





Latest results from the XENONnT experiment

Dr. Maxime Pierre

On behalf of the XENON collaboration

maxime.pierre@nikhef.nl

UNIVERSITEIT VAN AMSTERDAM Nikhef ×××







Light and Charge readout

- Prompt scintillation signal (S1)
- Secondary proportional scintillation signal in GXe from drifted electrons (S2)

Event reconstruction

- O 3D Position:
 - Z from drift time
 - (X, Y) from PMTs hit pattern \odot
- Energy $\rightarrow E = W \cdot (n_{ph} + n_e)$

Particle discrimination

Interaction type Nuclear Recoil (NR)/Electronic \bigcirc Recoil (ER) through S1/S2 ratio

$$\left(\frac{S2}{S1}\right)_{NR} < \left(\frac{S2}{S1}\right)_{ER}$$









Primary Goal is to search for direct detection of WIMPs interaction... but can do much more!

Maxime Pierre maxime.pierre@nikhef.nl

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Time [YYYY-MM, UTC]

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Time [YYYY-MM, UTC]





- Radon suppression milestone: distillation with combined gaseous and liquid xenon flow
- Stable detector response: achieving < 0.3%
 - (1.1%) variation in Light and Charge Yields





Time [YYYY-MM, UTC]

³⁷Ar Maintenance & distillation (S1-only)

2023.07

Science Run 1 Highlights

- Radon suppression milestone: distillation with combined gaseous and liquid xenon flow
- Stable detector response: achieving < 0.3%
 - (1.1%) variation in Light and Charge Yields
- Excellent electron lifetime $\sim O(10)$ ms









Time [YYYY-MM, UTC]

Maintenance & distillation (S1-only)

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Science Run 1 Highlights

- Radon suppression milestone: distillation with \bigcirc combined gaseous and liquid xenon flow
- Stable detector response: achieving < 0.3%
 - (1.1%) variation in Light and Charge Yields
- Excellent electron lifetime $\sim O(10)$ ms
- Calibration plan follows SR0 strategy with three new features:
 - New ER calibration source: 222 Rn (β) and 232 Th (χ)
 - New low-energy NR calibration source: ⁸⁸YBe

















New Result from XENONnT

Next science case focus?

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New Result from XENONnT



Entering the Neutrino Fog Coherent Elastic neutrino Nucleus Scattering

Maxime Pierre

maxime.pierre@nikhef.nl

Never measured in a DM direct detection experiment Never measured in a xenon target **Never** measured from astrophysical source





The ⁸B Solar Neutrino CE_vNS Gate

$CE_{\nu}NS$ search in XENON:

Nearly indistinguishable from a ~6 GeV WIMP with $\sigma_{SI} = 4.4 \text{ x}$

 $10^{-45}\,\mathrm{cm}^2$

Boost in cross section from \bigcirc coherent effect...but low energy recoil (< 1.5 keV_{NR} in LXe)





⁸B CE_vNS - Signal Region of Interest Maxime.pierre@nikhef.nl

Boost sensitivity by lowering our energy threshold

⁸B event rate in the conventional WIMP "3-fold analysis" (SR0): ~1% detection efficiency \rightarrow 0.2 events / (t x yr)

Lowering our S1 and S2 threshold → improve our expected event rate to 3.7(3.3) events / (t x yr) in SR0(1)





Boost sensitivity by lowering our energy threshold

Blinded Region of Interest

S1 ROI: 2 or 3 hits ; A hit corresponds to a recorded photon by PMT+DAQ+software **S2 ROI:** [120 - 500] PE \rightarrow Reject high rate of isolated S2 background signal



Main Source: Accidental Coincidences

- Accidental Coincidence (AC): Random pairing of isolated S1 and **isolated S2**, whose exact origin is under investigation. Current culprit:
 - **Isolated S1 signals**: from pile-up induced single PMT hits, misclassified single electrons,...
 - **Isolated S2 signals**: from few-electron pile-up events, notably following high-energy interactions,...

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Background Model

Accidental Coincidences Suppression:

1-Time Shadow

- Use space/time correlation with previous high-energy interaction \bigcirc
 - → Isolated S1 rate: $15 \text{ Hz} \rightarrow 2.3 \text{ Hz}$
 - **Isolated S2 rate**: 150 mHz \rightarrow 25 mHz

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Background Model

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2- Two Boosted Decision Tree (BDT)

- **S1 BDT:** leverage S1 pulse shape and spatial \bigcirc distribution across the PMT arrays.
- **S2 BDT:** check that S2 pulse shape correlated with the diffusion of the drifting electron cloud law.

Expected # of AC events:

 7.5 ± 0.7 (SR0) and 17.8 ± 1.0 (SR1)

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Yields model from ⁸⁸YBe Calibration Maxime Pierre maxime.pierre@nikhef.nl

Constrain Yield Models:

- Great agreement between data and model.
 - Background originating from Accidental Coincidences (AC) are modelled with data-driven simulation framework.
- Light (LY) and Charge Yields (CY) were extracted down to 0.5 keV_{NR} at XENONnT electric field of 23 V/cm with latest NEST parametrisation.
- Yield model uncertainty leads to ~ 30% signal rate uncertainty.

 $\gamma(^{88}\text{Y}) + ^{9}\text{Be} \rightarrow n + ^{8}\text{Be}$

⁸B CE_vNS search - Prediction

- (4.1) tonnes in SRO (SR1) leading to a total exposure of 3.51 t x yr

48% to observe > 3\sigma significance

Component	Expectation	Best-fit
AC (SR0)	$7.5~\pm~0.7$	
AC (SR1)	$17.8~\pm~1.0$	
\mathbf{ER}	$0.7~\pm~0.7$	
Neutron	$0.5\substack{+0.2 \\ -0.3}$	
Total background	$26.4^{+1.4}_{-1.3}$	
⁸ B	$11.9\substack{+4.5 \\ -4.2}$	
Observed		

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• **B neutrino flux**: $4.6^{+3.6}_{-2.3} \times 10^{6}$ cm⁻² s⁻¹ at 68% C.L. no tension with literature value

Background only hypothesis rejected with 2.73σ significance

Strong evidence of CEvNS Interaction

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⁸ B	$11.9\substack{+4.5 \\ -4.2}$		$10.7\substack{+3.7 \\ -4.2}$
Observed		37	

⁸B CE_vNS search - Unblinding

- **B neutrino flux**: $4.6^{+3.6}_{-2.3} \times 10^6$ cm⁻² s⁻¹ at 68% C.L. no tension with literature value

with 2.73σ significance

With constrain from SNO flux \rightarrow Measure the flux-weighted CE_vNS cross section: $1.1^{+0.8}_{-0.5} \times 10^{-39}$ cm²

A reminder that CE_vNS is a background for DM search

- Focus on light DM models such as asymmetric and self-interacting DM
- Benefit from the work done for the 8B CE ν NS analysis

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Strong evidence for CEvNS interaction in XENONnT: A first step into the neutrino fog

First Light DM results near the neutrino fog

More results to come:

- Stay tuned for the **incoming WIMP results** with SR0+SR1 datasets
- We continue to take data, SR2 ongoing with Gd-loaded Water in nVeto to improve our neutron tagging efficiency.
 - → Improved significance for ⁸B CE ν NS with increased exposure
 - Broad physics program with active analyses ongoing whose scope goes beyond WIMP search

- R0+SR1 datasets loaded Water in
- eased exposure ongoing whose

Back-Up

Background Model

Subdominant Source:

Background Model

Subdominant Source:

Nuclear Recoil:

- Radiogenic neutrons (fission, α-n),
 simulation and data-driven (NV) model
 - Projection SR0: 0.13 ± 0.07 events
 - Projection SR1: 0.33 ± 0.19 events

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Electronic Recoil:

- Dominated by ²¹⁴Pb flat β-spectra
 - Projection SR0: 0.13 ± 0.13 events
 - Projection SR1: 0.56 ± 0.56 events

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Background Model

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Surface Event:

- ER from ²¹⁰Pb plate out at detector walls
 - Data-driven model SR0/1: <0.3 events</p>

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Yields model from ⁸⁸YBe Calibration Maxime Pierre maxime.pierre@nikhef.nl

New calibration source:

- Low-energy NR calibration using external photoneutron source.
- Quasi-monoenergetic 152 keV \bigcirc neutrons produced via:

 $\gamma(^{88}\text{Y}) + ^{9}\text{Be} \rightarrow n + ^{8}\text{Be}$

⁸B CE_vNS search - Unblinding

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Yields uncertainties

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 $\left(t_{ly},t_{qy}\right)$ two morphers of the yields: uncertainties of the emission model

 $LY(t_{ly}) = \langle LY \rangle + t_{ly} \cdot \sigma_{LY}(sign(t_{ly}))$ $QY(t_{qy}) = \langle QY \rangle + t_{qy} \cdot \sigma_{QY}(sign(t_{qy}))$

with: $t_{ly} \sim N(0,1); t_{qy} \sim N(0,1)$

- Waveform-feature-based S1 BDT differentiates isolated S1 signals from random PMT hit clustering.
- Input features: double photo-electron emission, S1 pulse shape, S1 hit counts, PMT channel distribution of S1.
- Trained with a data-driven sample of isolated S1 and simulated ⁸B S1
- S1 area in the largest-contributing PMT is the most important feature due to the signal-only double photoelectron emission (DPE), where a single photon striking the PMT photocathode produces two photoelectrons with $p \approx 0.2$.
- Enhances signal vs. background discrimination but is significantly weaker than the S2 BDT.

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BDT trained using simulated signal and datadriven AC background, with each feature rigorously validated between data and simulation.

and data.

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AC Validation

VALIDATION OF AC MODEL

arXiv:2408.02877

- Validated by AC sideband unblinding (events that failed S2 BDT cuts) Validated by ³⁷Ar L-shell 0.27 keV ER calibration data
- The difference (<10%) is considered when determine systematic uncertainty Constrained ER light yield with 1598 observed events

Maxime Pierre

maxime.pierre@nikhef.nl

Credit slide: Langing Yuan

- Unlike WIMP, the B8 FV was not optimized based on signal and bkg predictions. It was selected to:
 - top/bottom → no areas with limited detector modelling
 - radius → minimize surface bkg to a negligible level.

Credit slide: Matteo Guida

- Events near wires are excluded from analysis due to insufficient simulation fidelity.
- S2 pulse shape varies near perpendicular wires, causing systematic errors if S2 BDT (trained on simulation) is applied.

Signal and Bac	kgrou
AC: uncertainty from discrepancy in derived the sideband unblinding.	d by
B8: 35% uncertainty from yields & efficience Flux is a free parameter.	с у. Е
ER: Electronic recoil background with flat spectrum 0-10 keV Conservative 100% uncertainty from yields.	Neutro
RG: Radiogenic neutron background, 58% uncertainty derived from the sideband.	Total backgrour
Surface background: not included in likelihe Fiducial volume such that it can be neglected	ہ ood. ed.
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From XENON1T to XENONnT

