## IMPERIAL



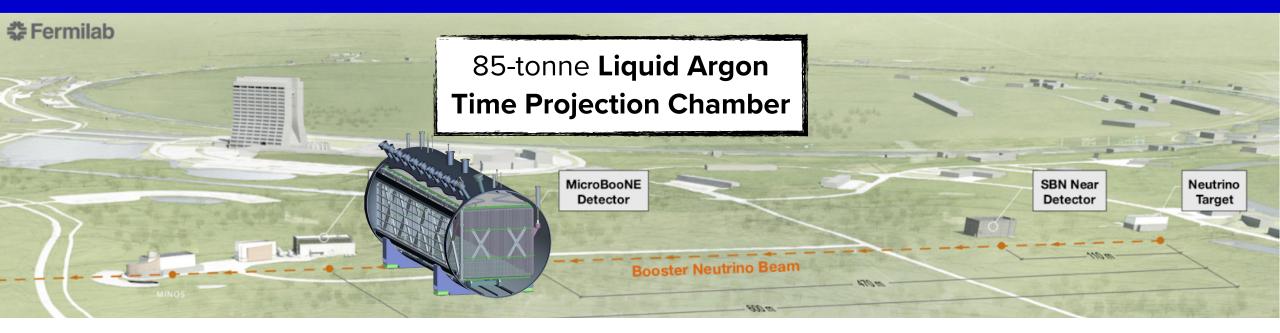
### Searches for physics beyond the Standard Model with the MicroBooNE Experiment

### Anyssa Navrer-Agasson,

On behalf of the MicroBooNE Collaboration

EPS-HEP - 9th July 2025

## **The MicroBooNE Detector**

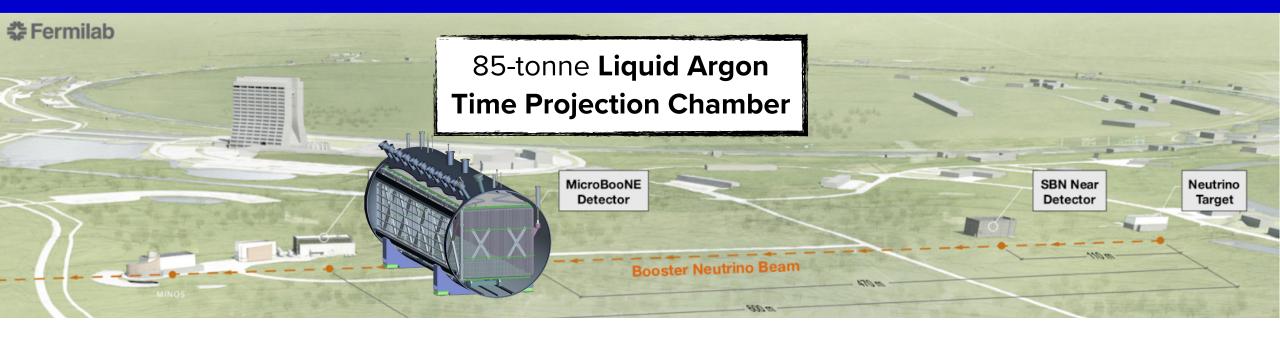


- Neutrino cross-section measurements
- Detector physics, R&D, tool development
- Beyond Standard Model physics → This talk!

#### IMPERIAL



## **The MicroBooNE Detector**

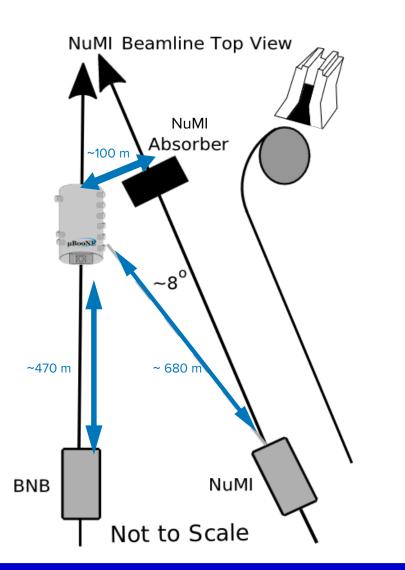


- Neutrino cross-section measurements
- Detector physics, R&D, tool development
- Beyond Standard Model physics → This talk!



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### **BSM Physics @ MicroBooNE: beamlines**



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#### **MicroBooNE** observes flux from two neutrino beams

#### **Booster Neutrino Beam (BNB)**

- 8 GeV protons
- Target ~470 m from MicroBooNE
- On-axis

#### Neutrino at the Main Injector (NuMI)

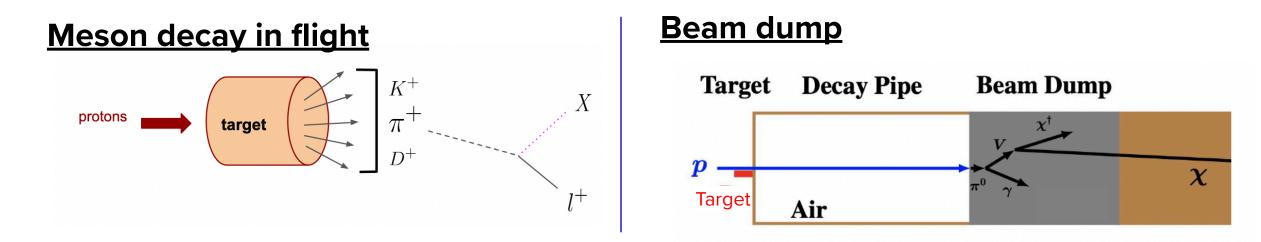
- 120 GeV protons
- ~680 m away from target
- ~8° off-axis
- Absorber ~100 m from MicroBooNE





### **BSM Physics @ MicroBooNE: beamlines**

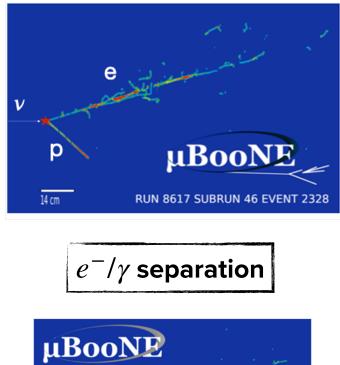
- Large flux of charged/neutral mesons from high intensity proton beams
- New particles can be produced from meson decays
- **Proximity to the NuMI absorber** → particles survive long enough to reach MicroBooNE

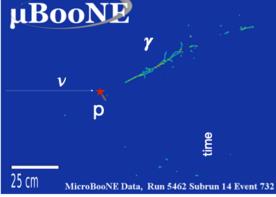


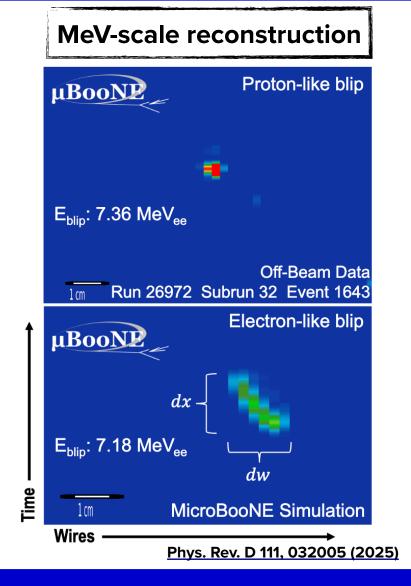
#### IMPERIAL



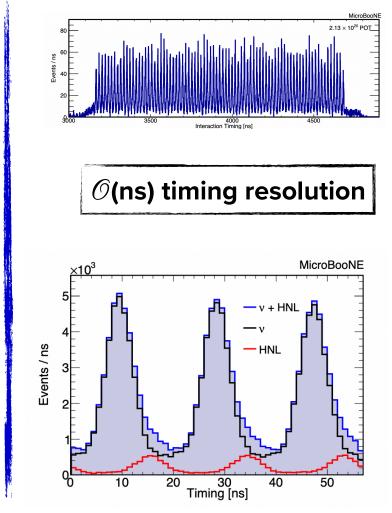
### **BSM Physics @ MicroBooNE: LArTPC**







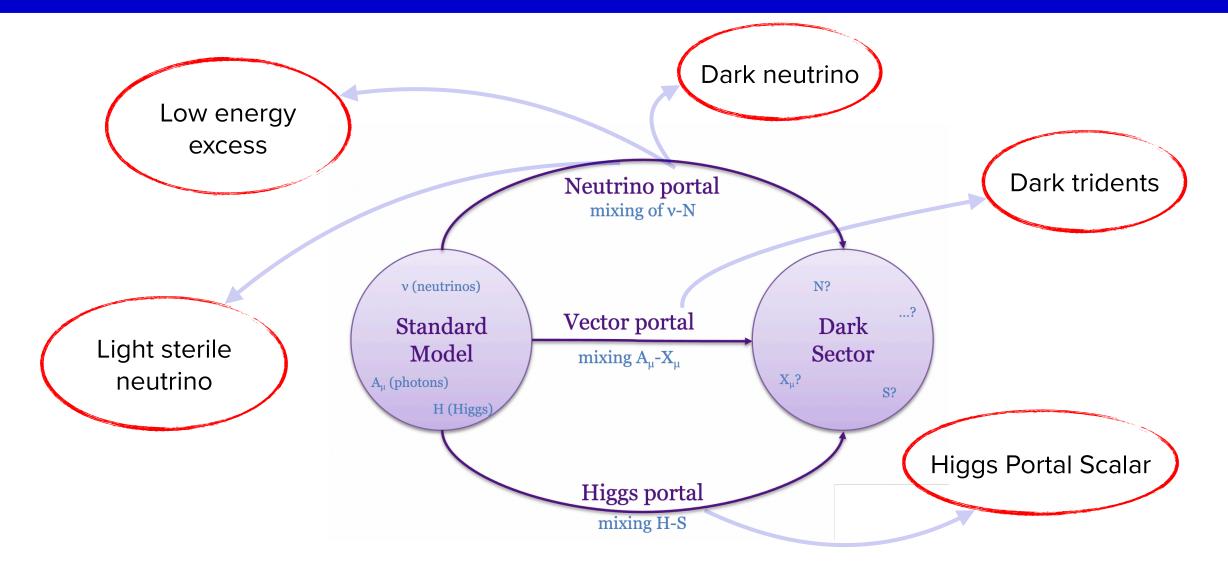
#### <u>Phys. Rev. D 108, 052010</u>



**IMPERIAL** A. Navrer-Agasson - EPS-HEP - 9<sup>th</sup> July 2025



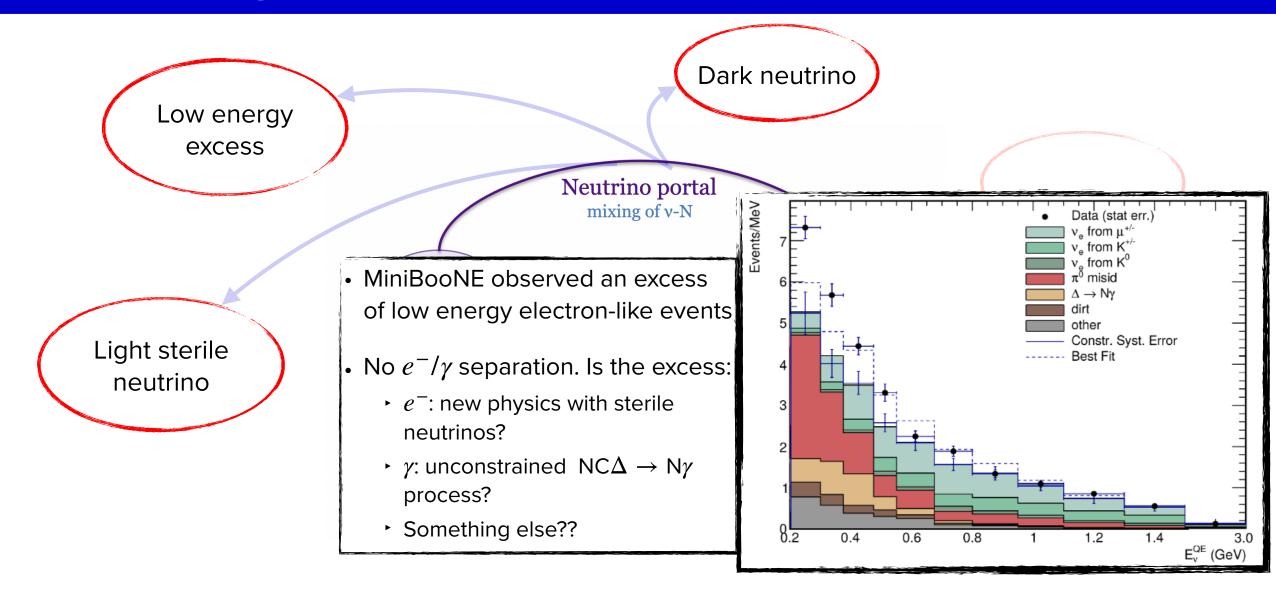
### **Exploring the dark sector**



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### **Exploring the dark sector**

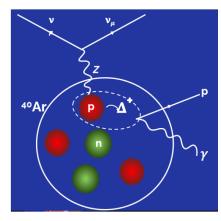


#### IMPERIAL A

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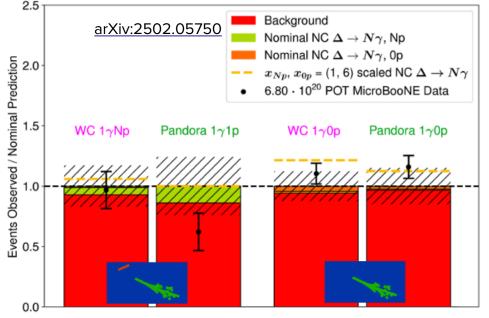
μBooNE

## Low energy excess: photon-like



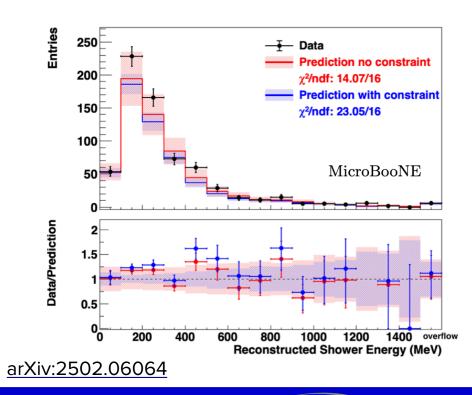
#### Exclusive

• Rejection of the NCA  $\rightarrow$  N $\gamma$  process at 94.4% CL



#### Inclusive

- Overall data/MC agreement across the board
- $2\sigma$  excess in the 0 proton final state

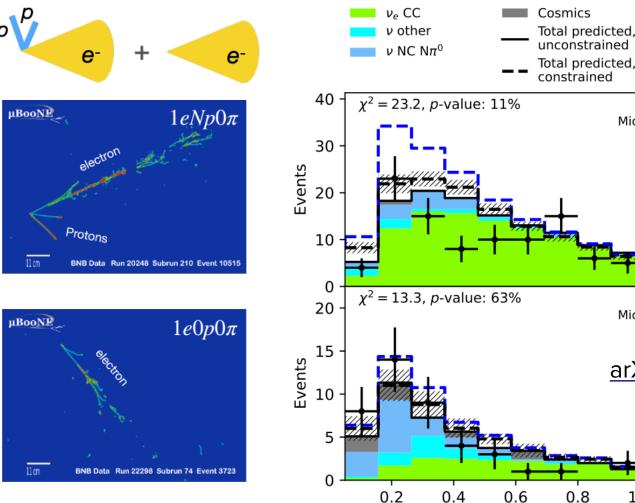


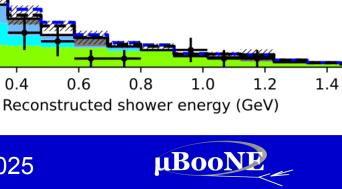
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# Low energy excess: electron-like

- No sign of excess in analysed data
- Reject the electron-like hypothesis as explanation for the MiniBooNE result at > 99% CLs
- Accepted for publication in PRL





11/1/ Uncertainty

Data

 $1eNp0\pi$  selection

 $1e0p0\pi$  selection

MicroBooNE,  $1.11 \times 10^{21}$  POT

arXiv:2412.14407

MicroBooNE,  $1.11 \times 10^{21}$  POT

LEE signal

model 2

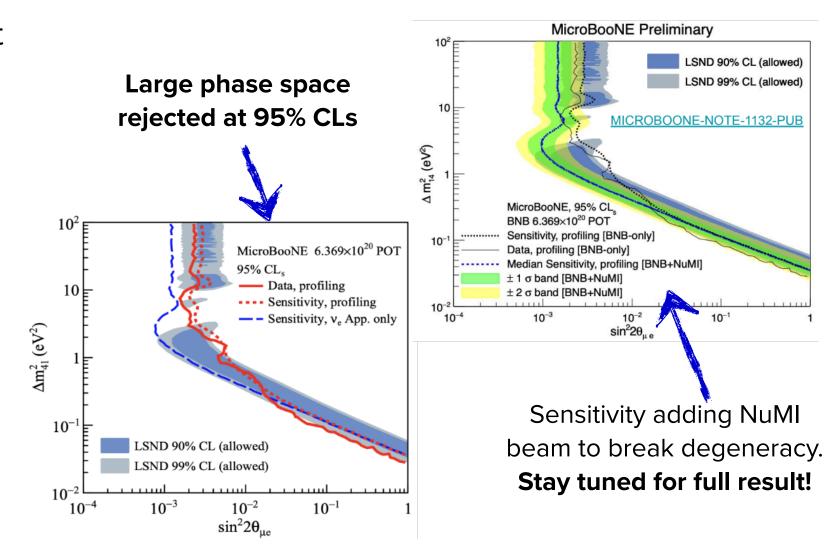
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## **Sterile neutrino search**

Inclusive search for a light
 sterile neutrino

 $\begin{pmatrix} \nu_{e} \\ \nu_{\mu} \\ \nu_{\tau} \\ \nu_{s} \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ \end{pmatrix} \begin{pmatrix} \nu_{1} \\ \nu_{2} \\ \nu_{3} \\ \nu_{4} \end{pmatrix}$ 

• Simplest 3 + 1 model



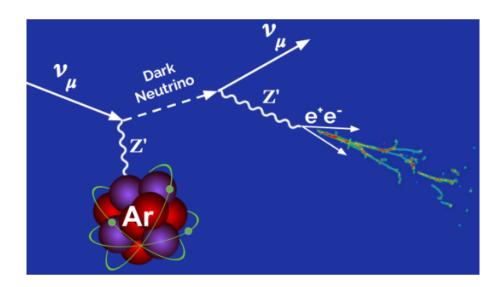


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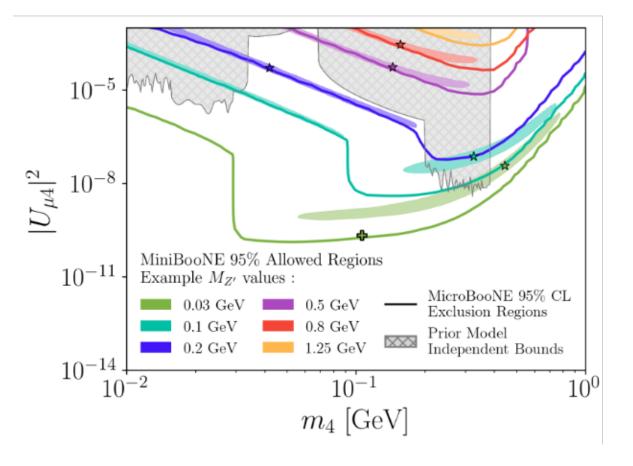
uBooNE

## **Dark neutrino**

- Active neutrinos from the BNB scattering to produce a dark neutrino,  $\nu_4$
- . Look for the  $\nu_4 \rightarrow \nu_\mu (Z' \rightarrow e^+ e^-)$  decay
- MiniBooNE excess as this model excluded at  $\geq$  95% CLs

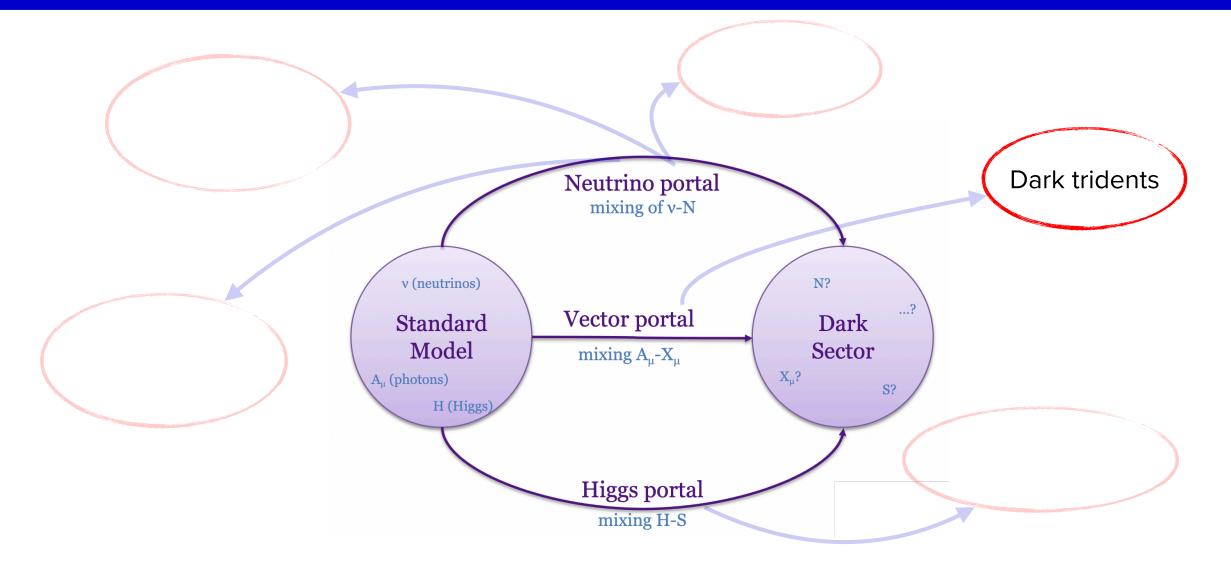


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### **Exploring the dark sector**



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μBooNE

- Dark matter produced in the beam via dark photon mixing
- Scatter off argon nucleus, accompanied by a dark photon radiation subsequently decaying to  $e^+e^-$

~30

Target

NuMI Beamline Side View

π

~8

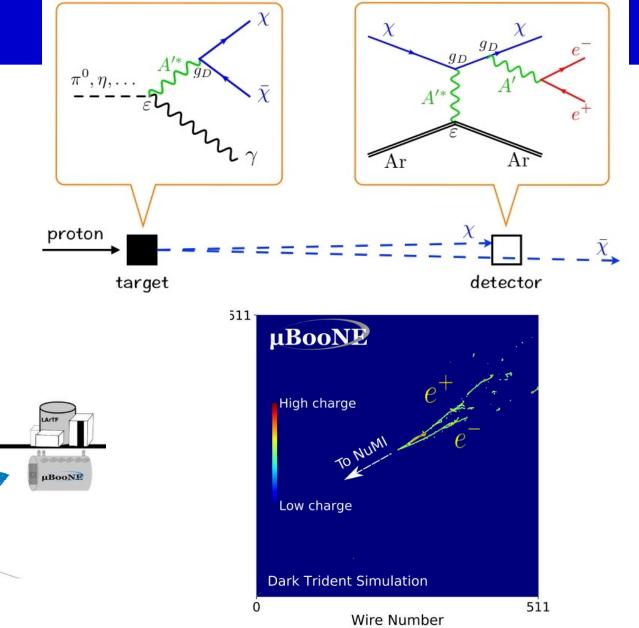
Horns

- **NuMI off-axis search**: neutrino background reduction
- First search in a LArTPC!

Main Injector

120 GeV

Beam



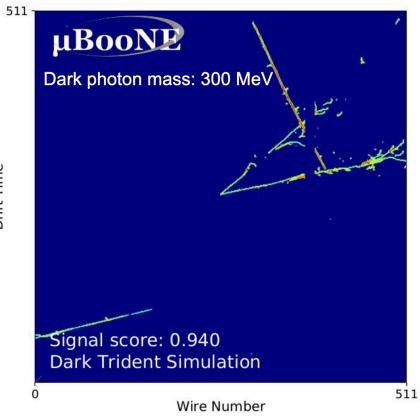
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Neutral mesons are not focused by

magnetic horns

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### Selection done with a Convolutional Neural Network



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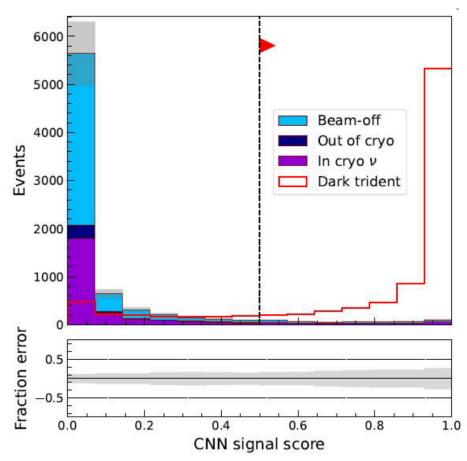


14

Drift Time

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### Selection done with a Convolutional Neural Network



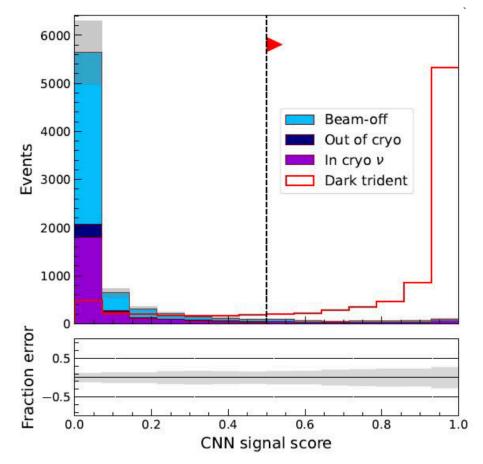
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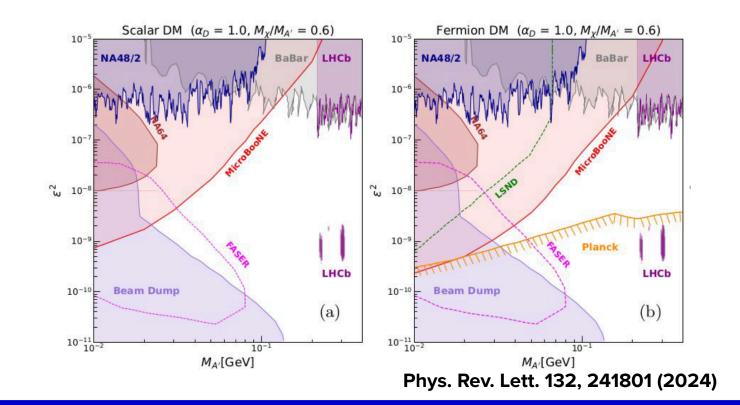


14

Selection done with a Convolutional Neural Network



- Consider **both scalar and fermion** dark matter
- Several couplings and mass ratios

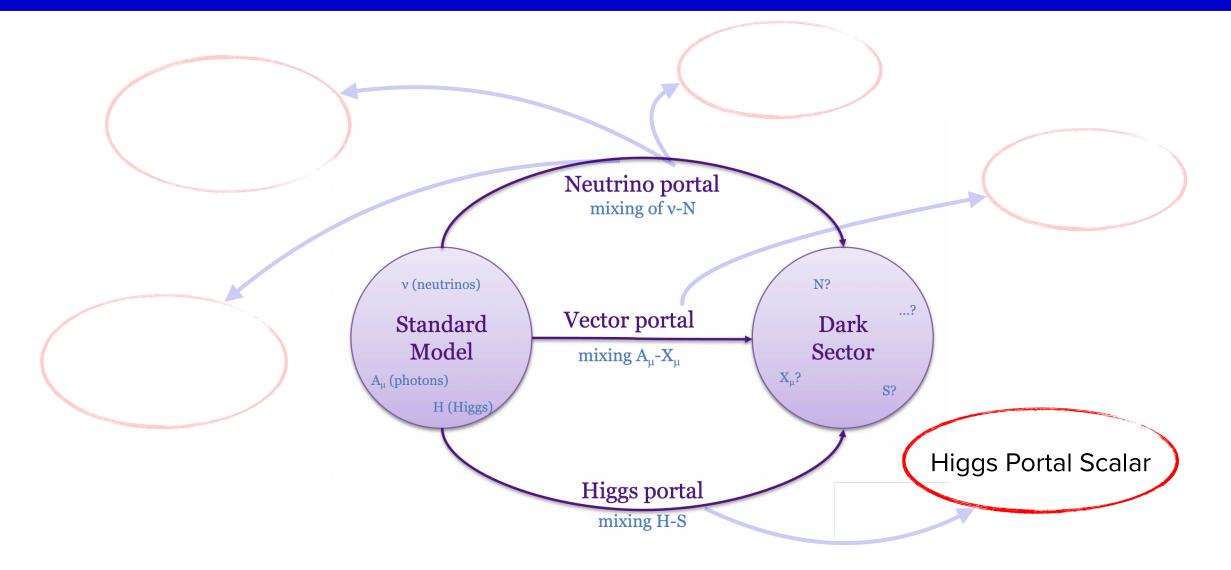


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14

### **Exploring the dark sector**

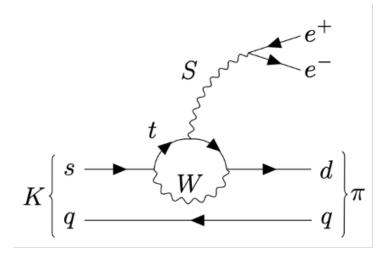


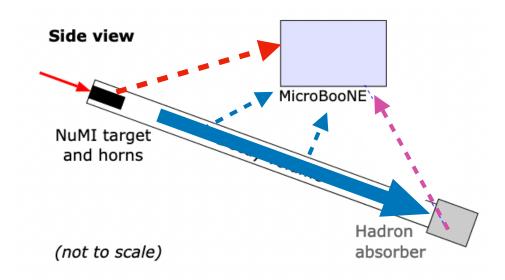
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# **Higgs Portal scalars**

- Neutral scalar singlet S, mixing angle  $\theta$  with the Higgs boson
- Production from **kaon decays**
- Signature:  $e^+e^-$





#### Kaons decaying

- at rest in the NuMI target
- at rest in the NuMI absorber
- In flight

#### IMPERIAL

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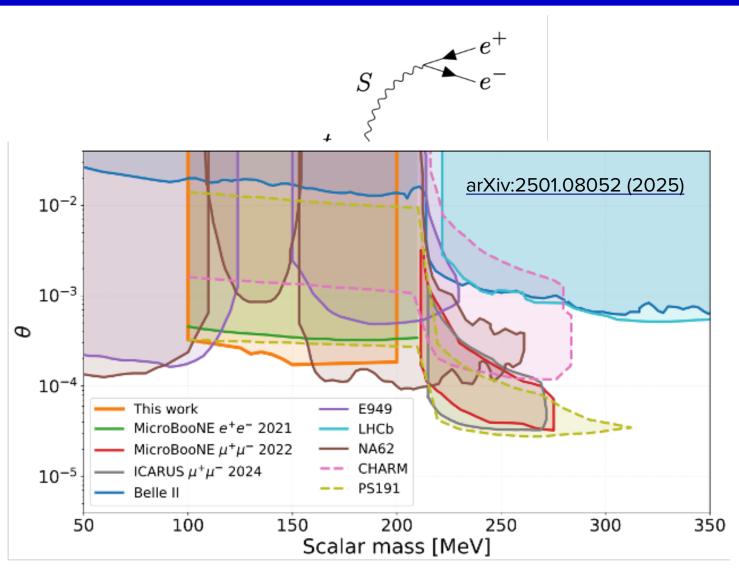
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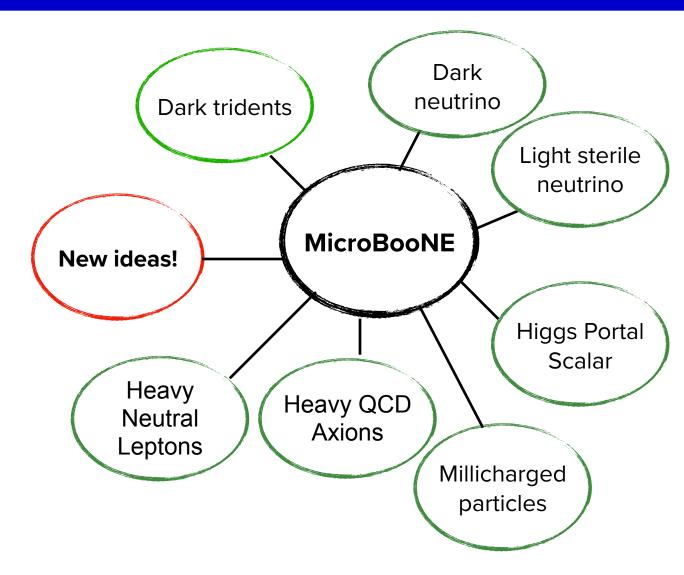


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# Summary

#### MicroBooNE completed 5 years of data taking

- Extensive physics programme including cross-sections and BSM physics
- Pioneering reconstruction techniques
- Exploits data from two neutrino beams (BNB and NuMI)
  - Access to a wide range of dark sector models
- Many new full dataset analyses upcoming!



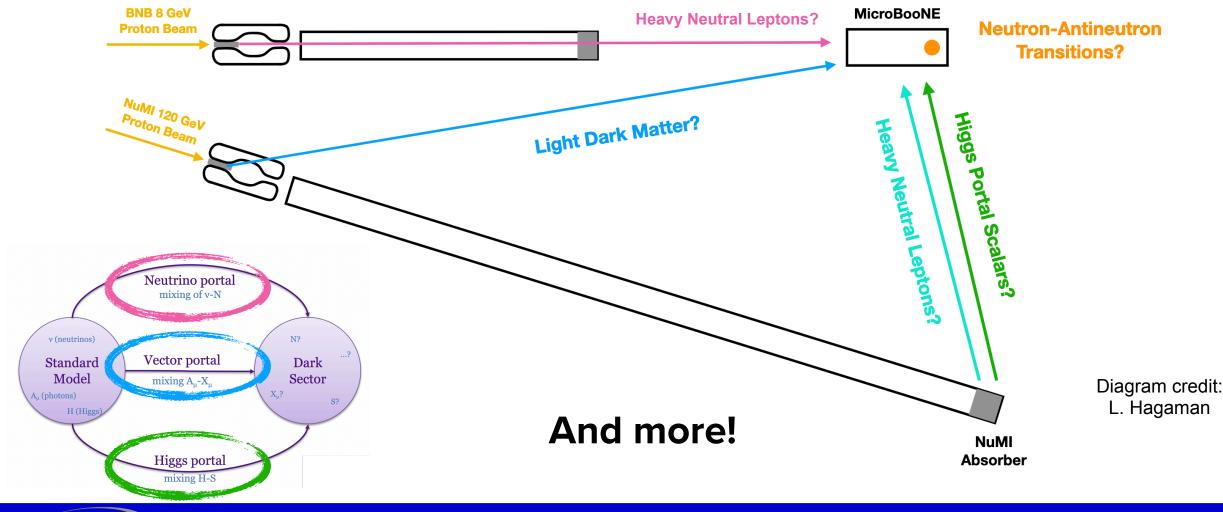
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uBooNE

# **Backup slides**

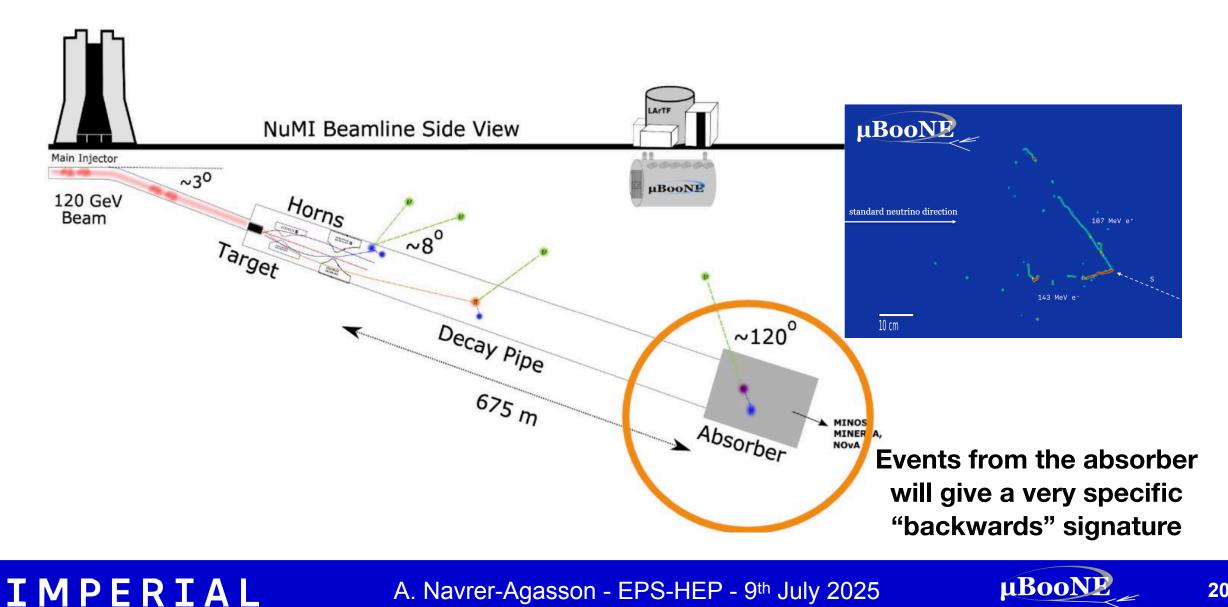
# **BSM Physics @ MicroBooNE**



μBooNE

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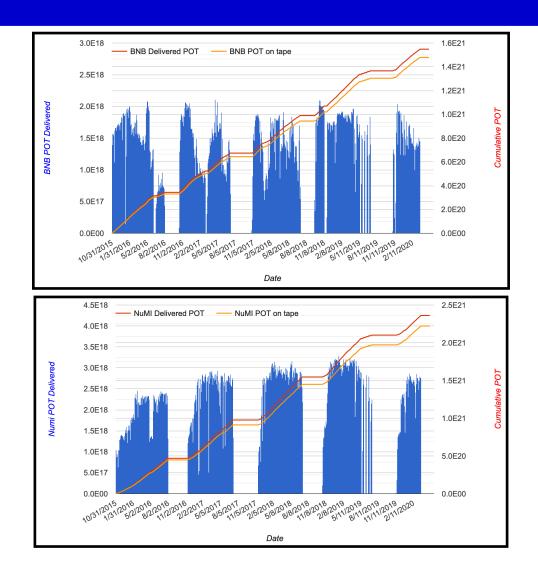
## The NuMI beam absorber





# **Data collected by MicroBooNE**

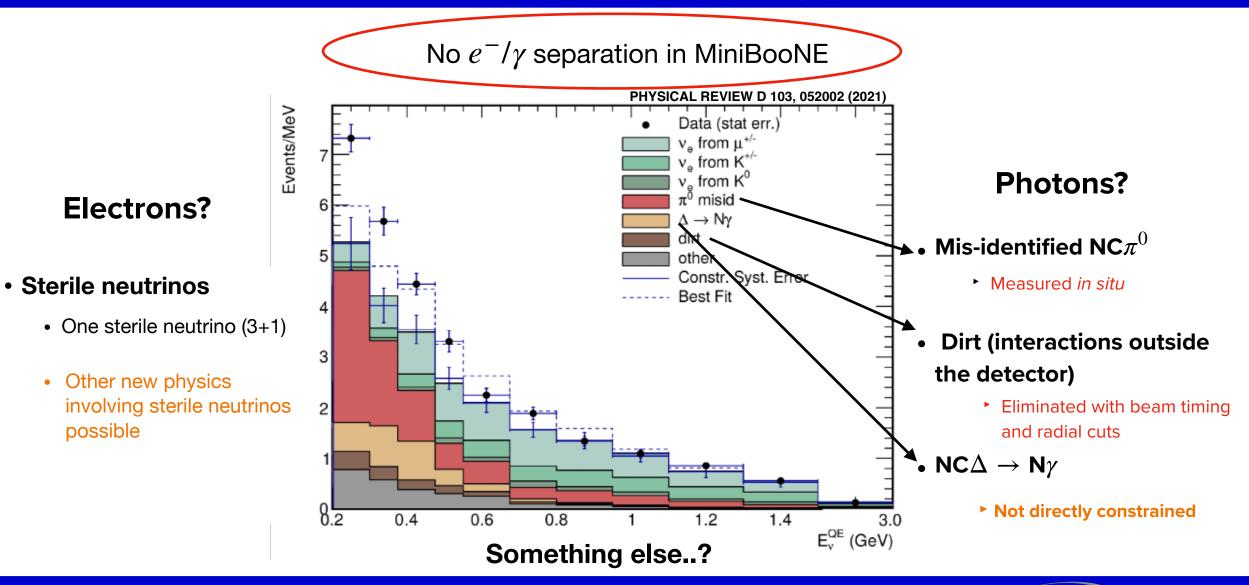
- MicroBooNE collected BNB and NuMI data between
   2015 and 2021 split into five runs
- Full dataset:
  - $1.1 \times 10^{21}$  POT (BNB)
  - ►  $2.37 \times 10^{21}$  POT (NuMI)



\*POT: Protons on Target



# Neutrino portal: low energy excess



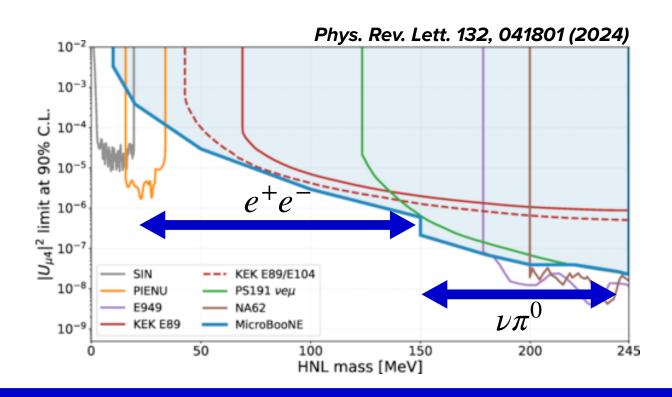
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# **Heavy Neutral Leptons**

- Search for heavy neutral leptons in BNB with mass O(100 MeV)
- Produced by **kaon decays in the beam**, then travel and interact in the detector



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HNL mass	0	105 140	211 <b>245</b>	<u>381</u>	<u>3</u> 493
Production mode				**	MEV
				$\underline{K \to \mu N}$	
					$K \rightarrow eN$
	$N \rightarrow eev$				
		$N \rightarrow \mu$	ev		
Decay mode		N -	$\rightarrow e\pi$		
			$N \rightarrow \mu \mu$	lV	
			$\underline{N}$ –	<u>&gt; μπ</u>	

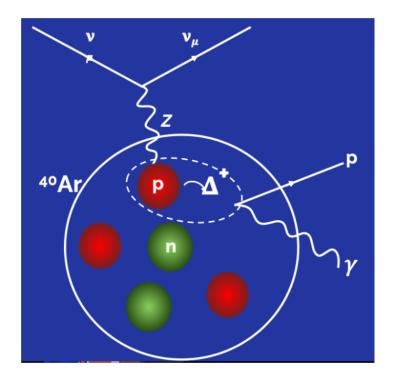
- BDT-based analysis
- **No excess** observed in signal region

**µBooNE** 

Set limits on |U<sub>µ4</sub>|<sup>2</sup> as a function of HNL mass



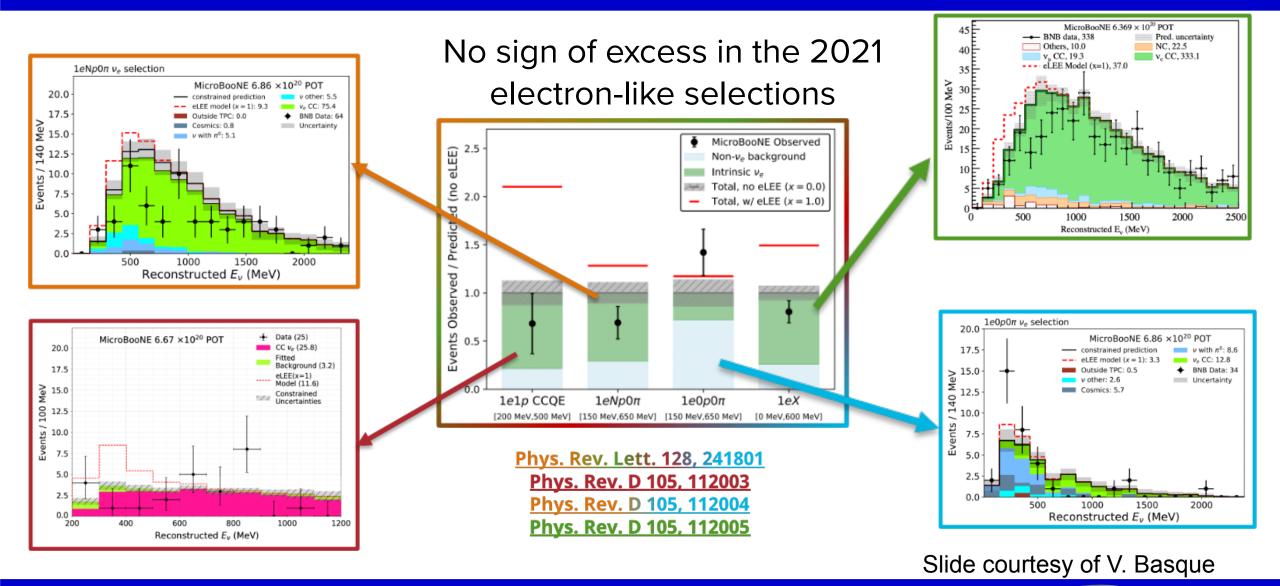
- Rare process: never directly observed
- Physics modelled with
   GENIE
- Needs x3 scaling to match MiniBooNE excess



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## 2021 results



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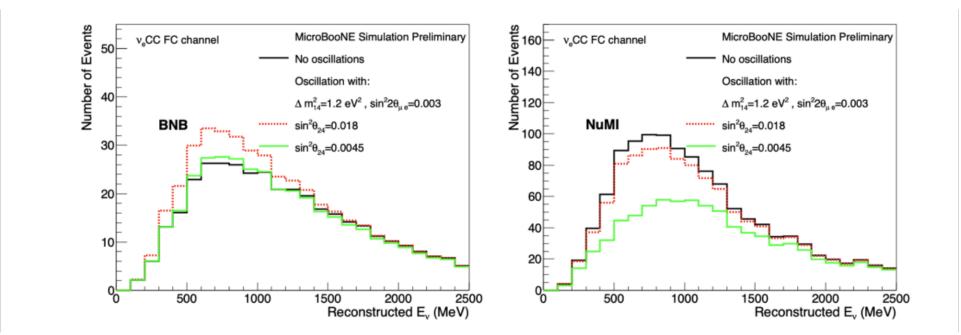
25

## Breaking the degeneracy with two beams

- Oscillation effect in 3+1 model can be hidden in appearance/disappearance degeneracy when using only one beam
  - BNB alone not sensitive to some values of  $heta_{24}$
- Add NuMI data to break it:

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• For the same mixing angles, large effect in NuMI





# **Ongoing searches**

Millicharged particles

- Particles with a fraction of electric charge
- Scatter off atomic electrons and cause "blips" of ionisation in LAr

 $10^{2}$ 

 $\frac{1}{2}$  10<sup>3</sup>

 $10^{2}$ 

 $f_a[\text{GeV}]$ 

• Leverages MeV-scale reconstruction

 $10^{-3}$ 

 $10^{-4}$ 

 $10^{-8}$ 

 $10^{-9}$ 

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NA62

Codominance

 $10^{-2}$ 

 $10^{-1}$ 

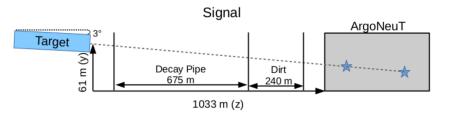
 $m_a[\text{GeV}]$ 

 $10^{0}$ 

 $10^{1}$ 

 $1/f_G[{
m GeV}^{-1}]$ 

 $g_{agg} =$ 



Heavy QCD Axions

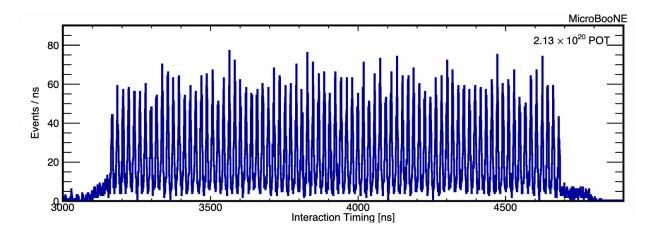
- Axions produced via mixing with neutral mesons
- Decay to **di-photon pairs** in MicroBooNE

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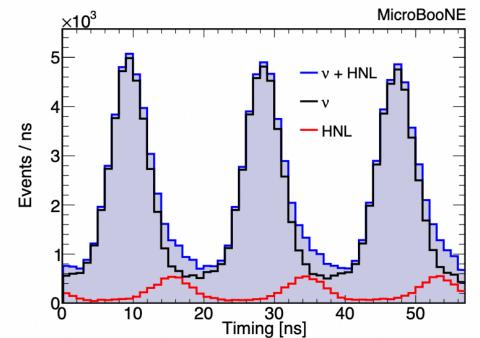


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# **O(1 ns) timing resolution**



- Demonstrated **1.73** ± **0.05** ns resolution
   on neutrino interaction time
- Allows to probe beam structure!



- In-bunch neutrino searches:
  - Reduction of cosmic background
- Out-of-bunch BSM searches
  - Reduction of neutrino background



### The NuMI beam absorber

