Composite Operators in Quantum (super)gravity

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• QFT setting – no strings or other non-QFT structures

$$Z = \int_{\Omega} Dg_{\mu\nu} D\phi^a e^{iS[\phi,e] + iS_{EH}[e]}$$

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- QFT setting no strings or other non-QFT structures
- Diffeomorphism is like a gauge symmetry [Hehl et al.'76]
 - Arbitrary local choices of coordinates do not affect observables – pure passive formulation

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Standard gravity

- Integration variable currently arbitrary choice
 - Manifold topolgies when factoring out diffeomorphisms
 - Other choices (e.g. vierbein) possible when integrating over diffeomorphism orbits



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- Setup of FRG/Asymptotic safety, CDT, EDT

 $\langle O \rangle = \int_{\Omega} Dg_{\mu\nu} D \phi^a O e^{iS[\phi,e] + iS_{EH}[e]}$

$\bigcup_{\mathbf{A}} \neq \langle O \rangle = \int_{\Omega} Dg_{\mu\nu} D \phi^a O e^{iS[\phi,e] + iS_{EH}[e]}$ non-zero

$0 \neq \langle Q \rangle = \int_{\Omega} Dg_{\mu\nu} D \phi^a O e^{iS[\phi,e] + iS_{EH}[e]}$ Needs to be invariant to be non-zero

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Needs to be invariant

- Locally under Diffeomorphism
- Locally under Lorentz transformation

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- Locally under Diffeomorphism
- Locally under Lorentz transformation
- Locally under gauge transformation
- Globally under custodial,... transformation to be non-zero

[Ambjorn et al.'12,19, Maas'19]

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 - E.g. again curvature scalars
 - Arguments/distances need to be invariants
 - E.g. geodesic distances [Ambjorn et al.'12, Schaden '15, Maas'19]

Causal Dynamical Triangulation [Ambjorn et al.'12,19]

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 $\langle O \rangle = \int_{\Omega} D d(X, Y) O e^{iS_{EH}[e]}$ Restricted to foliable, pseudo-Riemannian manifolds with fixed global structure

Causal Dynamical Triangulation



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Wick rotation allows use of standard Monte-Carlo (lattice) techniques

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Systematic errors due to statistics, finite volume and discretization

Space-time in CDT



Space-time in CDT

Maximum volume constrained by finite size



Space-time in CDT



Periodicity by boundary conditions

Space-time in CDT



Space-time in CDT

[Ambjorn et al.'12,'19 Maas, Plätzer, Pressler'25]



Space-time in CDT



Space-time in CDT



Curvature scalar R can be numerically Estimated from Q

Space-time in CDT



Space-time in CDT



Only cosmological constant, no inflation!

Space-time in CDT



An average gauge-fixed metric would be close to the de Sitter-like

Like in classical general relativity

Space-time in CDT Maas, Plätzer, Pressler'25] Per configuration histogram of the size - relatively similar, no large fluctuations Spatial global observables ∛#Simplices 10 5 <Q> 0.5 100 25 Ħ $<|\Delta Q|>$ 50 20 0.1 -40-20 0 2 15 ³√Simplices -20 10 0 τ An average gauge-fixed metric 20 would be close to the de Sitter-like

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 - Spin of particles can be locally changed spin cannot be an observable

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Composite operators need to be completely neutral

[Maas'23]

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 - Consider supergravity multiplet and scalar multiplet
 - Linear representation

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 - But what is spin?

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 - Expansion of composite yields effectively a fermion

$$S = \begin{vmatrix} \phi \\ \omega \\ D \end{vmatrix}^{+} \Gamma D \begin{vmatrix} e \\ \psi \\ F \end{vmatrix} \stackrel{e \sim \delta, \psi \sim \Delta, \phi \sim v}{=} \operatorname{const} + \omega^{+} \Gamma$$

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 Yields an new view on supersymmetry ...and why it may escape conventional detection