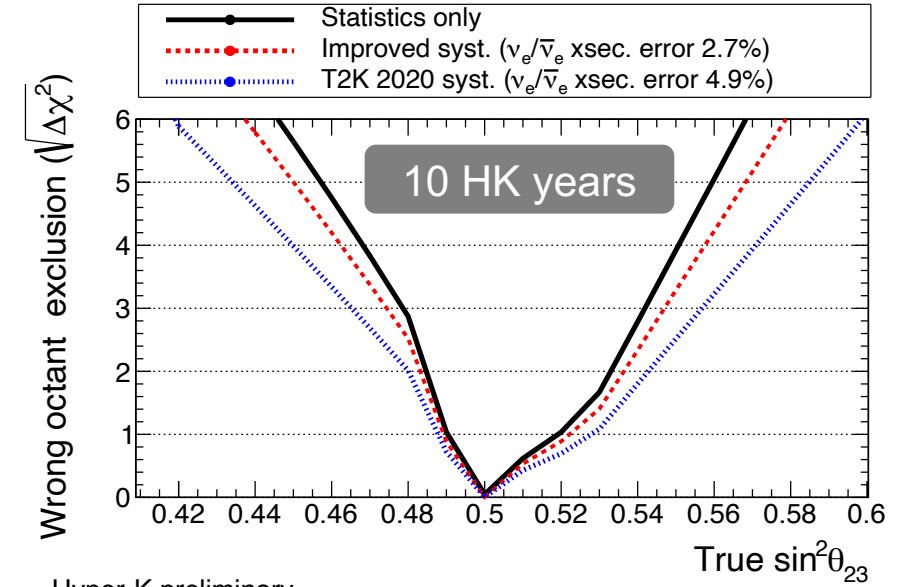
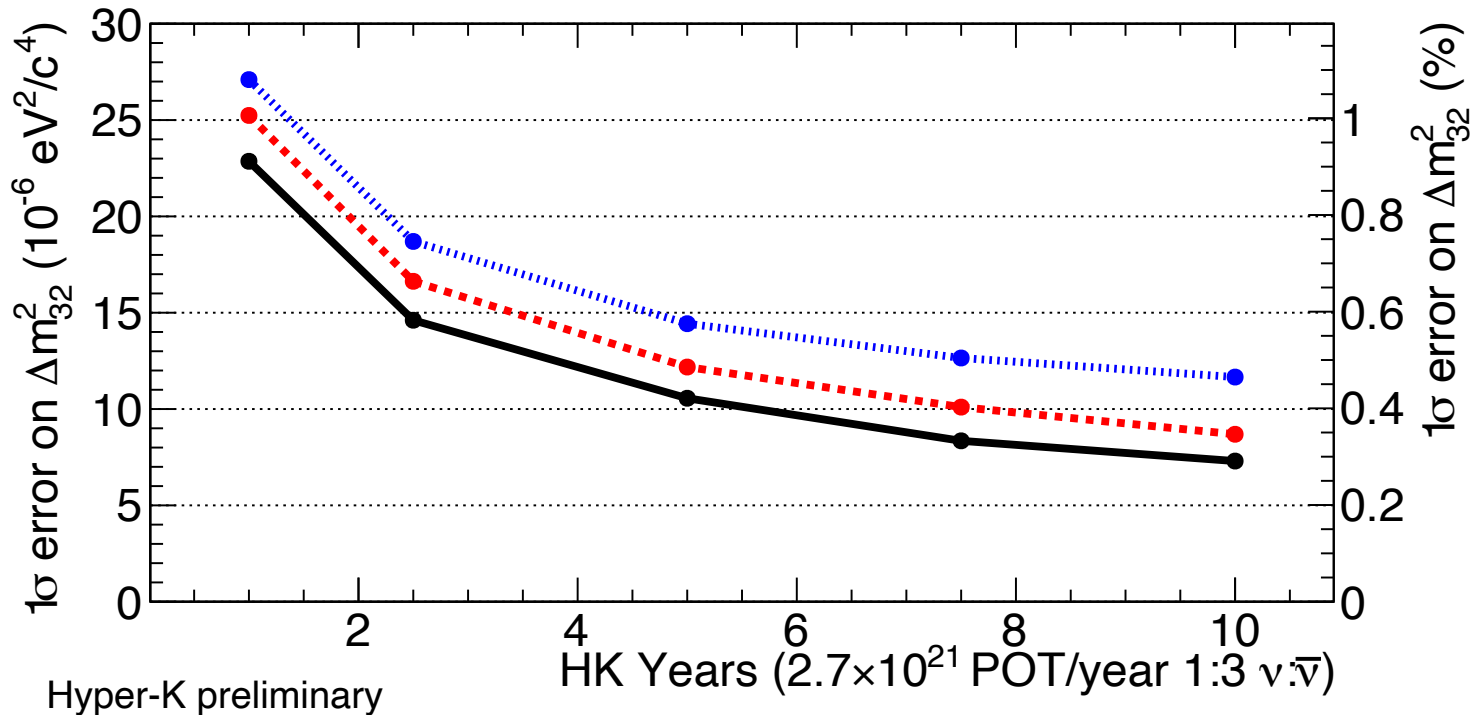
$\sin^2 \theta_{13} = 0.0218 \pm 0.0007$ ,  $\sin^2 \theta_{23} = 0.528$ ,  $\Delta m_{32}^2 = 2.509 \times 10^{-3} \text{eV}^2/c^4$ ,  $\delta_{CP} = -1.601$

# Long-baseline program: atm. parameters



## Long-baseline experiment provides precise measurements on atmospheric parameters

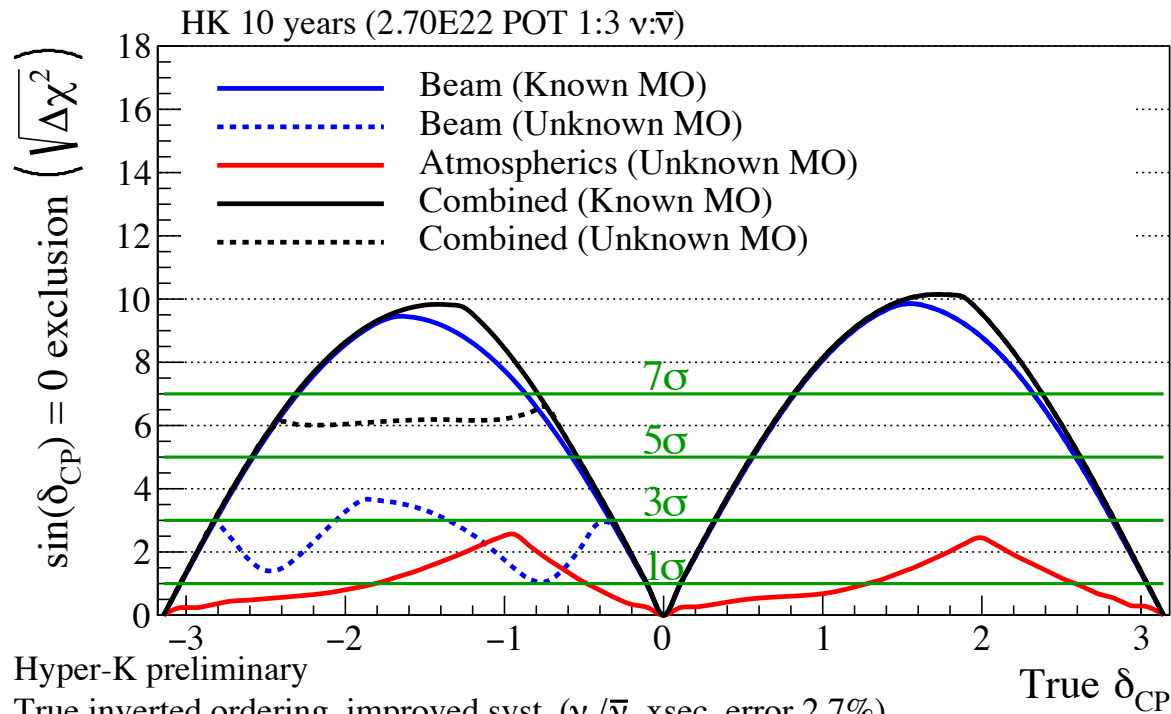
- $5\sigma$  wrong octant rejection for  $\sin^2 \theta_{23} < 0.45$  and  $\sin^2 \theta_{23} > 0.57$
- $\sim 0.5 - 3.5\%$  res. for  $\sin^2 \theta_{23}$  (depending on true  $\sin^2 \theta_{23}$ )
- $\sim 0.5\%$  res. for  $\Delta m_{32}^2$



# Beam + atmospheric

- Previous slides assumed known mass ordering
- MO unknown  $\rightarrow$  degeneracies can degrade sensitivity to  $\delta_{CP}$  and octant (depending on true MO and  $\delta_{CP}$ )

New Beam+Atm analysis is in progress basing on T2K+SK joint fit (arxiv/2405.12488)



	$\sin^2 \theta_{23}$	Atm. only	Atm. + Beam
Mass ordering	0.4	2.2 $\sigma$	3.8 $\sigma$
	0.6	4.9 $\sigma$	6.2 $\sigma$
$\theta_{23}$ octant	0.45	2.2 $\sigma$	6.2 $\sigma$
	0.55	1.6 $\sigma$	3.6 $\sigma$

Hyper-K preliminary  
 True inverted ordering, improved syst. ( $\nu_e/\bar{\nu}_e$  xsec. error 2.7%)  
 $\sin^2(\theta_{13})=0.0218$   $\sin^2(\theta_{23})=0.528$   $|\Delta m_{32}^2|=2.509 \times 10^{-3} \text{ eV}^2/c^4$  Based on older version of T2K analysis

Experimental setup:

# Hyper-Kamiokande detectors

# Hyper-Kamiokande detector

- Hyper-Kamiokande is third generation of Water Cherenkov detector in Japan
- To meet the physics goals mentioned above numerous improvements are incorporated:
  - **Increased Detector Volume:** 188.4 kton FV (x8.4 of SK FV)
  - **Improved photo detector system (20k 50 cm PMT, ~1k mPMT)**
  - **Improved electronics and readout system**
  - **New calibration methods**

**Hyper-K PMT**  
Hamamatsu R12860

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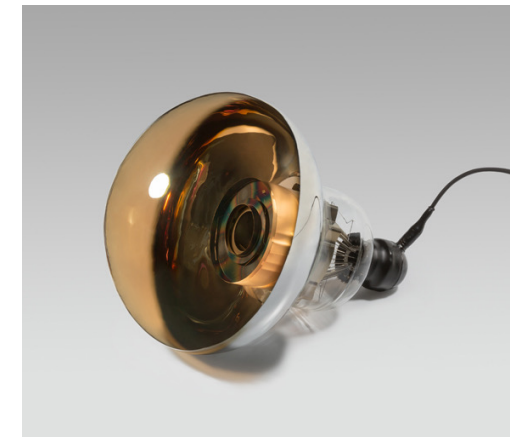
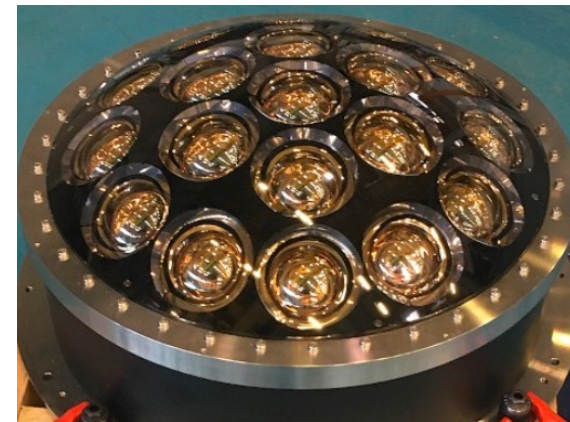
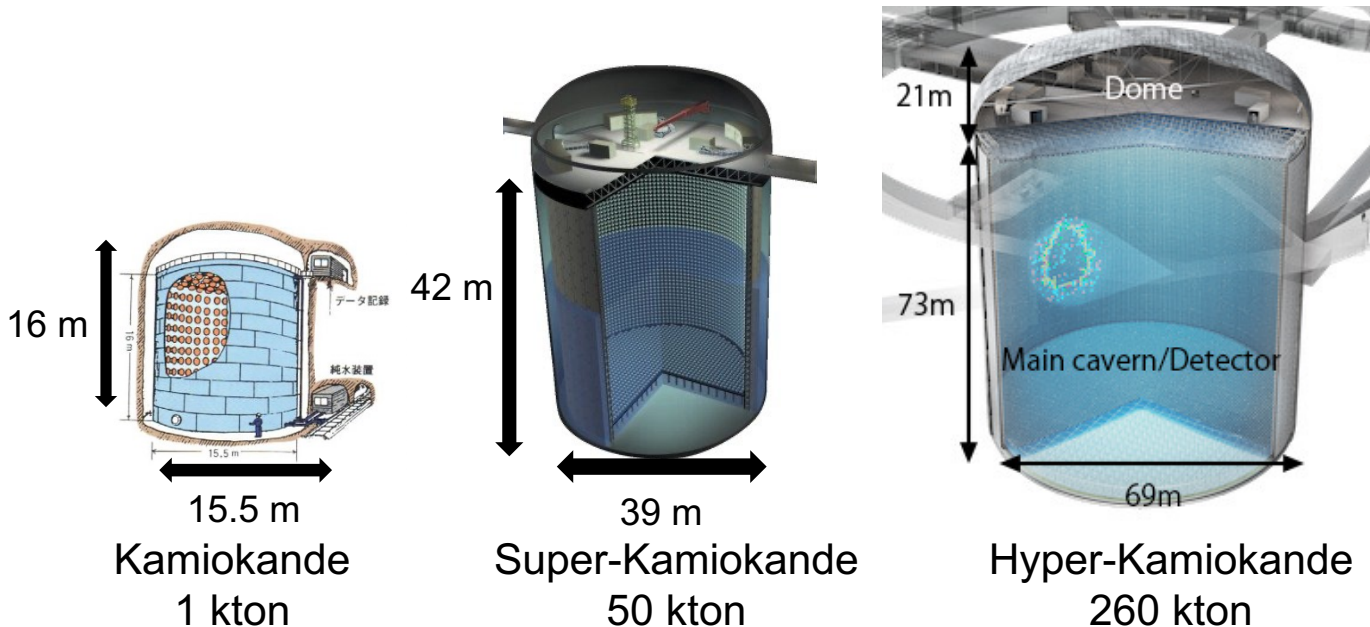
- Box-and-line dynode
- QE = 35 – 40%

- x2 photo-detection eff.
- x2 timing resolution
- x2 charge resolution

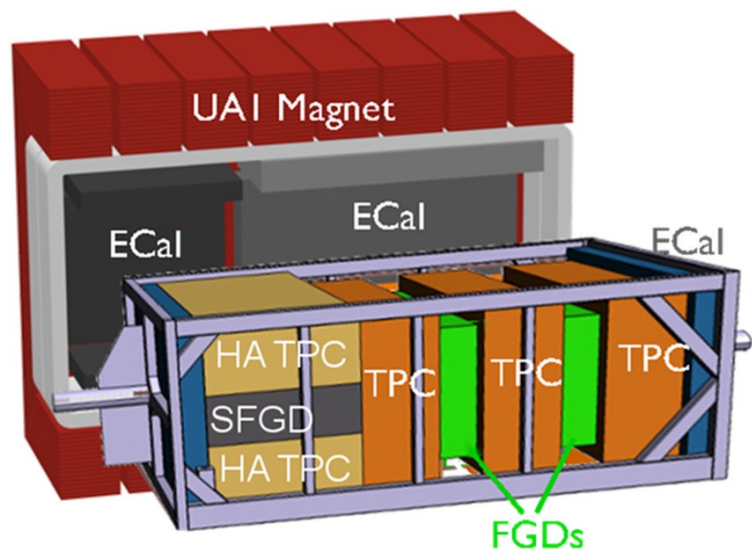
}

wrt to  
SK

Details in backup

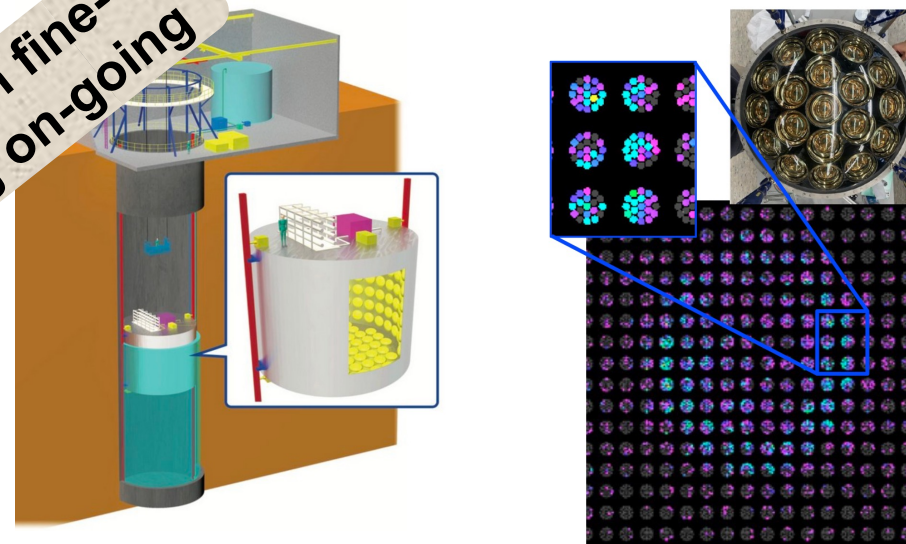


# Near Detector complex



ND280 IWCD

Design fine-tuning on-going



- Off-axis near detector ( $2.5^\circ$ )
- **Constrains flux and x-sec** syst. in the oscillation analysis
- **Upgraded in 2024**, with **successful operation** demonstrated in the T2K experiment.

- Additional “near” detector  $\sim 850\text{m}$  from target
- 600 ton Water Cherenkov detector instrumented with mPMTs
- Movable detector covering off-axis angles  $1.5^\circ - 4^\circ$

- Same target material and  $4\pi$  acceptance as FD
- Probe different neutrino energy spectra

Crucial reduction of systematics uncertainties

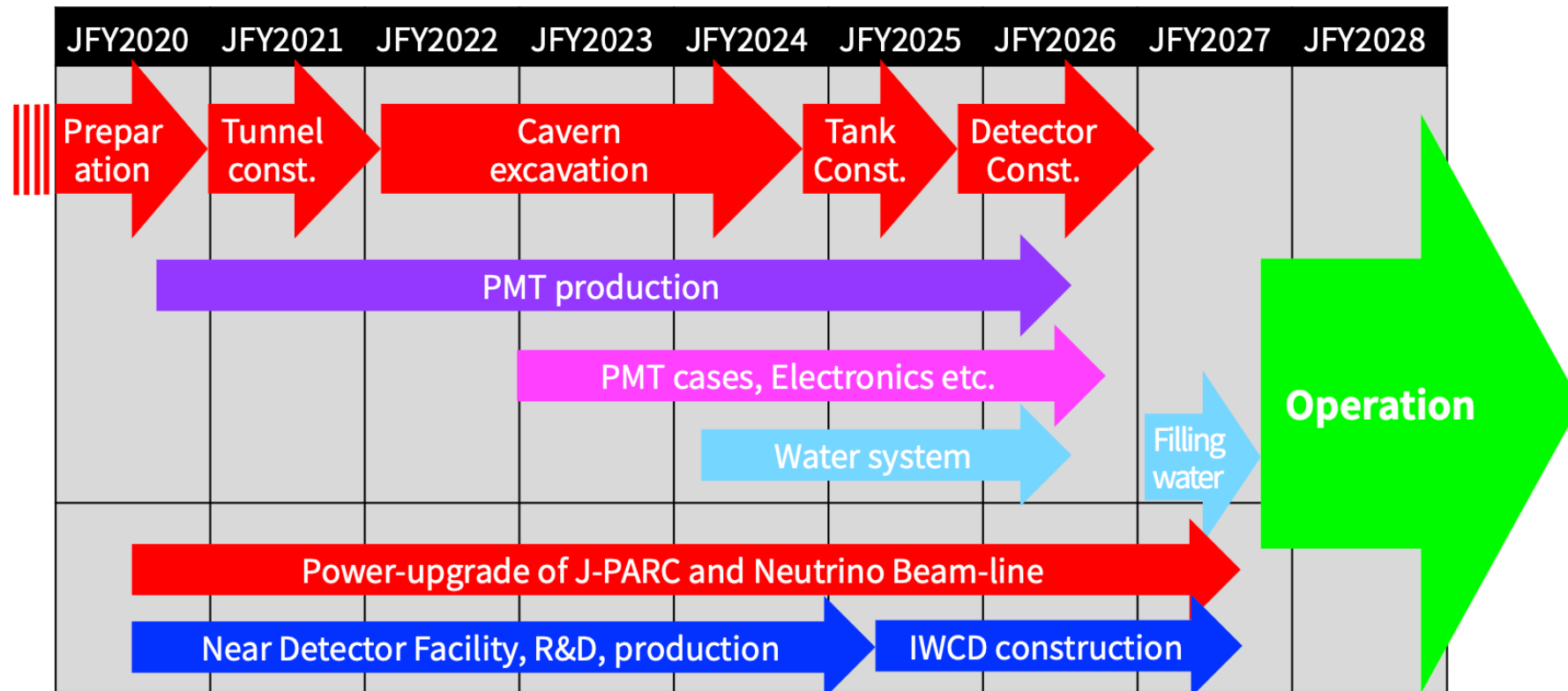


# Current status

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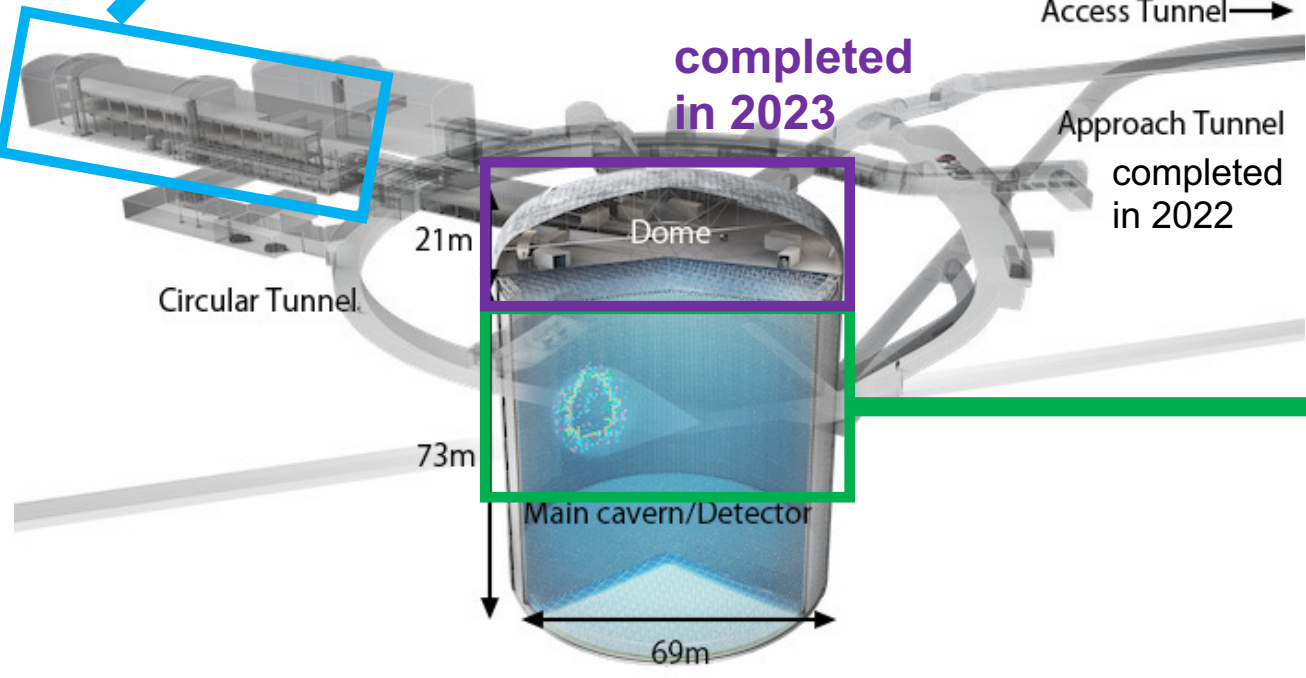
# Current status

- Cavern excavation and PMT delivery&tests are ongoing
- Construction phase extended by 6 months, mainly due to changes to the top structure of the detector
- End of detector construction and start water filling May 2027
- **Start of operations Dec. 2027**



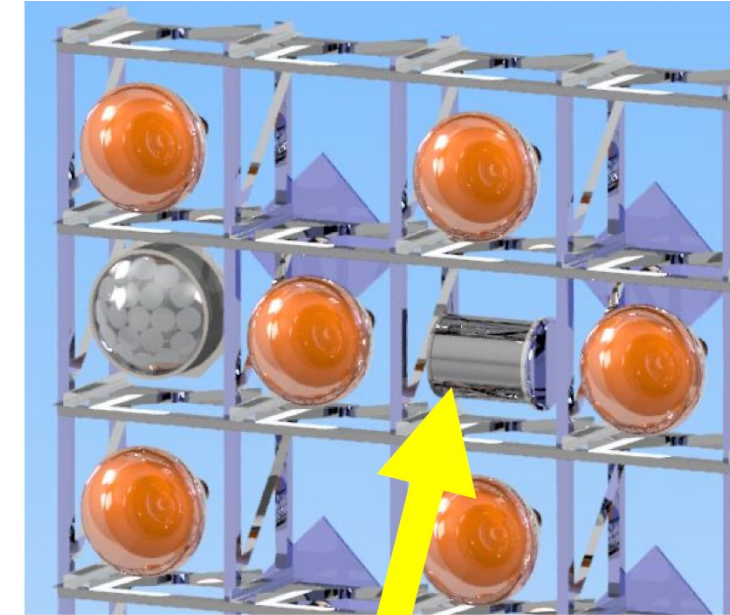
Note: JFY=Japanese Fiscal Year starts on April 1st

# Current status: Far detector

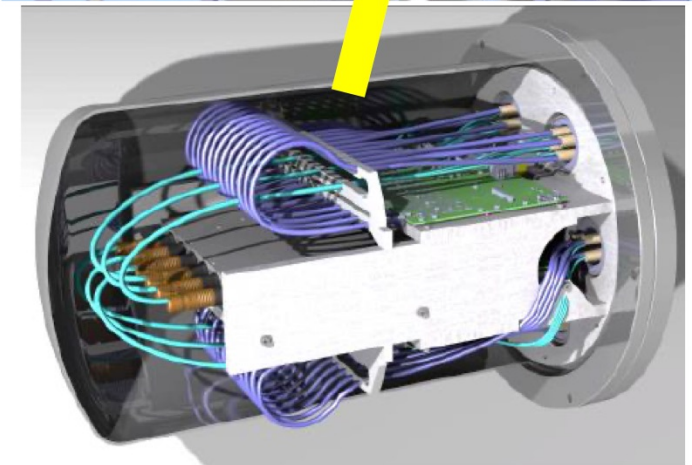
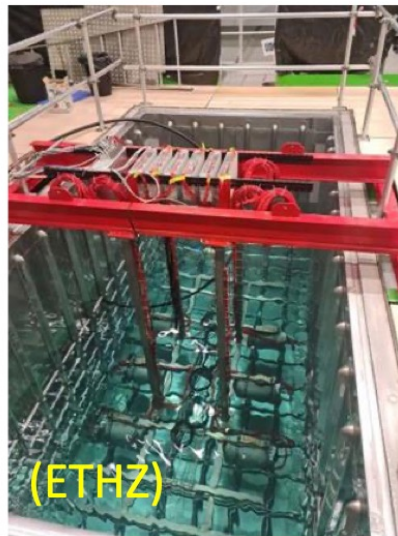
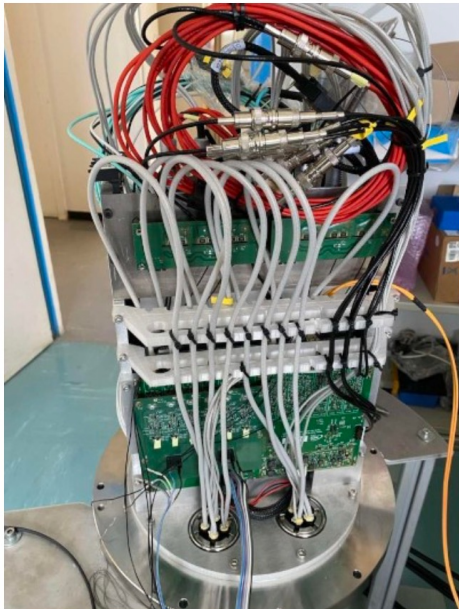


# Current status: Electronics

- Electronics will be underwater, in pressure vessels
- The development of individual components is nearly complete (digitiser, LV, HV, timing/synchronisation, calibrator etc)
- Fully assembled module tests in the water are on-going at CERN and Kamioka
- Calibration and assembly of mass produced components from Summer 2025 at CERN



10-unit test @ CERN

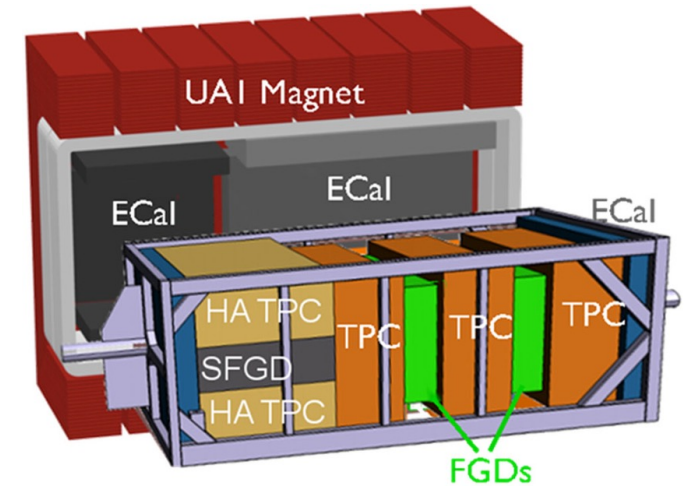


# French contribution

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## Hardware

- Contribution and expertise to the T2K Near Detector complex which will become a part of Hyper-Kamiokande experiment
  - ND280 HA-TPC and SFGD maintenance
  - Discussions started for further upgrade ND280++
  
- Direct contribution to the HK far detector electronics



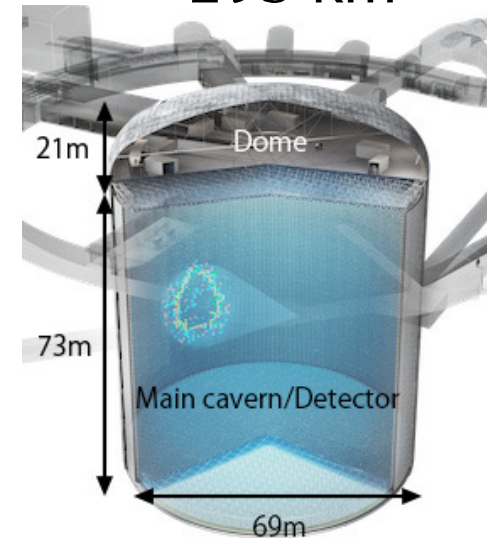
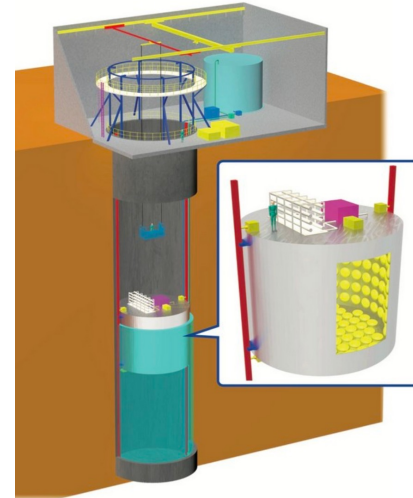
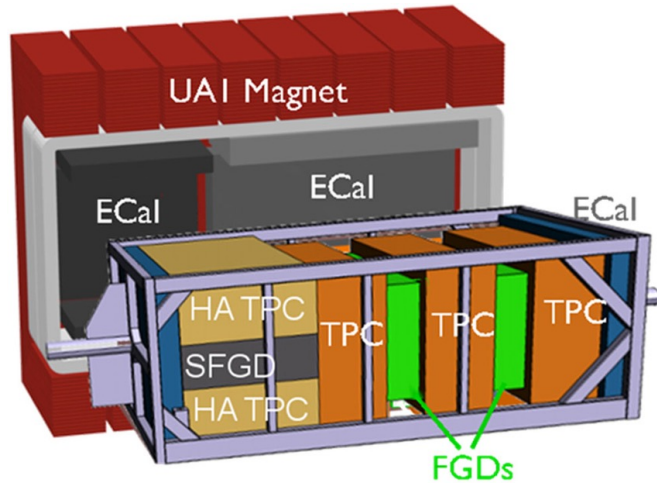
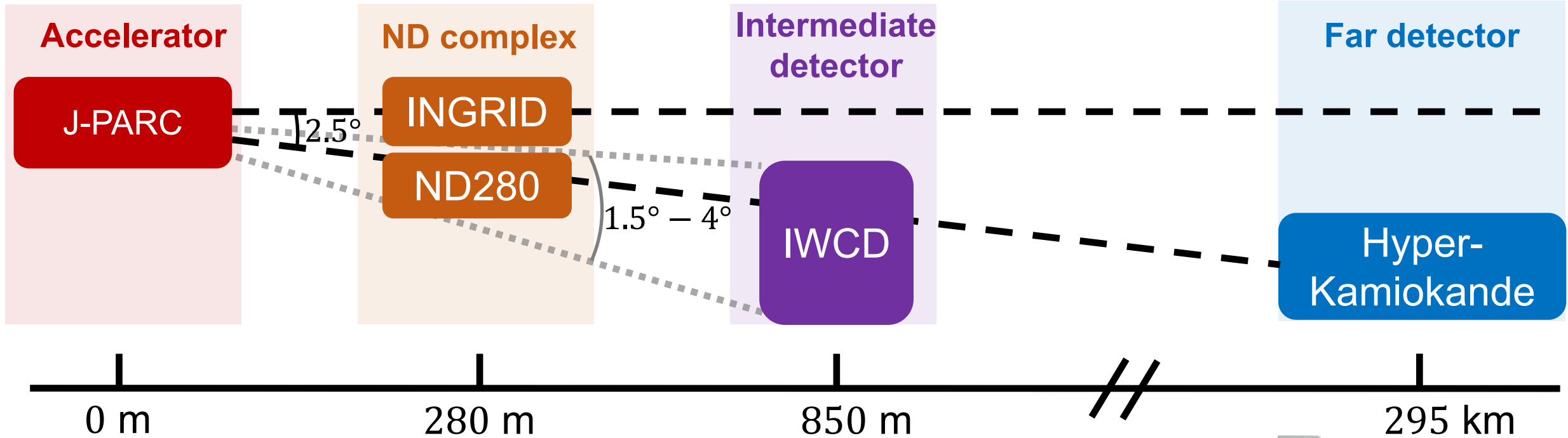
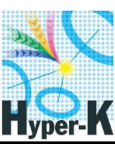
Contribution	Purpose	Current status
The time generation and distribution system	<ul style="list-style-type: none"> <li>• Time synchronization of all detector components</li> <li>• Global timing reference</li> </ul>	<ul style="list-style-type: none"> <li>• Final prototype is almost completed</li> <li>• The time distribution module mass production should start in early 2025 and finish by end 2025</li> </ul>
Electronics testbench	<ul style="list-style-type: none"> <li>• Perform precise calibration of ~1800 digitisers at CERN</li> <li>• Qualify underwater unit before shipment to Japan</li> </ul>	<ul style="list-style-type: none"> <li>• Tests at French testbench almost finalised</li> <li>• Selection of the waveform generator (critical step) has been completed</li> </ul>
Vertical slice tests	<ul style="list-style-type: none"> <li>• Validation of all subsystems compatibility in electronic chain</li> </ul>	<ul style="list-style-type: none"> <li>• Tests are ongoing at CERN and Kamioka</li> </ul>

- Large statistics will allow high precision studies of the oscillation of atmospheric, accelerator and solar neutrinos, as well as searches for new physics (proton decay in particular)
- For  $\delta_{CP}$  favoured from T2K data, **CP-V discovery in less than 3 years**
- Can **exclude CP conservation** at  $5\sigma$  in neutrino oscillations for **62% of true  $\delta_{CP}$  values** in 10 years of operation
- Detector construction and PMTs delivery&tests on-going, excavation of the far detector cavern will be completed by the end of 2024
- French institutes are actively contributing to various aspects of Hyper-K: hardware for FD, maintenance of ND, physics analysis, software development and computing
- Start of operation planned for **December 2027**

**BACKUP**

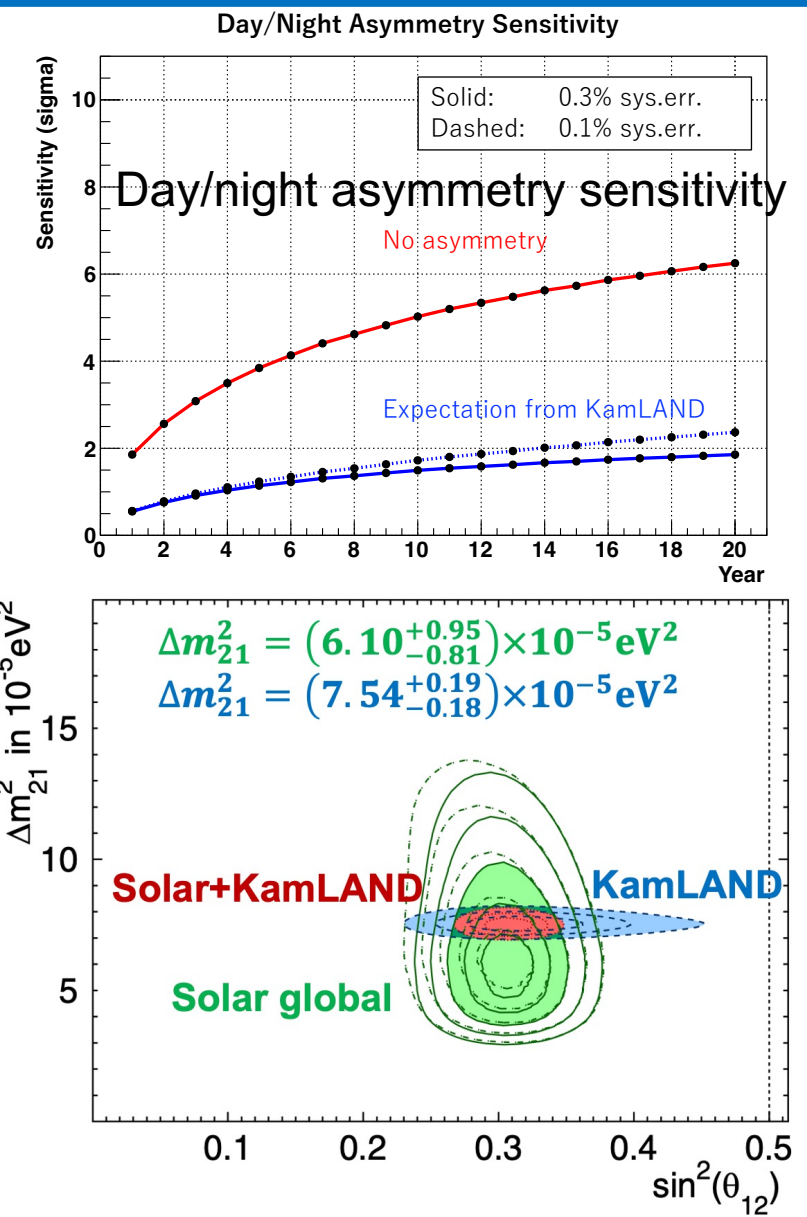


# Long-baseline program





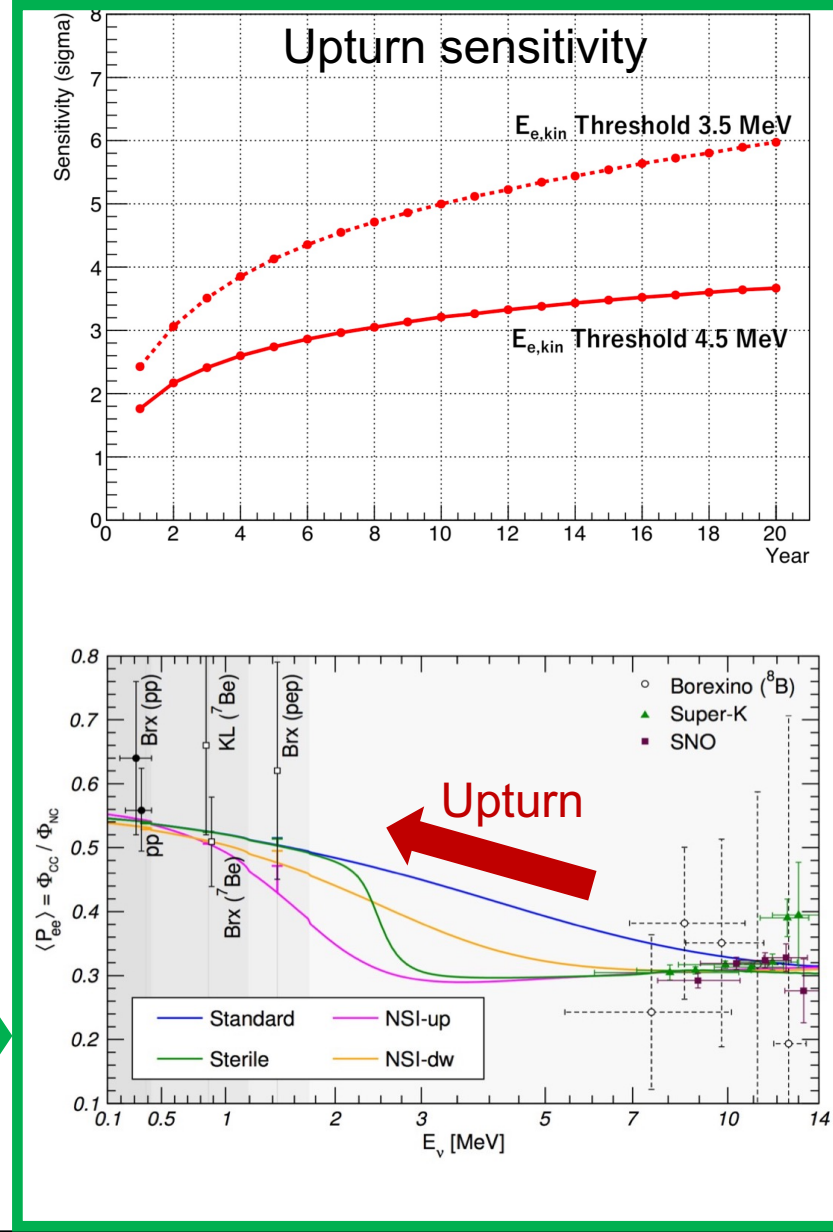
# Solar neutrino



Hyper-K can shed light on remaining questions on solar neutrino oscillations

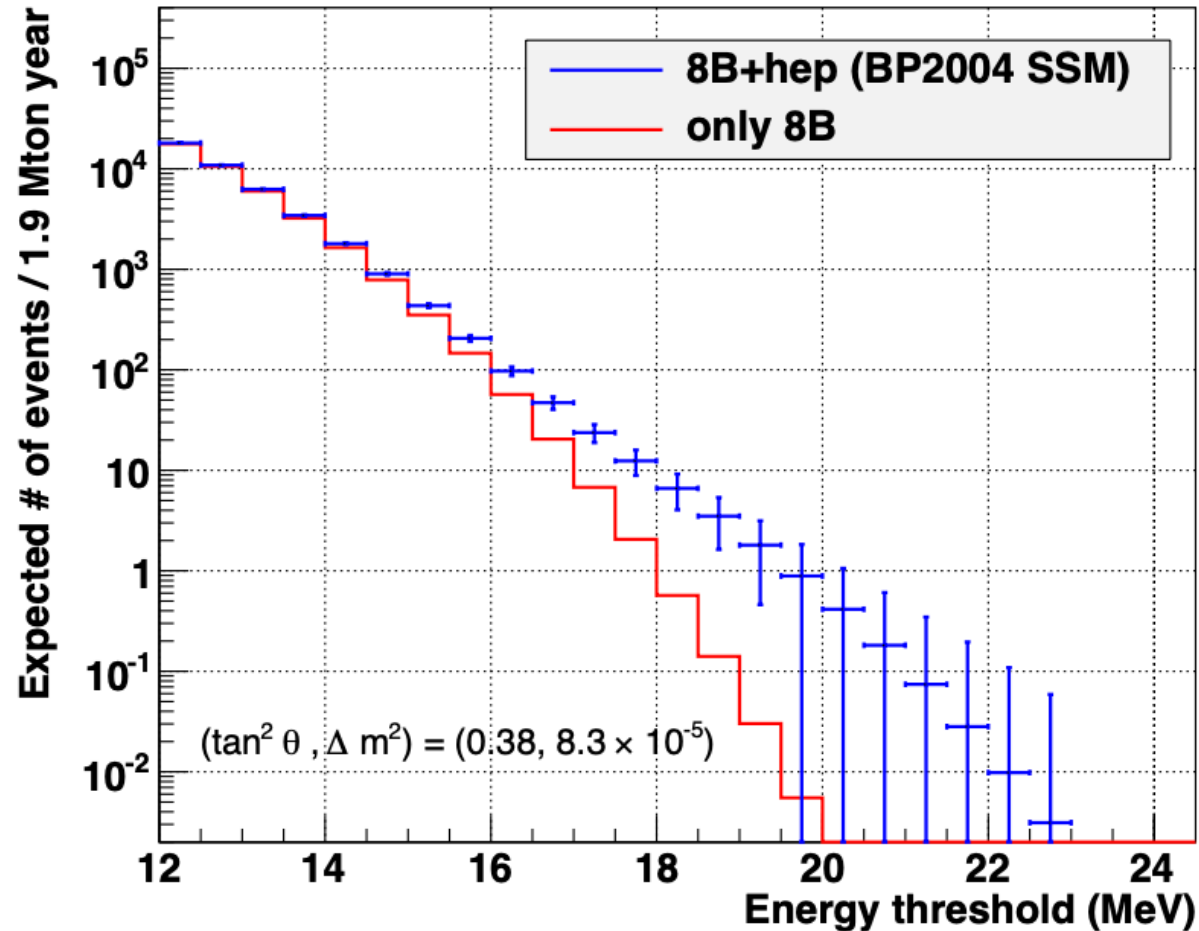
- 1.5 $\sigma$  tension between KamLAND ( $\bar{\nu}_e$ ) and solar global fit ( $\nu_e$ ) (CPT invariance? BSM physics?)
- 5 $\sigma$  significance of a non-zero day/night asymmetry after 10 years
- 2 $\sigma$  day/night sensitivity expected for the difference in  $\nu_e / \bar{\nu}_e$  osc. in 20 yrs.

- > 3 $\sigma$  sensitivity to upturn [3-7] MeV region after 10 HK yrs. ( $E_{th} = 4.5$ )
- Sensitive to sterile neutrino, NSI



# Solar neutrino

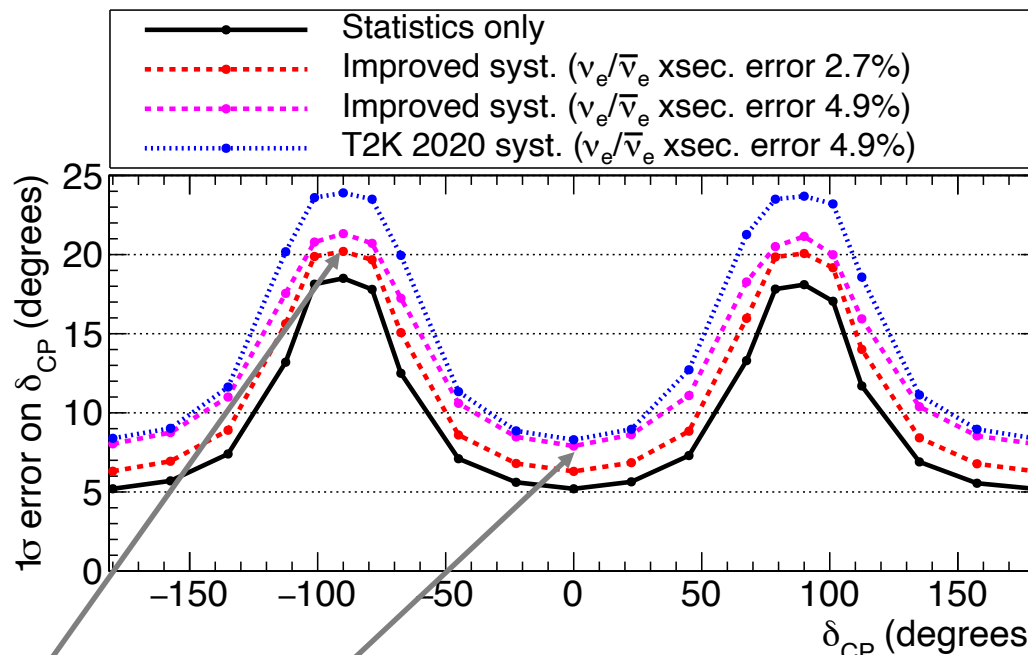
- One of other goals of solar neutrino program is to detect hep neutrinos



## Why **water** is used for proton decay search ?

- Easy to construct larger detector.
  - Much cheaper than iron or gas.
  - You can find large water tank everywhere (common technology).
- High efficiency and low uncertainty.
- $H_2O$  has two hydrogens which are not affected by nuclear effect . They are regarded as “free” proton.
- Free protons contribute high selection efficiency and low uncertainty.

# Long-baseline program: CP-violation



Hyper-K preliminary  
 True normal ordering (known), HK 10 Years ( $2.7 \times 10^{22}$  POT 1:3  $\nu \bar{\nu}$ )  
 $\sin^2 \theta_{13} = 0.0218 \pm 0.0007$ ,  $\sin^2 \theta_{23} = 0.528$ ,  $\Delta m_{32}^2 = 2.509 \times 10^{-3} \text{eV}^2/c^4$

CP-asymmetric term

CP-symmetric term  
 Impacts spectrum shape

$$P = A \sin \delta + B \cos \delta + C$$

$$\frac{\partial P}{\partial \delta} = A \cos \delta - B \sin \delta$$

$$\frac{\partial P}{\partial \delta}$$

$$\frac{\partial P}{\partial \delta}$$

$(\delta \sim 90^\circ) \sim B \rightarrow$  CP-even term domination  $\rightarrow$

$(\delta \sim 0) \sim A \rightarrow$  CP-odd term domination  $\rightarrow$

resolution dominated by  
 the measurement of the  
 energy spectra

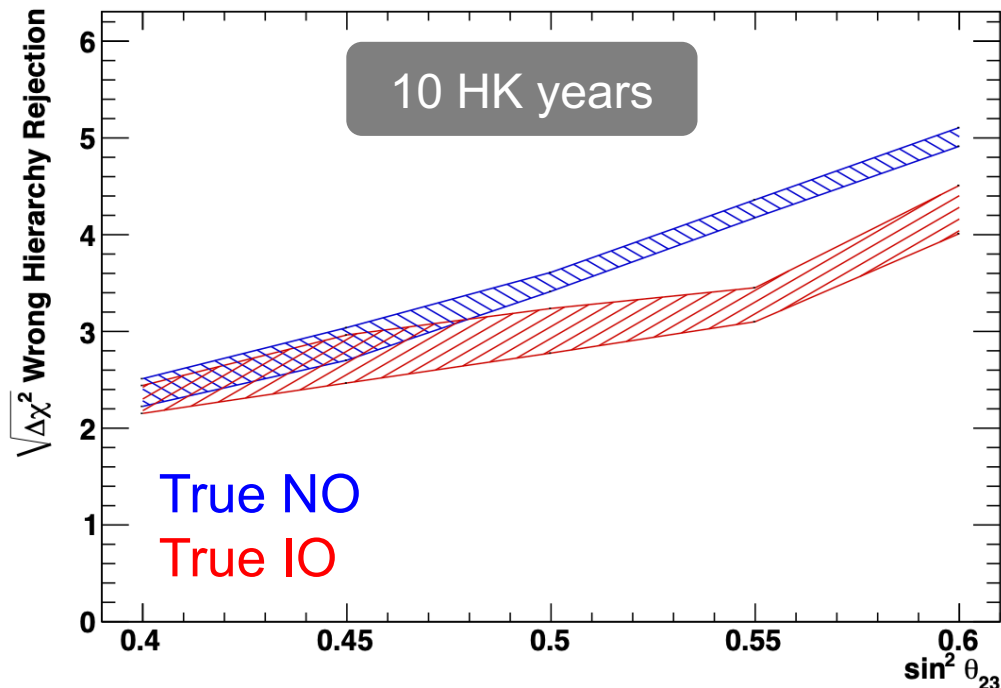
resolution dominated by  
 $\sigma(\nu_e)/\sigma(\bar{\nu}_e)$  error

- $> 3\sigma$  mass hierarchy sensitivity for  $\sin^2 \theta_{23} > 0.53$
- $> 3\sigma$   $\sin^2 \theta_{23}$  octant sensitivity for  $\sin^2 \theta_{23} < 0.43$  and  $\sin^2 \theta_{23} > 0.6$

Sensitivity studies based on SK analysis:

- Scaled SK MC
- No improvement of Super-K systematics assumed

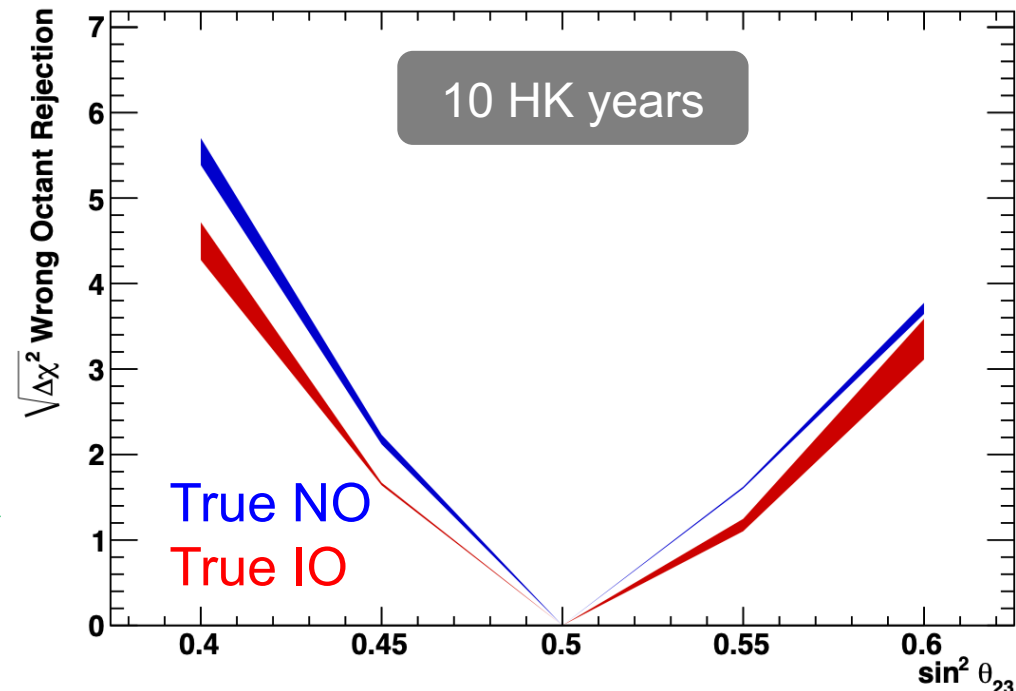
## Mass ordering sensitivity



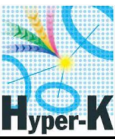
Asymmetry between  $\nu$  and  $\bar{\nu}$

Appearance and disappearance interplay

## $\sin^2 \theta_{23}$ octant sensitivity



# Hyper-Kamiokande detector: photo-detection system



- HK will use **20k 50cm** PMTs to instrument **ID** (20% coverage)
- Improved model compared to SK

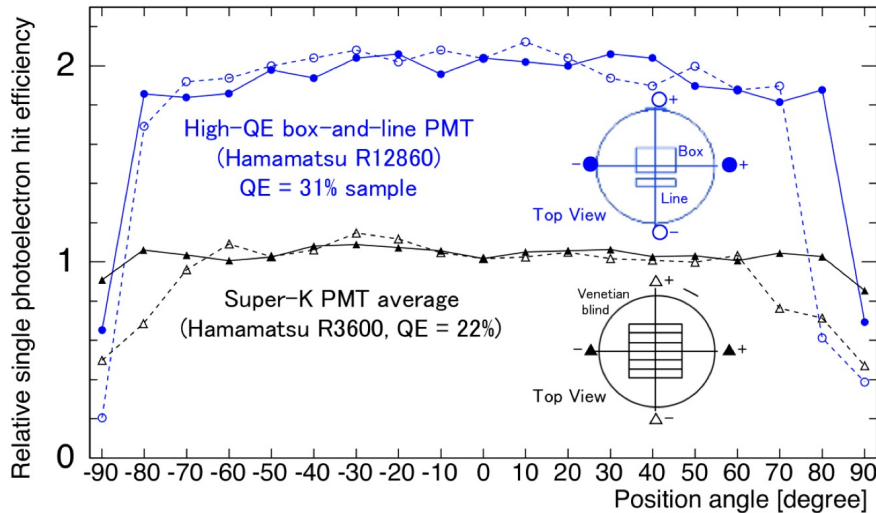
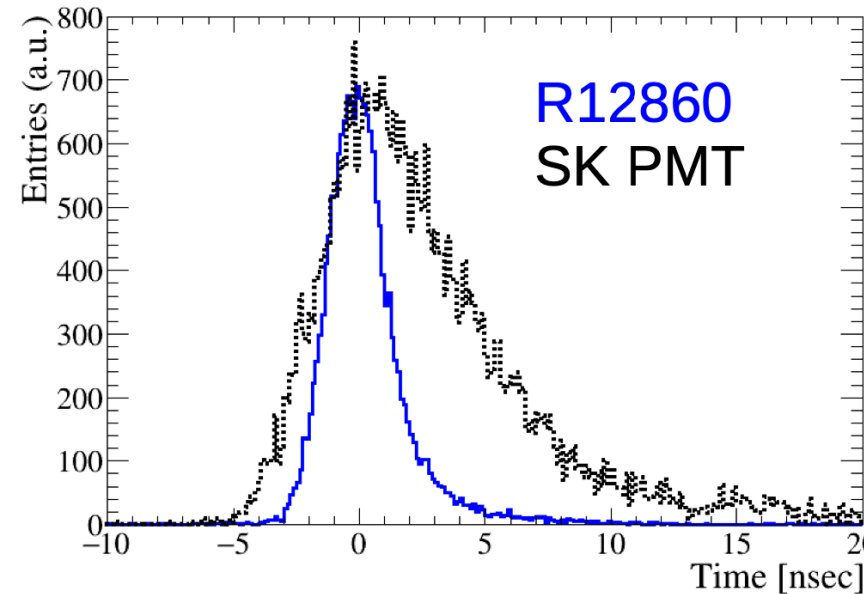
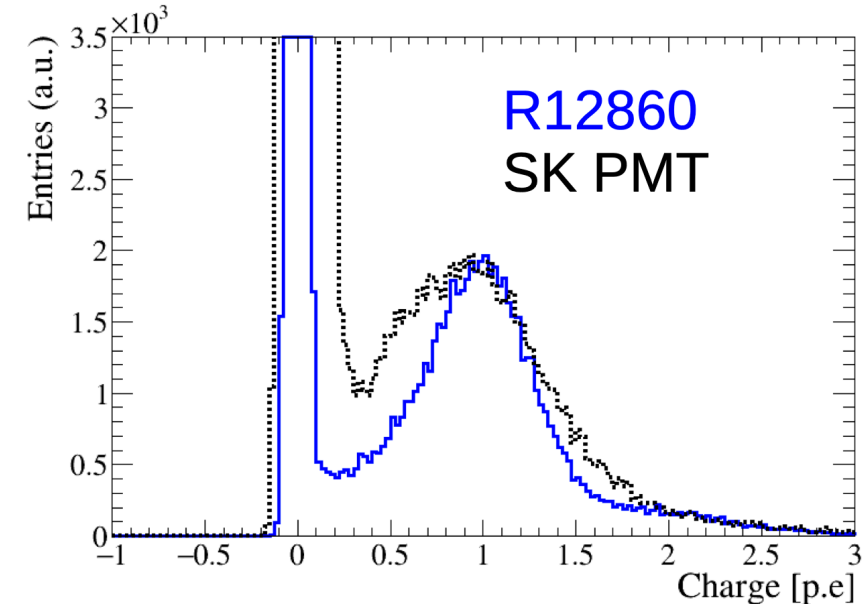
**Super-K PMT**  
Hamamatsu R3600

- Venetian blind
- QE = 20 – 25%



**Hyper-K PMT**  
Hamamatsu R12860

- Box-and-line
  - higher collection efficiency
  - more uniform gain
  - more uniform electron drift path → (better timing and charge res.)
- QE = 35 – 40%



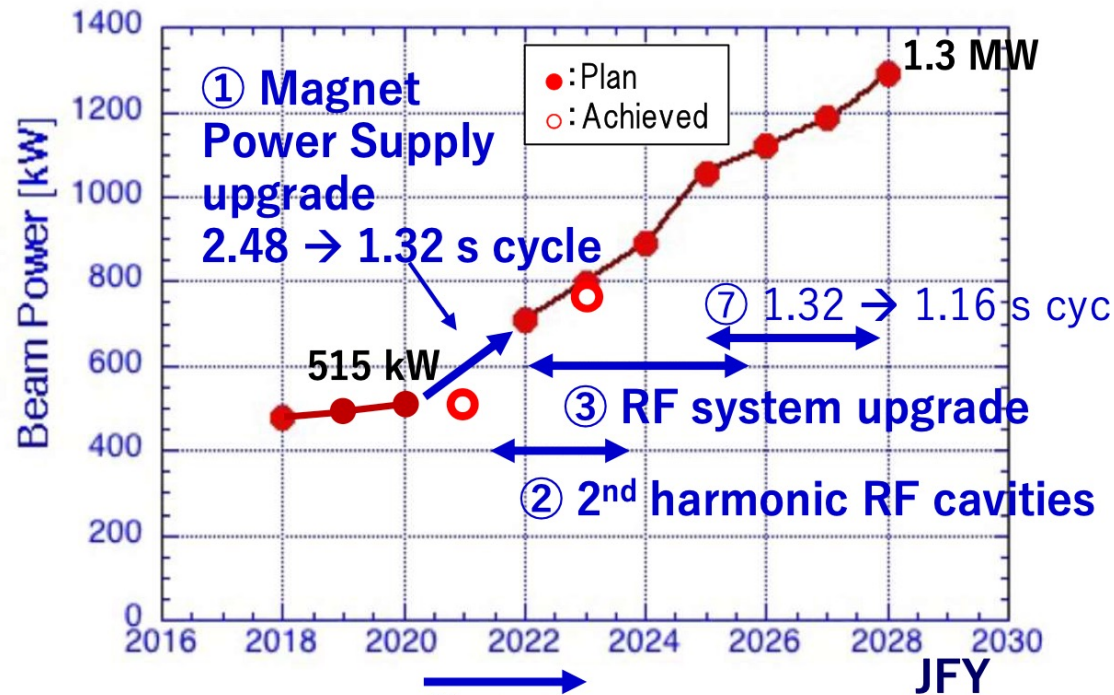
- ~2x photo-detection efficiency
- TTS: 6.73 ns → 2.59 ns (FWHM)
- Charge resolution: 60.1% → 30.8%



# Power Projection in MR Upgrade Plan

$$\text{Beam Power} = \text{Energy (30GeV)} \times 1/T_{\text{rep}} \text{ (pulse/s)} \times \# \text{ of protons (/pulse)}$$

## Original power projection in MR Upgrade Plan



- ④ Collimator system
- ⑤ Injection/FX system
- ⑥ Beam Monitors (BPM circuits)

S. Igarashi, *et. al.*,  
PTEP vol 2021,  
Issue.3,p33

JFY2021	<b>515 kW</b>	2.48 s	$2.66 \times 10^{14}$ ppp
JFY 202*	<b>&gt; 940 kW</b>	<b>&lt; 1.36 s</b>	$2.66 \times 10^{14}$ ppp

- Consecutive demonstration of 760 kW in Dec. 2023
- 800 kW reached in Summer 2024
- 1.3 MW by JFY2028
  - “1.36 → 1.16 s cycle” & “More protons/pulse”

```

/home/daqkun/workspace/develop/jnu_beam
MR Run#          91
MR Shot#         2448782
                 (2024/06/14 09:33:58)
NU Run#          910576
Event#           61240
Spill#           8358153
Deliv. p#        3.88838e+20
                 (this J-PARC run)
Deliv. p#        4.21035e+21
                 (2010/Jan/1~)
                
```

June 2024

Last shot MR Power is **800.9** [kW]  
(2024/06/14 09:33:58)

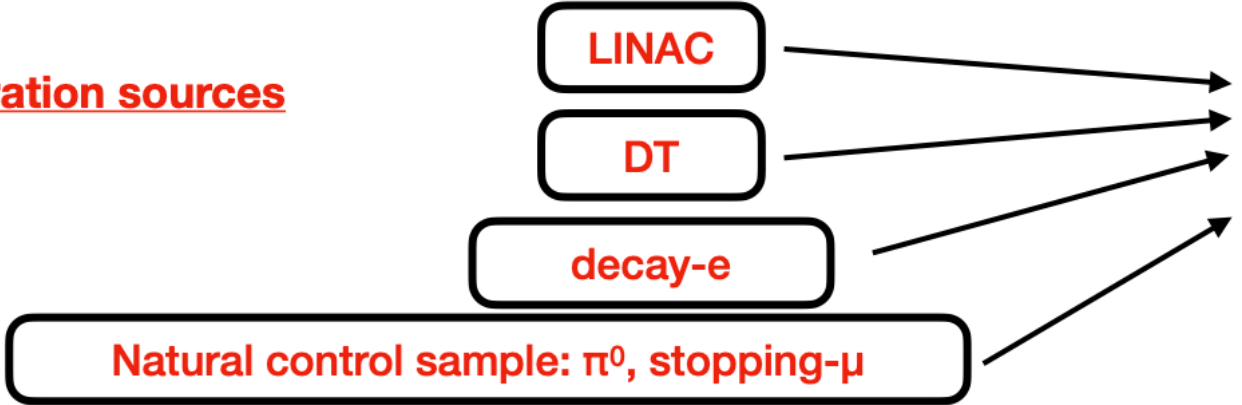
MR DCCT\_073\_1 measurement : 2.2657e+14 [protons per spill]  
NU CT01 measurement : 2.2628e+14 [protons per spill]

Parameter values :  
 LI current: 60.02 [mA]  
 MR micro pulse: 400 [usec]  
 MR chop width: 455 [nsec]  
 MR thinning: 110 /128  
 MR # of bunch: 8

Prediction from parameter values :  
**Expected PPP : 2.1075e+14**  
**Expected PPB : 2.6343e+13**  
**!!! Expected Power : 783 [kW] !!!!**

# Calibration methods overview

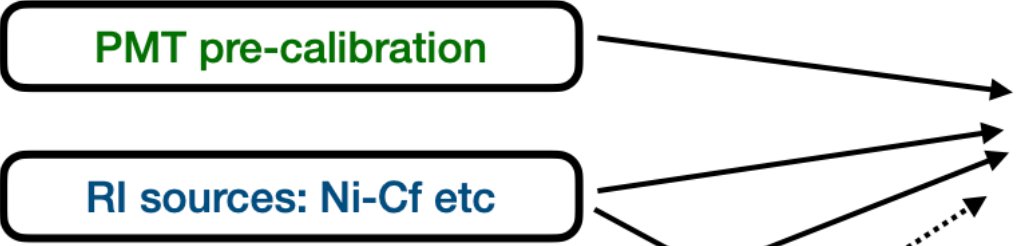
High-level calibration sources



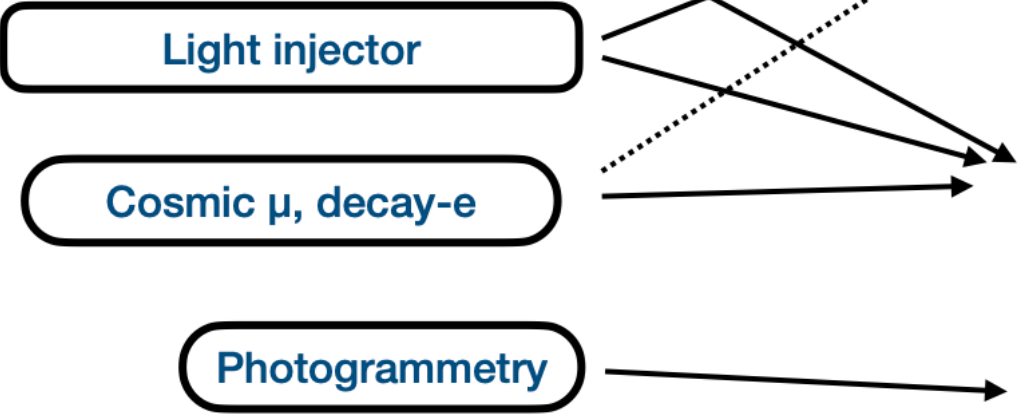
**Reconstruction performance**

- Vertex
- Direction
- Energy
- PID
- Ambient background ID

Ex-situ basic calibration



In-situ basic calibration



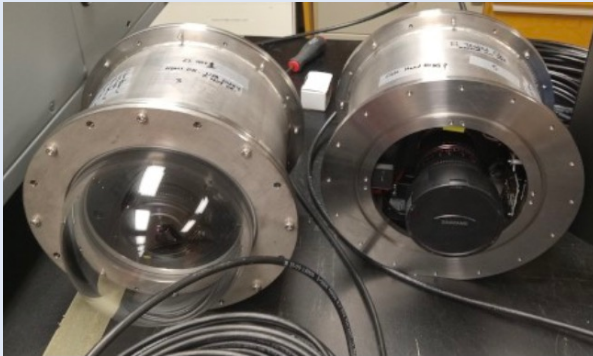
**Detector response**

- PMT related:
  - Gain
  - QE x CE (x HE) (x,y, $\theta$ , $\phi$ )
  - Timing offset (x,y)
  - Linearity
  - After-pulse
  - Dark-rate
- Light propagation
  - Absorption and scattering
  - Position dependence of the above
  - Reflection on the PMTs/BS
- Geometry

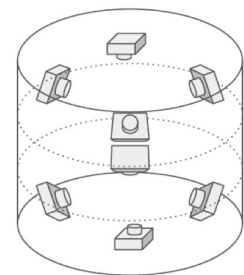
Taken from [here](#)

## Photogrammetry

- Cameras for precise PMT position determination
- Illumination from LED system inside mPMT



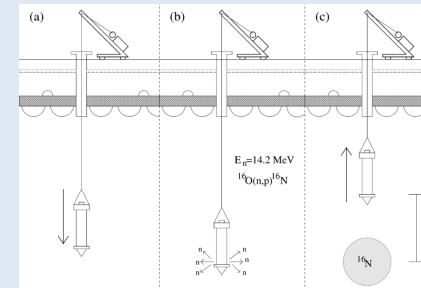
Total coverage = 436 %



8 cameras config B

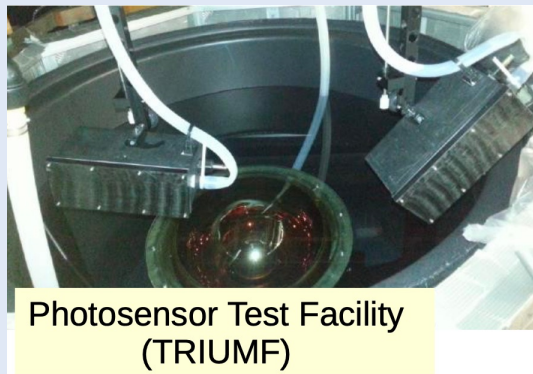
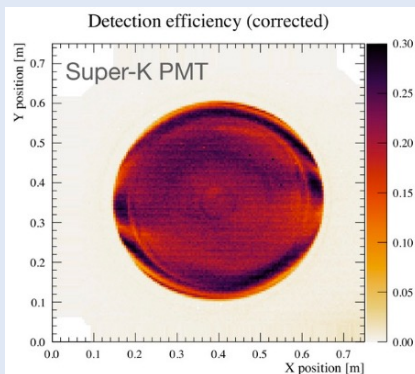
## Radioactive sources

- Measure detector response, PMT efficiency
- $^{16}\text{N}$  source ( $\beta\gamma$  source with well understood spectrum)
- NiCf source (uniform Cherenkov light )



## Precise PMTs pre-calibration

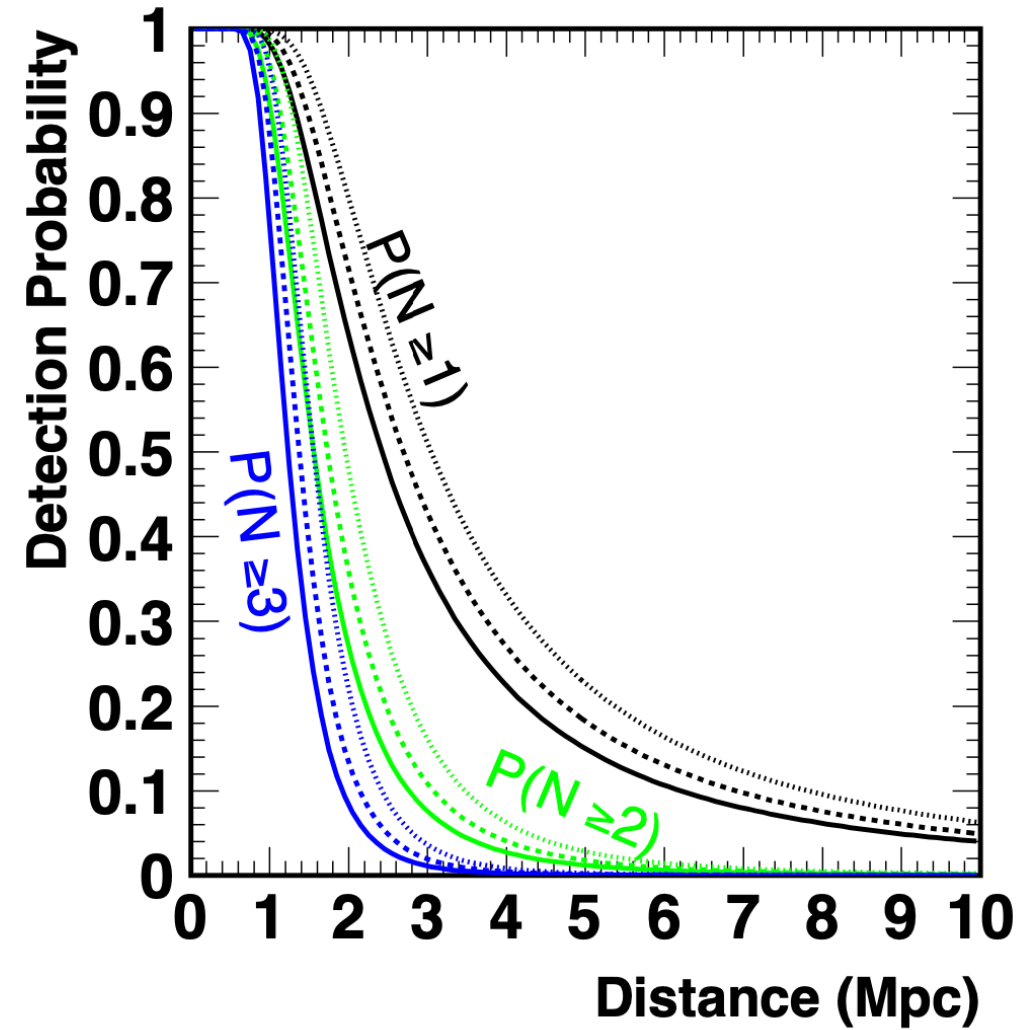
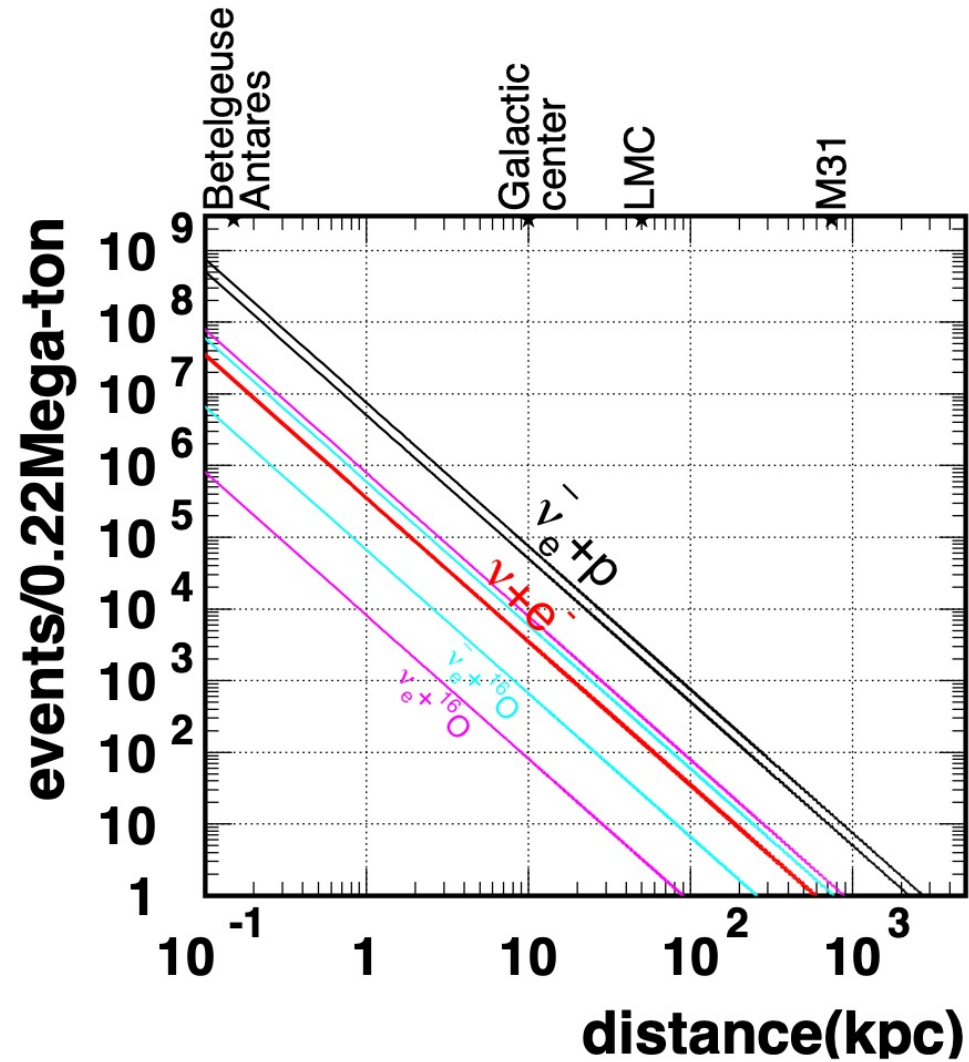
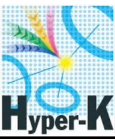
- Make detailed measurement of responses prior to the installation for subset of the PMTs



## mPMTs

- Cross-calibration
  - Disentangle PMT angular response and light traveling direction (granularity)
  - Better separation of indirect photons (precise timing)
- LED source in 200 mPMT
  - 300 nm for Raman scattering

# Supernovae neutrino

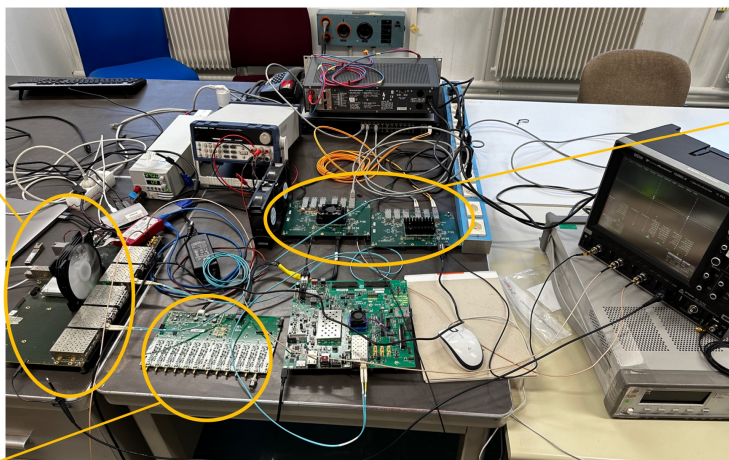


# French contribution (plots)

Test Setup

TDS

TDM

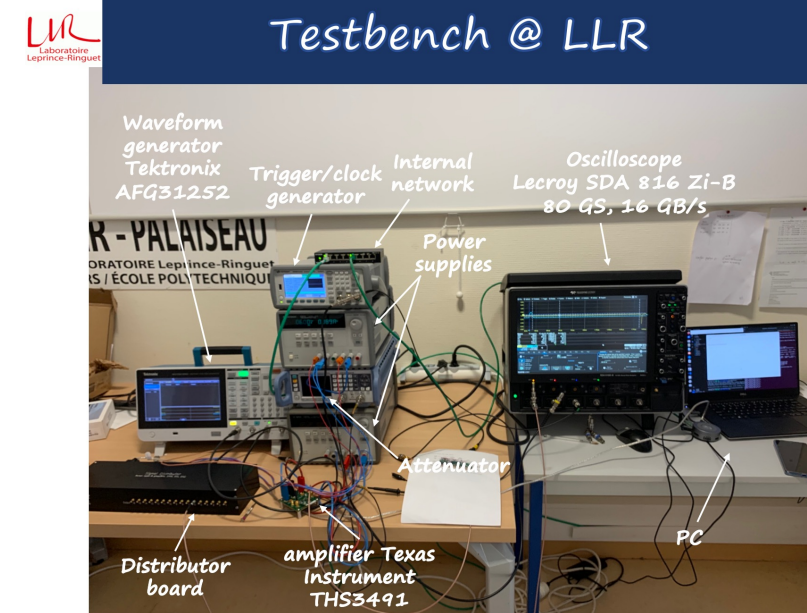
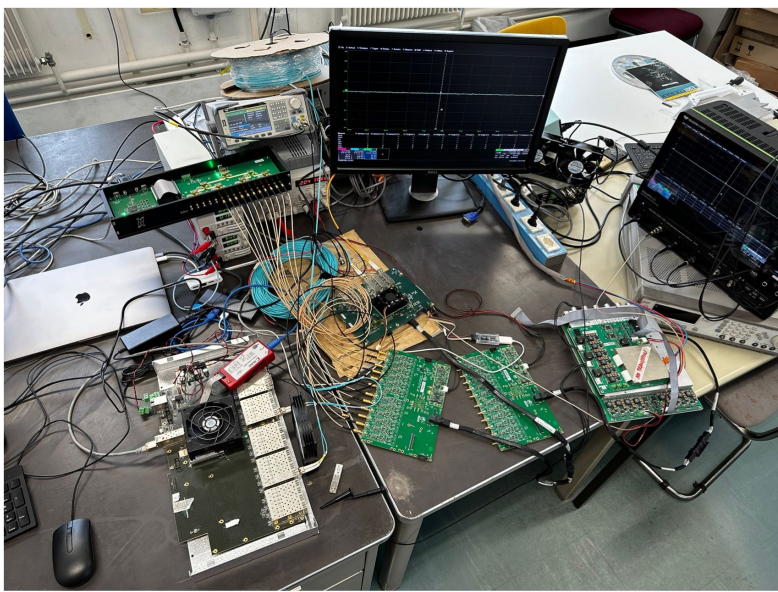


DPB (2)

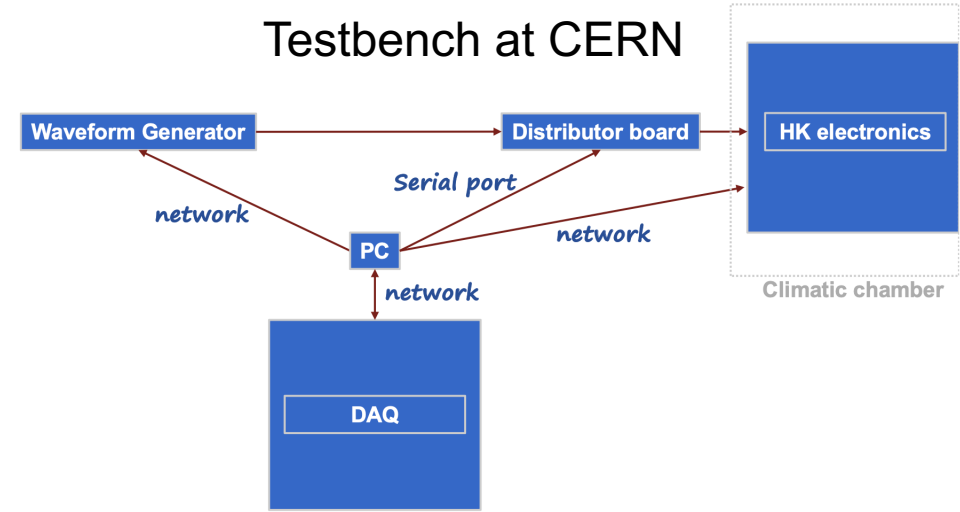
- Lab instruments:
- 24V PS
  - 12V PS
  - Scope Wavepro 254HD-MS
  - Patter gen.
  - Ethernet switch

Digitizer

VST



Testbench at CERN



- Mass production of 50cm PMTs started in 2020
- Production suspended in 2022 due to higher than expected failure rate
- PMT delivery restarted in May 2023, with sampling test of delivered PMTs at Kamioka
- So far, in line to complete delivery of 20.5k by Sep. 2026
  
- For mPMT: design complete outside of LED part. Production not started yet. (19 3" PMTs)
  
- 3600 PMTs for OD will be produced

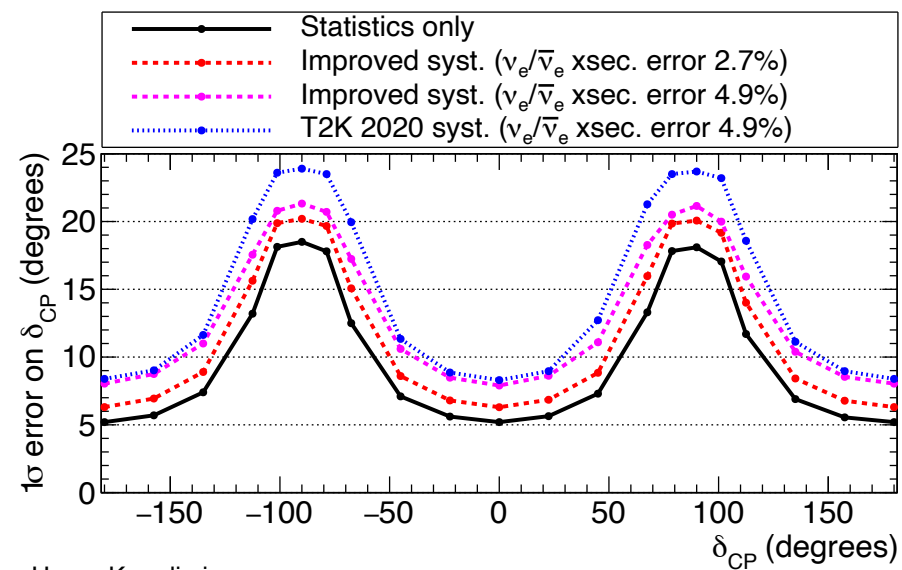
## Computing

- Make CC-IN2P3 a Tiers-1 site of HK experiment
  - Host HK database
  - Discussions with collaboration are ongoing

Detector	MC (HS06 CPU.h)	MC Storage (TB)
INGRID	0.13M	7
ND280	19.2M	2,250
IWCD	97M	52
Far detector	20M	500
<b>Total</b>	<b>136.33M</b>	<b>2,824</b>

## Software

- Oscillation analysis (beam neutrinos, joint beam+atm)
  - Porting T2K and SK tools and accommodating them for Hyper-K
  - Provide comprehensive osc. sensitivity results
  - Develop more robust systematic model
- Hyper-K events reconstruction
  - Porting SK standard algorithm
  - Develop ML approach
- DSNB
  - Phenomenological studies
  - Preparation of new sensitivity studies



Hyper-K preliminary  
 True normal ordering (known), HK 10 Years ( $2.7 \times 10^{22}$  POT 1:3  $\nu \bar{\nu}$ )  
 $\sin^2 \theta_{13} = 0.0218 \pm 0.0007$ ,  $\sin^2 \theta_{23} = 0.528$ ,  $\Delta m_{32}^2 = 2.509 \times 10^{-3} \text{eV}^2/c^4$