



The LIM Super Sample Covariance

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What Is the Super Sample Covariance and Do We Care?

- **Nonlinear** evolution of matter → **non-Gaussianities**
- Contributions of tri spectrum to covariance
- Super survey modes
 - different evolution of nonlinearities
 - Super sample covariance

What Is the Super Sample Covariance and Do We Care?

- Well studied for galaxy redshift surveys
- Effects larger for **smaller** surveys
- Especially important at **small scales**

⇒ Study the effect for typical LIM surveys

The Sample Covariance without a Window

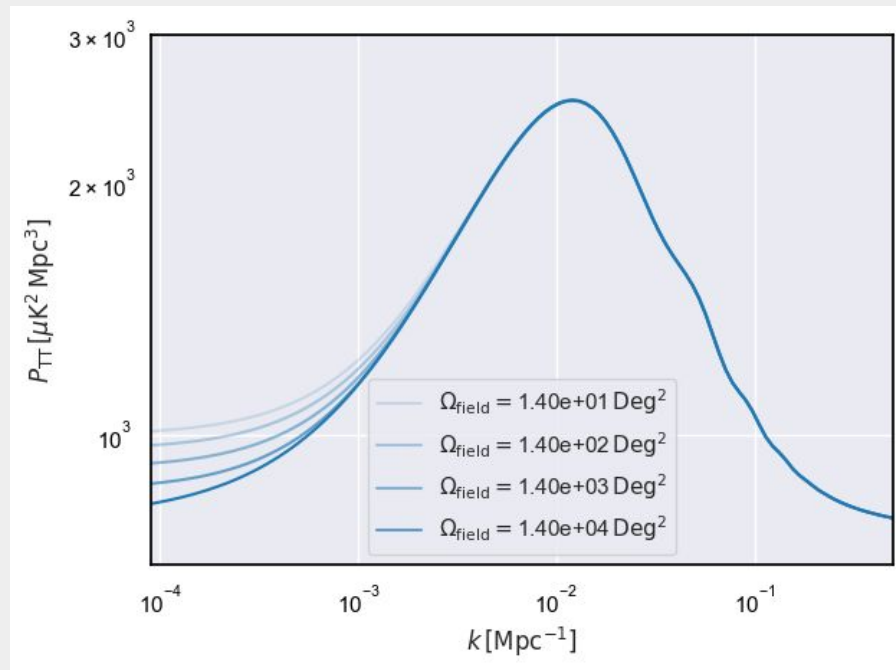
The non-Gaussian part is related to
the **connected 4-point function**

$$\text{Cov}(k_1, k_2) = \text{Gaussian}(k_1) \delta_{k_1, k_2} + \text{nonGaussian}(k_1, k_2)$$

Effect of the Survey Window

In Fourier space:

- Perturbations are smoothed over characteristic scale
- Smaller surveys have larger smoothing scales
- Effect on power spectrum



Effect of the Survey Window

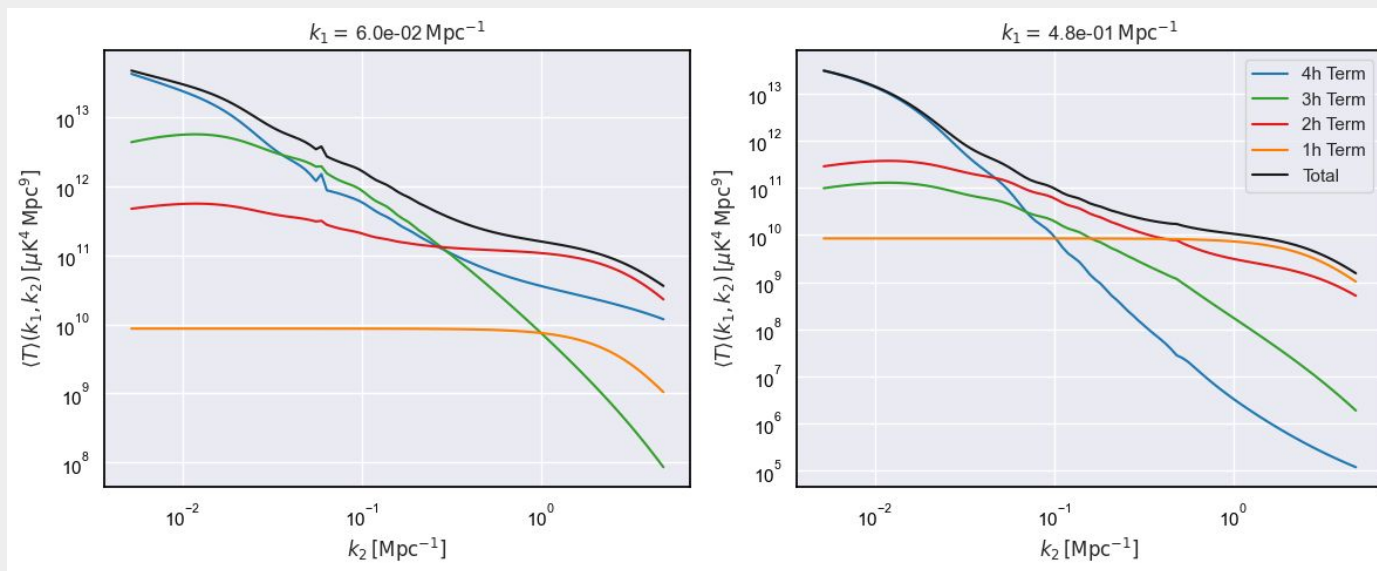
For the tri spectrum:

We **expand** the non-Gaussian covariance
in k_i over smoothing scale

$$\widehat{\text{nonGaussian}} = \text{nonGaussian} + \text{SSC}$$

Non Gaussian Covariance

Compute the tri spectrum using
LIM halo model $\times$ standard perturbation theory



The Super Sample Covariance

$$\text{SSC}(k_1, k_2) = \frac{\sigma^2}{V_{\text{field}}} R(k_1) P_{\text{TT}}(k_1) R(k_2) P_{\text{TT}}(k_2)$$



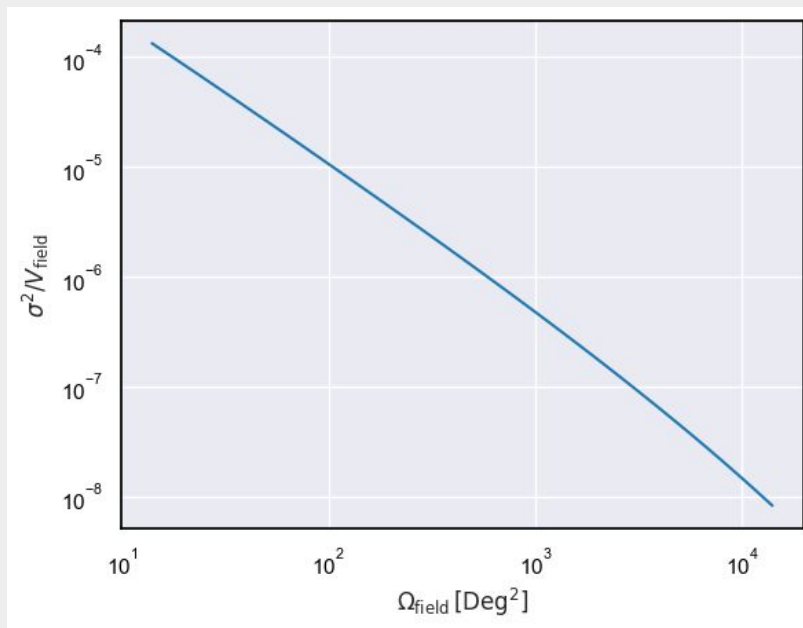
Variance of δ smoothed
over the window



Response of P to large
scale modes

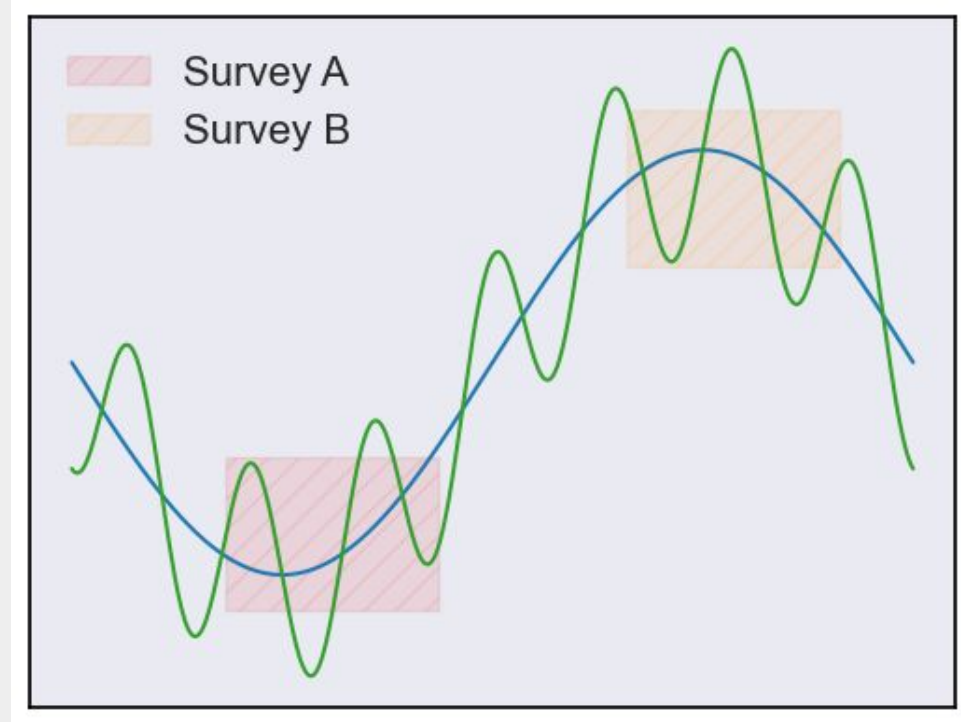
The Super Sample Covariance

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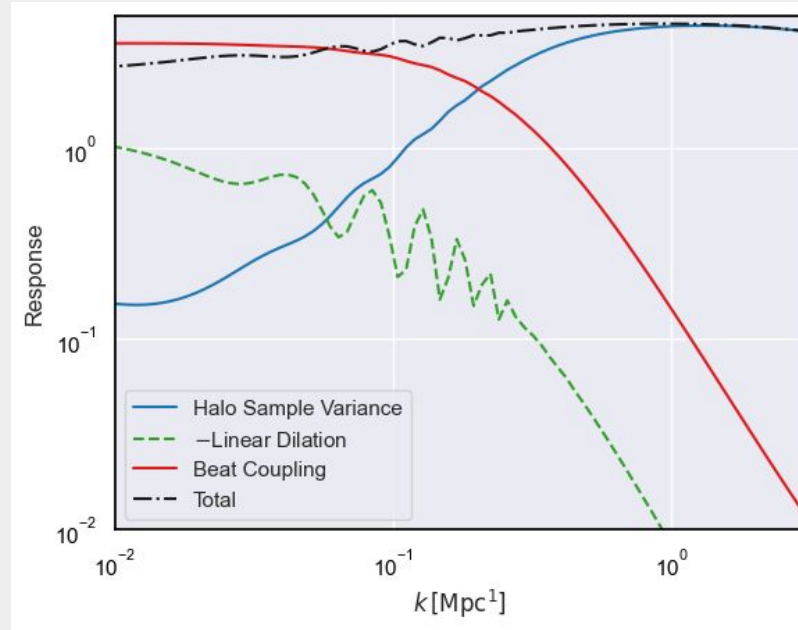
The Super Sample Covariance

The LIM **observables** depend on the particular **large wavelength** modes within a particular survey



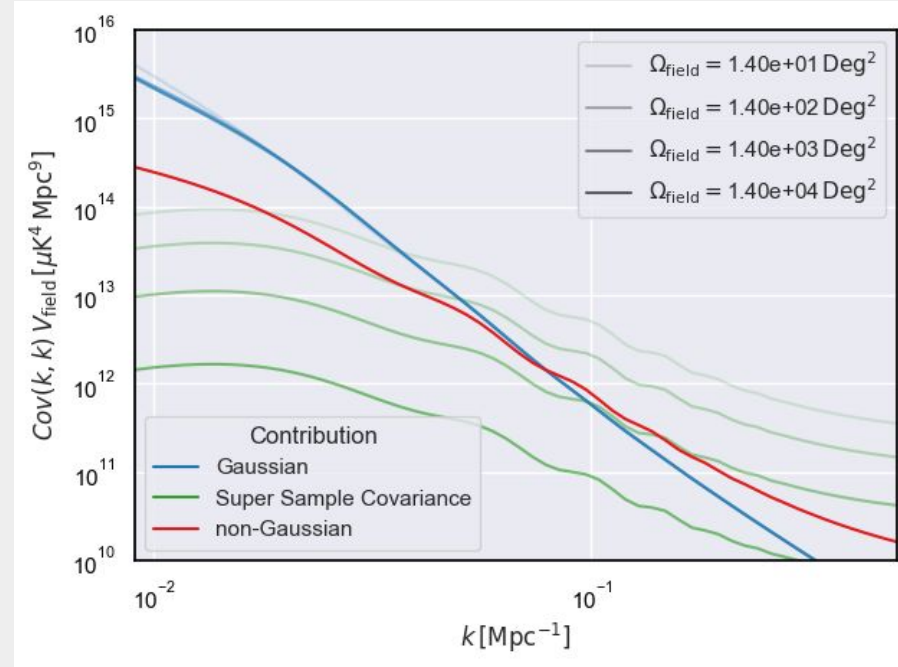
The Super Sample Covariance

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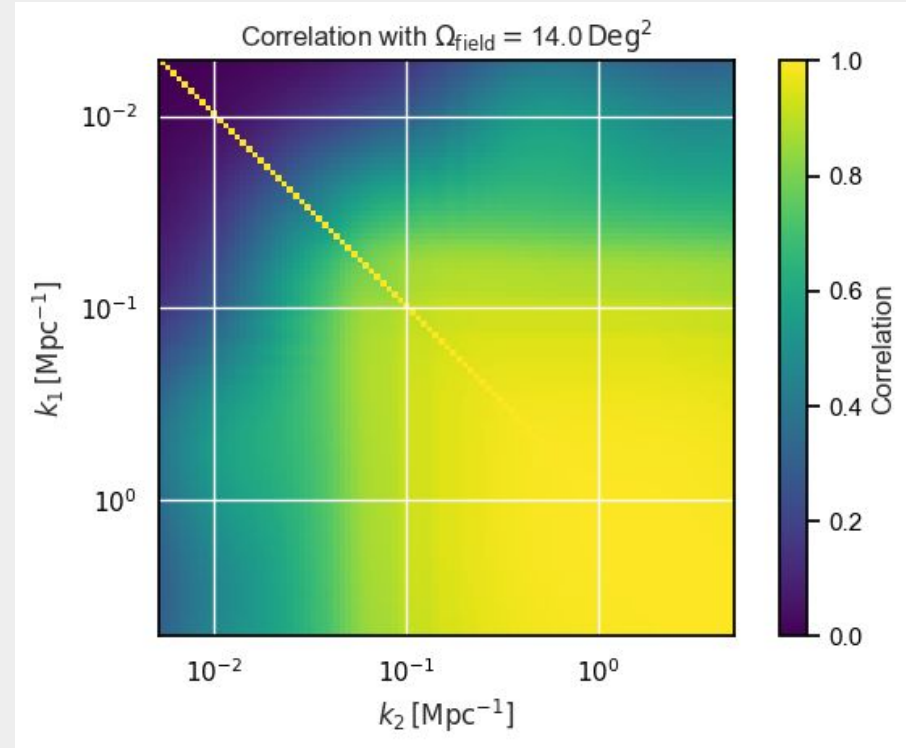
The Full Sample Covariance

1. Small survey volumes
→ Dominant SSC already at large scales
2. Larger survey volumes
→ tri spectrum contribution becomes significant



The Full Sample Covariance

3. Non-Gaussianities
introduce significant
off diagonal terms



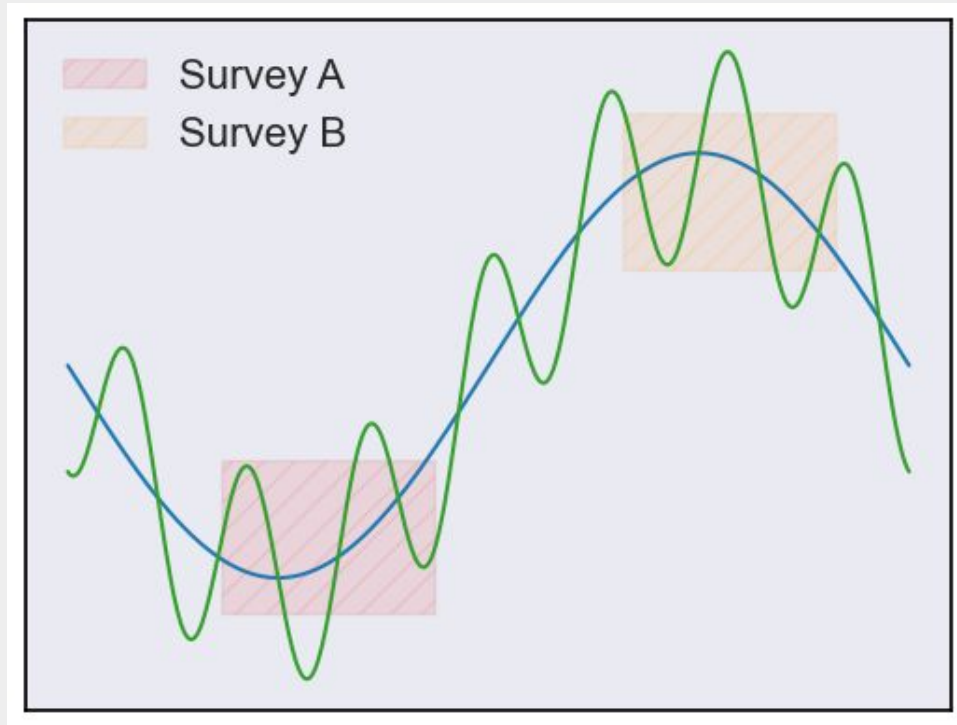
Non-Analytic Covariance Estimates

- Only (non-)Gaussian can be estimated from data
- For the SSC super survey modes have to be considered
- Can be extracted from simulations (Field-to-Field Variance)

Interpretation

Detections: SSC is measure for deviation from cosmological average

Parameter inference: SSC adds additional uncertainty to measurements



Key Takeaways

Nonlinear evolution leads to strong **non-Gaussianities**

Small volumes have additional source of super survey uncertainty

Could be important vor **LIM!**

With the detection era on the horizon we want to be ready to interpret the data

Find the code at Github →



<https://github.com/Sefa76/SSLimPy>