Direct neutrino mass measurement at KATRIN

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- Motivation for KATRIN
- Measurement principle
- Results





Unveiling new physics with neutrino mass

- Neutrino flavor oscillations → neutrinos are not massless
 ⇒ necessary extension of the SM
- The most abundant known massive particle in the universe
- Neutrino mass helps to understand:
 - Formation of cosmological structures
 - Neutrino mass generation models
 - Neutrino mass hierarchy which neutrino is the lightest?







Ways to access the neutrino mass

			Ho Ho Ho Ho
	Cosmology	Search for 0vββ	β-decay & electron capture
Observable	$M_ u = \sum_i m_i$	$m_{etaeta}^2 = \left \sum_i U_{ei}^2 m_i ight ^2$	$m_{\mathcal{V}}^2 = \sum_i U_{ei} ^2 m_i^2$
Present upper limit	0.12 ⁽¹⁾ eV (0.064 ⁽²⁾ eV)	0.036 – 0.156 ⁽³⁾ eV	0.45 ⁽⁴⁾ eV
Model dependence	Multi-parameter cosmological model	 Majorana v Nuclear matrix elements 	Direct, only kinematics; no cancellations in incoherent sum

⁽¹⁾ Planck Coll., A&A 641 (2020) A6
 ⁽²⁾ DESI Coll.,arXiv:2503.14738



KATRIN: International collaboration

Karlsruhe Tritium Neutrino Experiment

- Located in Karlsruhe because of the Tritum Laboratory Karlsruhe at KIT Campus Nord
- ~ 20 institutes from Germany, US, Italy, Czechia, Thailand, Spain, France
- ~ 150 people







Tritium β -decay kinematics

High precision spectroscopy of β -electrons

- Small rates \rightarrow high activity source
- Small effect → high precision and low background needed





Measurement principle



• Excellent energy resolution: $\frac{\Delta E}{E} \sim \frac{B_{\text{ana}}}{B_{\text{max}}} \sim 10^{-4}$

Low background with ultra-high vacuum of 10⁻¹¹ mbar

Modeling



Duration of ~ 3 h per 40 voltage points

Fermi theory with radiative corrections and molecular excitations

Spectrometer transmission and energy losses in source

$$\boldsymbol{R}(U) = \boldsymbol{A}_{\boldsymbol{s}} \cdot N_T \int_{qU}^{\boldsymbol{E}_0} \frac{d\boldsymbol{\Gamma}}{d\boldsymbol{E}} (\boldsymbol{E}, \boldsymbol{m}_{\boldsymbol{v}}^2) \cdot \boldsymbol{f}(\boldsymbol{E}, \boldsymbol{U}) d\boldsymbol{E} + \boldsymbol{R}_{\boldsymbol{b}\boldsymbol{g}}$$

4 free fit parameters: $A_s, E_0, m_{\nu}^2, R_{bg}$

KATRIN data taking

Stack data points with the same measurement conditions = hundreds of β -spectrum scans = campaign



> 190 Mio. electrons

2019: m_{ν} < 1.1 eV (90% CL) **2022**: $m_{\nu} < 0.8 \text{ eV} (90\% \text{ CL})$



Key optimizations since previous release



shifted analyzing plane

Results of KNM1-5 science runs



- 7 different configurations, 59 spectra, 1609 data points, parameter correlations across datasets
- 2 independent analysis frameworks
 Kleesiek et al. EPJ C 79 (2019) 204; Karl et al., EPJ C 82 (2022) 439

$$m_{
u}^2 = -0.14^{+0.13}_{-0.15} \,\mathrm{eV^2}$$

Combined best fit value



Confidence interval



$$m_{
m v} < 0.45~{
m eV}$$
 (90% CL)

using Lokhov-Tkachov construction Lokhov & Tkachov, Phys. Part. Nucl. 46 (2015) 347

- Upper limit from Feldman-Cousins construction $m_{\nu} < 0.31 \text{ eV}$ (90% CL) benefits from negative best-fit Feldman & Cousins, Phys. Rev. D 57 (1998) 3873
- Bayesian analysis results in preparation



RESEARCH

PARTICLE PHYSICS

Direct neutrino-mass measurement based

on 259 days of KATRIN data

KATRIN Coll., Science 388 (2025) 180



Science

Physics beyond the neutrino mass



β-spectroscopy at highprecision& high stability

Search for Lorentz violation through sidereal modulation

→ PRD **107** (2023) 082005



Search for **shape distortion** through exotic weak interactions

- General neutrino interactions
- → PRL **134** (2025) 251801
- New light bosons





Physics beyond the neutrino mass



β-spectroscopy at high precision & high stability

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Search for **shape distortion** through exotic weak interactions

- General neutrino interactions
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- New light bosons

Line search for capture of local cosmic relic v → PRL 129 (2022) 011806



eV-scale sterile neutrino search



Motivated by short-baseline oscillation anomalies, direct comparison with $\Delta m_{41}^2 \approx m_4^2 - m_\beta^2$

$$\sin^2(2\theta_{\rm ee}) = 4|U_{\rm e4}|^2 \left(1 - |U_{\rm e4}|^2\right)$$

 Δm^2_{41} (eV²)

10¹

 10^{-1}

Challenging the Neutrino-4 result

KATRIN bound dominates for $\Delta m_{41}^2 > 5 \text{ eV}^2$







Exotic weak interactions from spectrum shape distortion

soon

General neutrino interactions



- Effective field theory of weak interactions
- New interactions considered (S, P, V, AV, T couplings)
- Modification of the effective decay rate: $\frac{d\Gamma_{\rm GNI}}{dE} = \frac{d\Gamma_{\rm SM}}{dE} \sum_{\nu = R_N} \xi_k \left[1 \frac{b'_k}{E_0 E} \right]$
- Competitive coupling limits to the high-energy investigations

Light boson searches

- Consider additional new light bosons coupling to leptons
- Contribution to the decay rate: $\frac{d\Gamma}{dF} = \frac{d\Gamma_{SM}}{dF} + \frac{d\Gamma_X}{dF}$
- New parameters: boson mass m_X and coupling g_X
- Probe new physics at low energy scale of < 20 keV
- Complementary search for new light bosons







Summary

Current world-best direct neutrino mass limit by KATRIN

 $m_{
m v} < 0.45~{
m eV}$ (90% CL)

- Based on first 5 science runs
- Data taking continues until the end of 2025
- Final sensitivity of < 0.3 eV
- Rich physics program beyond the neutrino mass
- Future plans:

Julv 7. 2025

16

- Detector upgrade for keV sterile neutrino search: TRISTAN (2026-27) Mertens et al., J. Phys. G 46 (2019) 065203
- **next-generation** ν -mass experiment (R&D phase)
 - Differential detection and atomic tritium technologies











TRISTAN project

Aim: Measure the full tritium spectrum to search for keV steriles at ~10⁻⁶ admixture

Challenges: Energy range x 400, electron rate x 10⁶, high energy resolution, control of systematics

Technology: Silicon Drift Detector (SDD) array with > 1000 pixels





- \checkmark Capacity to handle high rates (> 10⁸ cps)
- \checkmark Excellent energy resolution (160 eV @ 5.9 keV)

replica beamline for system characterization

Timeline: Implement in KATRIN beamline for measurements during 2026-2027

Mertens *et al.*, J. Phys. G46 (2019) & G48 (2021); Gugiatti *et al.*, NIM A979 (2020) & NIM A1025 (2022), Biassoni *et al.*, EPJ Plus 136 (2021) 125, Siegmann *et al.*, J. Phys. G51 (2024) 8, 085202, Spreng *et al.*, JINST 19 (2024) P12009



KATRIN as R&D facility for differential detector



Goal: Sub-eV resolution for **differential** energy measurement. Added benefit: Immune to MAC-E filter backgrounds!

Several promising technology options



Demonstrated to improve MAC-E resolution by factor ~10



Sensitivity beyond KATRIN



KATRIN now: Integral, $\Delta E = 2.7 \text{ eV}$, bg = 0.1 cps More tritium would help, but ultimately need:

- Differential* measurement (FWHM < 1 eV) *) meaning: non-integrating spectroscopy
 - ✓ Better use of statistics
 - \checkmark Intrinsically low background
- Atomic tritium
 - ✓ Avoid broadening (~ 1 eV)
 - ✓ Avoid limiting systematics of T₂





Backup



Cosmological observables

Neutrinoless double β-decay

