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Joint Bayesian calibration and map-making for intensity mapping experiments

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What this talk is not

Not an end-to-end Bayesian model - by design

Reasons

- Component separation is not necessary/essential for gain/noise calibration
- No reliable ground-truth sky model \rightarrow model dependence becomes a liability

Not: "We model this way to minimise degeneracy"

Instead

- Model degeneracies when they exist
- Model ignorance with general, safe representations
- Stay true to what we exactly know

What this talk is about

A Bayesian conservative's framework for calibration and map-making

Timestream of Data
System Temperature

$$d(t_a) = G(t_a) T_{sys}(t_a) (1 + \hat{w}),$$

Instrumental gain
White noise: $\hat{w} \sim \mathcal{N}(0, 1/(\mathcal{T}\Delta\nu))$

Key ingredients

- Instrumental gain modelling (smooth component & 1/f noise)
- Multiplicative noise treatment (no noise model conditionals)
- Joint inference of smooth gain, 1/f noise, and system temperature components
- Experimental setup for demonstration: MeerKAT-type observation in mind

Instrumental Gain Modelling



$$\mathcal{P}_{\widehat{\epsilon}}(f) = egin{cases} 0, & |f| < f_{\mathcal{C}}, \ (f_0/|f|)^lpha\,, & |f| \geq f_{\mathcal{C}}. \end{cases}$$



Stochastic Gain Model



Flicker Noise (Complete Model)

Gaussianity and stationarity.

$$oldsymbol{P}_{\hat{\epsilon}}(f) = egin{cases} 0, & |f| < f_{\mathcal{C}}, \ (f_0/|f|)^lpha\,, & |f| \geq f_{\mathcal{C}}. \end{cases}$$



Stochastic Gain Model

But how to implement the 1/f noise covariance?

 Case 1: **Diagonal in DFT** space

• Case 2: **Diagonal in CFT** space



Stochastic Gain Model

But how to implement the 1/f noise modelling?

Case 2: **Diagonal in CFT** space

$$\begin{aligned} (\mathbf{N}_{\text{corr}})_{aa'} &= \xi(t_{a'} - t_a) \\ \xi(\tau) &= \frac{\Theta_0^{\alpha}}{\pi \tau} \text{Re} \left[\Gamma(\mu, i\Theta_c) \, \boldsymbol{e}^{-i\frac{\pi}{2}\mu} \right] \\ \mu &= \mathbf{1} - \alpha, \Theta_0 = \tau f_0, \Theta_c = \tau f_c \end{aligned}$$



System Temperature Model

Sky + Receiver + Ground + Atmospheric + ...

$$egin{aligned} T^{ ext{sys}}_{j,oldsymbol{
ho}}(t_a) &= \sum_{X} B^X_{oldsymbol{
ho}}(t_a) st T^{ ext{sky},X} \ &+ T^{ ext{el}}_{oldsymbol{
ho}}(t_a) + T^{ ext{nd}}_{j,oldsymbol{
ho}}(t_a) + T^{ ext{rec}}_{j,oldsymbol{
ho}}(t_a) \end{aligned}$$

Linear Model

$$ec{\mathcal{T}}_{\mathrm{sys}} = \mathbf{U}_{\mathrm{ce}}\,ec{\mathbf{p}}_{\mathrm{ce}} + \mathbf{U}_{\mathrm{res}}\,ec{\mathbf{p}}_{\mathrm{res}}$$



Joint Bayesian Analysis Framework

• The data model can be rewritten as

$$d(t_a) \simeq g(t_a) T_{\mathrm{sys}}(t_a) \left[1 + \hat{w} + \hat{\epsilon}\right].$$

• the likelihood function of the data model is given by

$$L(\boldsymbol{n}|\boldsymbol{g}, T_{\mathrm{sys}}, \boldsymbol{\mathsf{N}}) = (2\pi)^{-\frac{N}{2}} |\boldsymbol{\mathsf{N}}|^{-\frac{1}{2}} \exp\left\{\left[-\frac{1}{2}\boldsymbol{n}^{T}\boldsymbol{\mathsf{N}}^{-1}\boldsymbol{n}\right]\right\},$$

where $\mathbf{N} = \mathbf{N}_{w} + \mathbf{N}_{corr}$.

• Gibbs sampling steps:

$$egin{aligned} & g^{(i+1)} \leftarrow \mathcal{P}_{ ext{post}}(g|oldsymbol{d}, T^{(i)}_{ ext{sys}}, oldsymbol{\mathsf{N}}^{(i)}) \ & oldsymbol{\mathsf{N}}^{(i+1)} \leftarrow \mathcal{P}_{ ext{post}}(oldsymbol{\mathsf{N}}|oldsymbol{d}, T^{(i)}_{ ext{sys}}, g^{(i+1)}) \ & \mathcal{T}^{(i+1)}_{ ext{sys}} \leftarrow \mathcal{P}_{ ext{post}}(\mathcal{T}_{ ext{sys}}|oldsymbol{d}, g^{(i+1)}, oldsymbol{\mathsf{N}}^{(i+1)}) \end{aligned}$$

Tractable Implementation for Gain and Tsys samplers Iterative GLS sampler

 Conditional on noise, the data model can be generally written as

 $\boldsymbol{d} = (\boldsymbol{U}\boldsymbol{p} + \boldsymbol{\mu}) \circ (1 + \boldsymbol{n}).$

• Rearrangement - effective additive form:

$$oldsymbol{d}' = oldsymbol{U}oldsymbol{p} + oldsymbol{\epsilon}, \quad oldsymbol{\epsilon} \sim \mathcal{N}(0, \Sigma),$$

where $d' = d - \mu$ and the noise covariance structure is given by:

$$\Sigma = \text{Diag}(\mathbf{U}\boldsymbol{p} \!+\! \mu) \, \mathbf{N} \, \text{Diag}(\mathbf{U}\boldsymbol{p} \!+\! \mu).$$

Estimating $\boldsymbol{\Sigma}$ with iterative GLS:

1. Initialize $\boldsymbol{p}^{(0)}$ using OLS:

$$\boldsymbol{p}^{(0)} = (\mathbf{U}^{\top}\mathbf{U})^{-1}\mathbf{U}^{\top}\boldsymbol{d}'.$$

2. At iteration k, compute

$$\Sigma^{(k)} = \operatorname{diag}(\mathbf{U}\boldsymbol{p}^{(k)} + \mu) \operatorname{N} \operatorname{diag}(\mathbf{U}\boldsymbol{p}^{(k)} + \mu).$$

3. Update *p* by solving:

$$\left[\mathbf{U}^{\top}\left(\boldsymbol{\Sigma}^{(k)}\right)^{-1}\mathbf{U}\right]\boldsymbol{\rho}^{(k+1)}=\mathbf{U}^{\top}\left(\boldsymbol{\Sigma}^{(k)}\right)^{-1}\boldsymbol{d}'.$$

4. Repeat until convergence:

$$\|\boldsymbol{p}^{(k+1)} - \boldsymbol{p}^{(k)}\| < ext{tol}$$

Tractable Implementation for Gain and Tsys samplers Gaussian Constrained Realisation (GCR) equations

Per TOD set sampling:

$$\left(\mathbf{C}^{-1} + \mathbf{U}^{\top} \boldsymbol{\Sigma}^{-1} \mathbf{U}\right) \boldsymbol{\rho}_{\text{sample}} = \mathbf{U}^{\top} \left(\boldsymbol{\Sigma}^{-1} \boldsymbol{d}' + \boldsymbol{\Sigma}^{-\frac{1}{2}} \boldsymbol{\omega}\right) + \mathbf{C}^{-1} \bar{\boldsymbol{\rho}} + \mathbf{C}^{-\frac{1}{2}} \boldsymbol{\eta},$$

Multiple TOD sets joint fitting:

$$\left(\mathbf{C}^{-1} + \sum_{j} \mathbf{U}_{j}^{\top} \mathbf{\Sigma}_{j}^{-1} \mathbf{U}_{j}\right) \mathbf{p}_{\mathsf{sample}} = \sum_{j} \mathbf{U}_{j}^{\top} \left[\mathbf{\Sigma}_{j}^{-1} \mathbf{d}_{j}' + \mathbf{\Sigma}_{j}^{-\frac{1}{2}} \omega_{j}
ight] + \mathbf{C}^{-1} ar{\mathbf{p}} + \mathbf{C}^{-\frac{1}{2}} \eta,$$

Tractable Implementation for Sampling 1/f **Parameters** COMAT: An $O(n^2)$ implementation

$$\mathbf{N}_{corr} = \begin{pmatrix} \xi_0 & \xi_1 & \xi_2 & \cdots & \xi_{n-1} \\ \xi_1 & \xi_0 & \xi_1 & \cdots & \xi_{n-2} \\ \xi_2 & \xi_1 & \xi_0 & \cdots & \xi_{n-3} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ \xi_{n-1} & \xi_{n-2} & \xi_{n-3} & \cdots & \xi_0 \end{pmatrix}$$

• Levinson-Durbin Recursion

Perturbative Scenario

$$(\Lambda + \mathbf{P})^{-1} \simeq \Lambda^{-1} - \sum_{k=0}^{m-1} \Lambda^{-1} \mathbf{P} \left[-\mathbf{P} \Lambda^{-1} \right]^k \Lambda^{-1}.$$

$$\det egin{pmatrix} A & B \ C & D \end{pmatrix} = \det(A)\det(D-CA^{-1}B).$$

$$\begin{split} \ln[\det(\Lambda+\mathbf{P})] &\approx \sum_{k=0}^{m-1} \frac{(-1)^k}{k+1} \operatorname{Tr}\left[\left(\mathbf{P} \Lambda^{-1}\right)^{k+1} \right] \\ &+ \ln[\det(\Lambda)] \end{split}$$

3. Bayesian Workflow

3.2. 1/f Sampler

Tractable Implementation for Sampling 1/f Parameters



3.

Illustrating Examples Experimental Setup



(a) 1×TOD: Scan



(b) 2×TOD: Scan



(a) 1×**TOD**: Sky



(b) 2×TOD: Sky



(a) $1 \times \text{TOD}$: Integrated beam.



(b) $2 \times \text{TOD}$: Integrated beam_{14/20}

4. Demonstration

Illustrating Examples

Prior Setup

- DC mode: Prior STD $\sim 20\%$
- Other modes: Flat priors

1/f Prior

- Power law index:
 Flat prior
- **Reference (Knee) frequency:** Flat prior

System Temperature Prior (Rough Sky Knowledge + Cal. Src.)

- Prior STD per pixel: 10 K
- Flux scale calibration:
 - Sharp prior on one pixel ("1 CalSrc")
- Other components (receiver temperature, ground pickup, etc)
 - No prior! If you choose to model ignorance, that's because ignorance is your prior!



(a) Estimated (1×TOD; 1 CalSrc)



(b) Estimated (2×TOD; 1 CalSrc)





(a) Residual (1×TOD; 1 CalSrc)



(b) Residual (2×TOD; 1 CalSrc)



(a) Uncertainty (1×TOD; 1 CalSrc)



(b) Uncertainty (2×TOD; 1 CalSrc)

An illustrating Example

Histogram of the residuals across pixels



Figure: 1×TOD; 1 CalSrc

Figure: 2×TOD; 1 CalSrc

CONCLUSION & DISCUSSION

• Consider Model Ignorance:

Beyond end-to-end modelling with educated guesses, it's also valuable to explicitly model our ignorance—identifying and parameterizing what we don't know.

• LEVERAGE INTERNAL CONSTRAINTS:

Design experiments strategically to exploit internal constraints that limit the discriminability between different components in the data model.

• THINK LIKE THE TELESCOPE:

MeerKAT-like observations can distinguish system temperature components by their differing temporal behaviors—use this perspective to guide modelling.

• BAYESIAN MODELLING WITH IGNORANCE SHOULD EVOLVE: Building a pipeline is only the beginning—learning and adapting both the prior and the model itself is essential as understanding deepens.

Echoes from Earlier Sessions | Ad: SPYDUST

 an improved and extended Python implementation for modelling spinning dust emission

Zhang & Chluba JCAP03(2025)038

• A full Stokes implementation with reduced dimensionality (via moment expansion)

Zhang & Chluba in prep.

 Improved Fokker-Planck treatment better rotational statistics

Zhang & Chluba in prep.





Echoes from Earlier Sessions ||

"redshift-space distortions" (RSD)

This is CORRECT!

$$P_s(k,\mu) = (1 + \beta^2 \mu(k)^2)^2 P_r(k)$$

- Large scale, but smaller than the curvature scale: Cartesian plane wave synthesis
- $\mu(k)$: defined with the normal mode peculiar velocity
- Scalar perturbations $\vec{u}(\vec{k}) \parallel \vec{k}$
- $\mu(k)$: NOT directly given by LOS projection of real space peculiar velocity