



FORMOSA

**looking forward to millicharged
particles at the LHC**

**Juan Salvador Tafoya Vargas (UC Davis)
on behalf of the FORMOSA Collaboration**

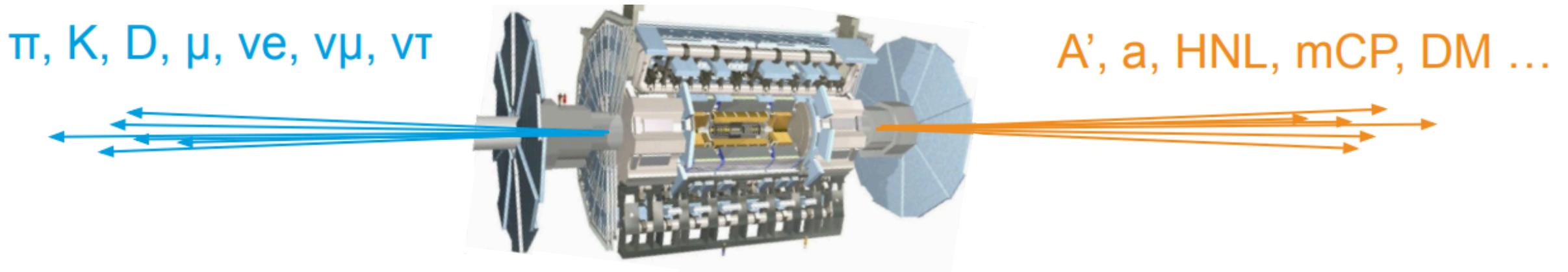
**EPS-HEP 2025
Marseille**

2025.07.09

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Searching for new physics

- New physics searches limited by the design of existing general-purpose experiments
- Take advantage of already-happening LHC collisions
 - forward region of e.g. ATLAS I.P. very rich in SM and potential new physics



- Physics collisions already happening → no considerable additional operational cost
 - Initial investment needed for the construction of the facility in the forward region
 - Allows for design and installation of dedicated detectors

New physics and dark matter → Hidden Valley

No signs of new physics
seen at the LHC (yet)

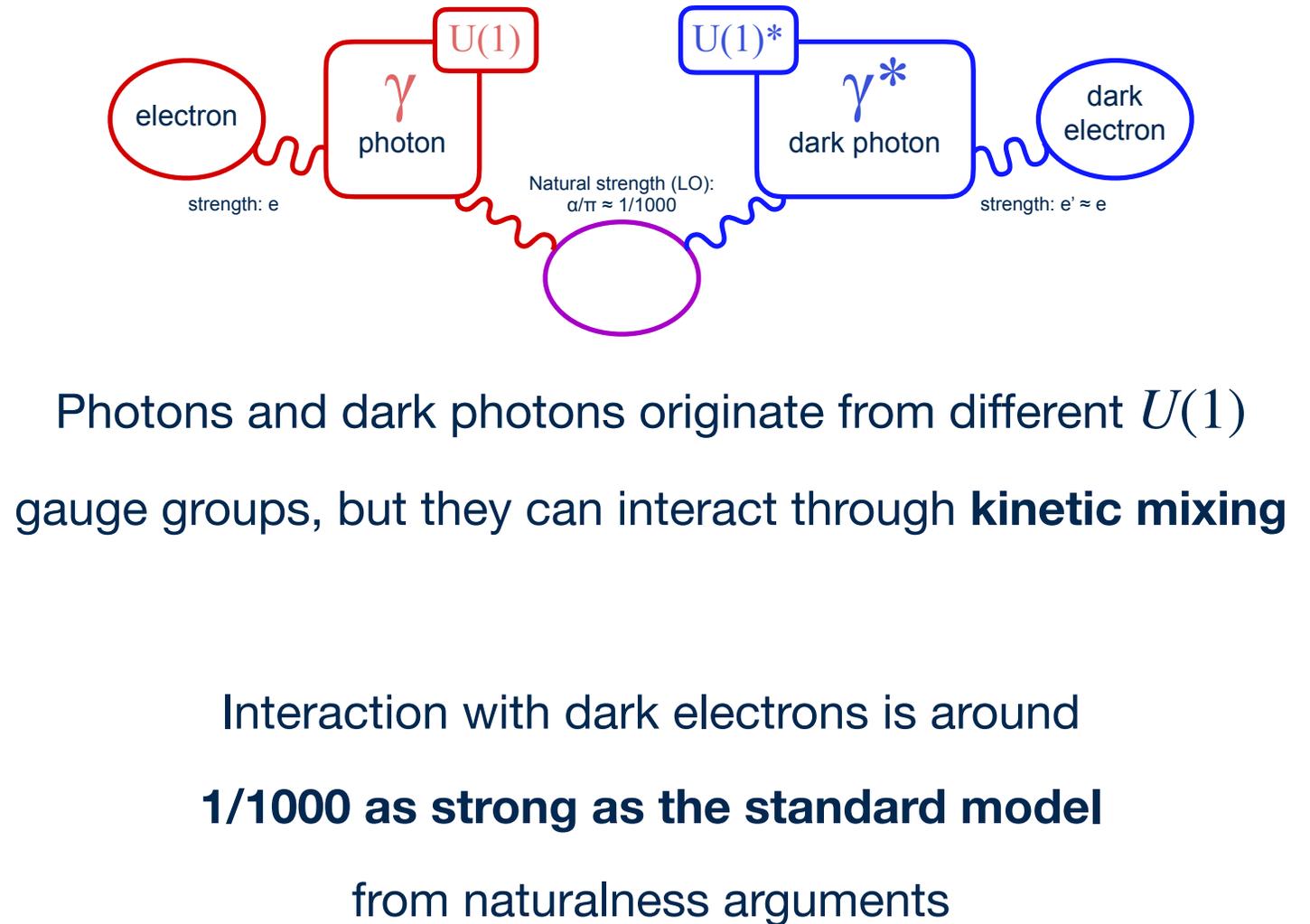
SM extensions that
include dark (or hidden)
sectors give very
plausible hint

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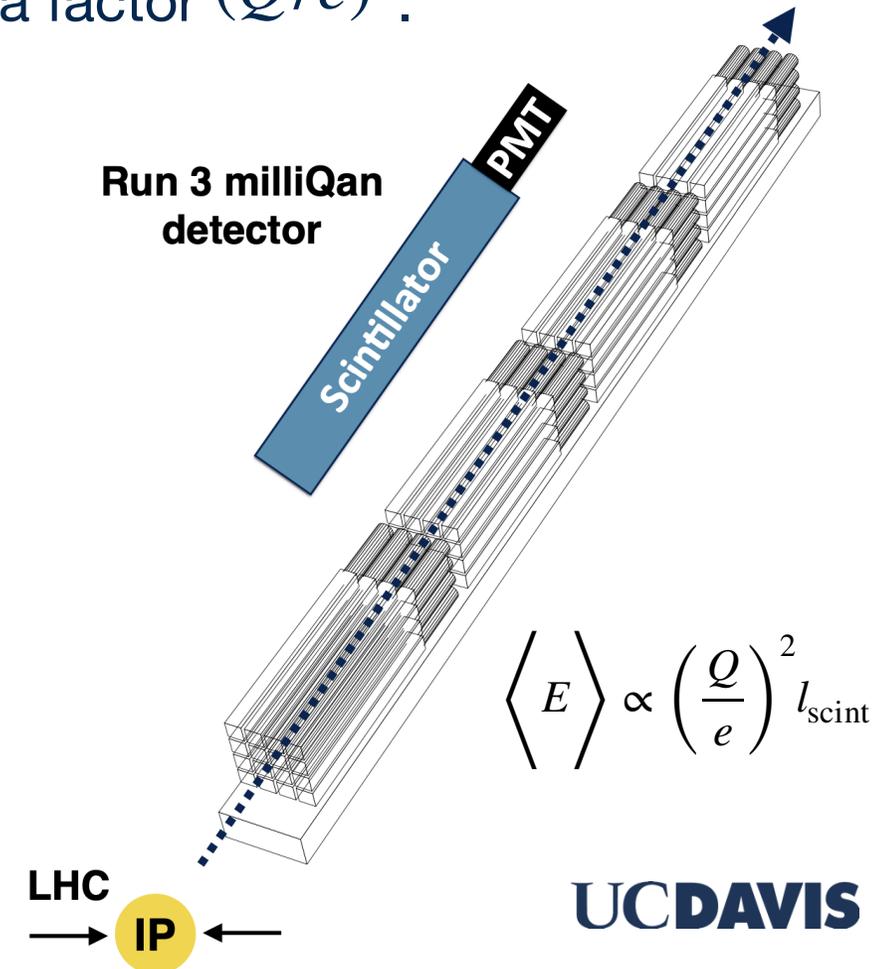
Millicharged particle searches at the LHC

Millicharged particles (**mCPs**) are well motivated in dark sector theories, but difficult to detect because the interaction strength is reduced by a factor $(Q/e)^2$.

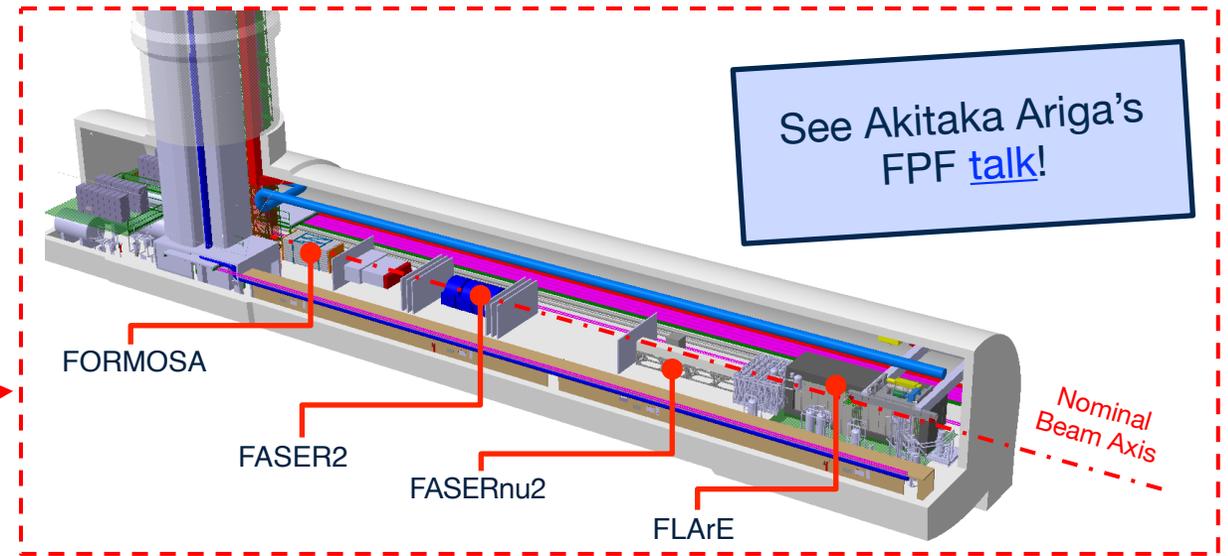
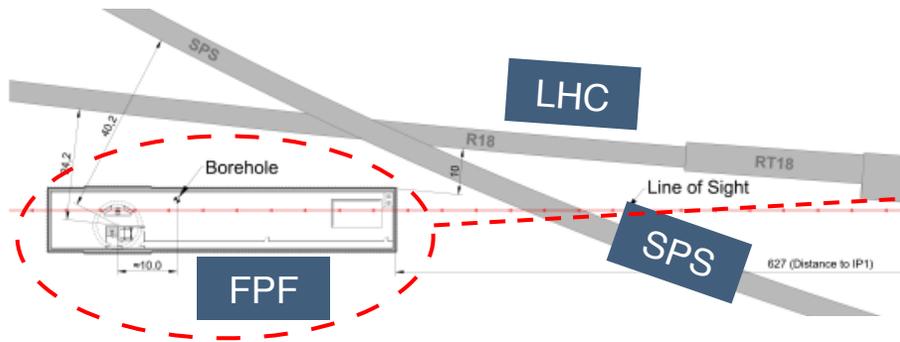
Core concept: Use array of efficient long scintillator bars + PMTs to detect ionisation from mCPs.

Challenges:

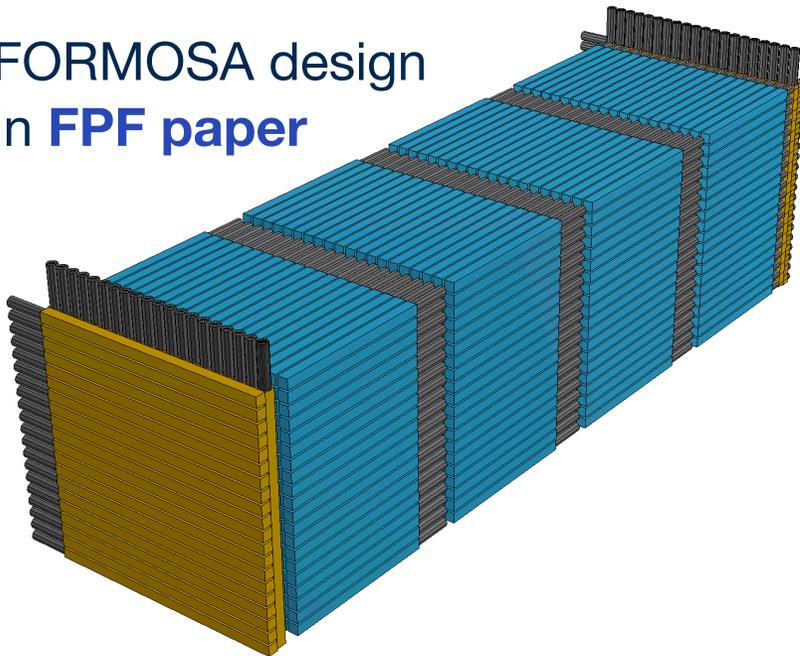
- Expect few scintillation photons to be produced
→ must be able to detect single scintillation photons
- Well controlled backgrounds → signatures “point” at the interaction point, triggering on sets of signals within small time windows (~15 ns)



FORMOSA at the FPF



FORMOSA design
in **FPF** paper



Aimed at the ATLAS I.P., located in the very-forward region ($\eta > 9$)

Expect to see **~250x rate** of millicharged particle detection in the forward region compared to the central one (i.e. milliQan, $\eta \sim 0$)

20 rows x 20 cols x 4 layers of bars for detection

Shielded from most SM particles produced at the I.P. by **~ 200m of rock**

Shielded from LHC beam radiation by **>10m of rock**

Main background: collision muons → segmented muon veto panels

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Main background at FPF: I.P. muons

Muon background/radiation flux simulated and validated (CERN FLUKA team)

- expected muon flux $O(1 \text{ Hz/cm}^2)$ within 1m of LOS
- affected by triplet magnets
- studying if flux can be reduced with sweeper magnet

Important to understand DAQ strategy to take into account muon-induced activity

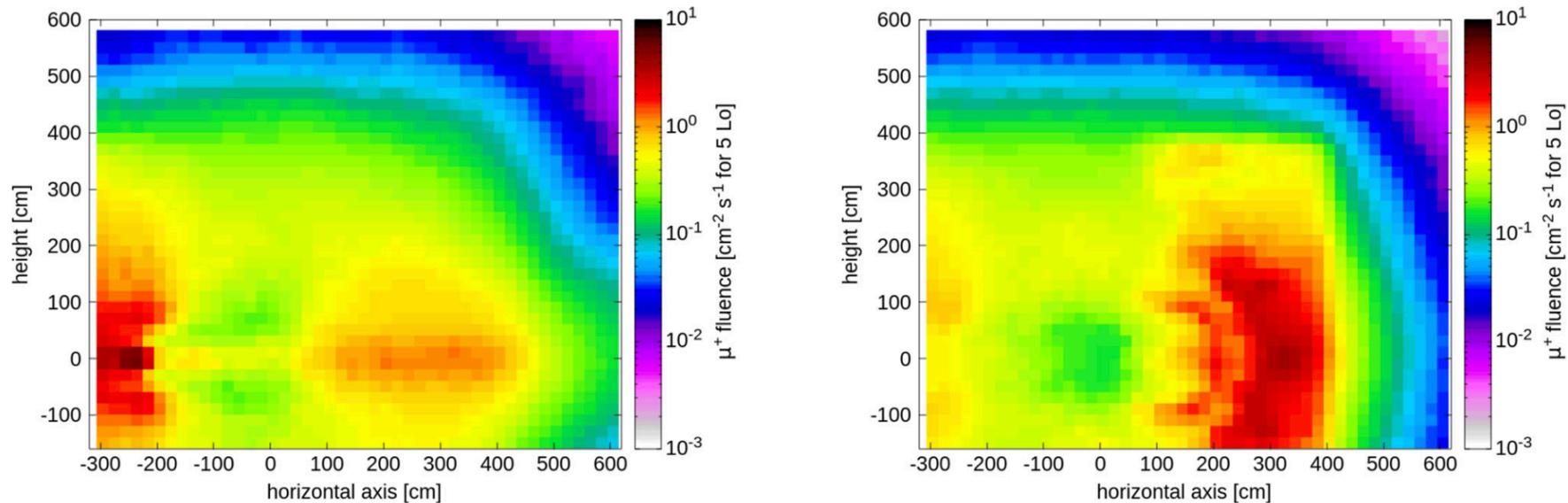


Fig. 14 The muon fluence rate for μ^- (left) and μ^+ (right) in the transverse plane in the FPF cavern for the HL-LHC baseline luminosity of $5 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$. The coordinate system is defined such that (0, 0) is the LOS, and -ve x is towards the center of the LHC ring

See FPF's input to [ESPPU](#) for more details

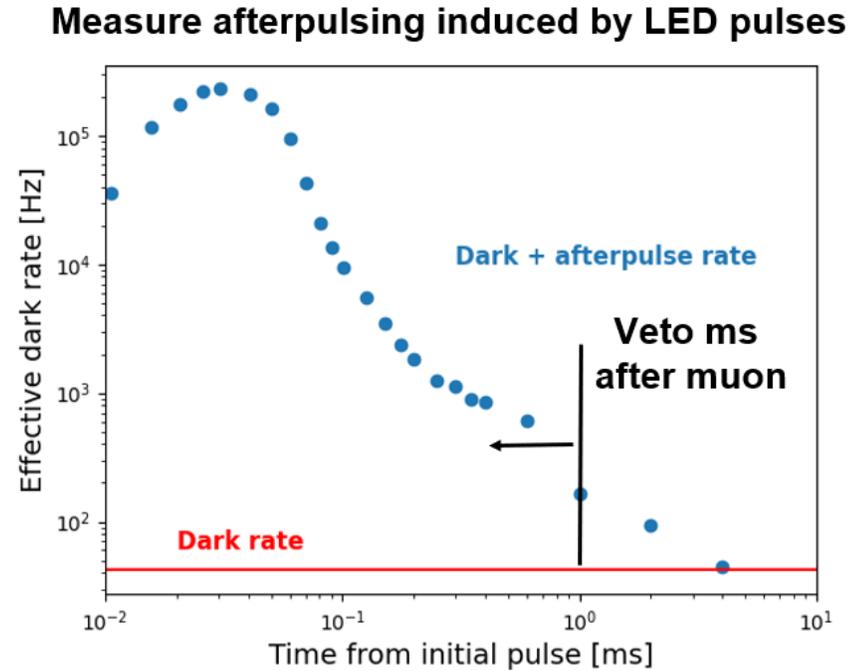
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- affected by
- studying if flu

Large induced signals (such as those left by muons) can lead to a temporary "increase" of the PMT's dark rate, making triggering unreliable.

Lab bench studies suggest we can work around it by applying a small cooldown deadtime.



to understand DAQ
take into account
induced activity

Fig. 14 The muon fluence rate for μ^- (left) and μ^+ (right) in the transverse plane in the FPF cavern for the HL-LHC baseline luminosity of $5 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$. The coordinate system is defined such that (0, 0) is the LOS, and -ve x is towards the center of the LHC ring

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The FORMOSA demonstrator

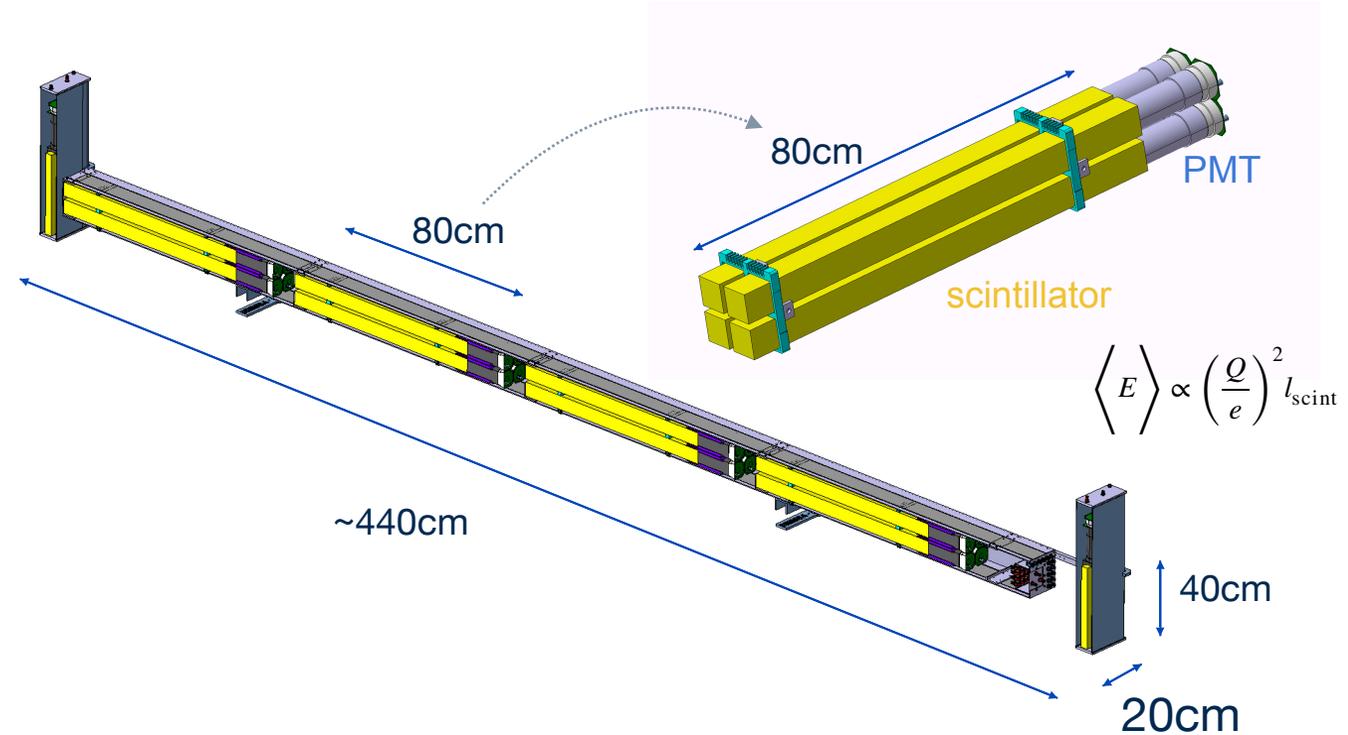
A small-scale version of the full FORMOSA was installed during YETS2023-2024 to **prove concept** and **target new phase space**

- 2 rows x 2 columns x 4 layers of bars
- front+back muon veto panels

Located in the UJ12 cavern (behind FASER), in the opposite side of the to-be FPF

Additional features added later:

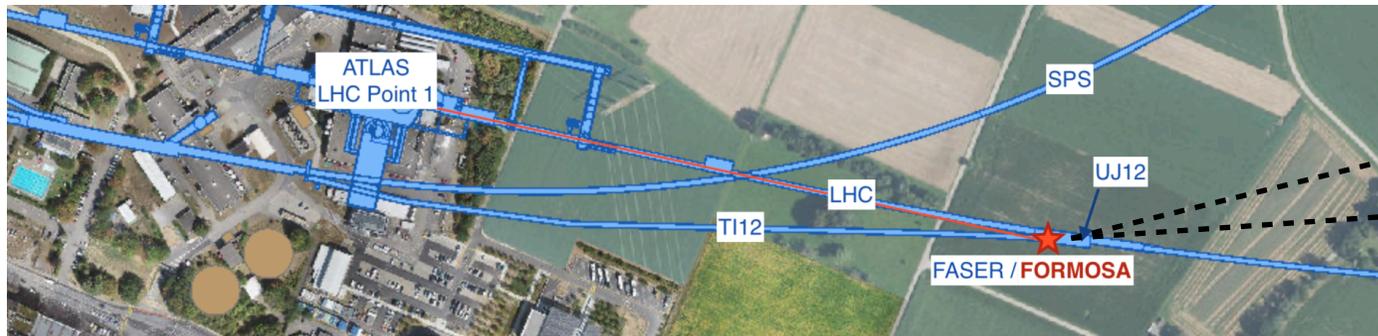
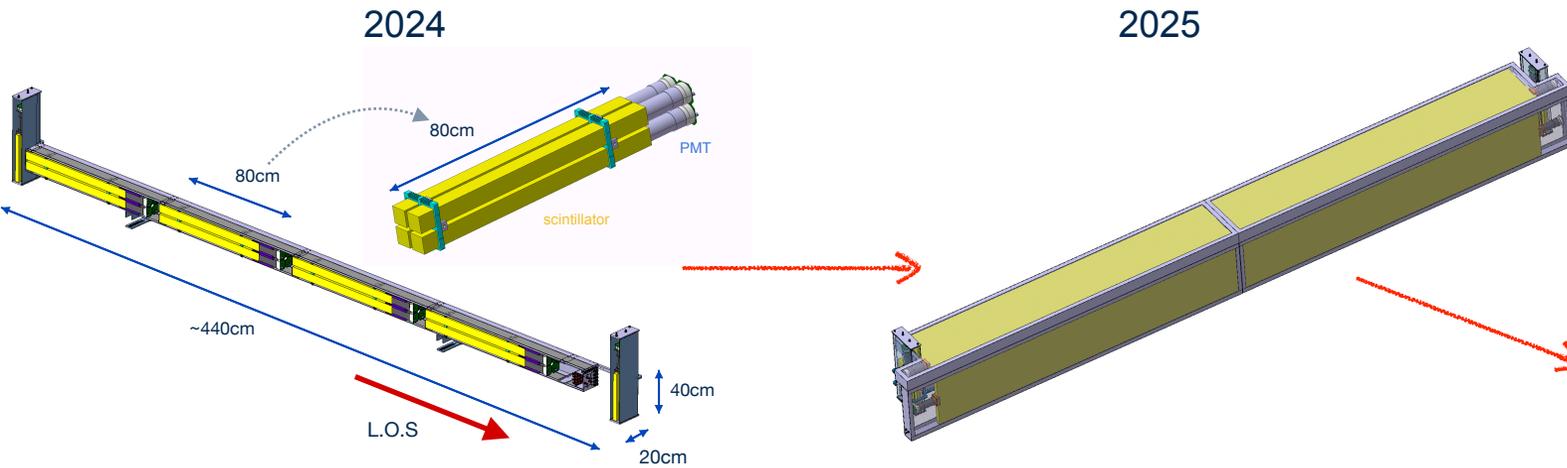
- CeBr3 test module
- Hermetic side+top panels



The FORMOSA demonstrator

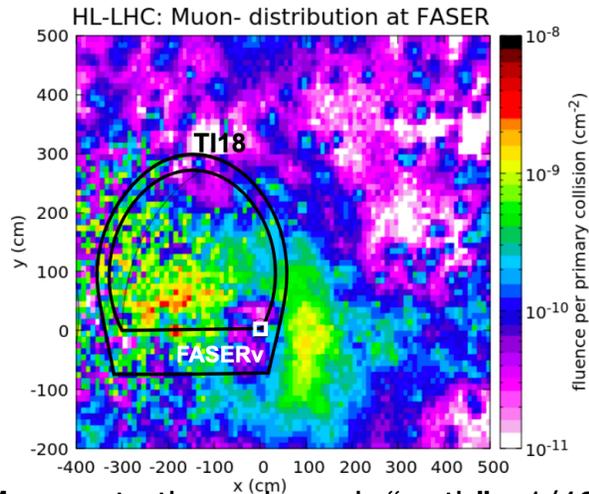
Proof of concept: collecting data and validating design/DAQs through Run 3

Design evolving to test new technologies and adapt to UJ12's environment



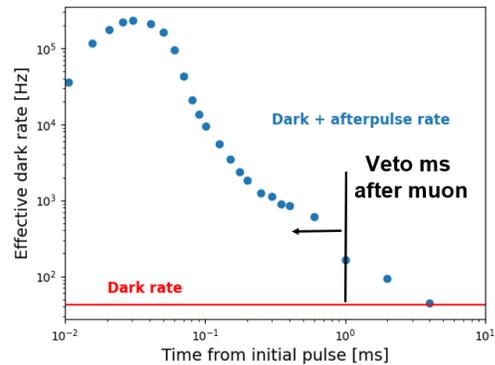
Main backgrounds at the demonstrator

Dominant expected background in the forward region:
afterpulses initiated by through-going muons



Muon rate through each “path” $\sim 1/40\text{ms}$

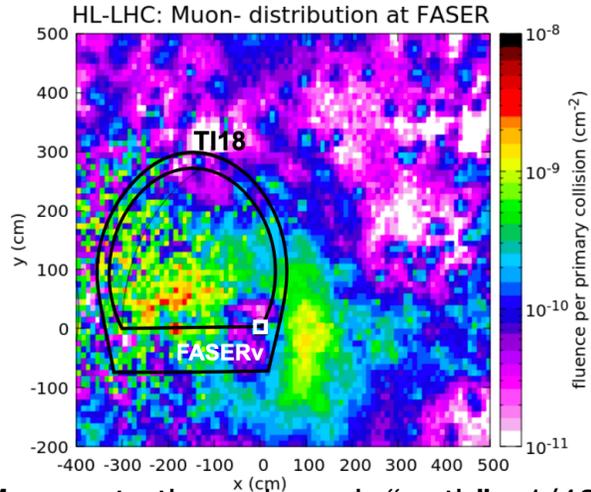
Measure afterpulsing induced by LED pulses



Lab study:
Can devote a ms of
deadtime to cool down
following a collision muon.
But muons are frequent
($\sim 100\text{Hz}$ for demonstrator),
so can't wait too long

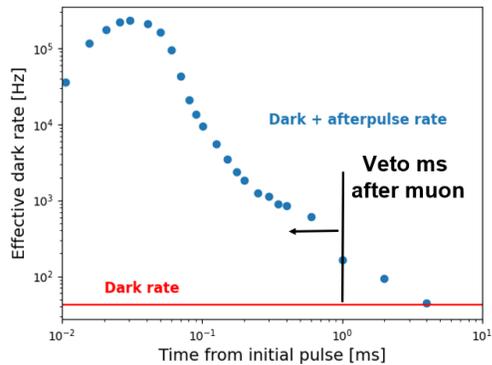
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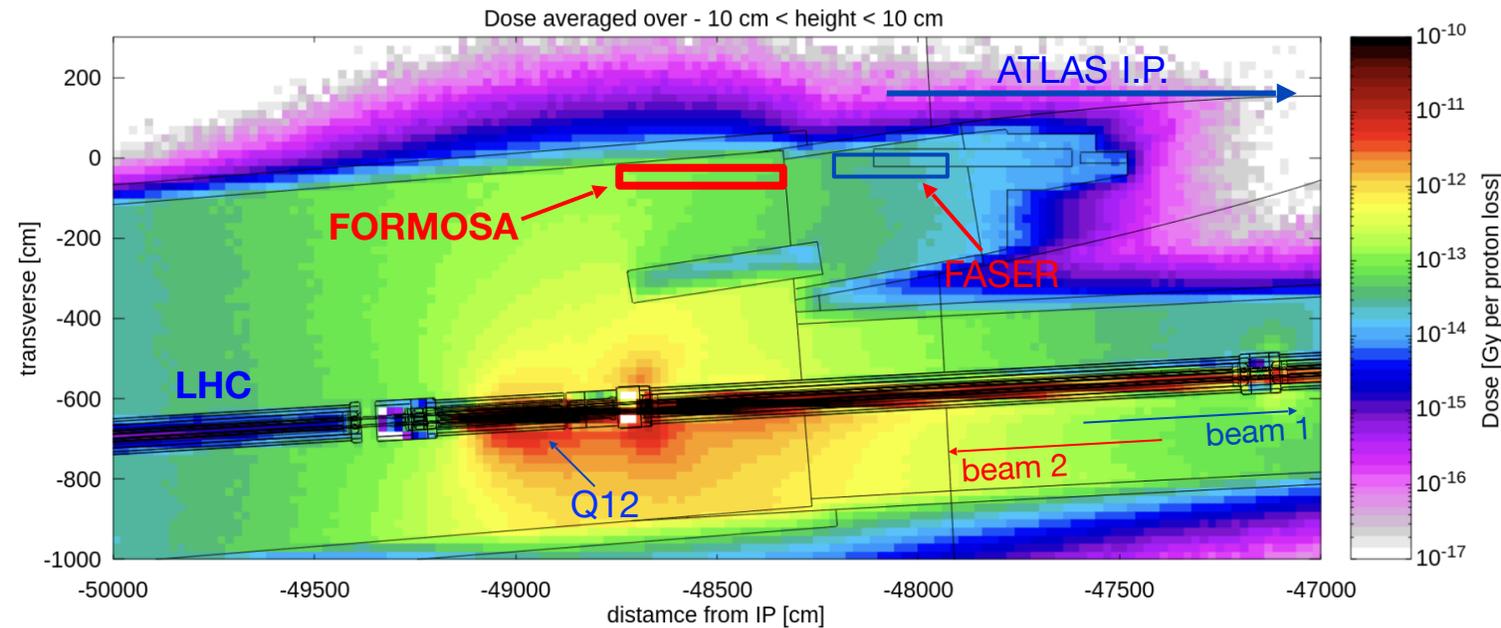
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An additional challenge at demonstrator: **Q12 beam background**
 (N.B. this background will NOT be present in the FPF)

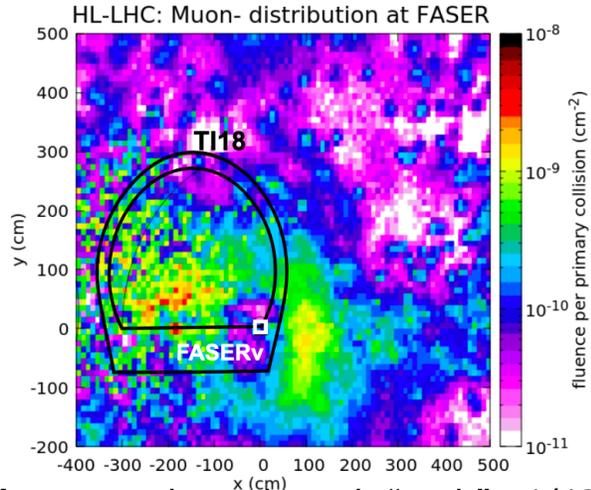


Absolute average dose due to proton losses at Q12 (simulation provided by FASER)

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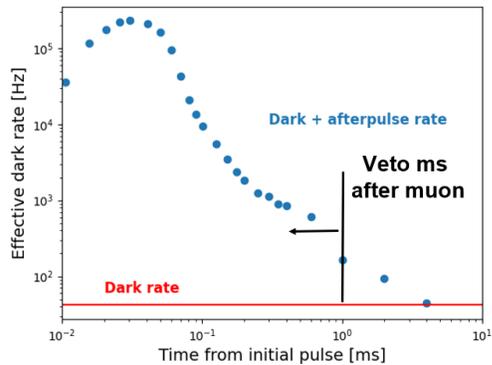
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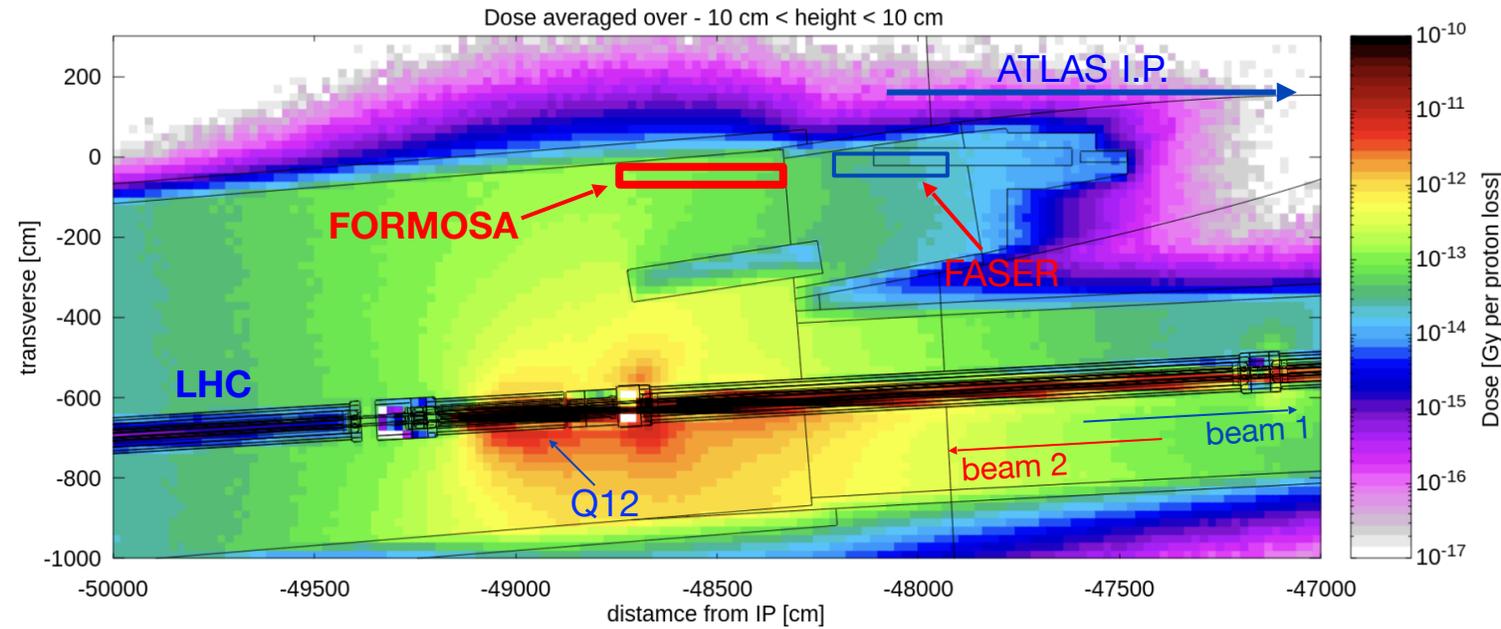


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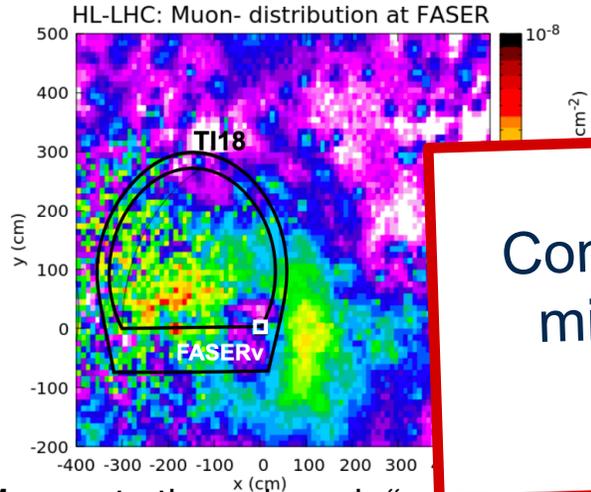
Side panels

Large front+back muon veto panels

Main backgrounds at the demonstrator

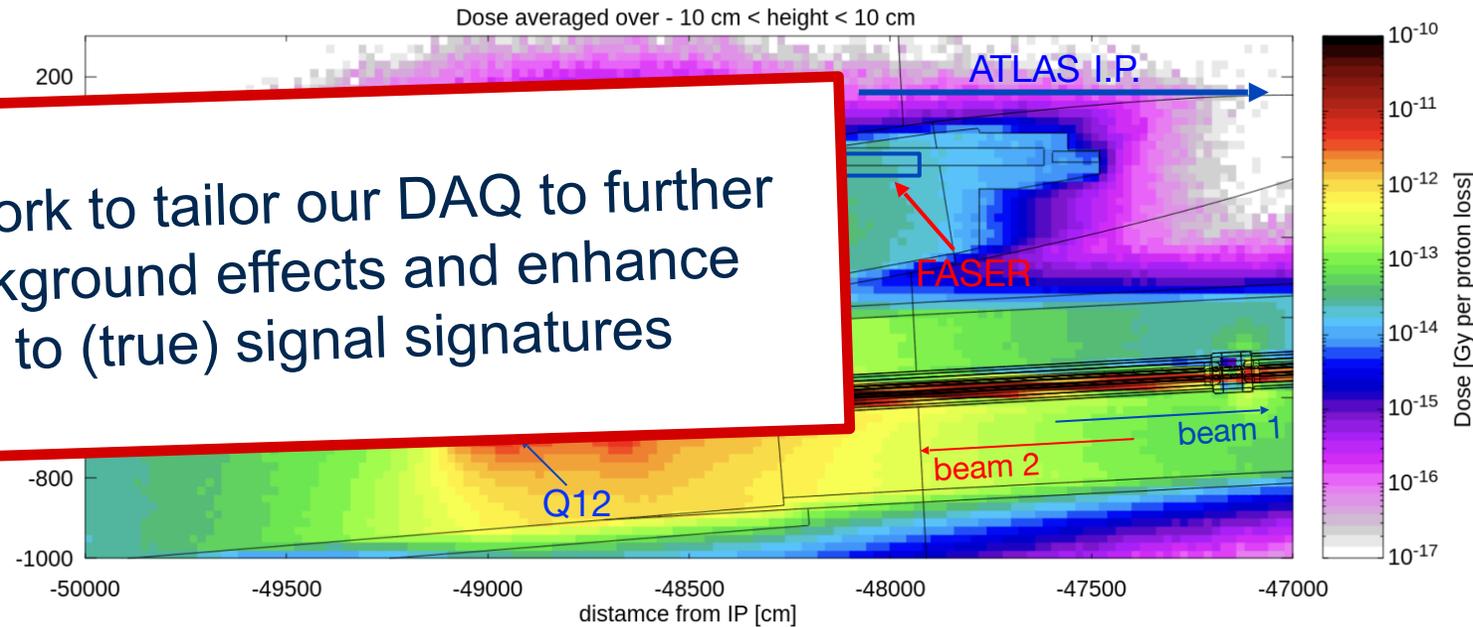
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Muon rate through each "pair" ~ 1740 ms

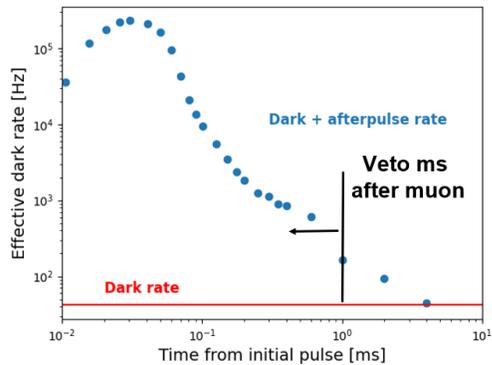
Continuous work to tailor our DAQ to further mitigate background effects and enhance sensitivity to (true) signal signatures



Absolute average dose due to proton losses at Q12 (simulation provided by FASER)

Side panels

Measure afterpulsing induced by LED pulses

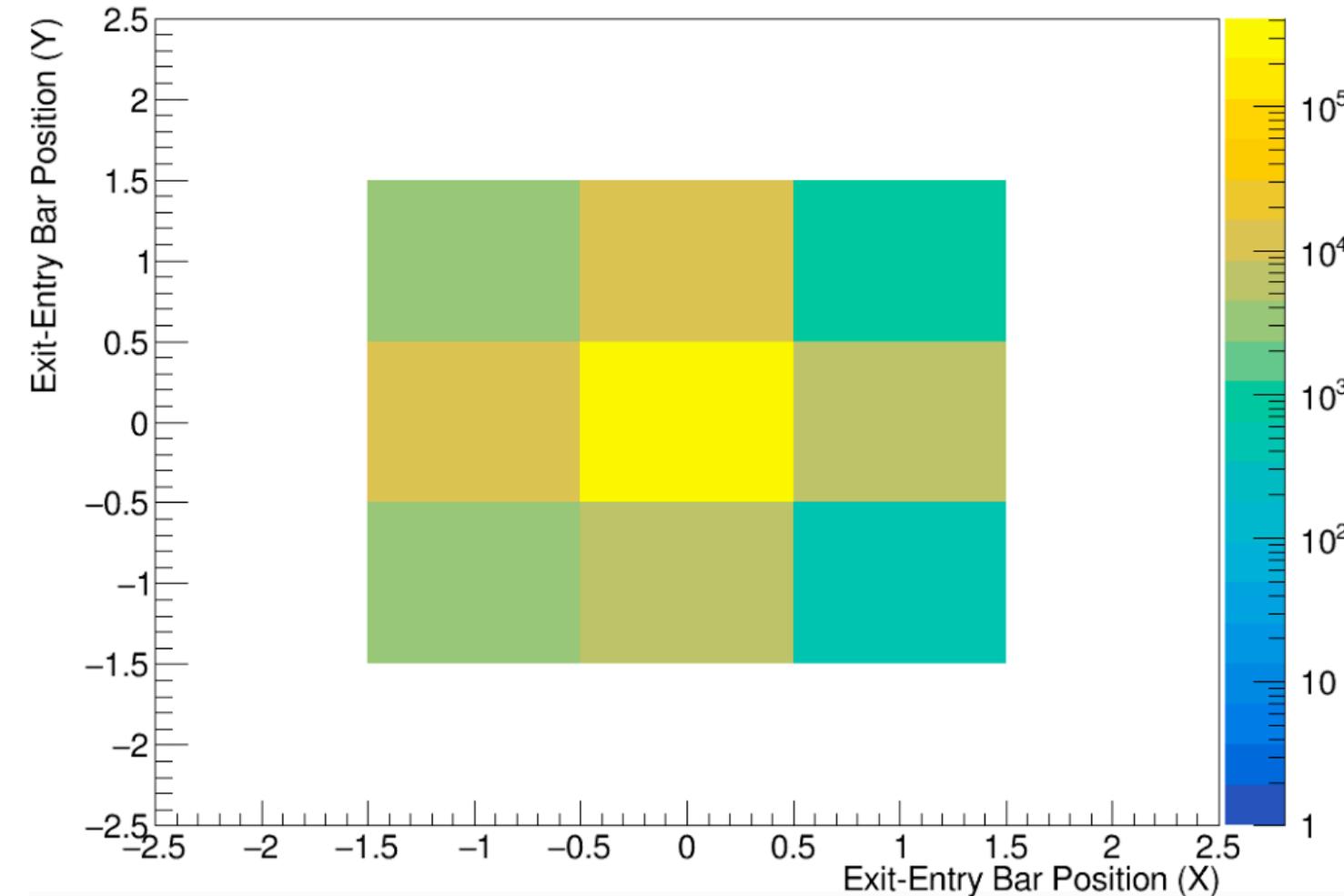


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Large front+back muon veto panels

Alignment validation

Muonic Angular Distribution

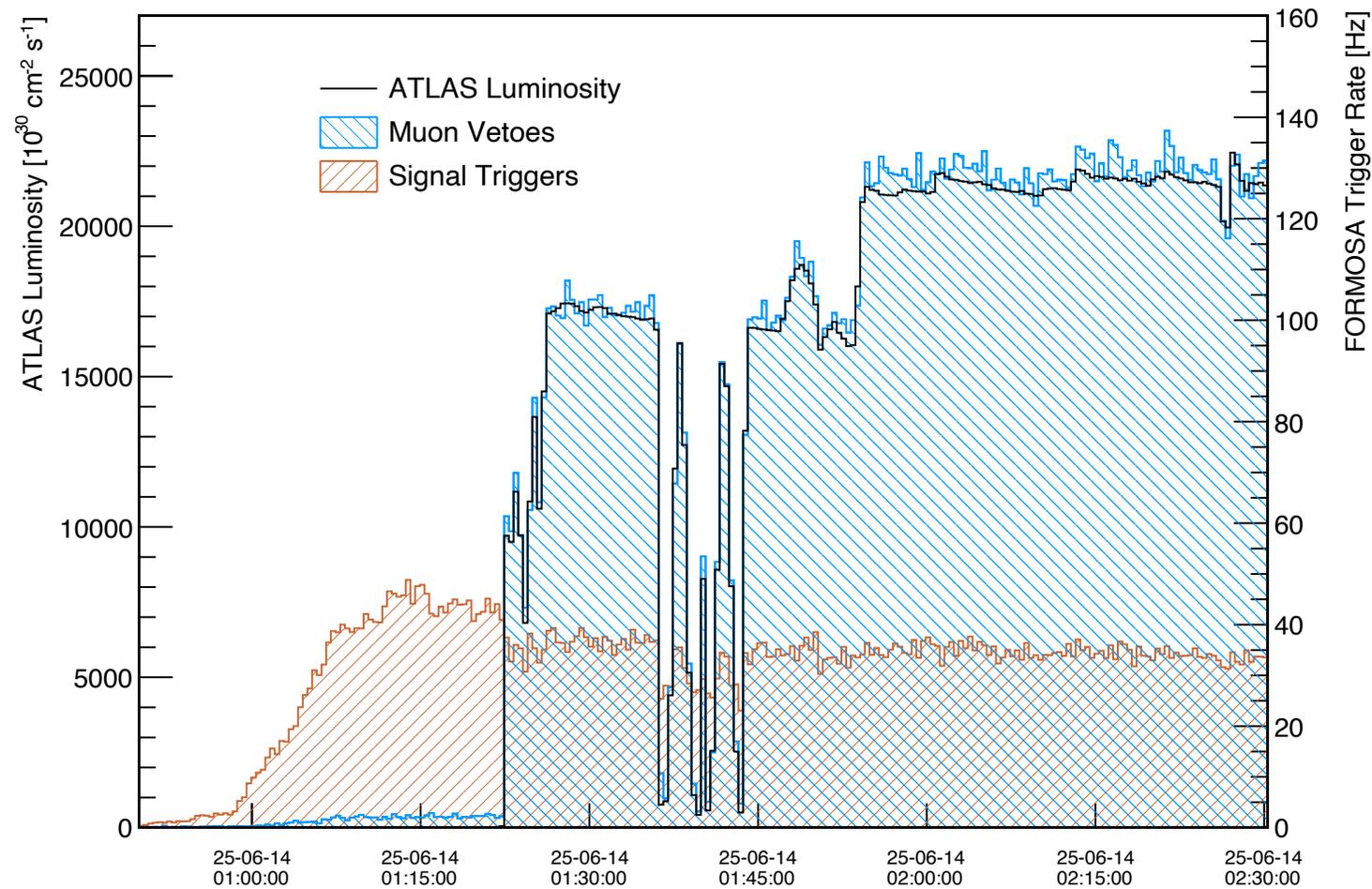


Checking the distribution of the “exit - entry” bar number of collision muon events

→ LHC collisional data confirms alignment!



DAQ validation

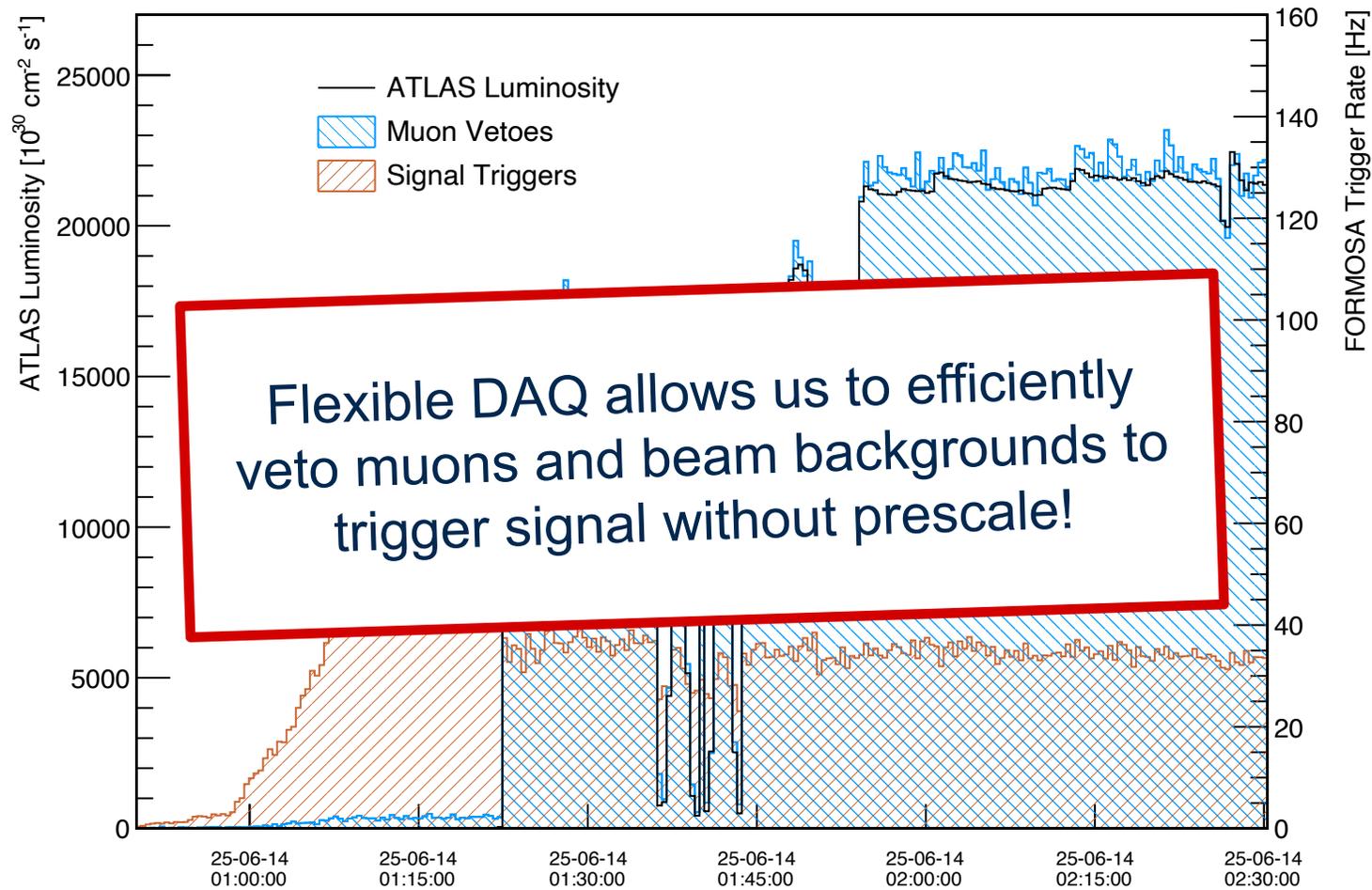


Data from the
14th of June 2025

**Signal triggers stably
collecting data**

**Muons very
effectively vetoed!**

DAQ validation



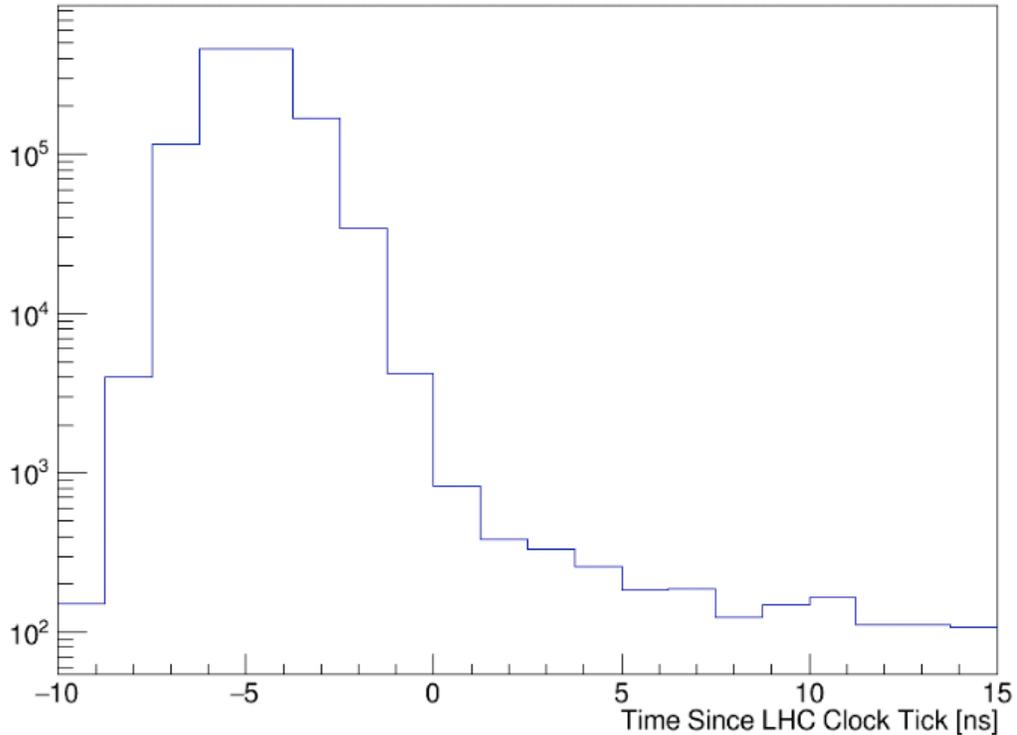
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Timing studies

Muon Intra-Clock Timing

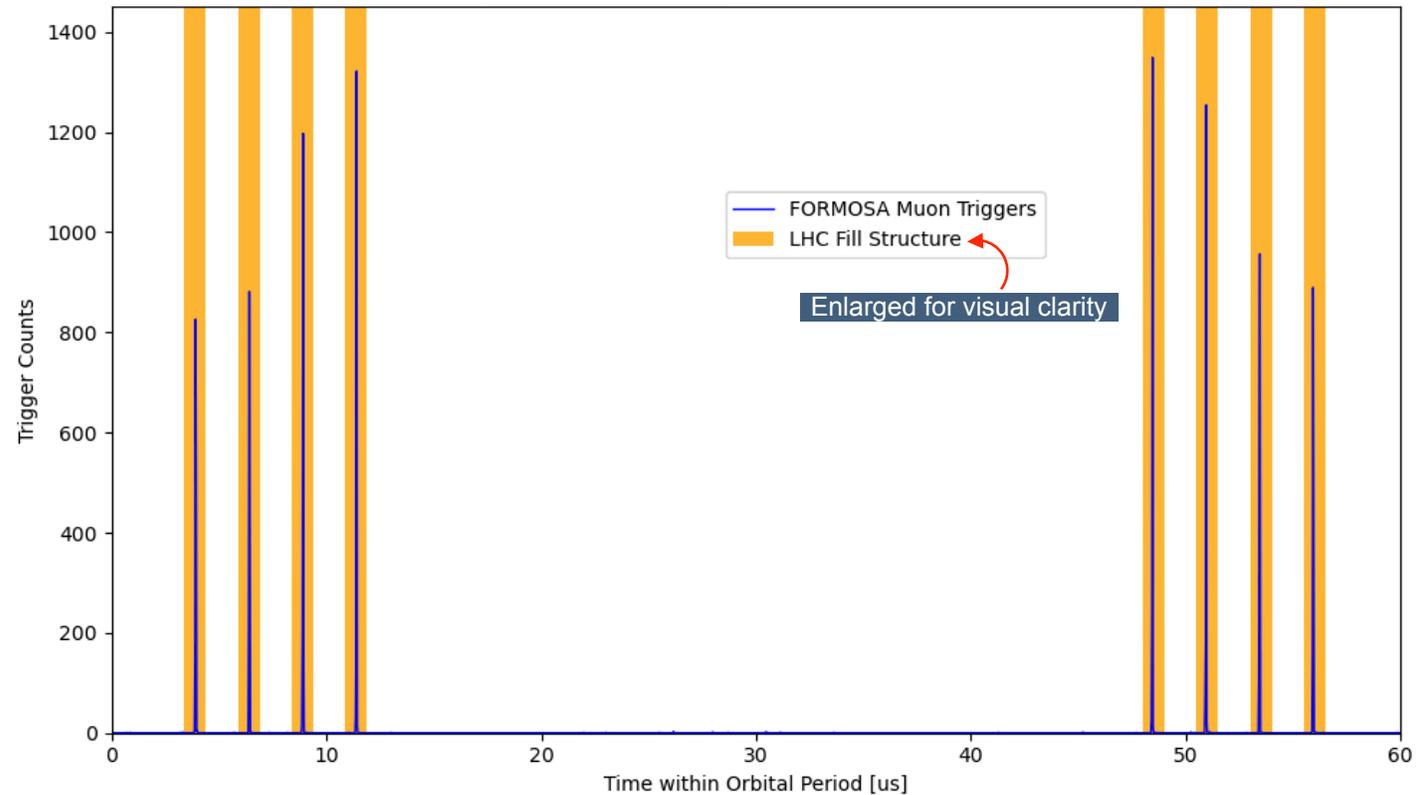


Muon arrival peak

Flat beam background

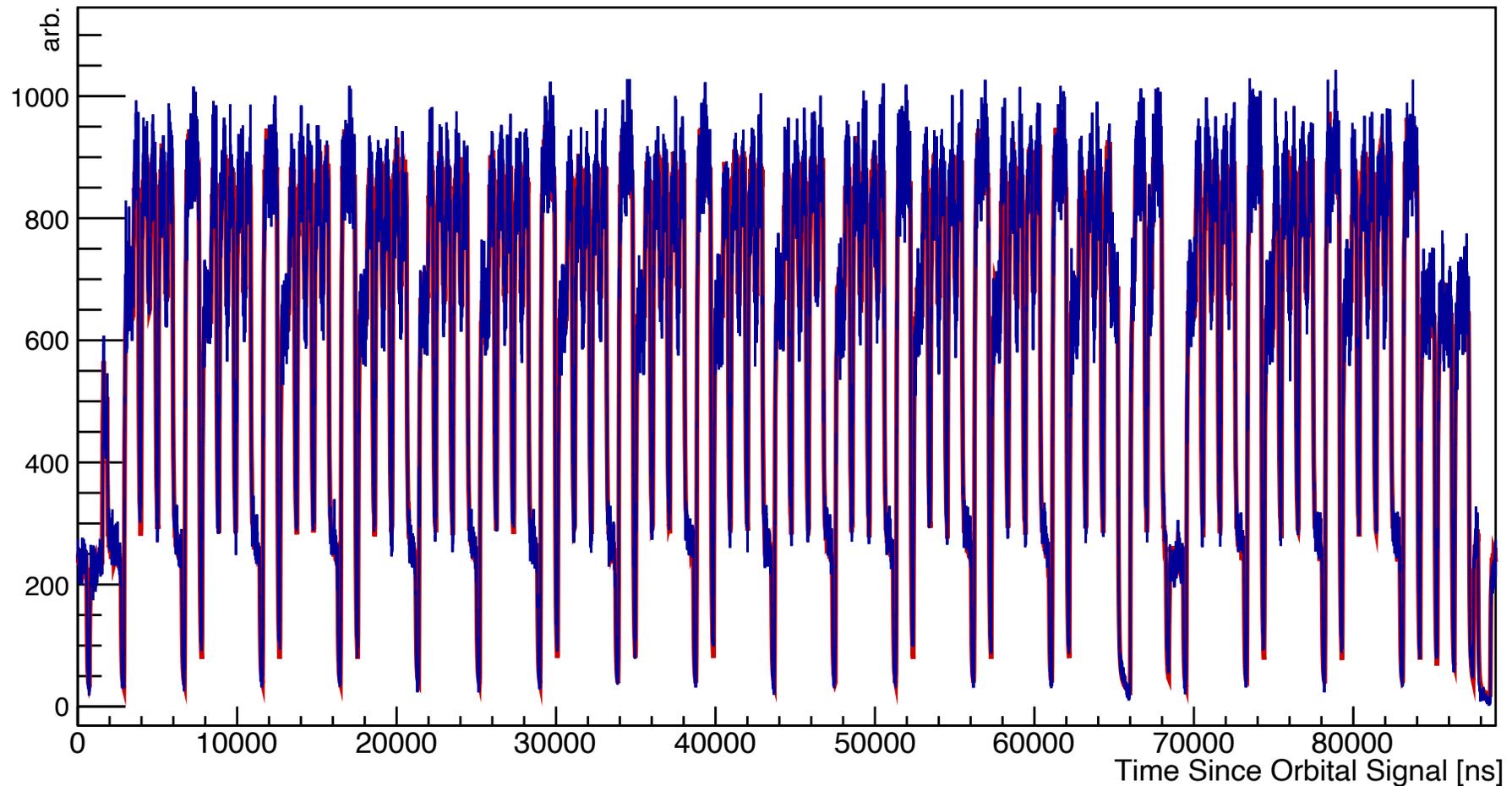
Looking *between LHC clock-ticks*

Timing FORMOSA Muon Triggers within LHC Orbital Period



Looking *within orbital* of a 12b fill with 8 colliding bunches

Timing studies: probing the LHC's fill scheme



We have excellent timing resolution!

Next step: to understand our trigger activity relative to the LHC's bunch structure

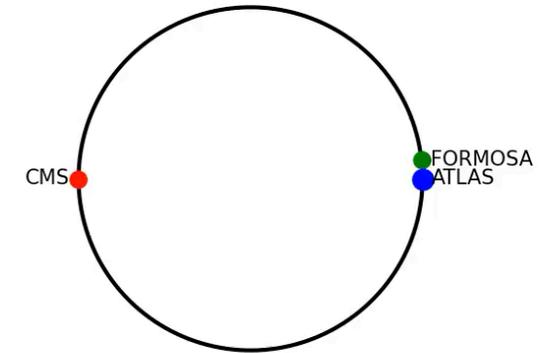
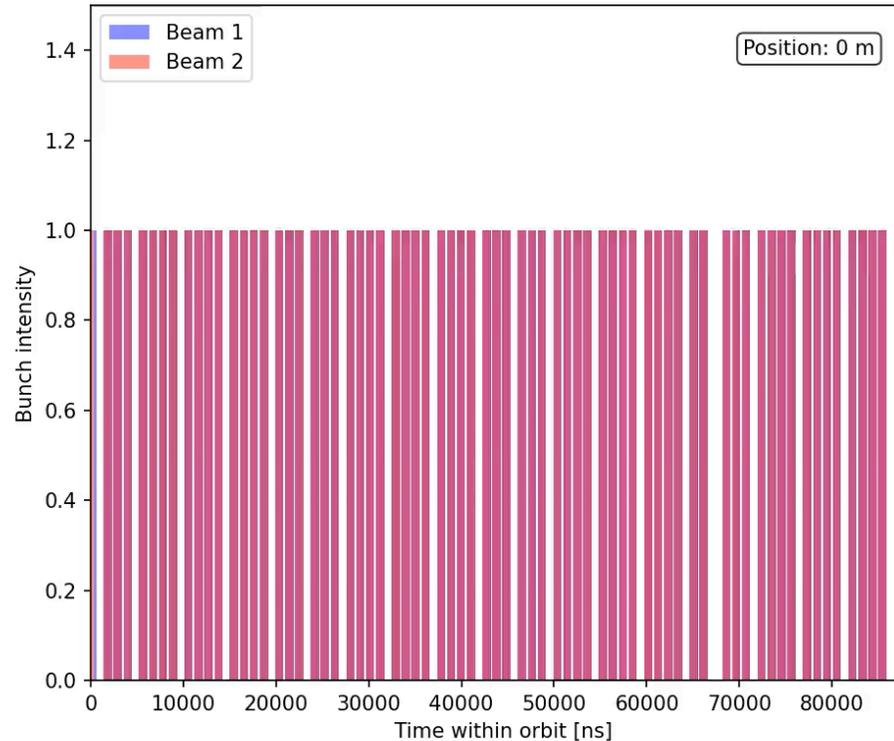
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Timing studies: probing the LHC's fill scheme

We want to compare our trigger times since last orbital signal against the LHC's fill scheme and I.P.1's collision scheme.

A first study:

- Track/move the position of all beam 1 and 2 bunches in their corresponding directions
- Each bunch/collision contributes with Landau-distributed trigger activity
- Place the demonstrator at 465m away from I.P.1



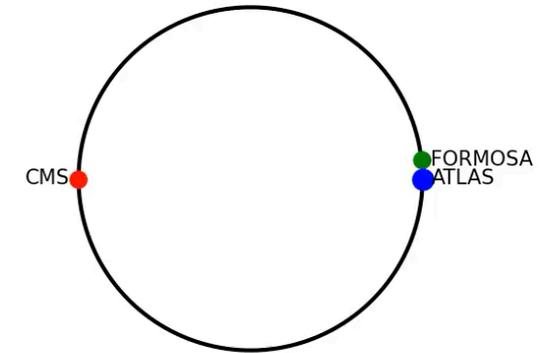
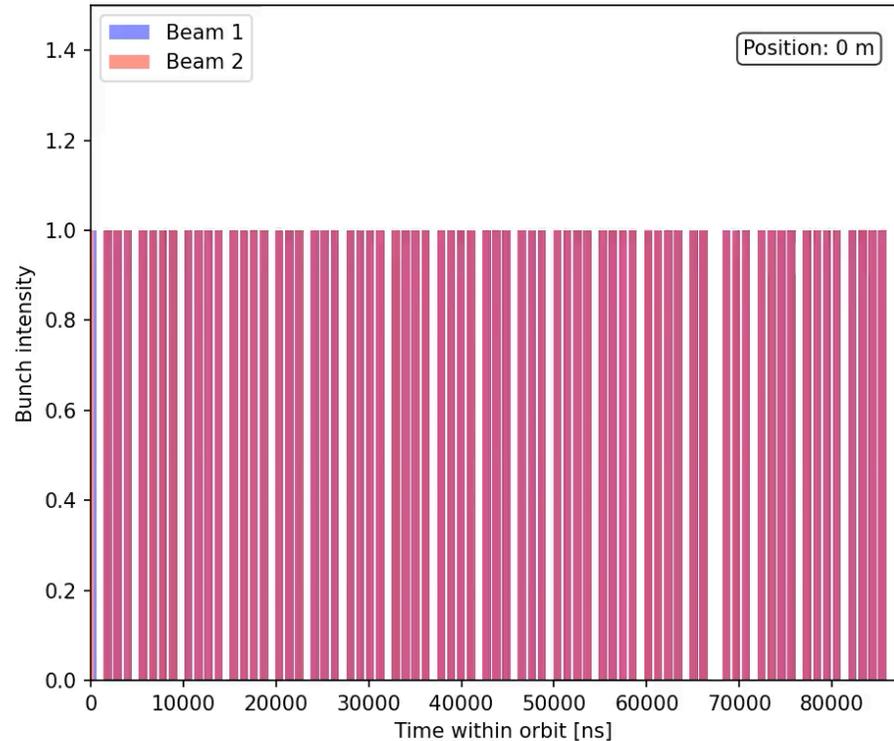
$$f(t) = A_1 * \sum_{i=0}^{Beam1} (I_{1,i} \cdot L(t, \mu_i + a_1, \sigma_1)) + A_2 * \sum_{j=0}^{Beam2} (I_{2,j} \cdot L(t, \mu_j + a_2, \sigma_2)) + A_3 * \sum_{k=0}^{Collisions} ((I_{1,k} \times I_{2,k}) \cdot L(t, \mu_k + a_3, \sigma_3)) + C$$

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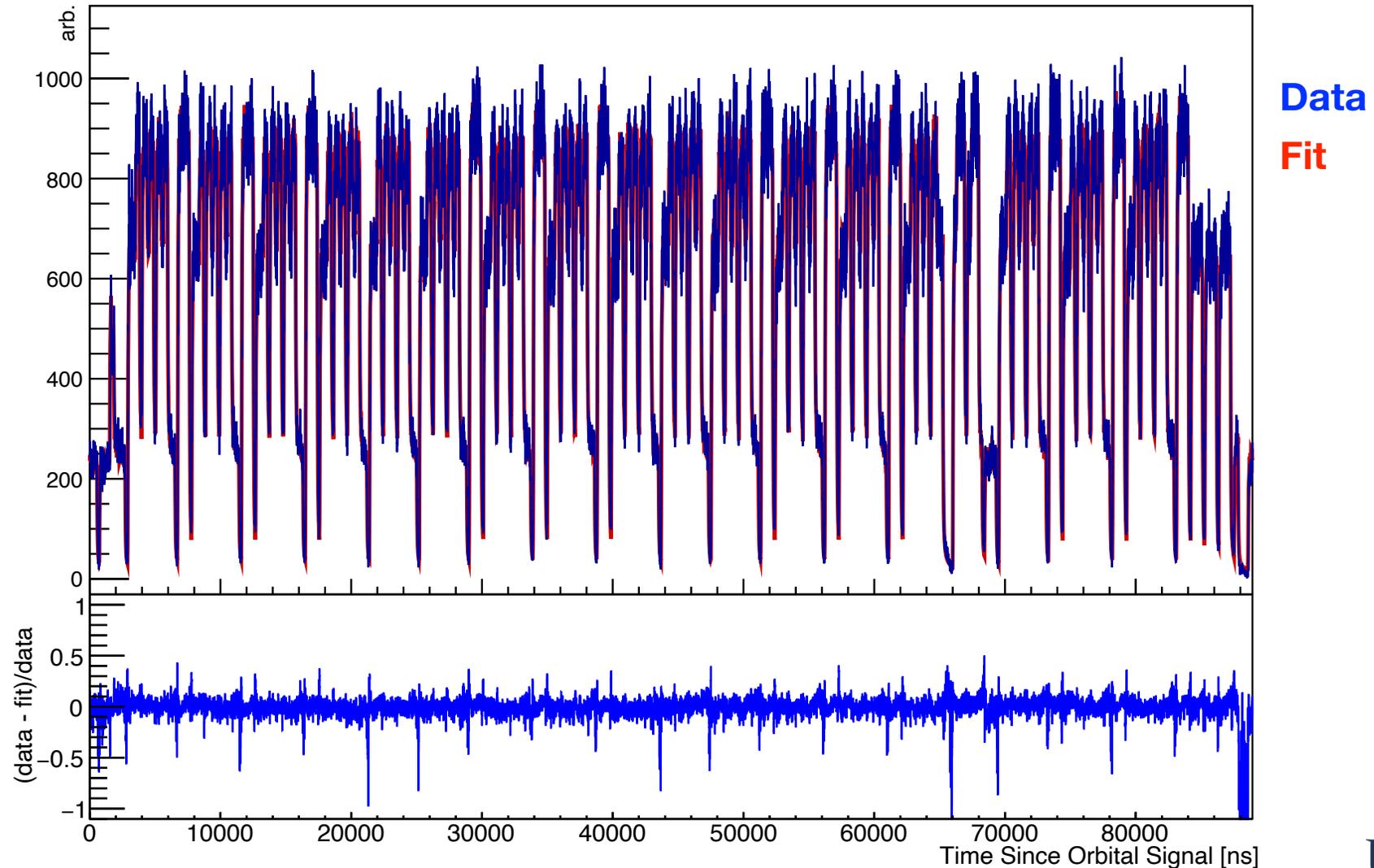
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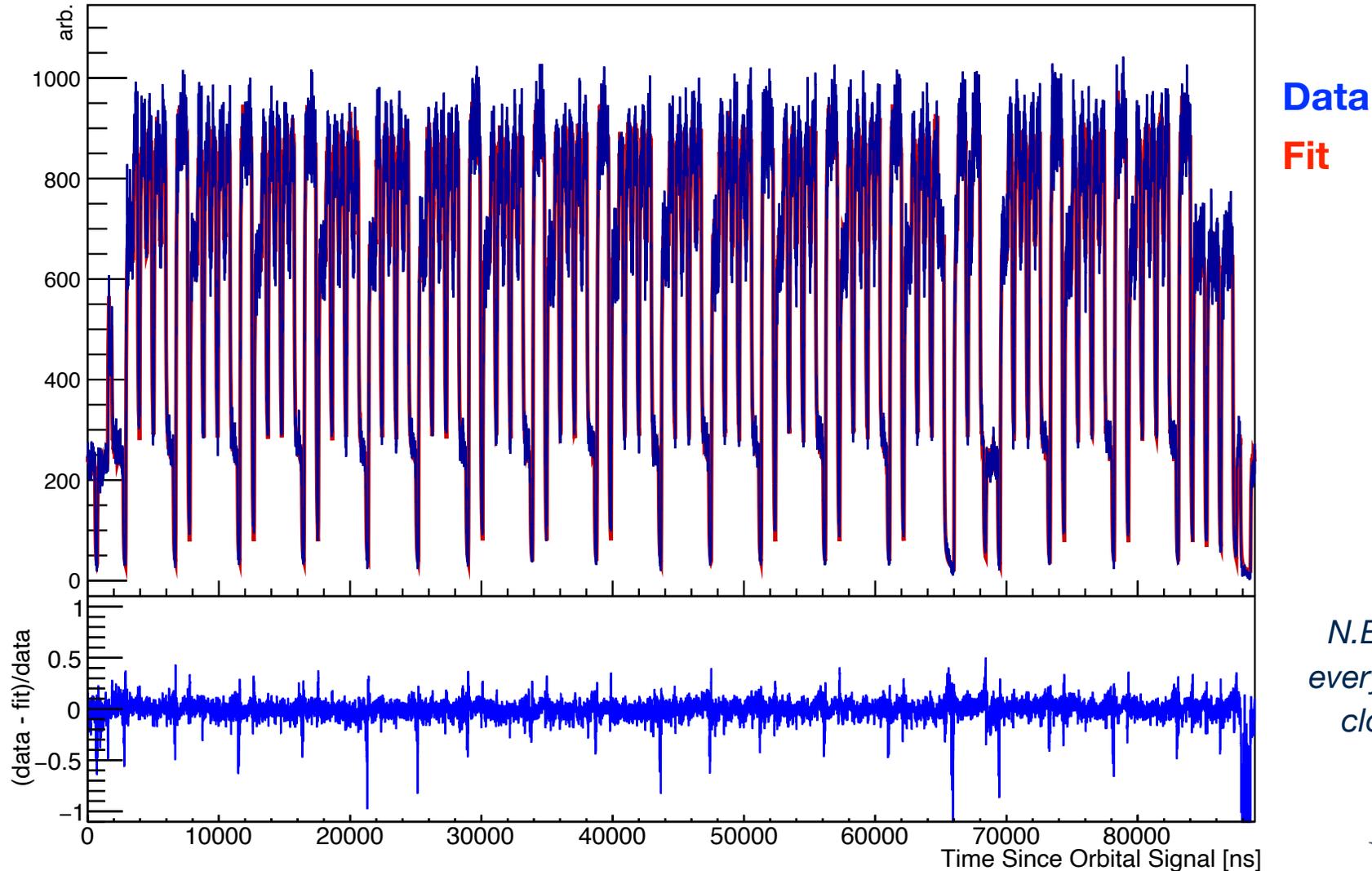
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Comparing FORMOSA Background Data to Circulating Beam Model



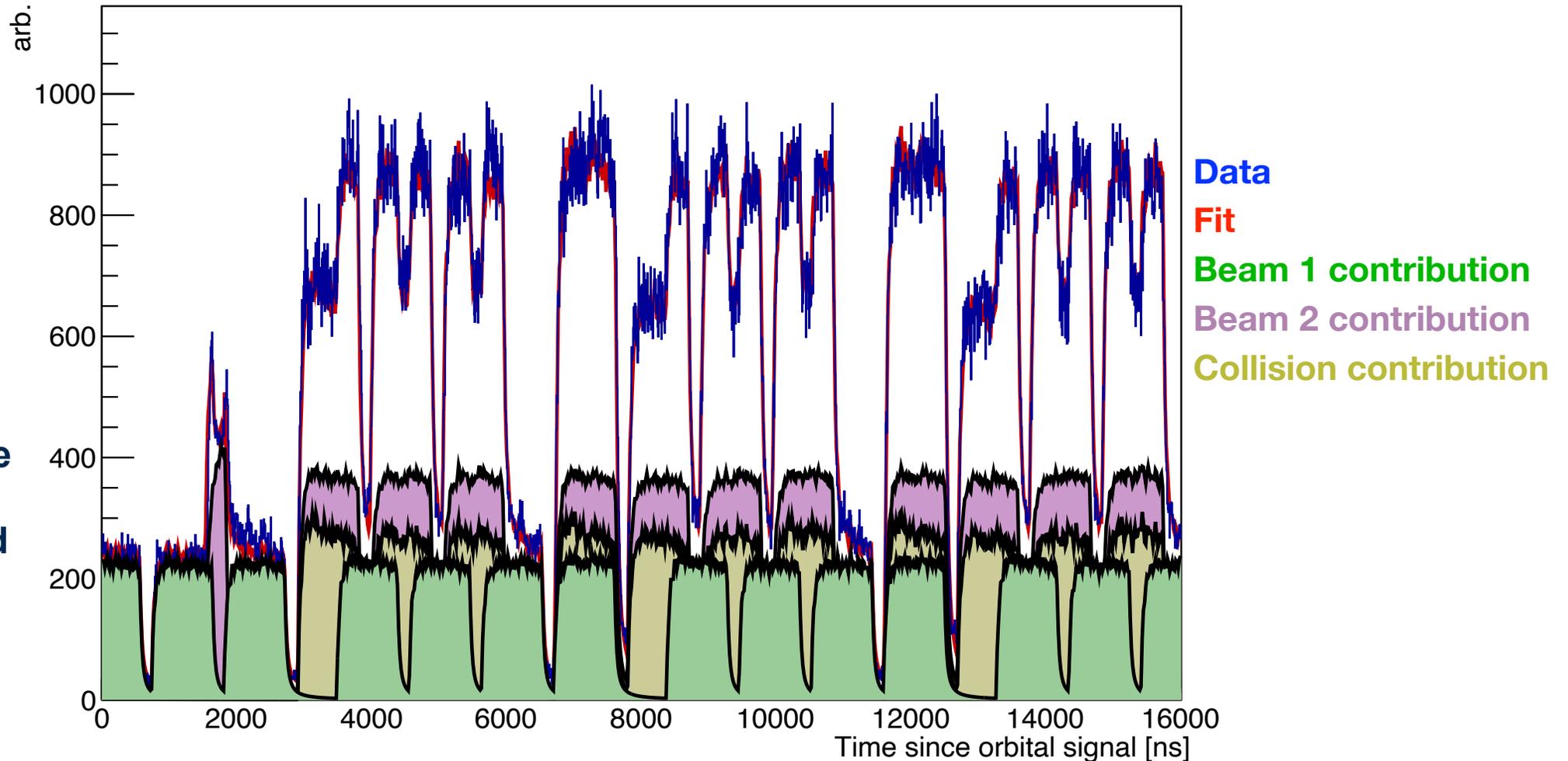
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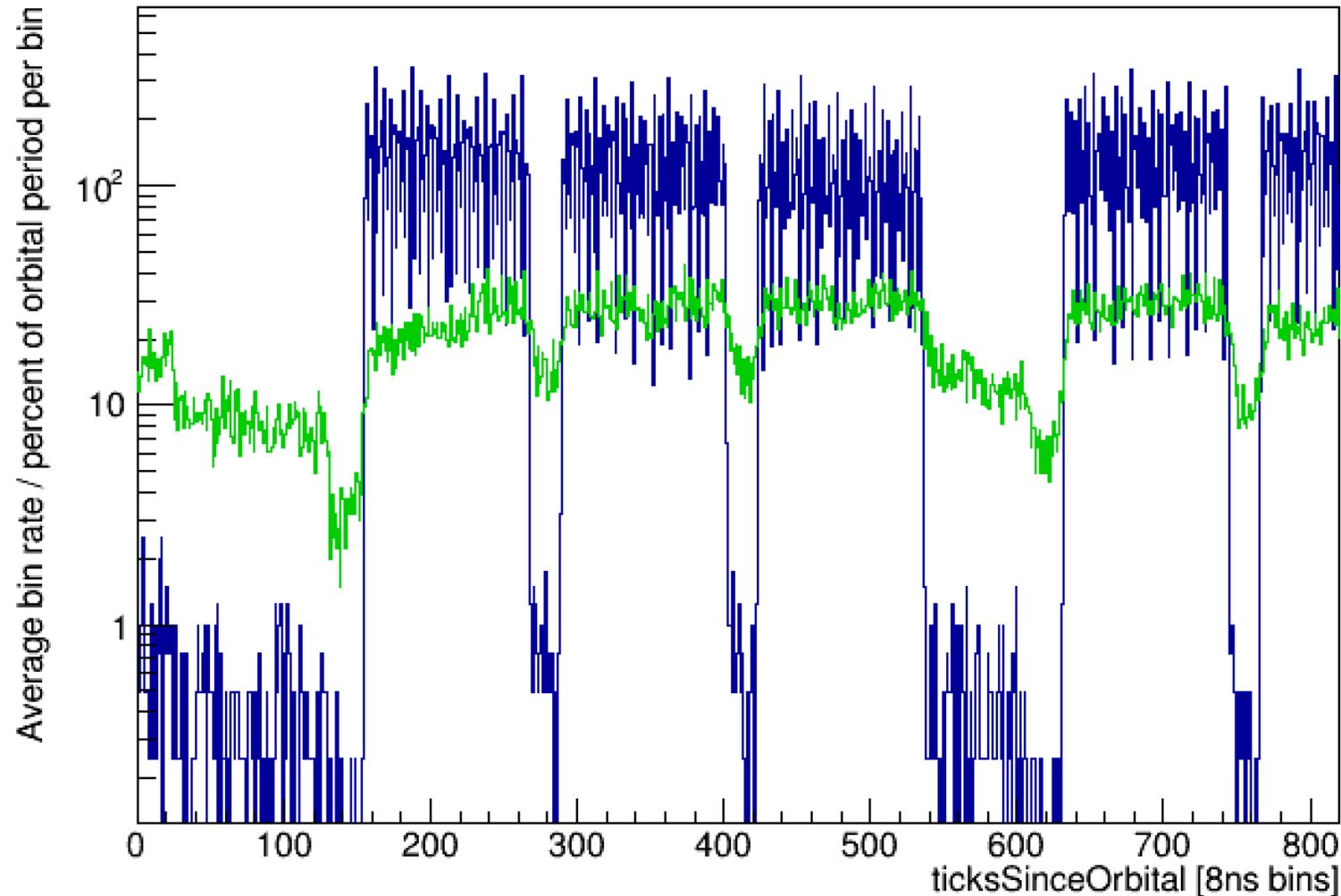


Zooming-in and drawing separate contributions hints origin of different signatures

Extremely valuable information for signal/background separation!

Muon veto studies

Comparing Triggers (Green) to Muon Vetoes (Blue)



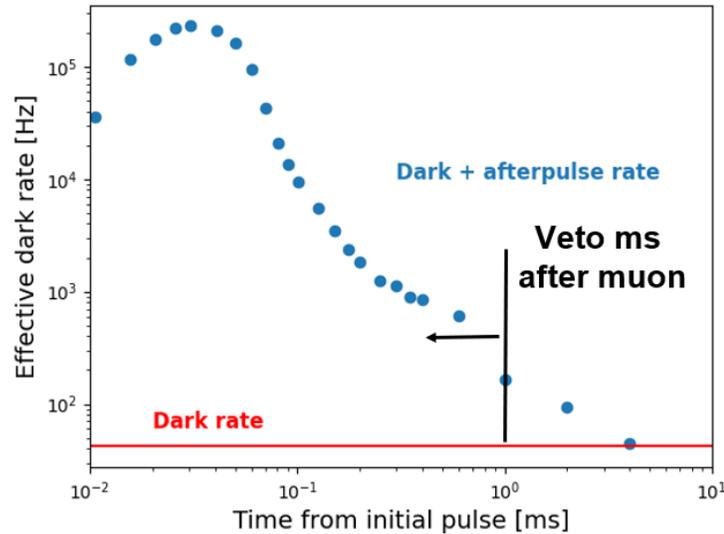
Fill with 2460b

N.B. LHC collisions every 25ns while DAQ clock runs at 8ns!

Crucial to determine when physics from collisions may be happening!

Muon veto and afterpulsing

Measure afterpulsing induced by LED pulses

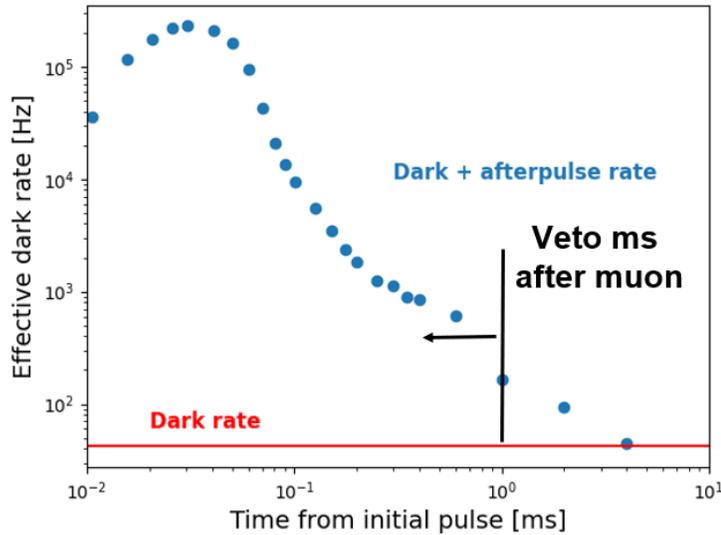


Study done for a single PMT suggests
1 ms of deadtime is enough

For ~ 100 Hz rate of muons, this
corresponds to 10% deadtime

Muon veto and afterpulsing

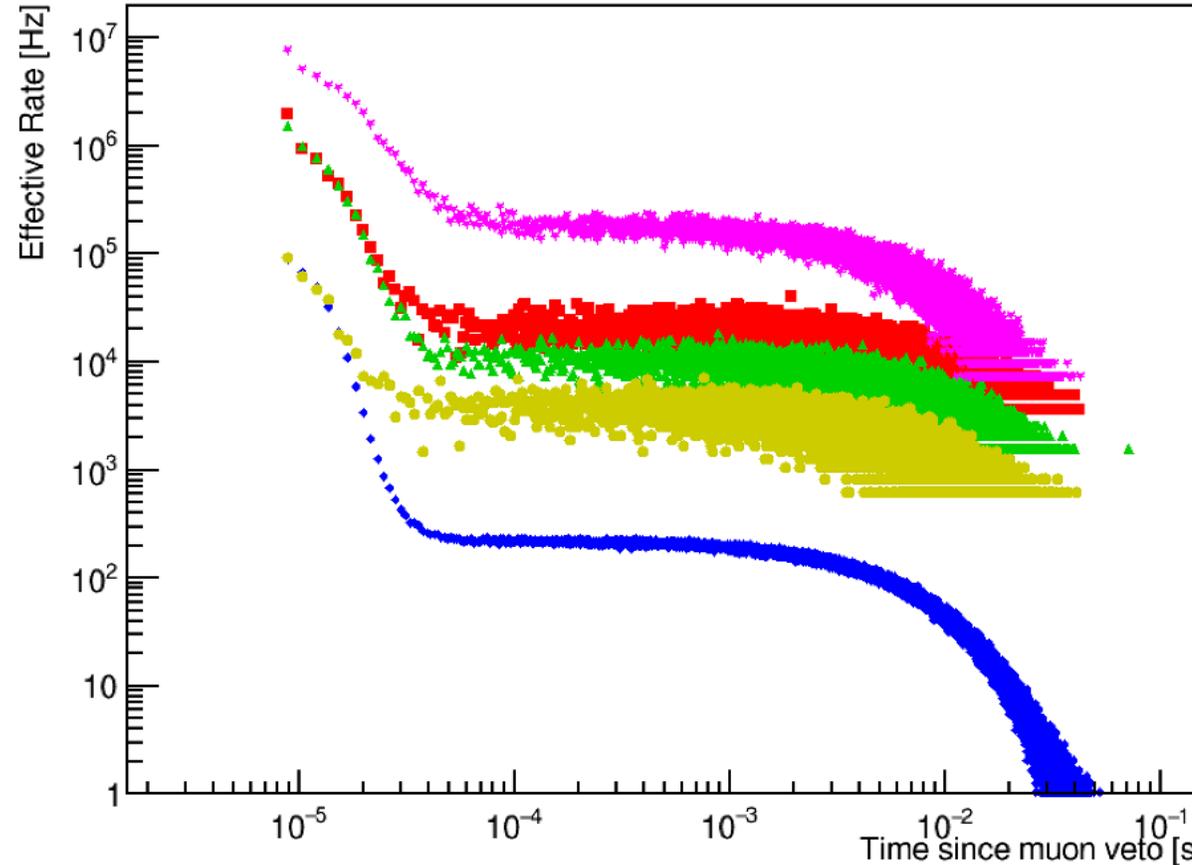
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Comparing Post-Muon-Veto Rates for FORMOSA Triggers



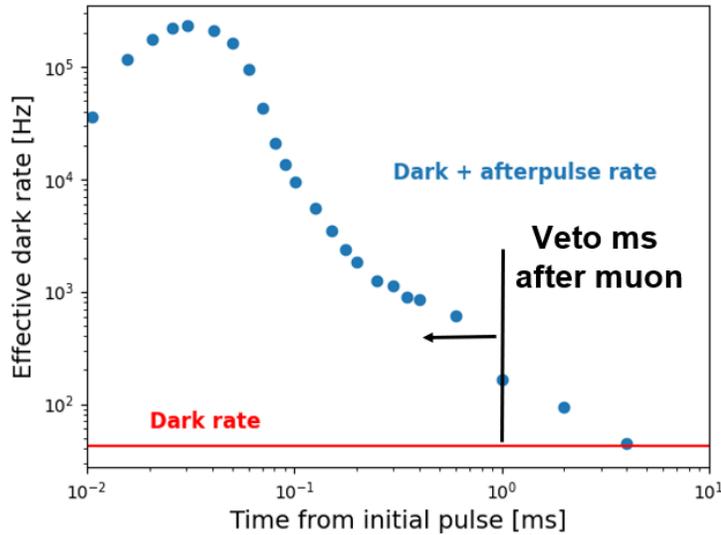
Signal triggers
4Layers
nLayers ≥ 3
nBars ≥ 3
2Layers (both)

Special run with
no muon deadtime

Overall scaling
still in progress

Muon veto and afterpulsing

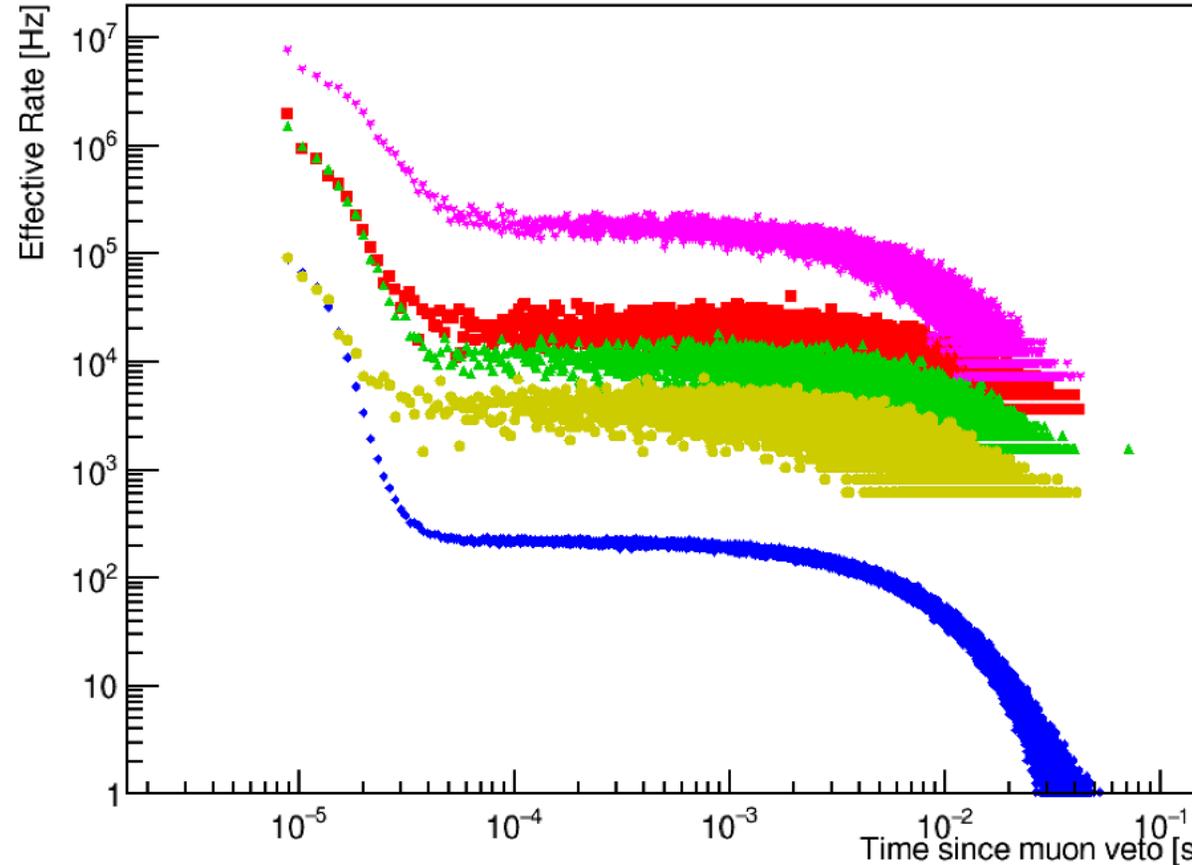
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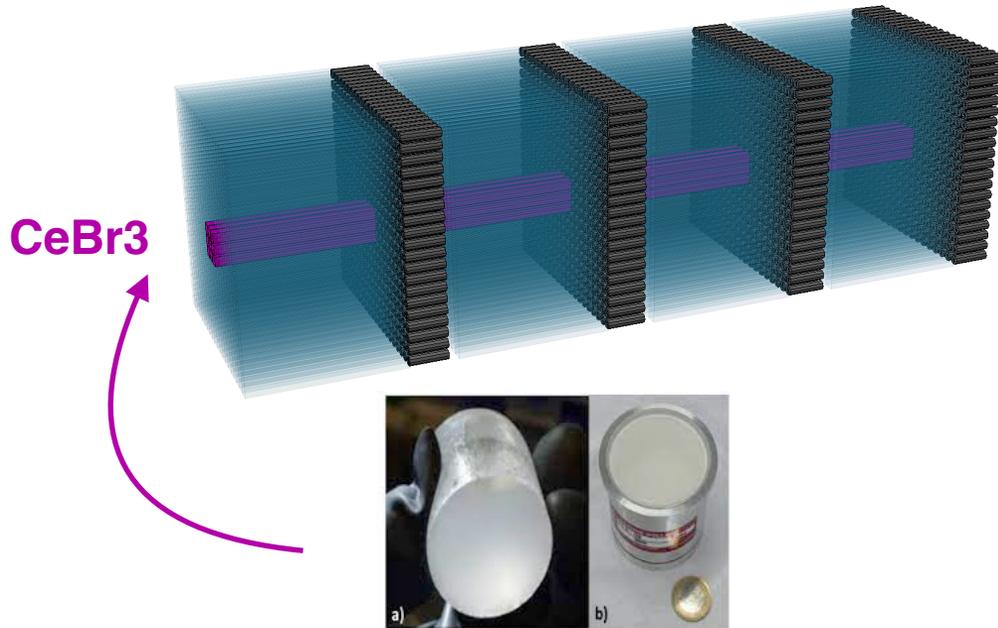
Signal triggers
4Layers
nLayers ≥ 3
nBars ≥ 3
2Layers (both)

Special run with
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Overall scaling
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**Data suggest 100 μ s deadtime is enough for signal triggers to recover back to nominal/healthy rates!
This corresponds to 1% deadtime**

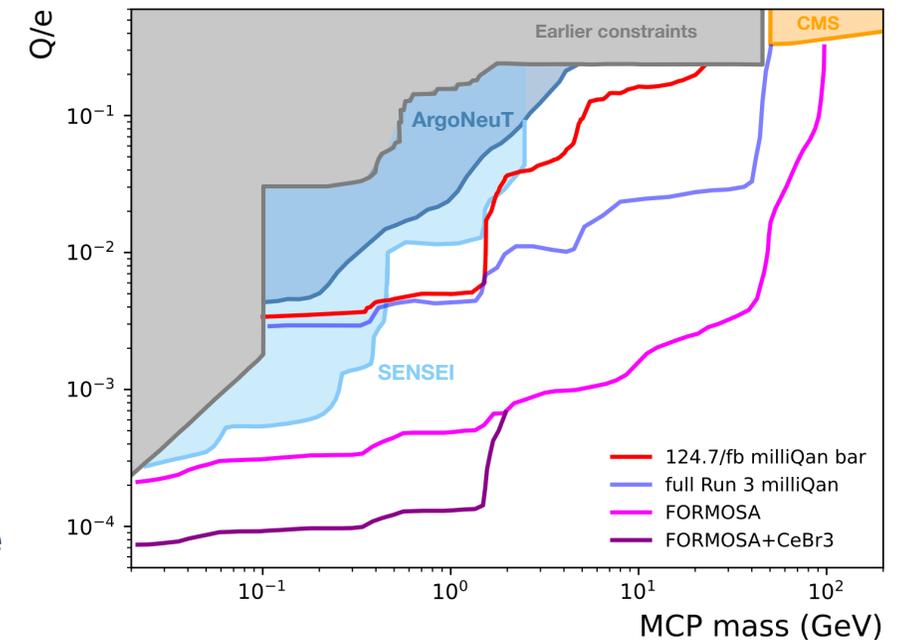
FORMOSA: CeBr3 module



Incorporated
CeBr3 module
into demonstrator
in Sept 2024

FORMOSA Simulation

2000 fb⁻¹ (13.6 TeV)

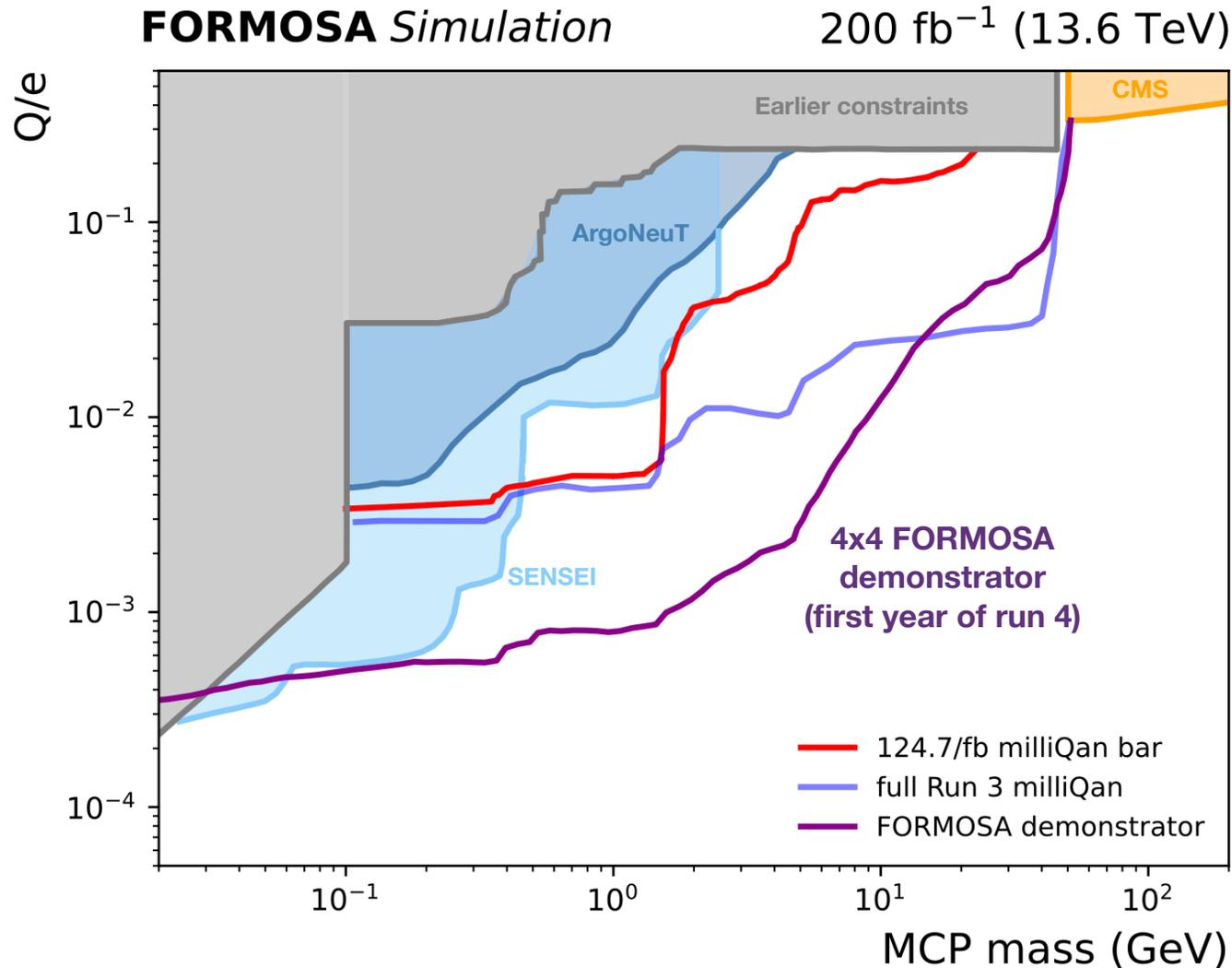


Investigate feasibility of FORMOSA subdetector made from CeBr3

- Factor ~35 larger light yield for same length compared to plastic, fast with low internal radioactivity

Considerable sensitivity gain possible! Studies ongoing using test module

FORMOSA demonstrator sensitivity

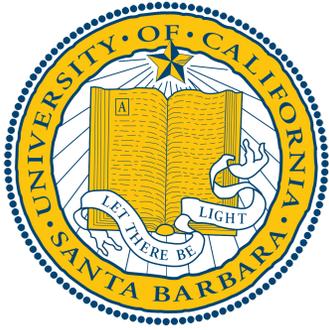


Hermetic rejection panels added in earlier this year: will attempt search with 2025 LHC data!

Expand demonstrator size for early HL-LHC if funding allows

The FORMOSA collaboration

Members from 8 institutions and growing!



This speaker supported by funding from DOE Office of Science

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Summary

- Demonstrator installed last year and taking data through 2024
- DAQ (hardware and firmware) validated
 - Alignment with L.O.S. to I.P. validated
 - Verified we are measuring physics correlated to LHC activity
 - Capable of identifying and vetoing online muons generated by I.P. collisions
 - Intra-clock timing proving useful for understanding activity and backgrounds
- Hermetic side and top panels installed to improve background understanding
- CeBr3 test module installed to carry out extended studies
- Demonstrator currently taking data stably!

End



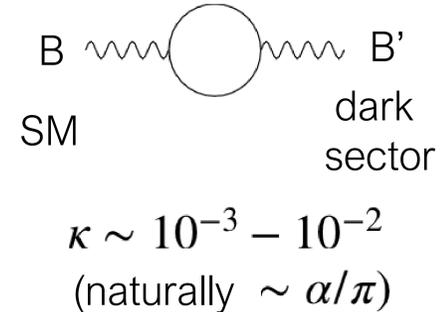
Backup

Why millicharged particles?

Standard motivation: Introduce new, hidden $U(1)$ with a massless field A' , a “dark photon” that couples to a massive “dark fermion” ψ'

$$\mathcal{L}_{\text{dark-sector}} = -\frac{1}{4} A'_{\mu\nu} A'^{\mu\nu} + i\bar{\psi}' (\underbrace{\gamma^\mu \partial_\mu + ie' \gamma^\mu A'_\mu + iM_{\text{mCP}}}_{\text{“dark fermion” with mass } M_{\text{mCP}}, \text{ charge } e'}) \psi' - \frac{\kappa}{2} A'_{\mu\nu} B^{\mu\nu}$$

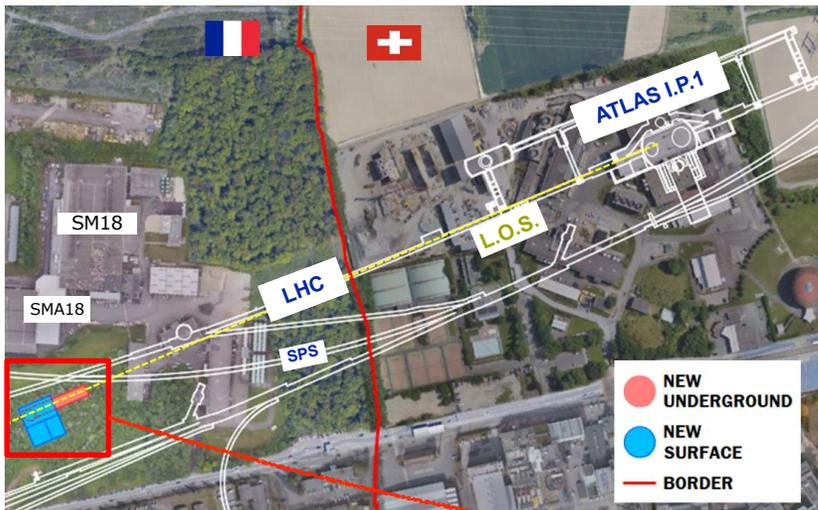
massless “dark photon”
mixing term



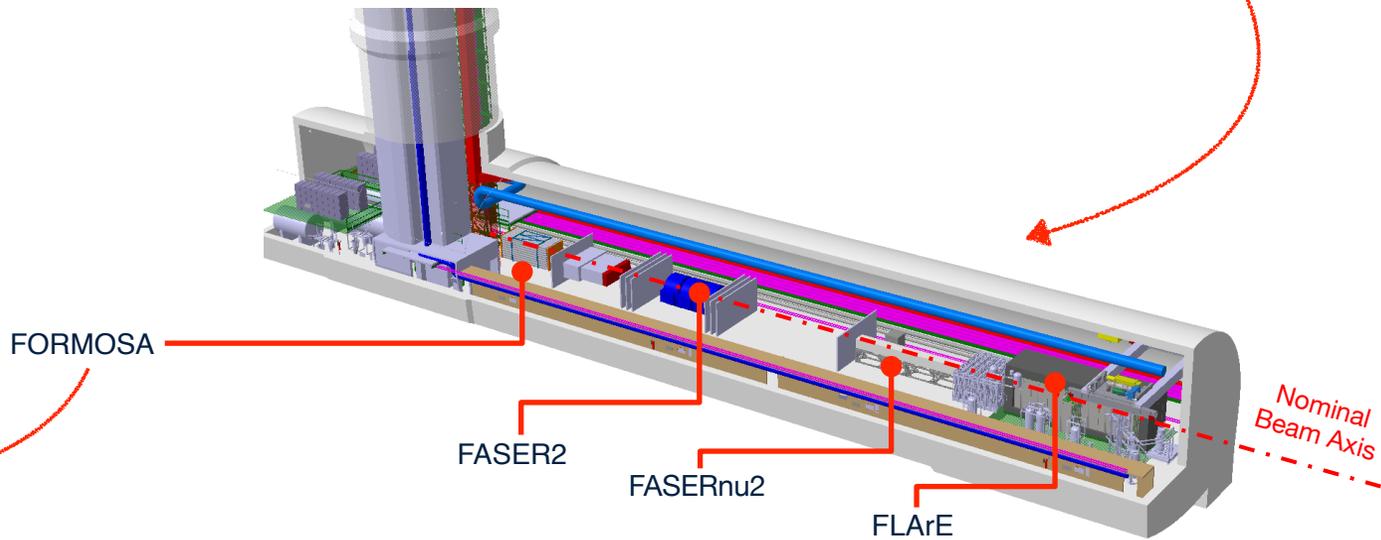
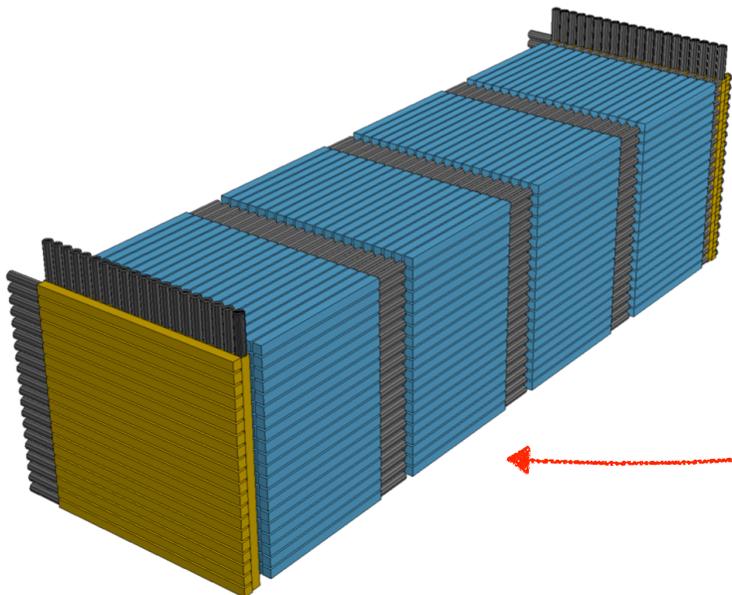
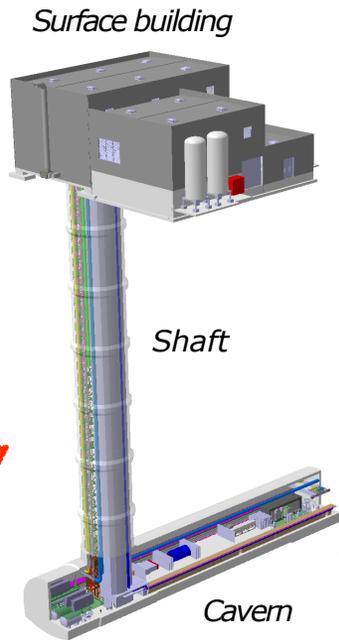
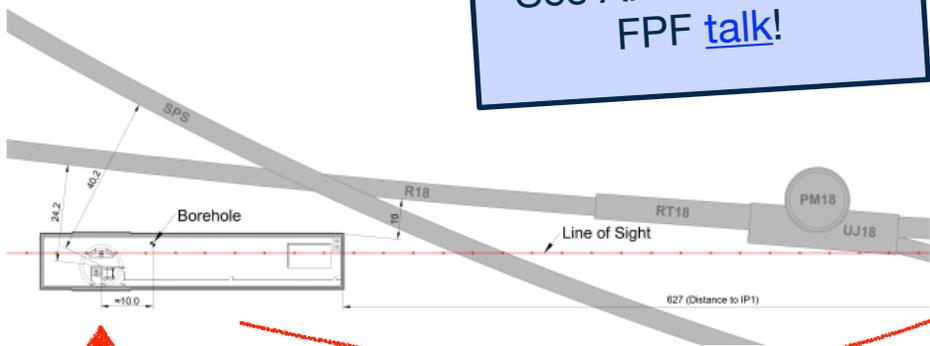
- ψ' has mass M_{mCP} and charge under the new $U(1)$ of e'
- Gauge transformation of $A'_\mu \rightarrow A'_\mu + \kappa B_\mu$ introduces coupling $\bar{\psi}' \kappa e' \gamma^\mu B_\mu \psi'$
- Conclusion: Coupling arises between dark fermion and SM photon of charge $\kappa e' \cos \theta_W$. **mCP parameters are entirely defined by their mass and charge**

see e.g. [arXiv:2104.07151v2](https://arxiv.org/abs/2104.07151v2) for more details

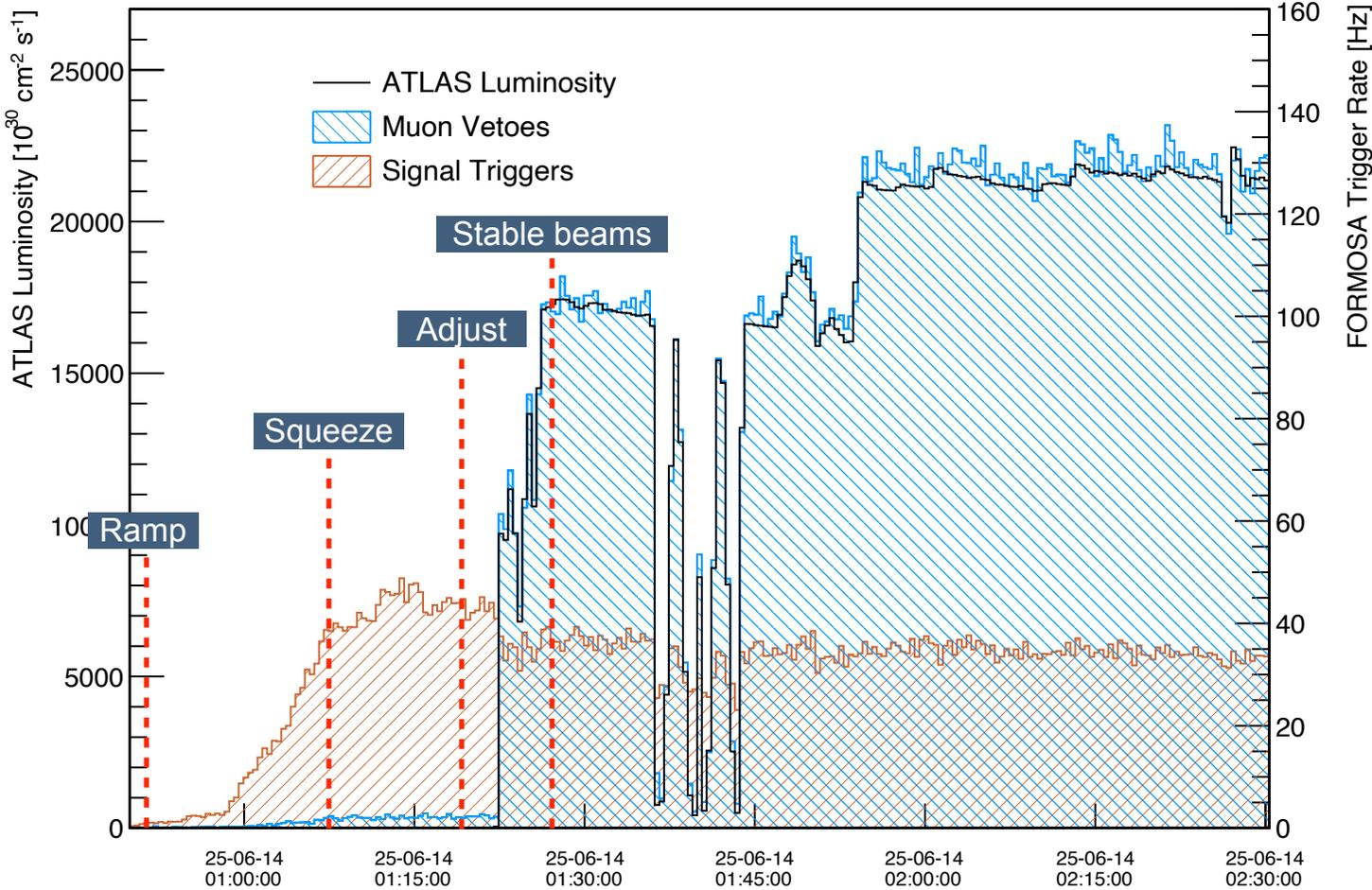
FORMOSA at the FPF



See Akitaka Ariga's FPF [talk!](#)



DAQ validation



Data from the
14th of June 2025

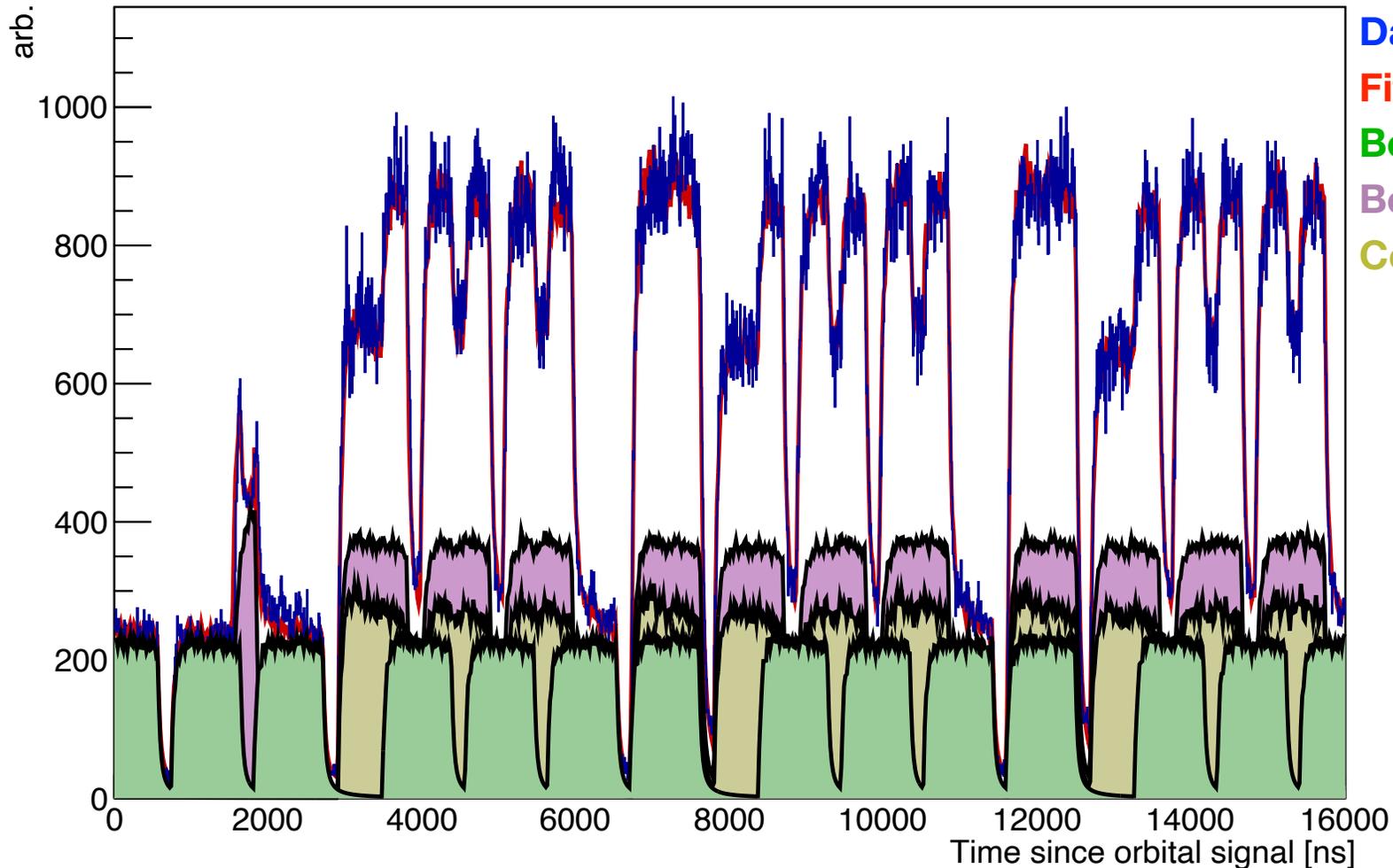
**Signal triggers stably
collecting data**

**Muons very
effectively vetoed!**



Timing studies: probing the LHC's fill scheme

Comparing FORMOSA Background Data to Circulating Beam Model



Data

Fit

Beam 1 contribution

Beam 2 contribution

Collision contribution

Fit function:

$$f(t) = A_1 * \sum_{i=0}^{Beam1} (I_{1,i} \cdot L(t, \mu_i + a_1, \sigma_1)) \\ + A_2 * \sum_{j=0}^{Beam2} (I_{2,j} \cdot L(t, \mu_j + a_2, \sigma_2)) \\ + A_3 * \sum_{k=0}^{Collisions} ((I_{1,k} \times I_{2,k}) \cdot L(t, \mu_k + a_3, \sigma_3)) \\ + C$$