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Inference of the Mass Composition of Cosmic Rays with energies between 3 and 100 EeV using the data of the Pierre Auger Observatory and Deep Learning

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One of the open questions of astrophysics is the mass composition of ultra-high-energy cosmic rays (UHECRs). The flux of UHECRs is extremely low, demanding large observatories for indirect measurements of cosmic-ray air showers, cascades of secondary particles created by interactions of the cosmic ray with the atmosphere. Located in Argentina, the Pierre Auger Observatory is the largest cosmic-ray observatory on Earth. The Observatory is a hybrid detector employing different detection principles to observe multiple components of air showers. The core part of the detector is the Surface Detector (SD), which comprises 1 600 water-Cherenkov detectors with 1.5 km spacing in an area of 3000 km². The highly sensitive Fluorescence Detector (FD) overlooks the area above the SD. Since the FD can only operate on moonless nights, its duty cycle is limited to approximately 15%.

The indirect nature of measurements of the Pierre Auger Observatory poses several challenges. For example, estimating the mass of a primary cosmic ray. The atmospheric depth of the shower maximum X_{\max} is a mass-sensitive observable. The FD observes the X_{\max} directly but can measure only a subset of the detected events due to its duty cycle.

On the contrary, the SD of the Pierre Auger Observatory, operating almost at 100% duty cycle, allows for a significant increase in the data. In this contribution, we present the X_{\max} reconstruction based on deep neural networks that extend the energy range and statistics. We probe the energy evolution of the mean and standard deviation of the reconstructed X_{\max} , which reflect the changes in the mass composition. The features found in the average X_{\max} rate suggest a heavier and purer mass composition with increasing energy.

Secondary track

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