

A limit on neutrino mass with the first **H_VLMES** dataset

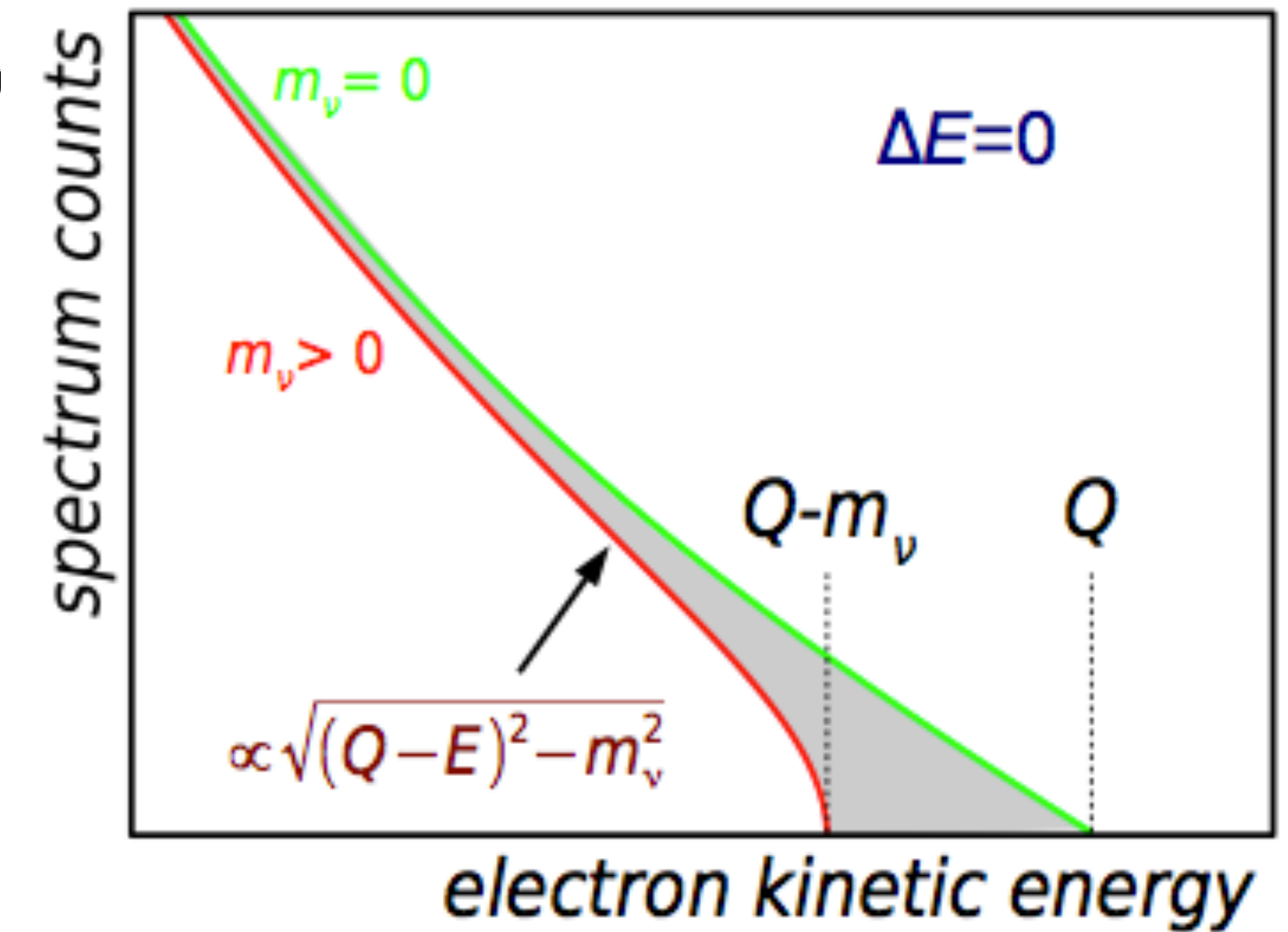


Outline

- Direct neutrino mass measurements
- ^{163}Ho and neutrino mass: the HOLMES experiment
- First results from HOLMES
- Next steps to improve m_ν sensitivity

Direct neutrino mass measurements

- **Low-Q β /EC decay isotopes** (^3H , ^{187}Re , ^{163}Ho) needed;
- **model independent**: relies only on E, p conservation
- ν mass appears as a **distortion at the end point of the electron energy spectrum**;
- 2 different approaches.



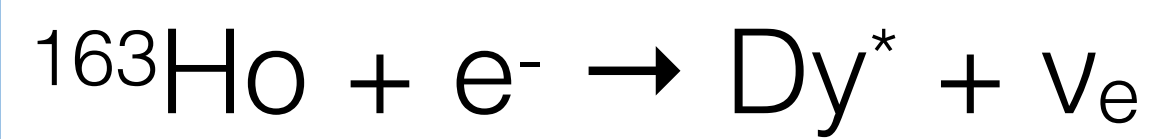
Spectrometry:

- source $\neq$ detector (**KATRIN** like).
- High statistics allowed, **no pile-up issue**.
- Main systematics sources:
 - **decay on excited states**;
 - **energy losses in source**.
- **Currently, gives the best limit: $m_\nu < 0.45$ eV**
- **Sensitivity down to 0.3 eV**

Calorimetry:

- source $\subset$ detector (**HOLMES** approach)
 - Remove systematics related to energy loss in the source, **but implies pile-up issue!**
 - Need a trade off between activity and time reso + pile-up rejection algorithm.
- **Could explore region below 0.3 eV**
- **Complementary to spectrometry**

^{163}Ho electron capture and ν mass



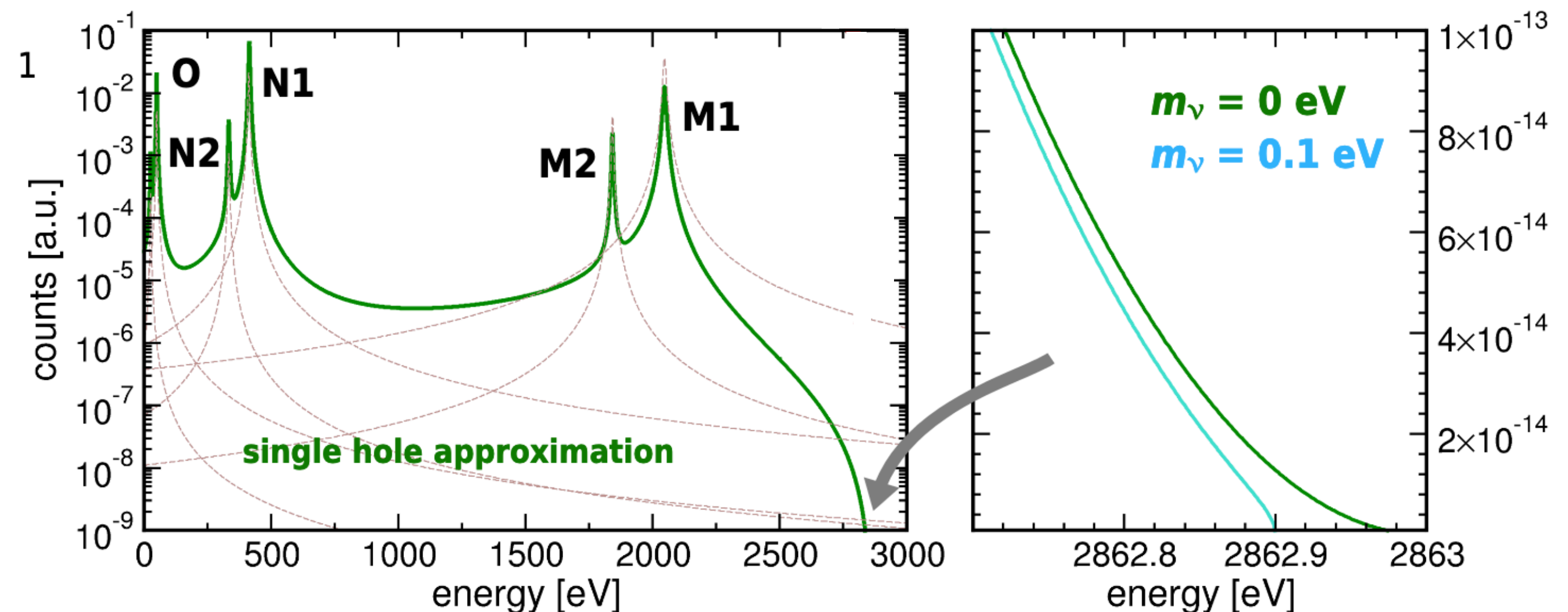
$$\frac{d\lambda_{EC}}{dE_c} = \frac{G_\beta^2}{4\pi^2} (Q - E_c) \sqrt{(Q - E_c)^2 - m_\nu^2} \times \sum n_i C_i \beta_i^2 B_i \frac{\Gamma_i}{2\pi} \frac{1}{(E_c - E_i)^2 + \Gamma_i^2/4}$$

Q~2.8 keV, capture only from shell $\geq$ M1
De Rujula & Lusignoli, Phys. Lett. B 118 (1982) 429

same phase space factor as β decay
(total de-excitation energy E_c instead of E_e)

Breit-Wigner shape

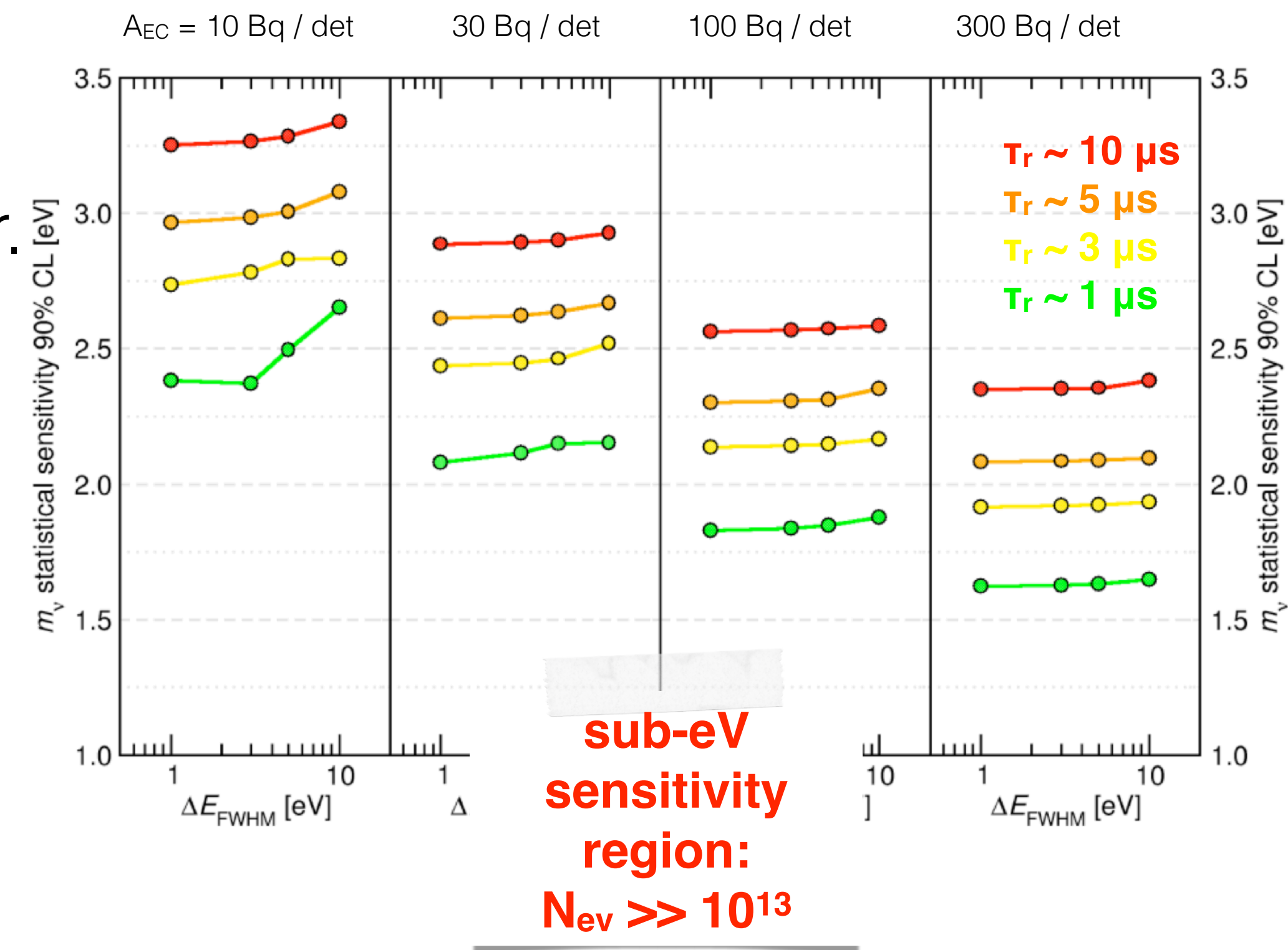
- Calorimetric measurement of Dy^* de-excitation spectrum;
- m_ν sensitivity depends on Q-value
- M1 line close to end-point could enhance “good” statistics?
- $\tau_{1/2} \sim 4570$ years $\rightarrow$ high specific activity, avoid to spoil detector performances by adding too much ^{163}Ho .



The HOLMES experiment

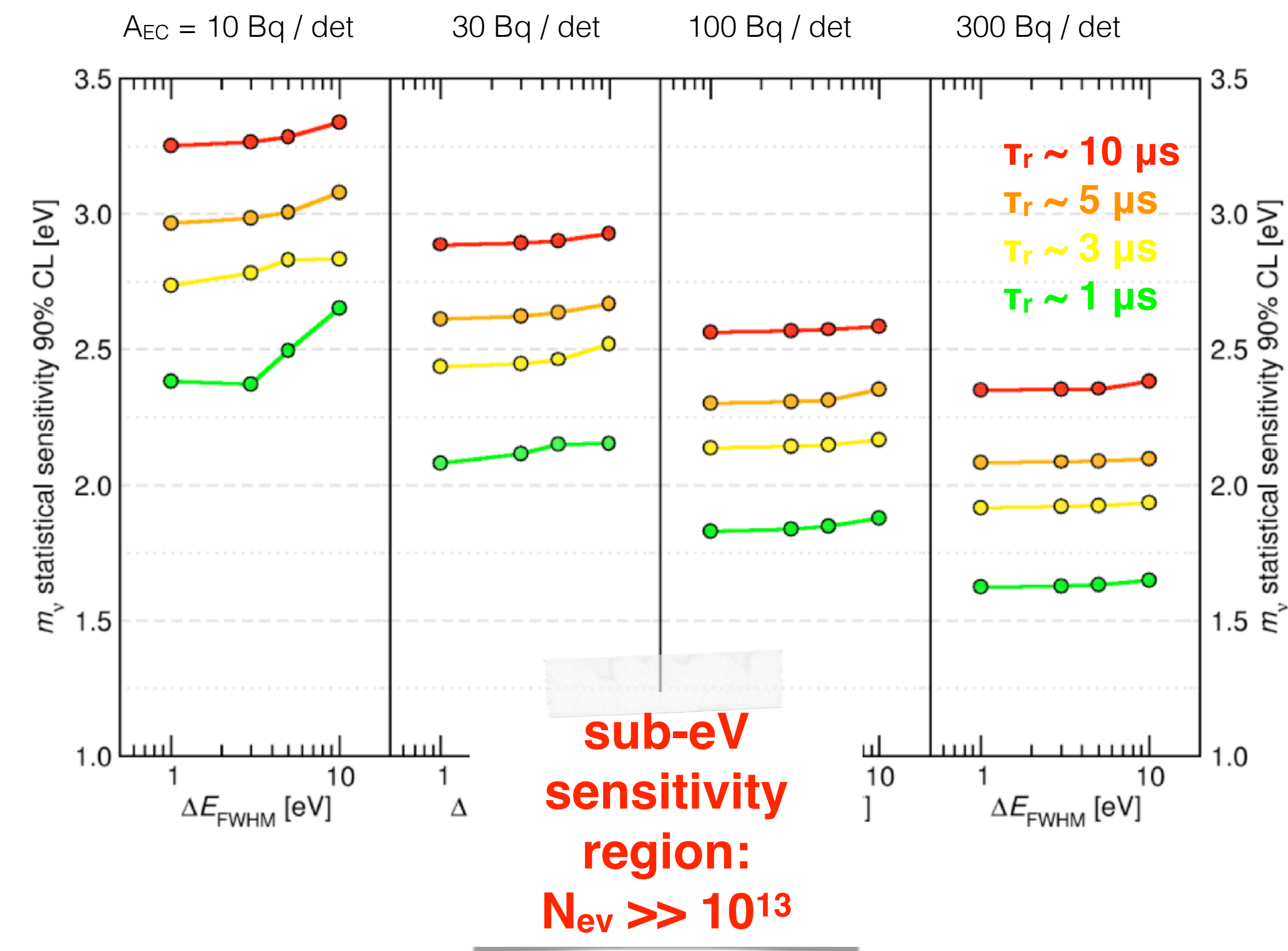
- Started as ERC Advanced Grant (2013)
- **Direct neutrino mass measurement** with statistical sensitivity around 1 eV
- Based on **Transition Edge Sensor (TES) micro-calorimeters** arrays with ^{163}Ho implanted Au absorber.
- Multiplexed readout
- Proposal:
 - 1k channels, $A_{\text{EC}} \sim 300 \text{ Bq / det}$
 - $\Delta E \sim \text{few eV} \rightarrow$ weak dependance from energy reso
 - $\Delta t \sim 1 - 10 \mu\text{s} \rightarrow$ to get rid of pile up
 - Needed at least $O(10^{13})$ events to explore eV region

Multi-step approach to prove the feasibility of the project...



The HOLMES experiment

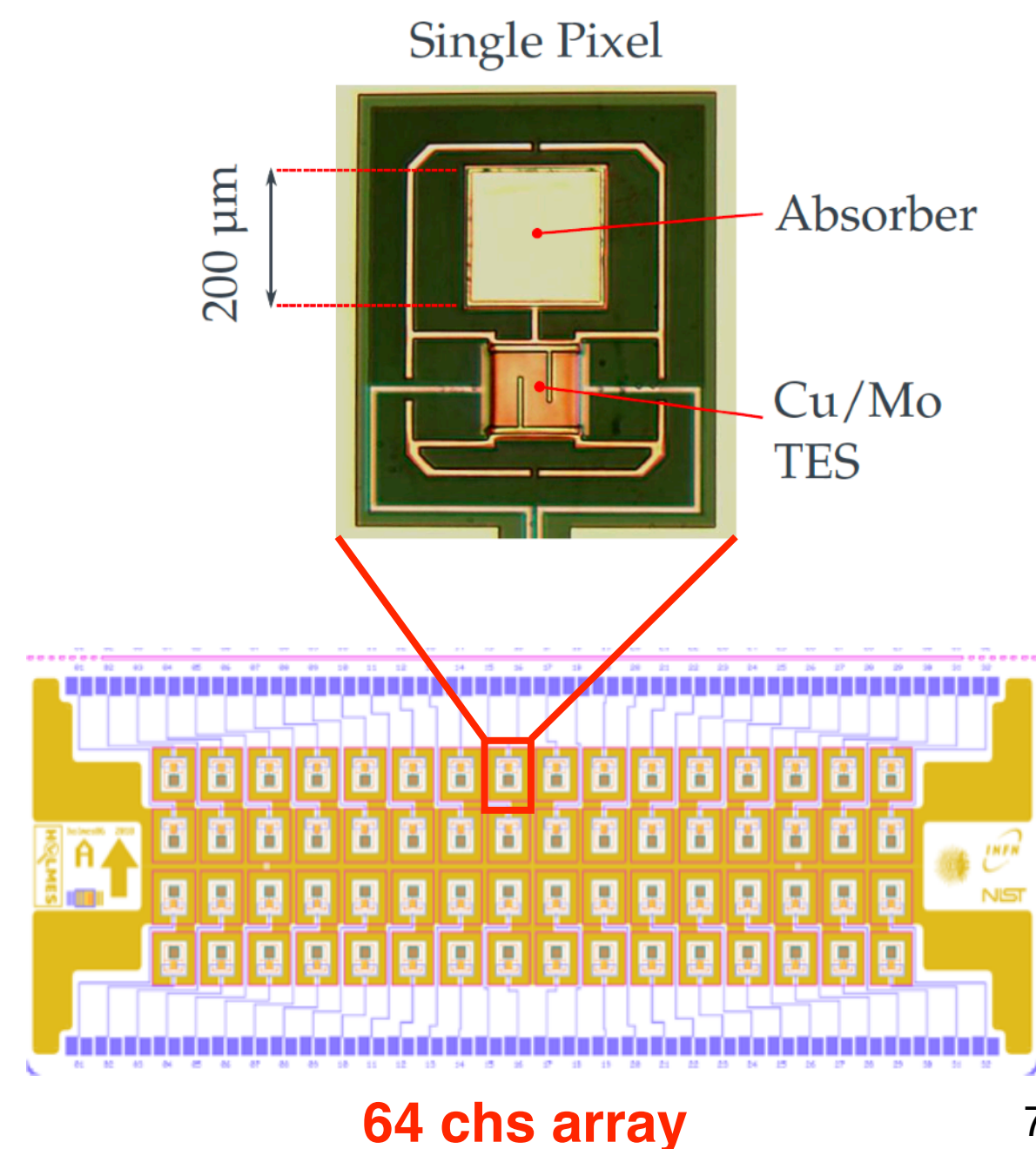
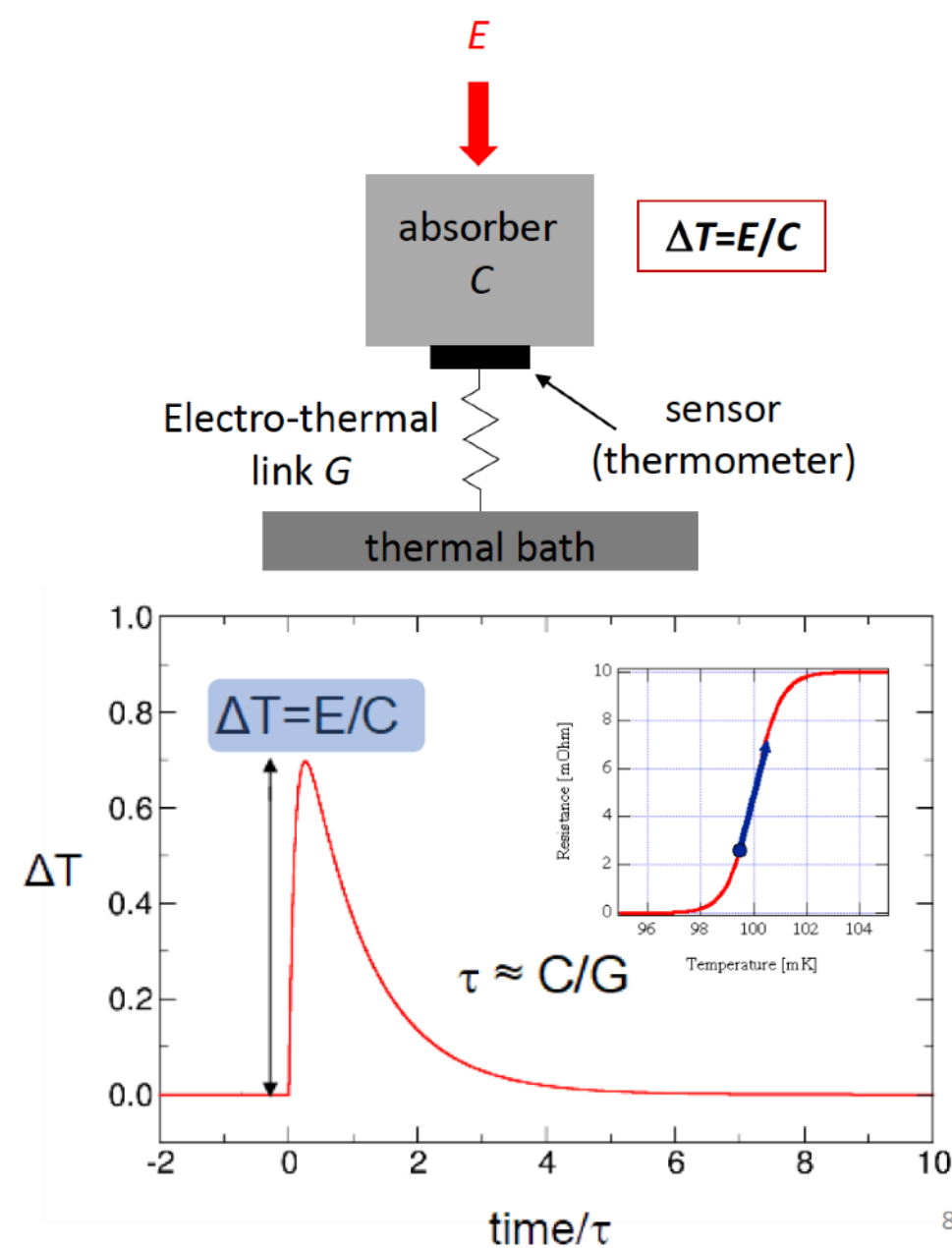
- **Multi steps approach:**
 - **64 chs prototype, ~ 1 Bq/ch (this talk)**
 - Probe of the full chain, assessment of implanted Ho effect on detector signals, first high statistic spectrum (analysis tool, spectrum evaluation...)
 - **256 chs, implanted with maximum achievable activity with current setup (1 - 3 Bq)**
 - $t_m = 1$ month, m_ν extraction with sensitivity $O(10 \text{ eV})$
 - **1k chs arrays with high activity $O(100 \text{ Bq/ch})$:**
 - 6.5×10^{16} total nuclei
 - $O(10^{13})$ events / year
 - $m_\nu \sim 1 \text{ eV}$



TES-based μ -calorimeters

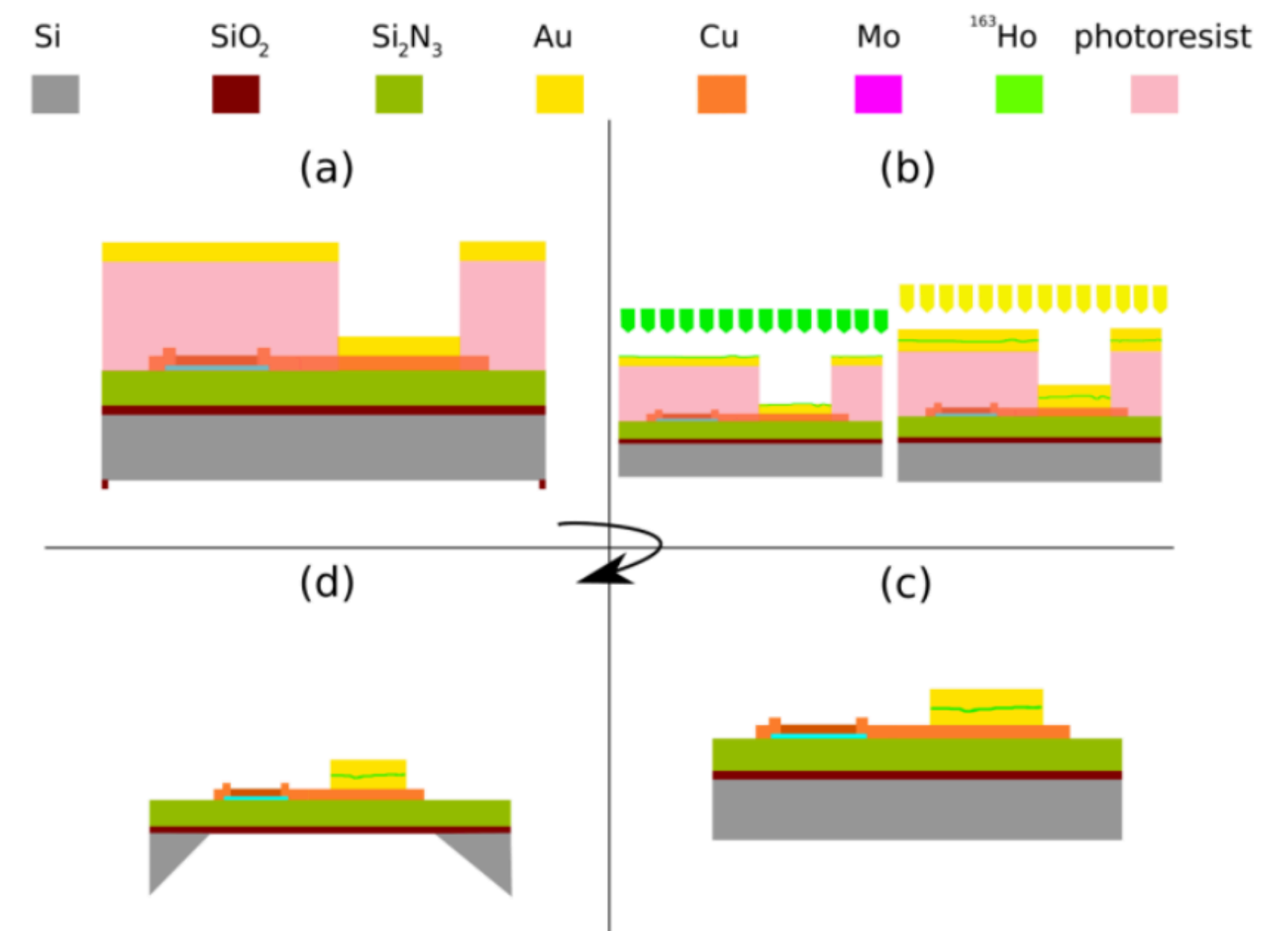
Transition edge sensors based μ -calorimeters:

- Gold absorber coupled to superconductive Cu/Mo sensor in transition region,
- energy release in absorber $\rightarrow$ temperature increase in TES $\rightarrow$ variation of TES resistance;
- 2 μm Au thickness for full energy absorption
- Operating point at 100 mK



Multi step production procedure:

- Up to first 1 μm of Au at NIST;
- Ho implantation in Genova;
- absorber completion, bonding and membrane release in Milano.



^{163}Ho embedding

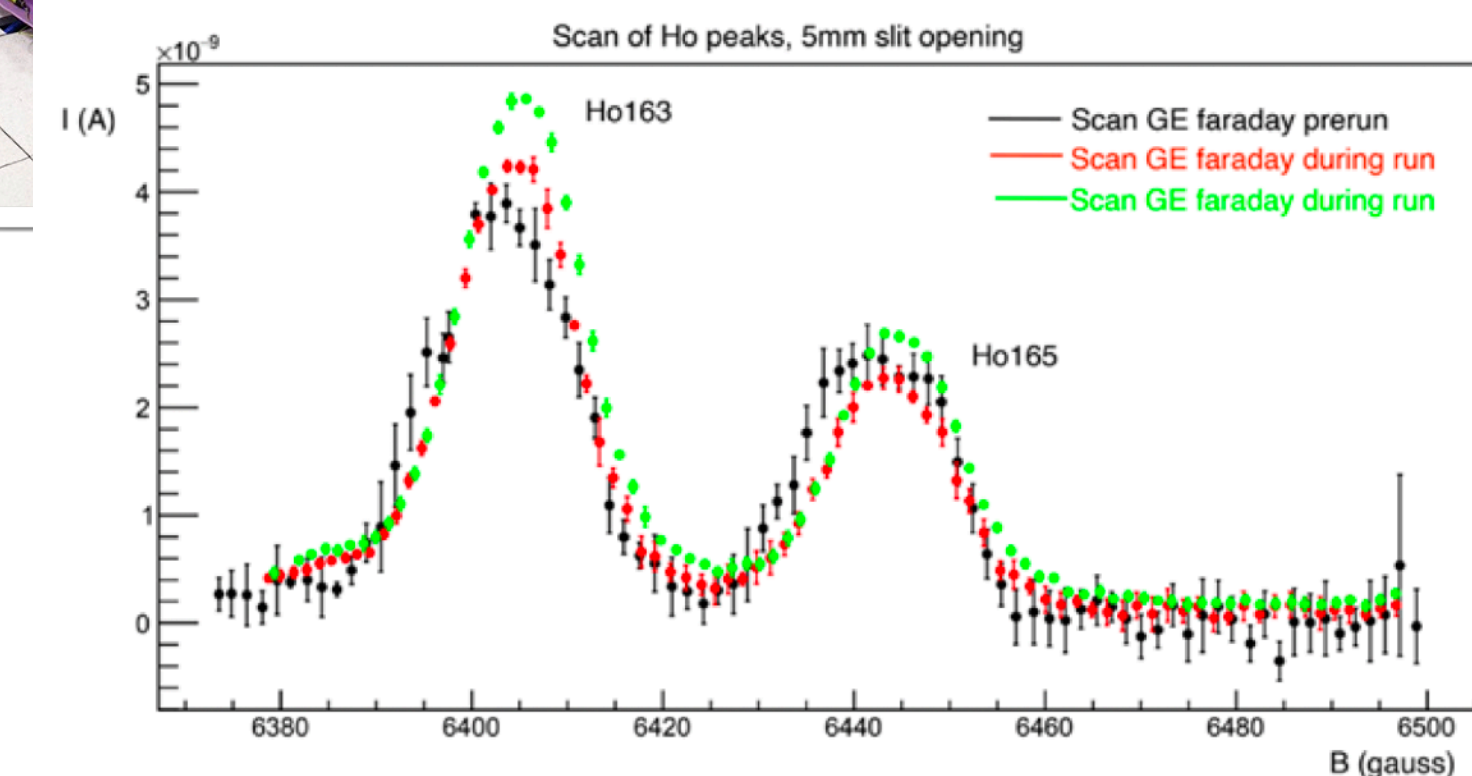
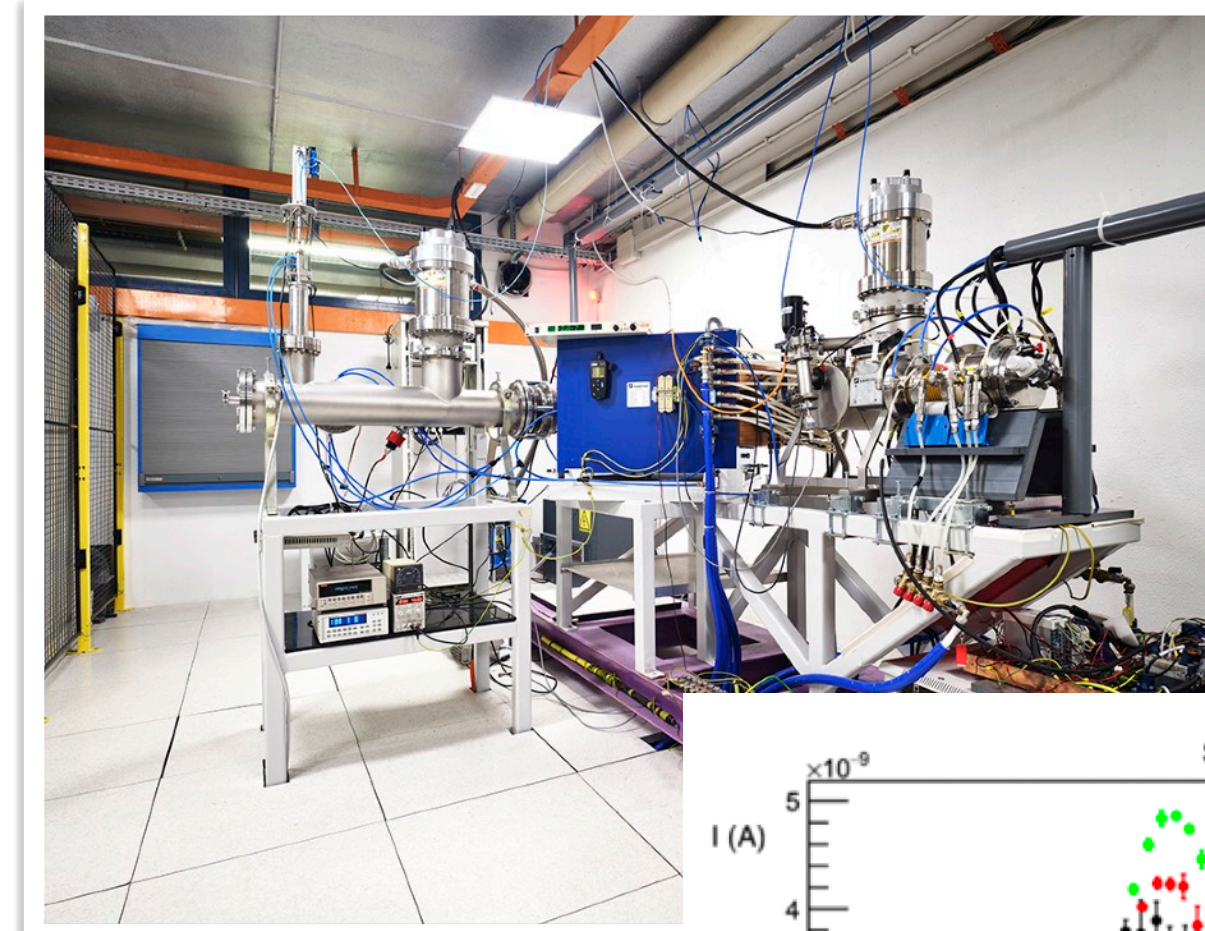
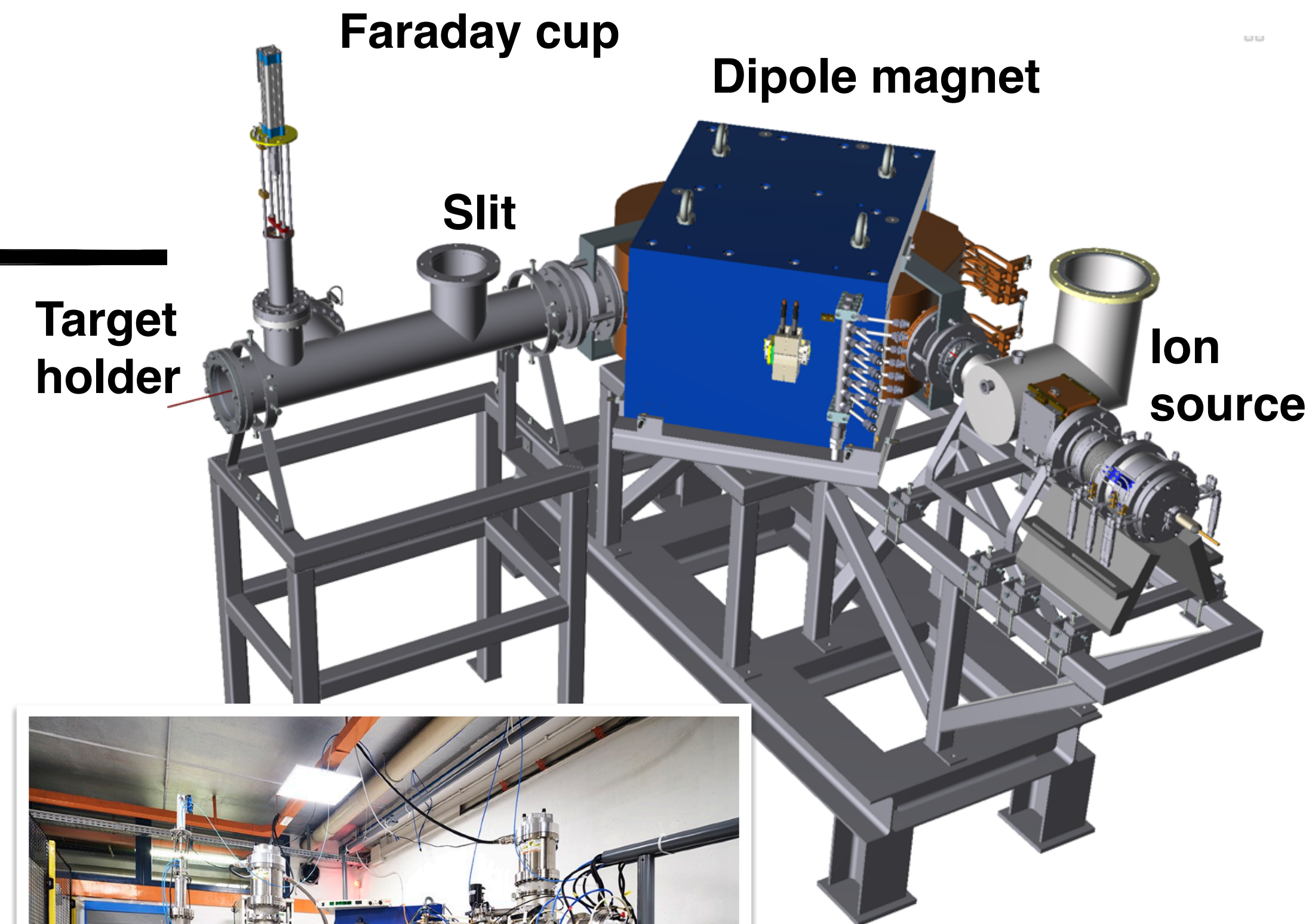
A **chemical purification** before and after the activation is needed to remove contaminants other than Ho.

An **ion implanter** has been designed and commissioned to embed ^{163}Ho in the μ -calorimeters and **to remove $^{166\text{m}}\text{Ho}$ residuals.**

With a quite simple setup (50 keV penning sputter ion source + dipole + slit and FC) we obtained:

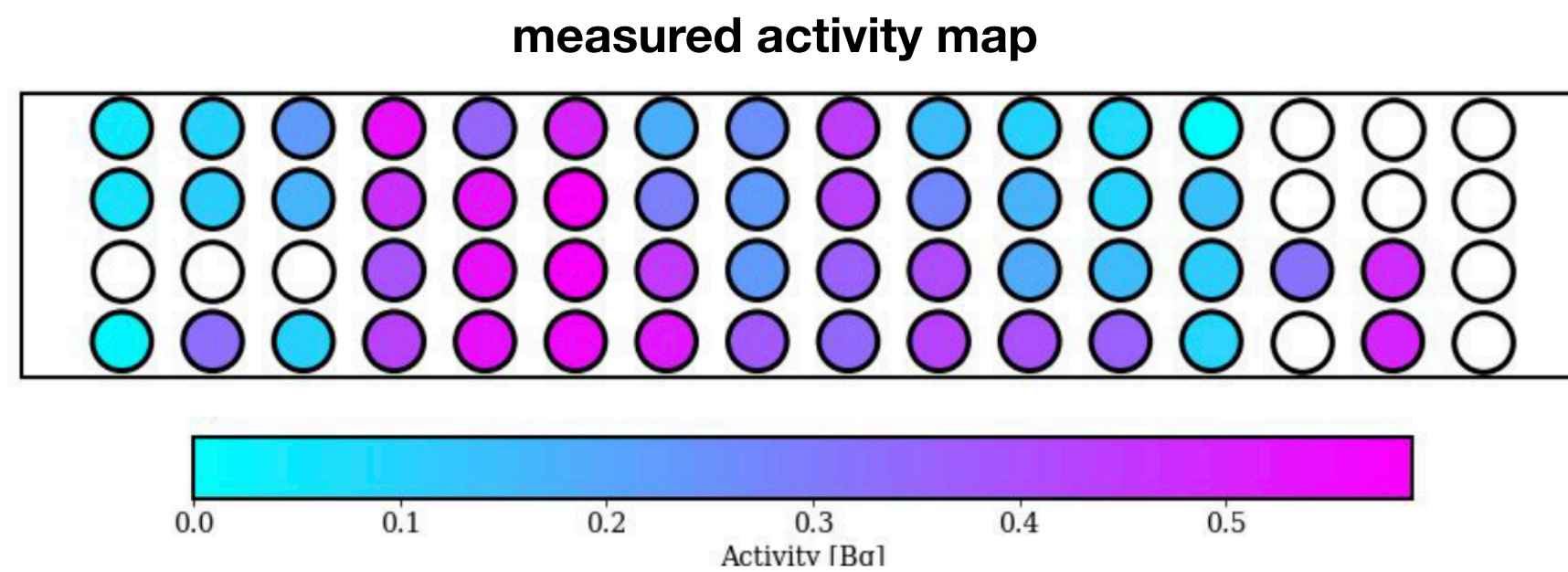
- beam size $\sigma \sim 1.5$ mm;
- 163 / 166 mass separation $> 5 \sigma$;
- efficiency $\sim 0.2\%$.

Even if efficiency is very low, it was still enough to implant the first array for physics DAQ!

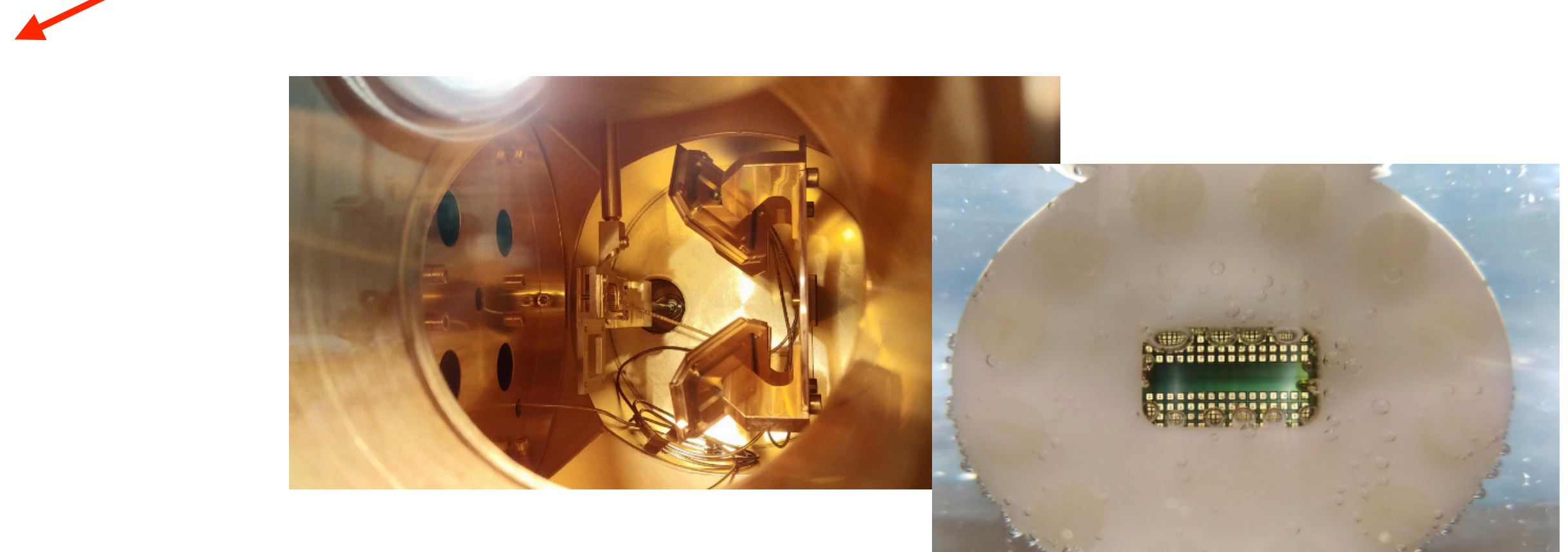
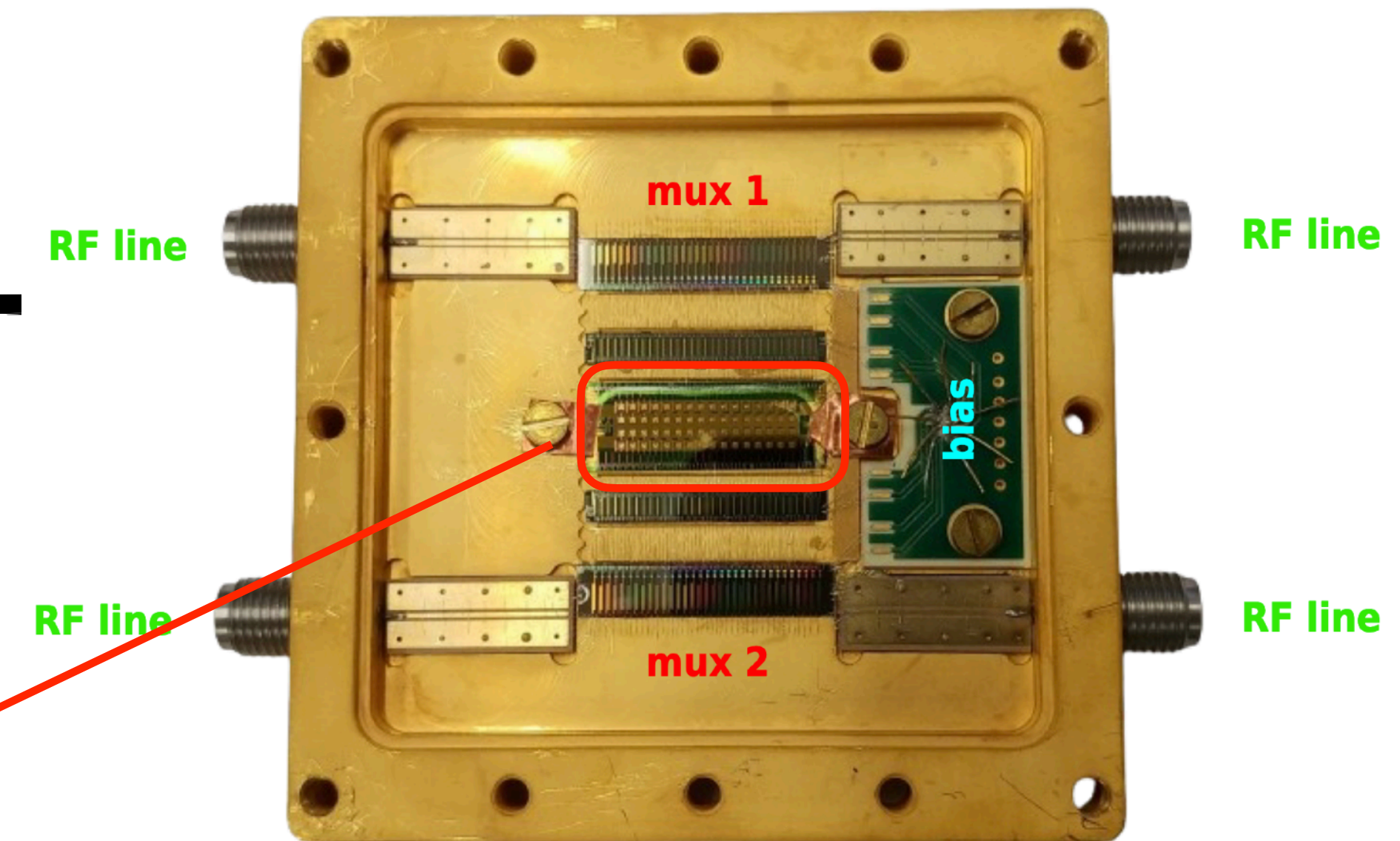


Arrays for physics data taking

After a long commissioning phase, we finally implanted our first arrays for physics data taking aiming for uniform $\sim 1\text{Bq}$ / detector activity...



At the end of the implantation run we got:
48 active detectors (activity above threshold);
average activity $\langle A \rangle = 0.27\text{ Bq}$;
total activity $A_{\text{tot}} = 13\text{ Bq}$;
peak activity $A_{\text{max}} \sim 0.6\text{ Bq}$

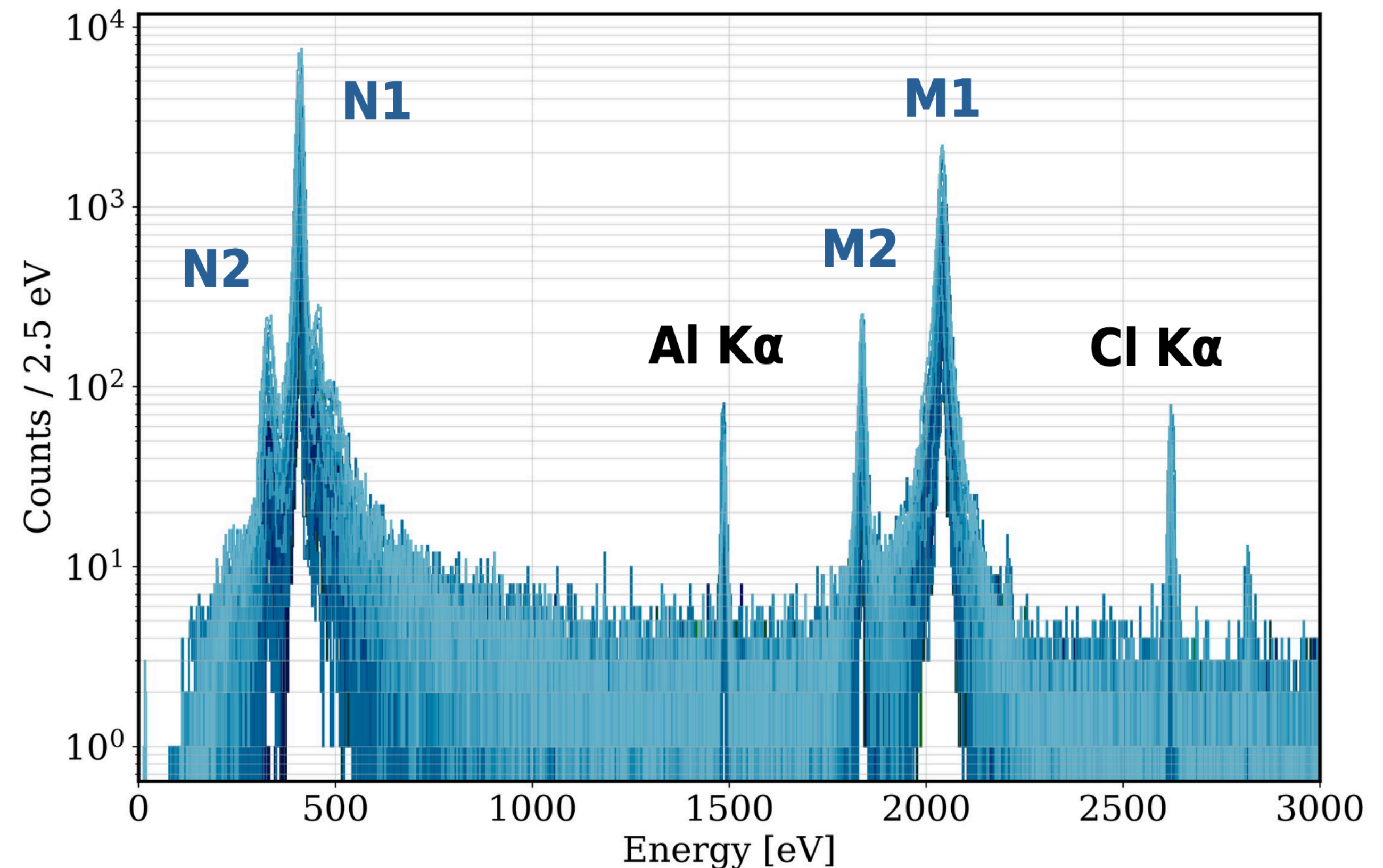


After implantation, the array is completed by adding a $1\text{ }\mu\text{m}$ Au layer to fully encapsulate the source, then membrane is released by means of KOH wet etching.

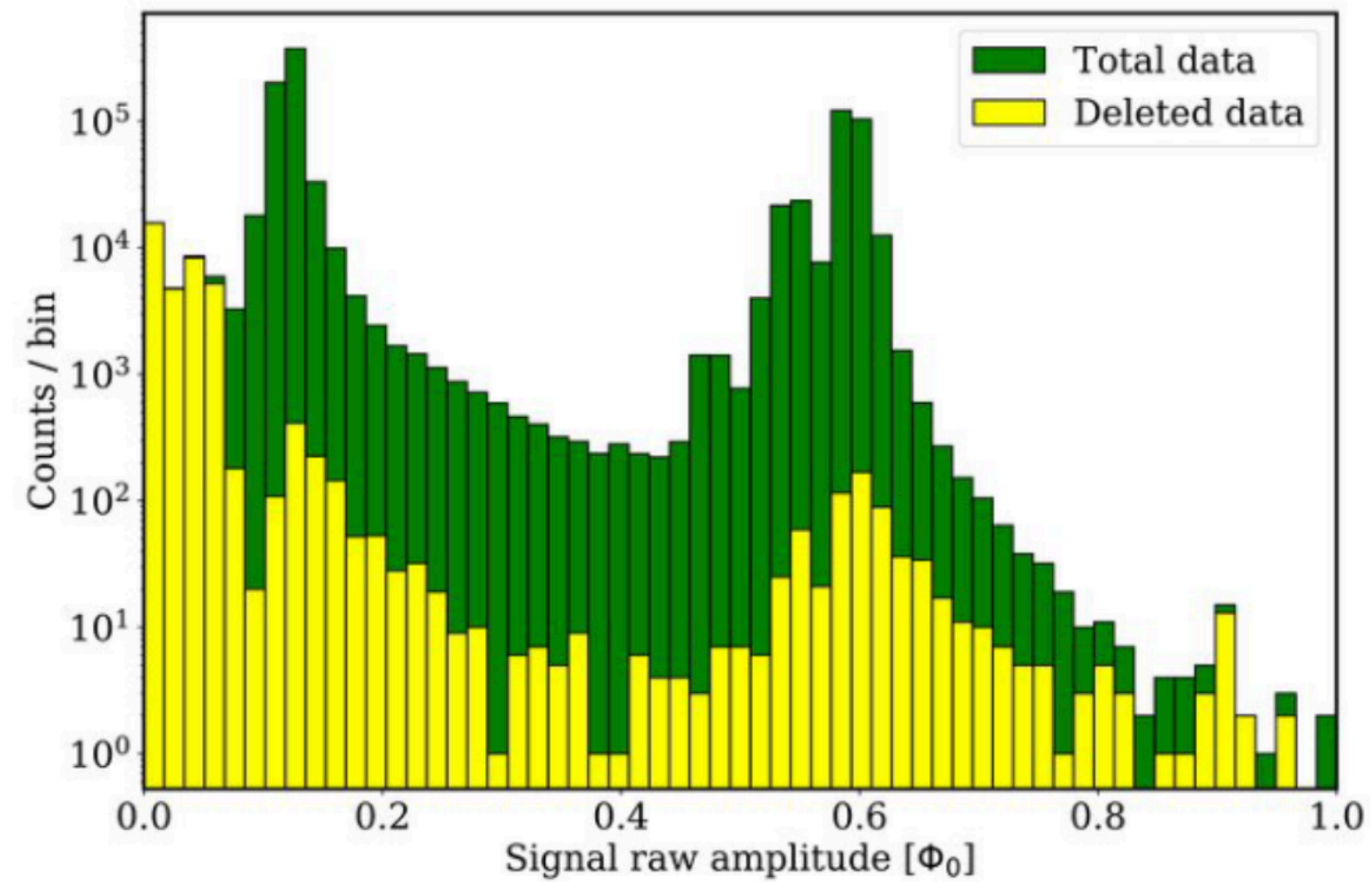
EC peak and detectors characterization

- Run with fluorescence X-ray source
- 48 active μ -calorimeters
 - $\Delta E_{FWHM} = 5.4 \sim 8.0$ eV
- 2nd order polynomial calibration
 - $E(A) = a_1 A + a_2 A^2$
- Find EC peak energies: calibration for physics run, needed to merge spectra from different μ -calorimeters

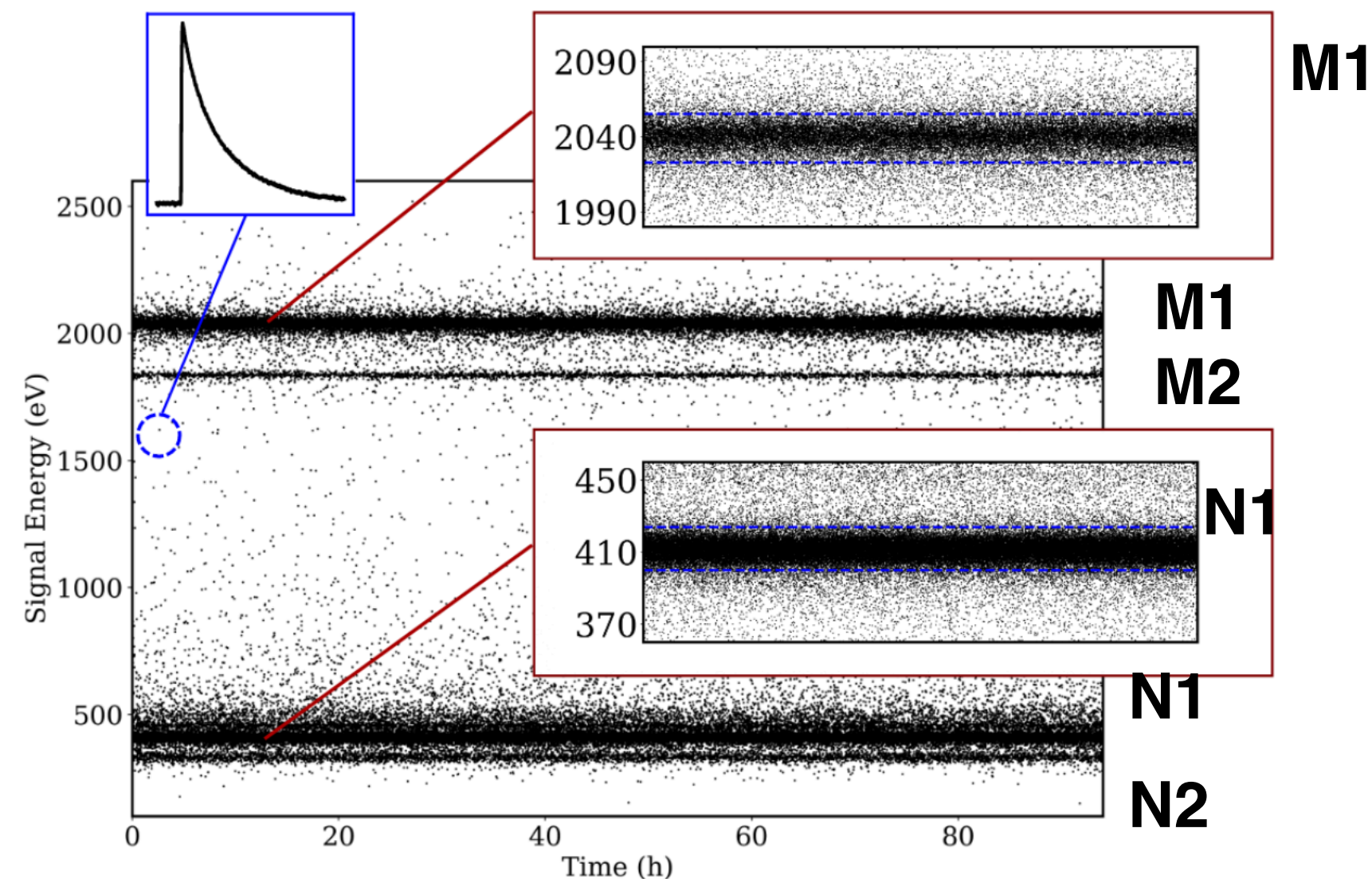
Peak	Position [eV]	Gamma [eV]	Asymmetry
M1	2040.8 ± 0.3	14.49 ± 0.05	1.306 ± 0.006
M2	1836.4 ± 0.8	8.2 ± 0.3	1.03 ± 0.05
N?	454.5 ± 0.1	22.3 ± 0.4	0.62 ± 0.02
N1	411.72 ± 0.1	5.57 ± 0.03	1.270 ± 0.008
N2	329.0 ± 0.1	16.4 ± 0.2	0.69 ± 0.01



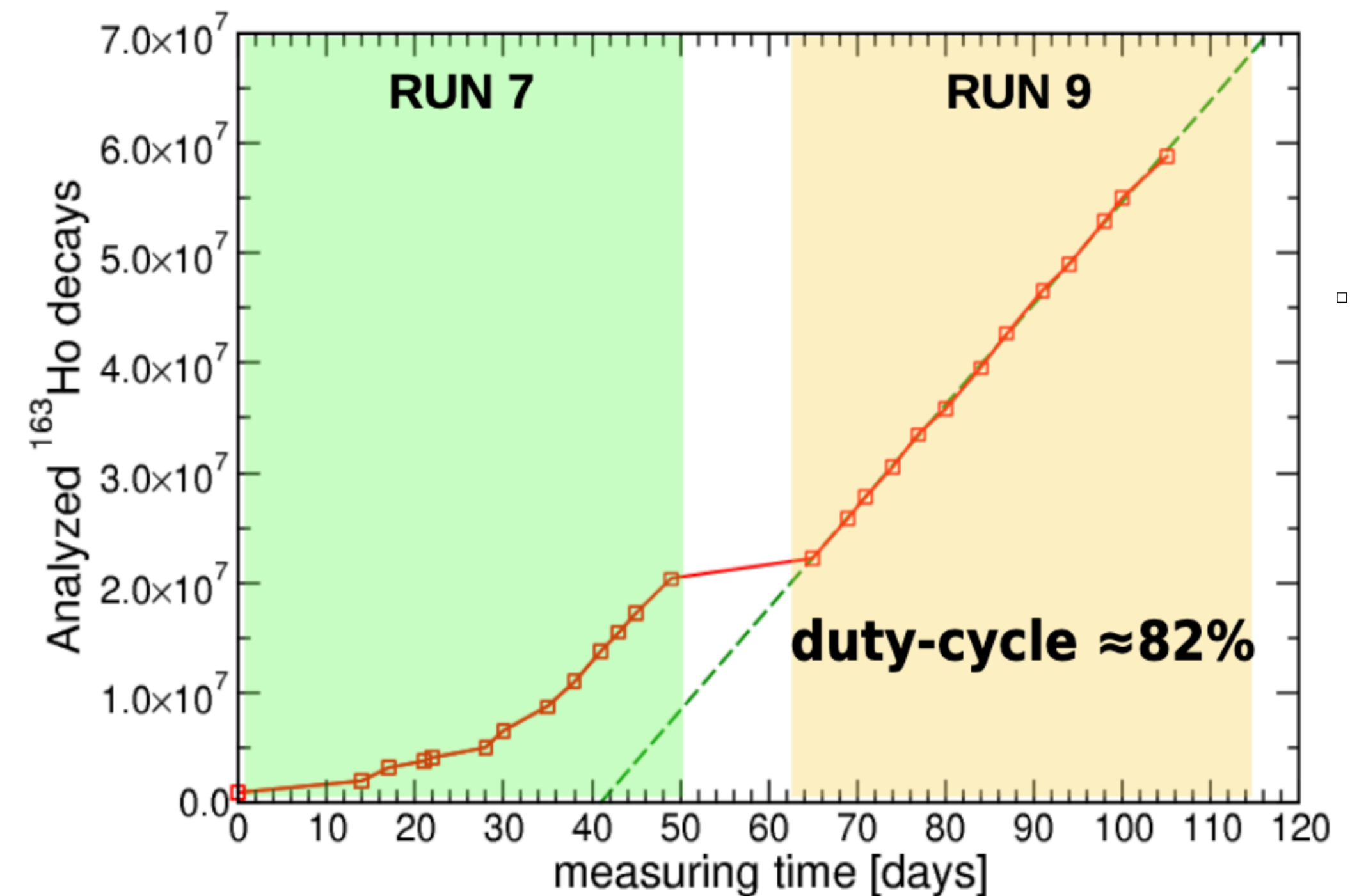
First ^{163}Ho high statistics spectrum



- **Physics data taking in 2024, about 2 months live time**
- $< 1\%$ signals discarded by first level analysis (based on pulse shape)

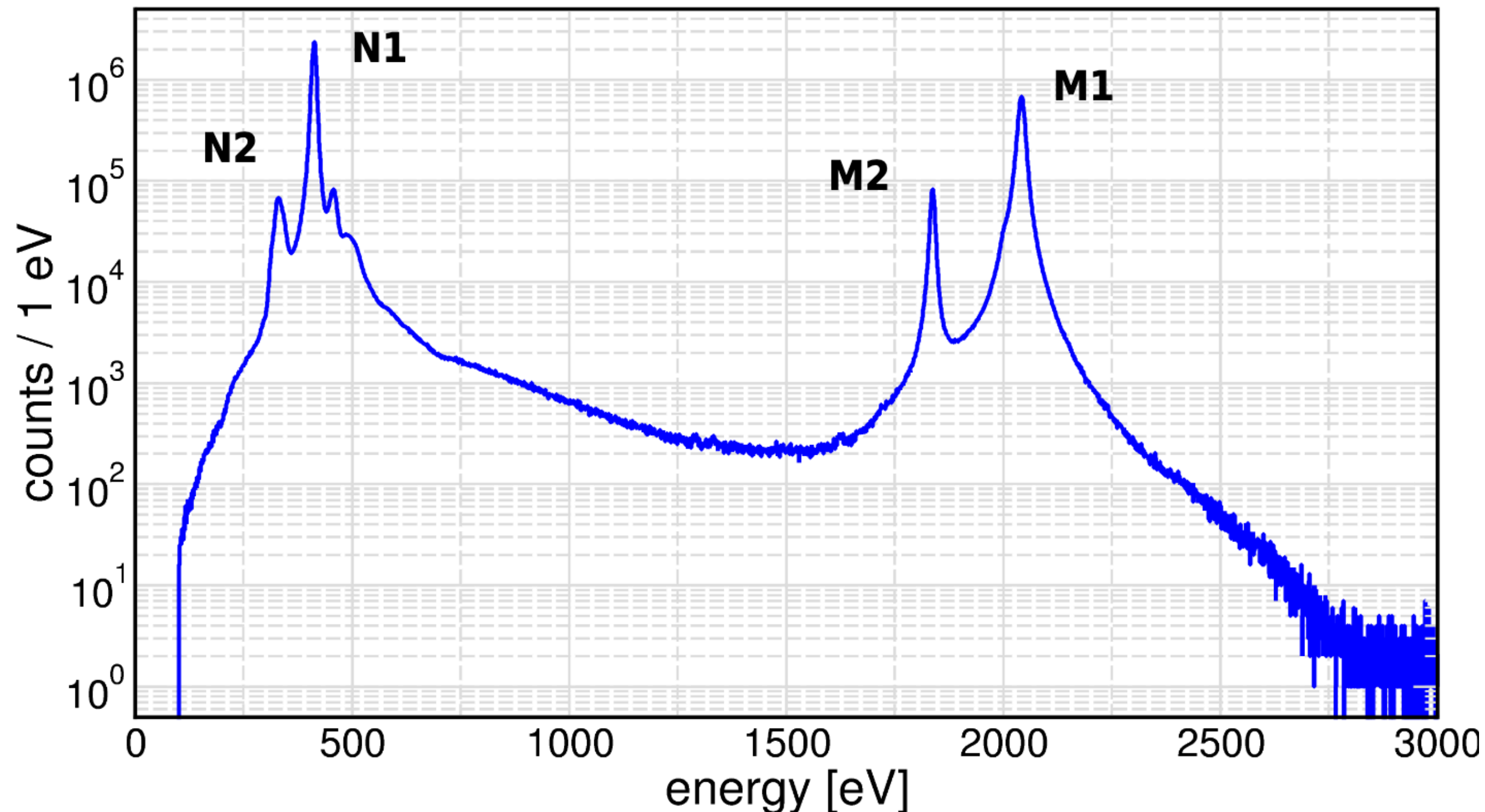


Gain drift correction



First ^{163}Ho high statistics spectrum

- 48 selected pixels
- Average activity $\langle A \rangle = 0.27 \text{ Bq} \rightarrow$ total activity $A_{\text{tot}} = 13 \text{ Bq}$ (3.2×10^{12} ^{163}Ho nuclei)



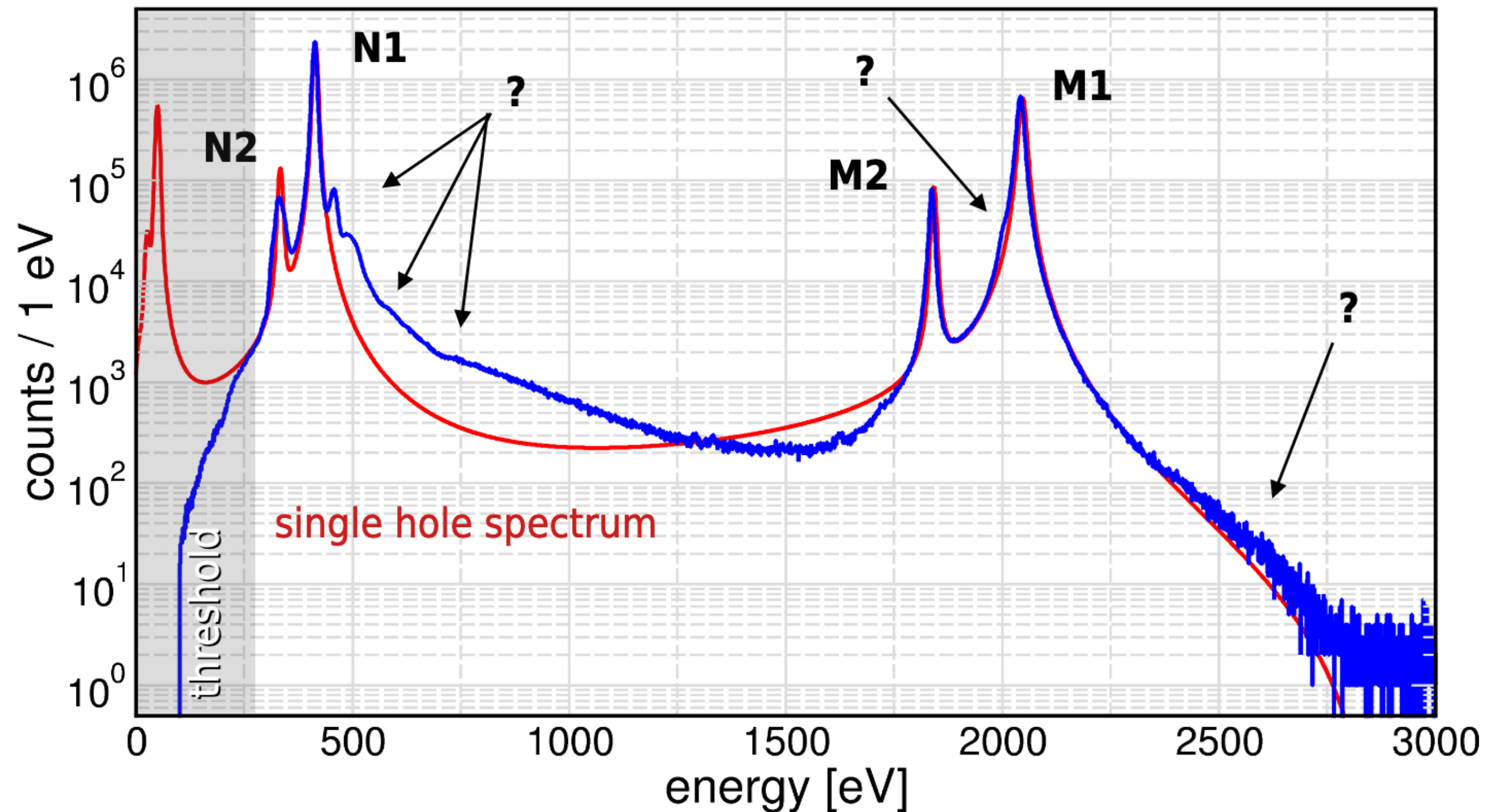
- $\Delta E_{\text{FWHM}} \sim 7 \text{ eV}$
- 7×10^4 detector x hour
- $\sim 7 \times 10^7$ ^{163}Ho decays

Sum of 1000 partial data sets
(48 det x 25 runs) with:

- Gain drift correction
- Energy calibrated using N1, M1 and M2 lines

First ^{163}Ho high statistics spectrum

- 48 selected pixels
- Average activity $\langle A \rangle = 0.27 \text{ Bq} \rightarrow$ total activity $A_{\text{tot}} = 13 \text{ Bq}$ (3.2×10^{12} ^{163}Ho nuclei)

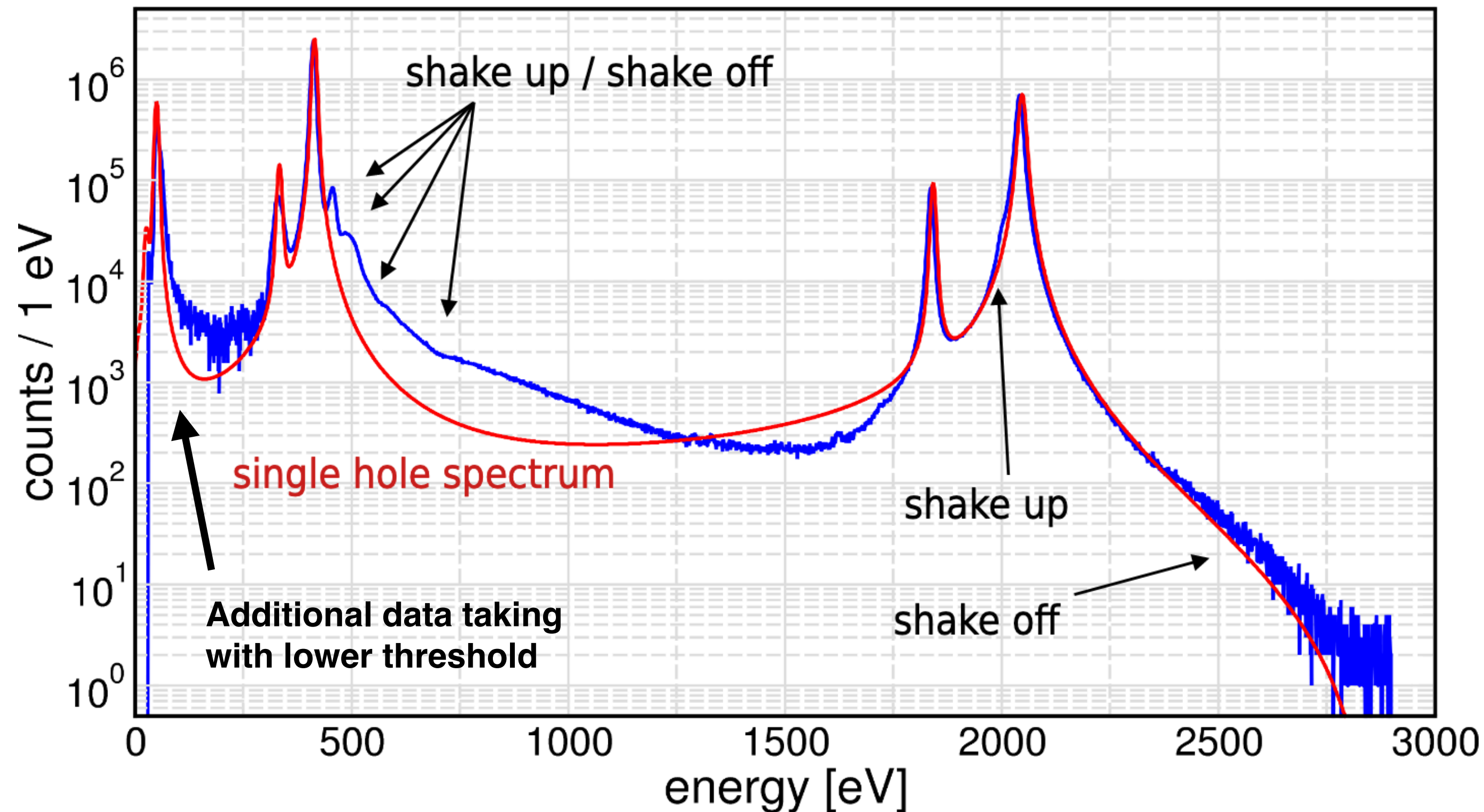


- $\Delta E_{\text{FWHM}} \sim 7 \text{ eV}$
- 7×10^4 detector x hour
- $\sim 7 \times 10^7$ ^{163}Ho decays

Single hole spectrum doesn't explain data...

First ^{163}Ho high statistics spectrum

- 48 selected pixels
- Average activity $\langle A \rangle = 0.27 \text{ Bq} \rightarrow$ total activity $A_{\text{tot}} = 13 \text{ Bq}$ (3.2×10^{12} ^{163}Ho nuclei)

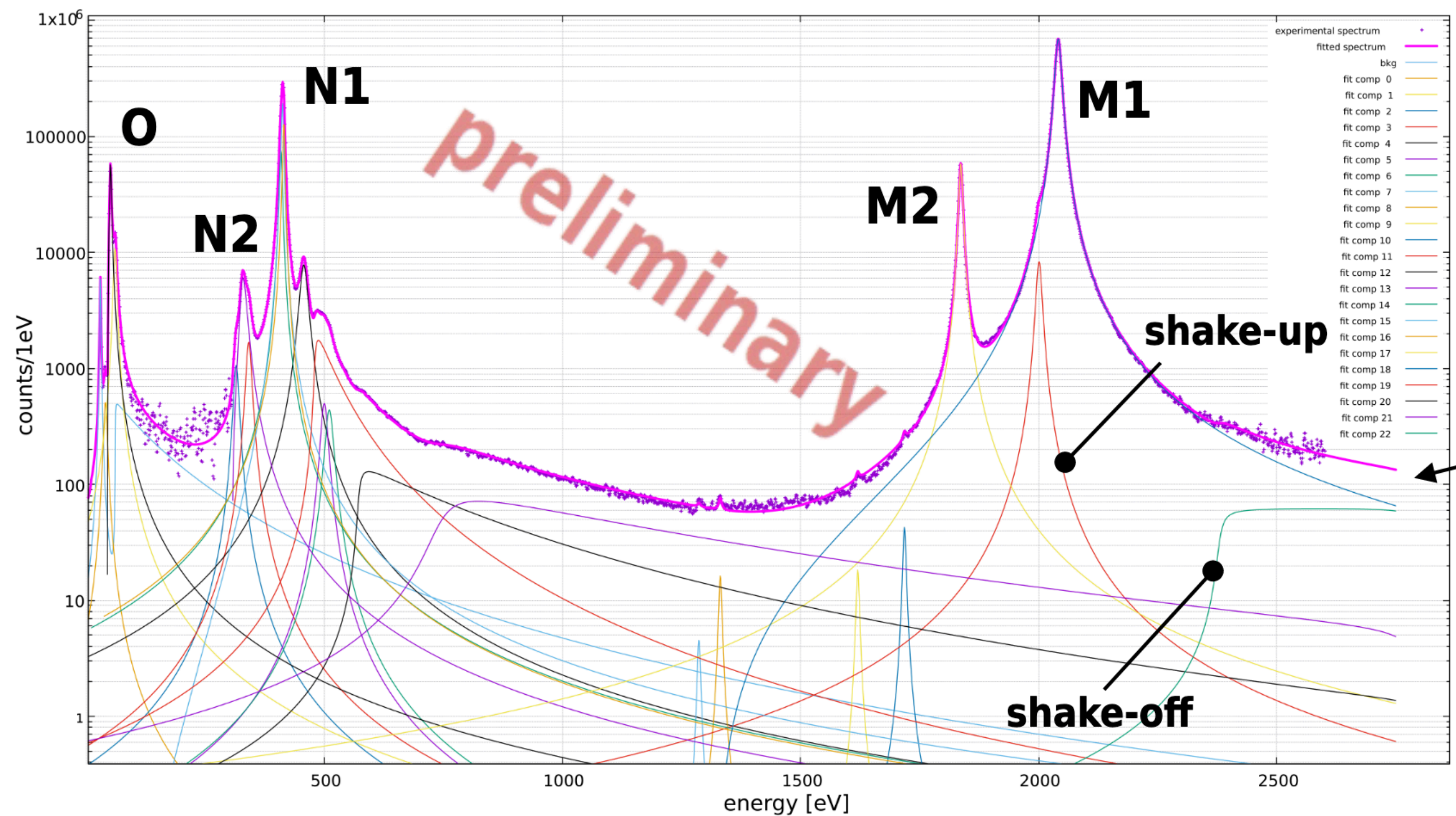


- $\Delta E_{\text{FWHM}} \sim 7 \text{ eV}$
- 7×10^4 detector x hour
- $\sim 7 \times 10^7$ ^{163}Ho decays

Single hole spectrum doesn't explain data...

2 holes excitation: the perturbation due to the $\text{Ho} \rightarrow \text{Dy}$ transition “shakes” one or more additional atomic electron to an upper bound state (**shake-up**) or to the continuum (**shake-off**).

Unfolding experimental ^{163}Ho spectrum...



Signal rate higher than single hole spectrum → impact on m_ν sensitivity?

Experimental EC spectrum deviates from theoretical predictions

→ phenomenological description of the EC spectrum is now under study

- shake-up peaks and shake-off spectra
- strongly asymmetric Lorentzians

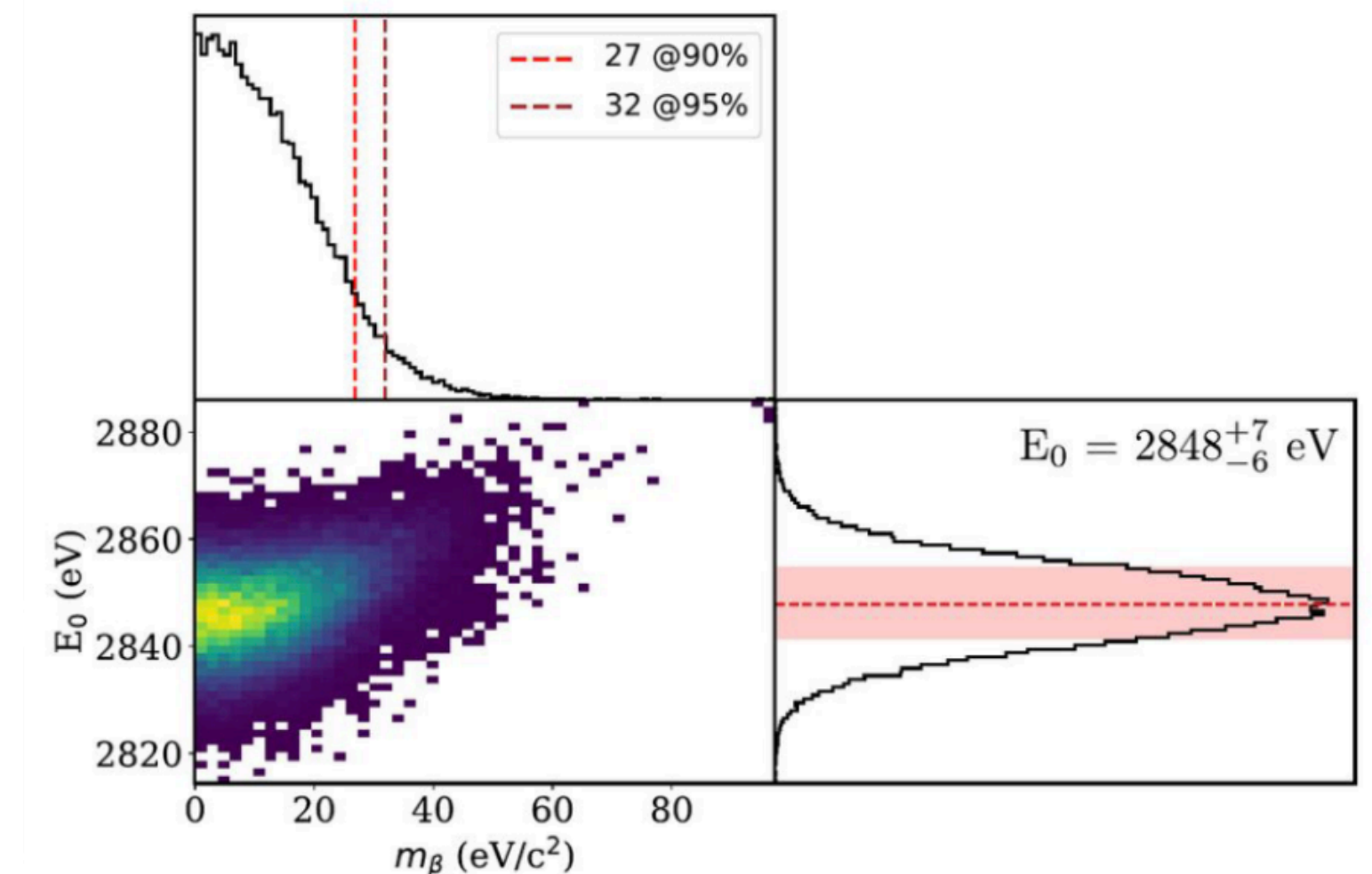
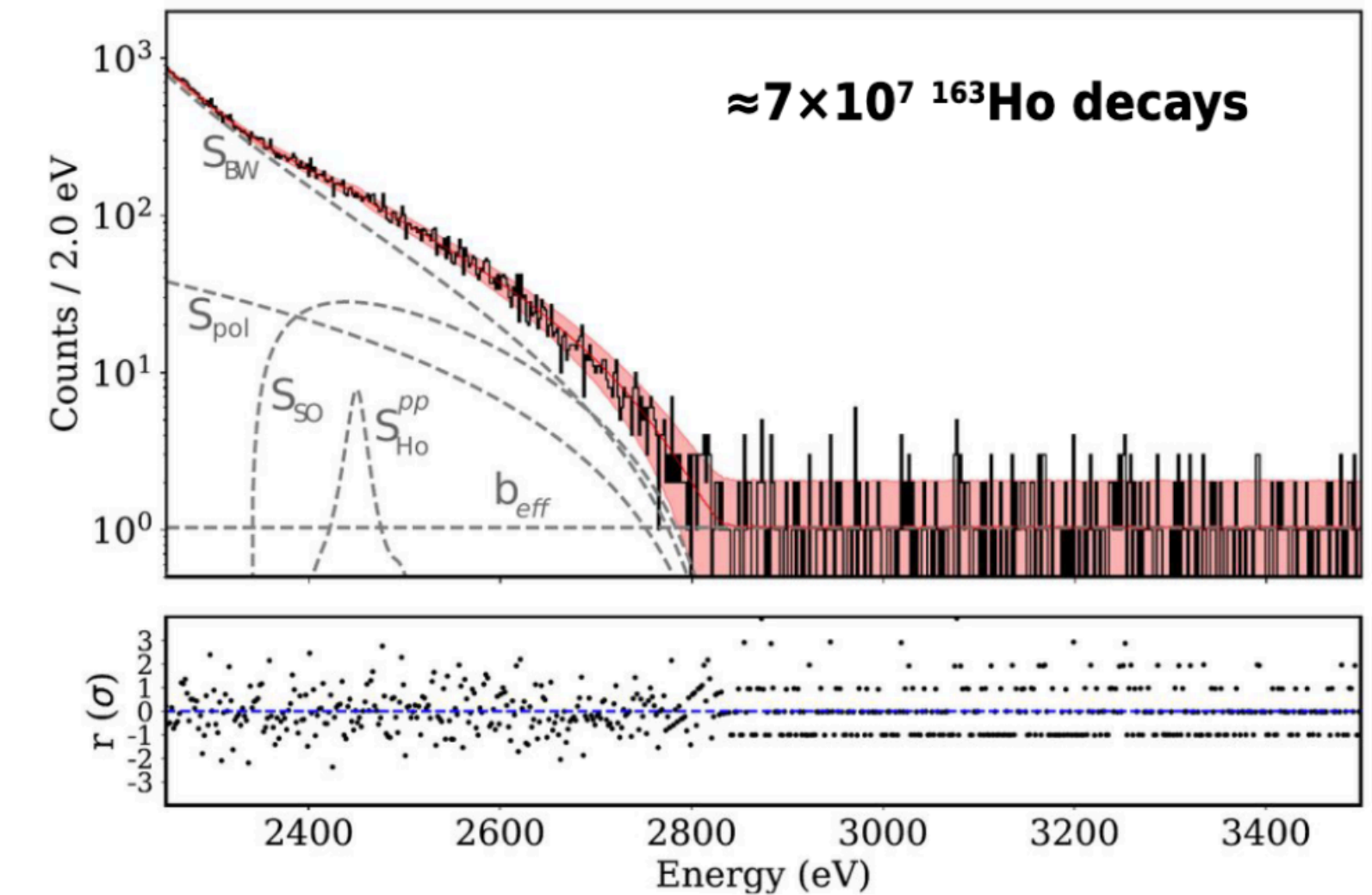
needed for assessing sensitivity of future ^{163}Ho experiments

Limit on neutrino mass

- **Bayesian analysis with 13 free parameters** performed in the ROI: (2250 - 3500) eV
- ^{163}Ho spectrum modeled as sum of few terms
- Posterior distributions explored through a Hamiltonian Markov chain Monte Carlo using STAN
- $\Delta E_{\text{FWHM}} \sim 7 \text{ eV}$, $f_{\text{pp}} < 10^{-5}$
- Background count rate $(1.7 \pm 0.1) \times 10^{-4} \text{ counts/eV/day/det}$
- **m_ν correlated only with E_0** (endpoint)
- Statistical sensitivity from toy MC $\sim 40 \pm 10 \text{ eV}$

$$m_\nu < 27 \text{ eV @ 90\% CI}$$

$$E_0 = 2848^{+7}_{-6} \text{ eV}$$

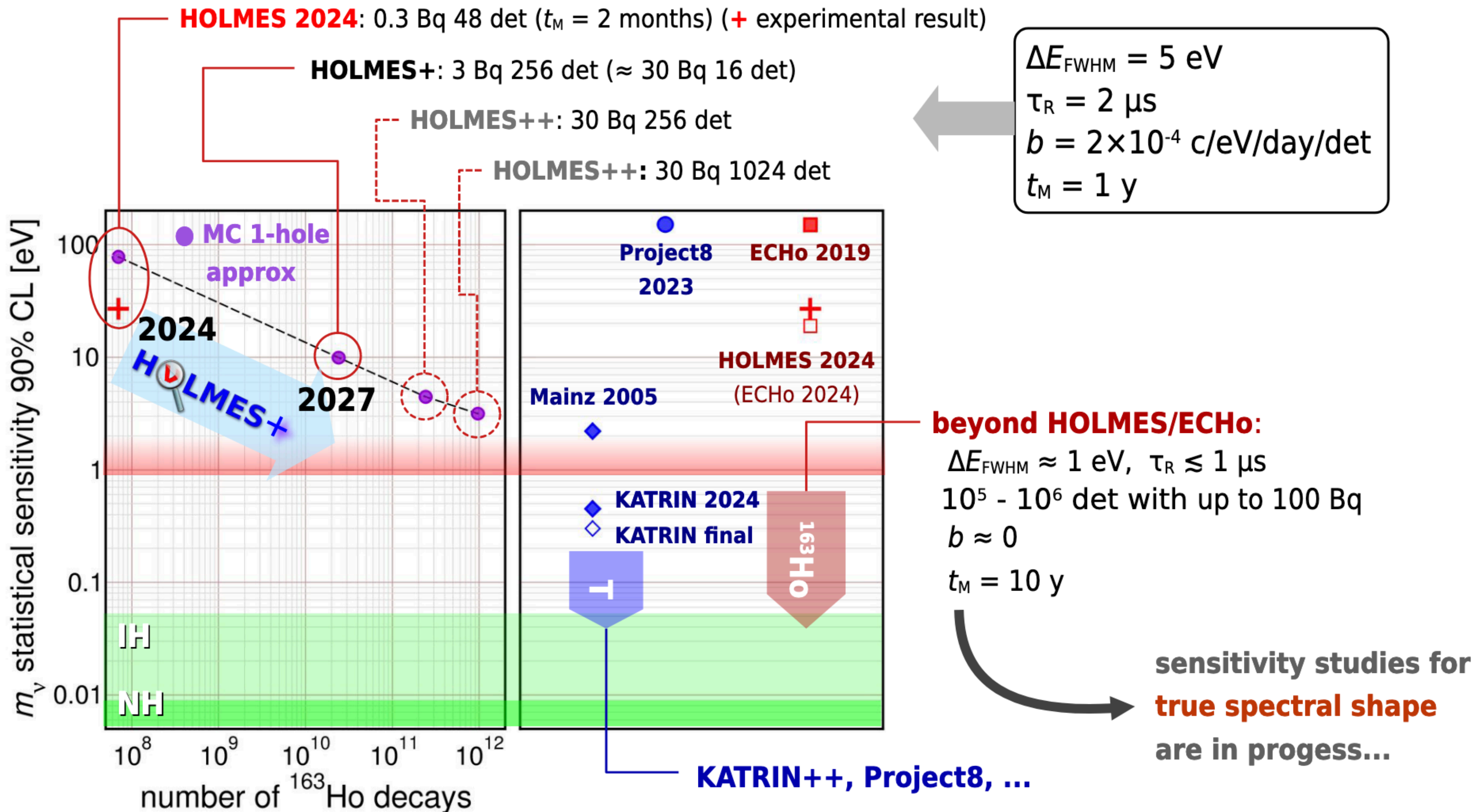


What next?

To reach the (sub) eV sensitivity we must:

- **Improve ion implanter for better beam control and higher efficiency**
 - Integrating 4poles + x-y stirring magnet + diagnostic elements + target co-evaporation chamber
 - Major upgrade of ion source → new (much) more efficient ionization technique is under study
- **Increase detector activity by 2 order of magnitude**
- **Reduce detector operating temperature down to 30 mK**
 - Addition of ^{163}Ho could spoil detector performance, must be recovered by lowering operating T
- **Increase the number of channel and reduce readout/DAQ cost → few €/ch**
- **0.1 eV sensitivity on m_ν requires scaling up HOLMES experiment by a factor $O(10^9)$...very aggressive but feasible!**

A possible roadmap towards sub eV sensitivity...



Summary

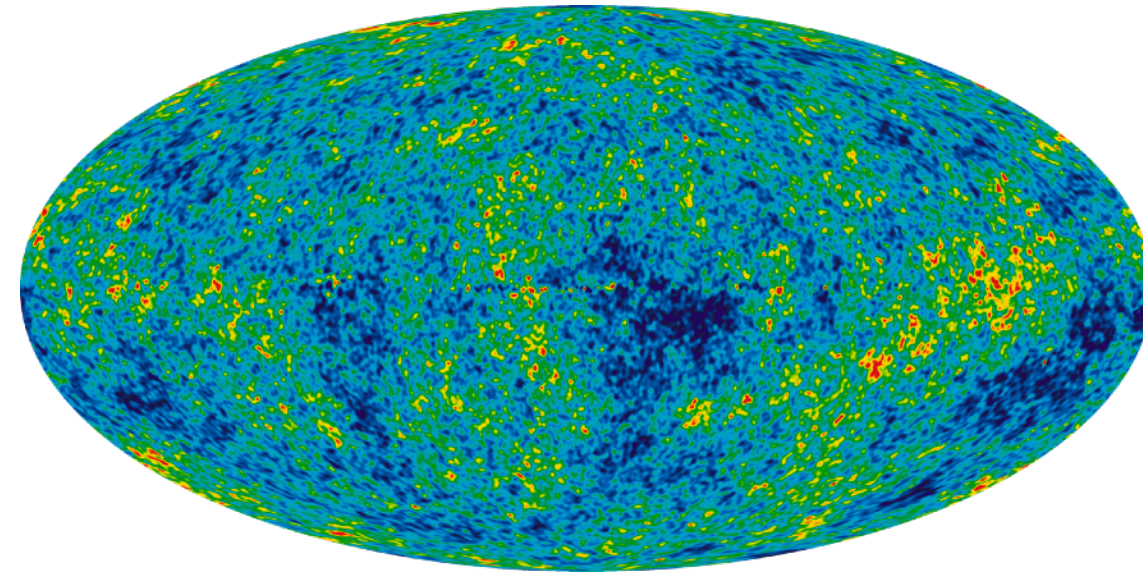
- HOLMES has successfully demonstrated the feasibility of the calorimetric measurement of ^{163}Ho EC decay spectrum.
- Even if still not competitive with the current best limit, this first result shows that the μ -calorimeters technology is ready and built the foundations for sub-eV sensitivity.
- In the next years, we plan to increase our sensitivity with a step by step approach
- New results will come (hopefully) soon!

Back-up

Neutrino mass measurements

Cosmological

$$M_\nu = \sum_i m_i$$



Best limit : 0.07 eV - analysis of CMB and structure formation data - model dependent

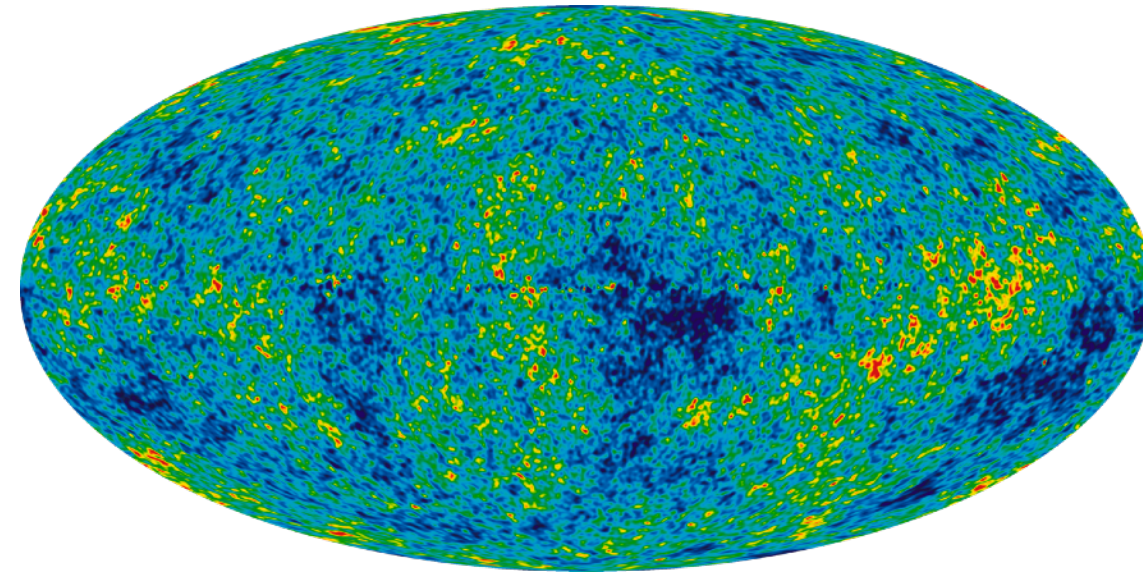
DESI + CMB + PLANCK

**neutrino mass
measurements**

Neutrino mass measurements

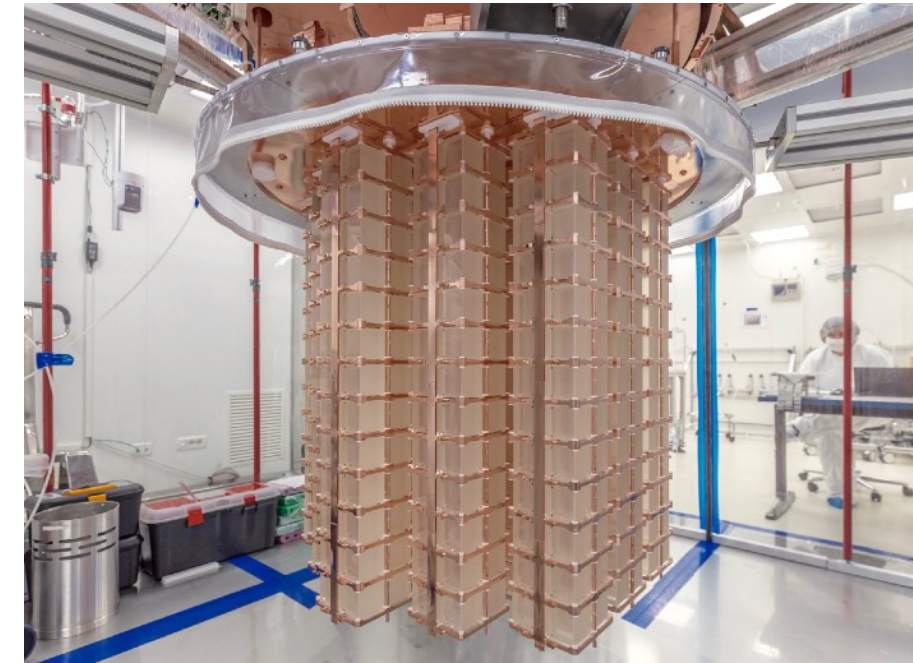
Cosmological

$$M_\nu = \sum_i m_i$$



Best limit : 0.07 eV - analysis of CMB and structure formation data - model dependent

DESI + CMB + PLANCK



$0\nu\beta\beta$

$$m_{\beta\beta}^2 = \left| \sum_i U_{ei}^2 m_i \right|^2$$

Best limit : 0.05 - 0.15 eV, inference neutrino mass from $0\nu\beta\beta$ decay, model-dependent

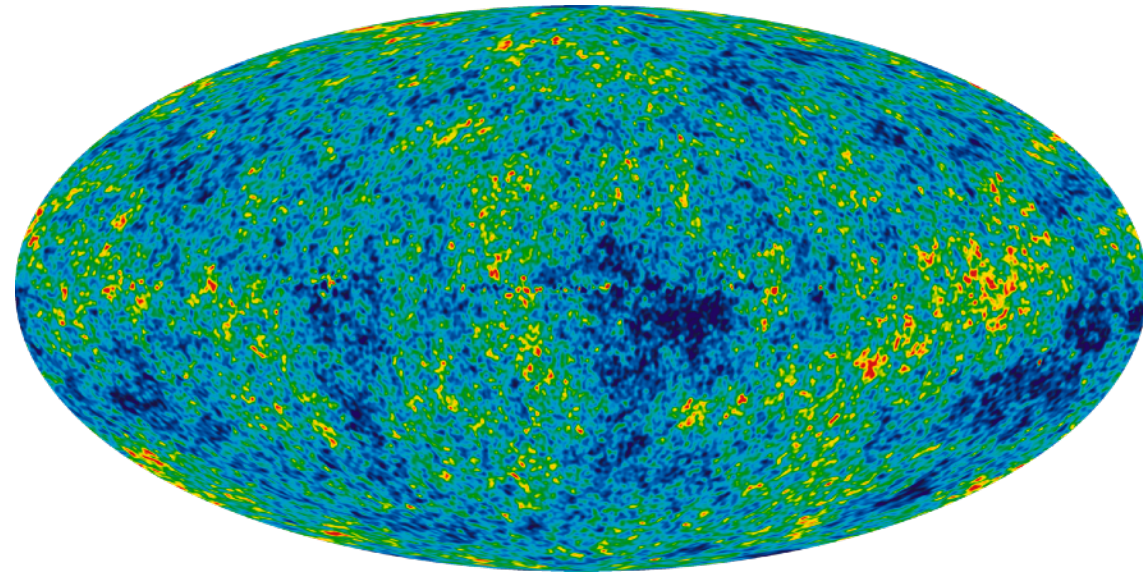
KamLAND-Zen, GERDA, CUORE, EXO, SNO+, Majorana, Nemo 3, COBRA...

neutrino mass measurements

Neutrino mass measurements

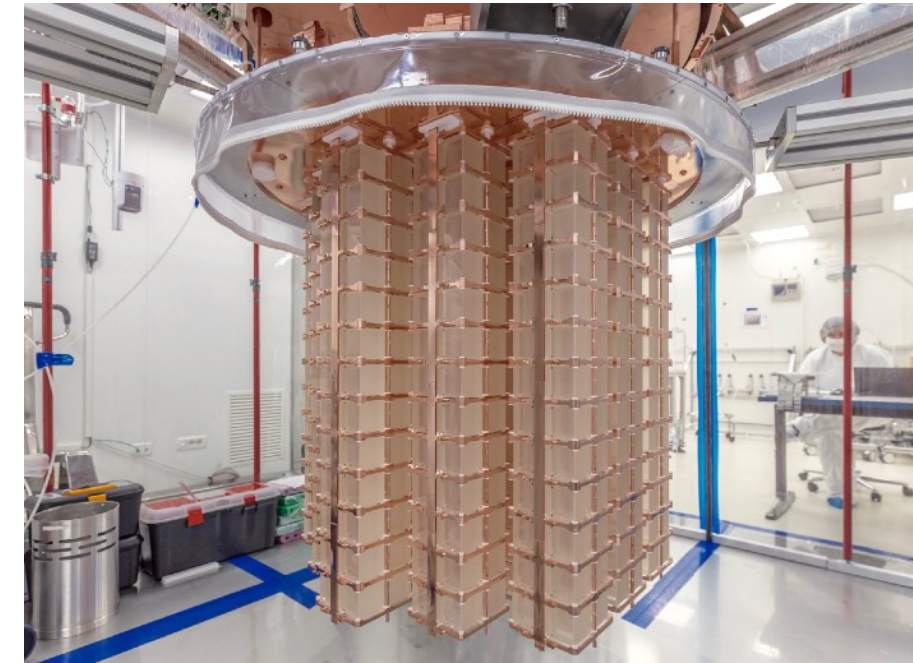
Cosmological

$$M_\nu = \sum_i m_i$$



Best limit : 0.07 eV - analysis of CMB and structure formation data - model dependent

DESI + CMB + PLANCK



$0\nu\beta\beta$

$$m_{\beta\beta}^2 = \left| \sum_i U_{ei}^2 m_i \right|^2$$

Best limit : 0.05 - 0.15 eV, inference neutrino mass from $0\nu\beta\beta$ decay, model-dependent

KamLAND-Zen, GERDA, CUORE, EXO, SNO+, Majorana, Nemo 3, COBRA...

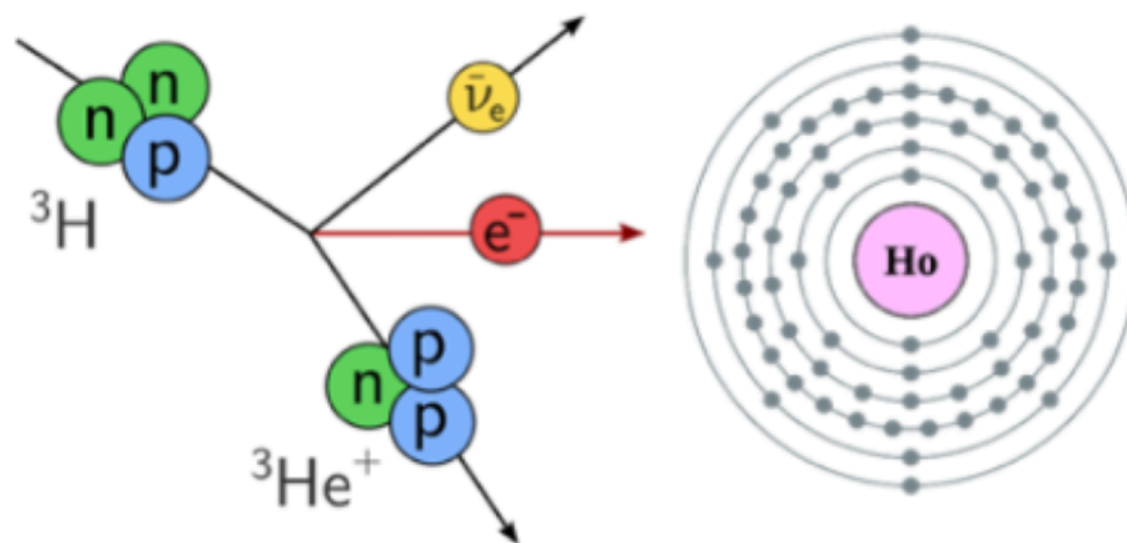
neutrino mass measurements

Direct measurements

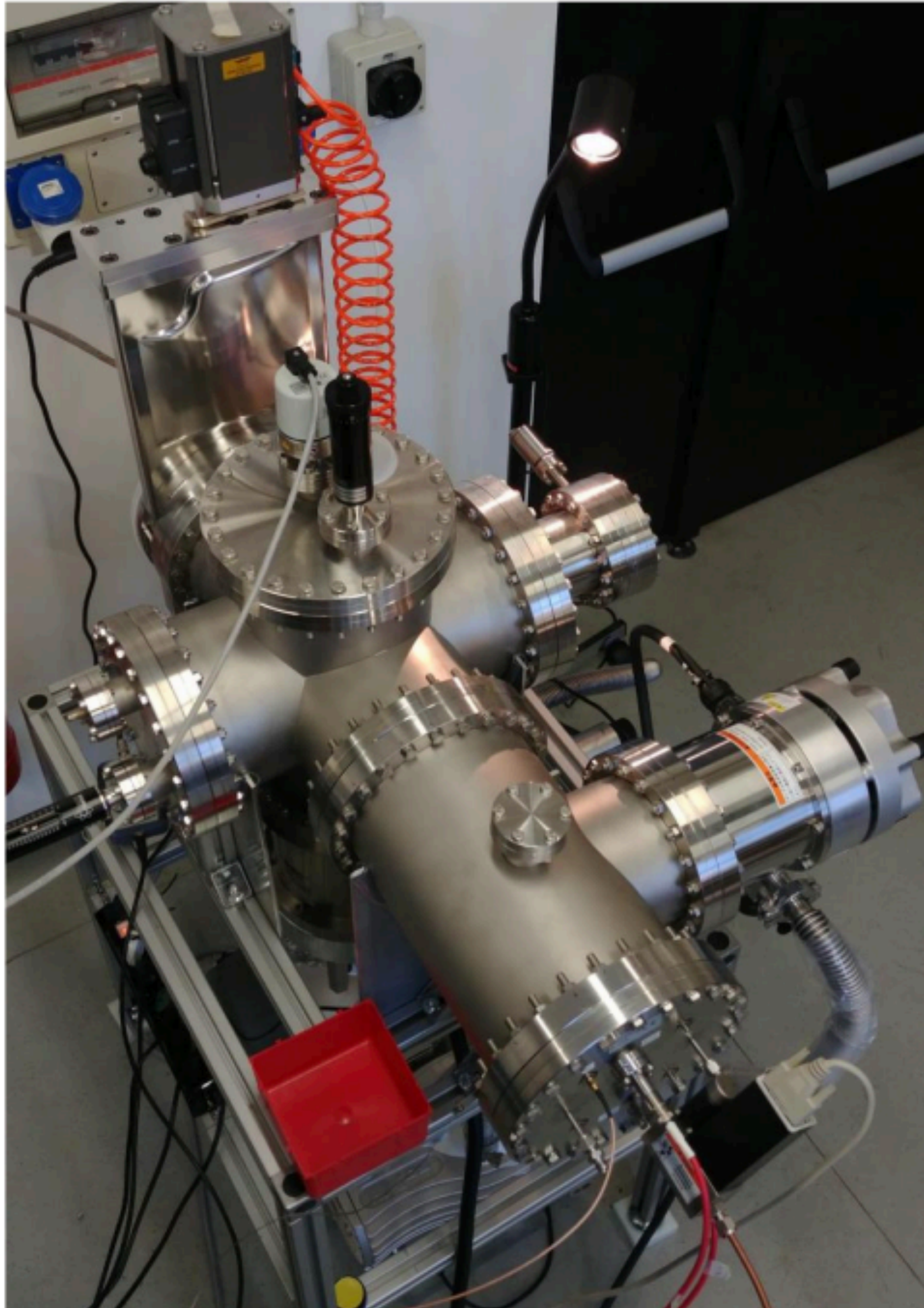
$$m_\beta^2 = \sum_i |U_{ei}|^2 m_i^2$$

Best limit: 0.45 eV, exploit kinematics of weak decays, model independent

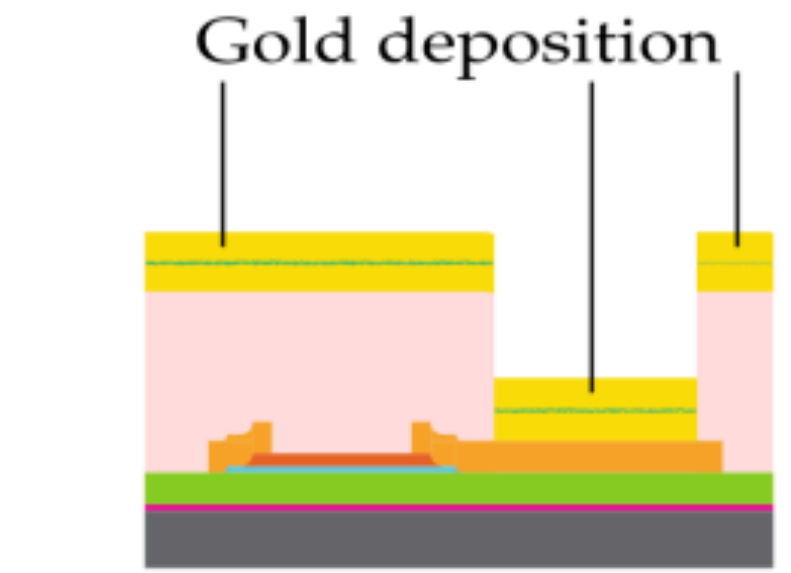
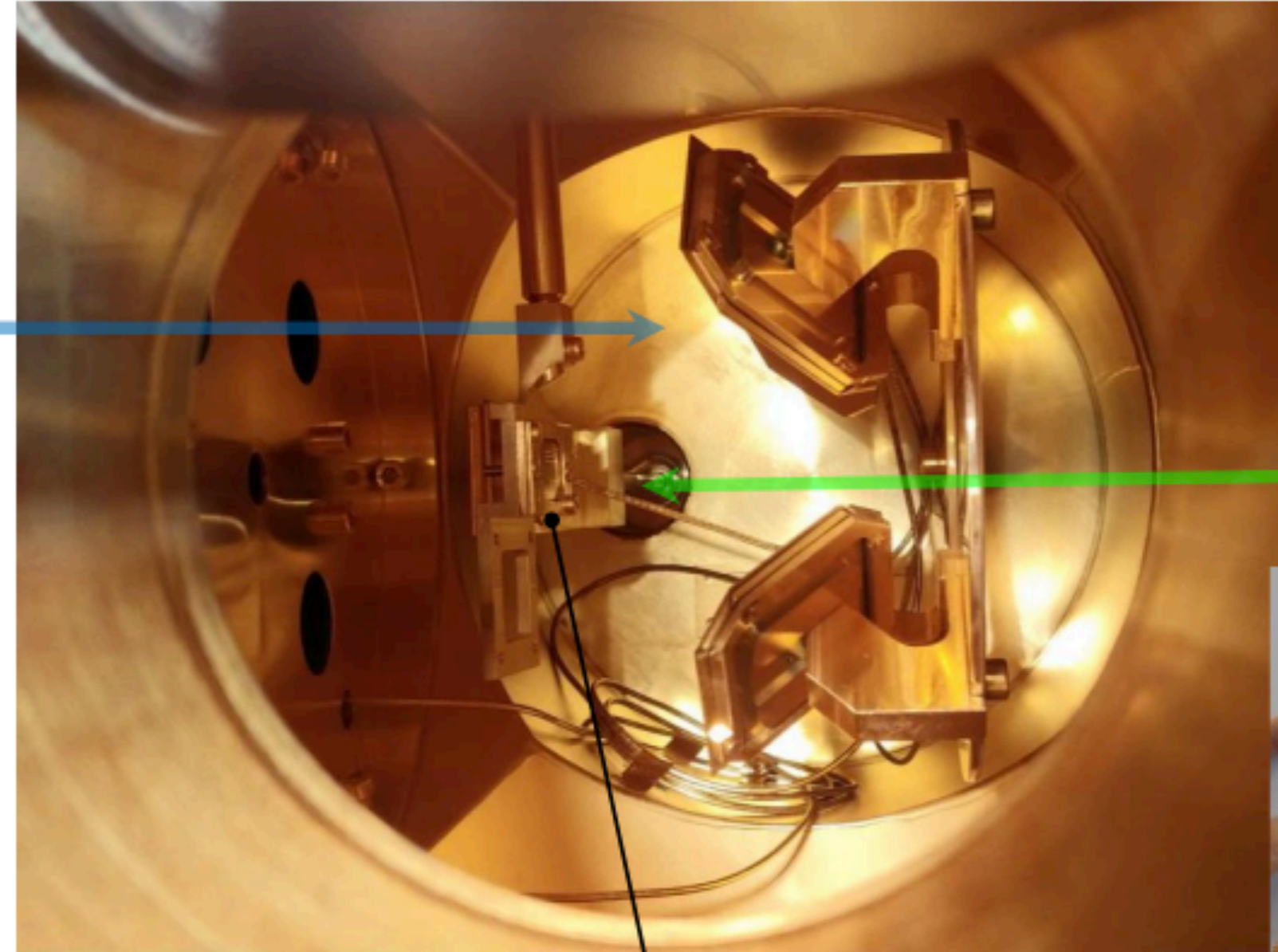
KATRIN, HOLMES, ECHO, Project-8, NuMECS...



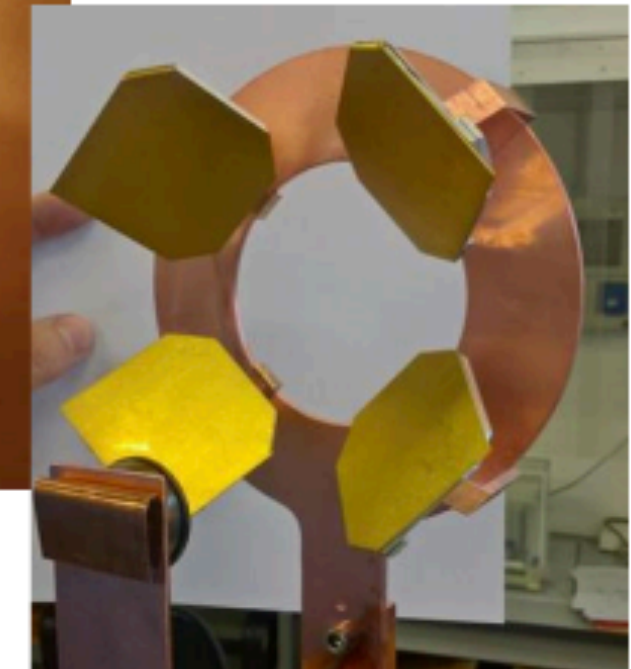
Full encapsulation @Milano-Bicocca



Ar ion
beam (for
sputtering)



^{163}Ho (not
implemented yet)

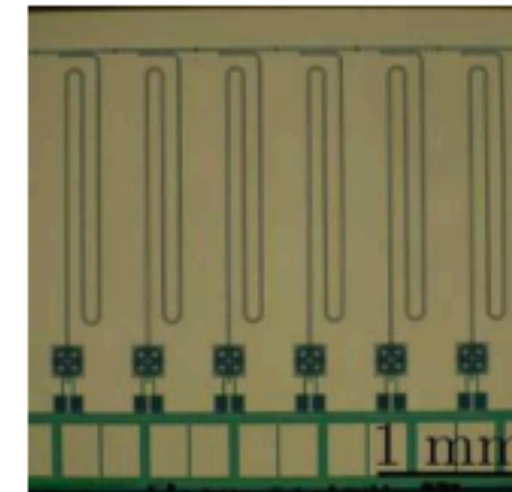
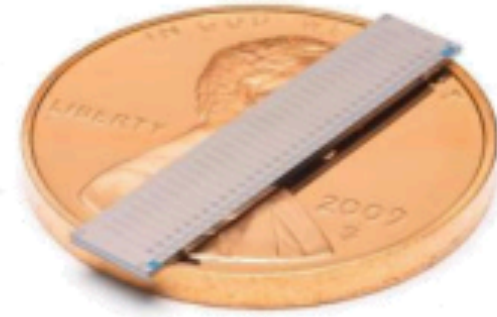
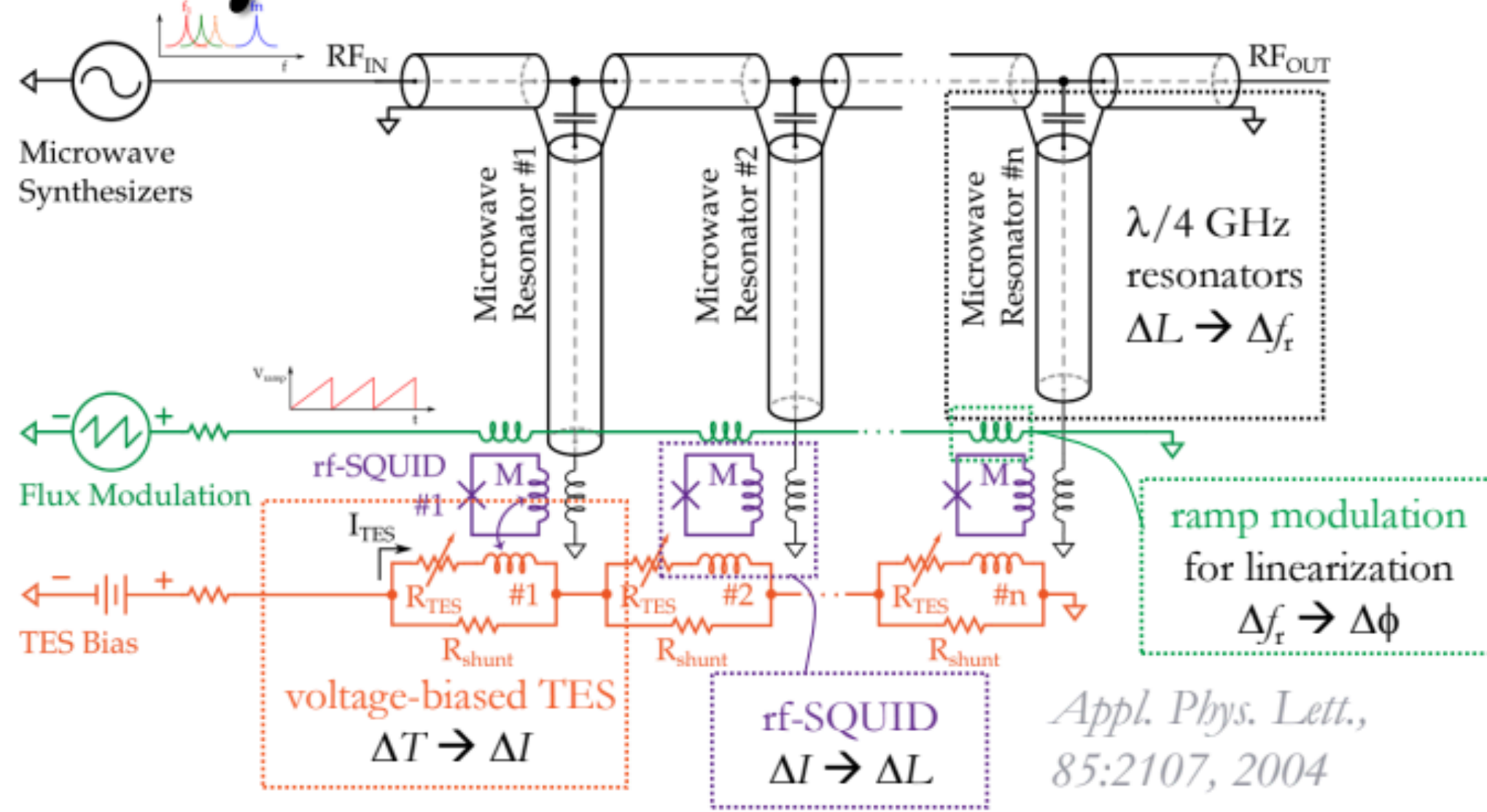


TES array

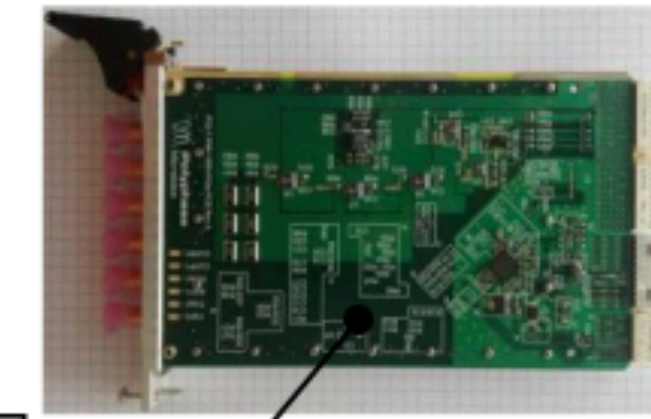
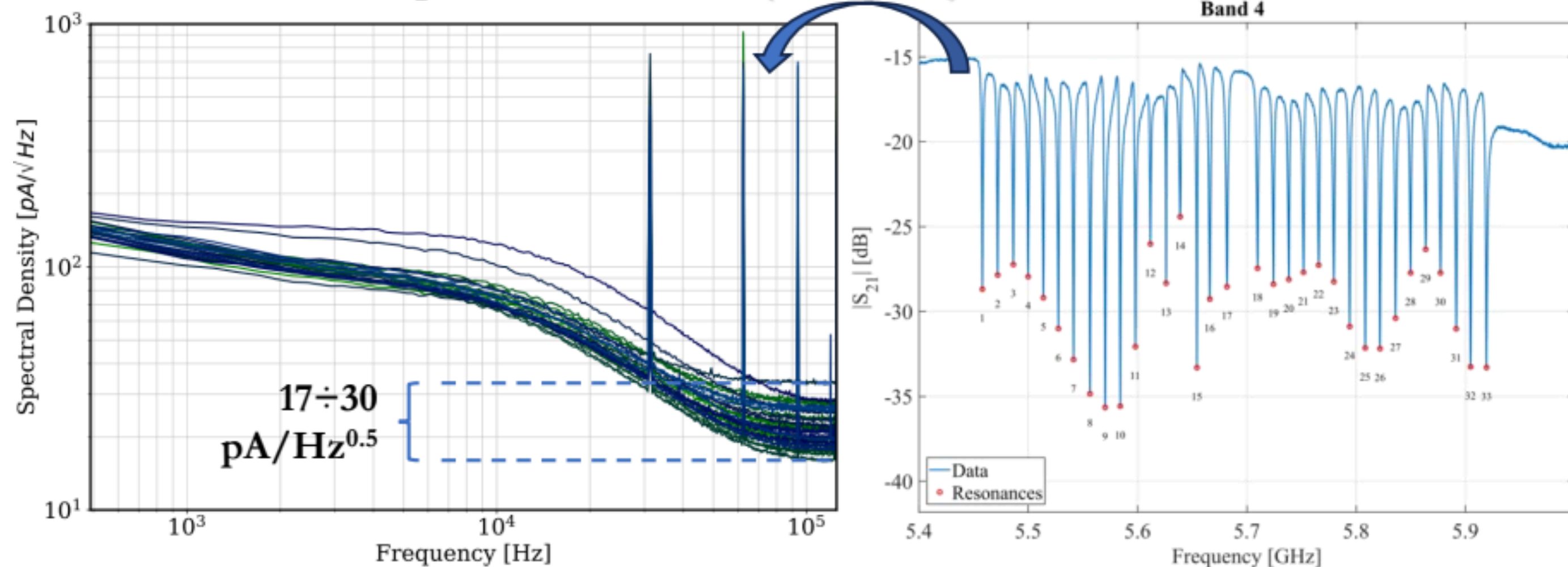
- 1 μm layer to fully encapsulate the ^{163}Ho
- ≈ 27 hours to complete the process
- soon will be integrated with the ion-implanter to compensate sputter and avoid oxidation



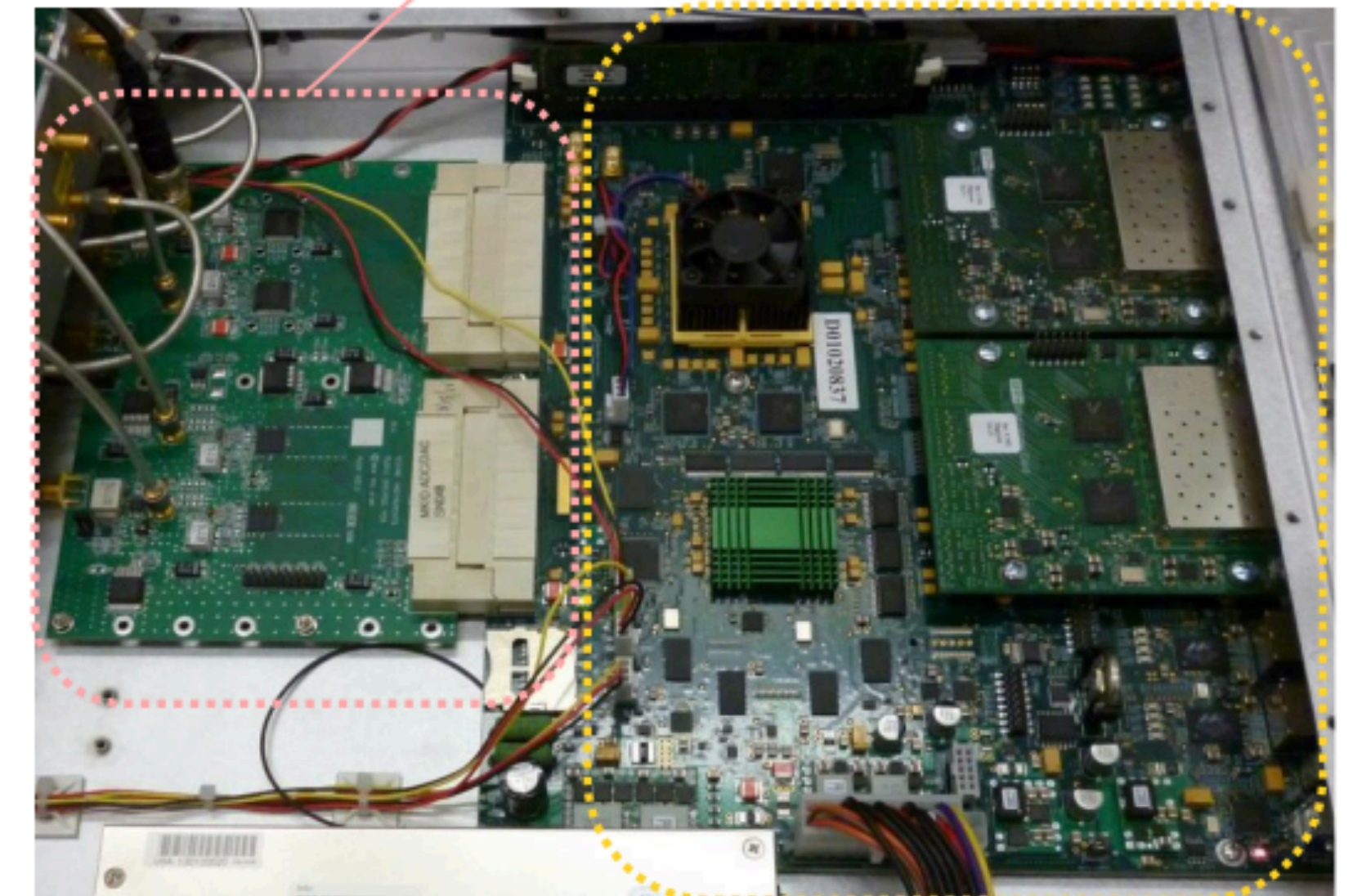
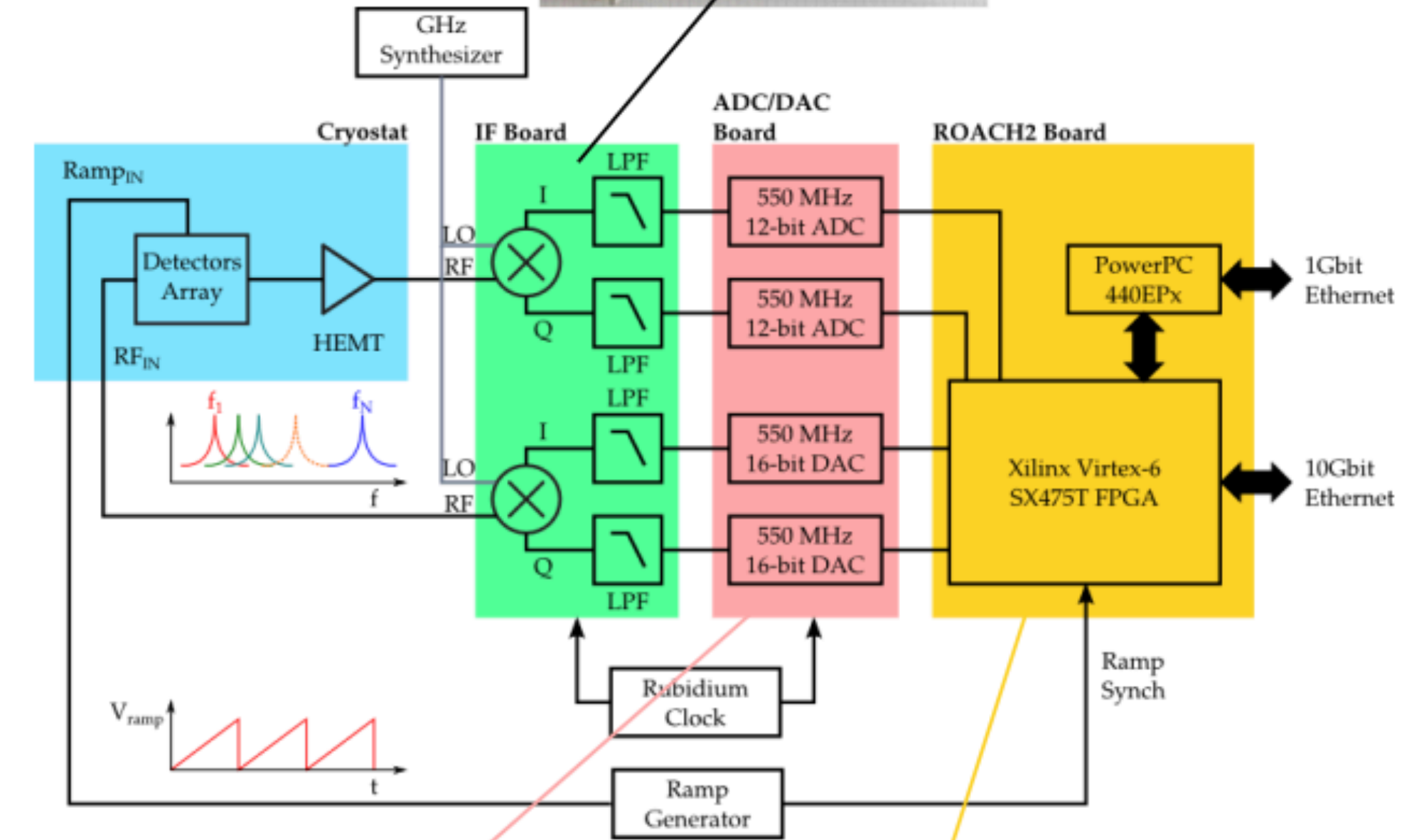
μ wave mux readout & DAQ



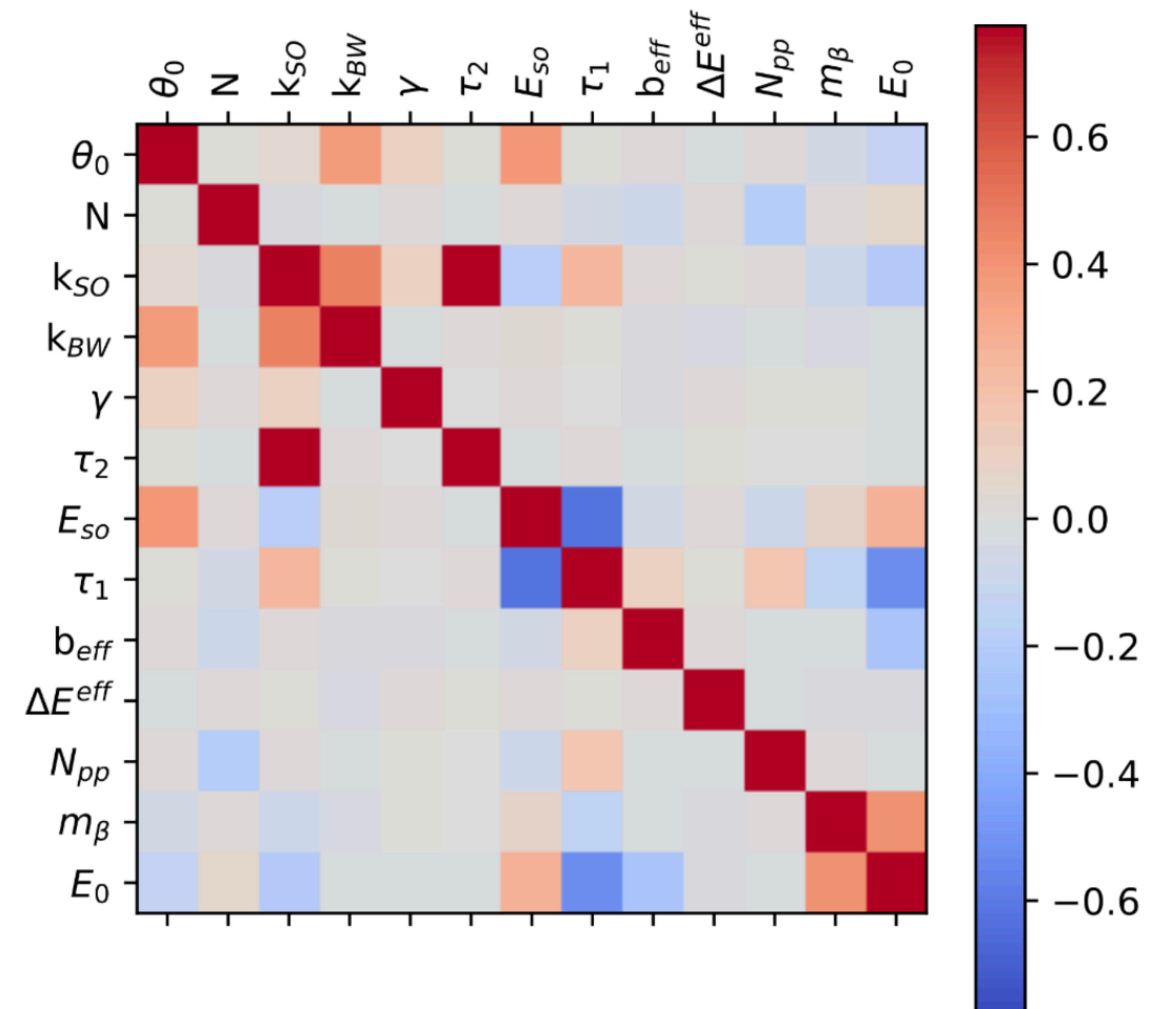
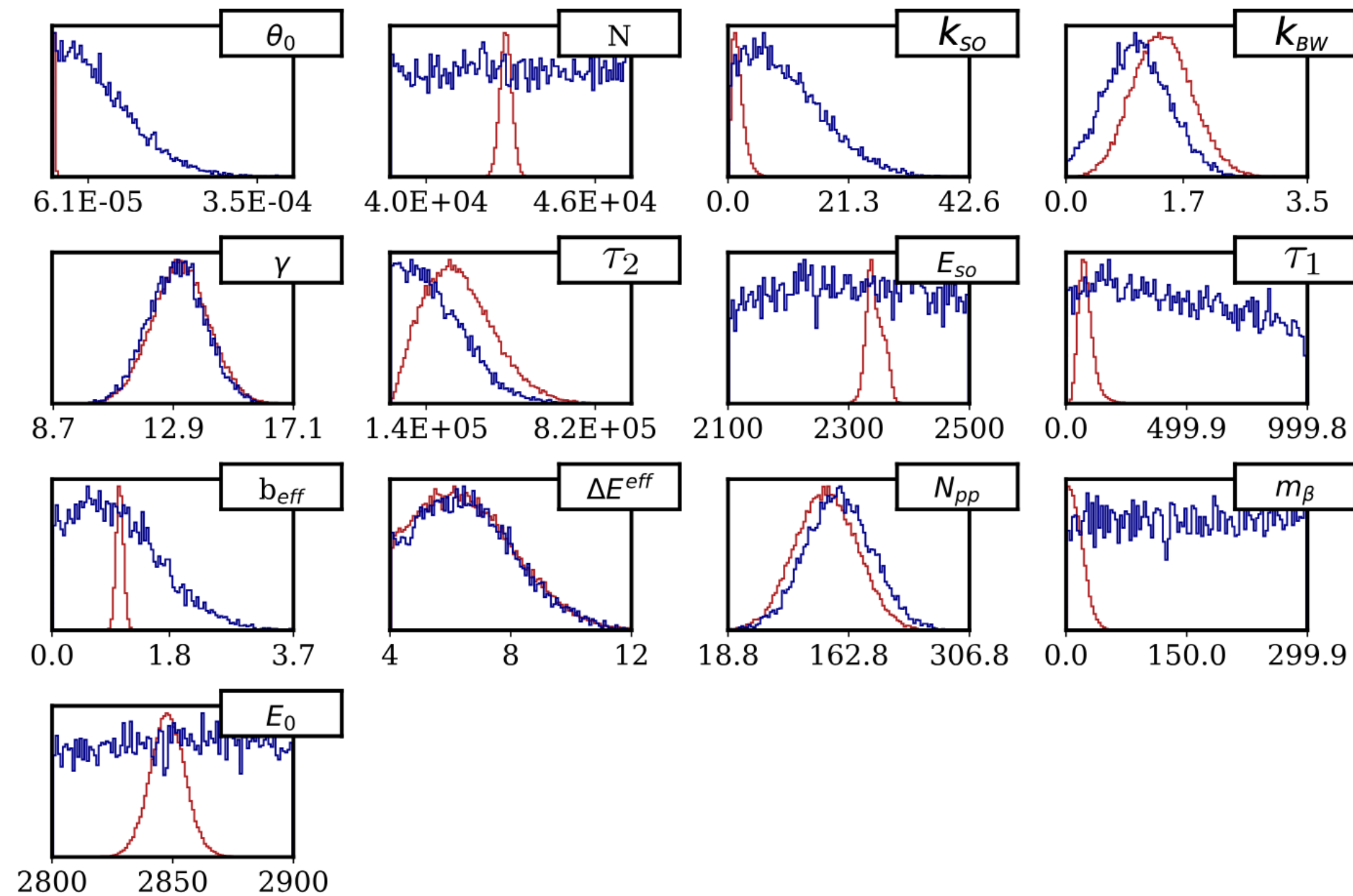
- $f_{\text{ramp}} = f_{\text{samp}} (= 500 \text{ kHz})$
- allows to readout many detectors with one common RF line and HEMT amplifier: **1 HEMT (4-8 GHz) = 256 TESs** JINST 14 P10035 (2019)



64 channels already available



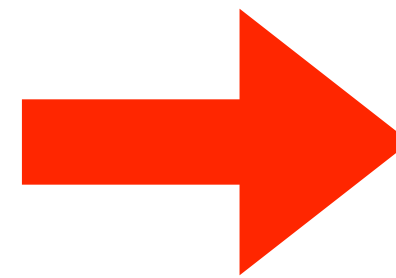
Parameters distributions and correlations



What next?



- Proved:
 - Multiplexed TES 64 pixel array
 - $\sim 1\text{Bq/det}$
 - $\Delta E_{\text{FWHM}} = 5 - 8\text{ eV}$
 - $\tau \sim 1\text{ }\mu\text{s}$
- m_ν stat sensitivity $\sim 20\text{ eV}$



- Mid term (3 years) plan for:
 - m_ν stat sensitivity $< 10\text{ eV}$
 - up to 256 detectors
 - up to 3 Bq / detector