

# Carte Blanche

James Binney

Rudolf Peierls Centre for Theoretical Physics

University of Oxford

# What have I learnt?

- So much about low-mass scalar fields!
  - (Pearl, Jens, Raquel, Marco)
  - A good route to a better understanding of wave-particle duality
- Fuzzy dark matter probably doesn't cut it
  - Over-heats disc galaxies (Jens)
  - Doesn't always provide enough power at high  $k$
- GC  $\gamma$ -ray excess probably from stars (Silvia Manconi)
- A dash of strong lensing strengthens weak lensing (Natalie Hogg)
  - and a geometrical measure of  $H_0$  is high

# JWST over-turning $\Lambda$ CDM?

- JWST sees stars & gas illuminated by stars (ALMA)
- The physics behind IGM  $\rightarrow$  stellar pop is
  - in principle understood
  - uncomputable
- $\Lambda$ CDM makes few reliable predictions
- It's exciting so look back to  $\sim 400$  Myr after BB,
  - just enjoy the movie

# Biggest takeaway?

- QCD axions properly motivated (Wilczek 1978)
  - They *should* exist (positrons, neutrinos, Higgs)
  - They *will be* found! (Fabrice Hubaut)
- The search reminds me of LIGO in 1980s
  - Senior (& v. smart) member of Princeton community said “NSF is wasting our money..”
- Dozens of colleagues laboured over decades rewarded only by small decreases in the noise
- Their heroism is inspiring
- axion detection will be an ever more heroic challenge
  - G-wave detectors knew precisely where they had to get
  - A-wave detectors don't know where the mountain top lies!

# Just a thought

- Rodrigo Vicente reminded us that GW detectors measure  $A$  not  $|A|^2$
- So Volume  $\sim$  sensitivity<sup>3</sup> not sensitivity<sup>3/2</sup>
- Surely it's possible to do wit emag what they do with gravity?

$$\psi = |\psi|e^{i\phi} \quad ; \quad \nabla\phi = \mathbf{p}/\hbar$$

$$\Delta\phi = \int d\mathbf{x} \cdot \nabla\phi = \frac{1}{\hbar} \int d\mathbf{x} \cdot \mathbf{p}$$

Emag wave polarised along  $x$  moving along  $z$

$$\frac{d\mathbf{p}}{dt} = e\mathbf{E} = -e \frac{\partial \mathbf{A}}{\partial t} \quad ; \quad \mathbf{A} = (A, 0, 0) \cos(kz - \omega t)$$

$$\Delta p_x = -eA \cos(\omega t)$$

Interferometer at  $z = 0$  with arms along  $x$  and  $y$  axes

$$\begin{aligned} \Delta\phi &= -eA \left( \int_0^L dx \cos(\omega t) + \int_L^0 dx \cos(\omega t) \right) \\ &= -eAv \left( \int_0^{L/v} dt \cos(\omega t) - \int_{L/v}^{2L/v} dt \cos(\omega t) \right) \\ &= \frac{eAv}{\omega} \left( \sin(\omega T) - [\sin(2\Omega T) - \sin(\omega t)] \right) \end{aligned}$$

e.g.,  $\omega T = \pi/2$  then

$$\Delta\phi = \frac{2eAv}{\omega}$$

# Finally

- Thank you Julien!
  - Thank you SOC!
  - Thank you LOC!
- I really enjoyed being here and we go away with good memories.

$$\psi = |\psi|e^{i\phi} \quad ; \quad \nabla\phi = \mathbf{p}/\hbar$$

$$\Delta\phi = \int d\mathbf{x} \cdot \nabla\phi = \frac{1}{\hbar} \int d\mathbf{x} \cdot \mathbf{p}$$

Emag wave polarised along  $x$  moving along  $z$

$$\frac{d\mathbf{p}}{dt} = e\mathbf{E} = -e \frac{\partial \mathbf{A}}{\partial t} \quad ; \quad \mathbf{A} = (A, 0, 0) \cos(kz - \omega t)$$

$$\Delta p_x = -eA \cos(\omega t)$$

Interferometer at  $z = 0$  with arms along  $x$  and  $y$  axes

$$\begin{aligned} \Delta\phi &= -eA \left( \int_0^L dx \cos(\omega t) + \int_L^0 dx \cos(\omega t) \right) \\ &= -eAv \left( \int_0^{L/v} dt \cos(\omega t) - \int_{L/v}^{2L/v} dt \cos(\omega t) \right) \\ &= \frac{eAv}{\omega} \left( \sin(\omega T) - [\sin(2\Omega T) - \sin(\omega t)] \right) \end{aligned}$$

e.g.,  $\omega T = \pi/2$  then

$$\Delta\phi = \frac{2eAv}{\omega}$$