Precision Luminosity Measurements in CMS with Run 2 and Run 3 Data



Peter Major on behalf of the CMS Collaboration

<u>EPS-HEP</u>, Marseille 2025. 07. 08.

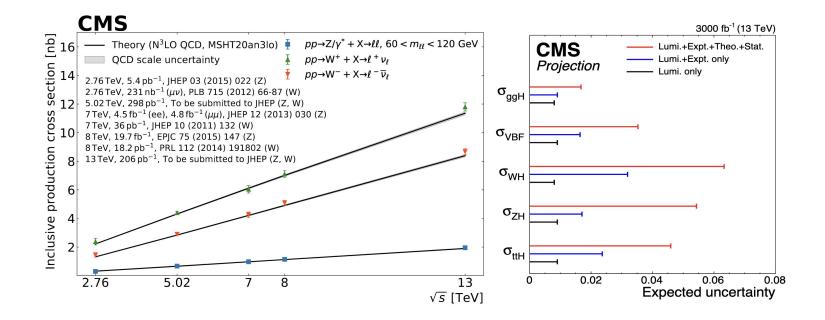
What is luminosity, why do we care?



Luminosity

- Is a measure of the accumulated data
- Connects theory and experiment: $\sigma_{\text{process}} L_{\text{int}} = \langle N_{\text{total}} \rangle \rightarrow \text{used in all xsec measurements}$
- Is amongst the leading sources of experimental uncertainties in SM precision measurements

In lepton colliders it is measured using benchmark physics processes like Bhabha-scattering ($\sigma_{process}$ very well known), but hadron colliders pose many challenges on account of the protons being composite particles (non-trivial PDFs) \rightarrow large production cross section uncertainties



Outline



Detectors

- Redundancy
- Diverse technologies
- Multiple ranges in occupancy

Calibration

- The van der Meer method
- Non-collision background
- Bunch intensity
- Beam-beam interactions
- Beam positions
- Lengthscale calibration
- Non-factorisation
- Emittance scan evolution
- Unknown biases

Integration

- Out-of-time corrections
- Emittance scans
 - Efficiency tracking
 - Non-linearity
- Residual effects
 - Consistency
 - Linearity
- Average luminosity

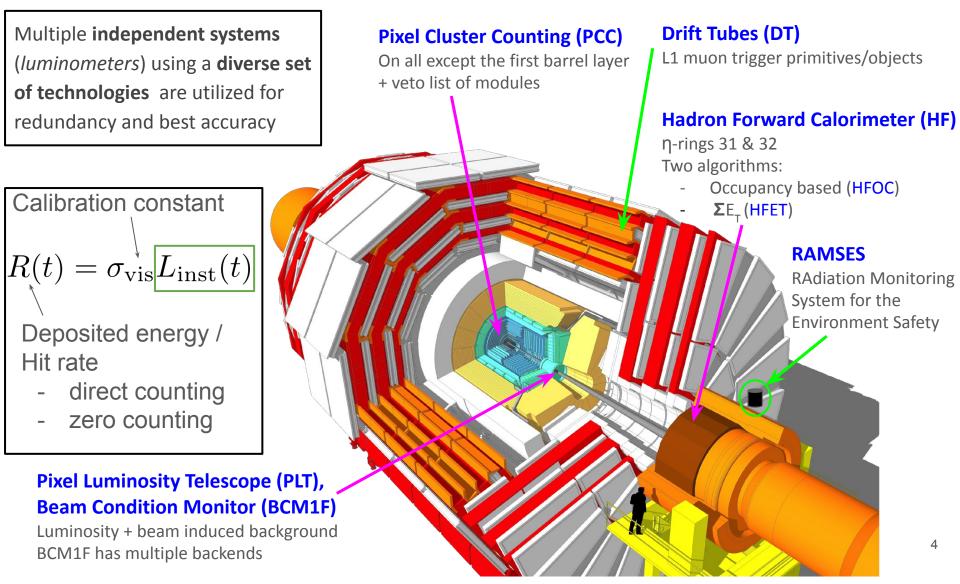
Standard candle proxies

- Z boson rate counting
- Muon pair production in ultraperipherial collisions

More general overview available in <u>ICHEP24</u> presentation.

What hardware is used?





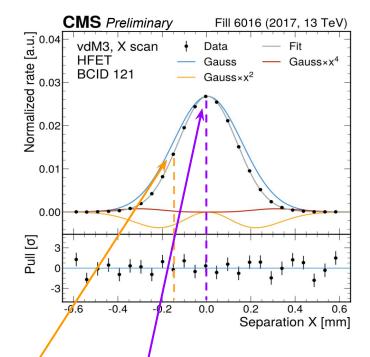


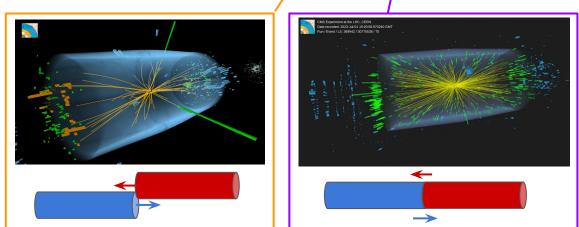
Calibration: Establishing absolute luminosity in well controlled conditions

The van der Meer method

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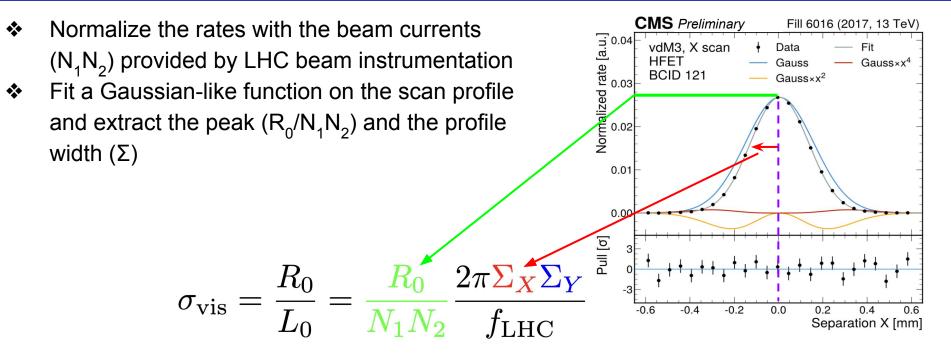
- Highly controlled special conditions:
 - Once a year
 - Wide beams finer relative control
 - Low PU reduced linearity effects
 - Isolated bunches reduced out-of-time
 - Tailored bunch tails in injector chain bunch distributions are approximately *factorisable*
- Perform an X and a Y beam separation scan

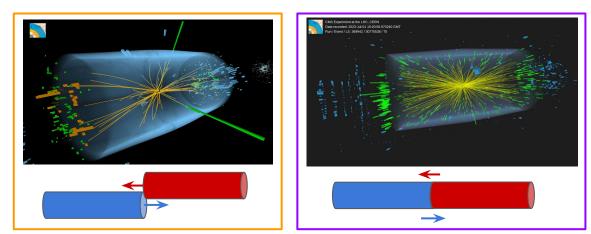




The van der Meer method



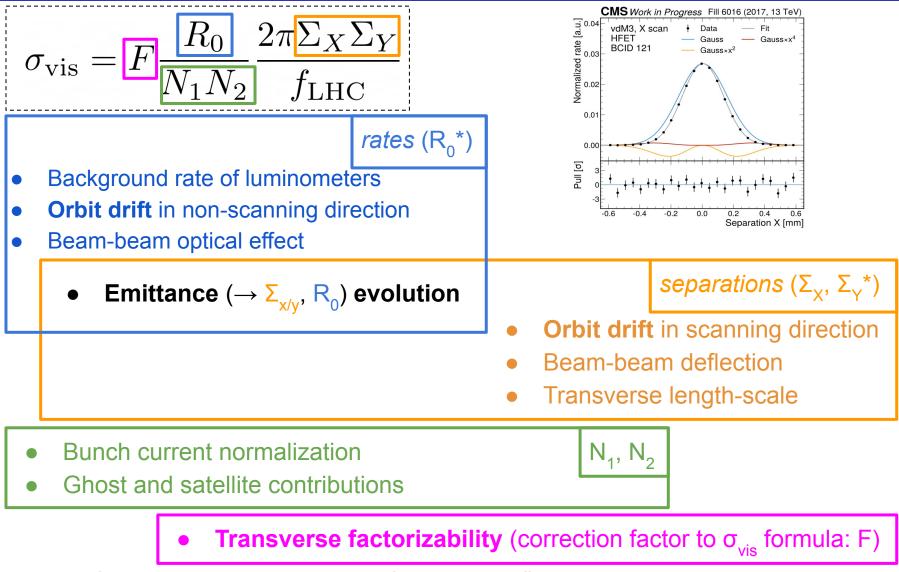




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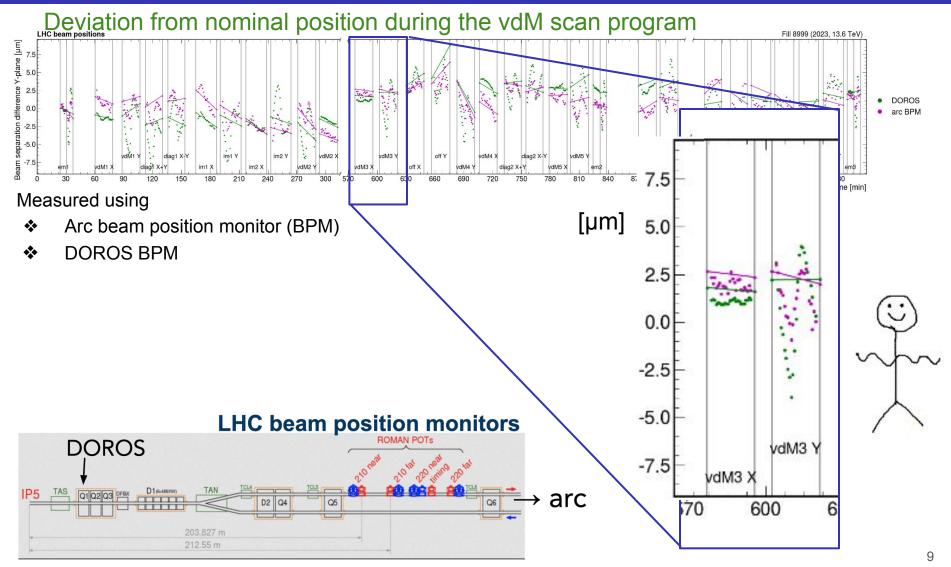
Corrections in the vdM procedure





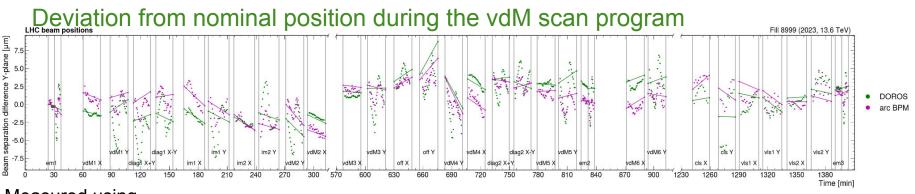
Orbit drift systematics





Orbit drift systematics





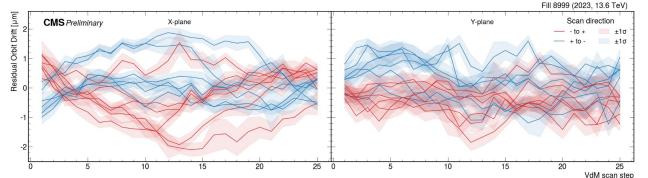
Measured using

- Arc beam position monitor (BPM)
- DOROS BPM

Contributes:

- Slow, linear orbit drift (estimated from before- and after-scan head-on readings)
- Beam-beam deflection (BB) (Bassetti-Erskine formula)
- Residual OD extracted as the residuals of the fit (only scanning plane fit shown):

 $\mathsf{BPM}_{x/y} - \mathsf{linOD}_{x/y} = \alpha \times \mathsf{Nominal}_{x/y} + \beta \times \mathsf{BB}_{x/y}(\Delta \mathsf{Nominal}_{x/y}) + \mathsf{c}_{x/y}$

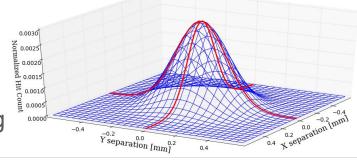


Fitted parameters: Lengthscale (BPM) BB dilution Constant

Typical OD uncertainty in 2022-2023: ~0.2% Large improvement since 2015-16 paper (0.5-0.8%)

Non-factorisation



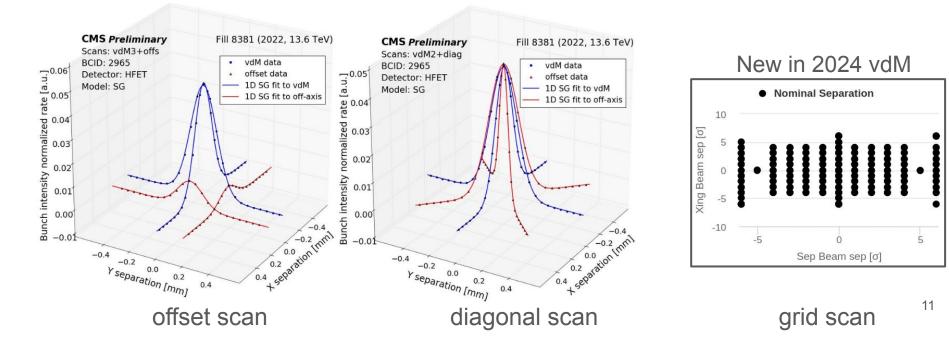


VdM method assumes R(x,y) = f(x)g(y)

 \rightarrow two scans are enough to get the integral of R(x,y)

2D scans

- Fits the bunch overlap shape directly
- Using complementary scans for off-axis sampling
- All BCIDs are used
- Modelling uncertainty dominates
- Luminous region analysis

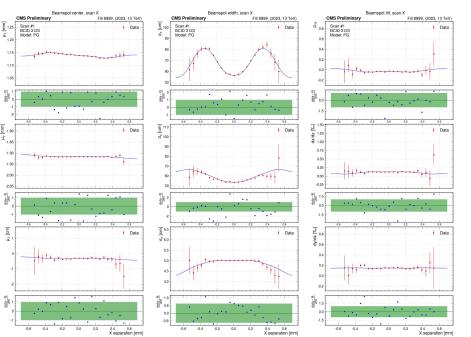


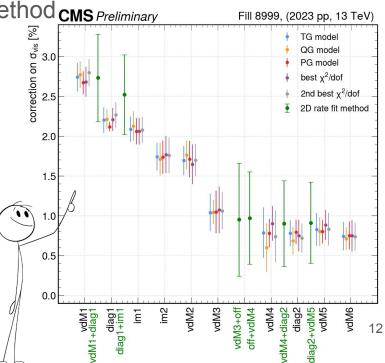
Non-factorisation



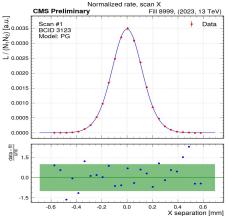
two scans are enough to get the integral of R(x,y)

- 2D scans
- Luminous region analysis 쇇
 - Fits the 3D bunch density function for the two beams
 - Using any scans
 - For few BCIDs with high rate vertex data
 - Uncertainty dominated by closure of the method CMS Preliminary





Uncertainty: 2022 (prelim): 0.8% 2023 (prelim): 0.7%





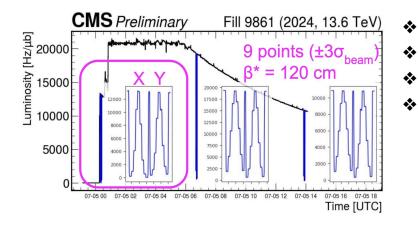


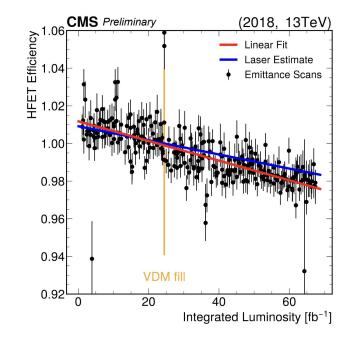


Integration: Measurement in high PU conditions

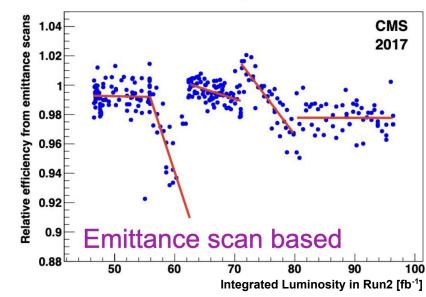
Rate corrections - Efficiency







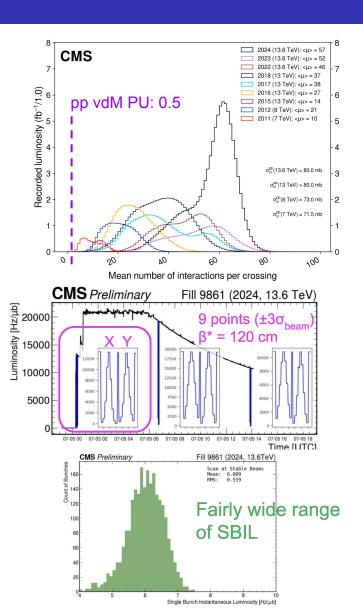
- Aging due to radiation
- Changing conditions (HV, temp, failing modules)
- All detectors potentially affected
 - Intrinsic correction: Emittance scan-based efficiency tracking (per-module for PLT, BCM1F)
 - Good agreement with alternative methods (Laser-based for HF, tracking-based for PLT)



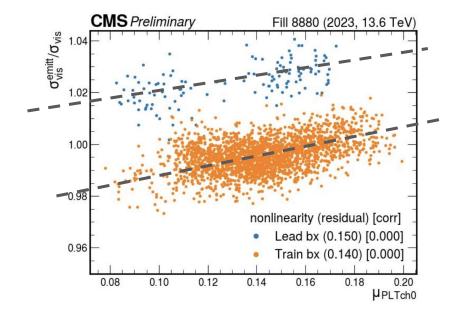
2017 PLT efficiency corrections

Rate corrections - Nonlinear response





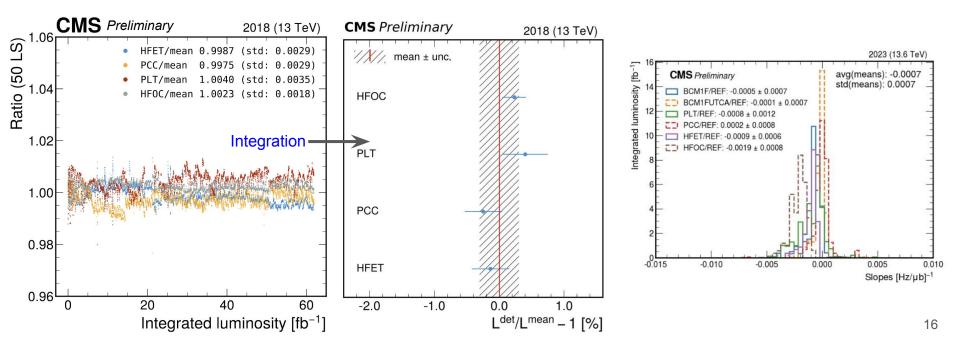
- VdM calibration performed at 1/100 of the datataking pile-up
- Model: $\mu_m = \mu_\ell (1 + \alpha \mu_\ell)$
- Mitigation correction of detectors based on intrinsic quantities:
 - Restrictive module selection based on noise levels and internal consistency (PCC)
 - Efficiency as a function of peak luminosity (SBIL) tracked via emittance-scans (per-module for PLT, BCM1F)



Closure: Consistency, non-linearity



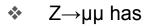
- Previously: Detectors ordered according to their dependability: The best available source provides the luminosity
- Current non-preliminary approach: Several detectors calibrated independently to a similar quality → use the average of the available sources
- Spread of detectors is tracked throughout the whole year
 - Uncertainty derived from the RMS the mean of all histograms
- Residual nonlinearity of the average lumi is evaluated with respect to DT and RAMSES, the more conservative estimate is used



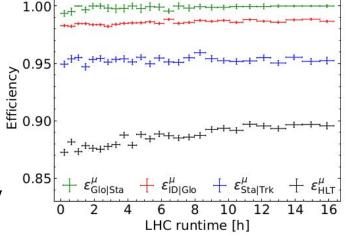
A pp standard candle: Z counting



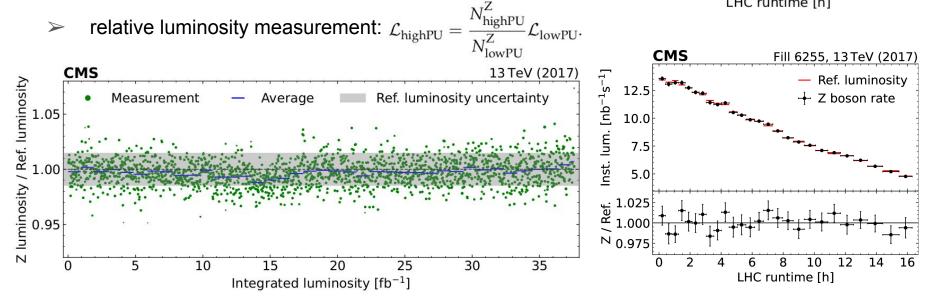
Fill 6255. 13 TeV (2017)



- a clean signature
- relatively large cross section (not enough for vdM)
- a not-too-well-known fiducial cross section (PDF)
- * Trigger and selection efficiencies are measured in situ every $20/pb \rightarrow$ intrinsic linearity and efficiency correction
- Primary use:
 - common ground for consistency checks at given energy



CMS



See: (Eur. Phys. J. C 84 (2024) 26) ¹⁷

Recent results



- Multiple independent luminometers relying on diverse technologies
- Several corrections applied in calibration and integration new approaches highlighted in table
 - Dominant sources: Factorisation, Integration, Beam-Beam, Orbit drift,
- Uncertainties treated as 100% or 0% correlated between years in combinations (see colors)
- Recent preliminary results approach 1% uncertainty, foreshadowing the upcoming Run2 precision result

| Results since 2020 | 2015 | 2016 | 2022 (prelim) | 2023 | pp ref 2017 | PbPb | PbPb |
|-----------------------------------|--------------------|------|------------------------|---------------------|------------------------|-------------------|------|
| | | | | (prelim) | (prelim) | 2015 | 2018 |
| | EPJ C81 (2021) 800 | | CMS-PAS- LUM-22-001 | CMS-DP-2 024-068 | CMS-PAS- LUM-19-001 | Submitted to EPJC | |
| Non collision rate | | | | | | 0.5 | 0.2 |
| Statistical | _ | _ | — | _ | <0.1 | 0.1 | 0.1 |
| Beam current | 0.1 | 0.1 | 0.2 | 0.20 | 0.2 | 0.2 | 0.2 |
| Ghost & satellite charges | 0.2 | 0.2 | 0.2 | 0.10 | 0.2 | 0.3 | 0.5 |
| Beam-beam effects | 0.5 | 0.5 | 0.4 | 0.34 | 0.8 | 0.2 | 0.3 |
| Linear (random) orbit drift | 0.2 | 0.1 | 0.1 | 0.02 | 0.3 | 0.5 | 0.1 |
| Residual (systematic) orbit drift | 0.8 | 0.5 | 0.3 | 0.16 | 1.0 | 0.2 | 0.2 |
| Length scale | 0.2 | 0.3 | 0.1 | 0.20 | 0.8 | 0.5 | 0.5 |
| Factorization bias | 0.5 | 0.5 | 0.8 | 0.67 | 0.8 | 1.1 | 1.1 |
| Scan-to-scan | | | 0.5 | 0.28 | 0.4 | _ | 0.5 |
| Bunch-to-bunch | 0.6 | 0.3 | 0.1 | 0.06 | 0.4 | — | _ |
| VdM consistency | | | 0.4 | 0.16 | 0.4 | 2.5 | 0.4 |
| Calibration | 1.3 | 1.0 | 1.2 | 0.89 | 1.9 | 2.9 | 1.5 |
| OOT (non coll. rate) | 0.3 | 0.4 | 0.2 | _ | <0.1 | 0.1 | 0.1 |
| Stability | 0.6 | 0.5 | 0.5 | 0.71 | 0.1 | 0.7 | 0.8 |
| Linearity | 0.5 | 0.3 | 0.5 | 0.59 | <0.1 | _ | - |
| Integration | 1.0 | 0.7 | 0.8 | 0.92 | 0.1 | 0.7 | 0.8 |
| Total | 1.6 | 1.2 | 1.4 | 1.28 | 1.9 | 3.0 | 1.7 |

Thank you!

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Zero counting



- In certain detectors directly counting individual hits is not feasible either due to resolution or bandwidth / computational limitations, but it is very possible to determine the the lack of a hit (the opposite of any number of hits)
- The hits follow a **Poisson distribution**: $P(n=k) = e^{-\lambda} \lambda^k / k!$
- The probability of zero hits is $P(n=0) = e^{-\lambda}$
- Therefore the mean hit count is $\lambda = -\ln(P(n=0))$
- **\diamond** Zeros are counted over several orbits before $\ln(n_0/n)$ is calculated
- At high pile-up zero-starvation can become a problem, as the logarithm explodes near 0, amplifying the noise of the detector and introducing a bias
- Occupancy measurement via zero counting in PLT, BCM1F*, HFOC

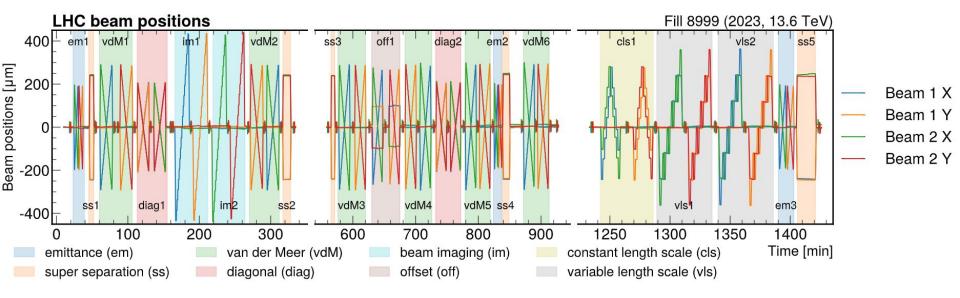
What is σ_{vis} ?



$$\begin{split} \mathsf{L}(\Delta \mathbf{x}, \Delta \mathbf{y}) &= \mathsf{n}_1 \mathsf{n}_2 \mathsf{f}_{\mathsf{LHC}} \int_{\mathbb{R}^2} \mathsf{d} \mathbf{x} \mathsf{d} \mathbf{y} \, \mathsf{b}_1 \Big(\mathbf{x} - \frac{\Delta \mathbf{x}}{2}, \mathbf{y} - \frac{\Delta \mathbf{y}}{2} \Big) \mathsf{b}_2 \Big(\mathbf{x} + \frac{\Delta \mathbf{x}}{2}, \mathbf{y} + \frac{\Delta \mathbf{y}}{2} \Big), \text{ for } \mathsf{b}_1, \mathsf{b}_2 \text{ bunch density functions} \\ \\ \int_{\mathbb{R}^2} \mathsf{d} \Delta \mathbf{x} \mathsf{d} \Delta \mathbf{y} \, \mathsf{L}(\Delta \mathbf{x}, \Delta \mathbf{y}) &= \mathsf{n}_1 \mathsf{n}_2 \mathsf{f}_{\mathsf{LHC}}, \\ \hline \\ \mathsf{f}_{\mathbb{R}^2} \, \mathsf{d} \Delta \mathsf{x} \mathsf{d} \Delta \mathbf{y} \, \mathsf{L}(\Delta \mathbf{x}, \Delta \mathbf{y}) &= \frac{1}{\sigma_{\mathsf{vis}}} \int_{\mathbb{R}^2} \mathsf{d} \Delta \mathsf{x} \mathsf{d} \Delta \mathbf{y} \, \mathsf{R}(\Delta \mathbf{x}, \Delta \mathbf{y}) = \mathsf{n}_1 \mathsf{n}_2 \mathsf{f}_{\mathsf{LHC}}, \\ \hline \\ \mathbf{\sigma}_{\mathsf{vis}} &= \frac{1}{\mathsf{n}_1 \mathsf{n}_2 \mathsf{f}_{\mathsf{LHC}}} \int_{\mathbb{R}^2} \mathsf{d} \Delta \mathsf{x} \mathsf{d} \Delta \mathbf{y} \, \mathsf{R}(\Delta \mathbf{x}, \Delta \mathbf{y}) \\ \exists f, g: R(\Delta x, 0) &= f(\Delta x) g(0) \\ R(\Delta x, 0) &= f(\Delta x) g(0), R(0, \Delta y) = f(0) g(\Delta y) \\ R(\Delta x, \Delta y) &= \frac{R(\Delta x, 0)}{g(0)} \frac{R(0, \Delta y)}{f(0)} = \frac{R(\Delta x, 0) R(0, \Delta y)}{R(0, 0)} \\ \int_{\mathbb{R}^2} R(\Delta x, \Delta y) &= \frac{1}{R(0, 0)} \int_{\mathbb{R}} R(\Delta x, 0) \int_{\mathbb{R}} R(0, \Delta y) \\ \int_{\mathbb{R}} R(\Delta x, 0) &= \sqrt{2\pi} R(0, 0) \Sigma_X \end{split}$$

The van der Meer method





Length scale calibration (LSC)



beam 1 beam 2

PbPb 2015 (5.02 TeV)

Slope: -0.011 ± 0.001 $\chi^2/d.o.f. = 0.4 / 2$

Backward

× beamspot

 1.4σ

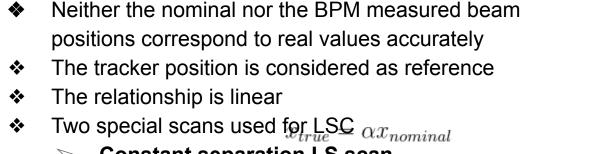
CMS Preliminary

- Forward

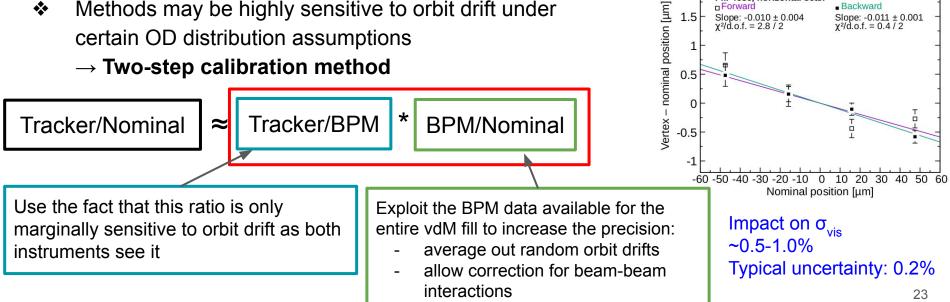
1.5

Fill 4689, horizontal scan

Slope: -0.010 \pm 0.004 χ^2 /d.o.f. = 2.8 / 2



- **Constant separation LS scan** \succ
 - Average LS for B1&B2
- Variable separation LS scan >
 - Separate LS for B1&B2
- Methods may be highly sensitive to orbit drift under * certain OD distribution assumptions
 - \rightarrow Two-step calibration method



Emittance evolution

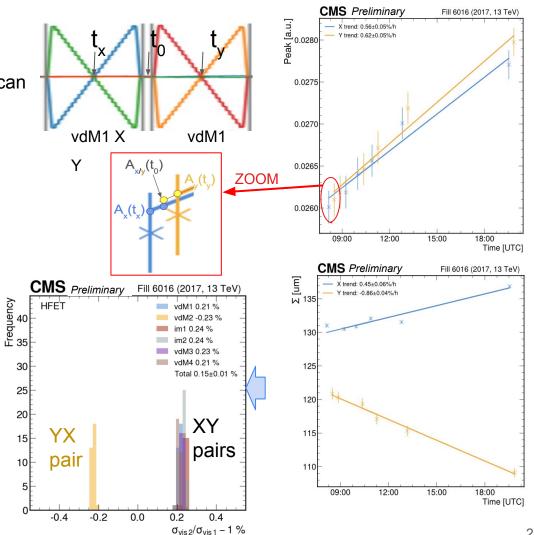
Issue:

The vdM profile parameters are constantly changing in time

 \rightarrow the parameters extracted in the X and Y scan are only approximately compatible

$$egin{aligned} \sigma_{vis1} &= 2\pi rac{A_x(t_x) + A_y(t_y)}{2} \Sigma_x(t_x) \Sigma_y(t_y) \ \sigma_{vis2} &= 2\pi rac{A_x(t_0) + A_y(t_0)}{2} \Sigma_x(t_0) \Sigma_y(t_0) \end{aligned}$$

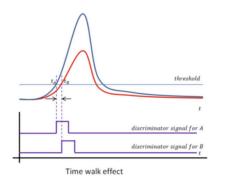
The impact is $\sigma_{vis2}/\sigma_{vis1}$ where the formulas both use the linear interpolation of the vdM parameters to capture the effect of the trend





Rate corrections - Out-of-time





CMS Run: 325170 LS: 51-100 (2018, 13 TeV) Before afterglow subtraction After afterglow subtraction 200

3450

3475

3500

BCID

40

30

20

10

3375

3400

3425

Out-of-time effects:

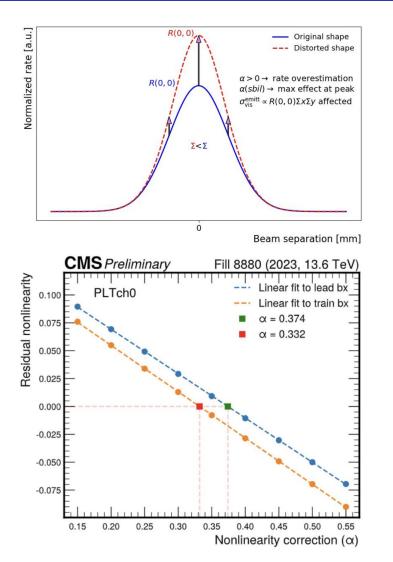
- Components:
 - > Type I: Signal spillover, Time walk
 - Type II: Material activation
- ✤ Affected: PCC, HFET, HFOC, BCM1F
- Template fit of single-bunch response functions for the two components



Rate corrections - Nonlinear response

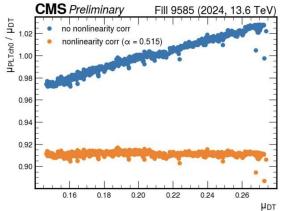
*





- * VdM calibration performed at 1/100 of the datataking pile-up
 - Model: $\mu_m = \mu_p (1 + \alpha \mu_p)$
- Mitigation correction of detectors based on intrinsic quantities: *
 - Restrictive module selection based on noise levels and \succ internal consistency (PCC)
 - Efficiency as a function of peak luminosity (SBIL) \succ tracked via emittance-scans (per-module for PLT, BCM1F)
 - Efficiency can not be used straightforwardly, as the scan curve is not uniformly distorted by the non-linearity
 - Use the Major-factor mildly profile dependent
 - Iterative- / interpolation-based procedure
- Highly linear detectors with a low occupancy are used as * reference to evaluate residual effects (DT, RAMSES)

For unc. only! \succ



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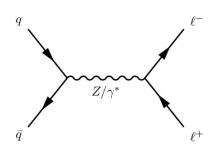
Standard candles: From cross-year consistency checks to the future of precision luminosity

Standard candle concept



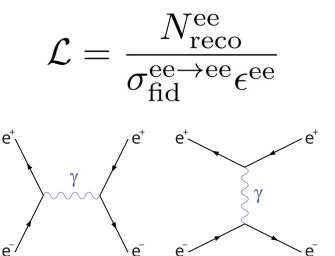
In e⁺e⁻ colliders: forward elastic (Bhabha) scattering

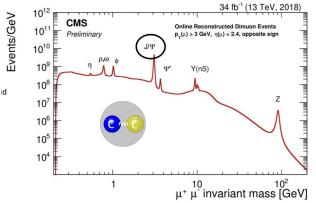
- Well known QED cross section
- Clean signature
- Only detector efficiency needs to be tracked
- LEP: 0.15% uncertainty



In pp collisions:

- $Z \rightarrow \mu \mu$ has
 - > a clean signature
 - relatively large cross section (not enough for vdM)
 - a not-too-well-known fiducial cross section
- ✔→μμ
 - > Much higher rate \rightarrow could be calibrated in vdM
 - Requires prescaled trigger in high PU
 - Low PU muons are difficult to handle
 - > Allows for transfering to $Z \rightarrow \mu \mu$ as well

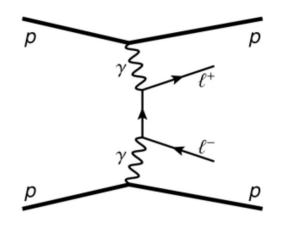


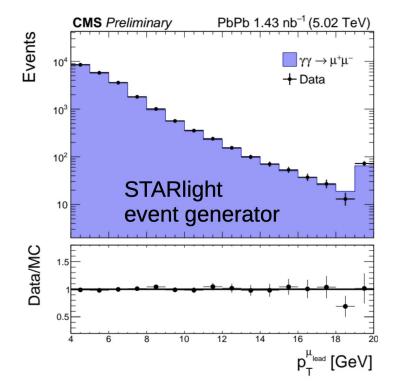


PbPb: Exclusive dimuon production



- * $\gamma\gamma \rightarrow \mu\mu$ in ultraperipherial collisions has
 - > a clean signature
 - well-known QED-based procedure BUT uncertainty from photon flux!
 - normalization to previous calibrations possible
- Publication in approval





Publications



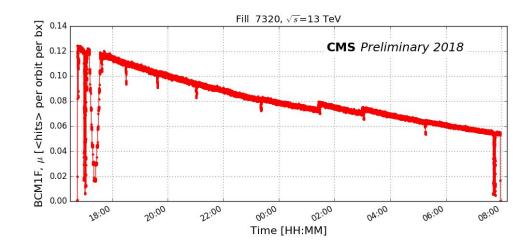
| 2015 | 1.6% | Publishe | d naper | |
|------|-------|---------------|-----------------|--|
| 2016 | 1.2% | | | |
| 2017 | 2.3% | <u>prelim</u> | Paper in | |
| 2018 | 2.5% | <u>prelim</u> | preparation | |
| 2022 | 1.4% | prelim | Paper in future | |
| 2023 | 1.28% | prelim | Paper in future | |

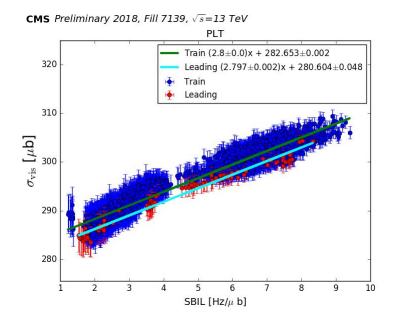
| | | | - | | |
|---|-----------------------------------|----------|--|--|--|
| Title | Date of approval | Pub | | | |
| pp@13 TeV (2015 + 2016) | 12 Nov 2020 (public: Apr 2021) | Paper | Available on the CERN CDS information server CMS PAS LUM-22-001 CMS Physics Analysis Summary | | |
| Z counting (2017) | 2 Mar 2023 (public: Sep 2023) | Paper | Contact: cms-pog-conveners-lum@cem.ch 2024/03/04 | | |
| pp@13 TeV (2017 + 2018) + Run 2 combination | In preparation | Paper | Luminosity measurement in proton-proton collisions at 13.6 TeV in 2022 at CMS | | |
| pp@13.6 TeV (2022) | 23 Feb 2024 (public: Mar 2024) | PAS | The CMS Collaboration | | |
| pp@13.6 TeV (2023) | This month | DPS note | The measurement of the integrated luminosity for the proton-proton collisions data- tablen preside at a contra-of mease energy of 13.6 TeV in 2021 with the CMS experiment at the CERS LLE: in specific and the contra-of-the measurement is collibrated from hearn-expansions swith the van der Meer scan method. The pre- cision of the collibration is limited by the knowledge of the factorization of the bunch | | |
| Run 2 (2015+2018) PbPb | In preparation | Paper | proton density during the van der Meer scans. Continuous rate measurements with various CM subdetectors provide a stable and linear luminosity measurement. Considering both calibration and integration sources, the integrated luminosity measurement has a total uncertainty of 14% . | | |

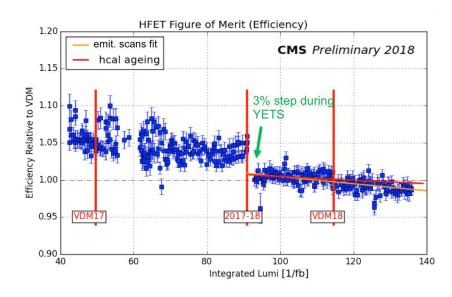
Emittance scans



- Luminometers are intrinsically corrected for all linearity affecting effects
- Emittance scans are treated like mini vdM calibrations
- Linearity and efficiency corrections



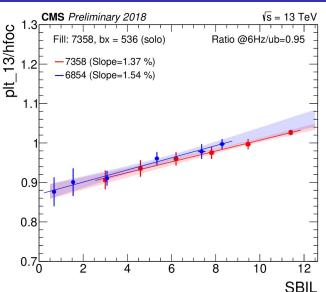


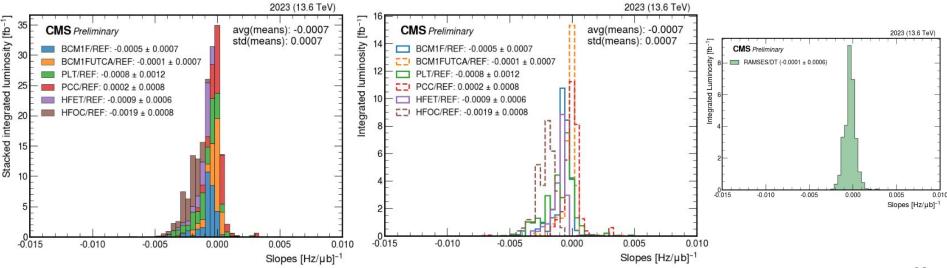


Linearity



- Luminometers are intrinsically corrected for all linearity affecting effects in situ
 - Data driven out-of-time corrections
 - Linearity from emittance scans
- Residual relative non-linearity is studied with respect to DT and RAMSES
 - Very low occupancy, highly linear detectors

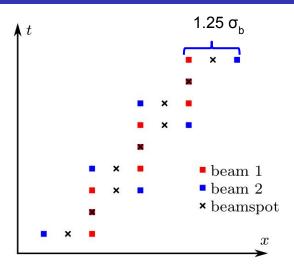


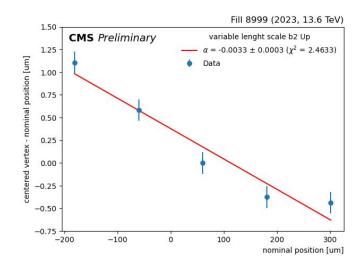


Length scale calibration (LSC)



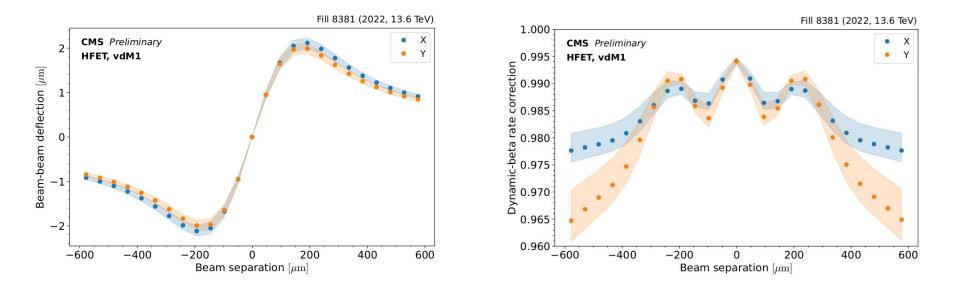
- Neither the nominal nor the BPM measured beam positions correspond to real values accurately.
- The tracker position is considered as reference
- * The relationship is linear $x_{true} = \alpha x_{nominal}$
- Two special scans used for LSC
 - Constant separation LS scan
 - Average LS for B1&B2
 - Variable separation LS scan
 - Separate LS for B1&B2





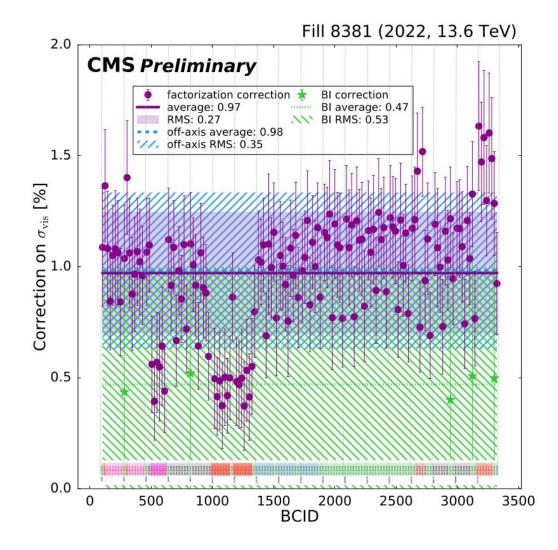
Beam-Beam effects





Non-factorisation BCID structure





Non-factorisation



- Imaging scan analysis
 - Fits the 2D bunch density function
 - Using a set of 4 special scans
 - For few BICDs with high rate VTX data

