



Search for dark sector at BESIII

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On behalf of the BESIII Collaboration

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- SM explains most facts but not all
- More than two dark clouds

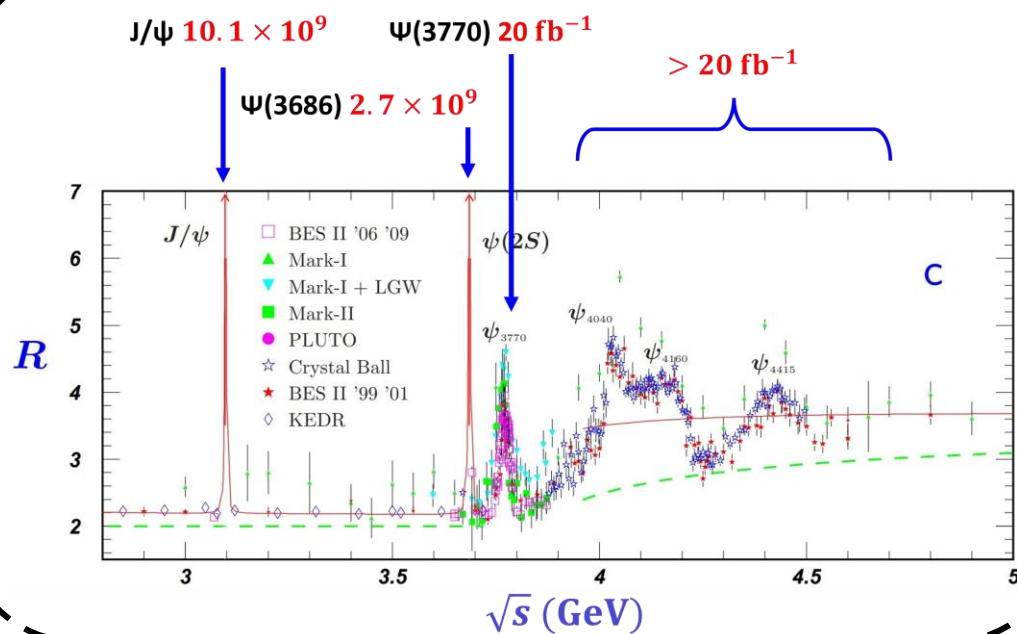
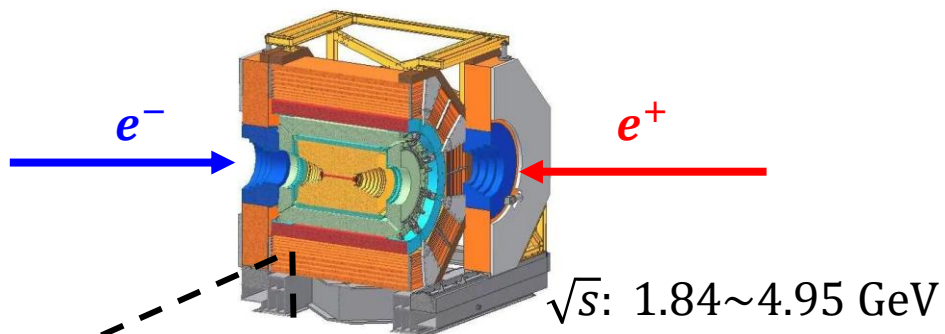
mass →	~2.3 MeV/c ²	~1.275 GeV/c ²	~173.07 GeV/c ²	0	~126 GeV/c ²
charge →	2/3	2/3	2/3	0	0
spin →	1/2	1/2	1/2	1	0
	u up	c charm	t top	g gluon	H Higgs boson
	~4.8 MeV/c ²	~95 MeV/c ²	~4.18 GeV/c ²	0	
	-1/3	-1/3	-1/3	0	
	1/2	1/2	1/2	1	
	d down	s strange	b bottom	γ photon	
	0.511 MeV/c ²	105.7 MeV/c ²	1.777 GeV/c ²	91.2 GeV/c ²	
	-1	-1	-1	0	
	1/2	1/2	1/2	1	
	e electron	μ muon	τ tau	Z Z boson	
	<2.2 eV/c ²	<0.17 MeV/c ²	<15.5 MeV/c ²	80.4 GeV/c ²	
	0	1/2	1/2	±1	
	1/2	1/2	1/2	1	
	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	W W boson	

A balance scale is shown against a light blue background. The scale is tilted, with the left pan (labeled 'matter') being lower and heavier than the right pan (labeled 'antimatter'). The left pan contains several small icons: a worm, a planet, and a cell. The right pan is empty. The scale is supported by a black central pillar with a stepped base.

Potentials to find new physics

This talk: New dark sector results at BESIII

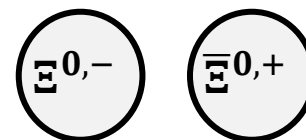
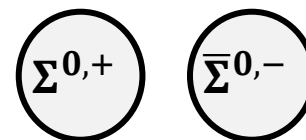
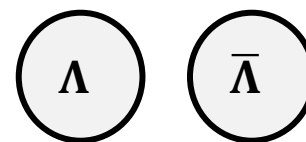
BESIII experiment



2025/7/9

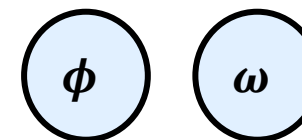
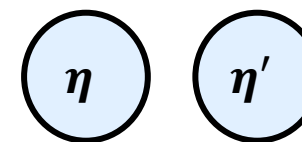
■ $10.1 \times 10^9 J/\psi$ events

• Hyperon



$\sim 10^7$

• Light meson

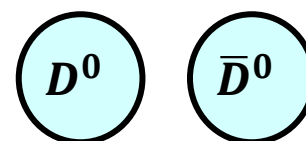


$\sim 10^7$

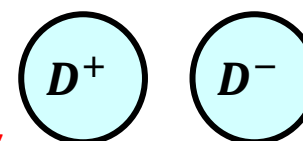


■ $20 \text{ fb}^{-1} \psi(3770)$

■



$\sim 10^7$

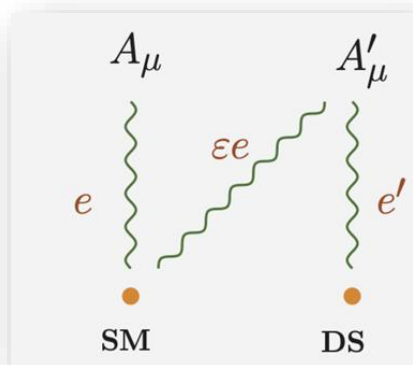
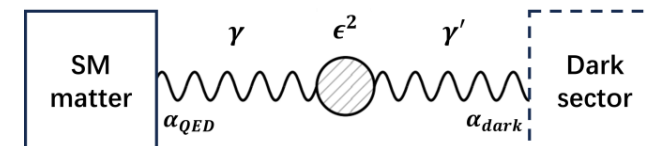


➤ Rich physics with clean background

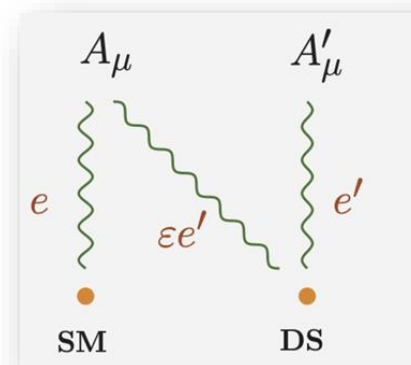
Massless dark photon



Simplest extension of the SM $\Rightarrow$ An extra Abelian gauge group, $U(1)_D \Rightarrow$ dark photon



- **Massive** dark photon
- Coupling with SM fermion
 - Strong constraint



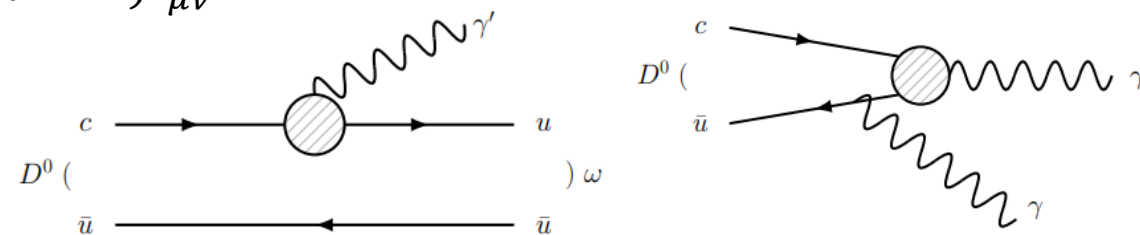
- **Massless** dark photon
- No direct coupling with SM fermion
 - Less constraint
 - Also important role in dark sector

Searching for the massless case can only be in the higher dimension operator:

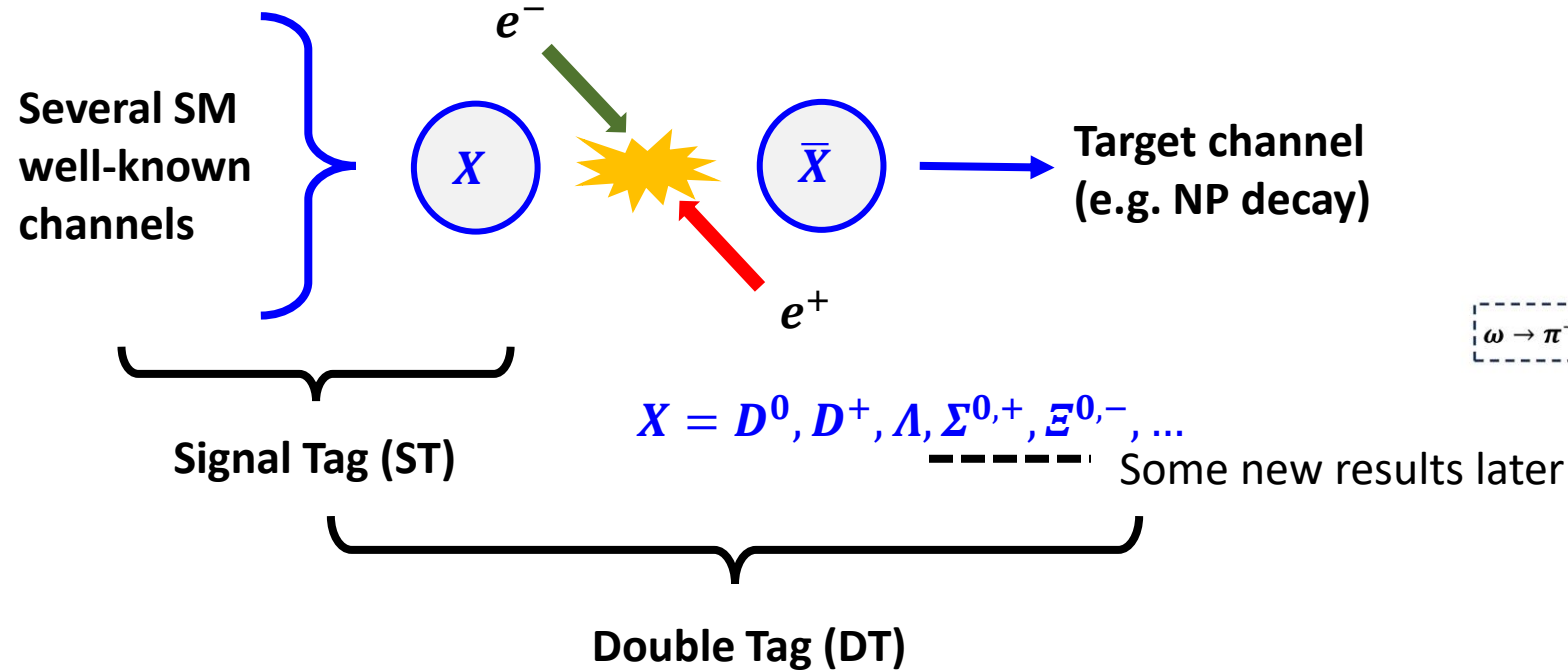
$$\mathcal{L}_{NP} = \frac{1}{\Lambda_{NP}^2} (C_{jk}^U \bar{q}_j \sigma^{\mu\nu} u_k \tilde{H} + C_{jk}^D \bar{q}_j \sigma^{\mu\nu} d_k H + C_{jk}^L \bar{l}_j \sigma^{\mu\nu} e_k H + h.c.) \bar{F}_{\mu\nu}$$

PRL 94, 151802 (2005)

- Naturally allow the **FCNC coupling**
- Less background and higher sensitivity

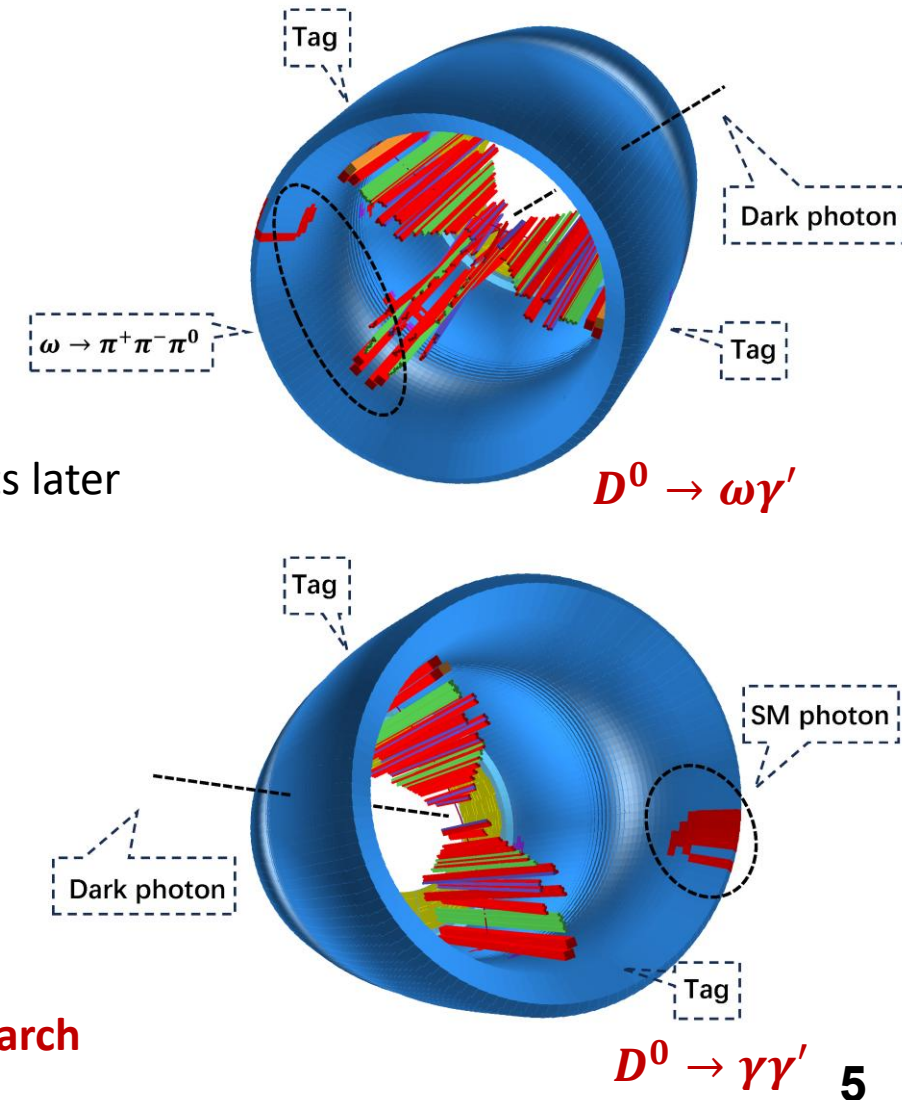


Double tag analysis at BESIII

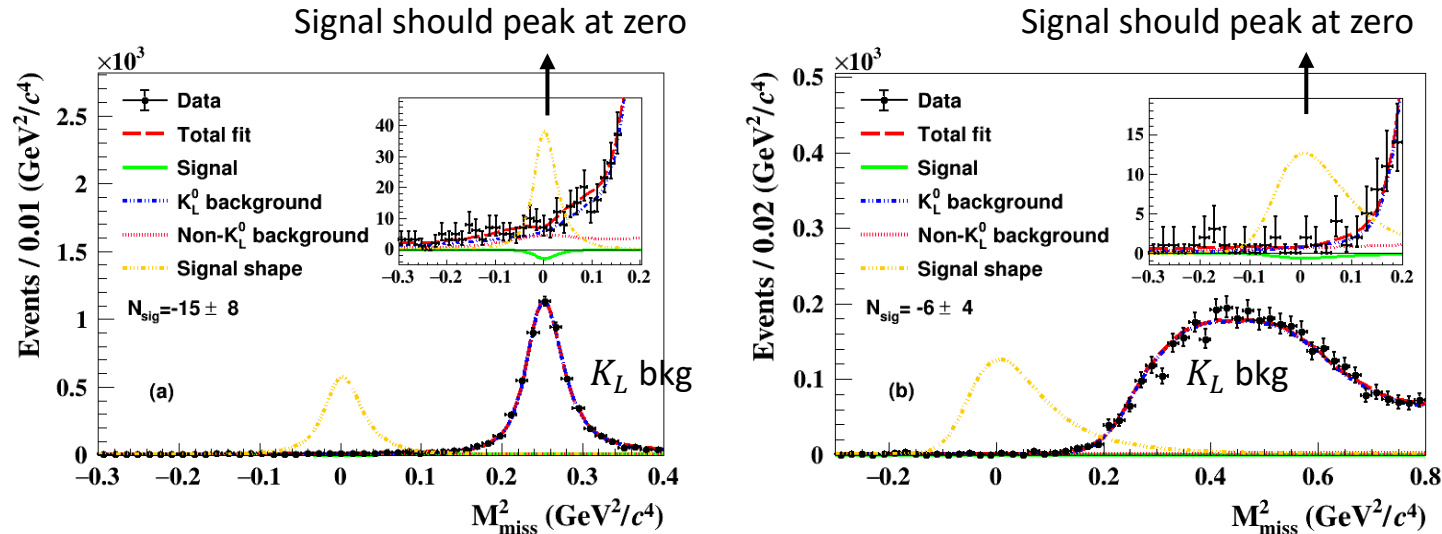


- Measure absolute branching fraction, $\mathcal{B} = \frac{N_{DT}/\epsilon_{DT}}{N_{ST}/\epsilon_{ST}}$
- **Clean background**
- **Recoil** missing particle, such as **invisible particle**
- Cancel systematic uncertainty

Active in dark sector search



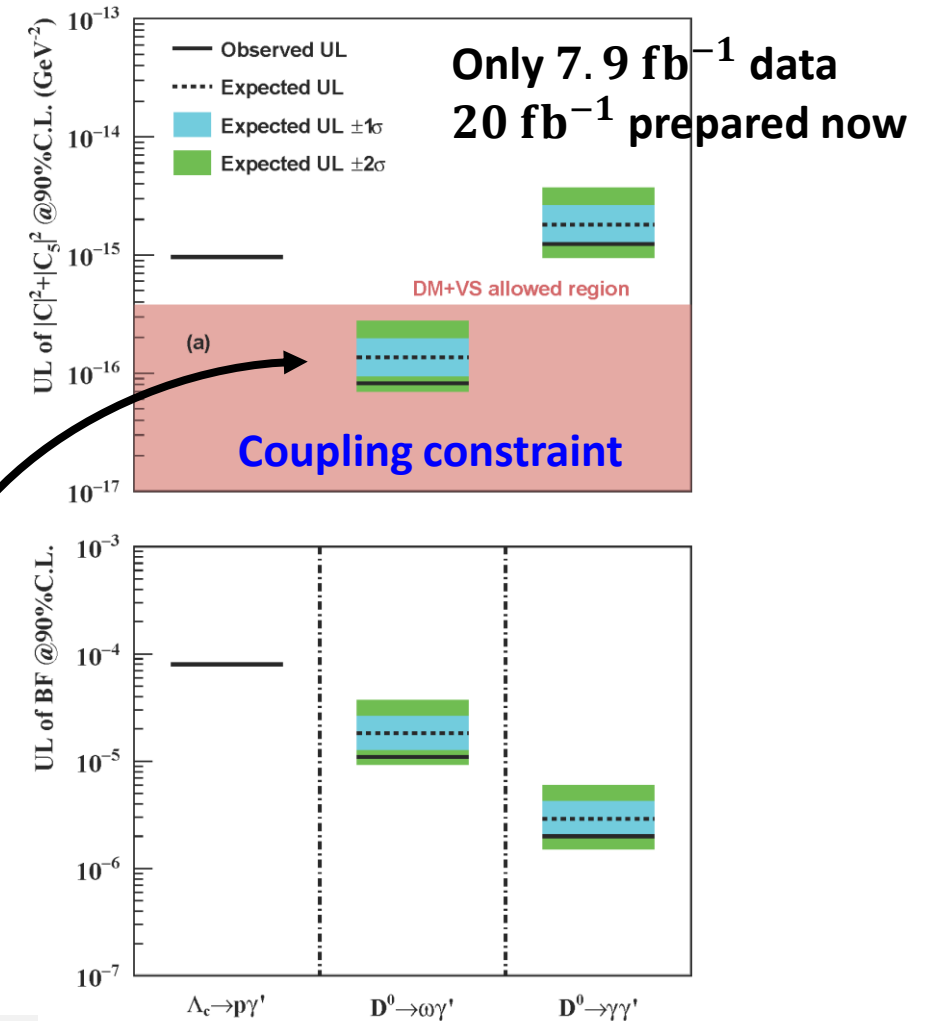
Search for $D^0 \rightarrow \omega\gamma'$ and $D^0 \rightarrow \gamma\gamma'$ at BESIII



Signal extraction of $D^0 \rightarrow \omega\gamma'$

Signal extraction of $D^0 \rightarrow \gamma\gamma'$

- **First search** for massless dark photon in D meson decay
- The constraint from $D^0 \rightarrow \omega\gamma'$ goes into the dark matter (DM) and vacuum stability (VS) allowed region for the first time, **improved by more than 1 order**



K_S invisible decay



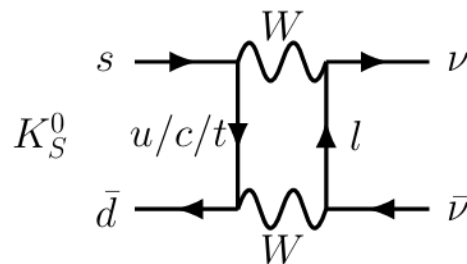
Invisible
decay of K_S^0

SM decay

Decay to DM

Ordinary-
mirror particle
oscillations

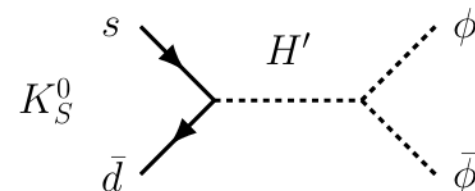
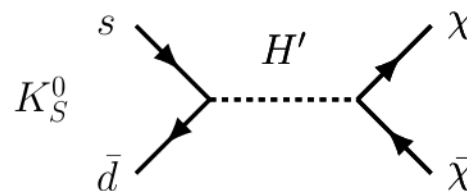
Input for
CPT test



$$\text{BF} < 10^{-16}$$

FCNC & Helicity suppression

Phys.Rev.D 91 (2015) 1, 015004



2HDM model

$$\text{BF} \sim \mathcal{O}(10^{-6})$$

Natural Sci.Rev. 1 (2024) 5

$$K_S^0 \rightsquigarrow K_S^{0'}$$

Mirror matter model

$$\text{BF} \sim \mathcal{O}(10^{-6})$$

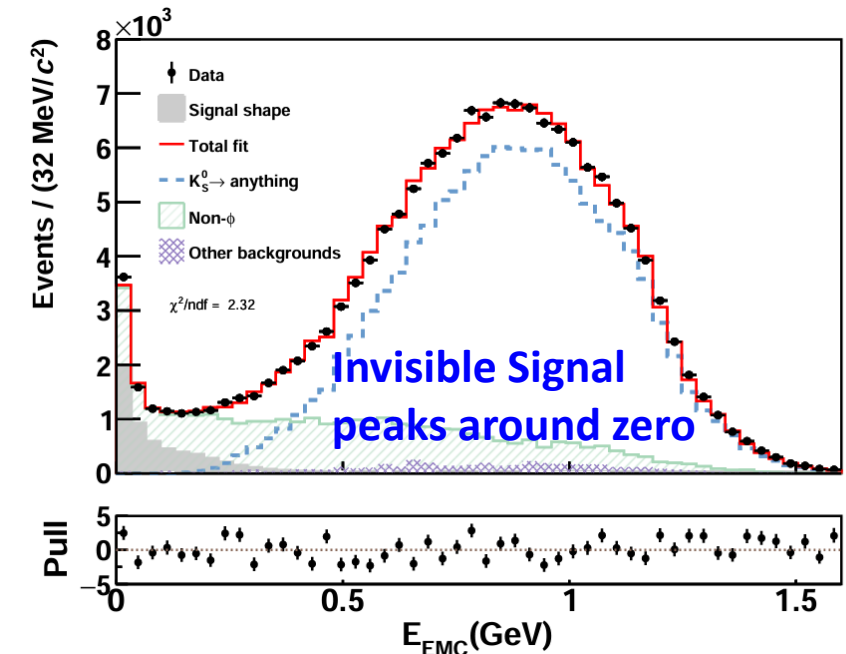
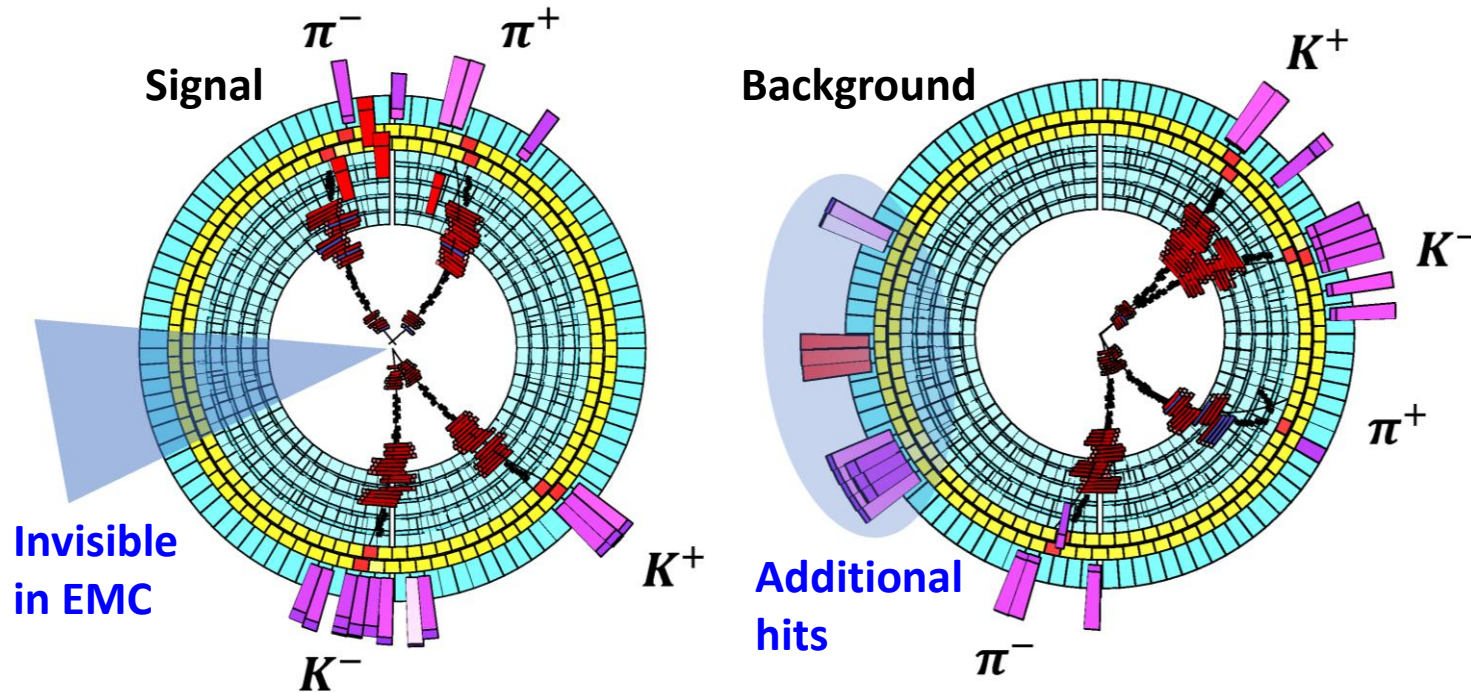
arXiv: 2006.10746

Bell-Steinberger relation **connects CPTV** to the amplitudes of all decay channels of neutral kaons.

BUT currently assumes no invisible modes

Phys.Rev.D 91 (2015) 1, 015004

Search for $K_S \rightarrow \text{invisible}$ at BESIII



- Using the deposited energy in EMC to identify the invisible signal
- An alternative method for invisible signal search at BESIII

- $J/\psi \rightarrow \phi K_S^0 K_S^0$ from $10^{10} J/\psi$
- $N_{\text{sig}} = 56 \pm 201$
- $\mathcal{B}(K_S^0 \rightarrow \text{invisible}) < 8.4 \times 10^{-4}$
@90% C.L.

First direct measurement of $K_S^0 \rightarrow \text{invisible}$; the UL still lies above the NP prediction

Dark baryon



Coincidence
problem

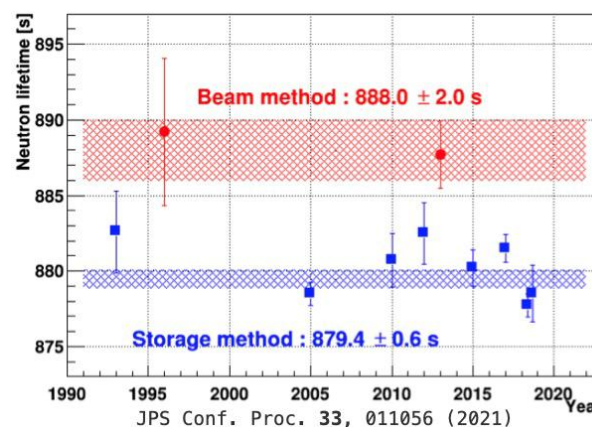


The baryon and dark matter
energy densities are fairly similar
 $\Omega_{DM}/\Omega_b \cong 5.36 \pm 0.06$

- Potential connection between their origins
- **DM may have non-zero baryon number**



Neutron
lifetime puzzle



$$\tau_n^{\text{beam}} = \frac{\tau_n^{\text{bottle}}}{\text{Br}(n \rightarrow p + \text{anything})}$$

$$\mathcal{B}(n \rightarrow \text{dark}) \sim 1\%$$

- **Motivates the existence of dark baryon**

B-Mesogenesis
mechanism



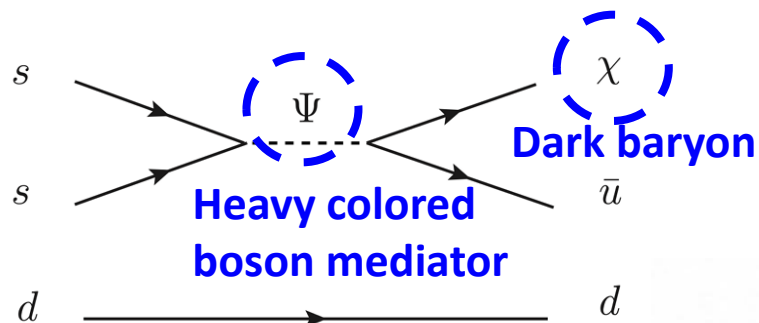
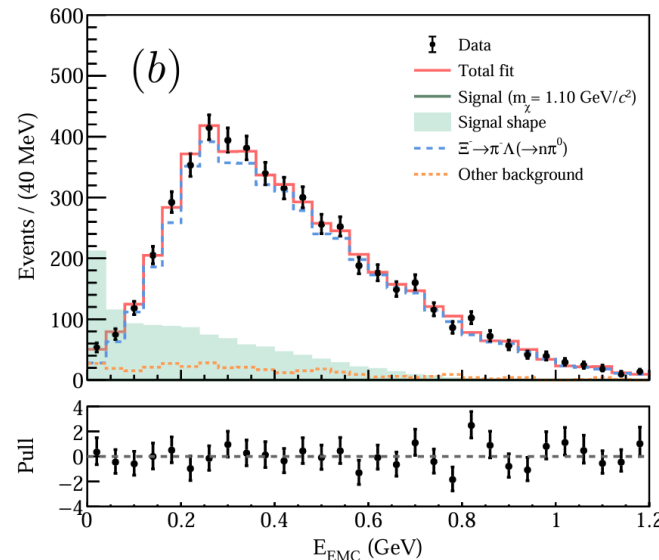
CP asymmetry
in $B \rightarrow \text{baryon} + \text{dark baryon}$

- Explain matter anti-matter asymmetry
- The origin of the dark matter

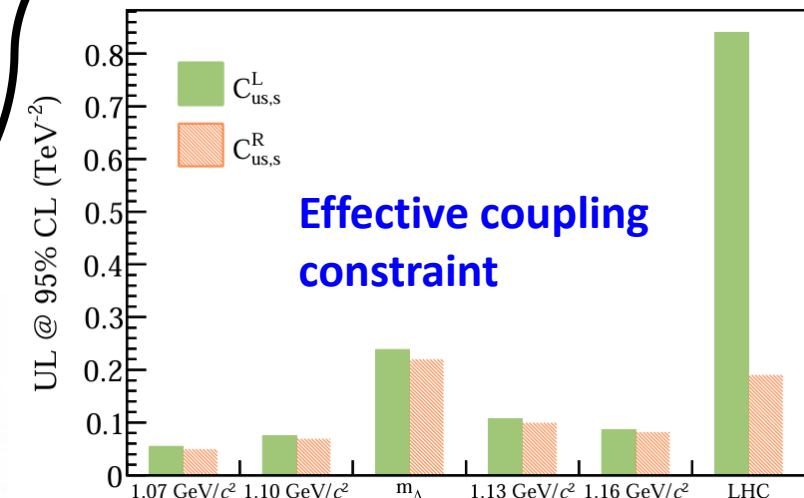
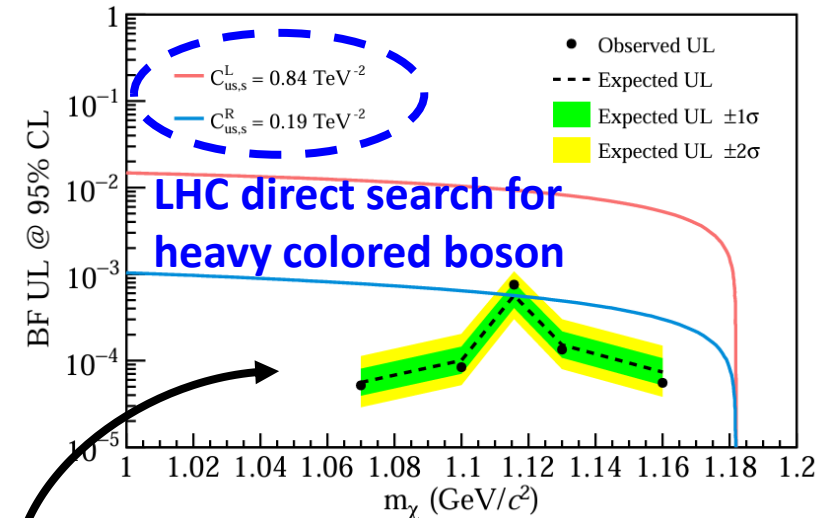
Naturally, dark baryon interacts with all SM quark flavor
 $\Rightarrow$ **Search for dark baryon in hyperon decay at BESIII**

Search for $\Xi^- \rightarrow \pi^- \chi$ at BESIII

- $J/\psi \rightarrow \Xi^+ \Xi^-$ from $10^{10} J/\psi$
- Double tag method
- Dark baryon χ with mass hypothesis of 1.07, 1.10, m_Λ , 1.13, 1.16 GeV
- Deposited energy in EMC used to extract the invisible signal

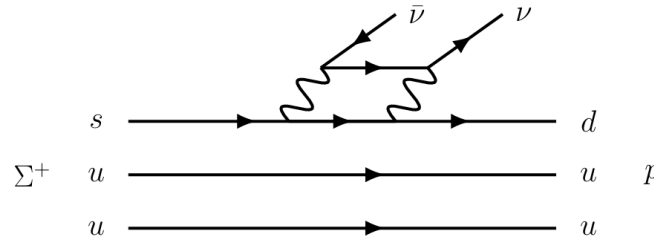


• Better effective coupling constraints



$\Sigma^+ \rightarrow p + \text{invisible and QCD axion}$

SM decay



- $s \rightarrow d\nu\bar{\nu}$, FCNC & GIM suppression
- $\text{BF} < 10^{-11}$

Phys.Rev.D 94 (2016) 11, 115013

Decay to BSM particles

- **Solution to strong CP problem**

➤ Naturally allow the FCNC coupling

- An excellent **dark matter**

$$\mathcal{L}_{a-f} = \partial_\mu a \bar{f}_i \gamma^\mu \left(\frac{1}{F_{ij}^V} + \frac{\gamma^5}{F_{ij}^A} \right) f_j$$

- $m_a \sim g \sim \frac{1}{F_a} \Rightarrow m_a \ll 1 \text{ eV};$
 g : coupling strength;
 F_a : decay constant

Vector coupling

Axial coupling

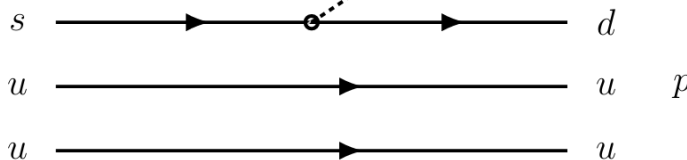
sensitive

$K^+ \rightarrow \pi^+ a$

Both sensitive

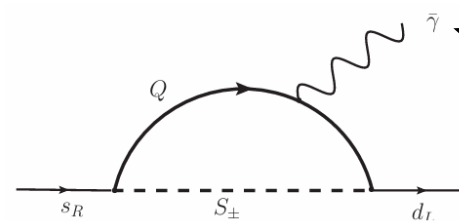
Hyperon decay

Decay to QCD axion



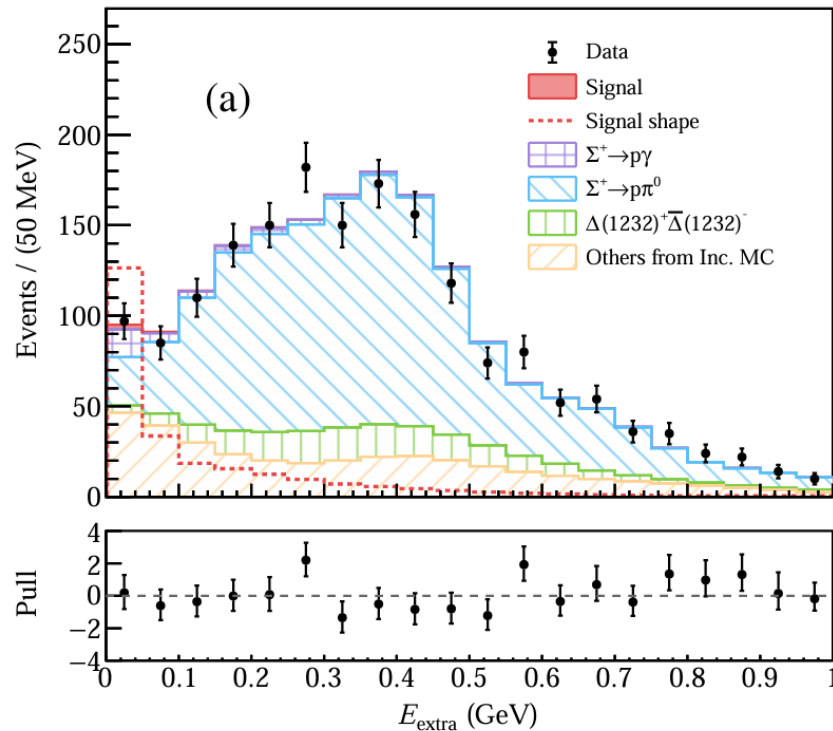
- Can also **Decay to massless dark photon**

- Maximum allowed $\text{BF} \sim 3.8 \times 10^{-5}$

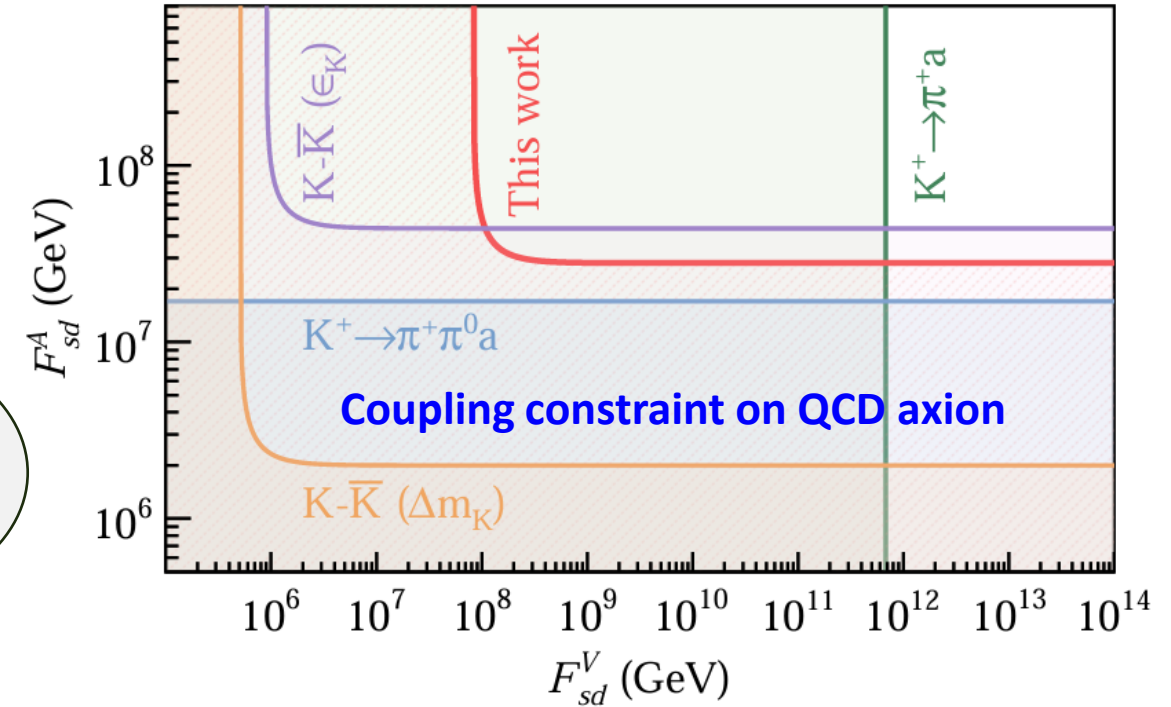


Search for $\Sigma^+ \rightarrow p + \text{invisible}$ at BESIII

- $J/\psi \rightarrow \Sigma^+ \bar{\Sigma}^-$ from $10^{10} J/\psi$
- Double tag method
- Invisible particle with mass hypothesis of zero
- Deposited energy in EMC used to extract the signal

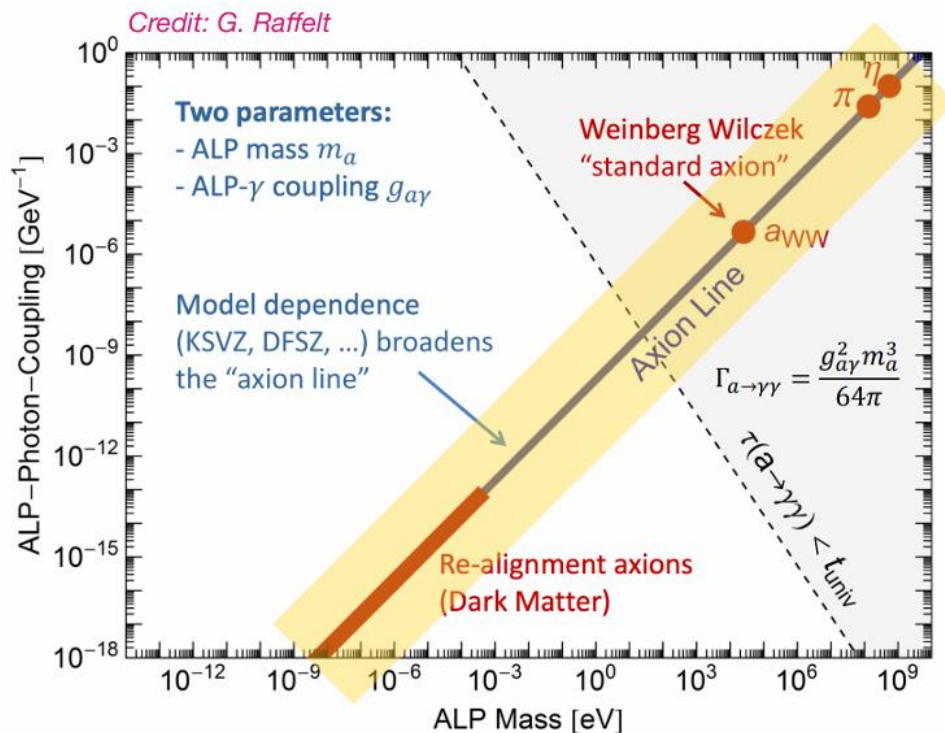


$\mathcal{B}(\Sigma^+ \rightarrow pa)$
 $< 3.2 \times 10^{-5}$
 @90% C.L.



- Competitive constraint on the axial coupling temp F_{sd}^A of QCD axion
- The BF UL lies below the maximum allowed BF of massless dark photon decay (3.8×10^{-5})

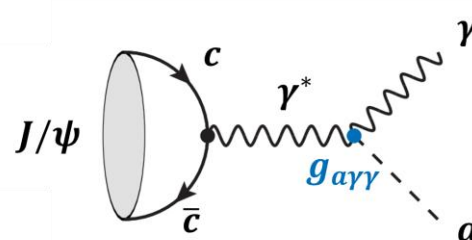
Axion-like particle



- **Axion-like particle (ALP)**
 - Similar to QCD axion
 - **arbitrary masses and couplings**

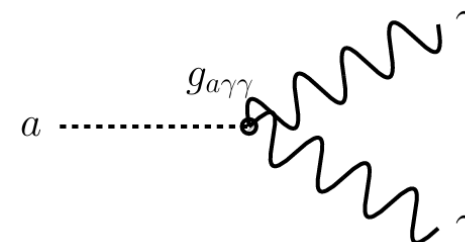
➤ Considering the ALP-photon interaction

- $\mathcal{L} \supset -\frac{1}{4} \mathbf{g_{a\gamma\gamma}} a F^{\mu\nu} \tilde{F}_{\mu\nu}$ JHEP 06 (2019) 091



- **Produced from heavy photon**

$$\begin{aligned} & \frac{\mathcal{B}(J/\psi \rightarrow \gamma a)}{\mathcal{B}(J/\psi \rightarrow e^+ e^-)} \\ &= \frac{m_{J/\psi}^2}{32\pi\alpha} g_{a\gamma\gamma}^2 \left(1 - \frac{m_a^2}{m_{J/\psi}^2}\right)^3 \\ &\Rightarrow \mathbf{BF} \sim \mathbf{g_{a\gamma\gamma}^2} \end{aligned}$$

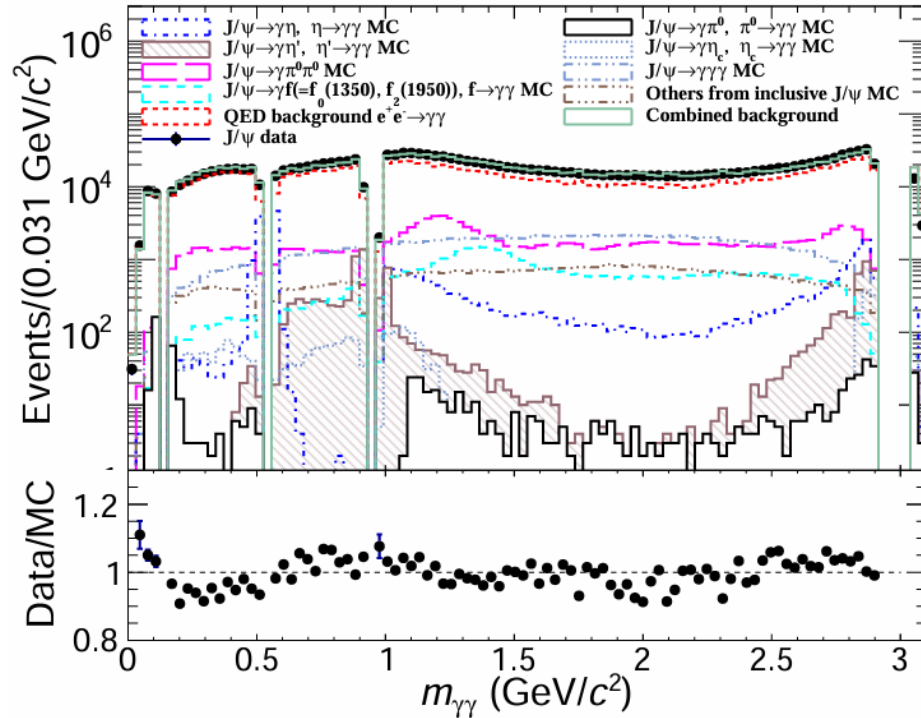


- **Decay to di-photons**

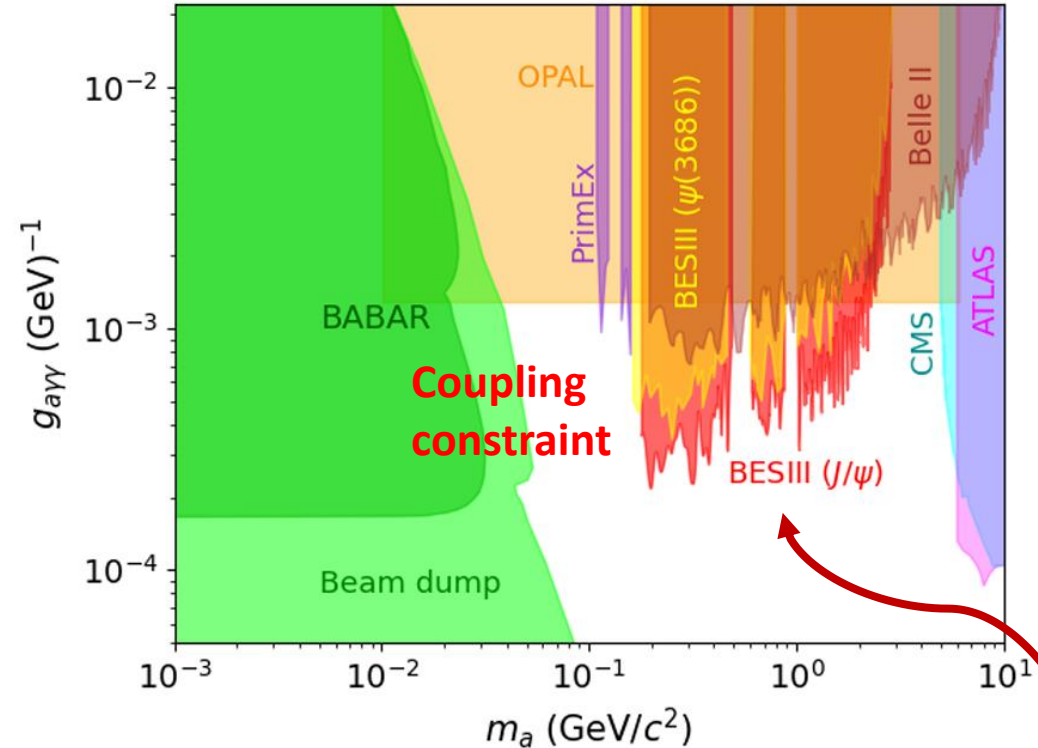
$$\begin{aligned} \Gamma_{a \rightarrow \gamma\gamma} &= \frac{g_{a\gamma\gamma}^2 m_a^3}{64\pi} \\ g_{a\gamma\gamma} &\sim 10^{-4} \text{ GeV}^{-1}, m_a \sim \text{GeV} \\ &\Rightarrow \mathbf{short-lived} \end{aligned}$$

New searches for ALP at BESIII: $J/\psi \rightarrow \gamma a, a \rightarrow \gamma\gamma$

Search for $J/\psi \rightarrow \gamma a_{\gamma\gamma}$ at BESIII



- 10^{10} J/ψ events
- Extract signal from $M_{\gamma\gamma}$ distribution
- Maximum signal significance: $< 3\sigma$

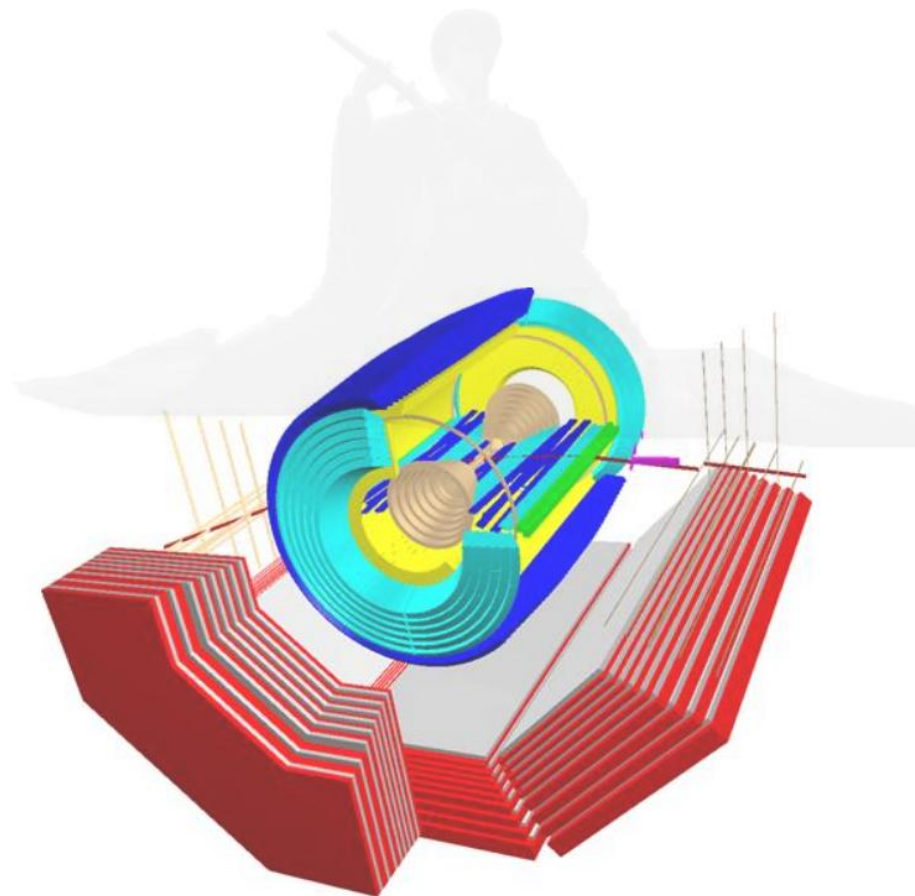


- UL on the BF of $\mathcal{B}(J/\psi \rightarrow \gamma a) \times \mathcal{B}(a \rightarrow \gamma\gamma)$
 $(3.6 \sim 53.1) \times 10^{-8}$ @ 90% C.L.
- New stringent constraints on ALP-photon coupling for $0.18 \leq m_a \leq 2.85$ GeV

Summary



- New results of **dark photon, dark baryon, axion** and **K_S^0 invisible decay** at BESIII
- **No evidence** and **New stringent constraint on NP**
- BESIII has collected $10^{10} J/\psi$, $2.7 \times 10^9 \psi'$, 20 fb^{-1} @ 3.77 GeV data ($D\bar{D}$) and more...
- **More & Better** results are coming soon



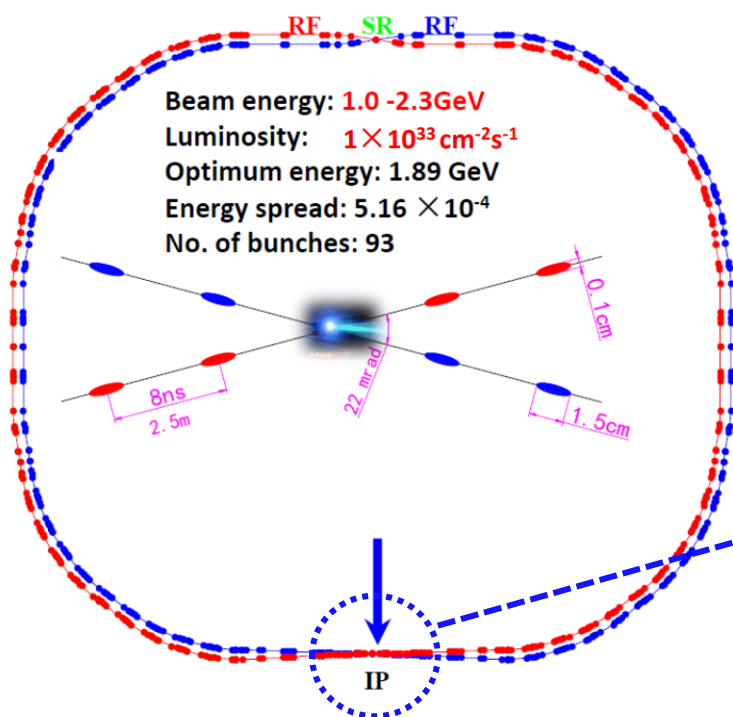
The future of **Dark Sector** is **Bright** !

THANK YOU



Appendix

BEPCH and BESIII

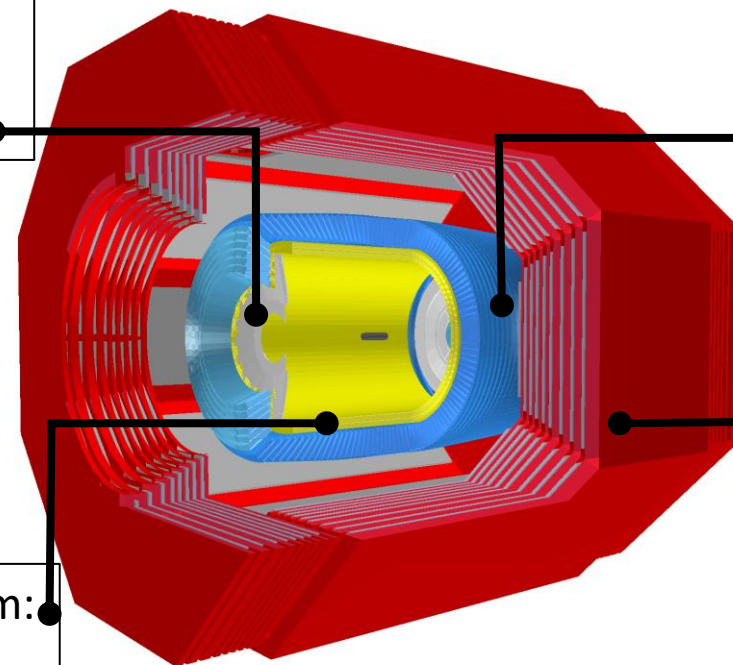


✓ **Beijing Electron
Positron Collider II**

Multilayer **D**rift
Chamber:
 $\Delta p/p = 0.5\%$ @ 1 GeV
 $dE/dx \sim 6\%$

Time-**O**f-**F**light system:
 $\sigma T = 68(60) \text{ ps}$
barrel(endcap)

BESIII



Electro**M**agnetic
Calorimeter:
CsI(Tl) 2.5%(5.0%)
Barrel(endcap)
@ 1 GeV

MUon **C**hamber:
RPC 9(8) layers
barrel(endcap)

✓ **BE**ijing
Spectrometer **III**

BESIII has collected large data samples in $\tau - c$ energy region
which can benefit the search for the hidden dark sector

Massless dark photon



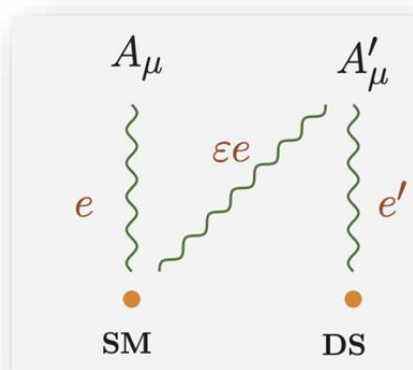
Massive dark photon

Symmetry broken spontaneously

$$A^\mu \rightarrow A^\mu + \epsilon A'^\mu$$

$$\mathcal{L} = eJ_\mu A^\mu + \underline{\epsilon e J_\mu A'^\mu} + e' J'_\mu A'^\mu$$

- Dark photon couples to the SM matter



Dark photon

Kinetic mixing: $\frac{\epsilon}{2} F'_{\mu\nu} F^{\mu\nu}$
 ϵ : mixing strength



SM photon

Massless dark photon

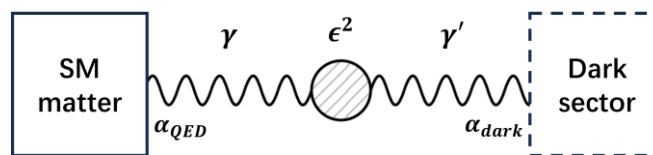
Symmetry remains unbroken

PRL 94, 151802 (2005)

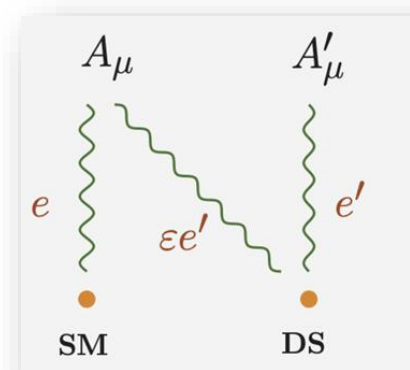
$$A'^\mu \rightarrow A'^\mu + \epsilon A^\mu$$

$$\mathcal{L} = eJ_\mu A^\mu + \underline{\epsilon' e' J'_\mu A^\mu} + e' J'_\mu A'^\mu$$

- SM photon couples to the dark sector particles



A portal to connect the SM matter and the dark sector

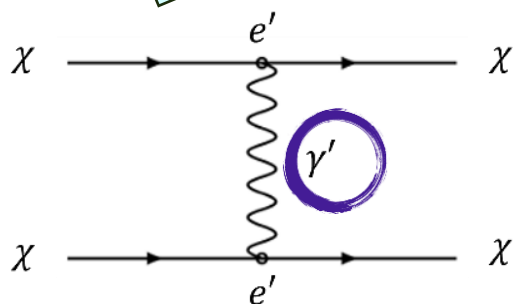


Massless dark photon has no interaction with the SM matter in the dimension-4 operator

Role of the massless dark photon



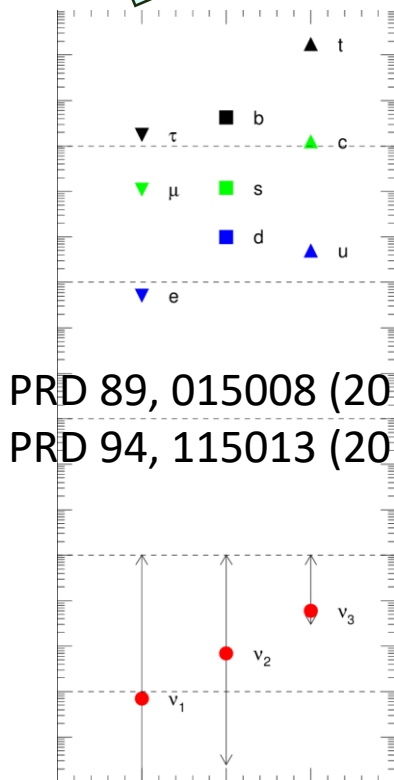
Provide a new
long-range
force of the DM



→ Explain the galaxy
formation and dynamics

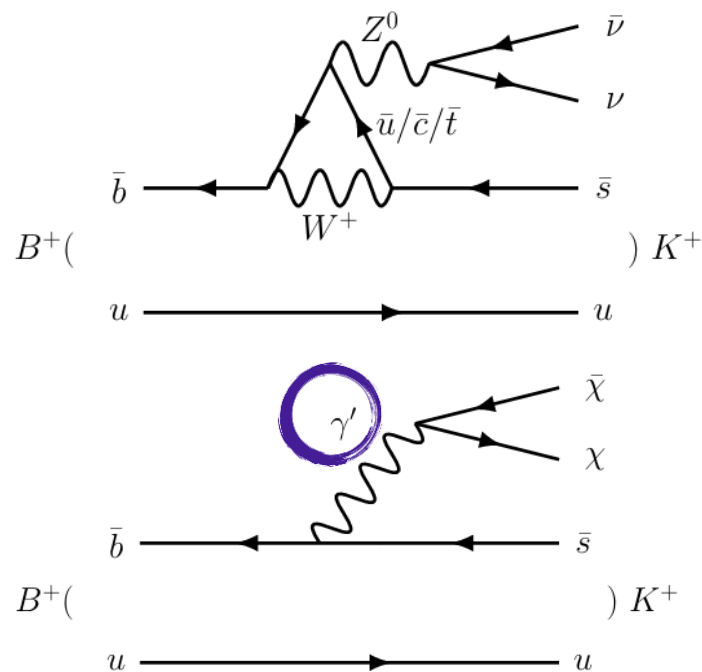
Phys.Rev.D 79 (2009) 023519
Phys.Rev.D 91 (2015) 023512
Phys.Lett.B 749 (2015) 236-241
JCAP 05, 022 (2017)
Phys.Rev.D 102 (2020) 8, 083009

Solution to the
Fermion mass
hierarchy



PRD 89, 015008 (2014)
PRD 94, 115013 (2016)

Explain excess of
 $B^+ \rightarrow K^+ \nu \bar{\nu}$
from Belle II



Phys.Rev.D 109 (2024) 11, 112006
Eur.Phys.J.C 84 (2024) 5, 460

More...

- Solution to the origin
of the CKM matrix
structure

- Solution to the
vacuum instability
problem in SM Higgs
Sector

- ...

PRD 101 (2020) 7, 075019
JHEP 01 (2022) 142



The interaction of γ'

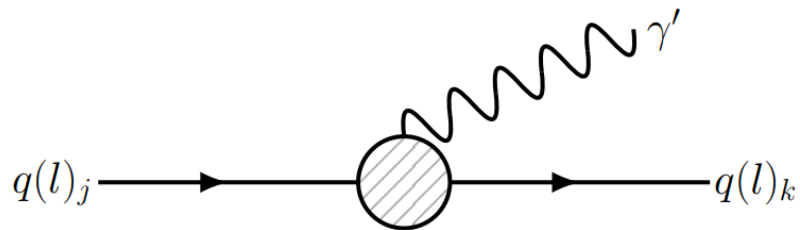
$$\mathcal{L} = eJ_\mu A^\mu + e'\epsilon J'_\mu A^\mu + e'J'_\mu A'^\mu \text{ (no interaction between } \gamma' \text{ and SM matter)}$$

PRL 94, 151802 (2005)

$$\mathcal{L}_{NP} = \frac{1}{\Lambda_{NP}^2} (\underbrace{C_{jk}^U \bar{q}_j \sigma^{\mu\nu} u_k \tilde{H}}_{\text{Up type quarks coupling}} + \underbrace{C_{jk}^D \bar{q}_j \sigma^{\mu\nu} d_k H}_{\text{Down type quarks coupling}} + \underbrace{C_{jk}^L \bar{l}_j \sigma^{\mu\nu} e_k H}_{\text{Charged leptons coupling}} + h.c.) \bar{F}_{\mu\nu}$$

Dimension-six operator Massless dark photon

Λ_{NP} : New physics energy scale
 C_{jk}^i : Dimensionless coefficient



Experimental search:

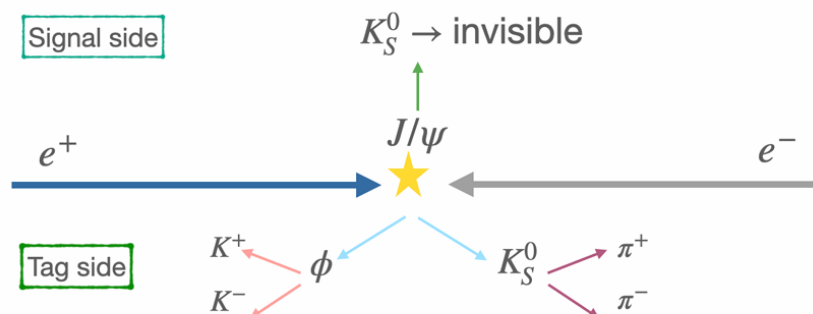
- $H \rightarrow \gamma\gamma'$
 - $\mu \rightarrow e\gamma'$
 - $\Lambda_c^+ \rightarrow p\gamma'$
- } • No signal observed
- The sensitivity still lie outside the theoretically predicted allowed region

Analysis strategy



K_S^0 source

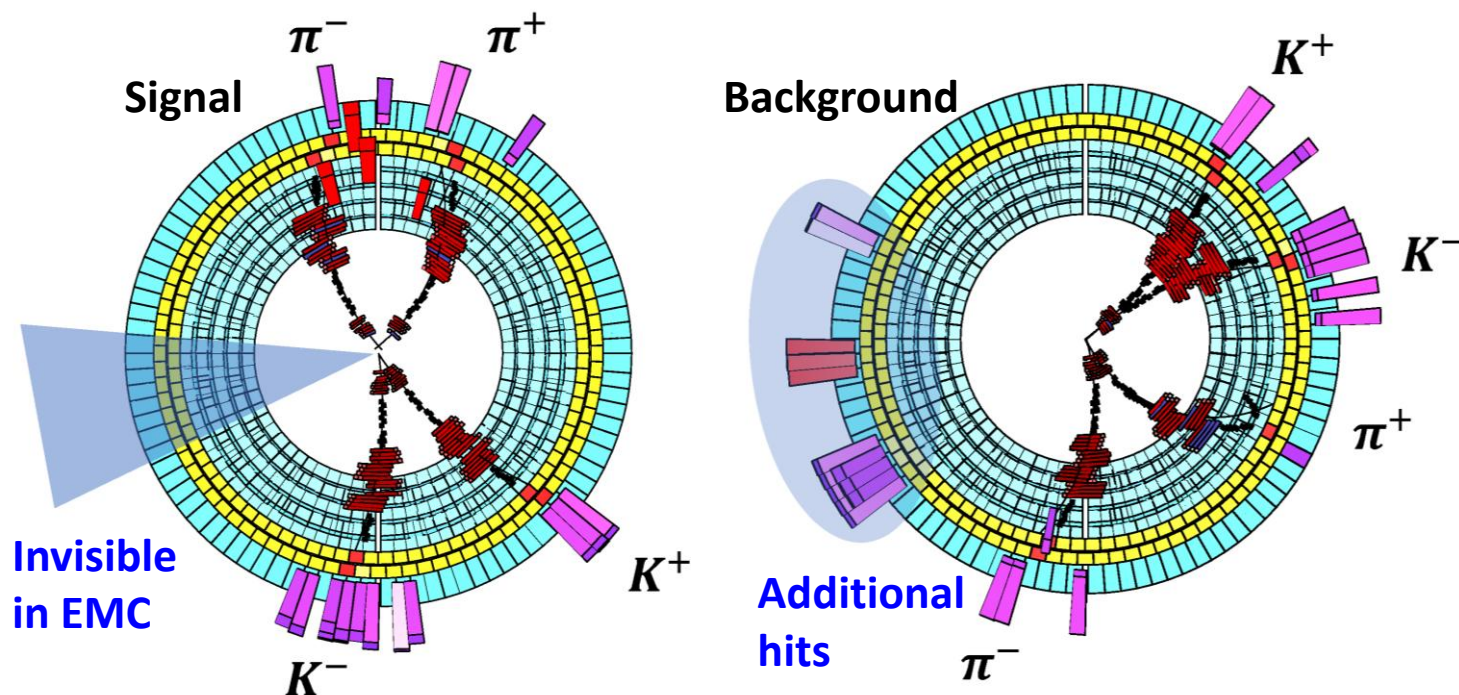
$J/\psi \rightarrow \phi K_S^0 K_S^0$ from $(10087 \pm 44) \times 10^6 J/\psi$



Why this channel?

- $J/\psi \rightarrow \phi K_S^0 K_L^0$ is forbidden with C Parity conservation, **lower background**
- But still have $J/\psi \rightarrow K^+ K^- K_S^0 K_L^0$

Signal feature



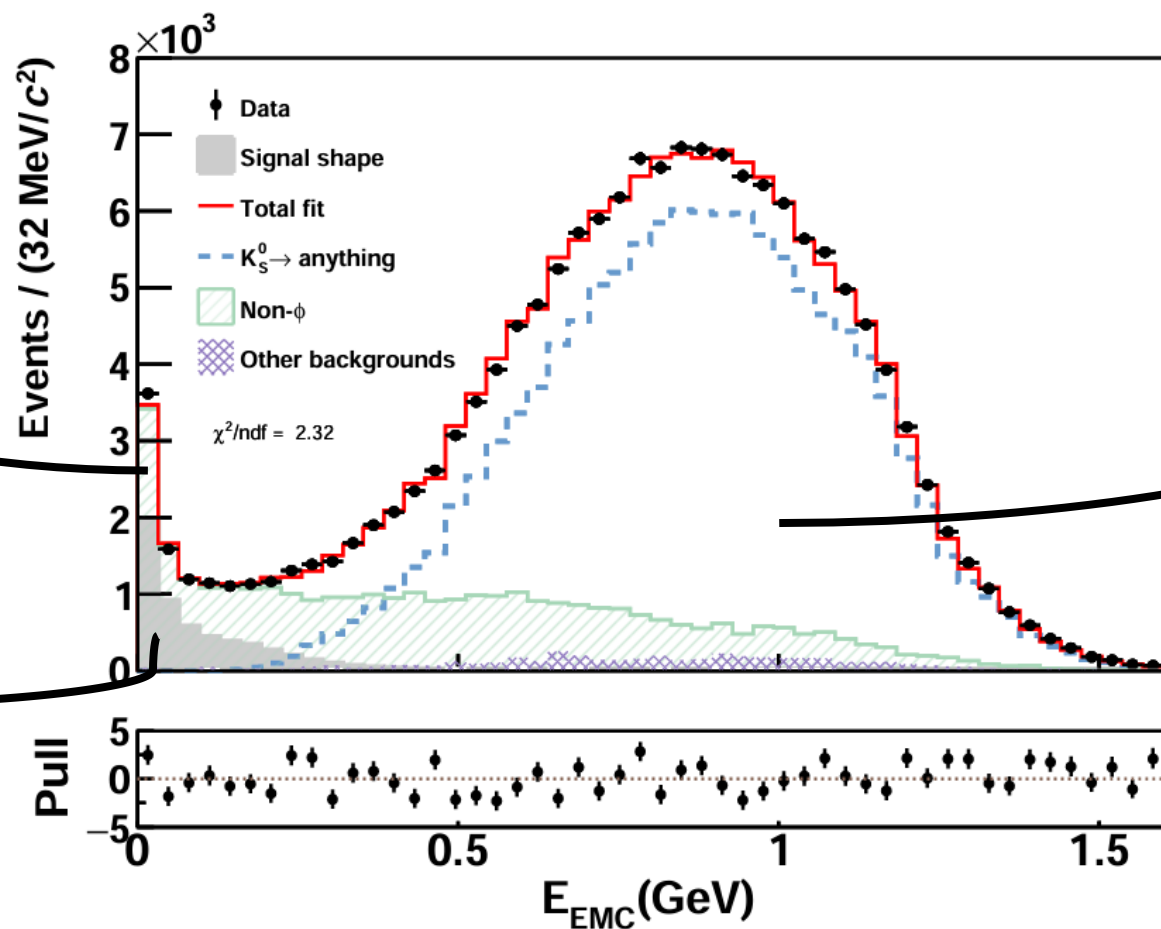
Using the deposited energy in EMC to identify the invisible signal

Signal extraction



$J/\psi \rightarrow K^+ K^- K_S^0 K_L^0$
peaking background
shape from ϕ sideband

Invisible Signal
peaks around zero



Other background modeled
with MC simulation, such as
 $K_S^0 \rightarrow \pi^0 \pi^0$

$N_{\text{sig}} = 56 \pm 201$
 $\mathcal{B}(K_S^0 \rightarrow \text{invisible}) < 8.4 \times 10^{-4}$
 (90% C.L.)

JHEP 05 (2025) 092

First direct measurement of $K_S^0 \rightarrow \text{invisible}$; the UL still lies above the NP prediction