Results from, and Status of, the LUX-ZEPLIN Experiment





Amy Cottle, UCL



LZ COLLABORATION

Thanks to our sponsors and participating institutions! 38 Institutions: 250 scientists, engineers, and technical staff

- Black Hills State University
- Brookhaven National Laboratory
- Brown University
- Center for Underground Physics
- Edinburgh University
- Fermi National Accelerator Lab.
- Imperial College London
- King's College London
- Lawrence Berkeley National Lab.
- Lawrence Livermore National Lab.
- LIP Coimbra
- Northwestern University
- Pennsylvania State University
- Royal Holloway University of London
- SLAC National Accelerator Lab.
- South Dakota School of Mines & Tech
- South Dakota Science & Technology Authority
- STFC Rutherford Appleton Lab.
- Texas A&M University
- University of Albany, SUNY
- University of Alabama
- University of Bristol
- University College London
- University of California Berkeley
- University of California Davis
- University of California Los Angeles
- University of California Santa Barbara
- University of Liverpool
- University of Maryland
- University of Massachusetts, Amherst
- University of Michigan
- **University of Oxford**
- University of Rochester
- University of Sheffield
- University of Texas at Austin
- University of Wisconsin, Madison
- University of Zürich







Science and **Technology Facilities Council**

Amy Cottle - EPS-HEP 2025



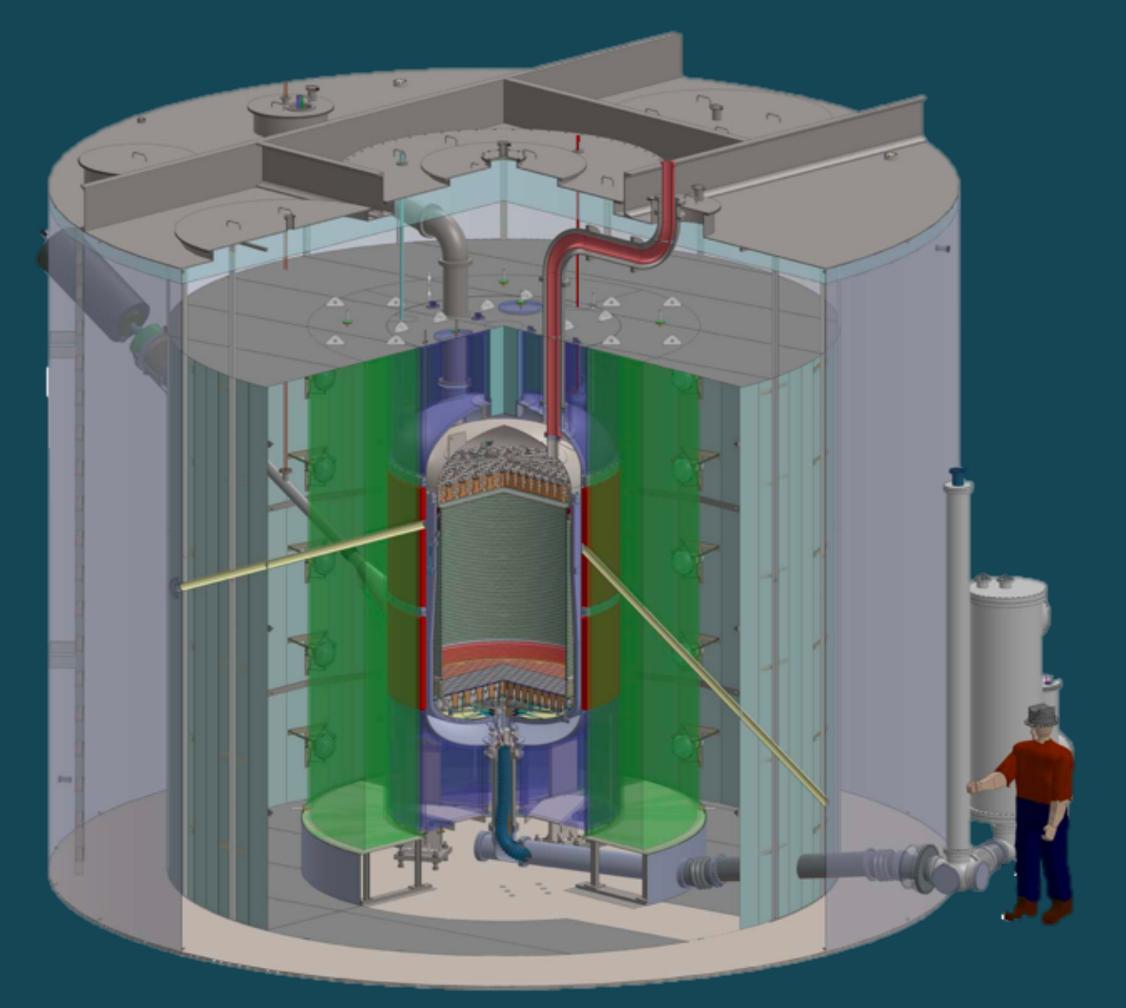


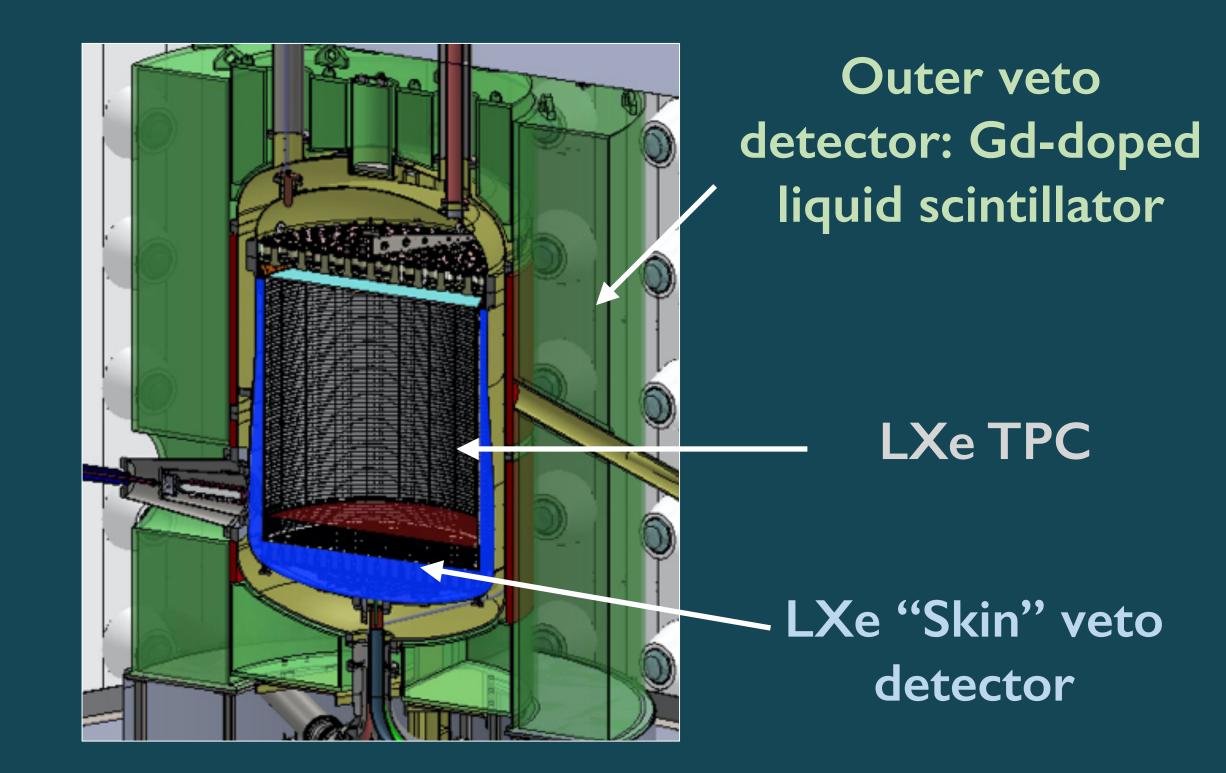






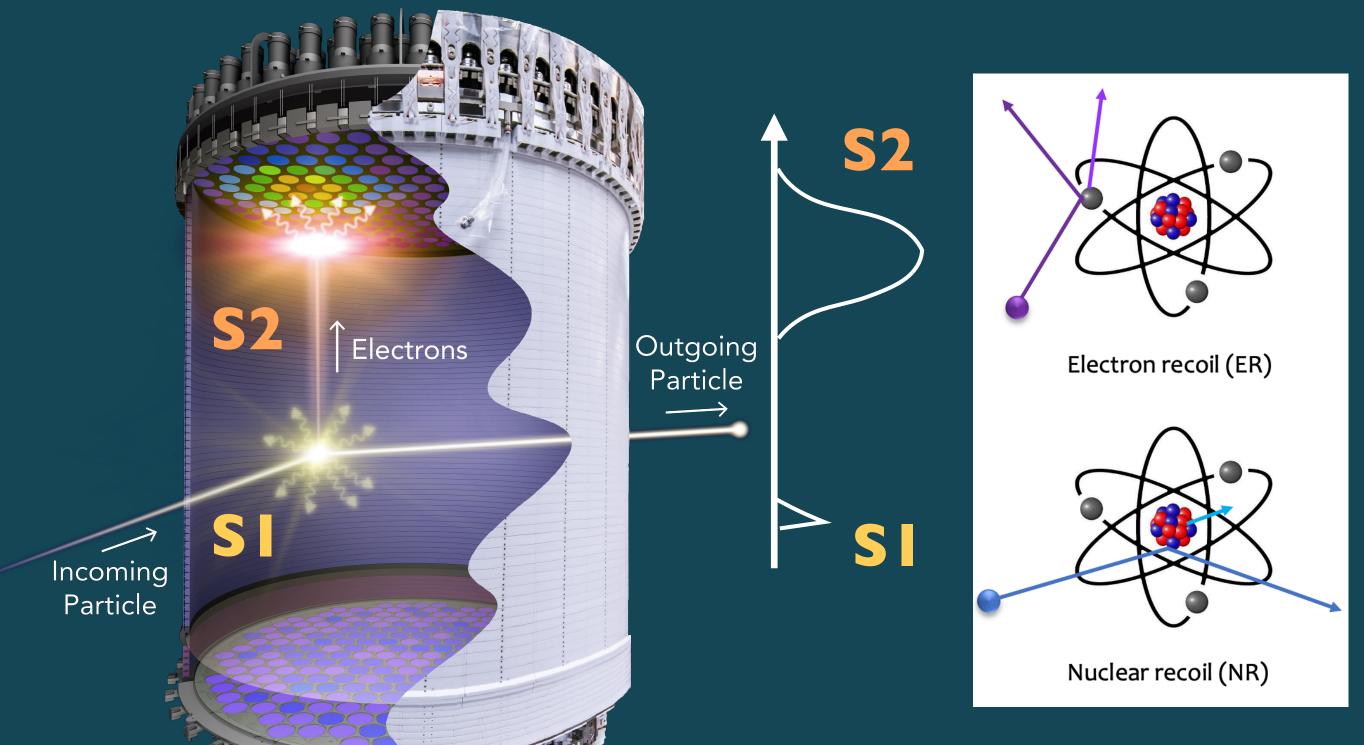
THE LZ EXPERIMENT

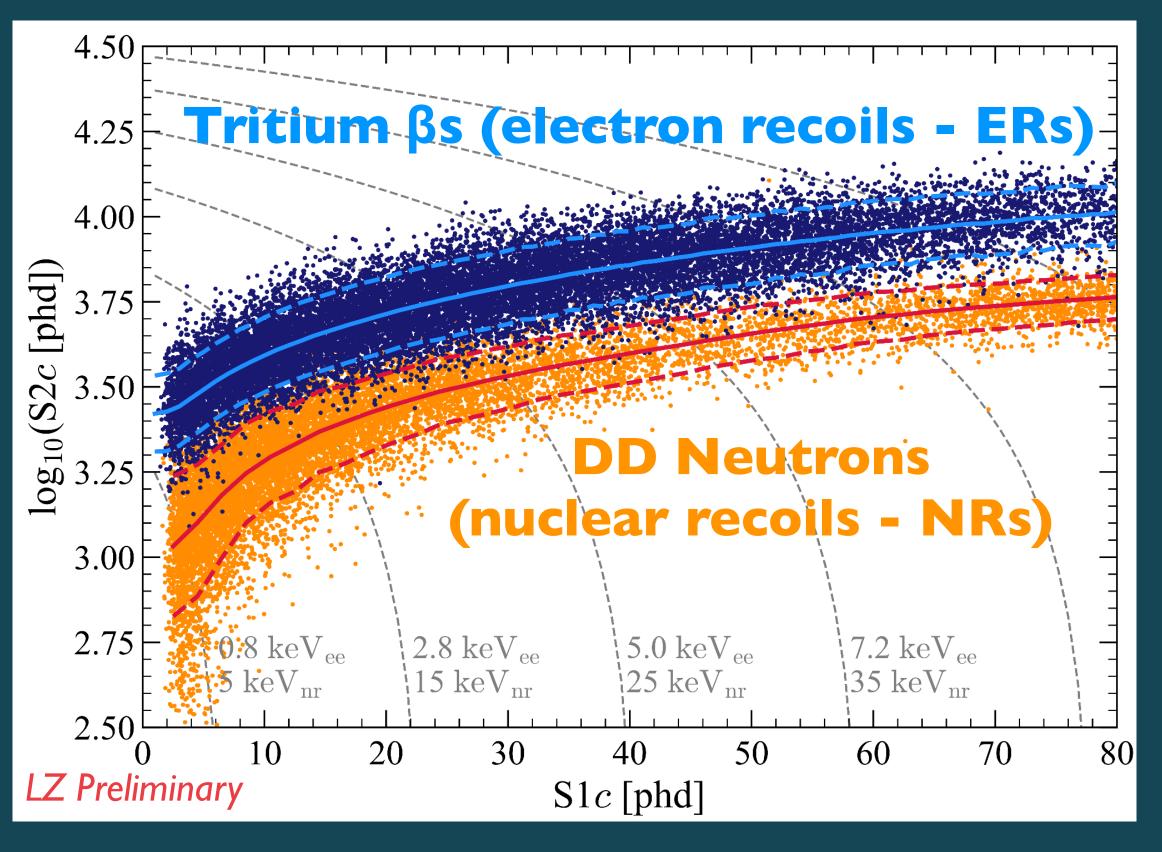




- ~1.5 km underground at the Sanford Underground Research Facility (SURF) in Lead, SD
- Operating world's largest 7 t active dual-phase xenon time projection chamber (TPC)
- TPC nested in Skin & Outer Detector (OD) veto systems \rightarrow tag neutrons and gamma rays

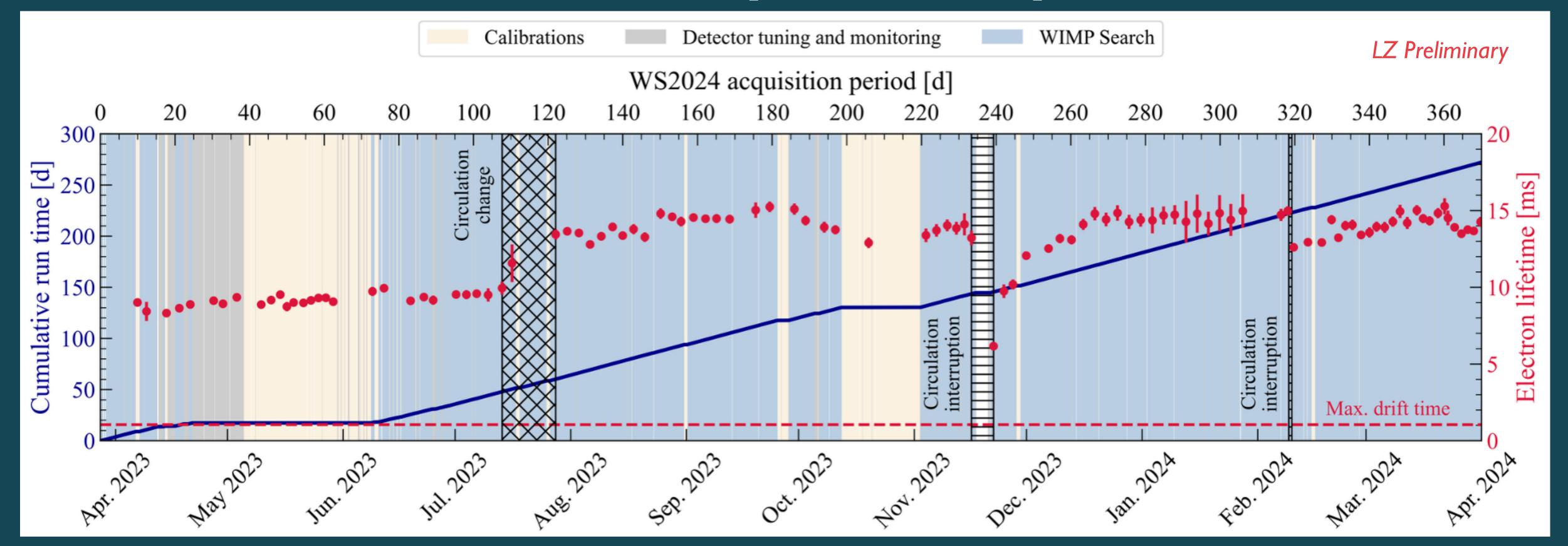
XENON TIME PROJECTION CHAMBER (TPC)





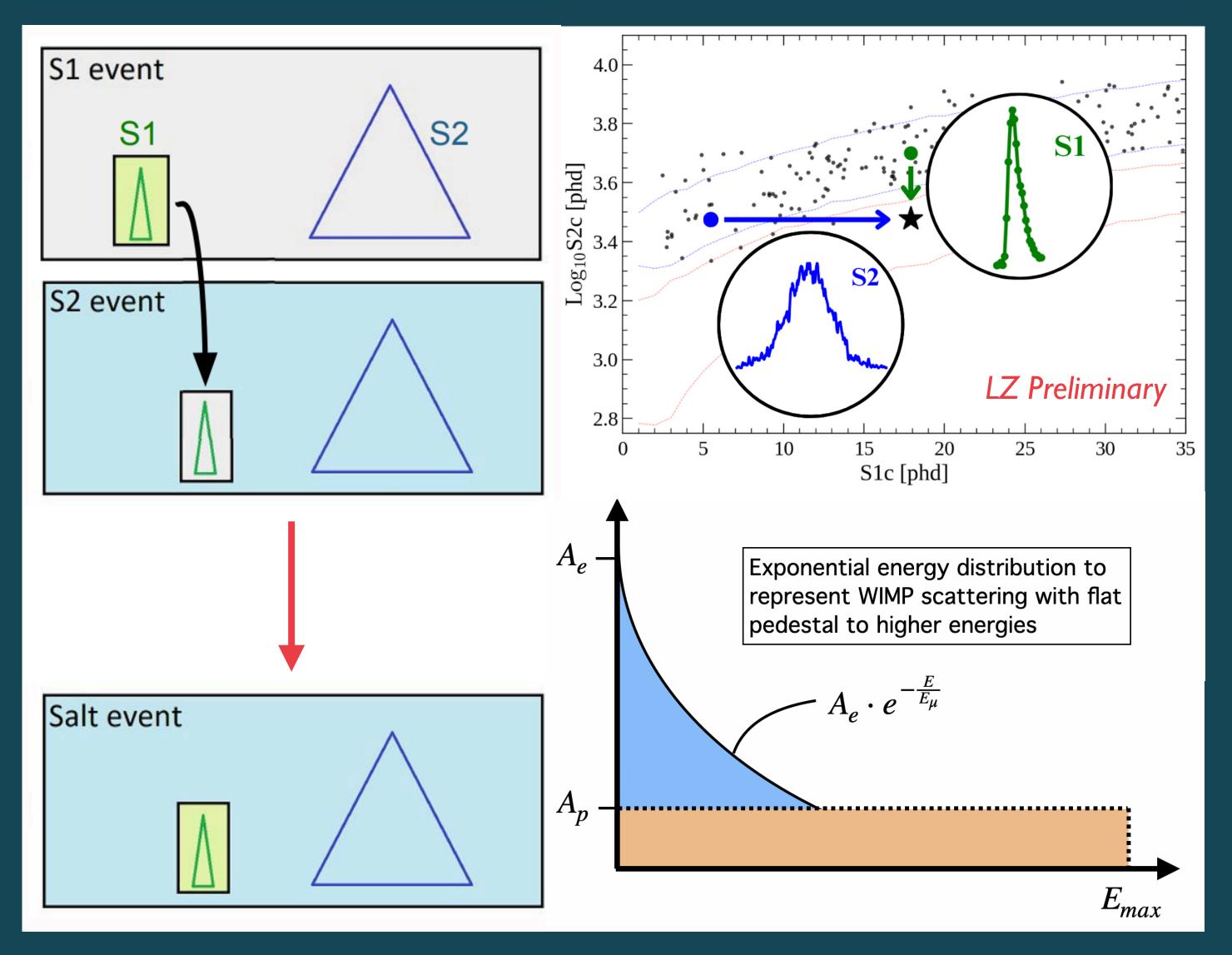
- Interactions give light (prompt scintillation \$1) and charge (proportional scintillation \$2)
- Excellent 3D position reconstruction (\sim mm) \rightarrow fiducialisation & distinguishing single scatters (SSs)
- Discrimination of electronic recoils (ERs) from potential WIMP nuclear recoils (NRs)

WIMP SEARCH 2024 (WS2024)



- 220 live-day exposure using data from March '23 to end March '24
- Major milestones: bias mitigation ("salting") began July 3rd; circulation state change July 12th

BIAS MITIGATION



- "Salting" fake signal events injected randomly during science data-taking
- Salt created using S1s & S2s from sequestered calibration data
- Parent distribution exponential
 WIMP recoil spectrum + flat pedestal
- Rate capped by WS2022 cross-section
- Parameters unknown when analysing
 → unsalting performed after all inputs

 are defined for statistical inference

XENON FLOW

 Circulation & cooling systems allow control of temperature & xenon flow

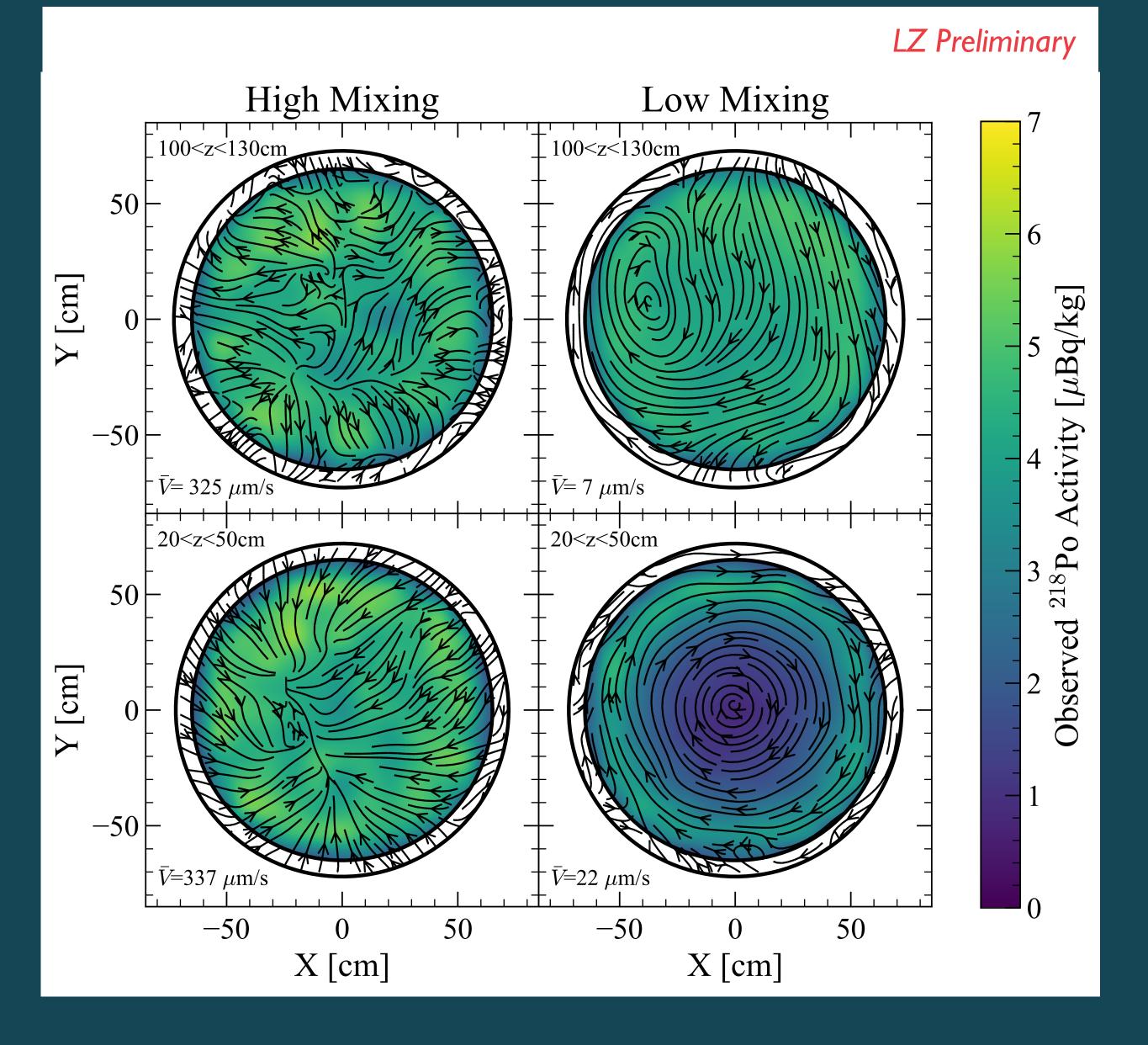
High-mixing state

More turbulent flow → uniform distribution of injected calibration sources

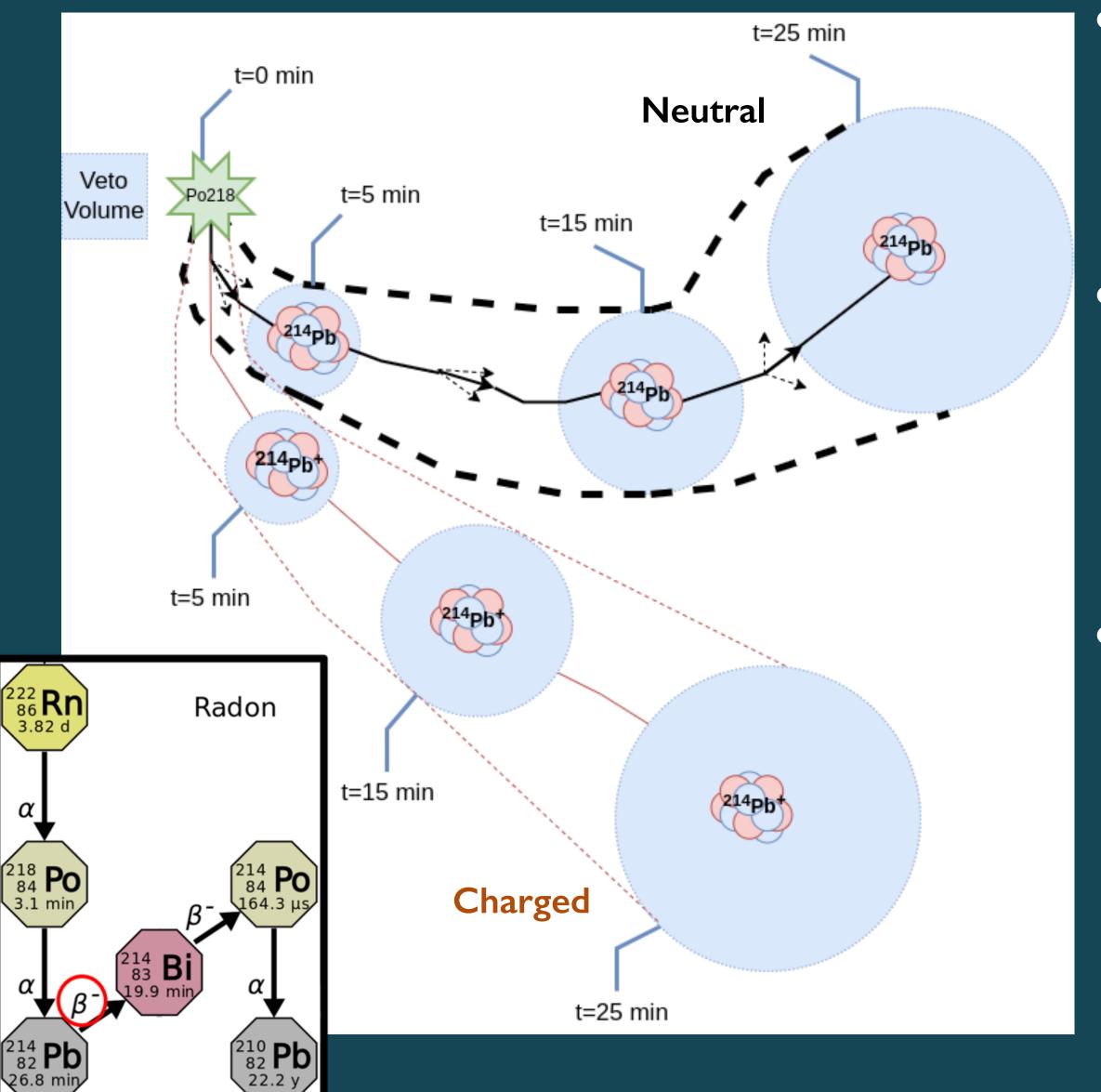
Low-mixing state

Slower, laminar flow

- ²²²Rn emanates from detector materials
- 222 Rn- 218 Po pairs ($\tau_{1/2} = 3.1 \text{ min}$)
 - → used to map the flow vectors



RADON TAG

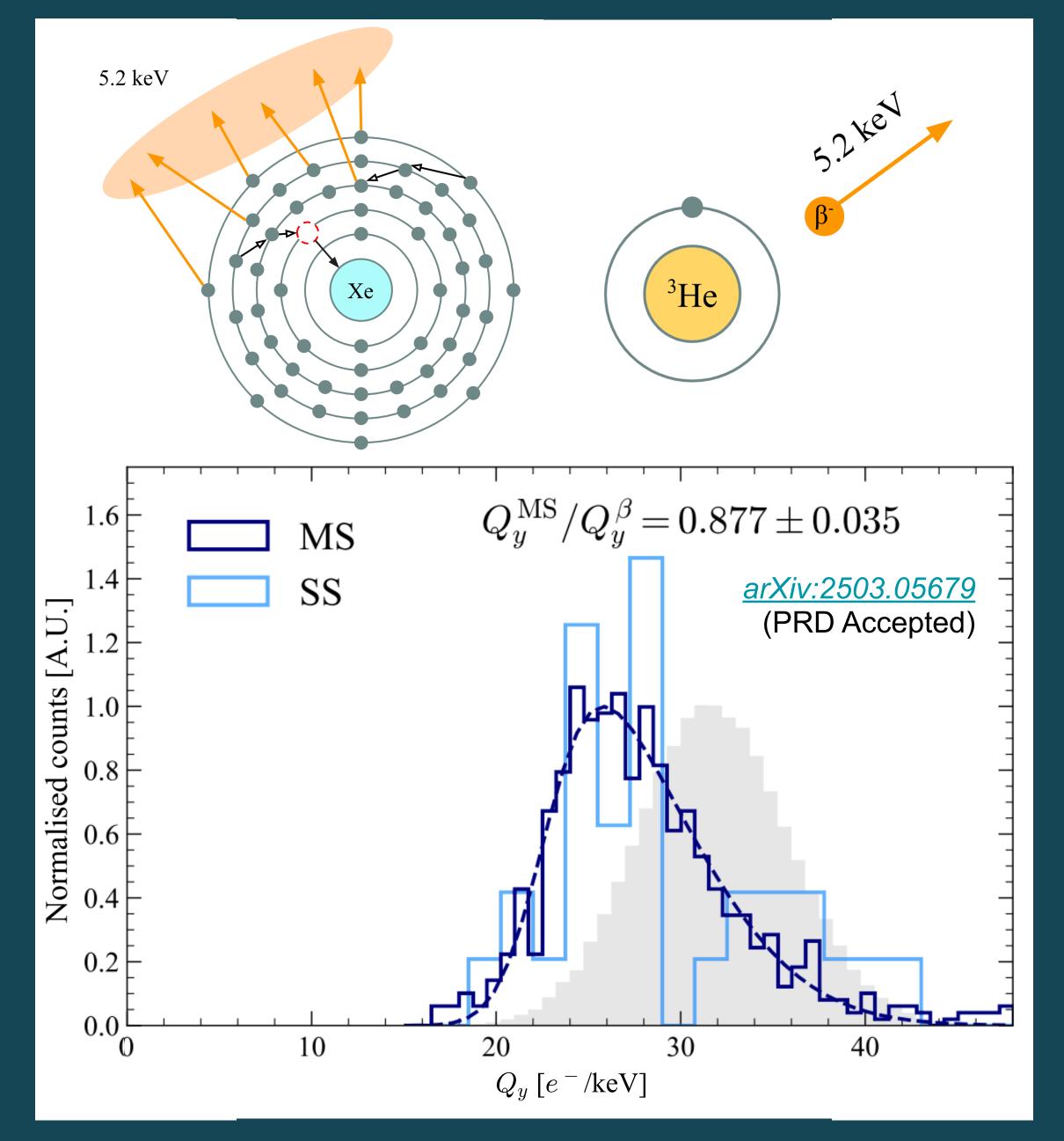


- "Naked" 214 Pb β decays = biggest ER background
- Simulations of neutral and charged ²¹⁴Pb movement using flow and field maps
 - → use to create a "radon tag" in low-mixing state
- Define co-moving volumes around "streamlines" where ²¹⁴Pb is likely to be found
 - Each volume active 81 mins (\sim 3x 214 Pb $\tau_{1/2}$)
- Tagged & untagged data both in WIMP analysis

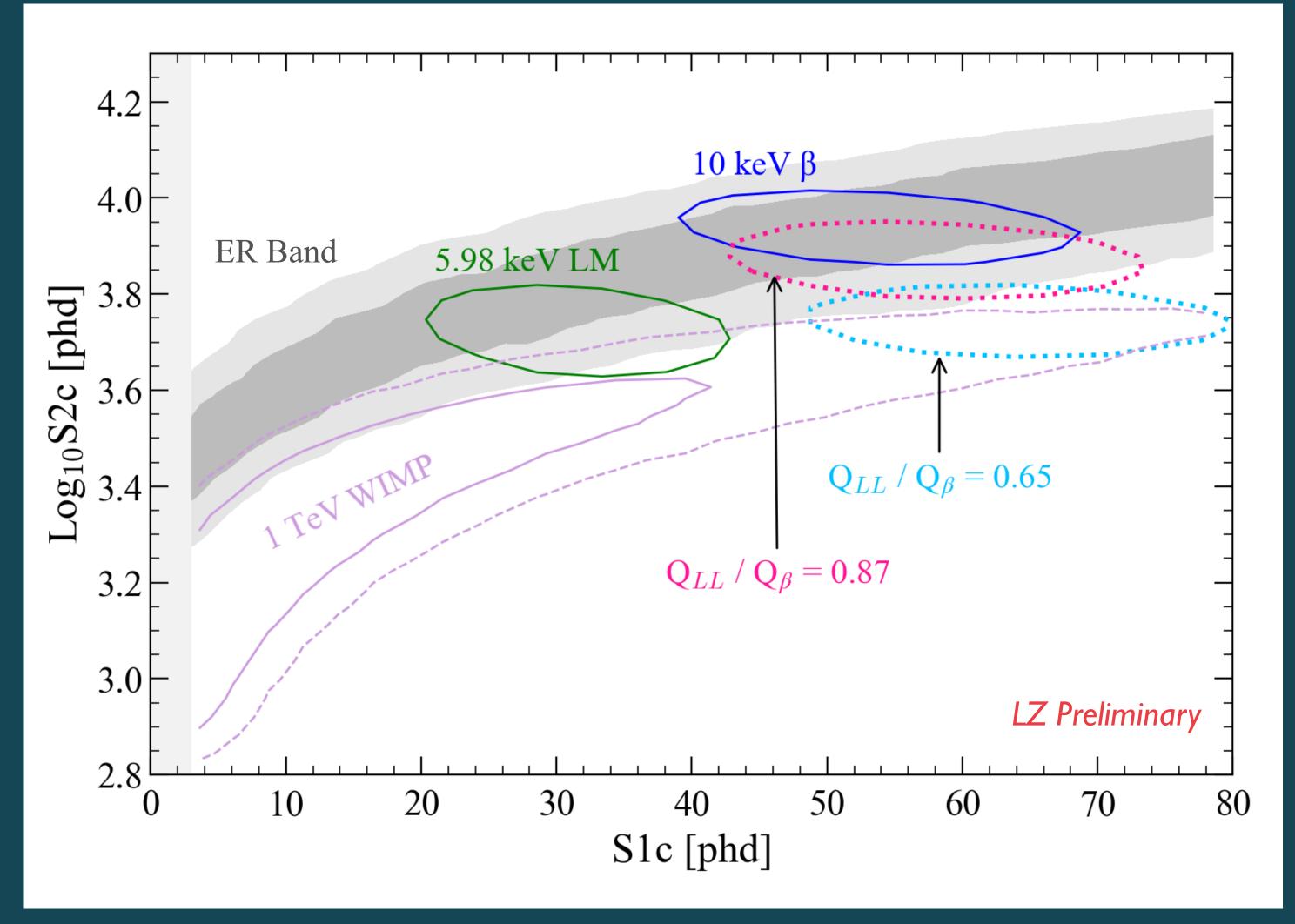
| | % ²¹⁴ Pb of Total | % Volume of Total |
|----------|------------------------------|-------------------|
| Tagged | 60 ± 4 | 15 |
| Untagged | 40 ± 4 | 85 |

ELECTRON CAPTURES

- 127Xe & 125Xe decay by electron capture (EC)
- L-shell EC (5.2 keV) relevant for WIMP search
 - Auger/X-ray cascade → more NR-like than β i.e. enhanced recombination
 - In-situ measurement in LZ for WS2024: $Q_{L}/Q_{\beta} = 0.877 \pm 0.035 \quad (arXiv:2503.05679)$
- How does this impact ¹²⁴Xe double electron capture (DEC)? "world's rarest decay"
 (LM (6.0) & LL-shell (10.0 keV) contributions)



MODELLING 124XE LM- & LL-SHELL DEC



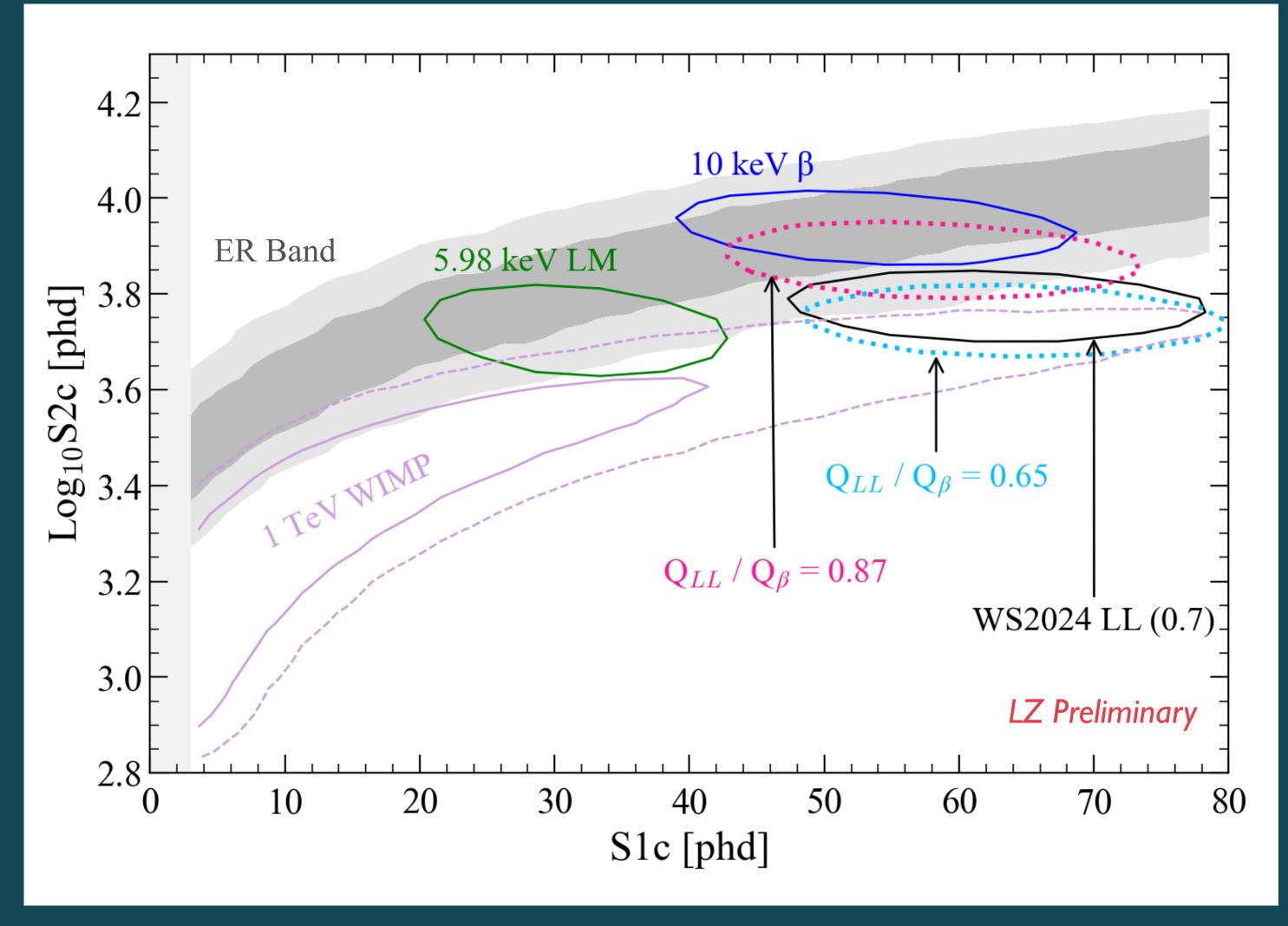
- Expect 7.1 (LM) + 12.3 (LL) = 19.4 counts with 20% uncertainty
- LM modelled with same as single
 L-shell charge suppression
- LL expected to be further charge-suppressed due to higher ionisation density i.e. Q_{LL}/Q_β < Q_L/Q_β
- Vary Q_{LL}/Q_{β} in fitting of our data:

$$0.65 < Q_{LL}/Q_{\beta} < 0.87$$

2x L-shell ionization density

 $\mathbf{Q}_{\mathsf{L}}/\mathbf{Q}_{\beta}$

MODELLING 124XE LM- & LL-SHELL DEC

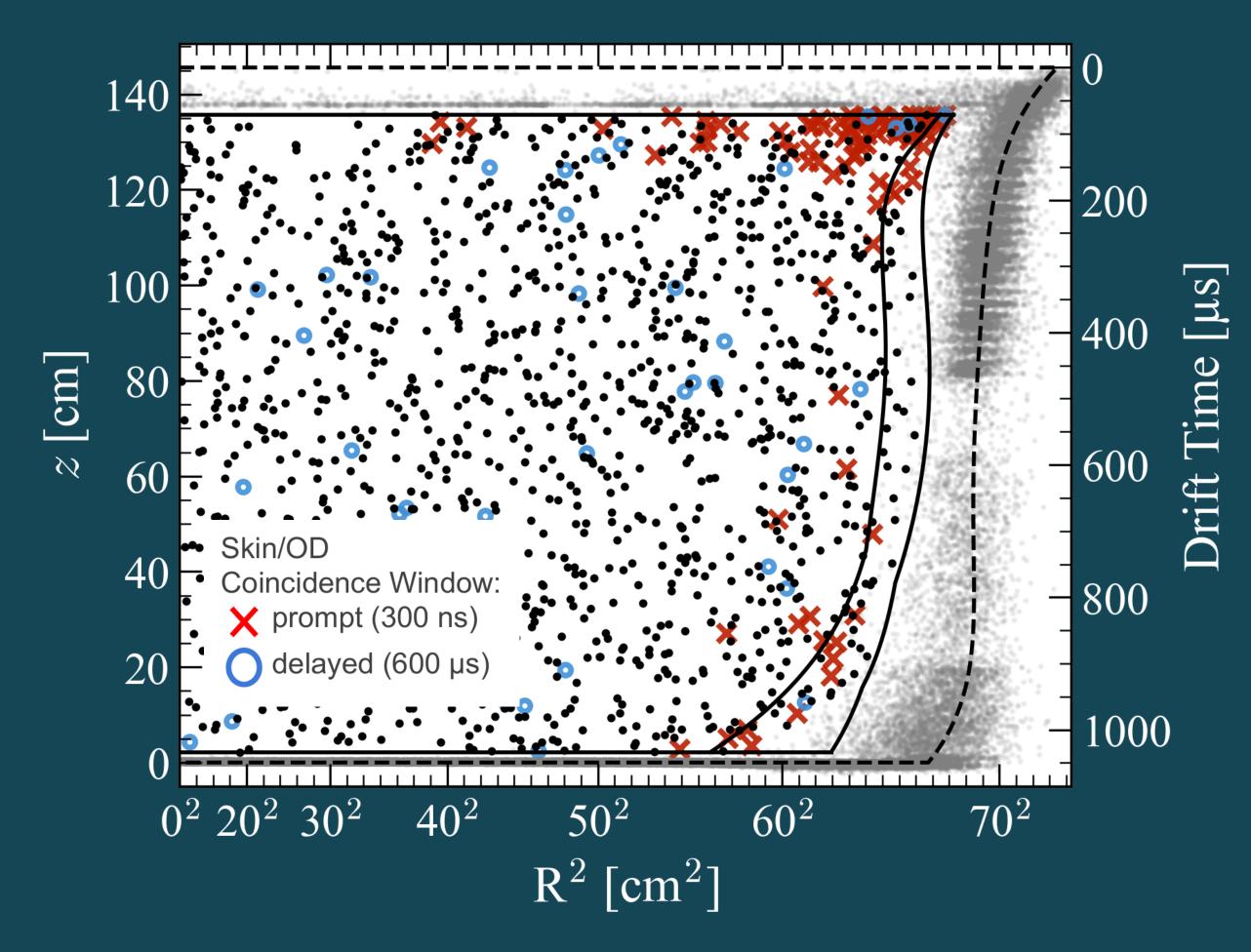


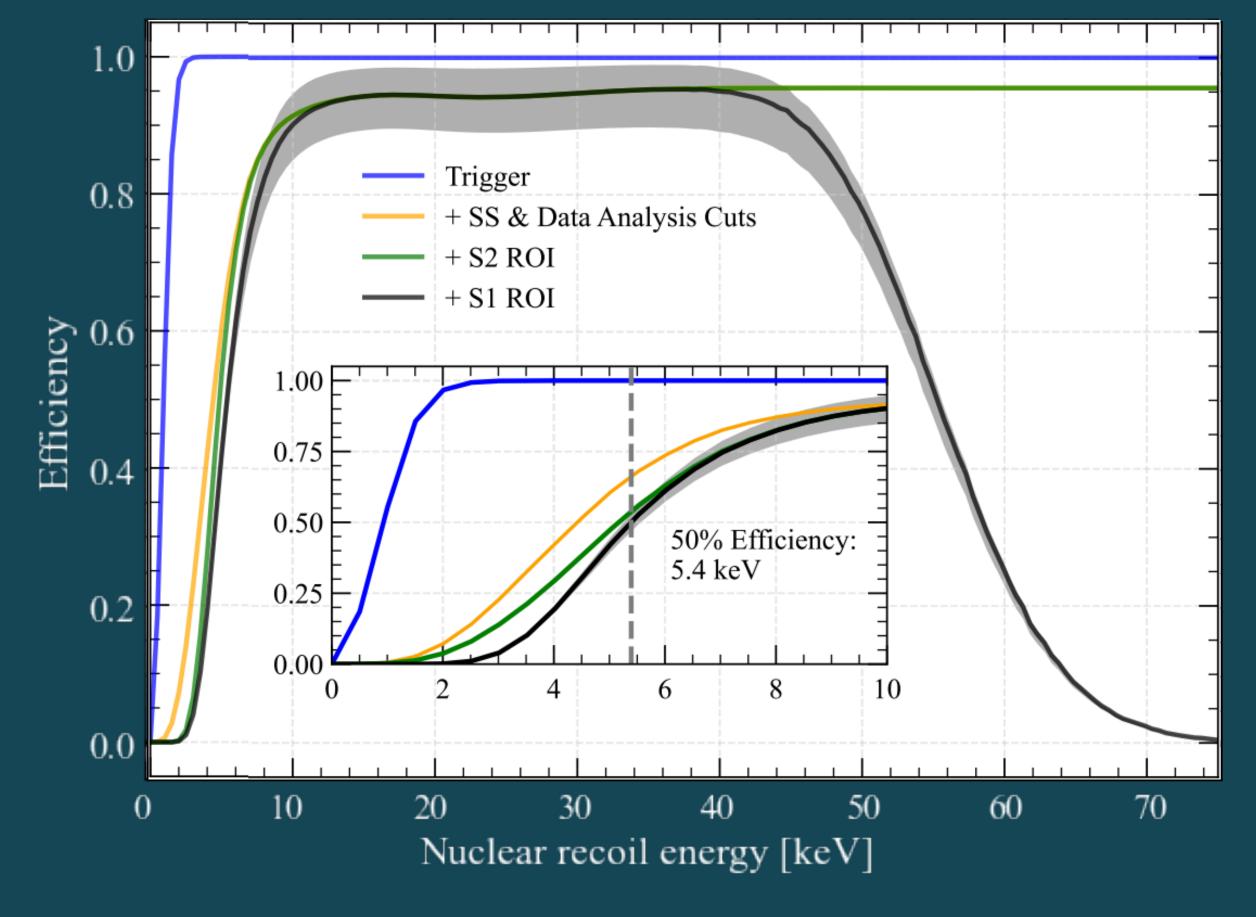
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- Vary Q_{LL}/Q_{β} in fitting of our data:

$$0.65 < Q_{LL}/Q_{\beta} < 0.87$$

Best-fit value of $Q_{LL}/Q_{\beta} = 0.70 \pm 0.04$

ANALYSIS CUT SUMMARY

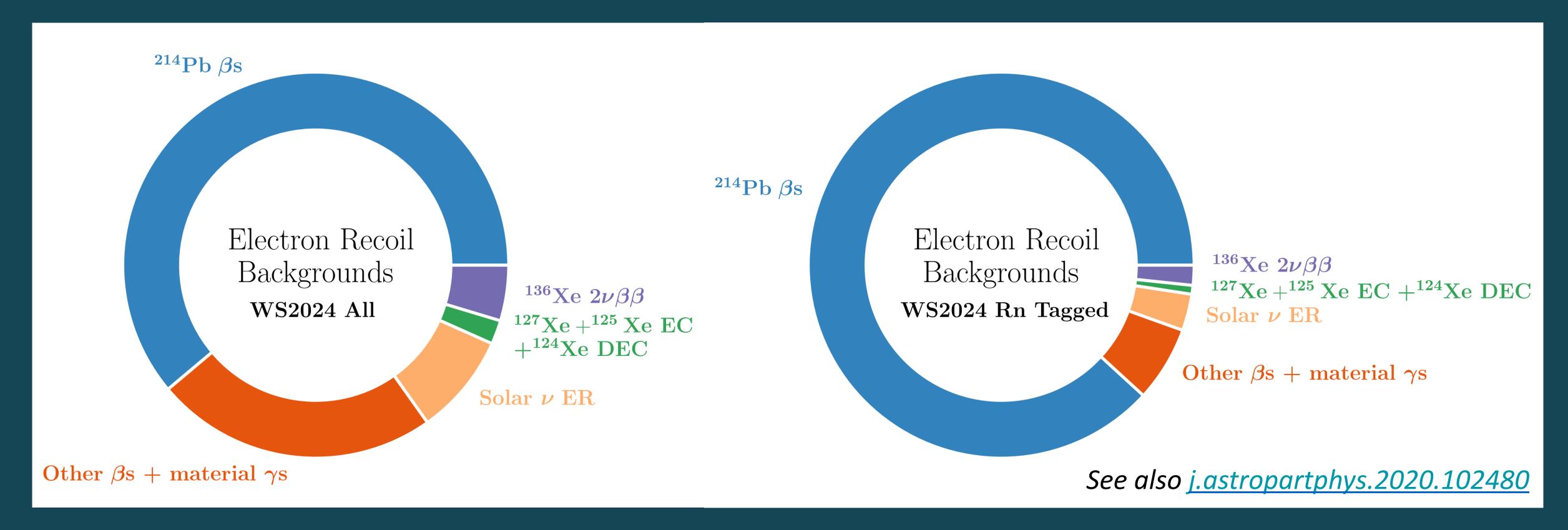




- Fiducial volume (5.5 ± 0.2 t)
- Veto detector anti-coincidence

- WIMP S2-S1 region of interest
- SI- & S2-based cuts (target accidentals)

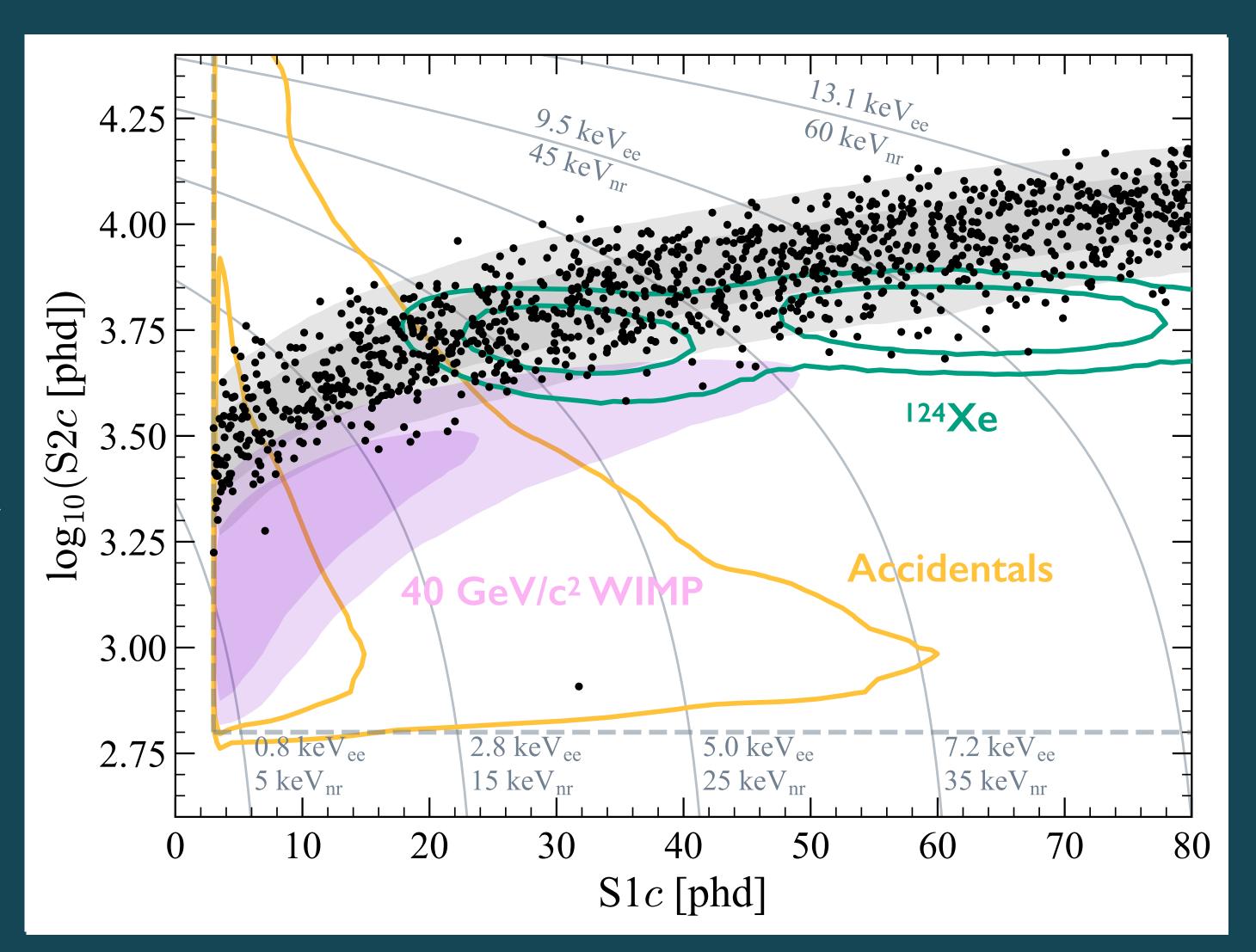
BACKGROUNDS MODEL EXPECTATIONS



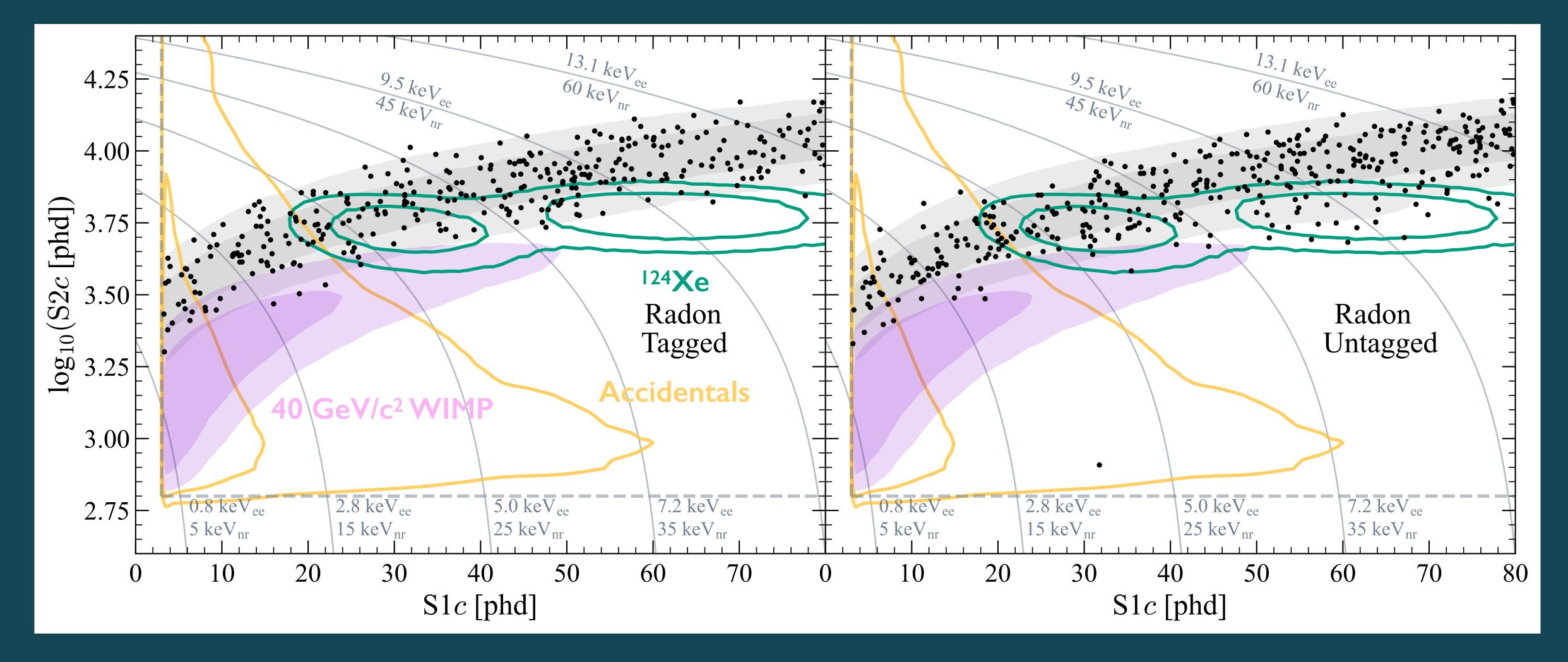
- Total expected ER counts for WS2024: I 207 (60% 214Pb) (327/88% 214Pb in radon-tagged data)
- Total expected NR counts in WS2024: 0.18 from CEVNS (no neutrons in-situ fit constraint)
- Accidentals (fake pairs of SI- and S2-identified pulses) controlled with cuts to 2.8 ± 0.6 counts

WS2024 DATA

- Final exposure of
 220 live days * 5.5 tonnes
 = 3.3 tonne years
- 7 salt events pass all analysis cuts out of 8 total injected in WS2024
 - → inline with evaluated signal efficiency
- 1220 events remain after unsalting
- Statistical analysis of these data in observed log₁₀(S2c)-S1c space
 - → no post-unsalting changes to model



WS2024 DATA - RADON TAGGED VS UNTAGGED

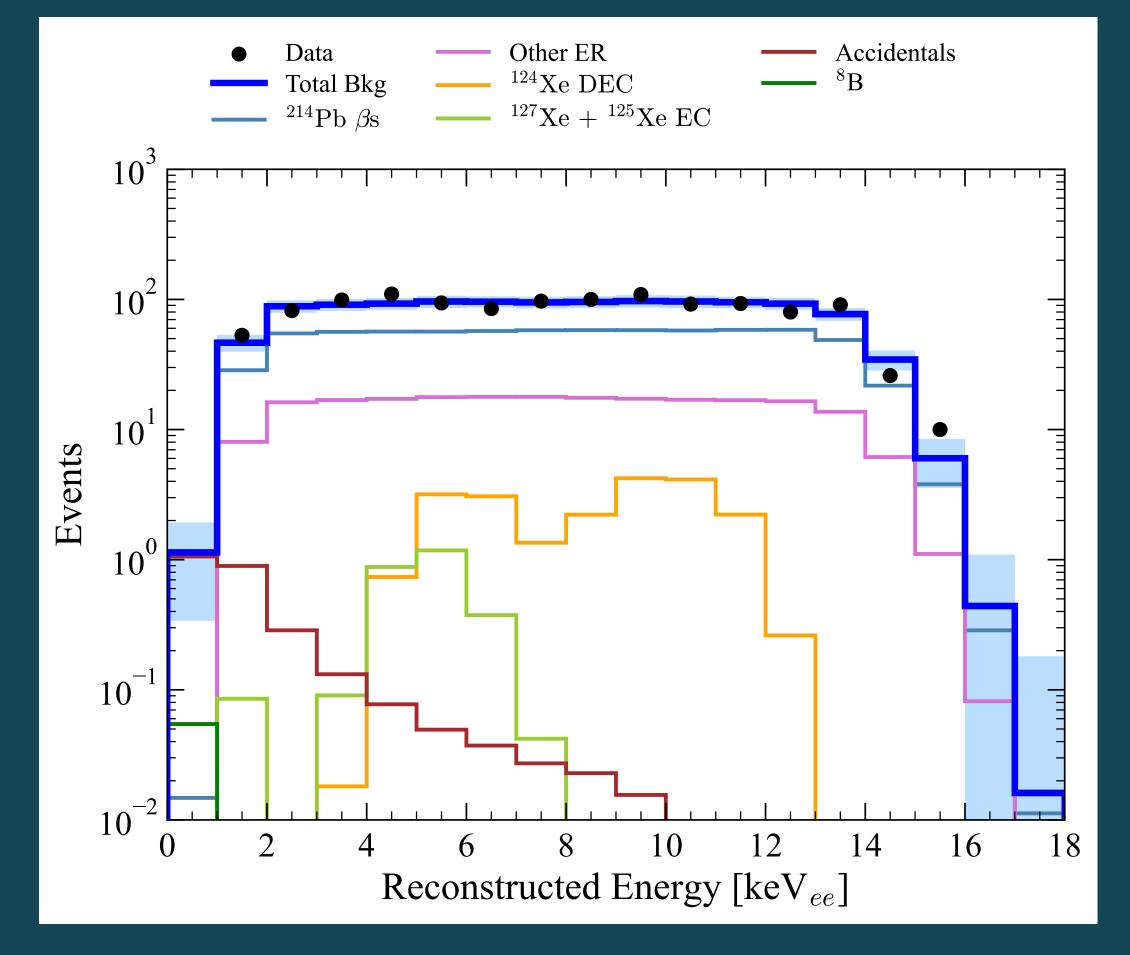


0.3 tonne years

1.8 tonne years

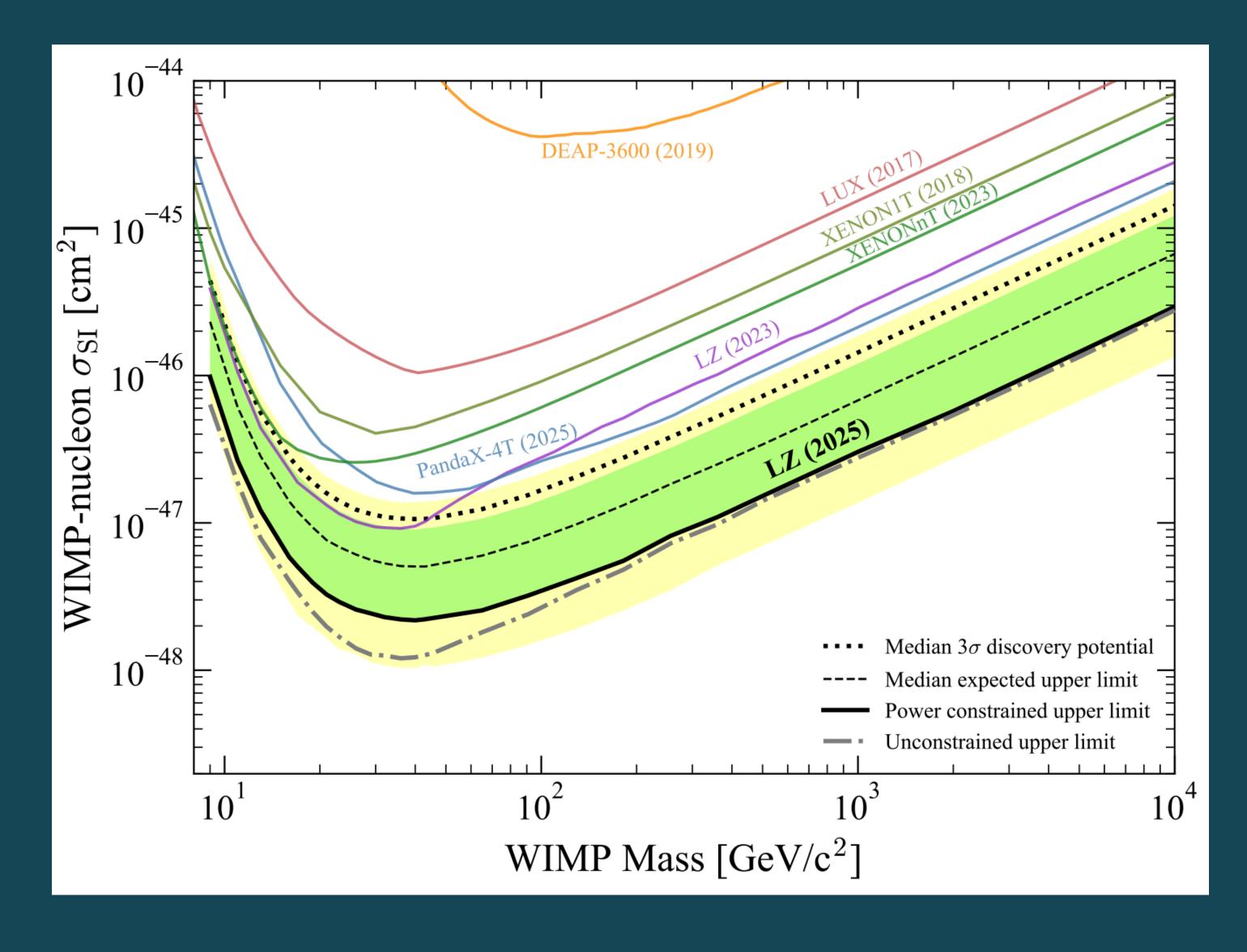
WS2024 FIT RESULTS

| Component | Expected Events | Best Fit Events |
|--|-----------------|---------------------|
| ²¹⁴ Pb β decays | 743 ± 88 | 733 ± 34 |
| ⁸⁵ Kr + ³⁹ Ar +detector γs | 162 ± 22 | 161 ± 21 |
| Solar v ERs | 102 ± 6 | 102 ± 6 |
| ²¹² Pb + ²¹⁸ Po β decays | 62.7 ± 7.5 | 63.7 ± 7.4 |
| ³ H + ¹⁴ C β decays | 58.3 ± 3.3 | 59.7 ± 3.3 |
| ¹³⁶ Xe 2vββ decay | 55.6 ± 8.3 | 55.9 ± 8.2 |
| ¹²⁴ Xe DEC | 19.4 ± 2.5 | 20.4 ± 2.4 |
| ¹²⁷ Xe + ¹²⁵ Xe EC | 3.2 ± 0.6 | 2.7 ± 0.6 |
| Atm. v CEvNS | 0.12 ± 0.02 | 0.12 ± 0.02 |
| 8B + hep v CEvNS | 0.06 ± 0.01 | 0.06 ± 0.01 |
| Det. Neutrons | | 0.0 ^{+0.2} |
| Accidentals | 2.8 ± 0.6 | 2.6 ± 0.6 |
| Total | 1210 ± 91 | 1202 ± 41 |



- Best fit of zero WIMPs at all masses tested (9 GeV/c² - 100 TeV/c²)
- Good agreement with backgroundonly hypothesis in all spaces examined

WS2024+WS2022 SPIN-INDEPENDENT LIMIT



- Combining 60 days of WS2022 with 220 days of WS2024
- Two-sided profile likelihood ratio test statistic
- Power constrained at -lσ as per recommended conventions

 EPJC 81, 907 (2021)
- Best limit from combined analysis of σ_{SI} = 2.2 × 10⁻⁴⁸ cm² for 40 GeV/c²
- Results & analysis in paper
 PRL 135, 011802 (2025)

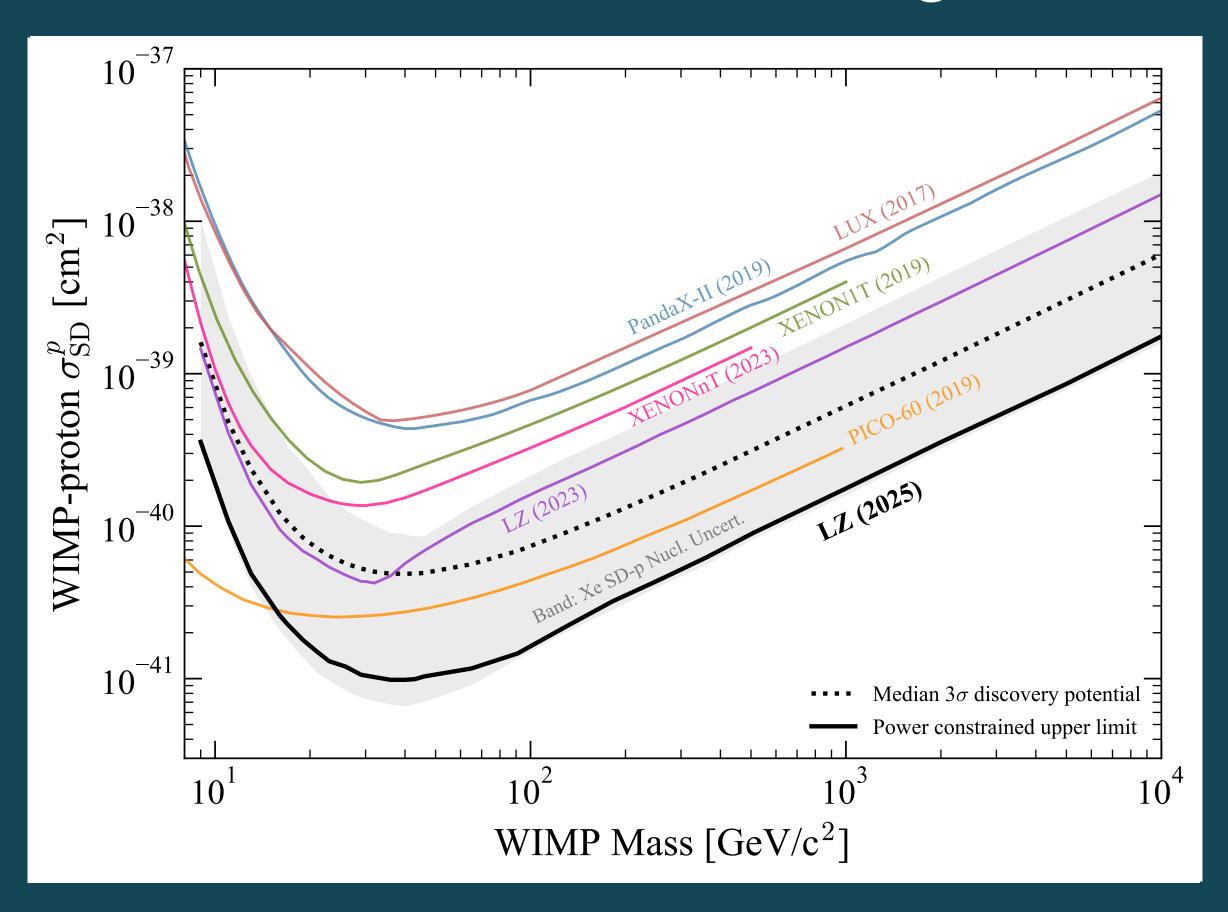
WS2024+WS2022 SPIN-DEPENDENT LIMITS

WIMP-Neutron Scattering

10⁻³⁹ 10⁻⁴⁰ 10⁻⁴¹ Rand: Xe SD-n Nucl. Uncert. Median 3σ discovery potential

WIMP Mass [GeV/c²]

WIMP-Proton Scattering

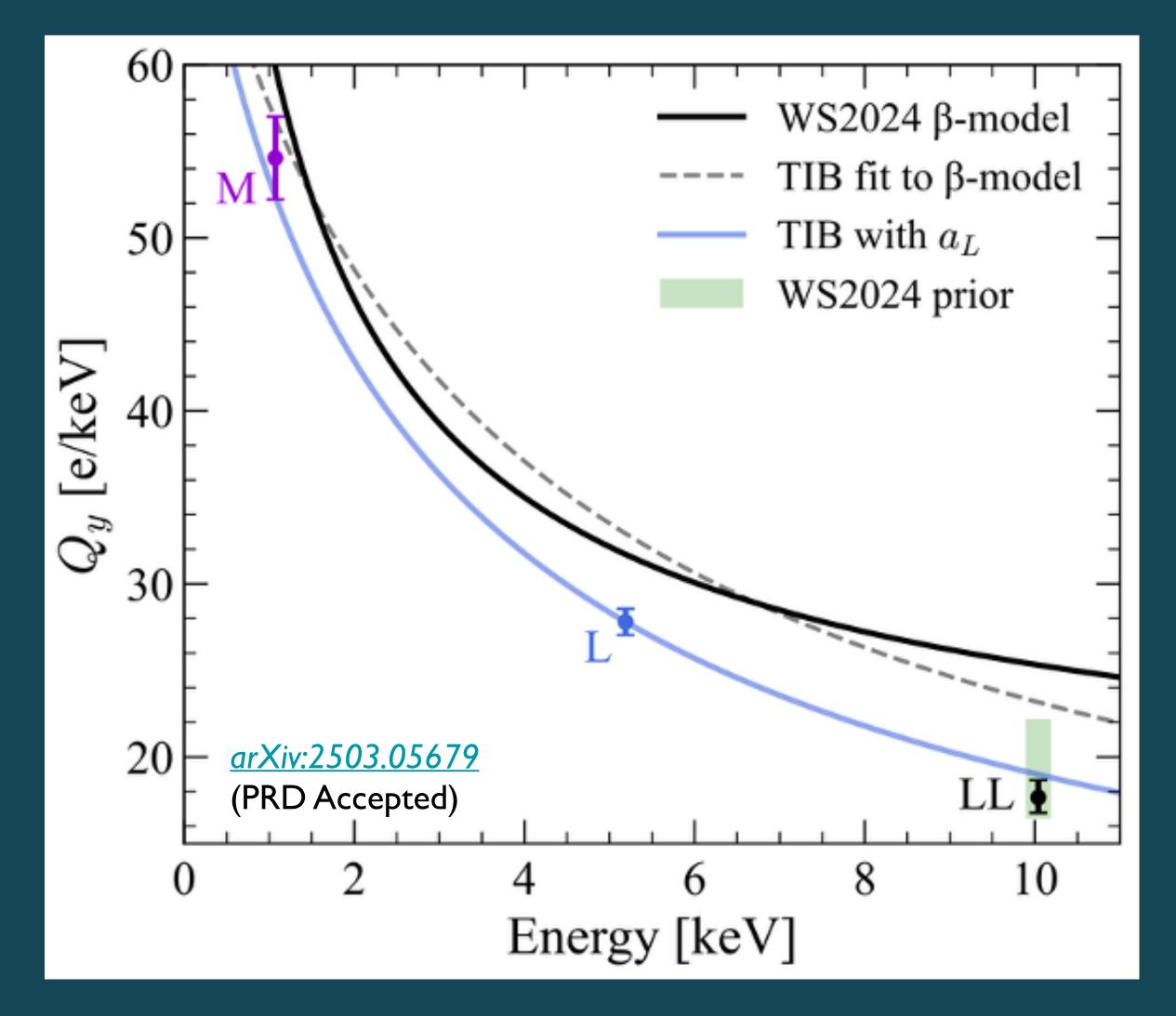


Uncertainty bands represent the theoretical uncertainty on the Xe nuclear structure factor

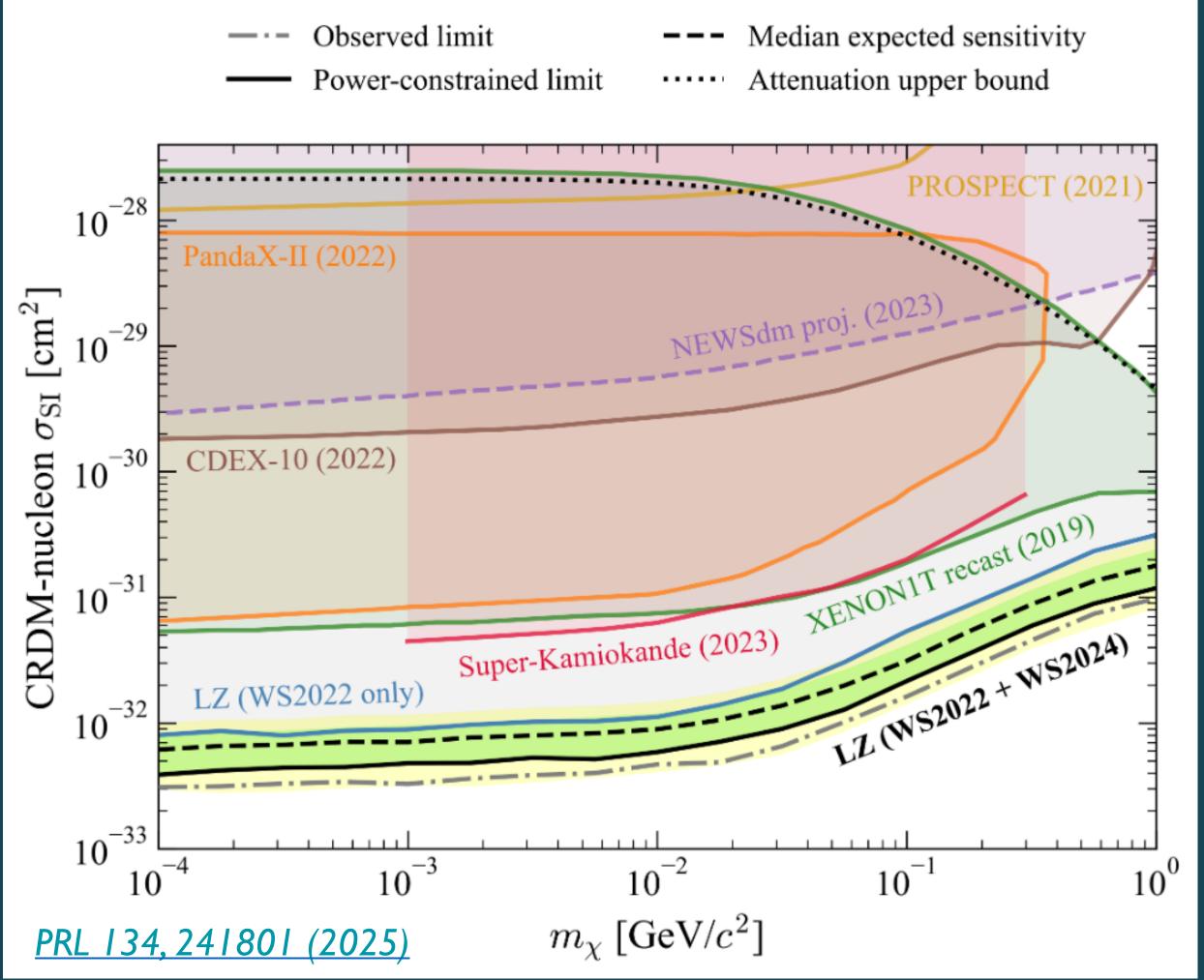
Power constrained upper limit

MORE RECENT PUBLICATIONS

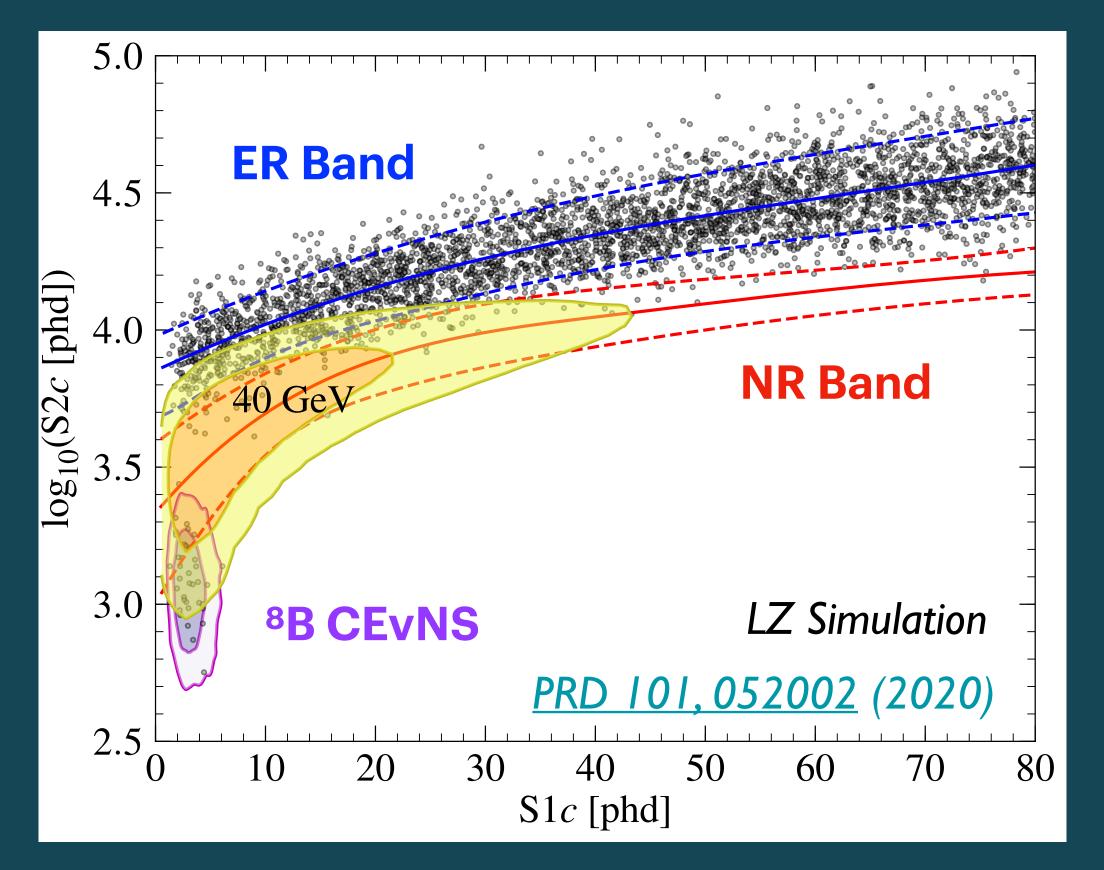
Measurements of Enhanced Recombination

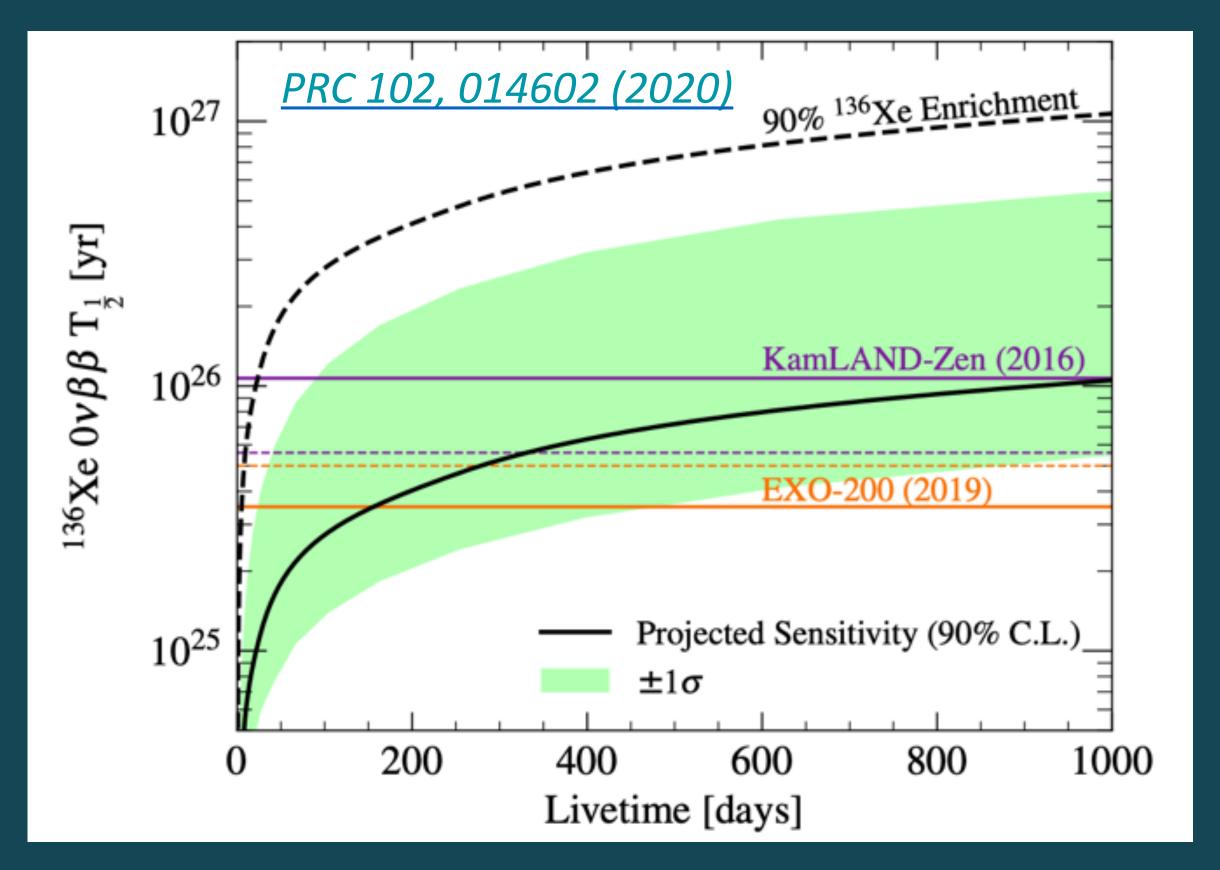


Cosmic Ray-Boosted Dark Matter



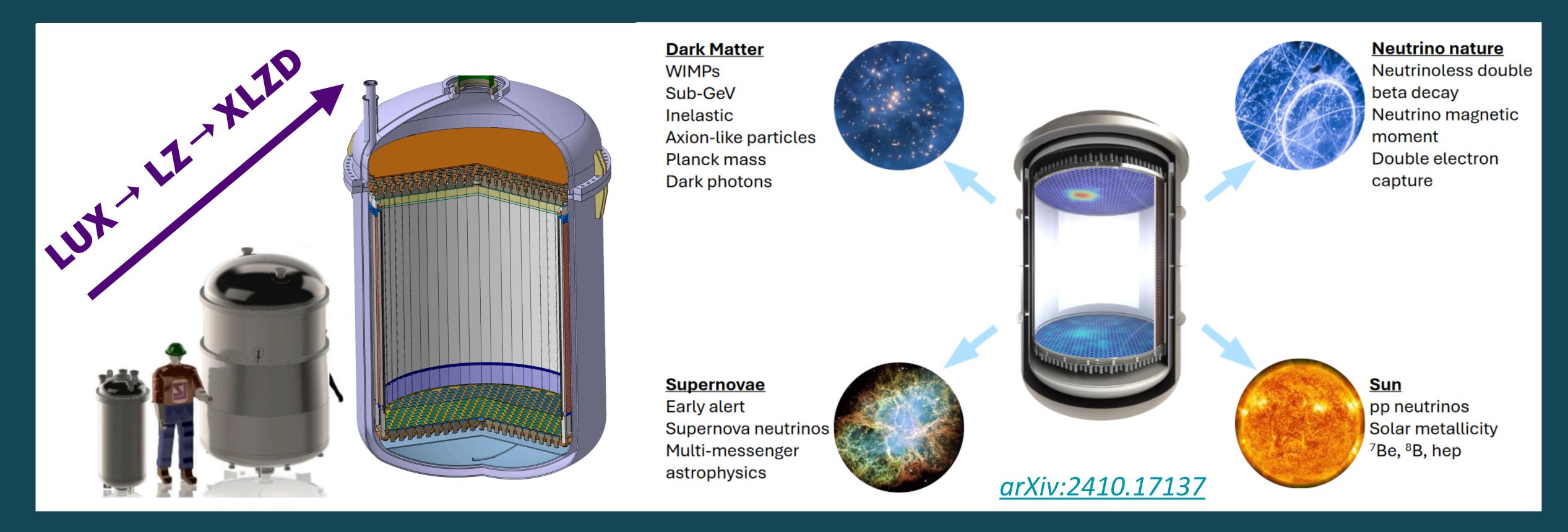
WHAT'S NEXT FOR LZ?





- LZ will continue flagship WIMP searches, but also show it is a multi-physics machine
- Observation of boron-8 solar neutrino CEVNS first for natural neutrinos
- 136Xe 0vββ decay search demonstrate competitiveness of this technology

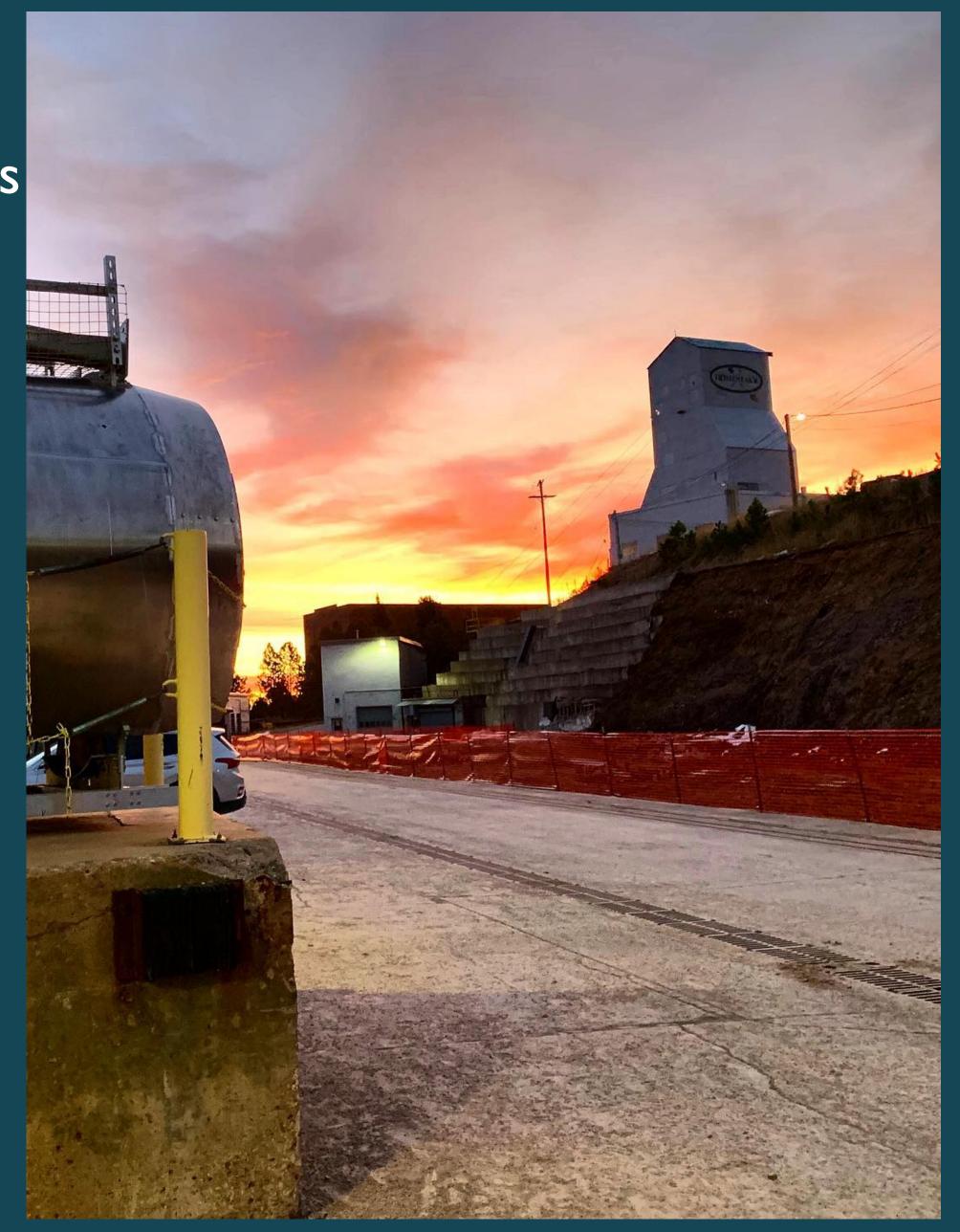
FURTHER INTO THE FUTURE - XLZD



- XLZD will use a 60-80 t xenon TPC (an order of magnitude bigger than current experiments)
- Merger of XENON, LZ, DARWIN \rightarrow ~500 people in 72 institutions, 17 countries
- A rare-event observatory, more than just the "definitive" WIMP search experiment

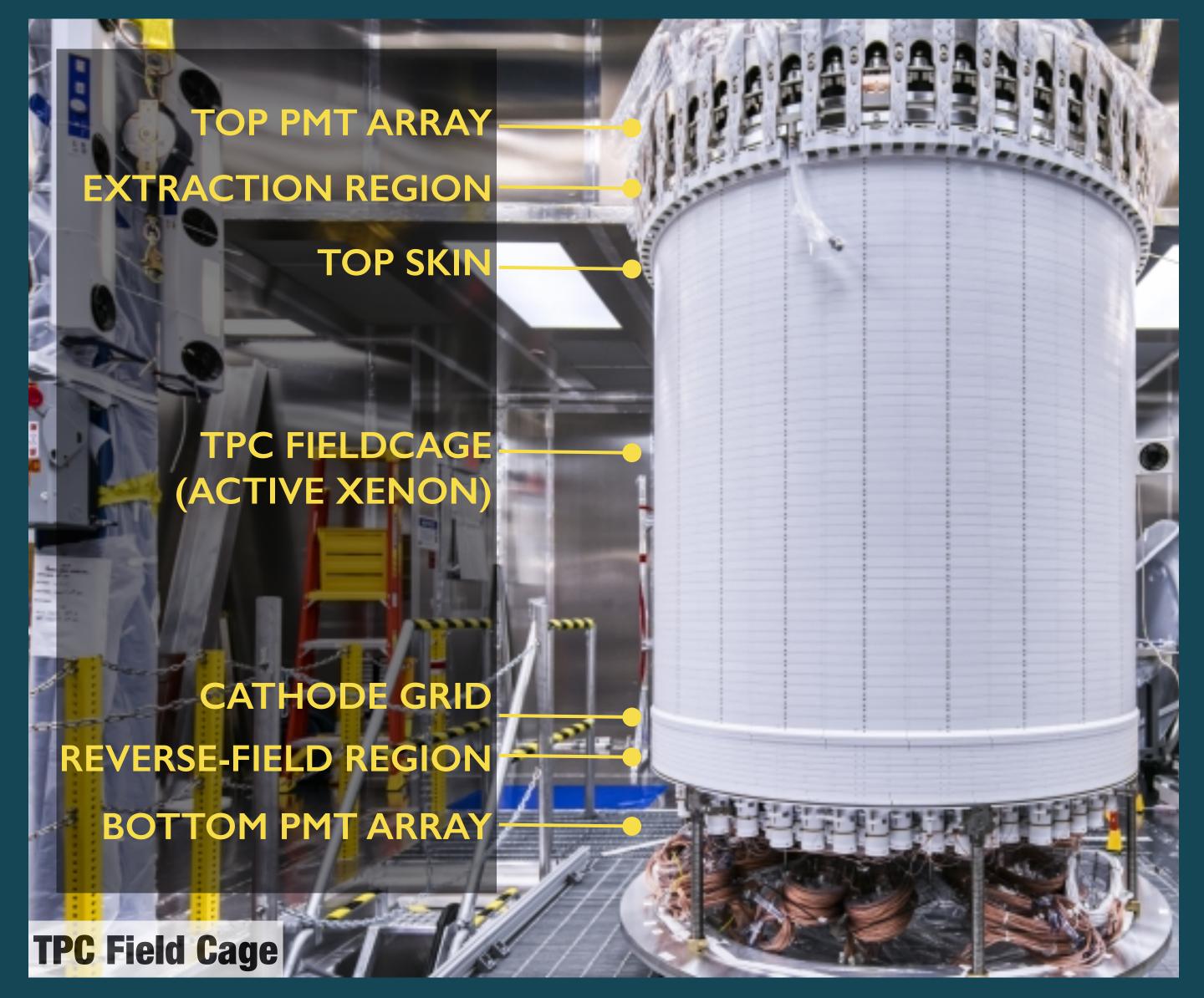
CONCLUSIONS

- New world-leading WIMP search limits exceeding previous best constraints by >4 times [PRL 135, 011802 (2025)]
- Radon tag developed and used for the first time:
 60% reduction in main ER background
 - First observation of charge-suppressed 124Xe DEC
- LZ will take data until 2028, towards 1000 live days
 - Multiple other physics channels to explore e.g. 8B CEvNS, neutrinoless double beta decay
 - LZ is discovery-ready for WIMPs
- Planning and R&D underway for next-generation, XLZD





THE LZ EXPERIMENT



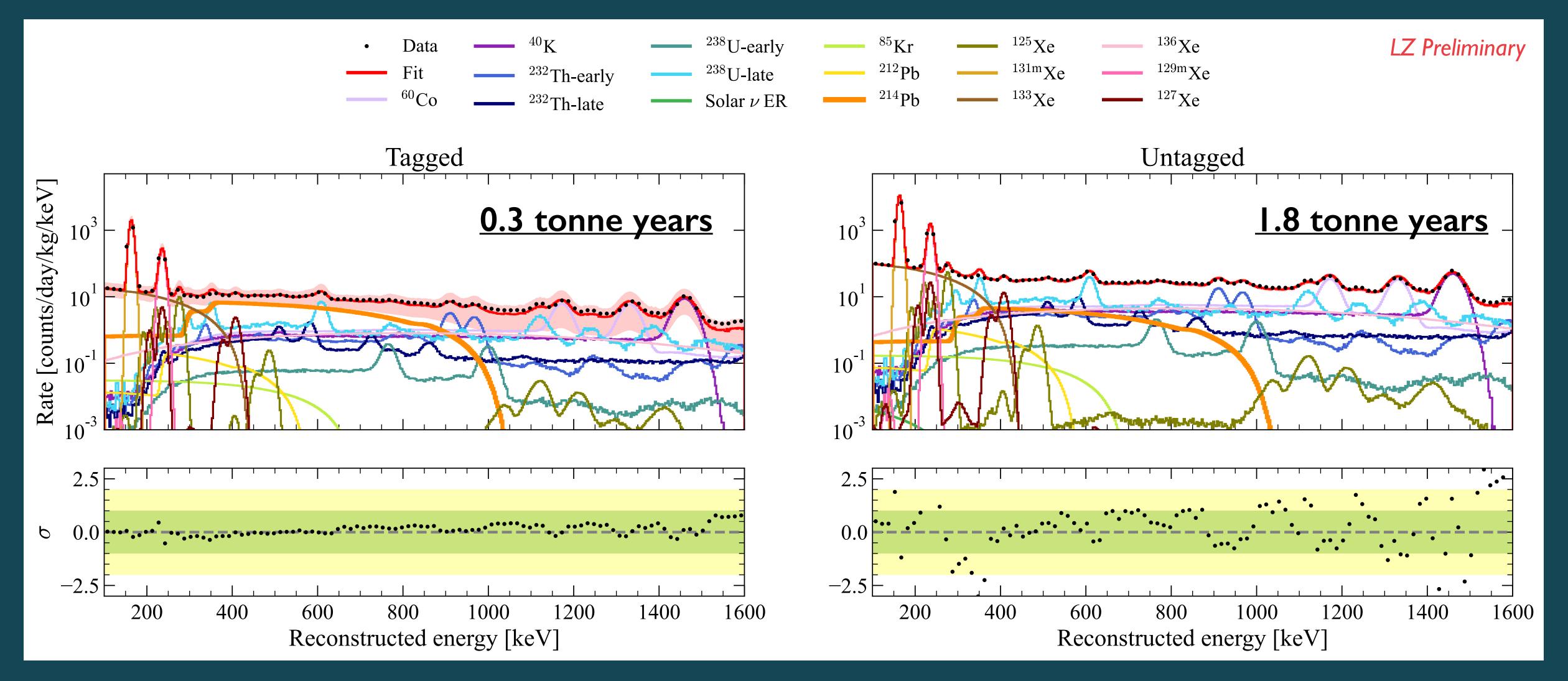








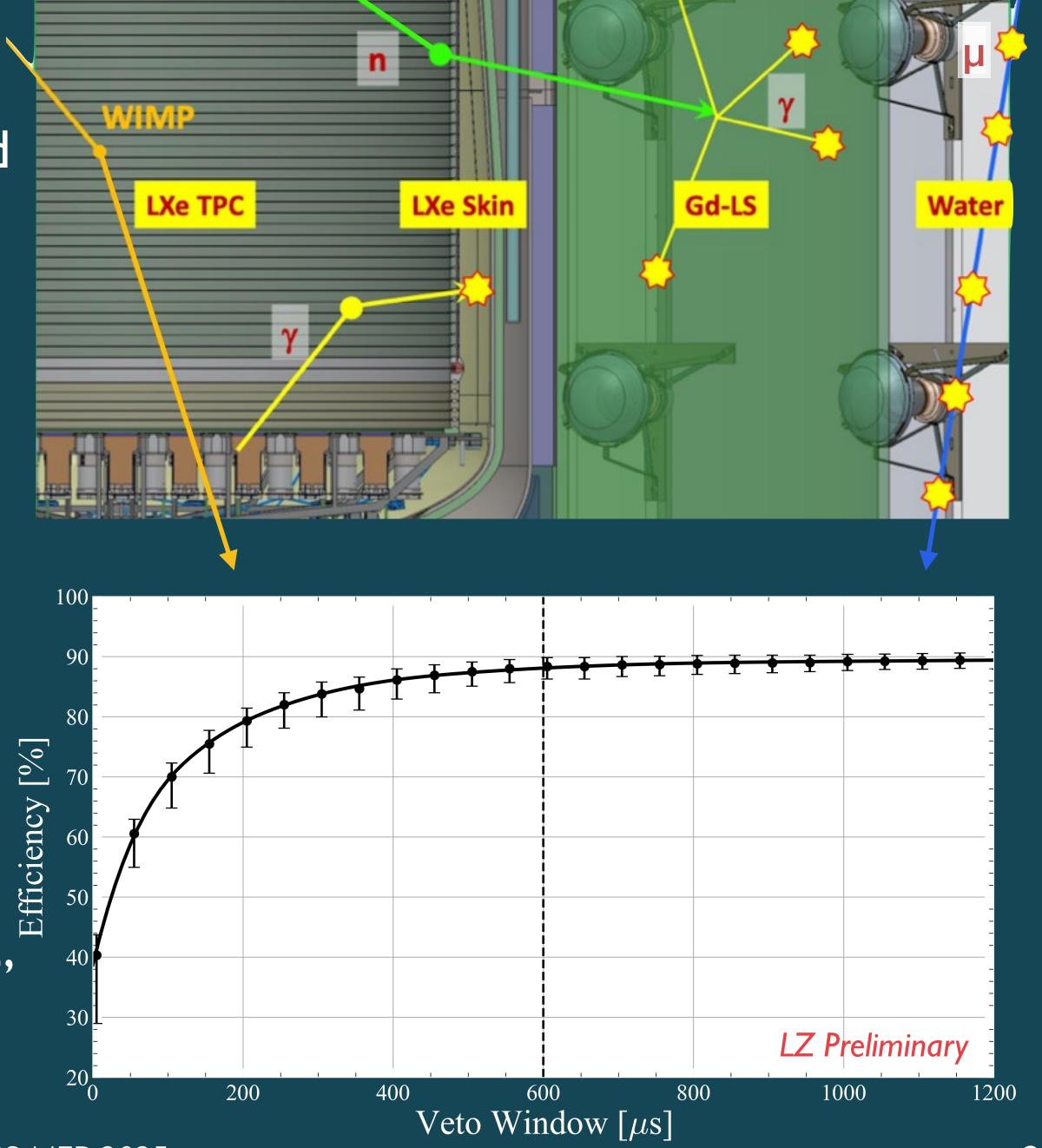
RADON TAGGED HIGH-ENERGY BACKGROUND FIT



Effective untagged 214 Pb activity of 1.8 \pm 0.3 μ Bq/kg (compared to 3.9 \pm 0.6 μ Bq/kg in total exposure)

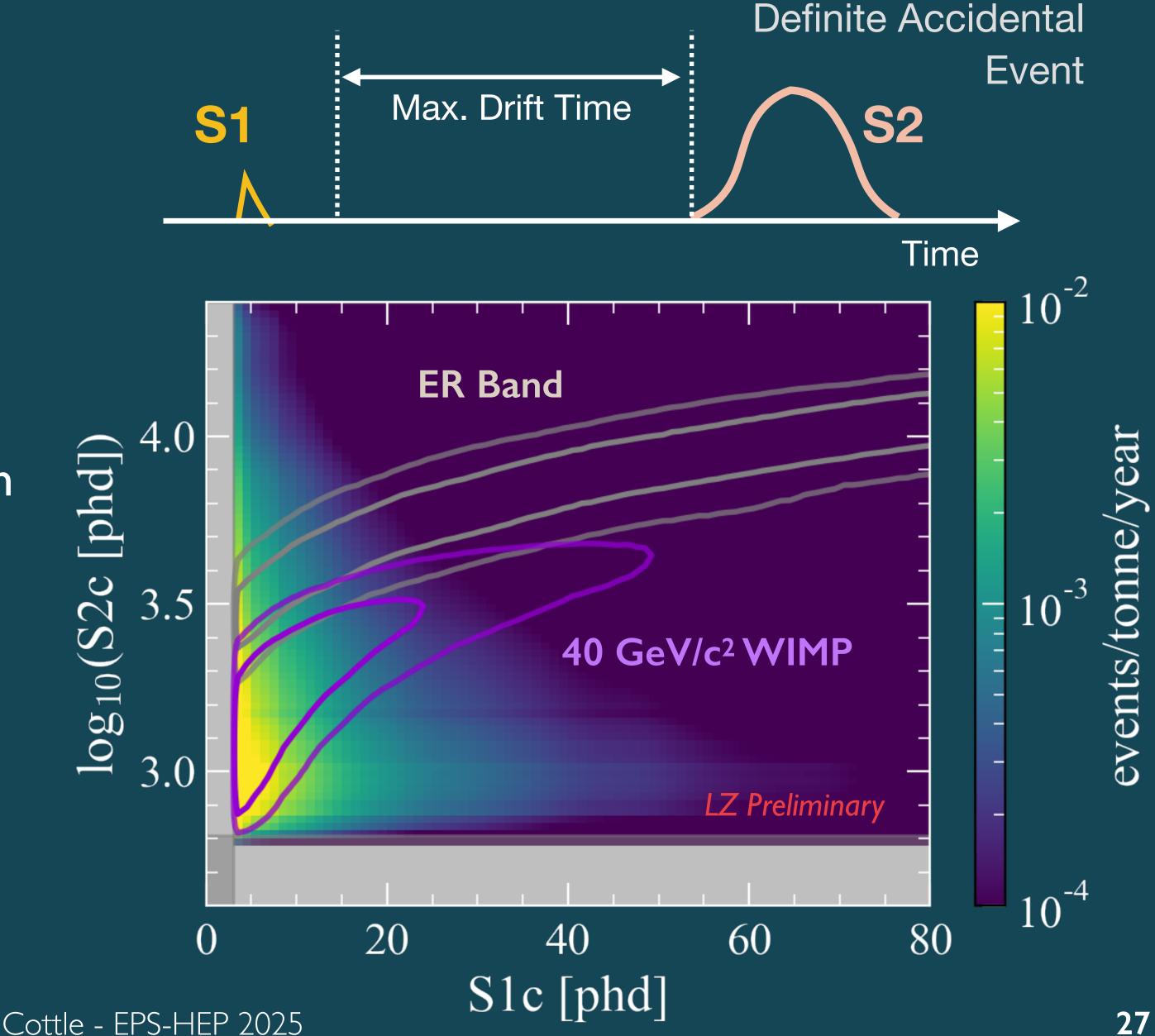
NEUTRONS & OD

- Neutrons induce NRs → dangerous background
- 17 tonnes Gd-loaded scintillator in OD
 - High thermal neutron capture cross-section
 - Release of ~8 MeV gammas from capture
- → delayed OD veto cut to reject neutrons
- AmLi neutron calibration-derived neutron veto efficiency = 89 ± 3 %
- Simulated neutron veto efficiency for radiogenic, background neutrons = 92 ± 4 %
 - → used for neutron constraint in final analysis

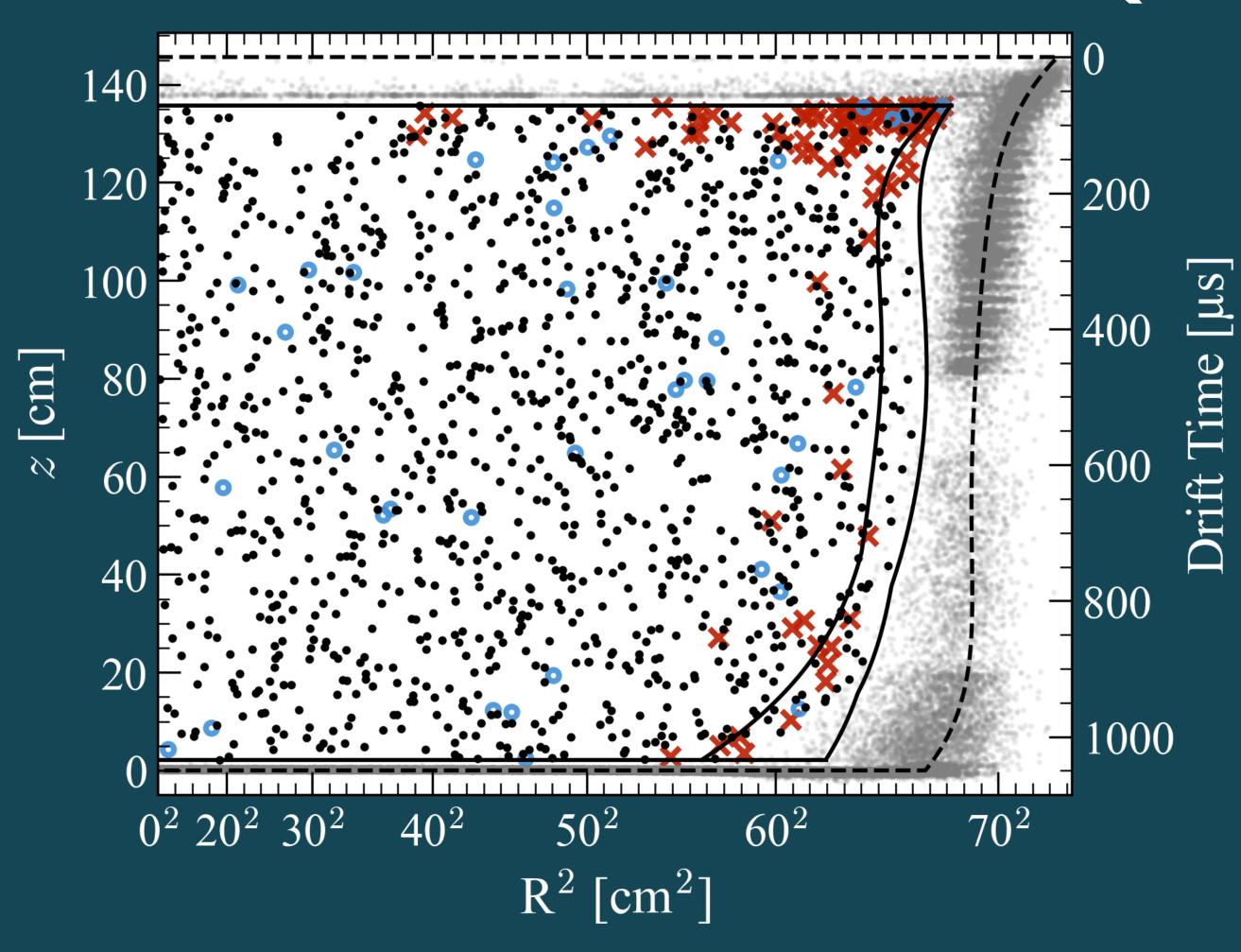


ACCIDENTAL COINCIDENCES

- Unrelated SIs & S2s can accidentally combine to produce single scatter events → could mimic a WIMP signal
- Rate: population of definite accidental events with unphysical drift time > 1 ms
- Distribution: fake events constructed from lone S1 & S2 pulse waveforms
- Analysis cuts developed to combat observed pulse/event pathologies
 - >99.5% rejection efficiency
 - WS2024 counts: 2.8 ± 0.6



FIDUCIAL VOLUME (FV)



Events prompt-tagged by vetoes

Events delayed-tagged by the vetoes

- FV defined to avoid higher background rates at TPC edges (self-shielding)
- TPC radial edge curved due to electric field
- FV definition:
 - $71 \mu s$ < drift time < $1030 \mu s$
 - Azimuthally & drift time-dependent
 radial cut chosen to ensure
 <0.01 wall background counts in the FV
- Calculated fiducial mass of 5.5 ± 0.2 t

COMBINED LIKELIHOOD

Exposures in Each Sample in Tonne Years

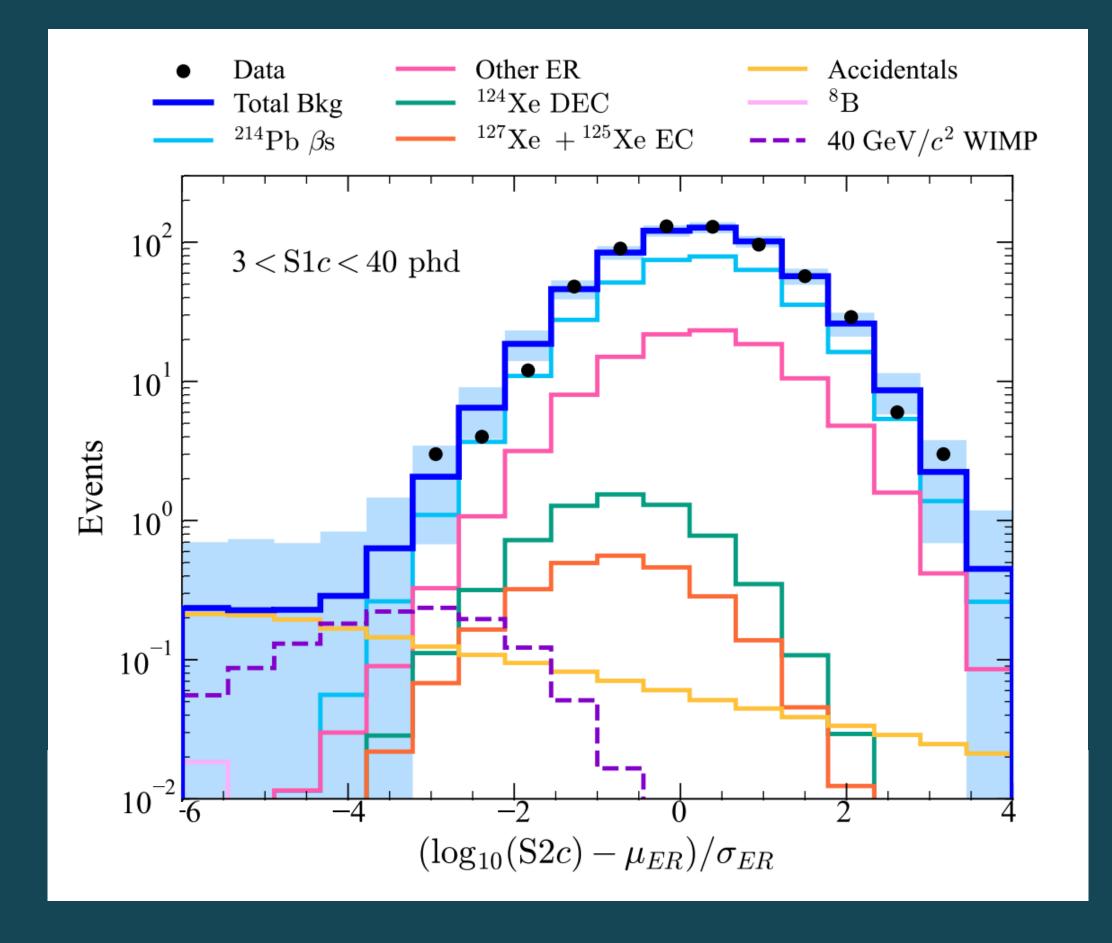
| | 2 | 3 | 4 | 5 | 6 |
|----------------------|-----------------------|--------------|-------------------|-------------------|--------|
| High-Mixing State | Radon Tag Inactive | Radon Tagged | Radon Untagged | Skin/OD Vetoed | WS2022 |
| 0.6 | 0.6 | 0.3 | 1.8 | n/a | 0.9 |

Six samples combined in likelihood for final statistical analysis

- WS2024 represented by samples 1-4
- Skin/OD vetoed sample (5) full 3.3 tonne years of WS2024, but failing veto coincidence cuts → provide a direct constraint on the neutron background rate
- WS2022 sample (6) unmodified since first result → push sensitivity further

WS2024 FIT RESULTS

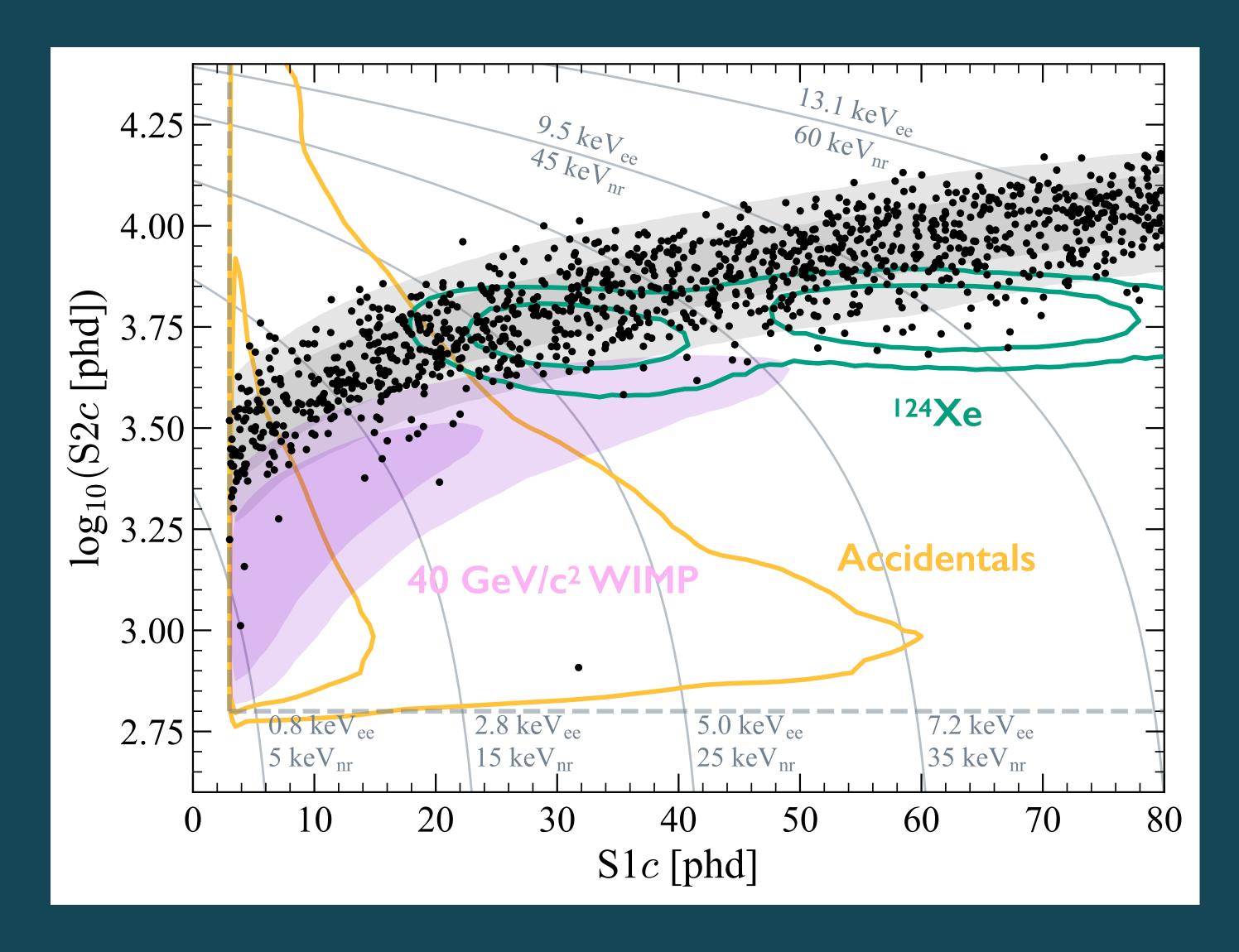
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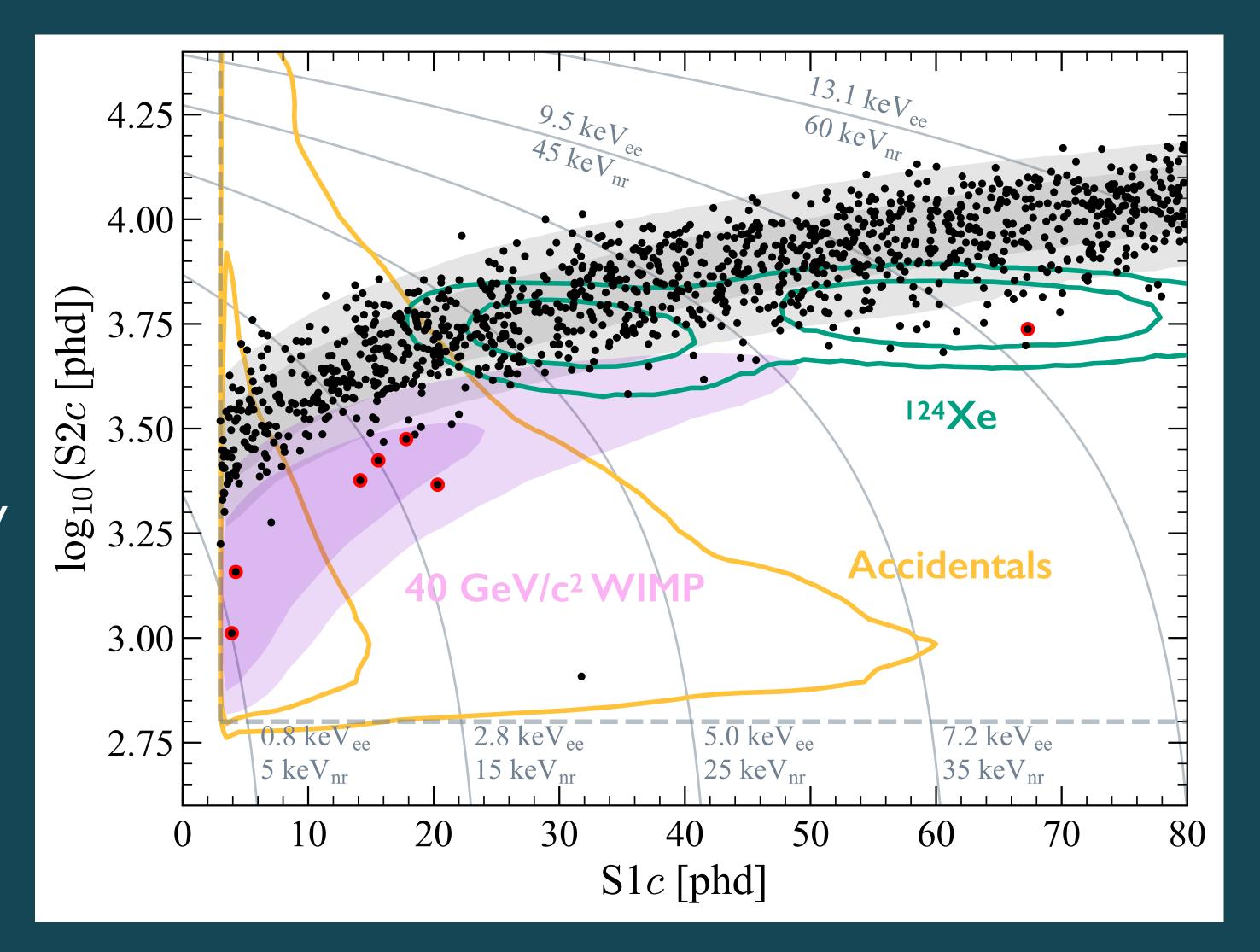
WS2024 DATA - SALTED

- Final exposure of
 220 live days * 5.5 tonnes
 = 3.3 tonne years
- 1227 events remaining

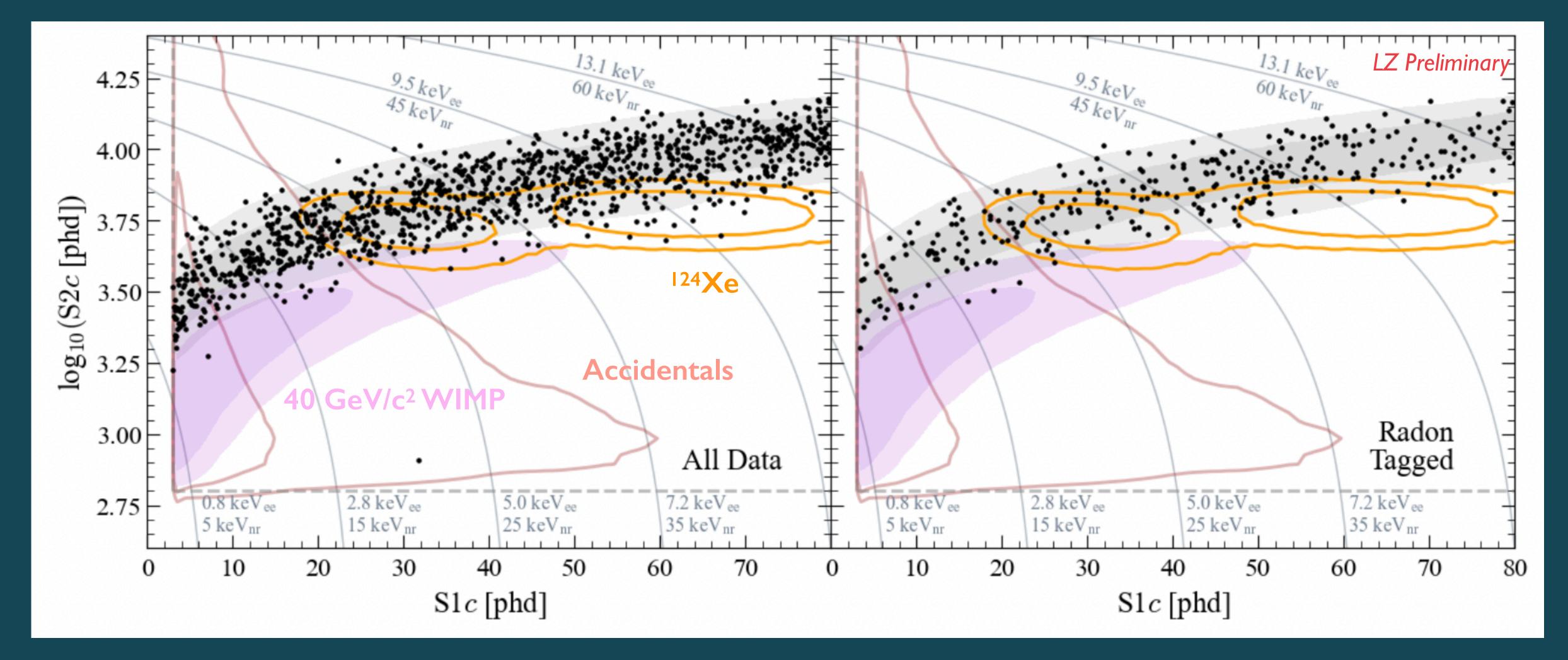


WS2024 DATA - SALTED

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- 7 salt events pass all analysis cuts out of 8 total injected in WS2024
 - → inline with evaluated signal efficiency



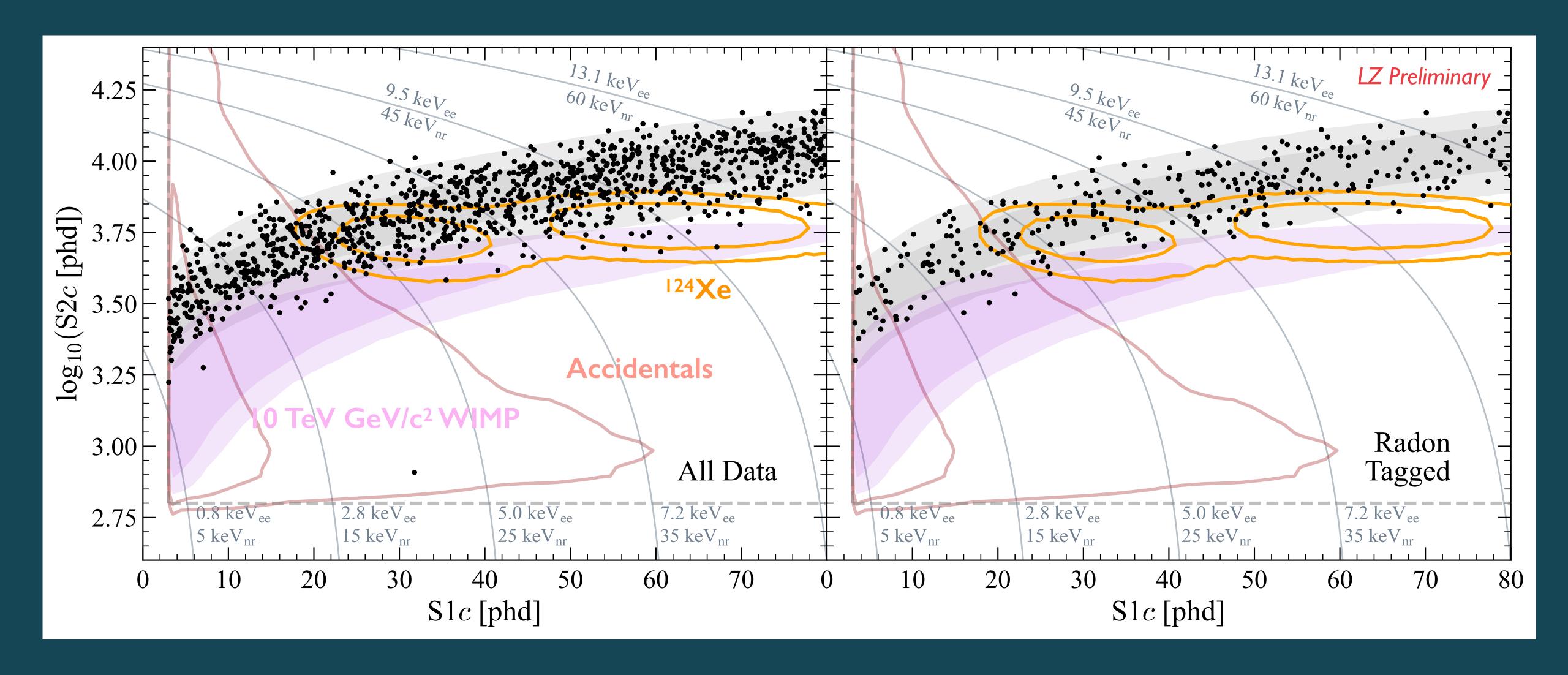
WS2024 DATA - UNSALTED; RADON TAGGED



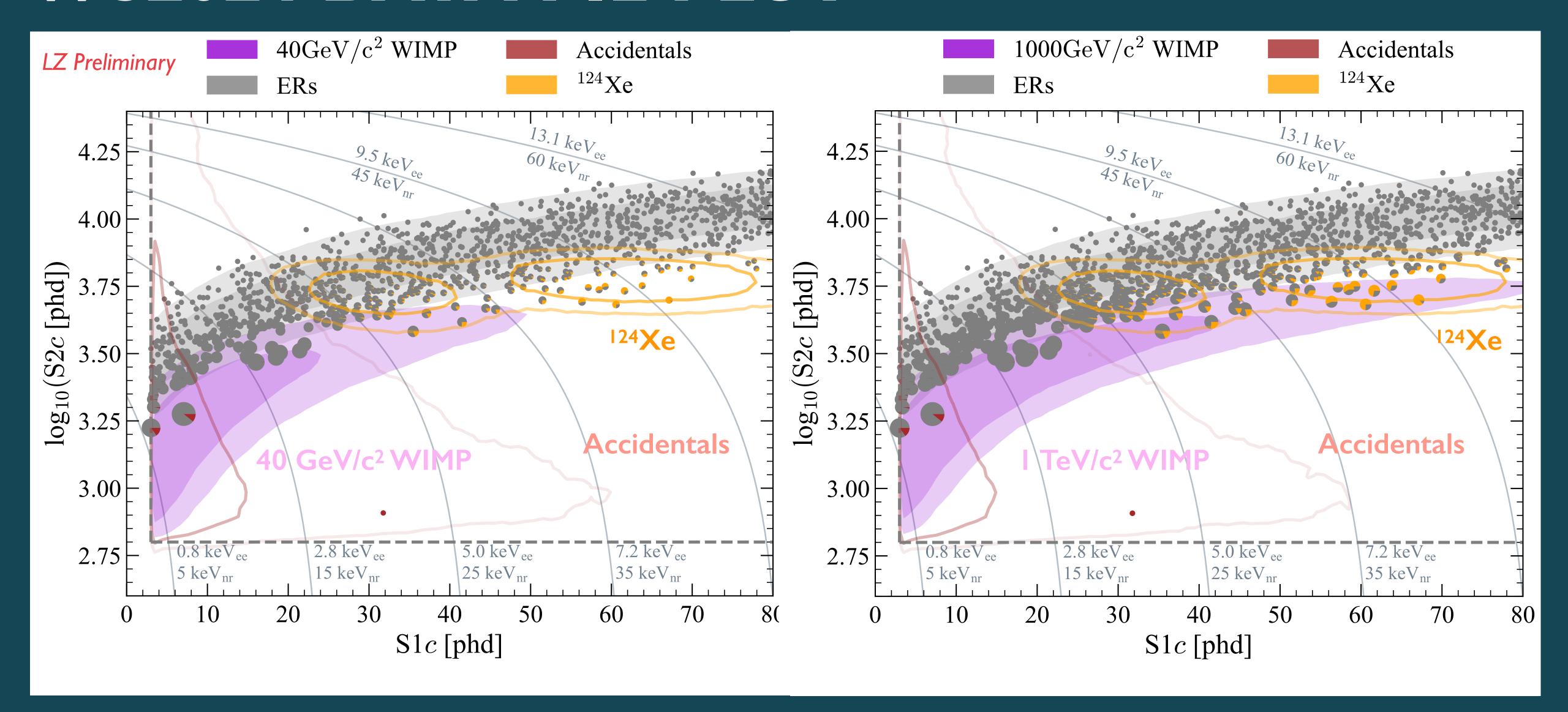
3.3 tonne years

0.3 tonne years

10 TEV/C² WIMP



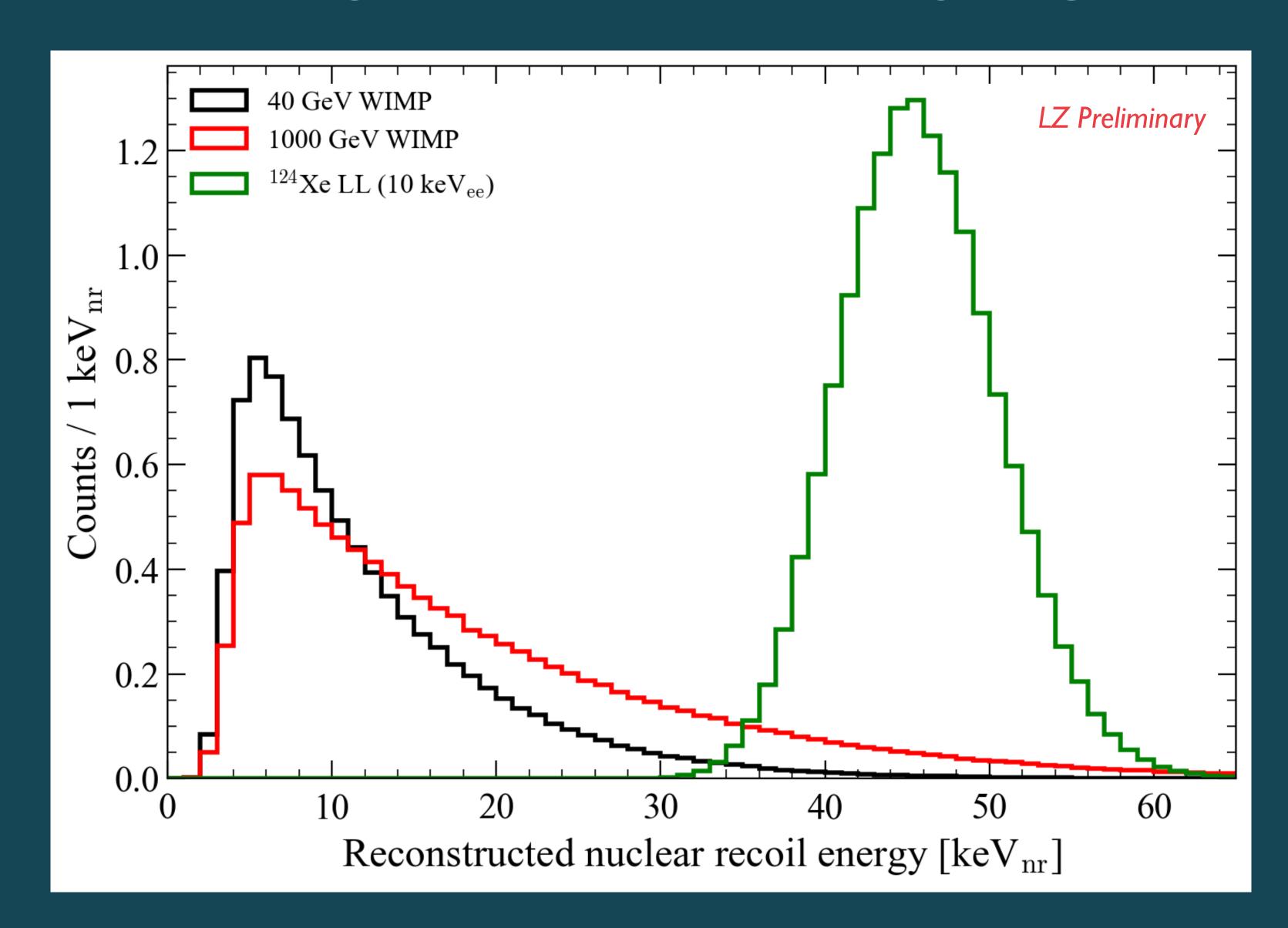
WS2024 DATA PIE PLOT



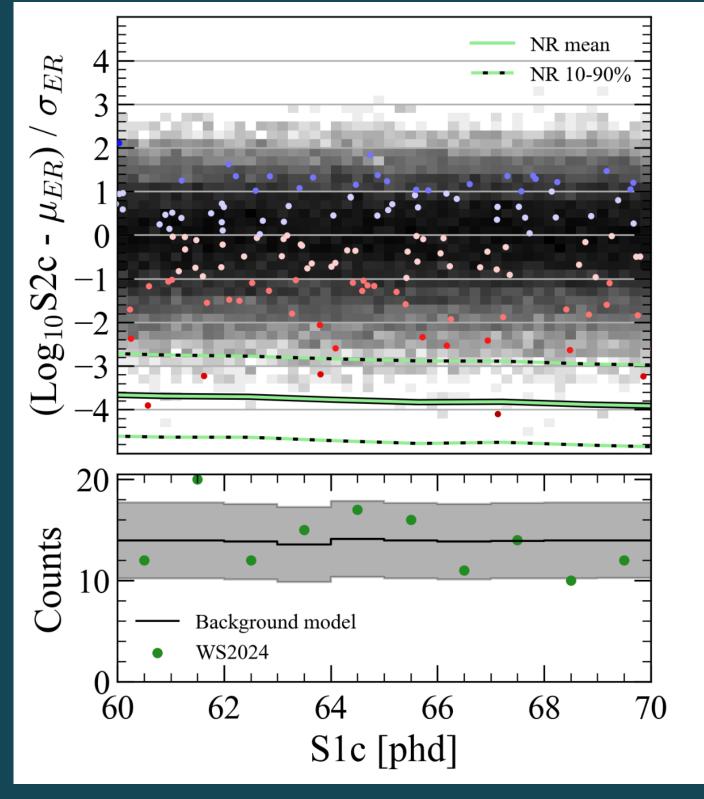
124XE LL-SHELL COMPARED TO DARK MATTER SPECTRA

WIMP spectra normalised to LZ's 4.2 tonne year median 3σ discovery potential:

- 9 events @ 40 GeV
- II events @ 1000 GeV



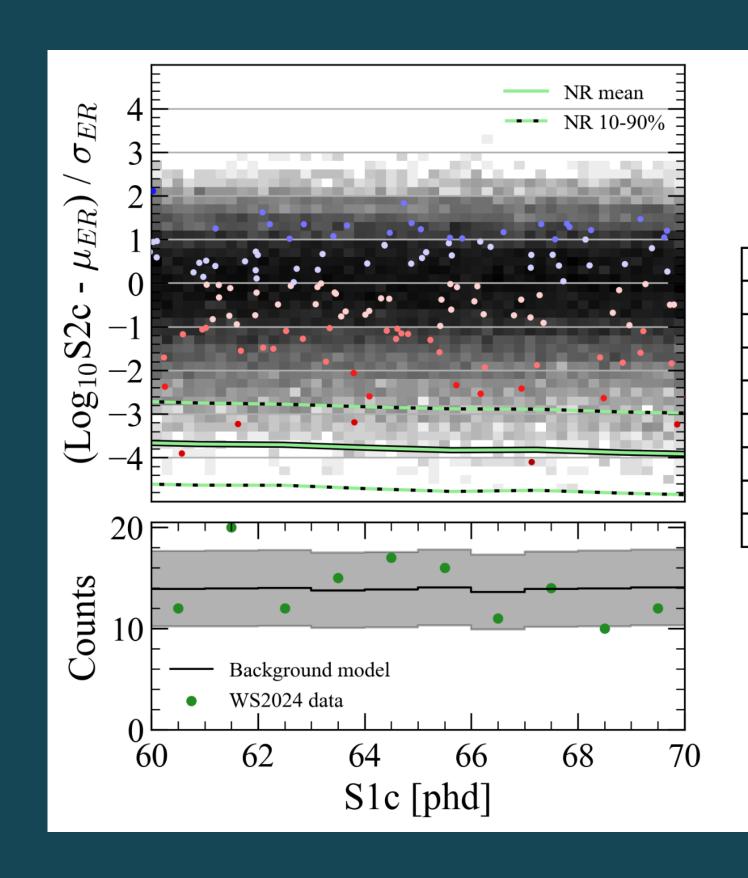
GOODNESS OF FITS IN KEY 124XE REGION



60 < S1c < 70 138 counts

| σ | Expt. | Obs. | p_{LR} | p_{MC} | p_{Pois} |
|----------|-------|------|----------|----------|------------|
| [-5, -4] | 0.11 | 1 | 0.1 | 0.092 | 5.2e-03 |
| [-4, -3] | 0.86 | 4 | 0.014 | 8.6e-03 | 2.0e-03 |
| [-3, -2] | 6.1 | 7 | 0.74 | 0.84 | 0.28 |
| [-2, -1] | 23 | 25 | 0.77 | 0.85 | 0.34 |
| [-1, 0] | 51 | 41 | 0.14 | 0.25 | 0.087 |
| [0, 1] | 51 | 37 | 0.030 | 0.063 | 0.02 |
| [1, 2] | 19 | 22 | 0.62 | 0.66 | 0.26 |
| [2, 3] | 2.4 | 1 | 0.31 | 0.45 | 0.31 |

LZ Preliminary



60 < S1c < 70 138 counts

| σ | Expt. | Obs. | p_{LR} | p_{MC} | p_{Pois} |
|----------|-------|------|----------|----------|------------|
| [-5, -4] | 0.44 | 1 | 0.47 | 1.0 | 0.073 |
| [-4, -3] | 2.6 | 4 | 0.42 | 0.54 | 0.12 |
| [-3, -2] | 7.4 | 7 | 0.89 | 1.0 | 0.54 |
| [-2, -1] | 21 | 25 | 0.52 | 0.52 | 0.22 |
| [-1, 0] | 50 | 41 | 0.17 | 0.33 | 0.10 |
| [0, 1] | 52 | 37 | 0.026 | 0.063 | 0.017 |
| [1, 2] | 19 | 22 | 0.63 | 0.68 | 0.26 |
| [2, 3] | 2.4 | 1 | 0.29 | 0.43 | 0.3 |

LZ Preliminary

$$Q_{LL}/Q_{\beta} = 0.87$$

(i.e. L-shell suppression)

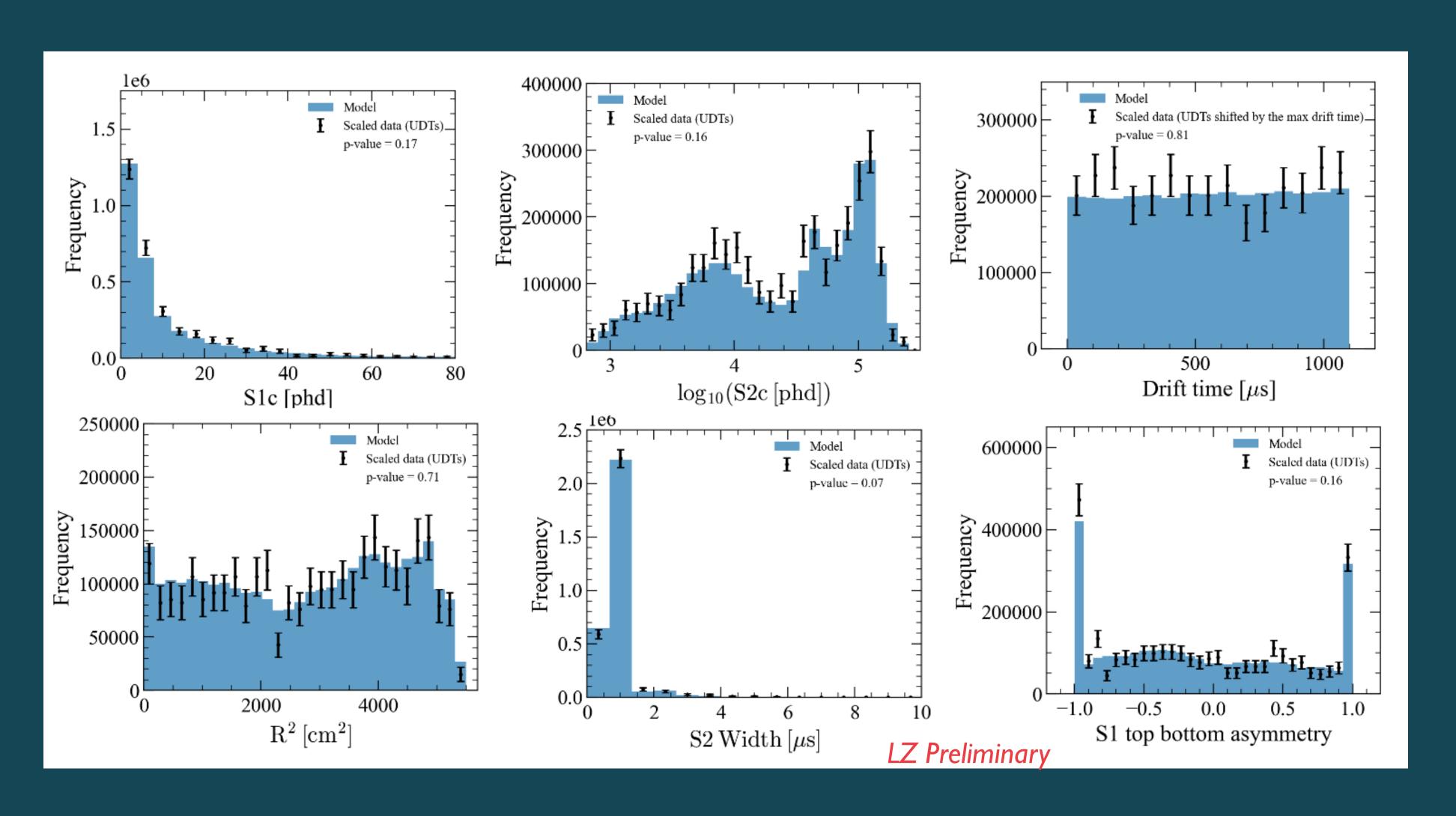
$$Q_{LL}/Q_{\beta} = 0.65$$

(i.e. double L-shell ionisation density)

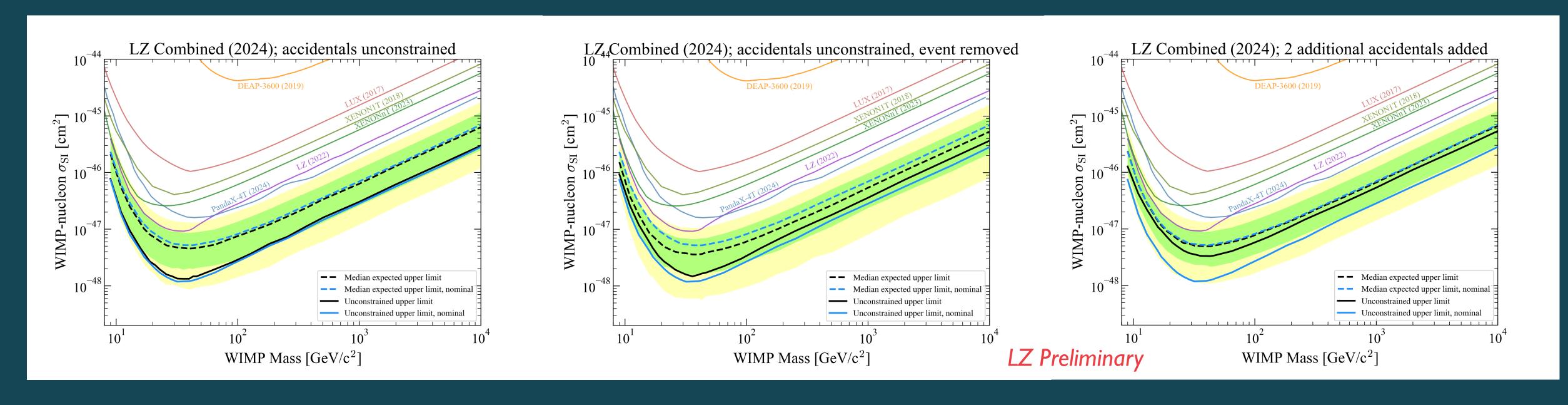
ACCIDENTALS MODEL & SIDEBAND COMPARISONS

Comparing manufactured accidental events and unphysical drift accidentals

Good agreement before application of SI- and S2-based cuts

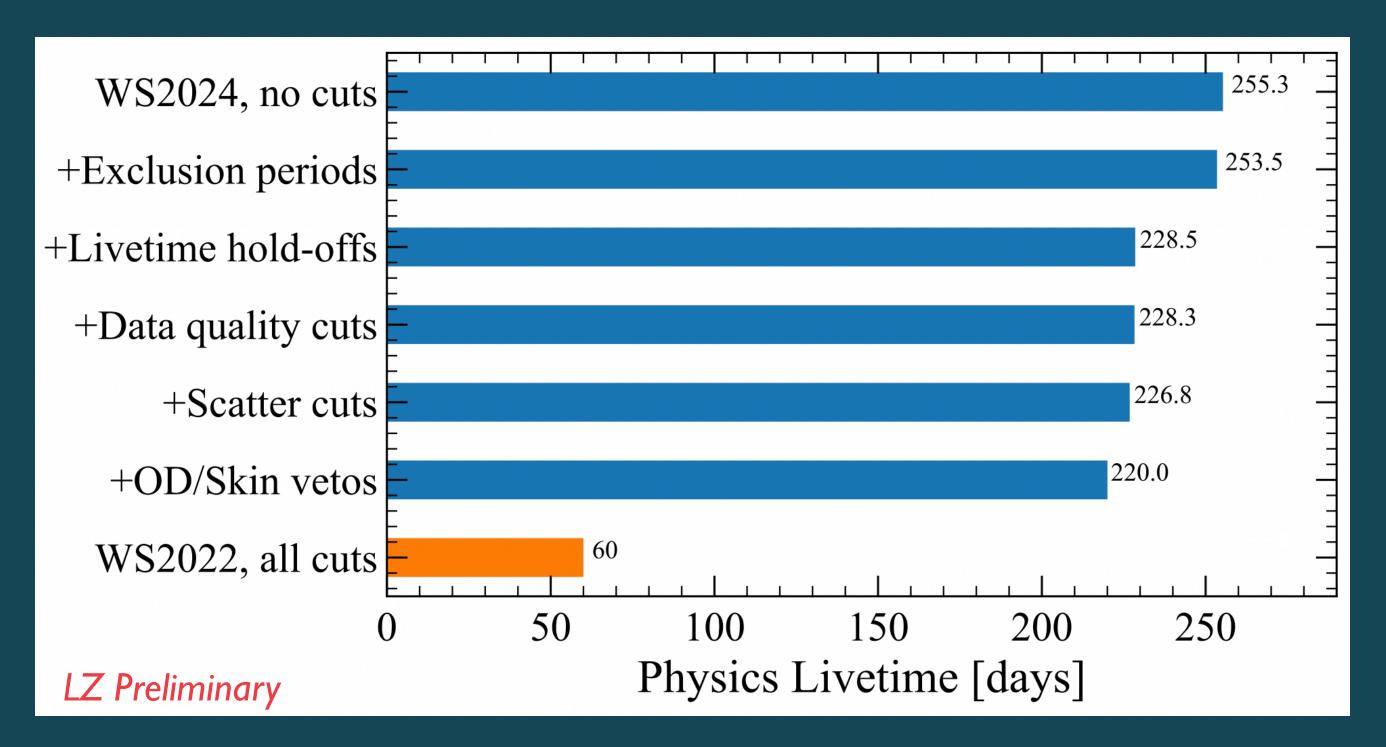


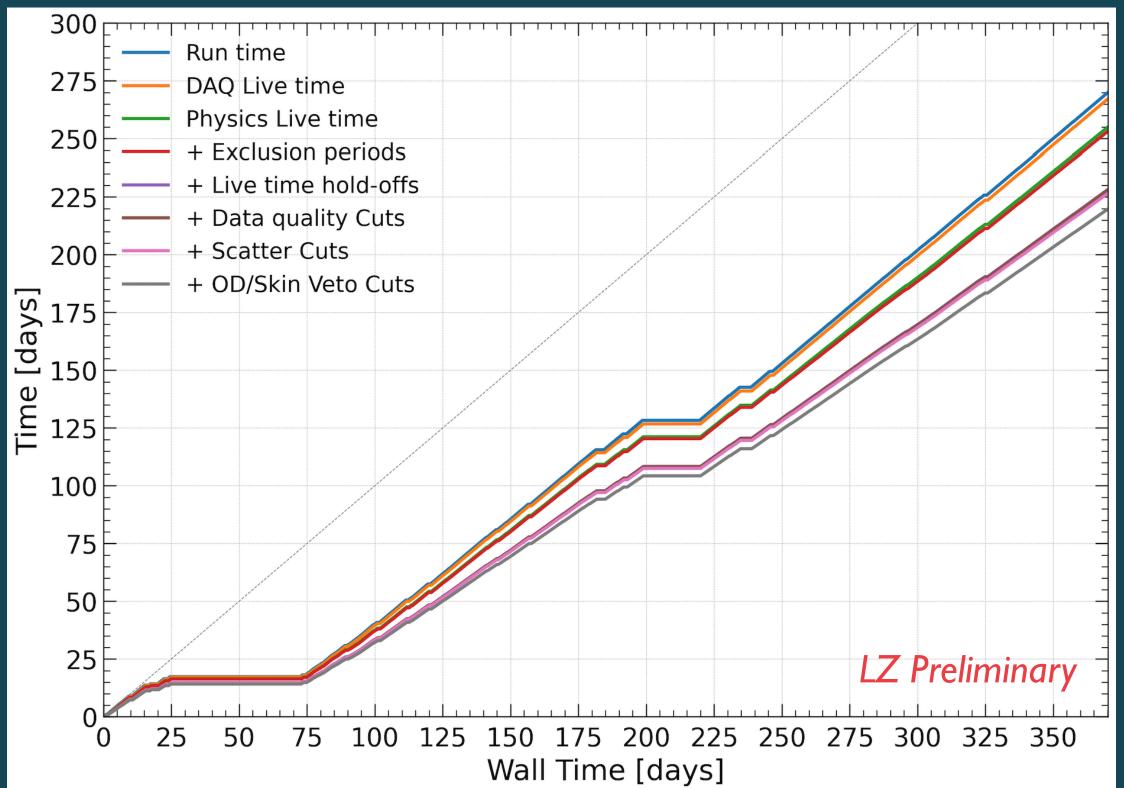
CHECKS OF ACCIDENTALS IMPACT ON LIMIT



- 1. Remove accidental rate constraint: best fit drops $2.6 \rightarrow 1.4$
- 2. Remove constraint & outlier event: best fit drops $1.4 \rightarrow 0$
 - Outlier event holds model up, over subtracting in the WIMP region
- 3. Adding fake events props limit back up
- → under-fluctuation of accidental events in the WIMP region

LIVE TIME

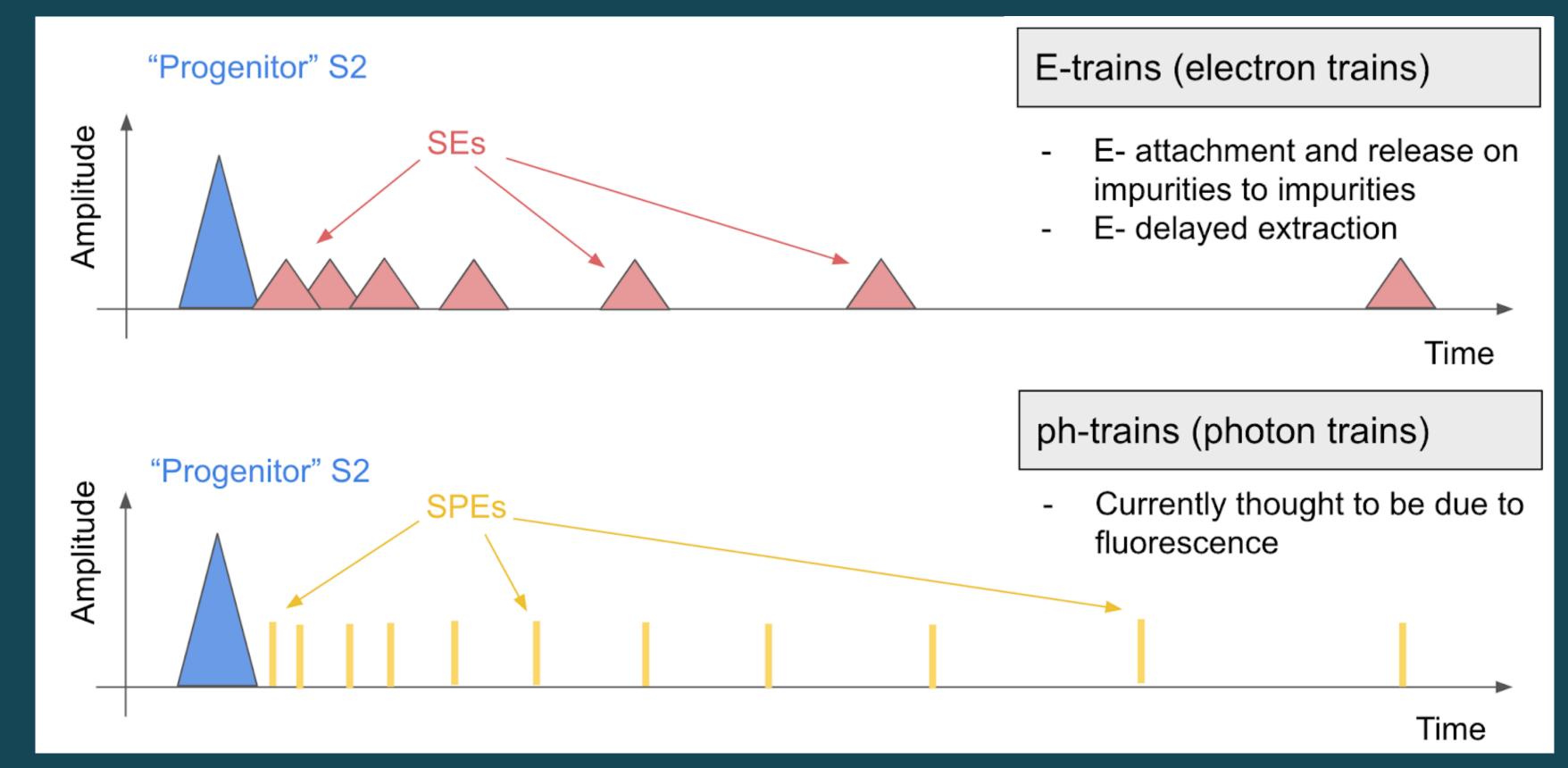




- Rejection of live time with detector instabilities, high TPC pulse rates
- 86% live time remaining after all analysis live time exclusions
 - → mainly driven by improved live time retention of e-train veto

E-TRAIN VETO

- Large S2s induce pulse "trains" lasting 100s of ms, much longer than the event window
- High pulse rates can lead to piled-up photon or electron pulses that mimic S1s and S2s, thus
 contributing to accidental coincidence backgrounds
- Removal of periods
 after S2s (e-/ph trains)
 excludes ~10% of our
 live time in WS2024
 (compared to ~30%
 for WS2022)
- Improvement due to optimisations & smaller S2s
 (= shorter exclusions)

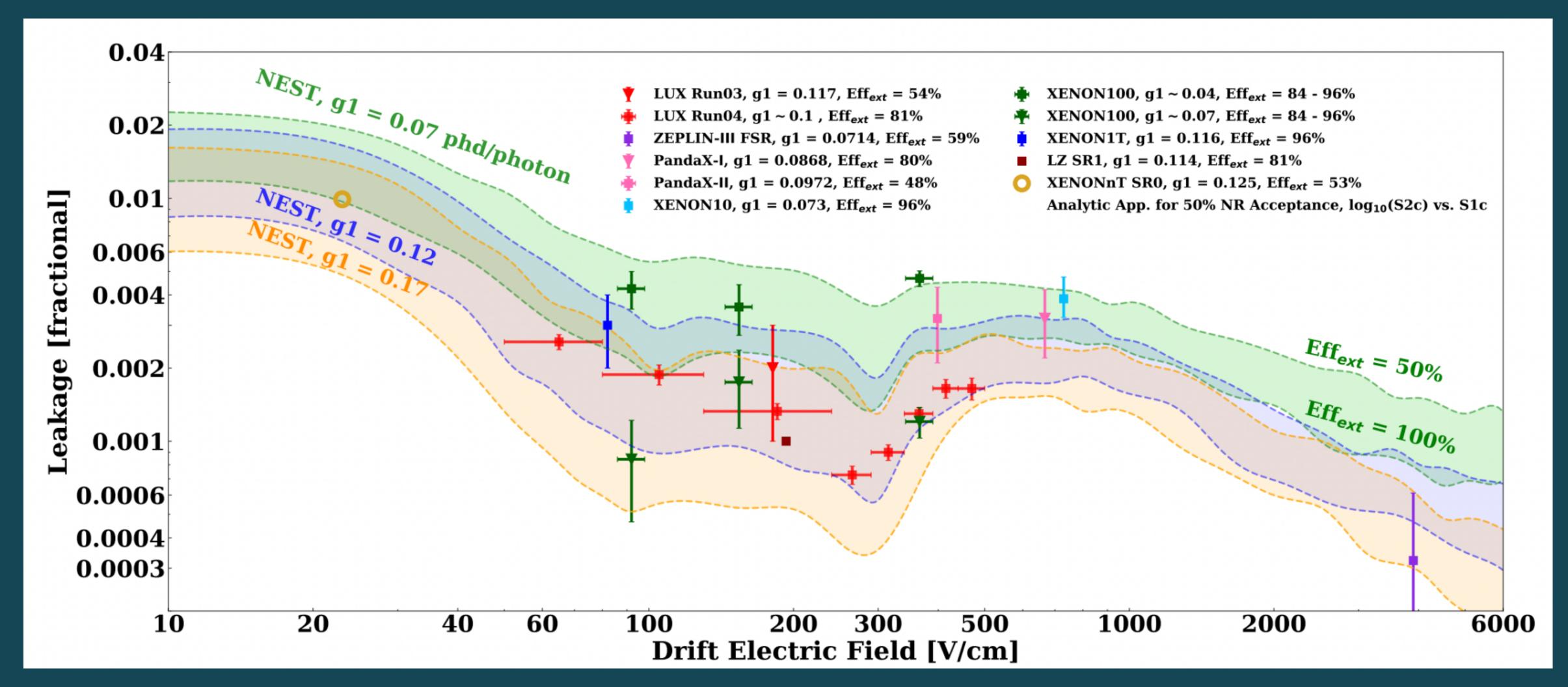


WS2024 VS WS2022 CONDITIONS

| | Analysis Live Time (Days) | Drift Field [V/cm] | Extraction Field (in liquid) [kV/cm] | Single Electron Size [phd] |
|---------|---------------------------|--------------------|--------------------------------------|-------------------------------|
| VVS2024 | 220 | 97 | 3.4 | 44.5 |
| WS2022 | 60 | 193 | 4.4 | 58.5 |

- Optimisations performed following WS2022: trigger configuration; electrode voltages; circulation
- Lowered gate-anode ΔV by 0.5 kV to reduce spurious electron emissions
- Optimised drift field to 97 V/cm to maintain similar ER/NR discrimination whilst enabling long-term, stable running of the detector

NEST MODEL OF ER LEAKAGE VS DRIFT FIELD



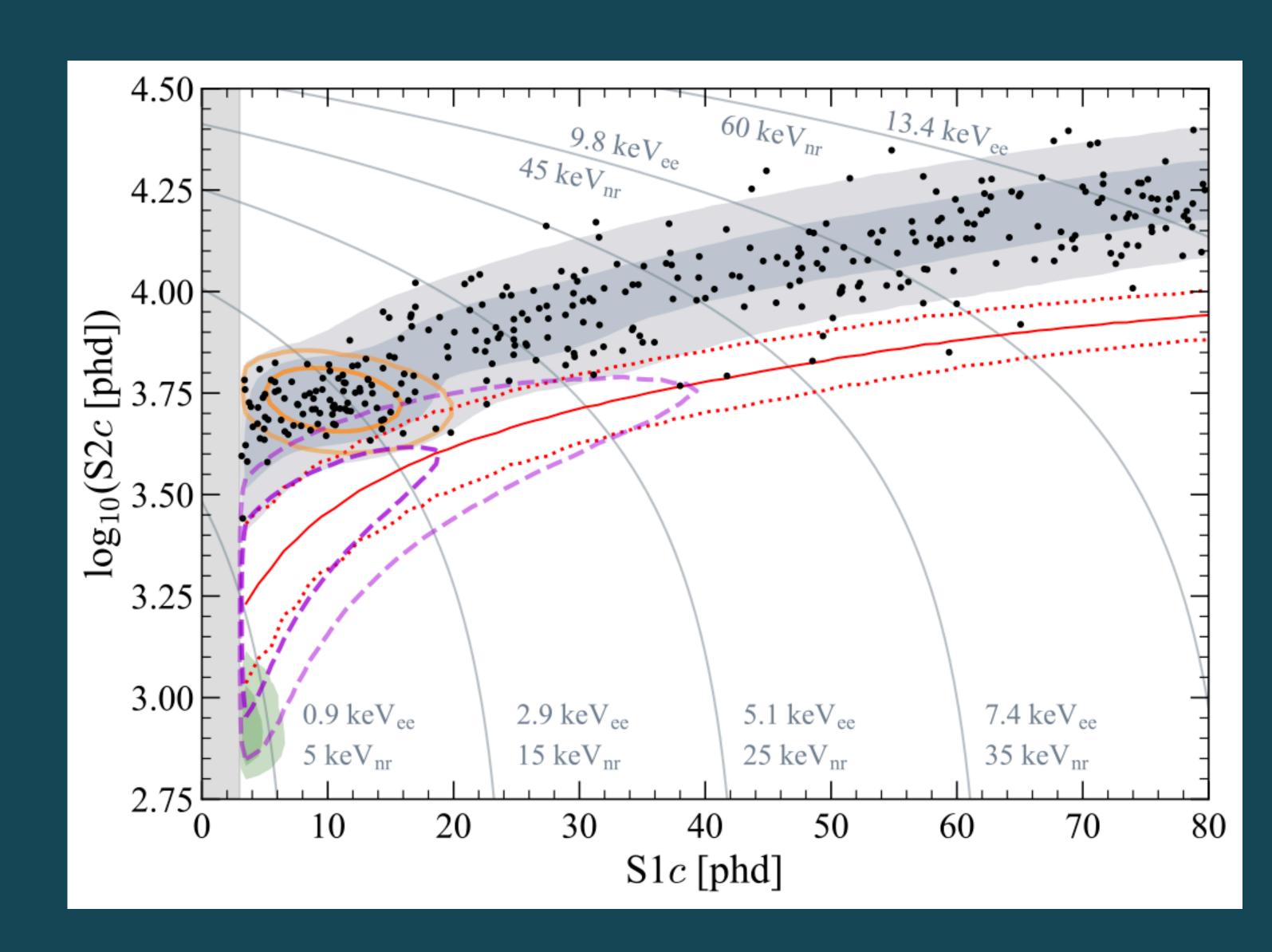
arXiv:2211.10726

WS2022 DATA

- 335 events after all cuts
- PDFs created with energy deposit + detector response simulations*
- Profile likelihood ratio (PLR) analysis

Key

- 1 & 2-Sigma Contours
- Post-fit total background distribution
- 37Ar
- 8B
- 30 GeV/c² WIMP
- NR band from DD

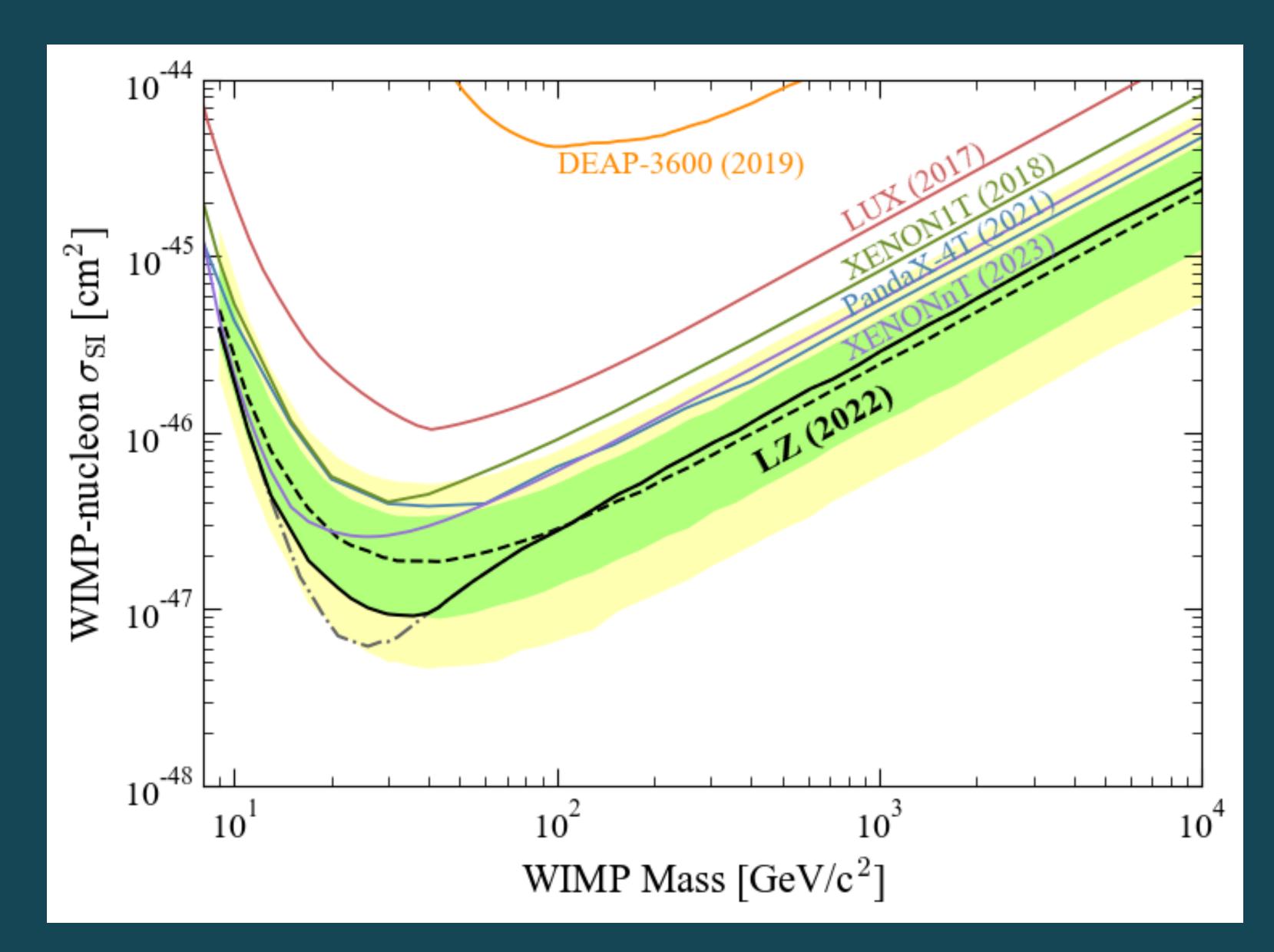


WS2022 LIMIT

- Two-sided PLR search with power-constrained limit defined using rejection power
- Minimum cross-section of $\sigma_{SI} = 9.2 \times 10^{-48} \text{ cm}^2 \text{ for WIMP}$ mass of 36 GeV/c²
- No evidence for WIMPs

<u>Key</u>

- Observed limit
- - Median expected sensitivity



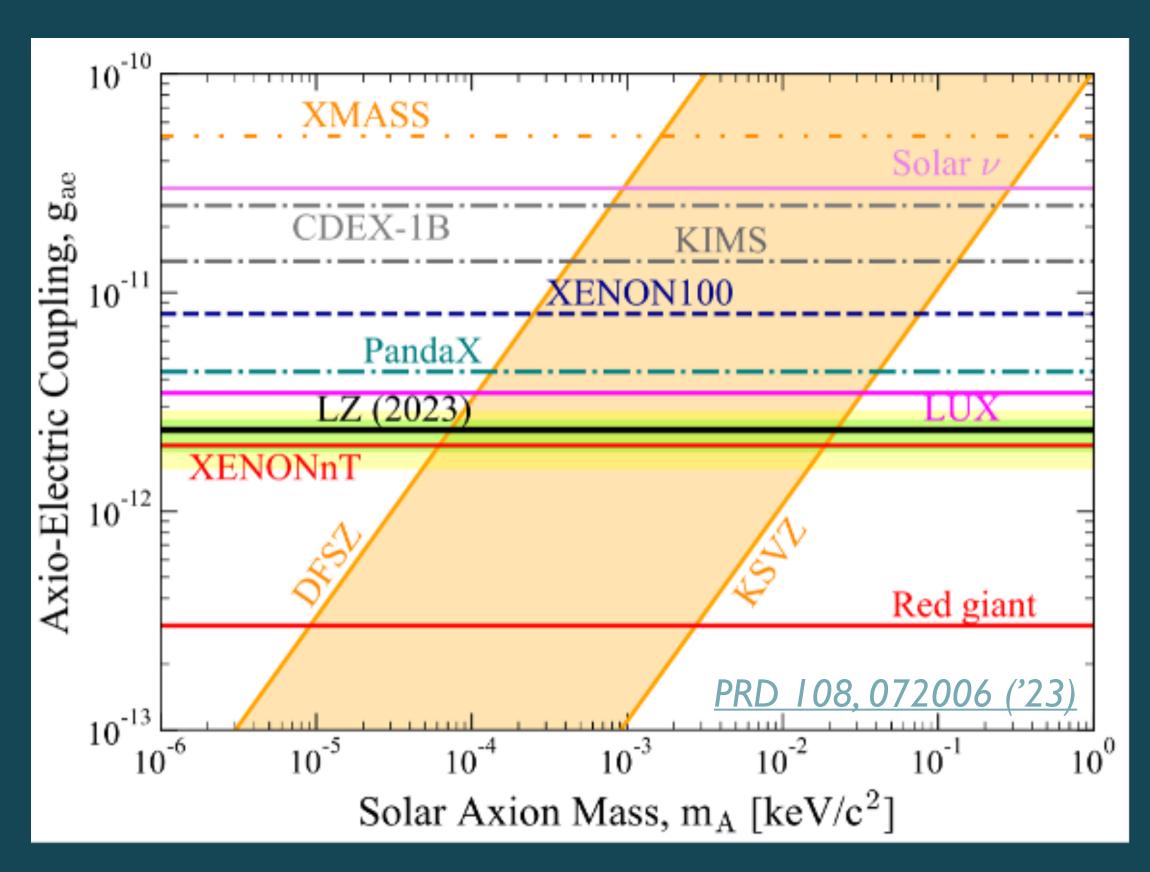
PHYSICS VIA ELECTRON RECOIL CHANNEL

Neutrino Magnetic Moment & Effective Millicharge

TEXONO Neutrino Effective MilliCharge δ_Q [e_0] XMASS PANDAX-II Neutrino Magnetic Moment $\mu_ u$ $[\mu_{ m B}]$ XENONIT (S2o) GEMMA Borexino TEXONO LZ **GEMMA** (2023)XENON1T (S1S2) XENONnT LZGlobular Cluster

PRD 108, 072006 ('23)

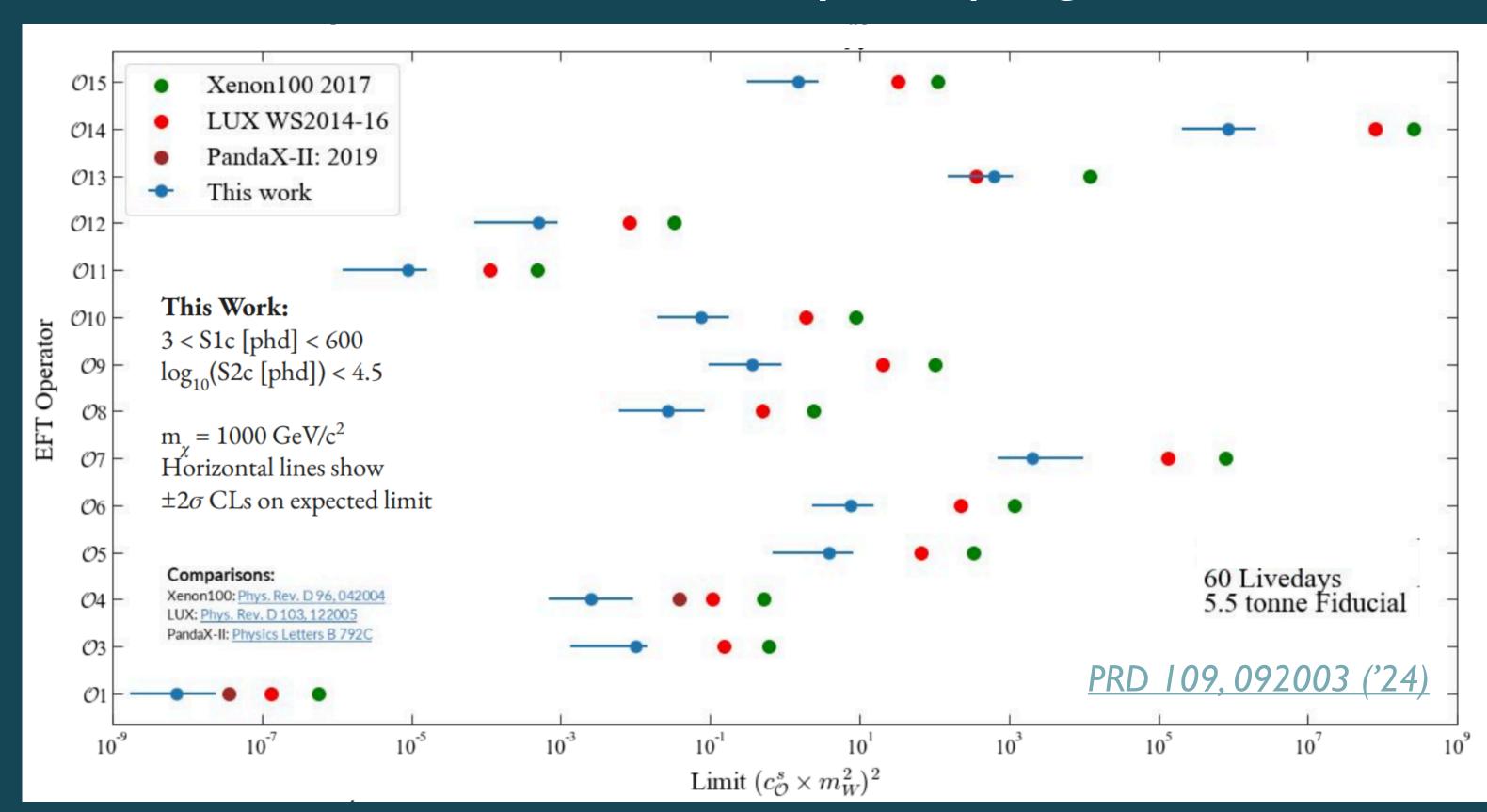
Solar Axion Interactions



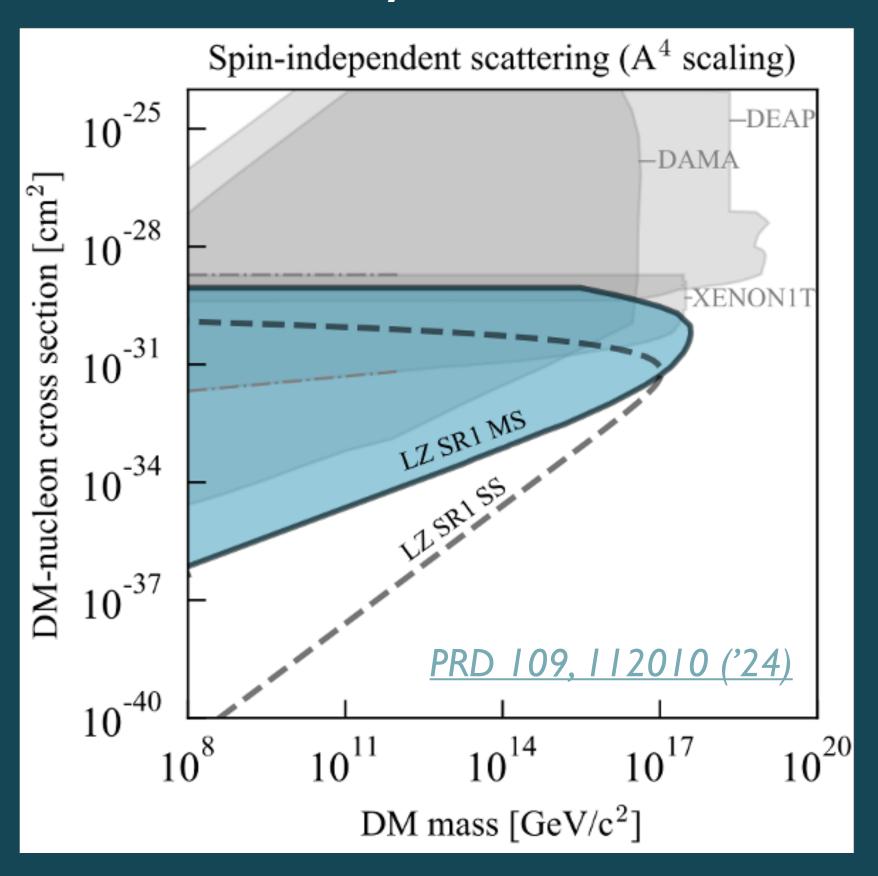
Neutrino interactions & properties; axions and ALPs; rare decays of xenon isotopes

MORE DARK MATTER RESULTS

Effective Field Theory Couplings



Ultra-heavy Dark Matter



New world-leading constraints on several EFT operators; high-energy MIMP interactions

→ broadband detectors, capable of searches across a wide range of candidate masses