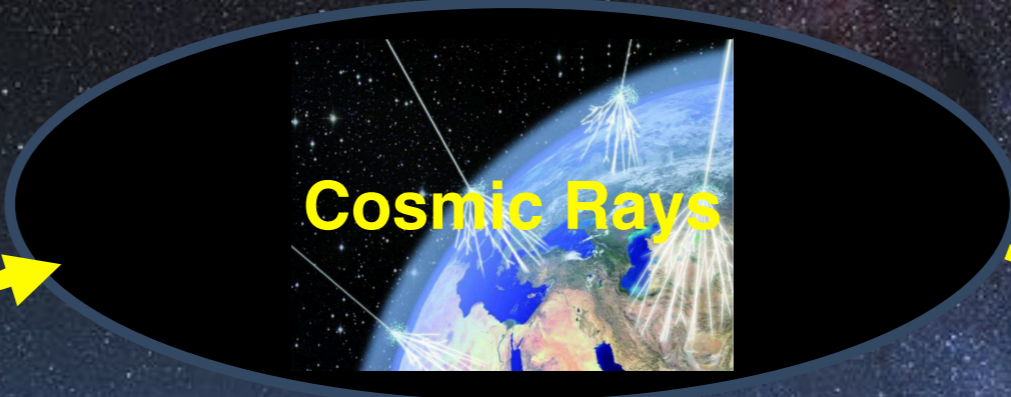
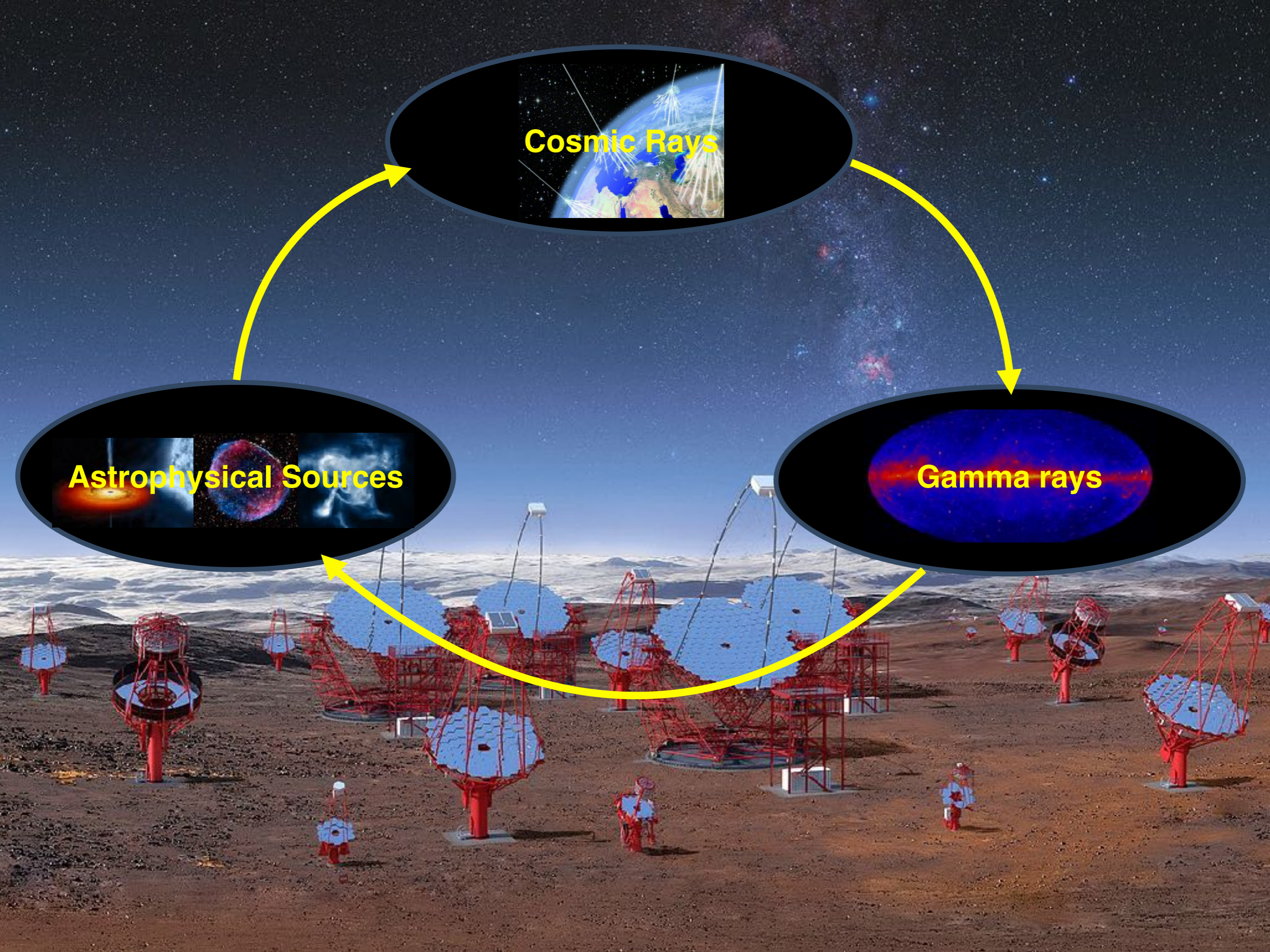


**Session**  
**Astroparticules**

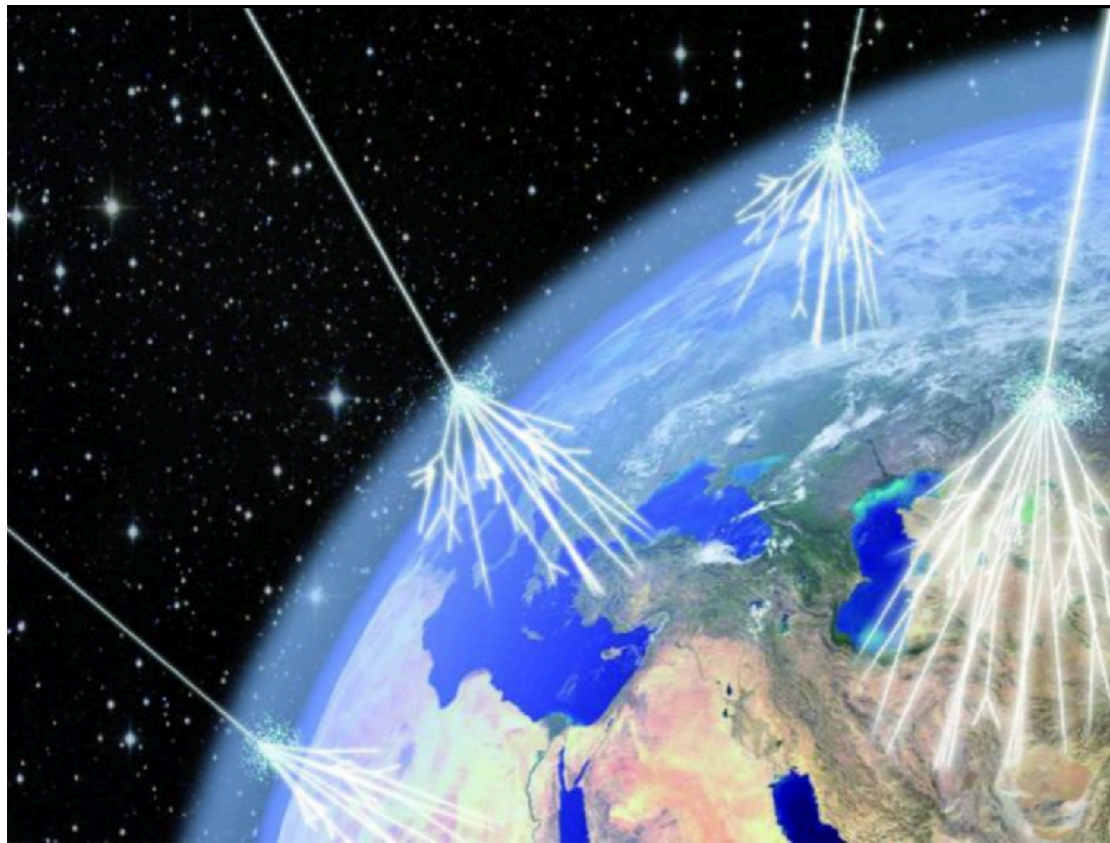
**Justine Devin**

[devin@lupm.in2p3.fr](mailto:devin@lupm.in2p3.fr)

**Laboratoire Univers et Particules de Montpellier (LUPM) - CNRS/IN2P3**

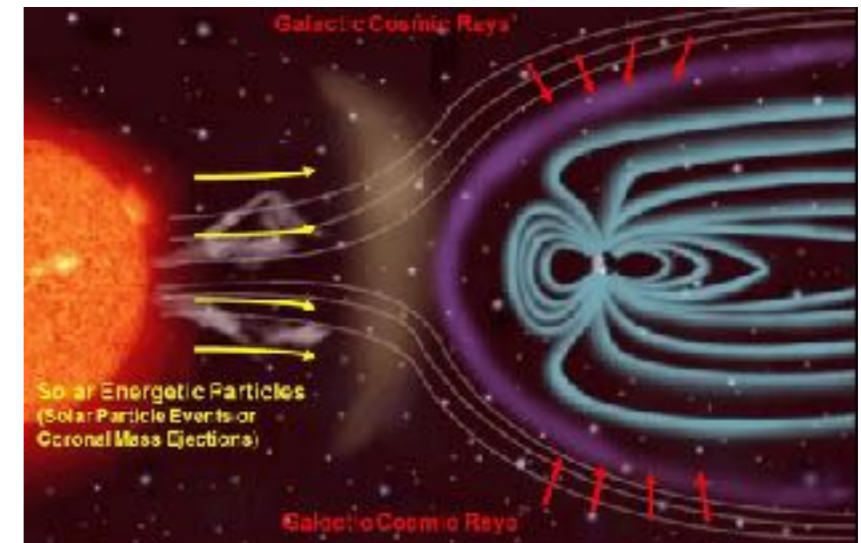


# Discovery of Cosmic Rays

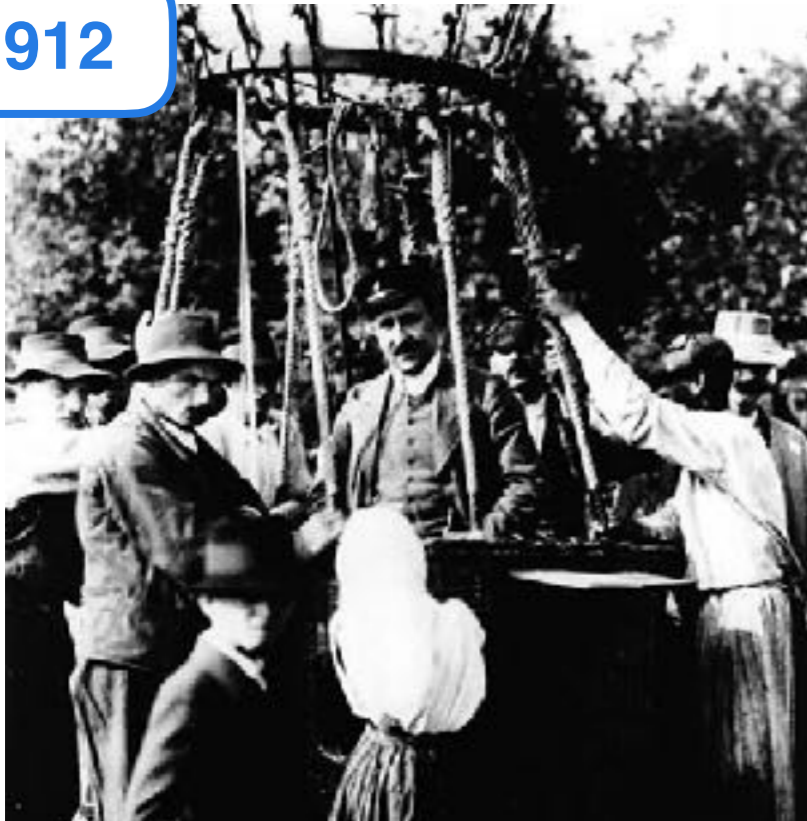


Earth continuously and isotropically strikes by **cosmic particles**

- Terrestrial magnetic field and atmosphere protect us



1912



Measure of an increase of the ionization rate with altitude in the atmosphere

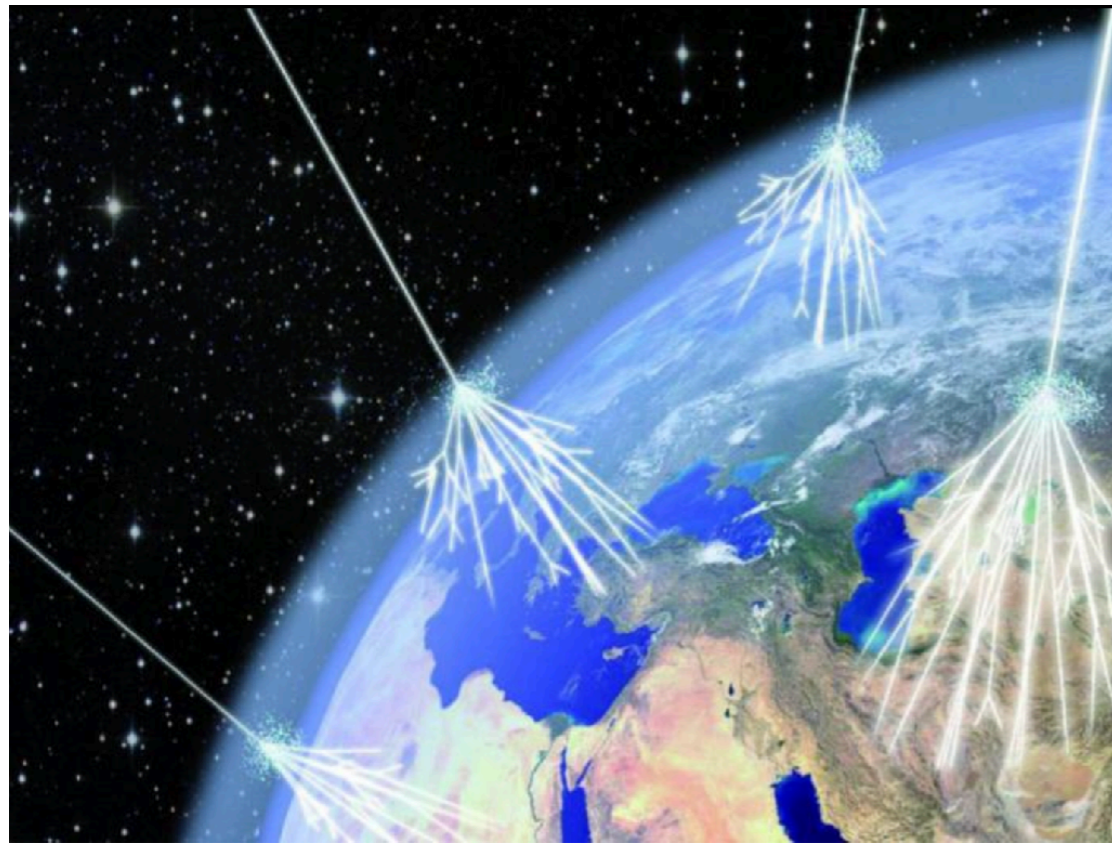
- ➔ Exists ionizing particles whose origin cannot be terrestrial

1936

**DISCOVERY OF COSMIC RAYS**

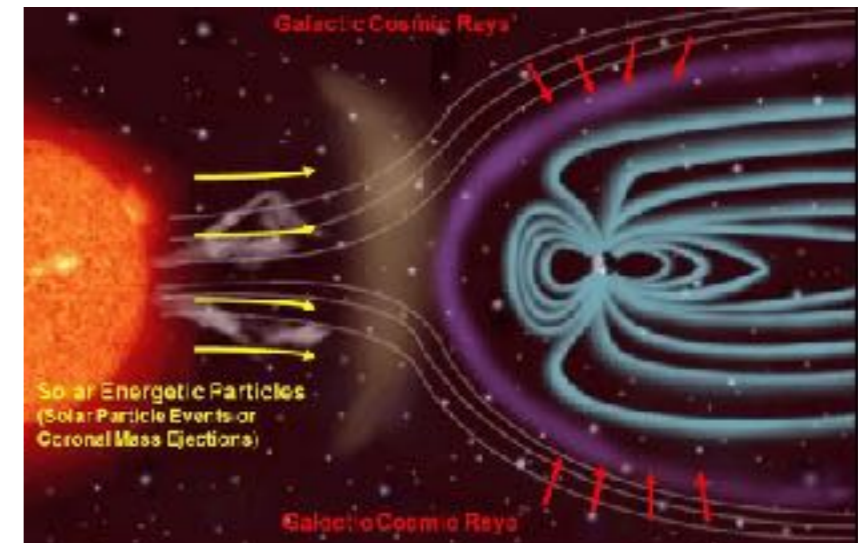


# Discovery of Cosmic Rays



Earth continuously and isotropically strikes by **cosmic particles**

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1912



Measure of an increase of the ionization rate with altitude in the atmosphere

- ➔ Exist ionizing particles whose origin cannot be terrestrial

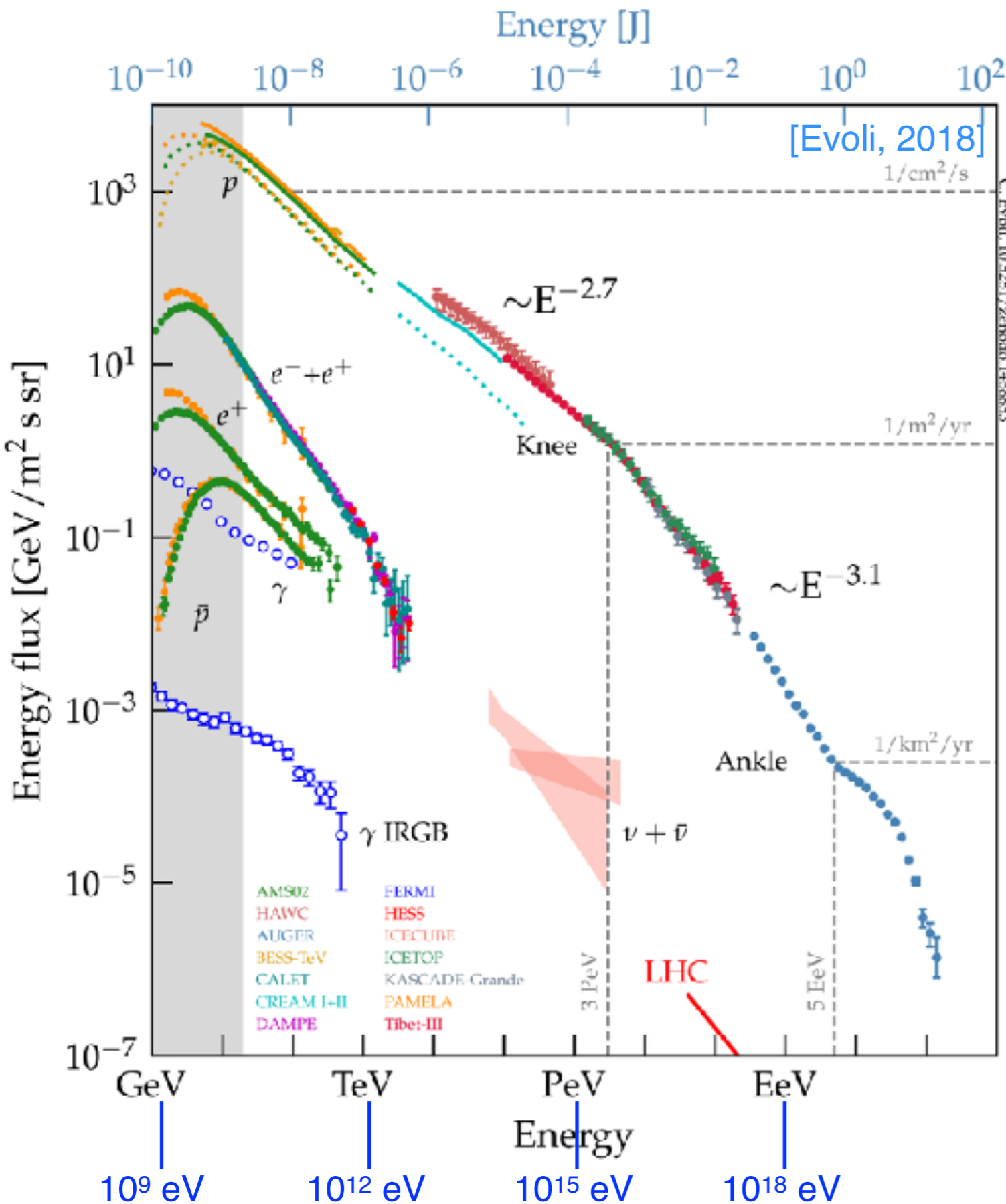
particles

**DISCOVERY OF COSMIC RAYS**

1936



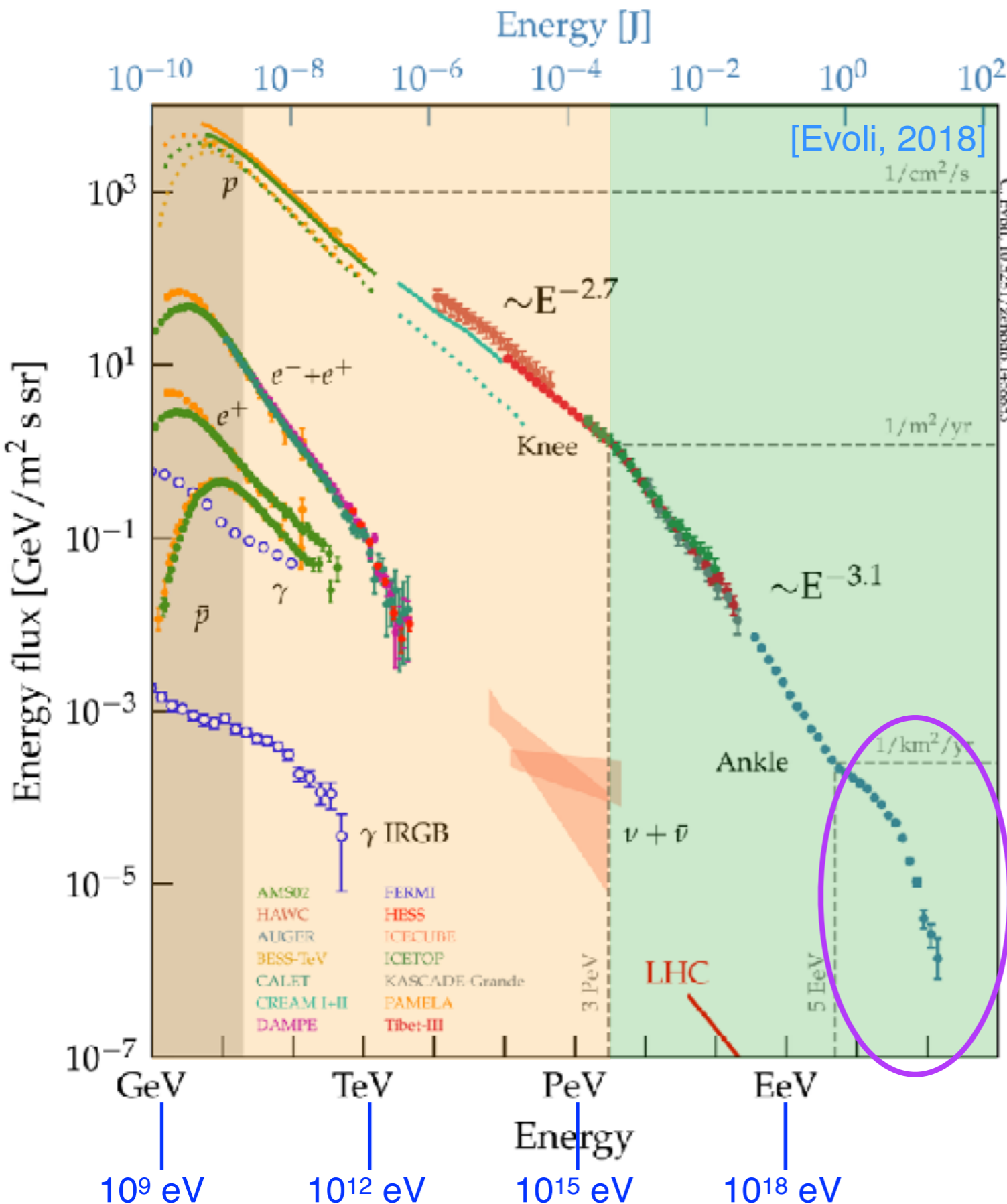
# What are cosmic rays?



## Composition:

- 90% protons
- 9% heavier nuclei
- 1% electrons

# Cosmic-ray spectrum



Origin:

**Galactic**

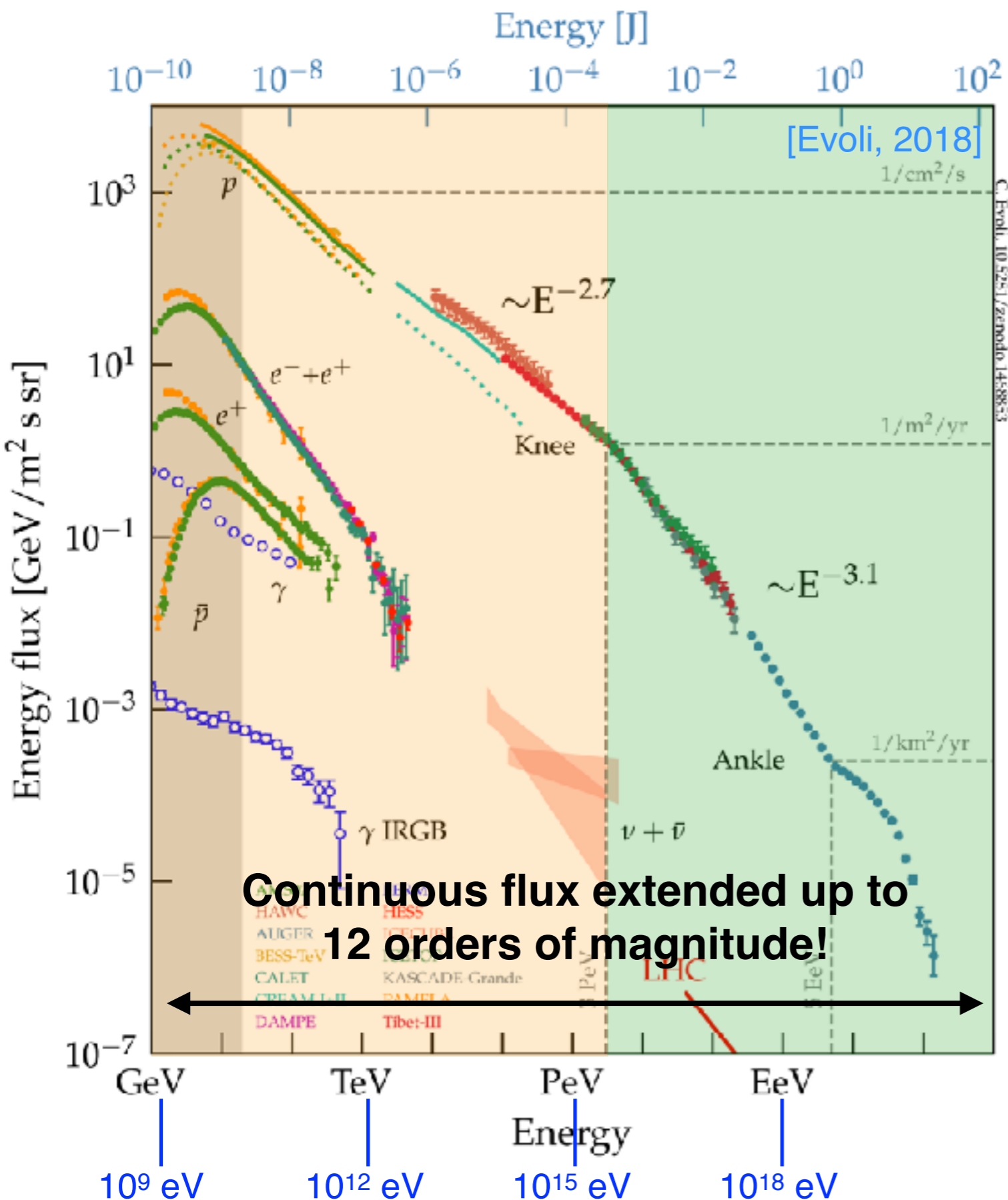
Larmor radius:  $r_L = \frac{E}{ZeB}$

Confined by the galactic magnetic field

**Extragalactic**

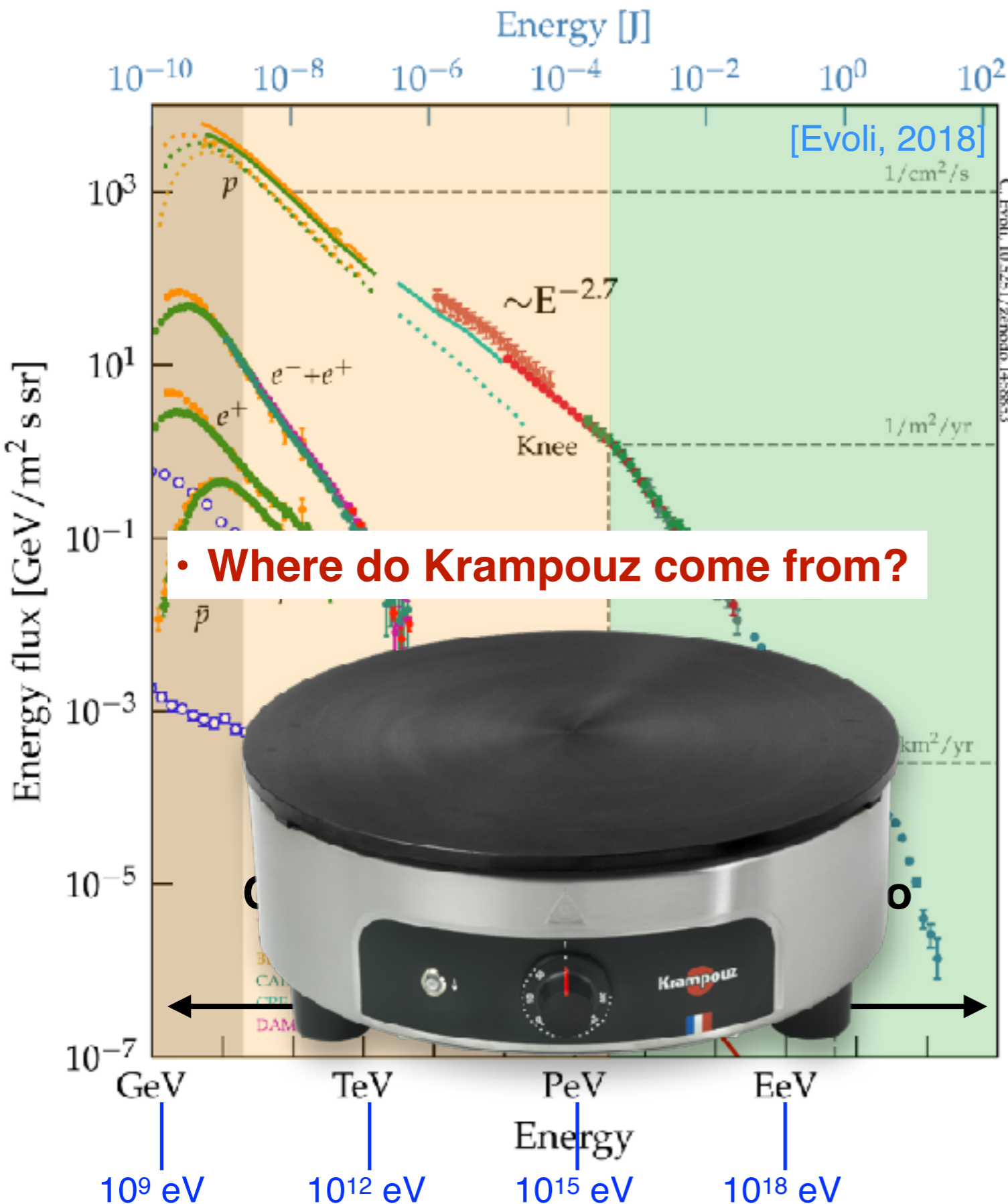
**Opacity:** highest-energy CRs interact with CMB photons

# Where do cosmic rays come from?



- Where did particles get such energy to become a CRs?
- Where and how are they accelerated in the Universe?
- **Where do they come from?**

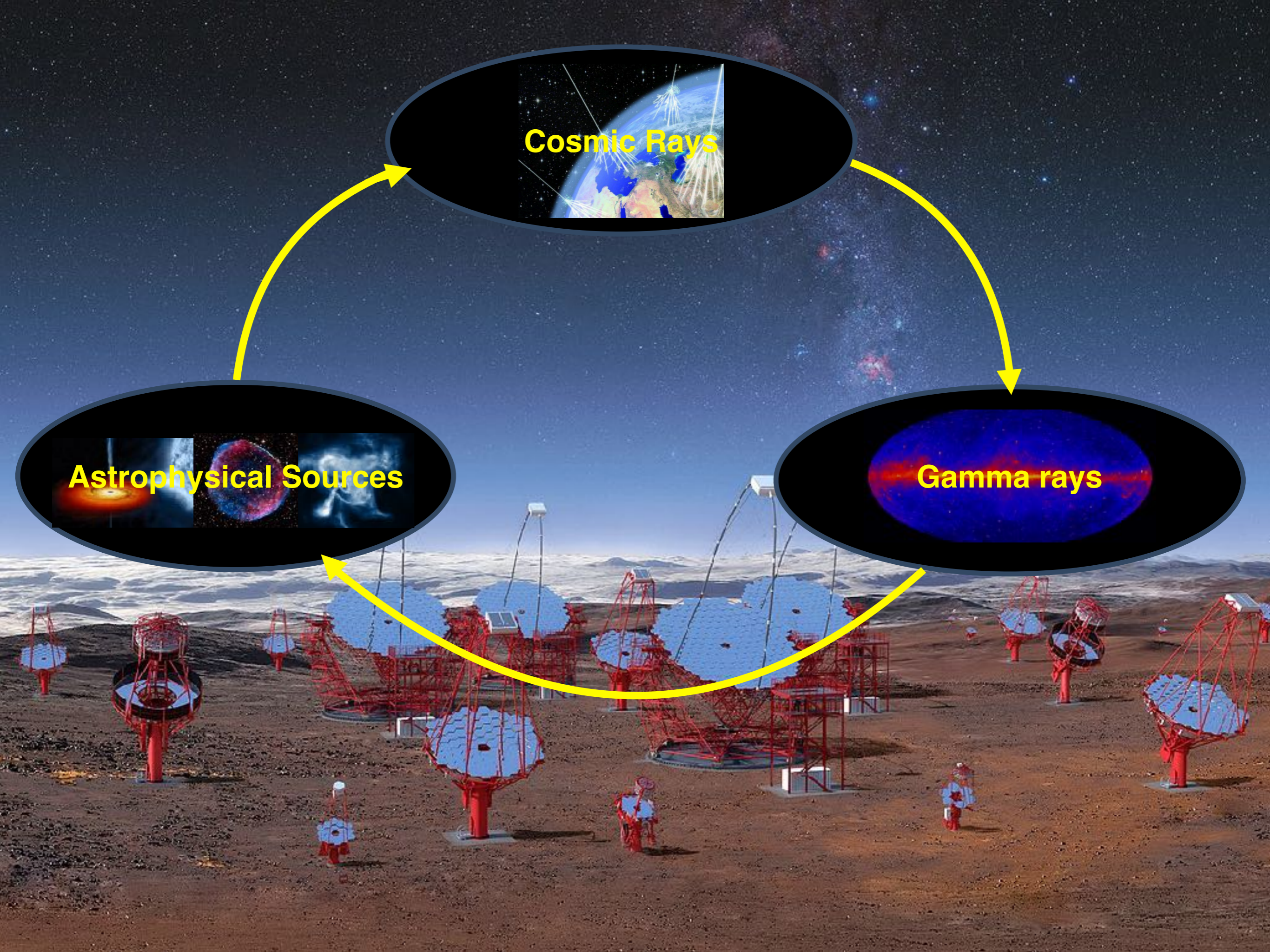
# Where do cosmic rays come from?



- Where did particles get such energy to become a CRs?
- Where and how are they accelerated in the Universe?
- **Where do they come from?**







**Cosmic Rays**

**Astrophysical Sources**

**Gamma rays**

# Our Galaxy: The Milky Way



Several hundreds of billion of stars mostly located on the spiral arms

And orbiting around the supermassive black hole Sagittarius A\*



[Event Horizon Telescope 2022]

Star forming rate (per year):  $1.7 \pm 0.2 \text{ Msun}$  [Licquia & Newman 2015]

**Large amount of activity in the Galaxy where stars born... and die**

**Create**      **Collapse**

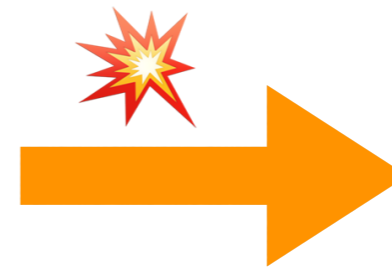
# Cataclysmic events

- $M < 8 M_{\text{sun}}$  : white dwarf + companion star

📌 Supernova remnant



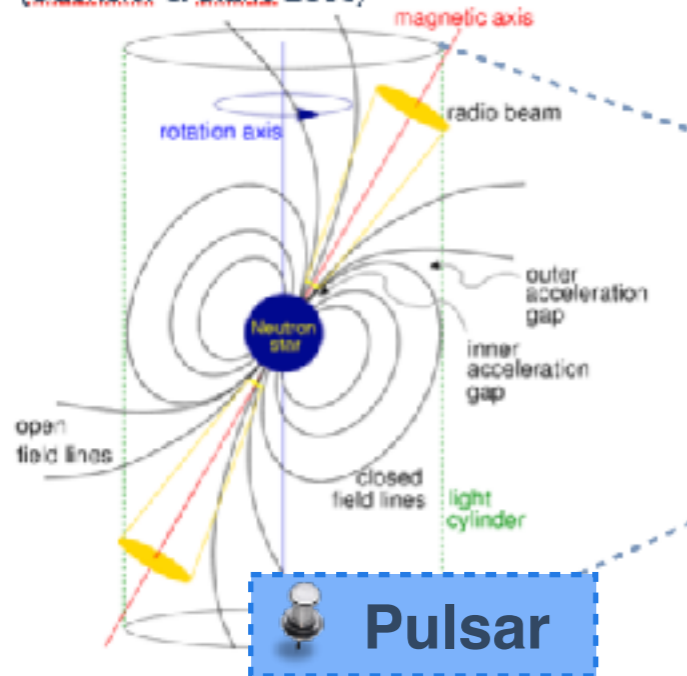
**SUPERNOVA**



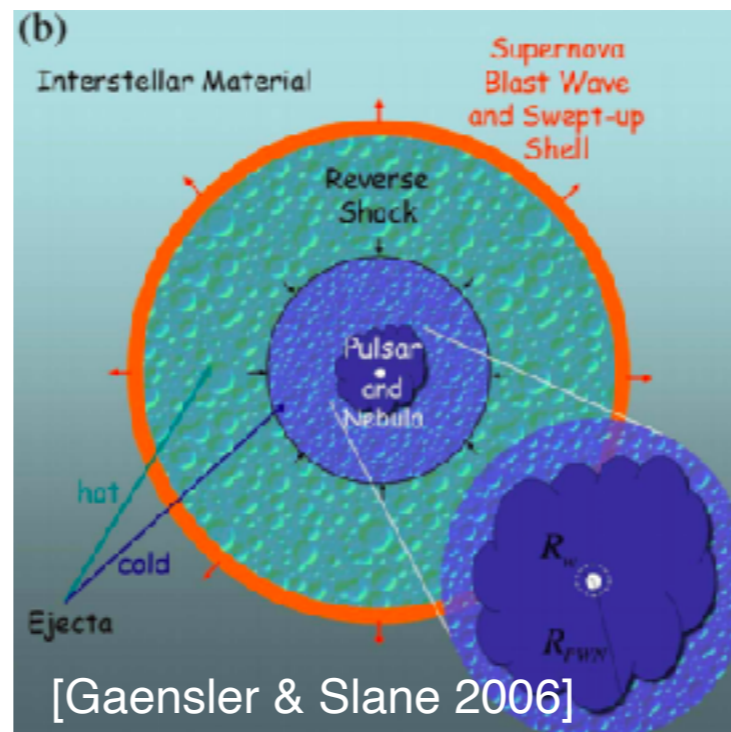
- $8 M_{\text{sun}} < M < 22 M_{\text{sun}}$  : neutron star / pulsar



(Gaensler & Slane 2006)



📌 Pulsar



[Gaensler & Slane 2006]

📌 Supernova remnant

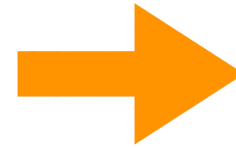


📌 Pulsar wind nebula

1 arcmin

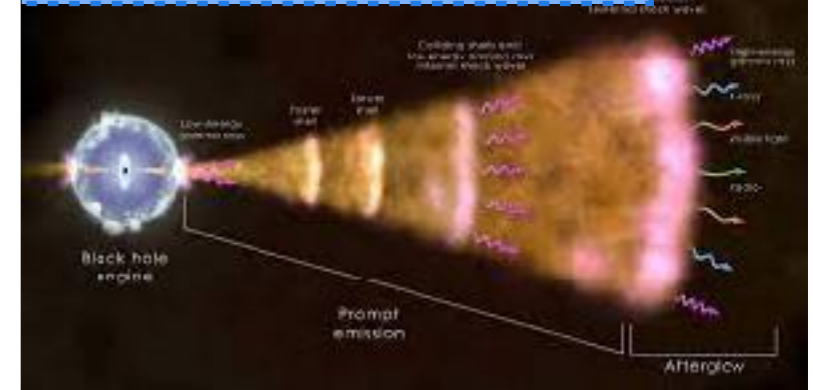
# Cataclysmic events

- $M > 22 M_{\text{sun}}$ : magnetar or black hole

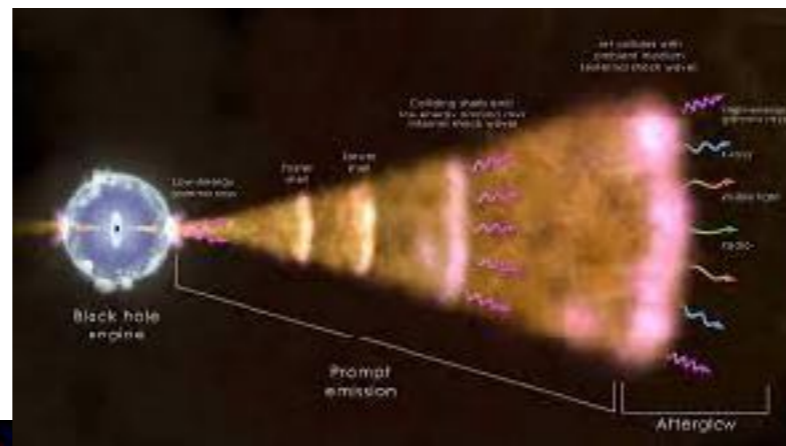
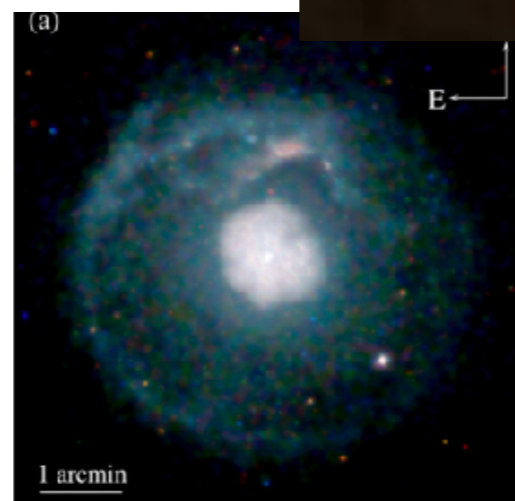


**SUPERNOVA**

## Gamma-ray burst



So  
What?



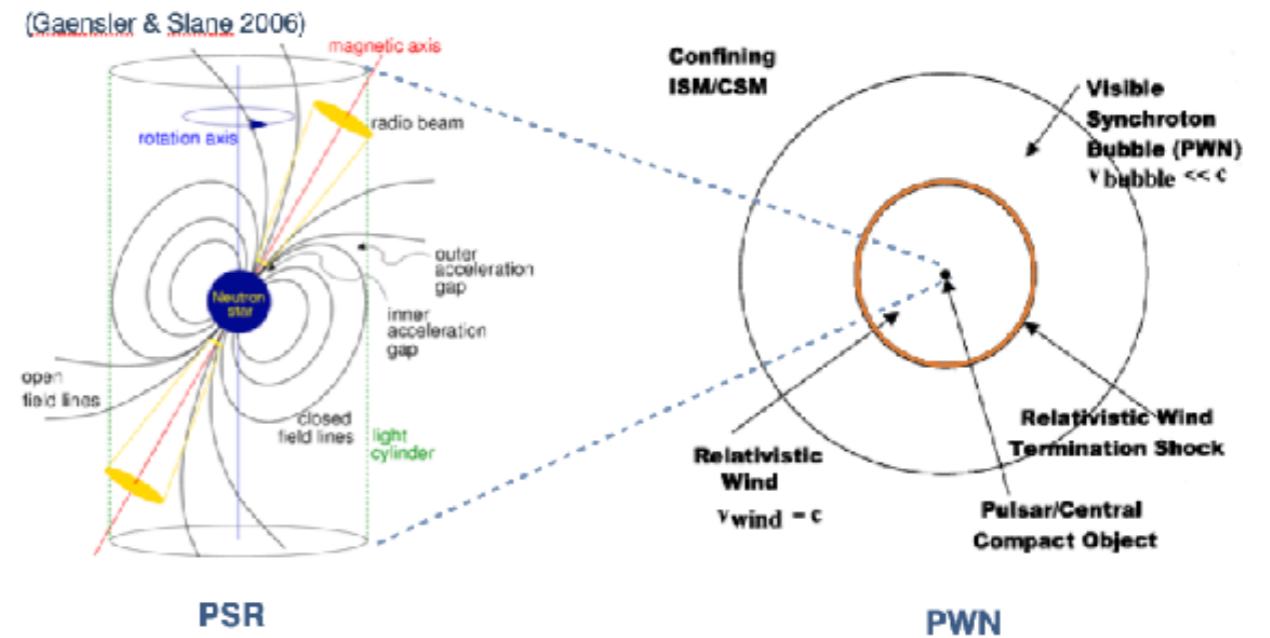
Acceleration of particles must occur within the most violent phenomena  
➔ **Supernova remnants: best candidates for acceleration of CRs!**

# Theory for acceleration of particles at astrophysical shocks

**SN 1006**  $V_{\text{shock}} \sim 5\,000 - 10\,000 \text{ km s}^{-1}$

**Protons or electrons ( $E_{\text{max}}$ )**

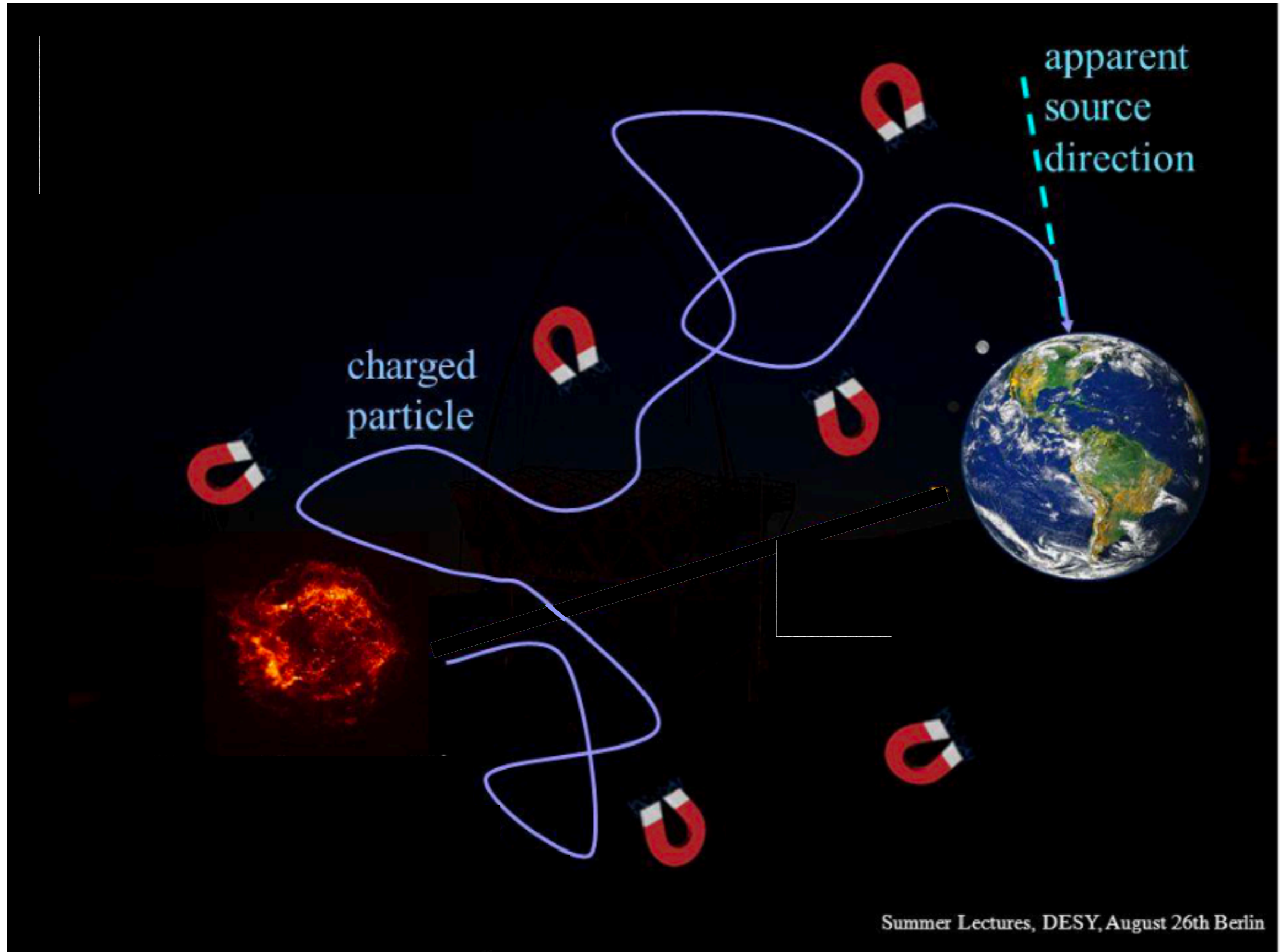
Energy gain up to  $E_{\text{max}}$  then **escape**



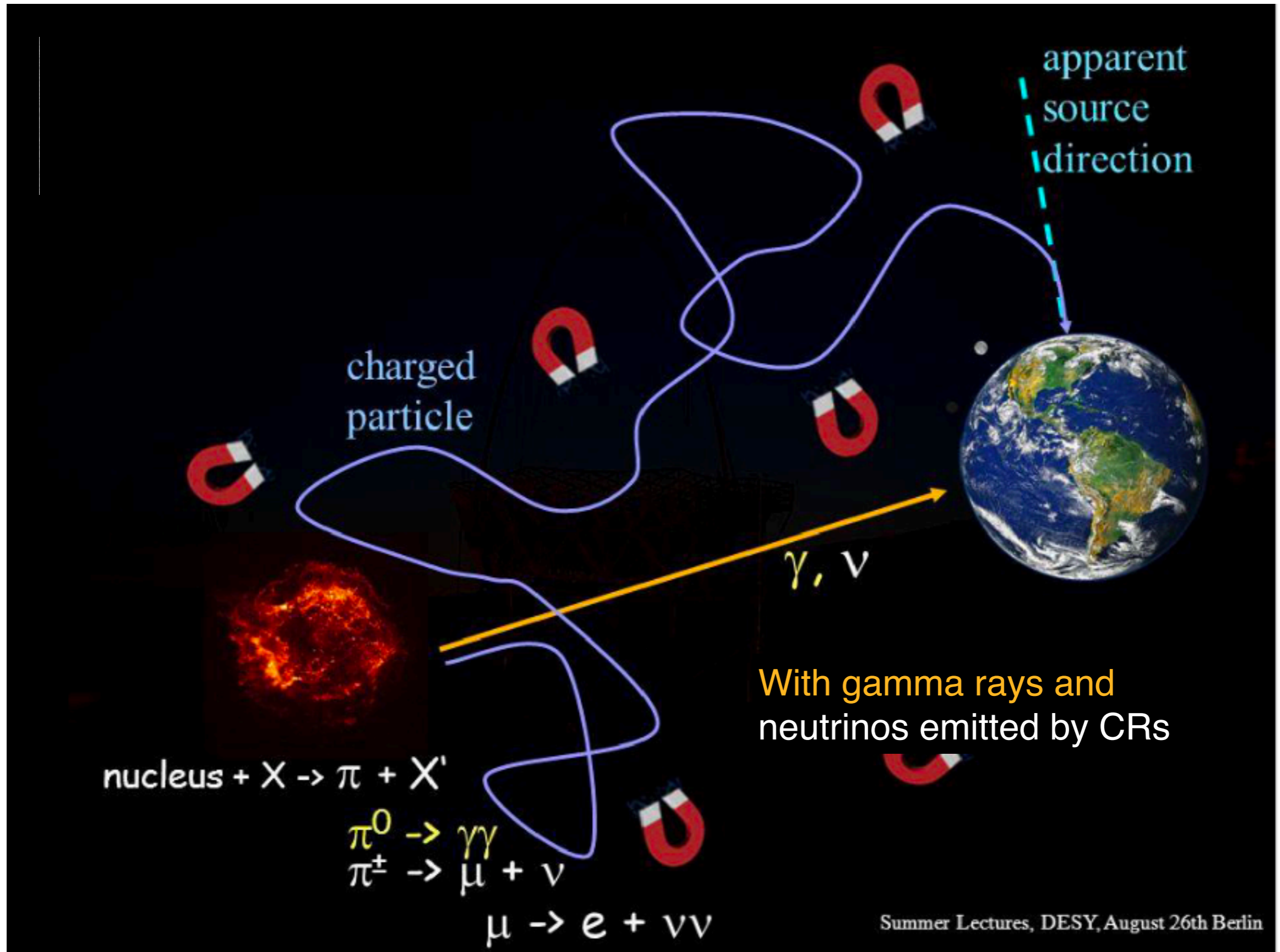
Applicable to any shock front in astrophysical sources

- ➔ Can these astrophysical sources be particle accelerators?
- ➔ And to such high energies measured in the CR spectrum?

# Chasing cosmic-ray sources with gamma rays



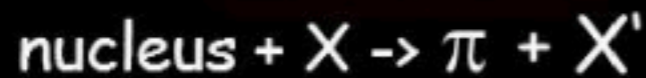
# Chasing cosmic-ray sources with gamma rays



# Chasing cosmic-ray sources with gamma rays

- + Bremsstrahlung (with matter)
- + Inverse Compton (with photon fields)

from accelerated electrons!

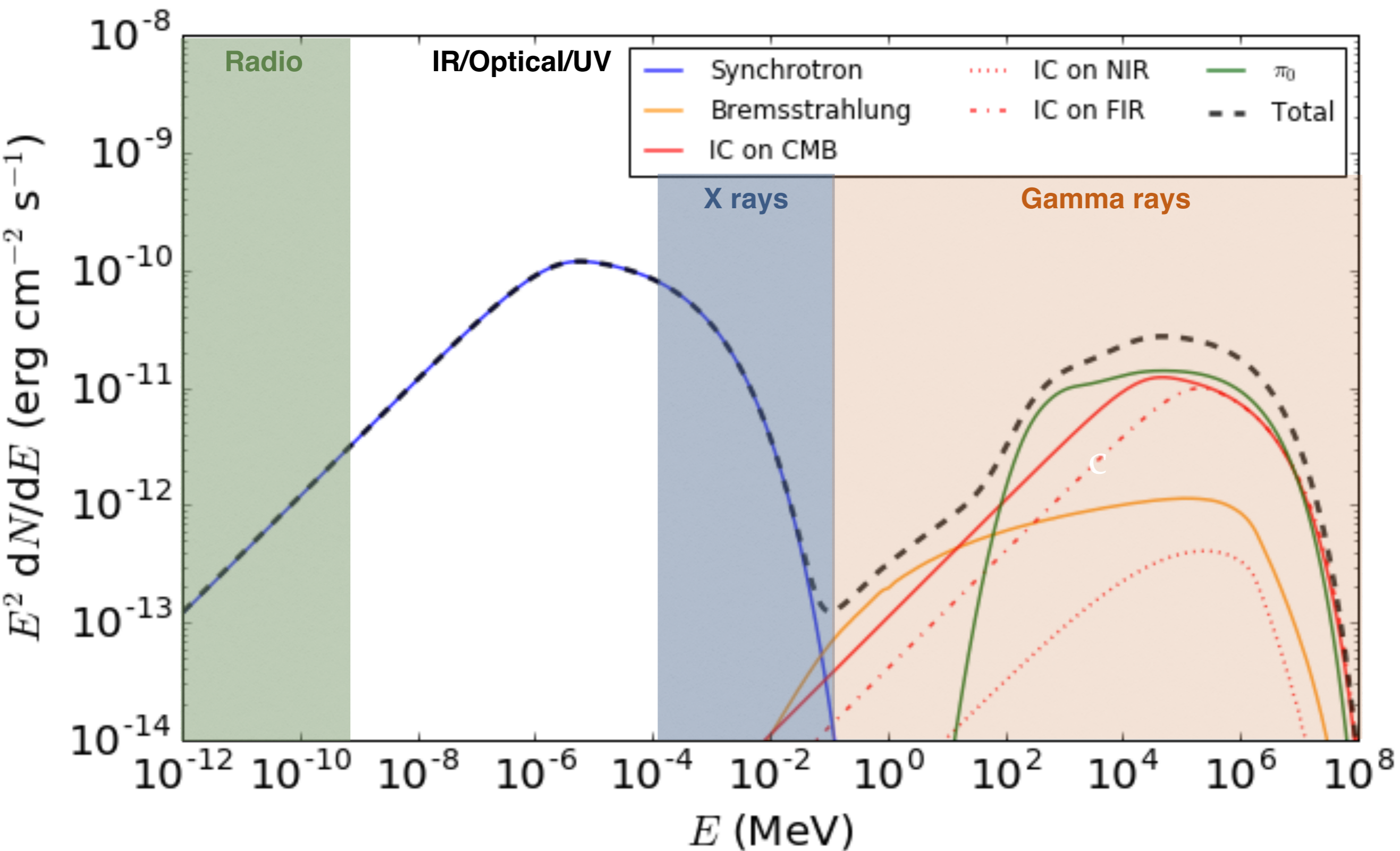


With gamma rays and neutrinos emitted by CRs

Summer Lectures, DESY, August 26th Berlin

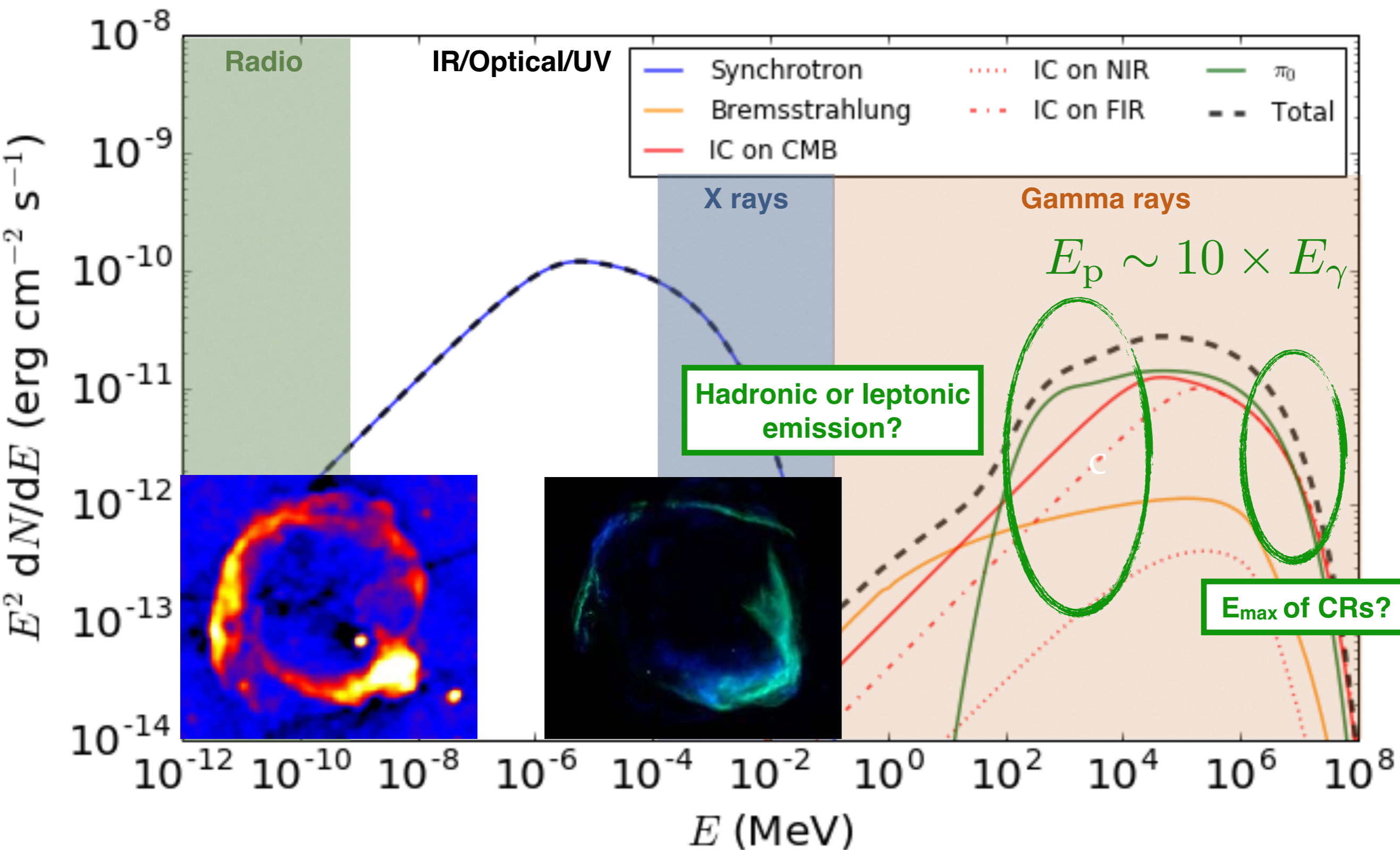


# Broadband modeling of the emission



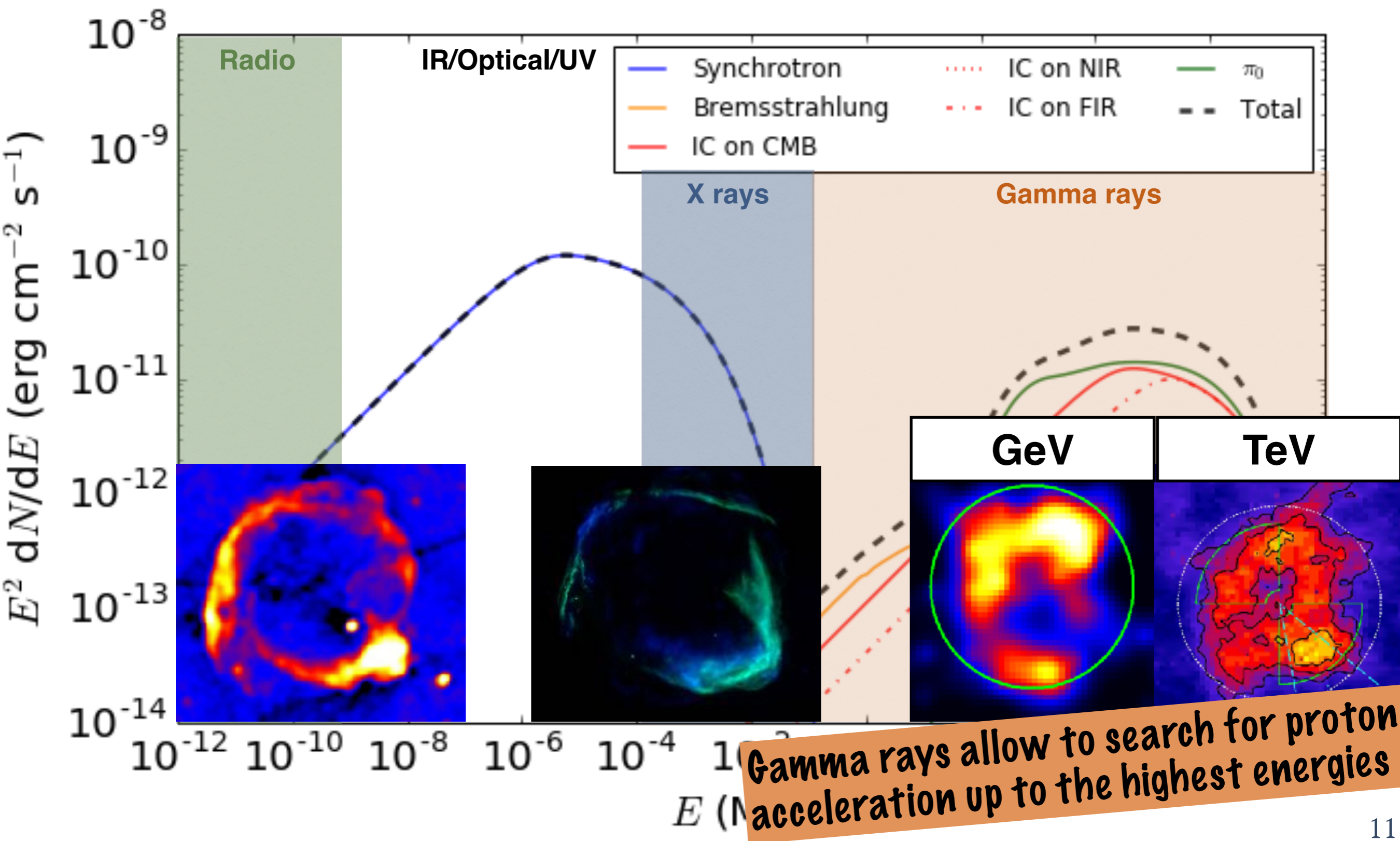
# Broadband modeling of the emission

Nature of the accelerated particles (protons or electrons)?  $E_{\text{max}}$  of accelerated CRs?

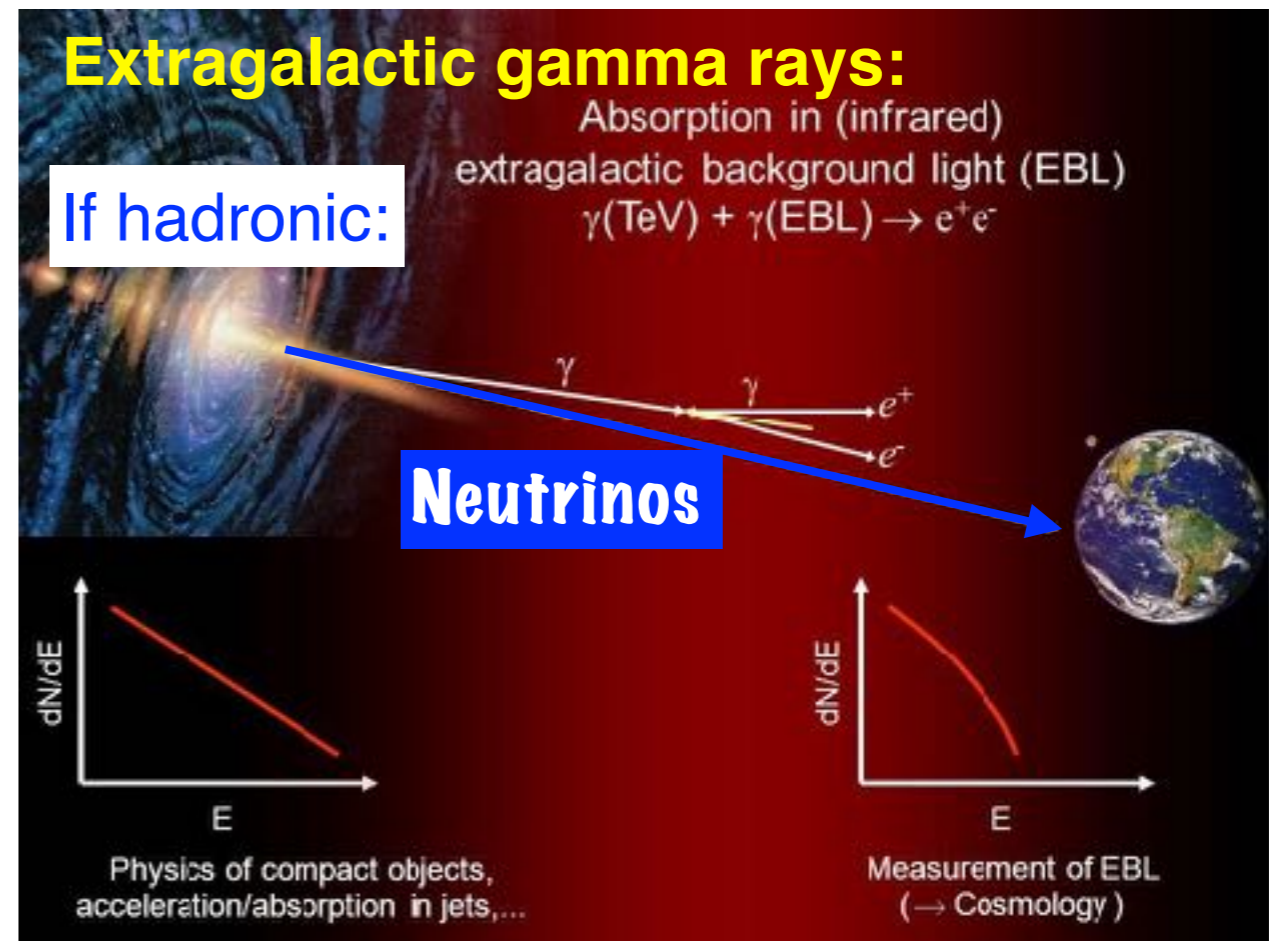
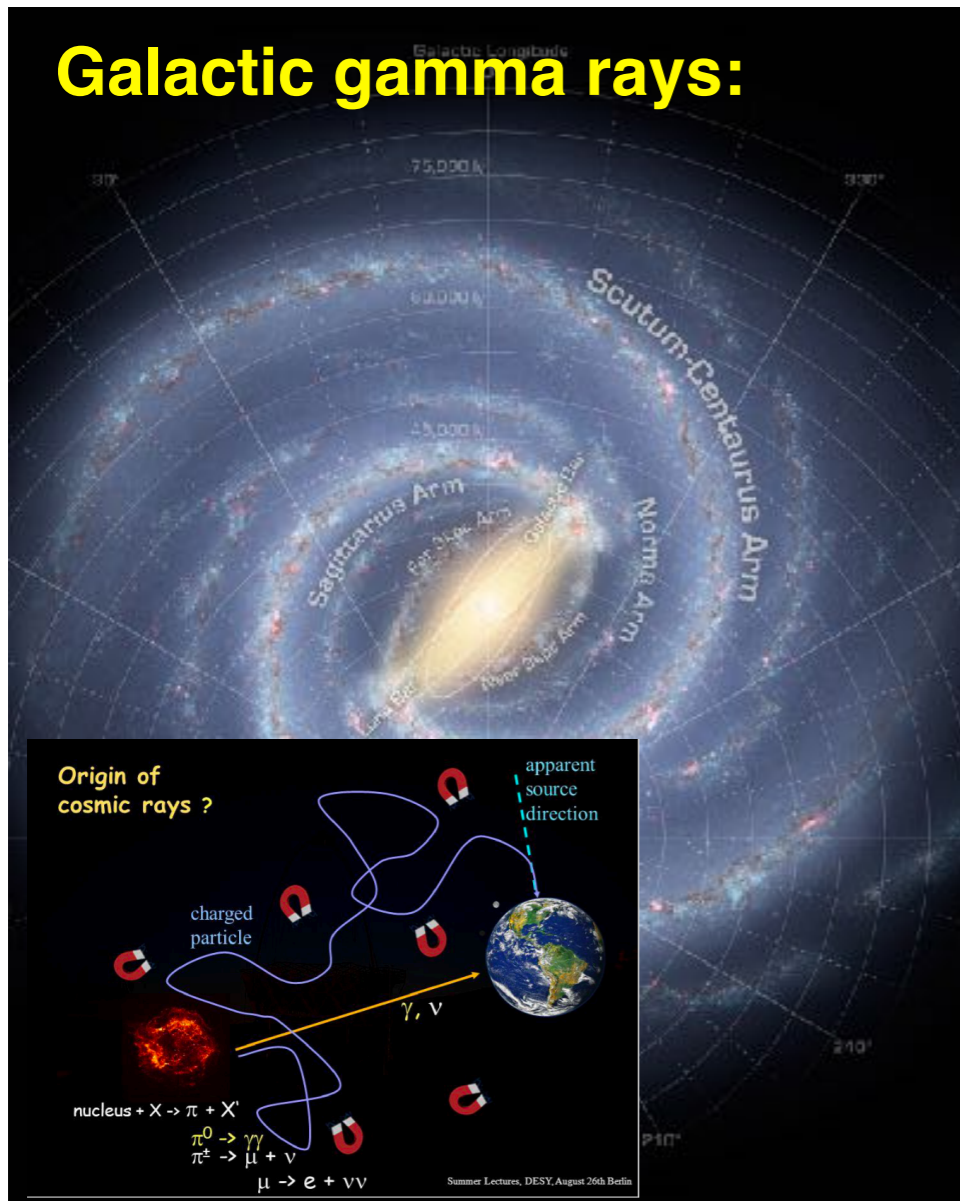


# Broadband modeling of the emission

Nature of the accelerated particles (protons or electrons)?  $E_{\text{max}}$  of accelerated CRs?



# Extragalactic gamma rays

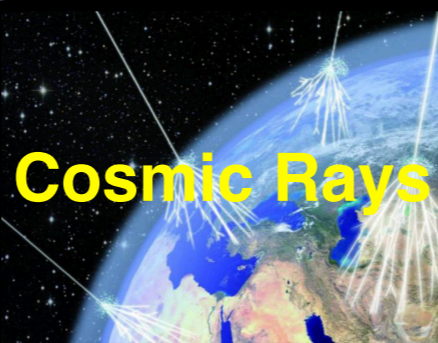


Extragalactic gamma rays absorbed by the Extragalactic Background Light

**Neutrinos are Great!**

- Irrevocable evidence for hadronic acceleration
- Can constrain the origin of extragalactic CRs (unlike gamma rays)

**Cosmic Rays**



**Let's Do Gamma-ray Astronomy!**

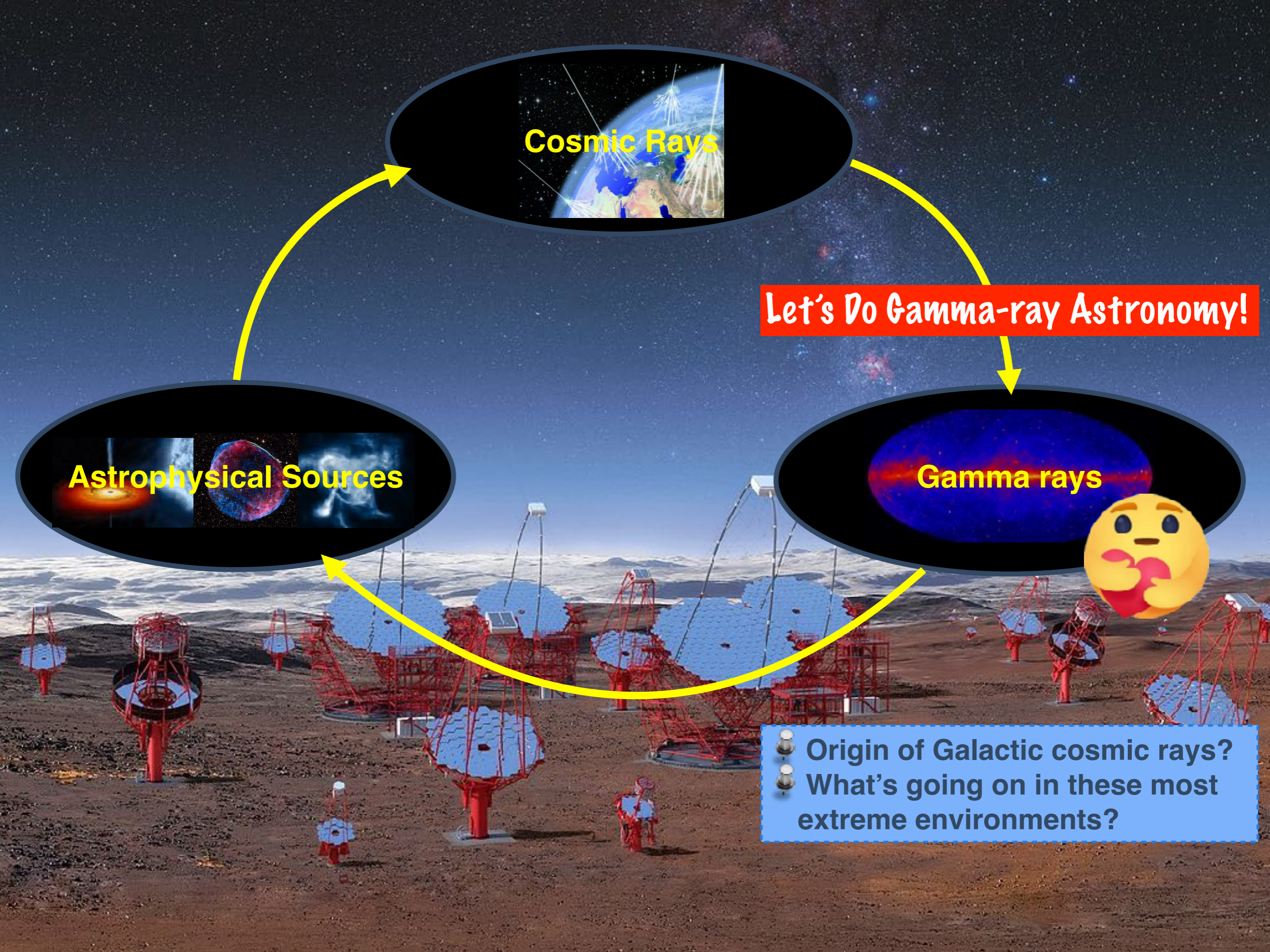
**Gamma rays**



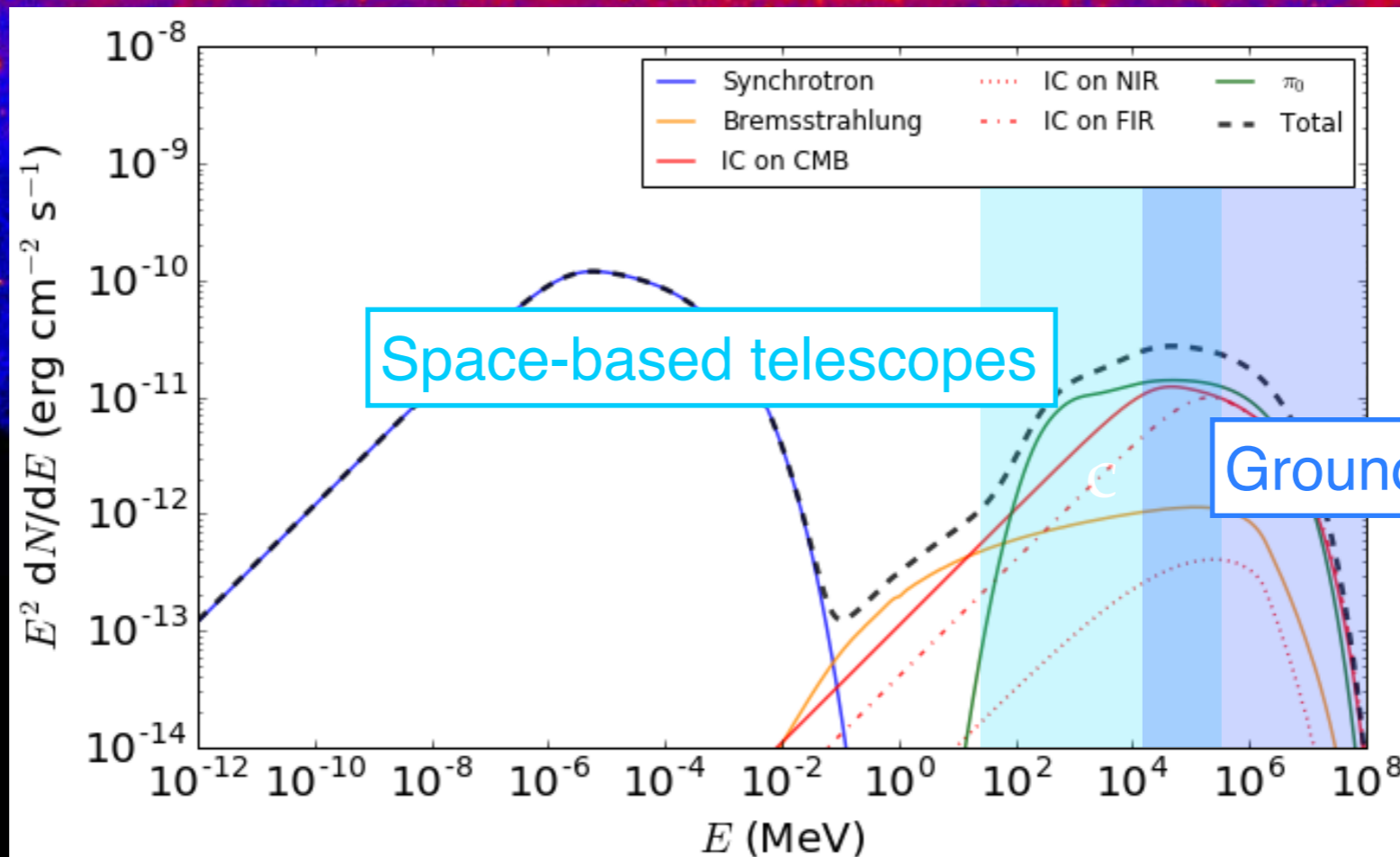
**Astrophysical Sources**



**Origin of Galactic cosmic rays?  
What's going on in these most extreme environments?**



# Gamma-ray Astronomy

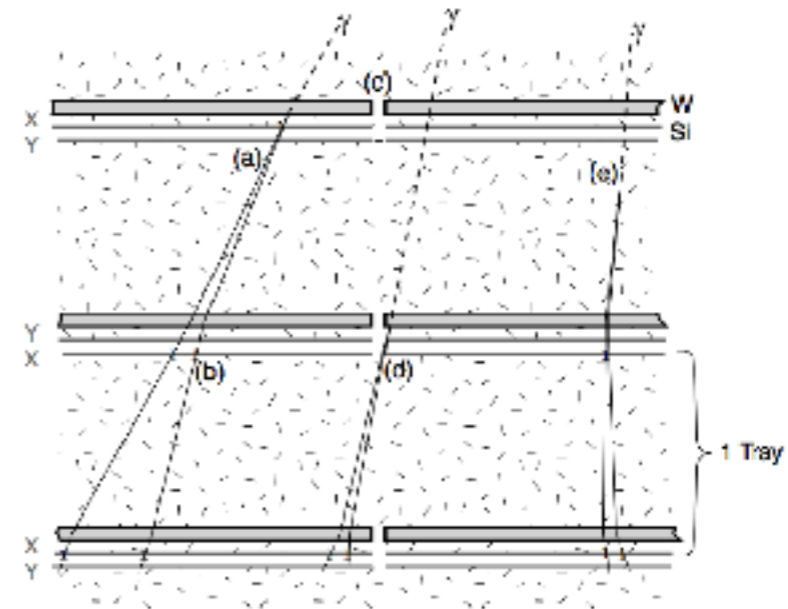
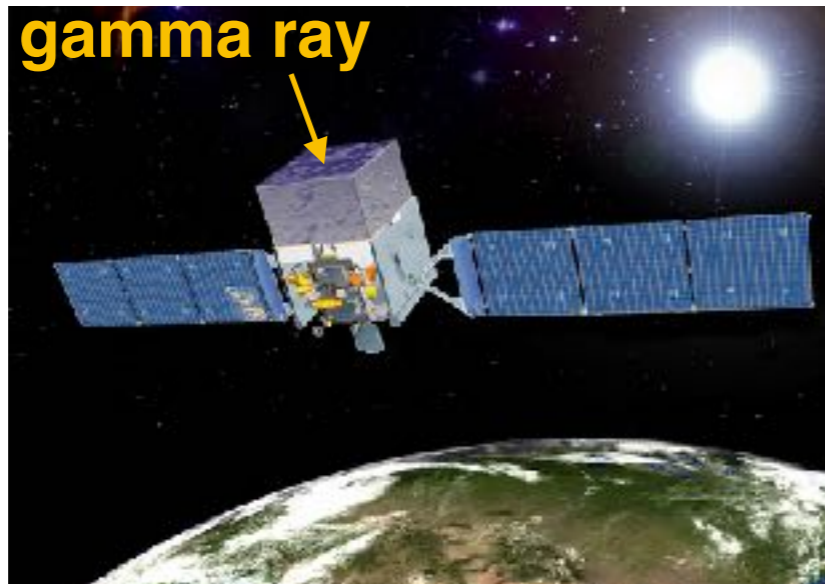


Space-based telescopes

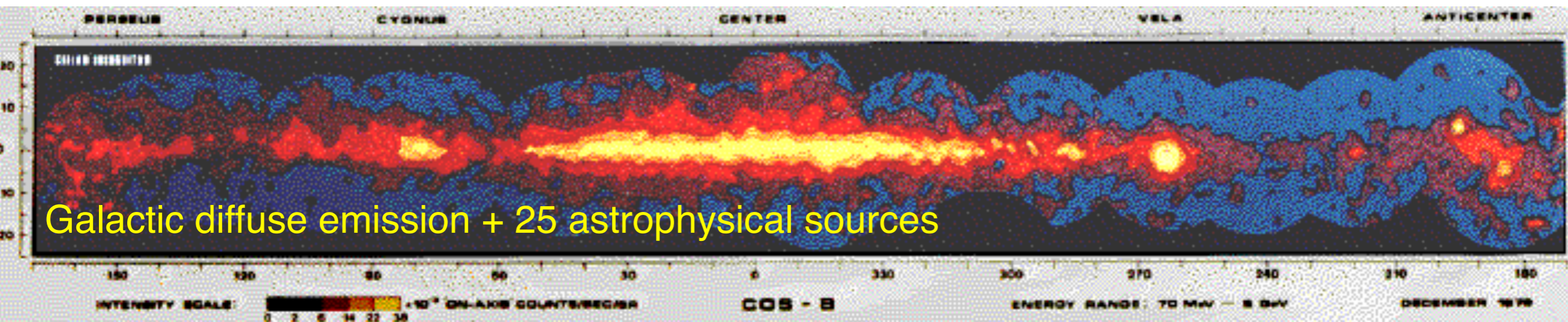
Ground-based telescopes

# How do we detect gamma rays?

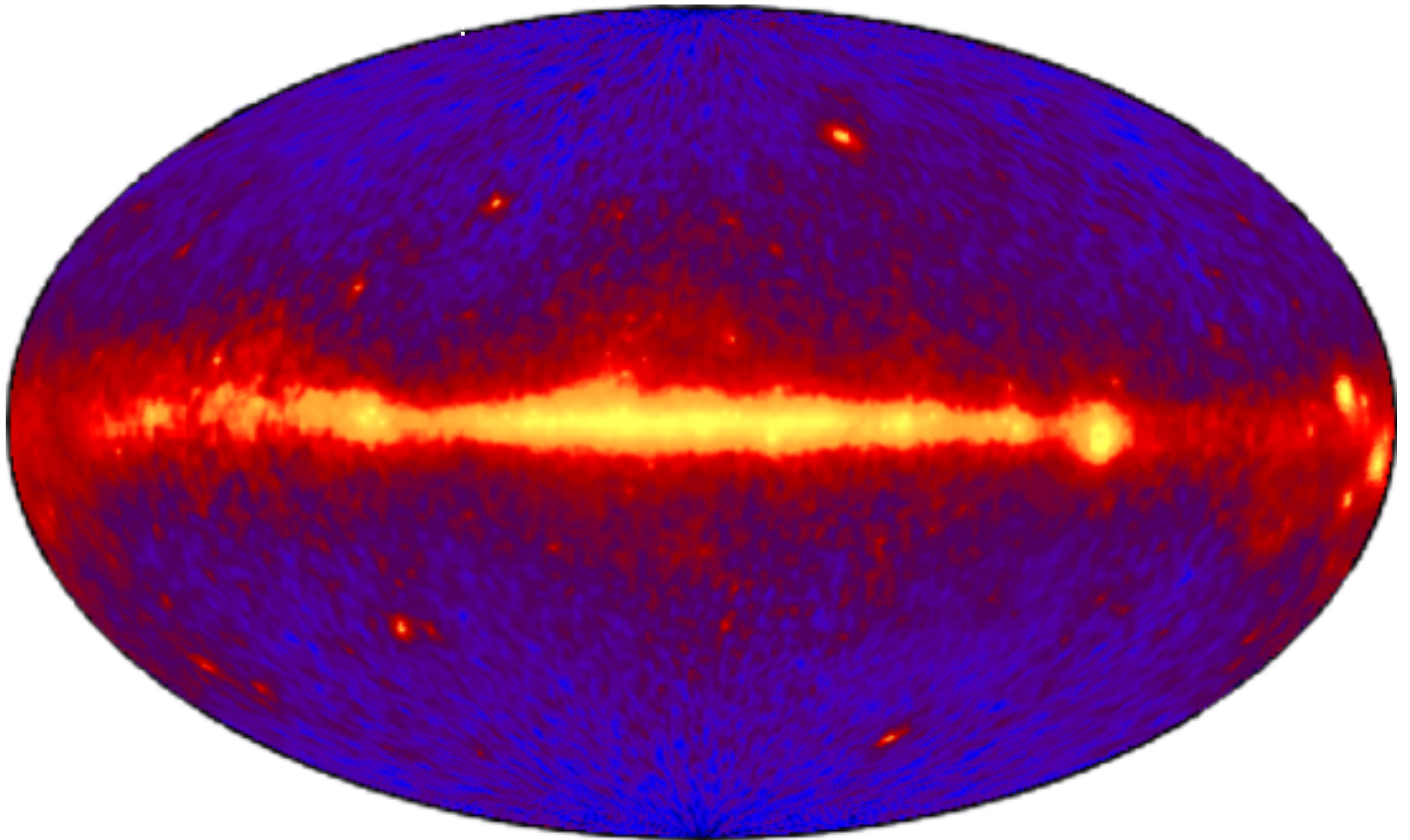
Interaction of gamma rays with the detector (creation of an  $e^-/e^+$  pair):



- 1961 : Explorer 11 (22 photons)
- 1968 : Orbiting Solar Observatory (621 photons)
- 1972 : Small Astronomy Satellite (first part of the sky)
- 1975 : COS-B (first full map of the Galaxy)



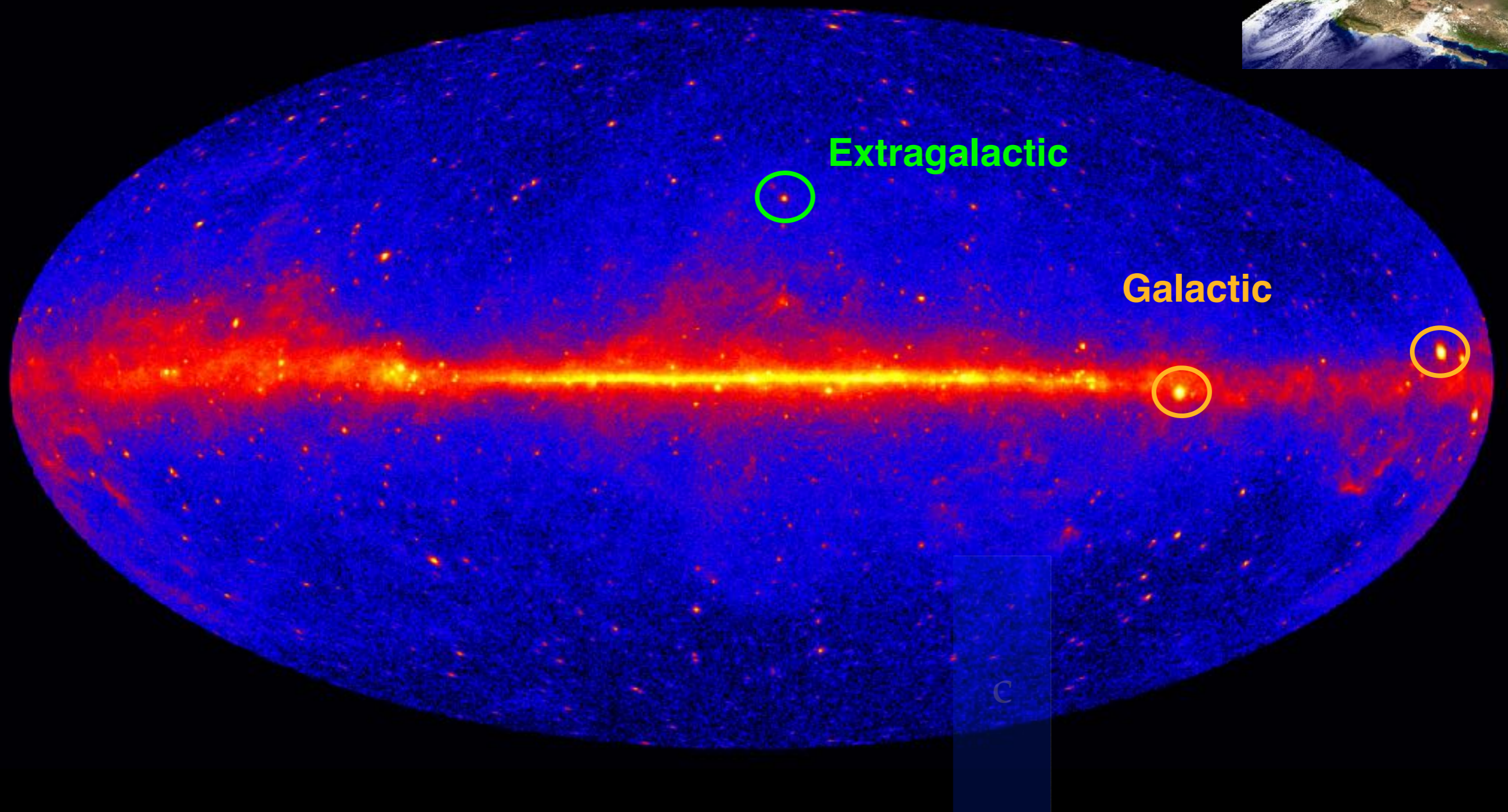
- **1991:** Compton Gamma-ray Observatory
- ➔ Galactic diffuse emission + 271 astrophysical sources





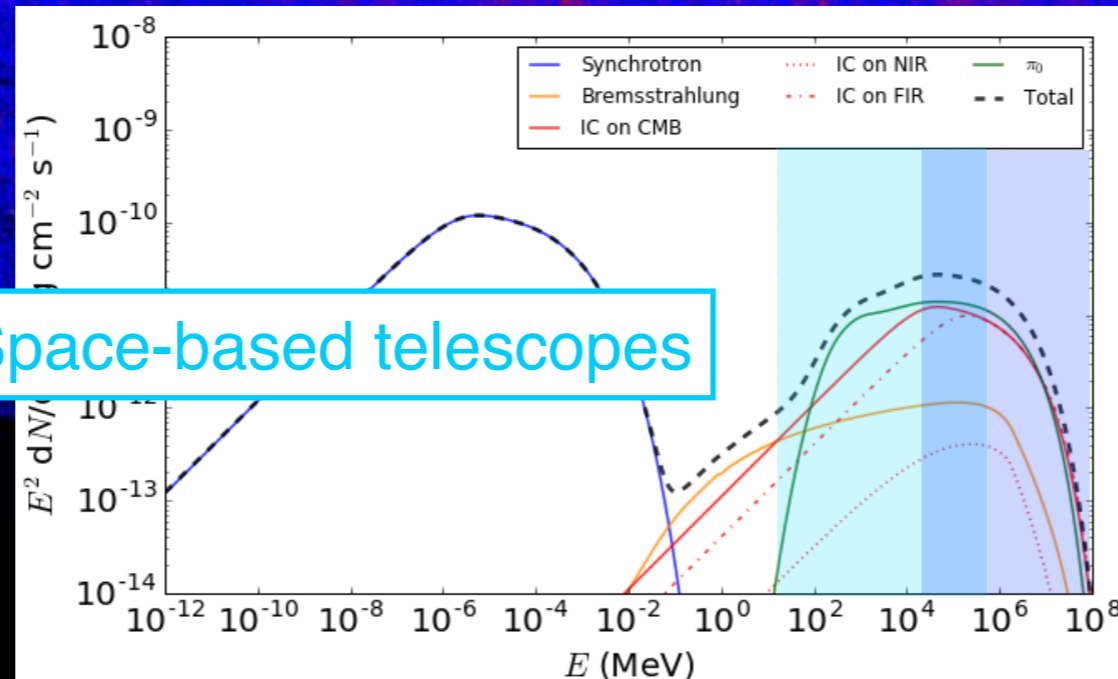
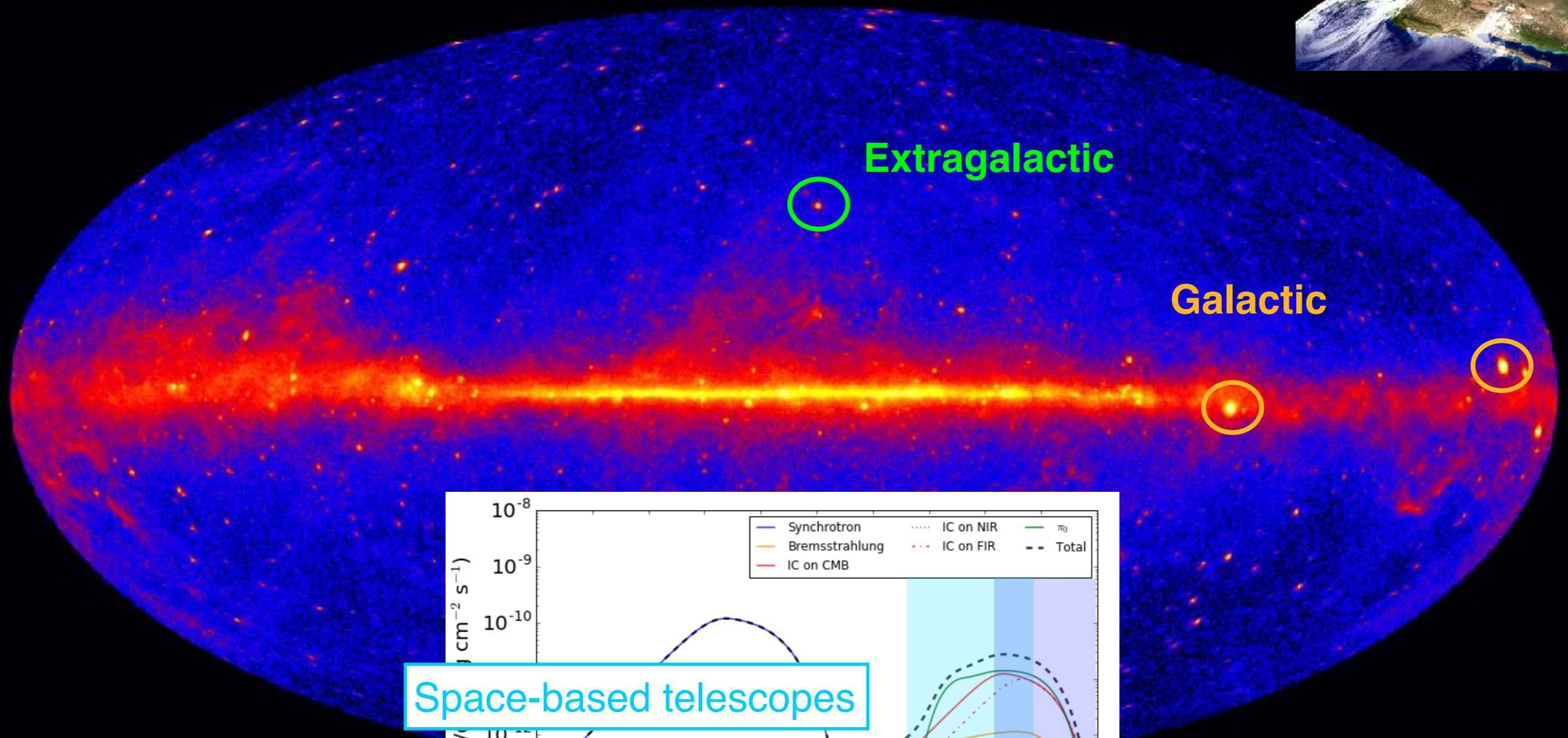
# The GeV Sky Today

- 2008 : Fermi, more than one billion of photons !
- ➔ Galactic diffuse emission + 5065 astrophysical sources



# The GeV Sky Today

- 2008 : Fermi, more than one billion of photons !  
→ Galactic diffuse emission + 5065 astrophysical sources

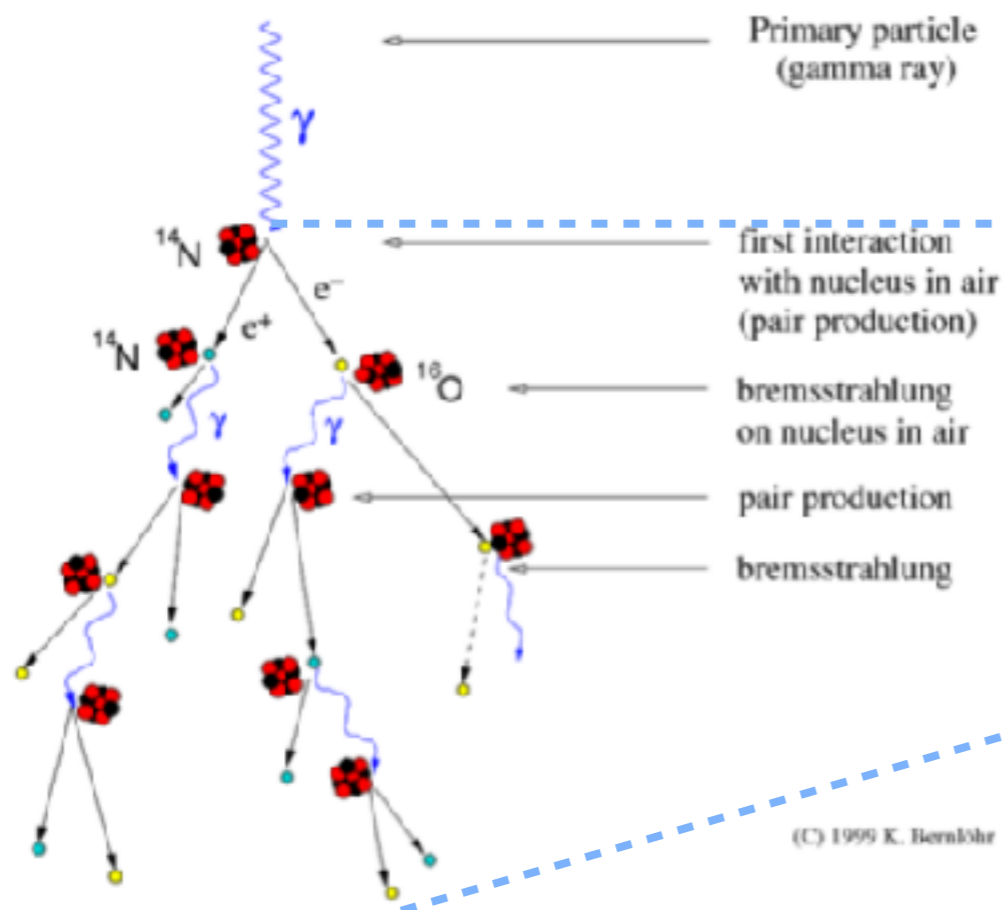


# TeV gamma-ray Astronomy

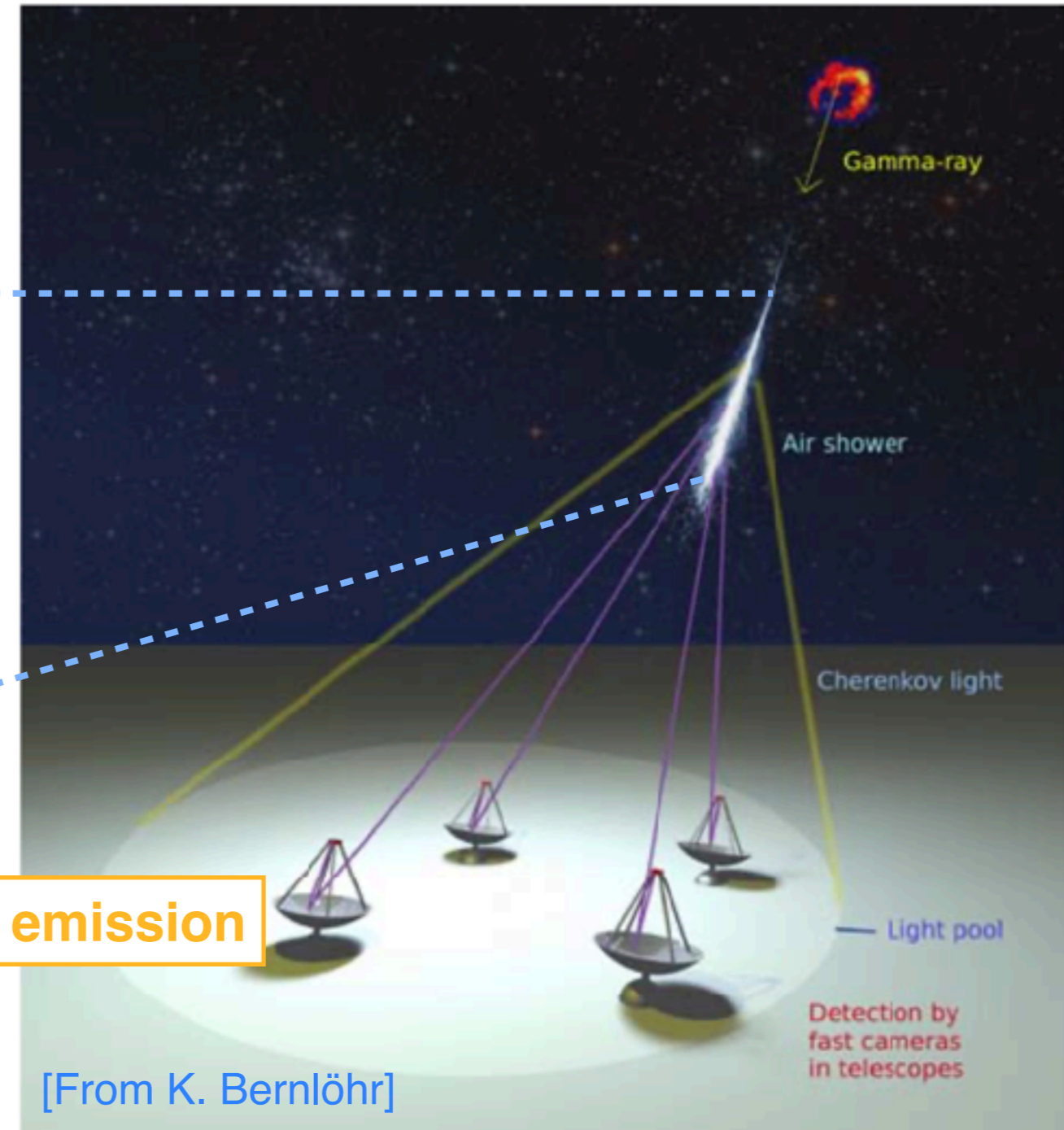
TeV photons more rare than GeV photons: need a larger collection area that can not be boarded on satellite

➔ **The atmosphere is used as a detector**

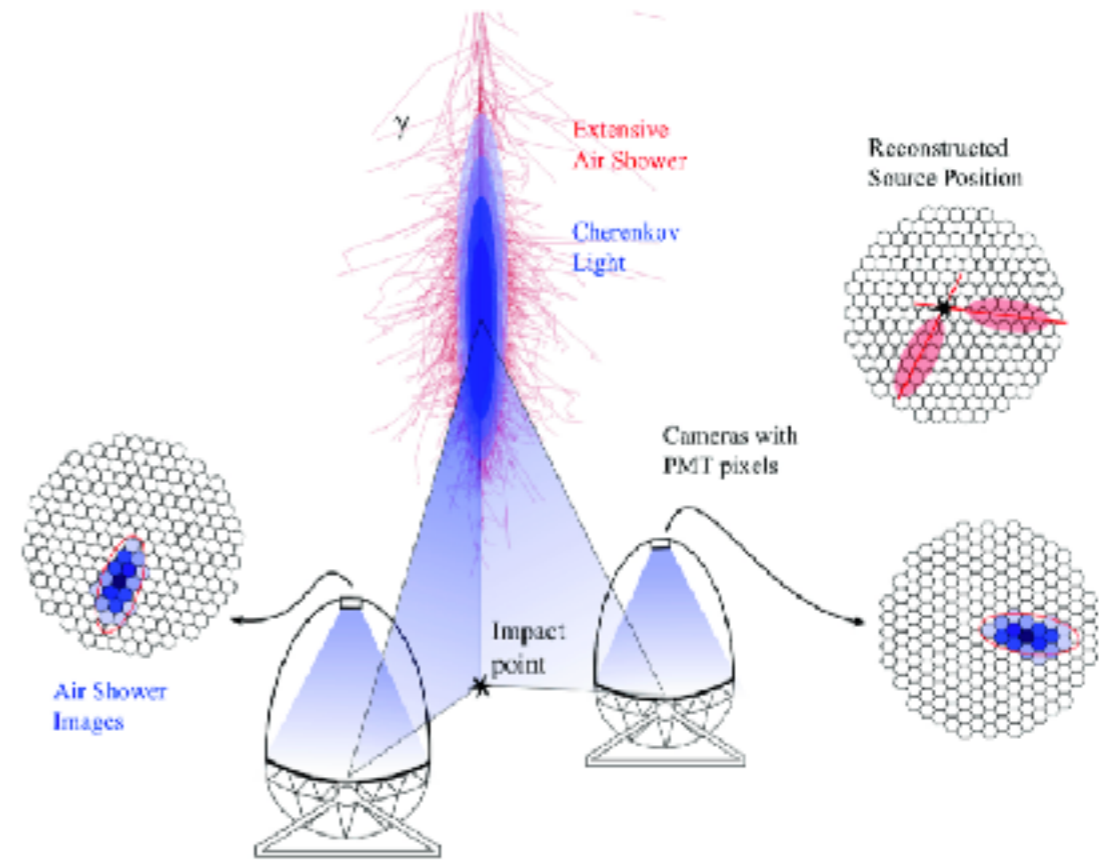
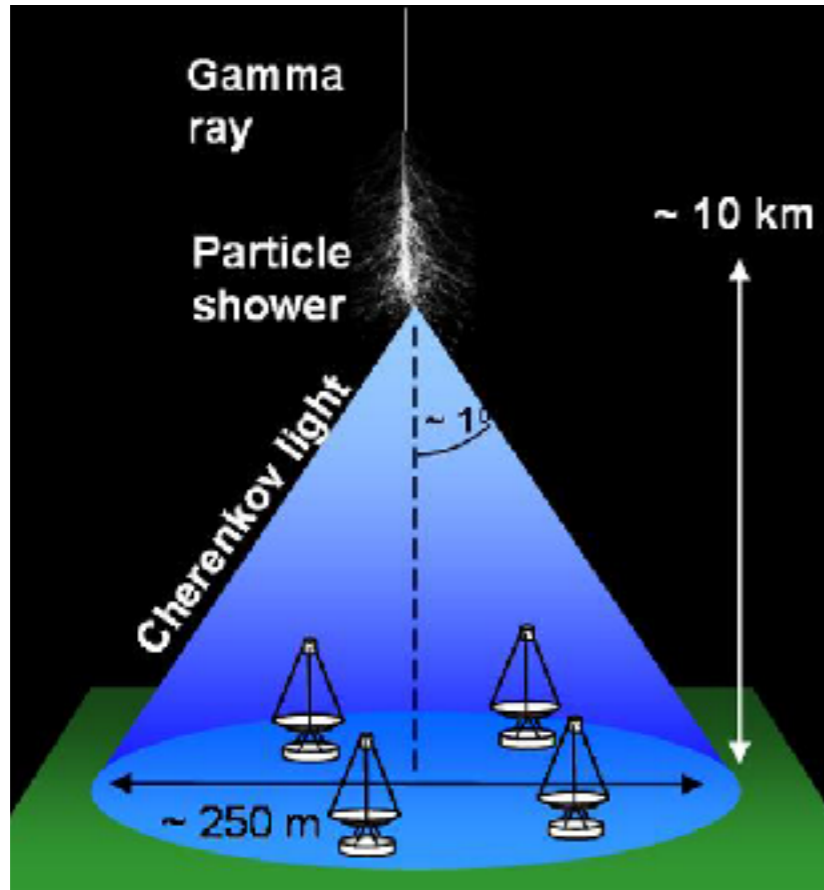
Development of gamma-ray air showers



**Cherenkov light emission**

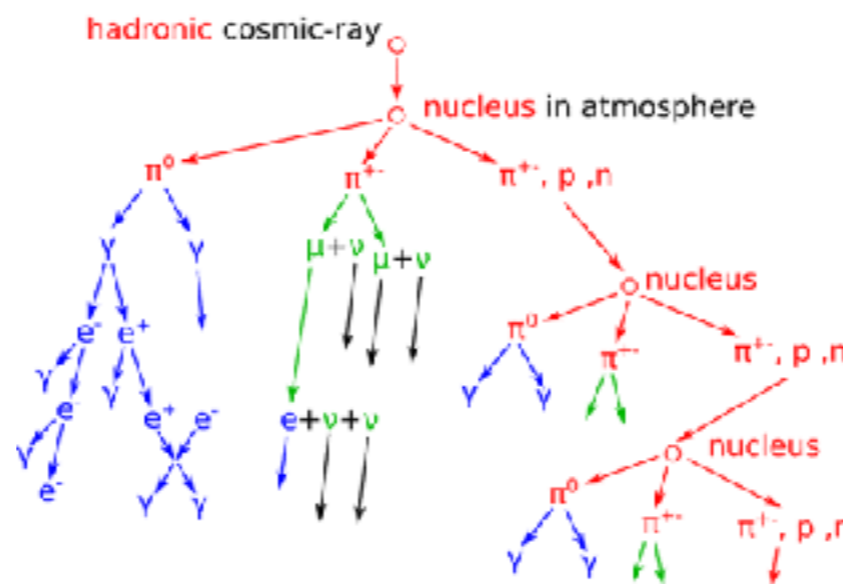


# Stereoscopic reconstruction

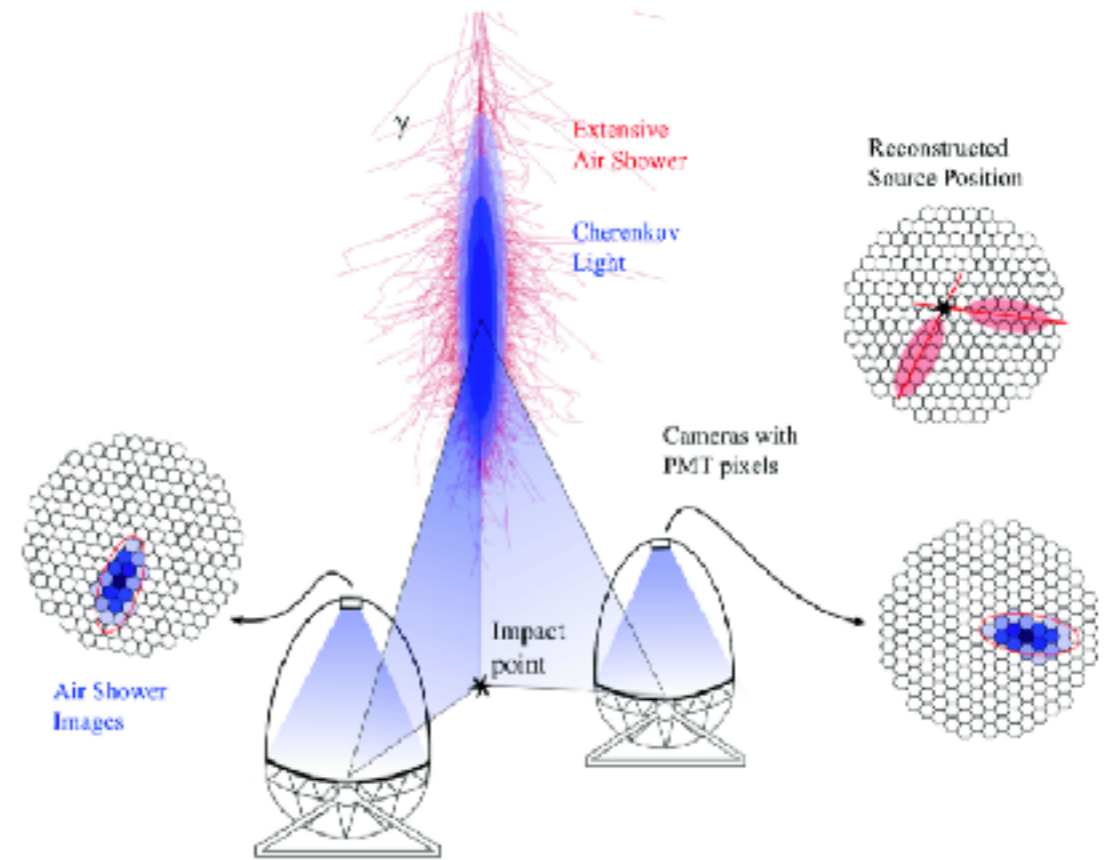
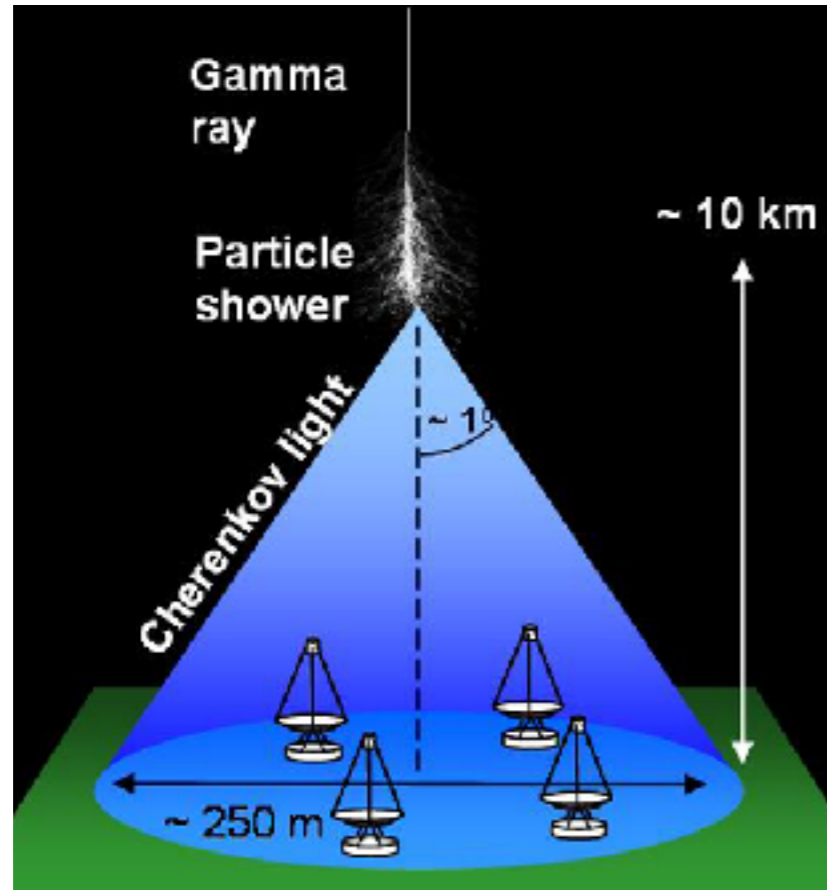


- Reconstructed arrival direction and energy of the primary gamma ray
- Limited field-of-view ( $\sim 3.5\text{--}5^\circ$ ): **pointed observations**

**Cosmic rays are Background here!**



# Stereoscopic reconstruction



- Reconstructed arrival direction and energy of the primary gamma ray
- Limited field-of-view ( $\sim 3.5\text{--}5^\circ$ ): **pointed observations**



# The TeV Sky Today

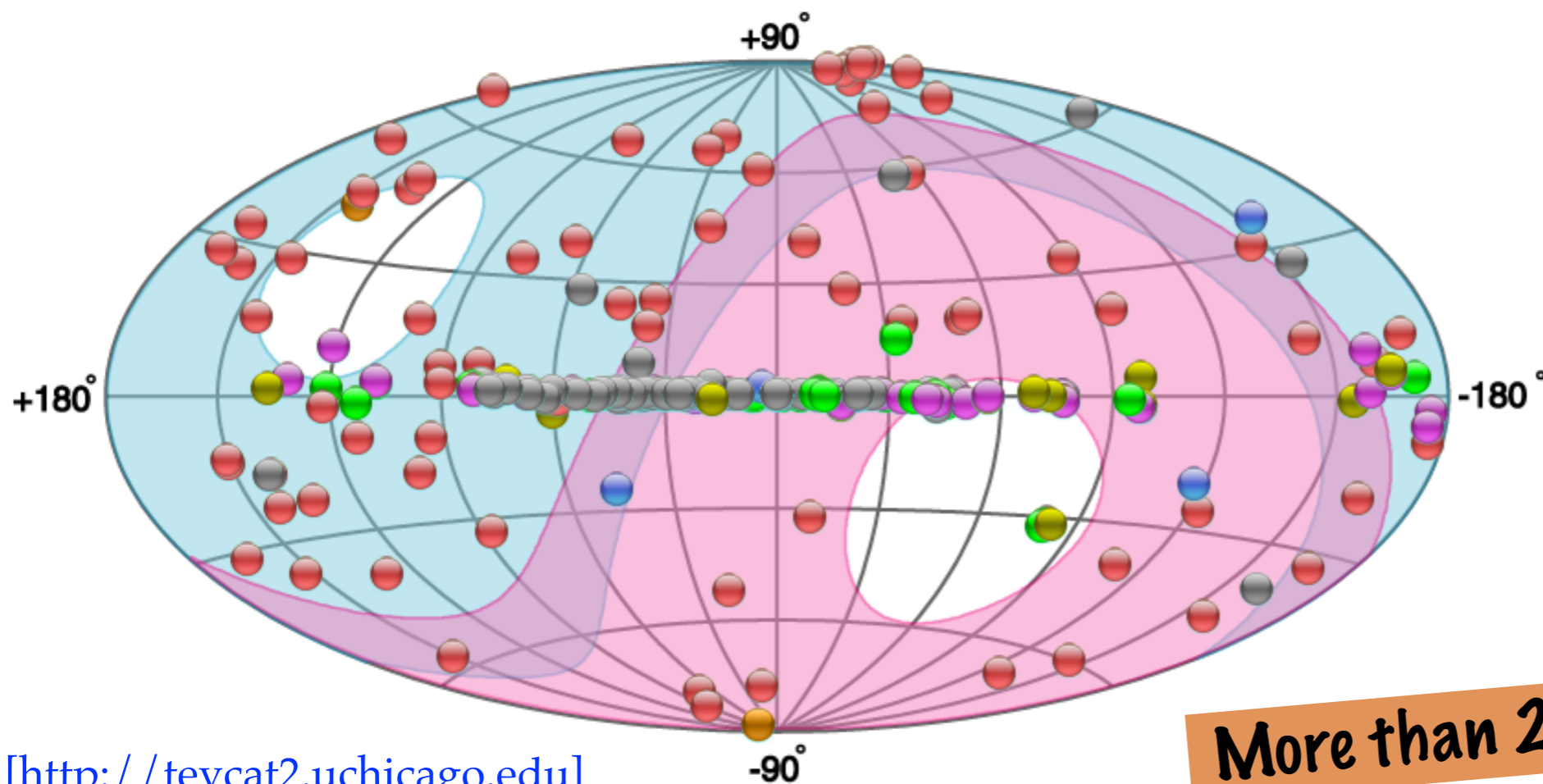
## Northern Hemisphere



## Southern Hemisphere



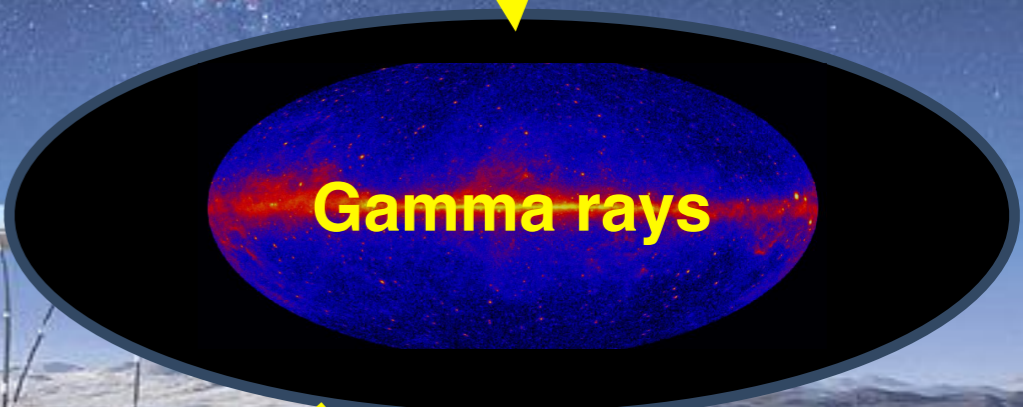
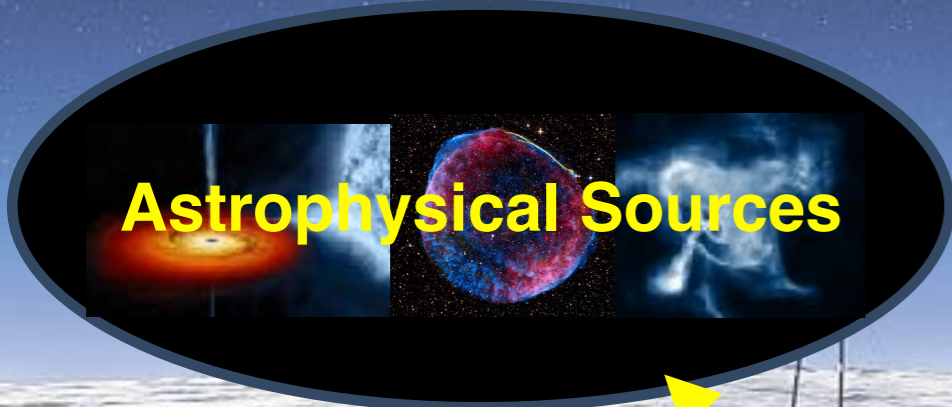
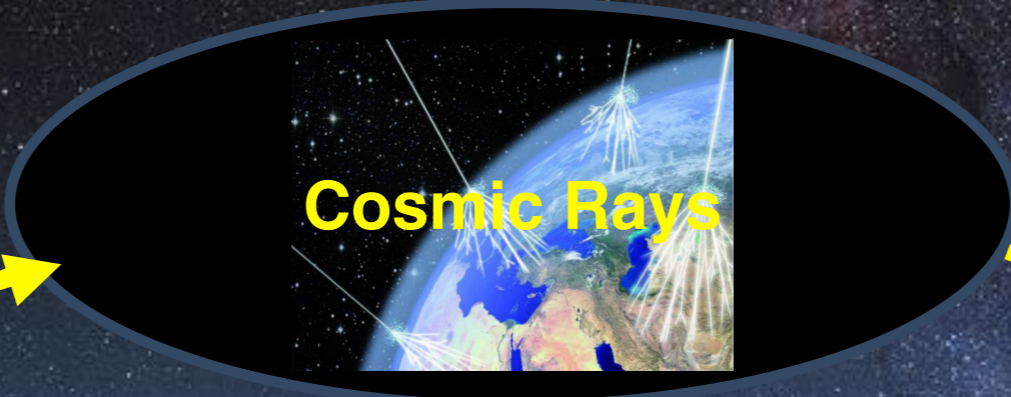
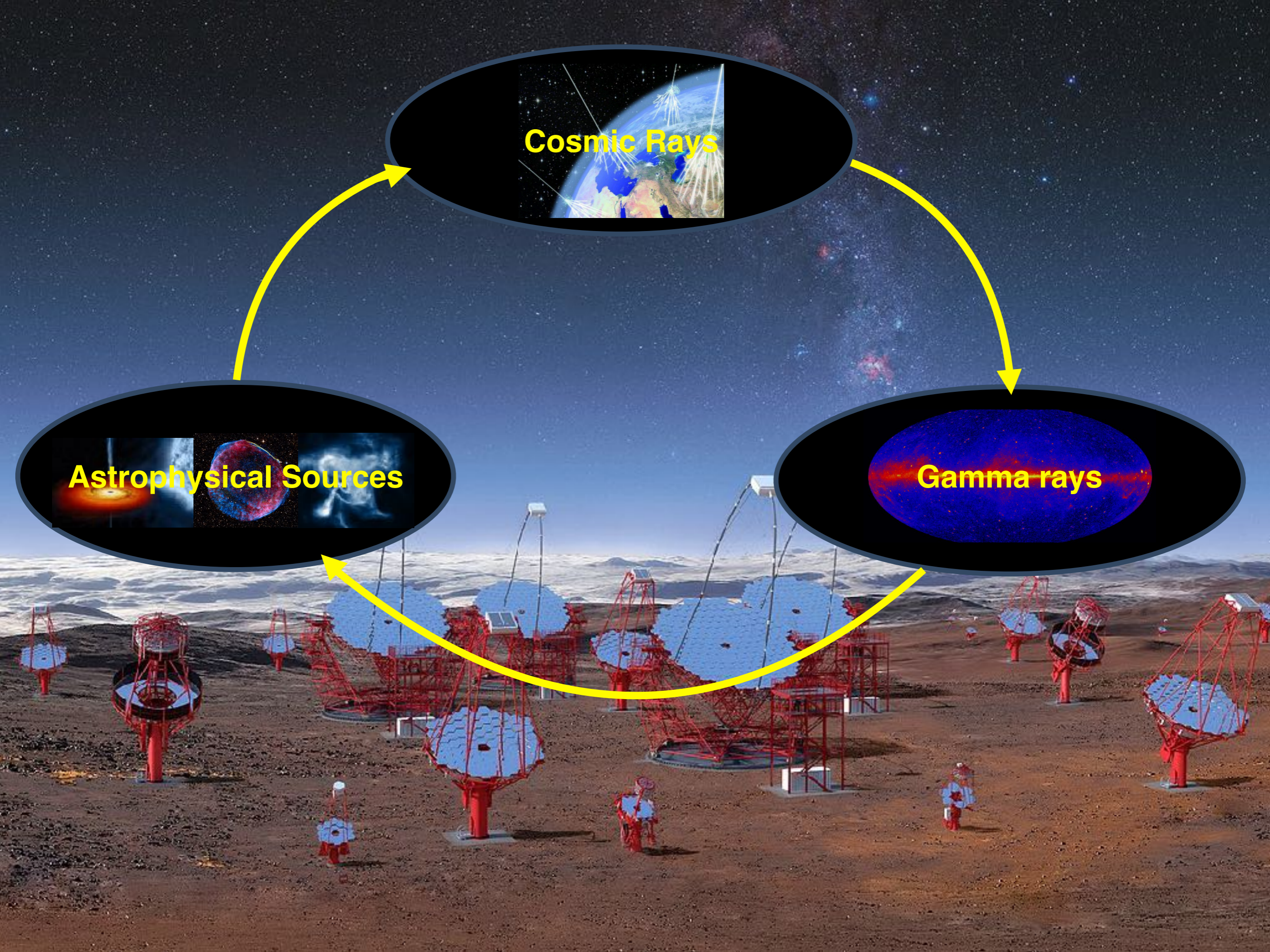
Wassim's Talk  
New Camera on CT5



### Source Types

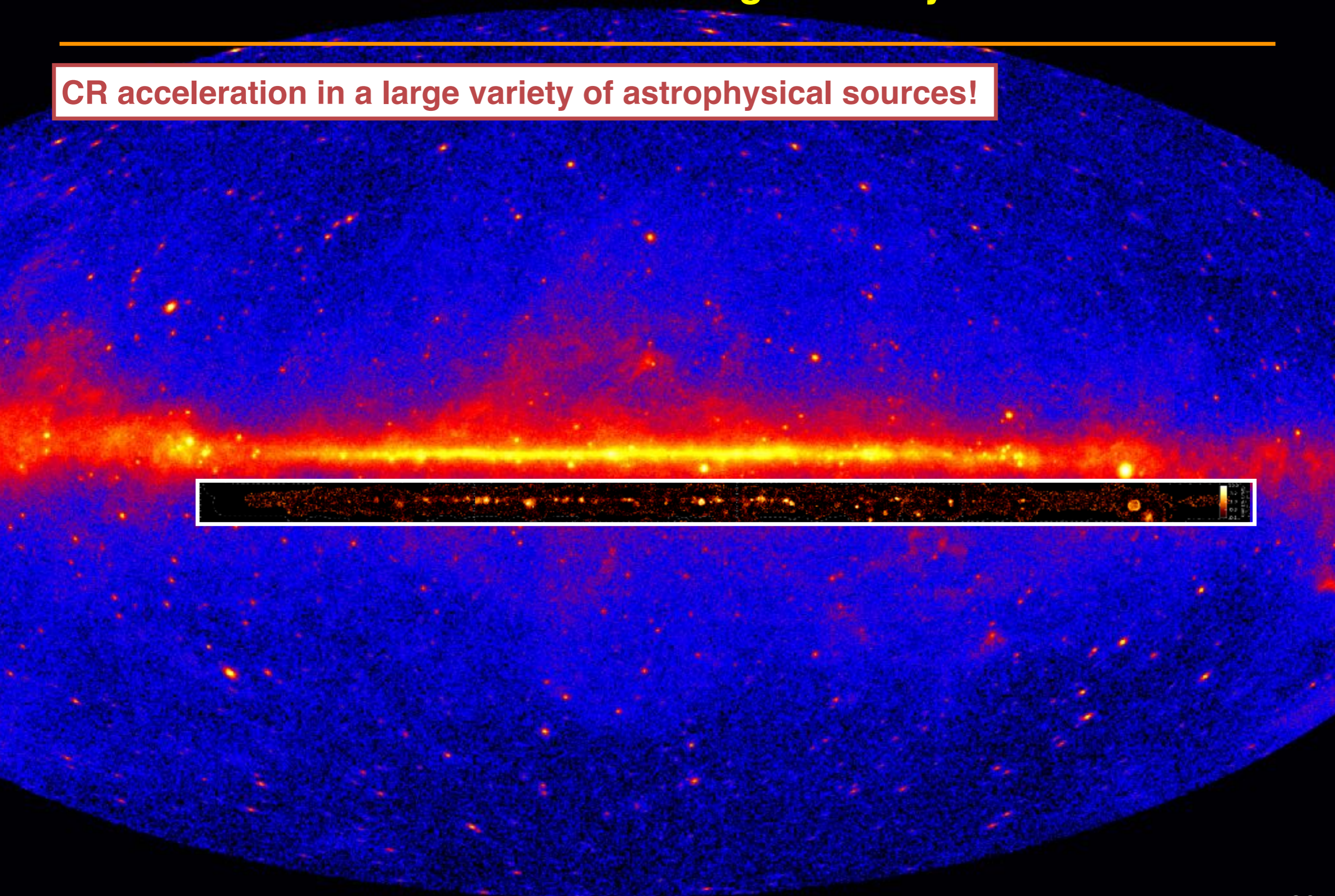
- Extended TeV Halo PWN
- Binary XRB PSR Gamma BIN
- HBL IBL FRI FSRQ  
Blazar LBL AGN  
(unknown type)
- Shell SNR/Molec. Cloud  
Composite SNR  
Superbubble
- Starburst
- DARK UNID Other
- uQuasar Star Forming  
Region Globular Cluster  
Cat. Var. Massive Star  
Cluster BIN PL

**More than 200 sources in total!**



# The Galactic Zoo in gamma rays

CR acceleration in a large variety of astrophysical sources!





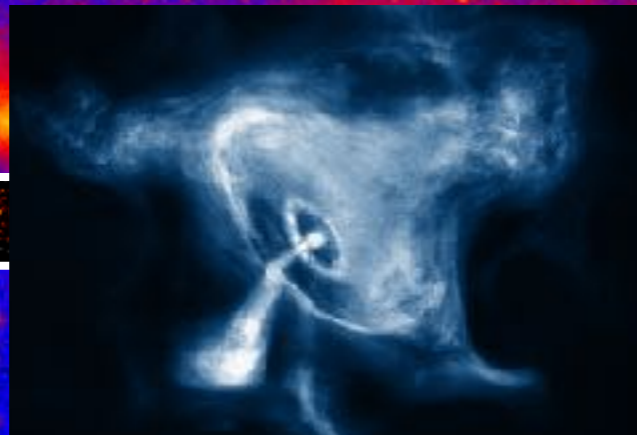
# The Galactic Zoo in gamma rays

CR acceleration in a large variety of astrophysical sources!

Supernova remnants



Pulsar and their nebula



Massive Stellar Clusters



Binary systems:

- Microquasar
- Colliding stellar winds
- ...



+ A large number of unidentified sources!



# The Galactic Zoo in gamma rays

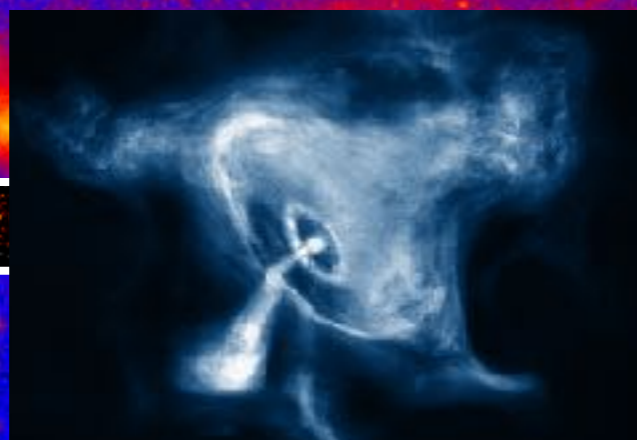
CR acceleration in a large variety of astrophysical sources!

● Maxime's talk:  
Pulsars at TeV energies

Supernova remnants



Pulsar and their nebula



HESS col. Nature 2023

Massive Stellar Clusters



HAWC col. Nature 2018

HESS col. Science 2022

Binary systems:

- Microquasar
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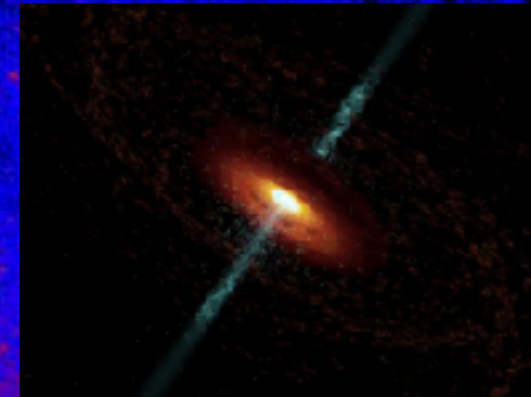


# The Extragalactic Zoo in gamma rays

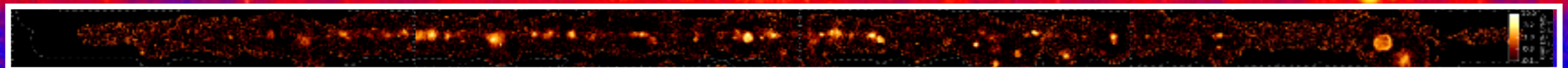
CR acceleration in a large variety of astrophysical sources!

Active Galactic Nuclei

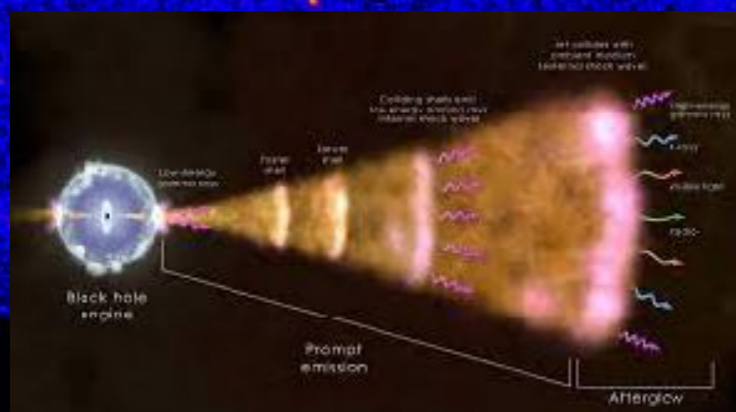
● Samantha's talk:  
VER J0521+211 with Fermi-LAT & VERITAS



VER J0521+211



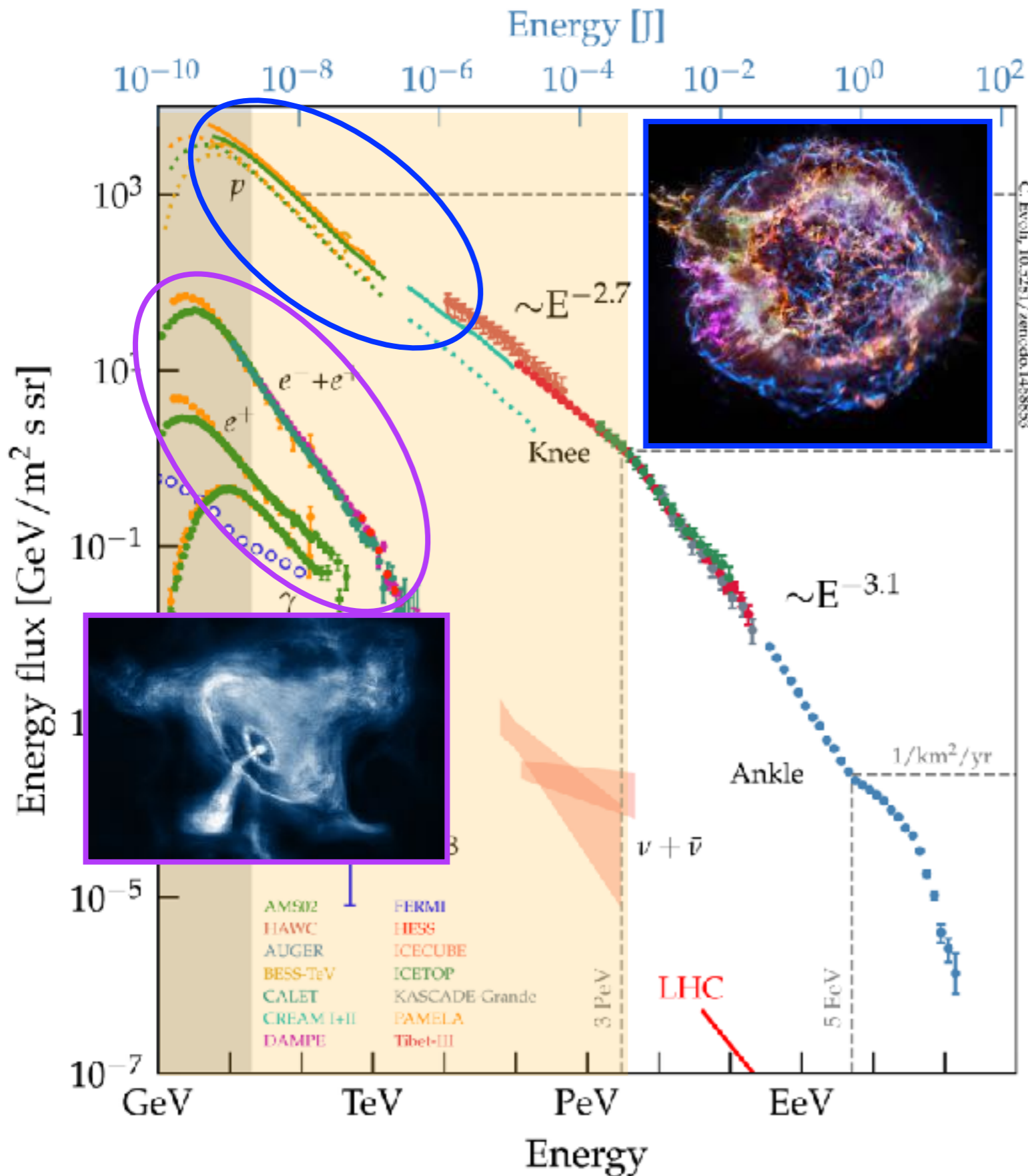
Gamma-ray Bursts



MAGIC col., Nature, 2019  
HESS col., Nature, 2019

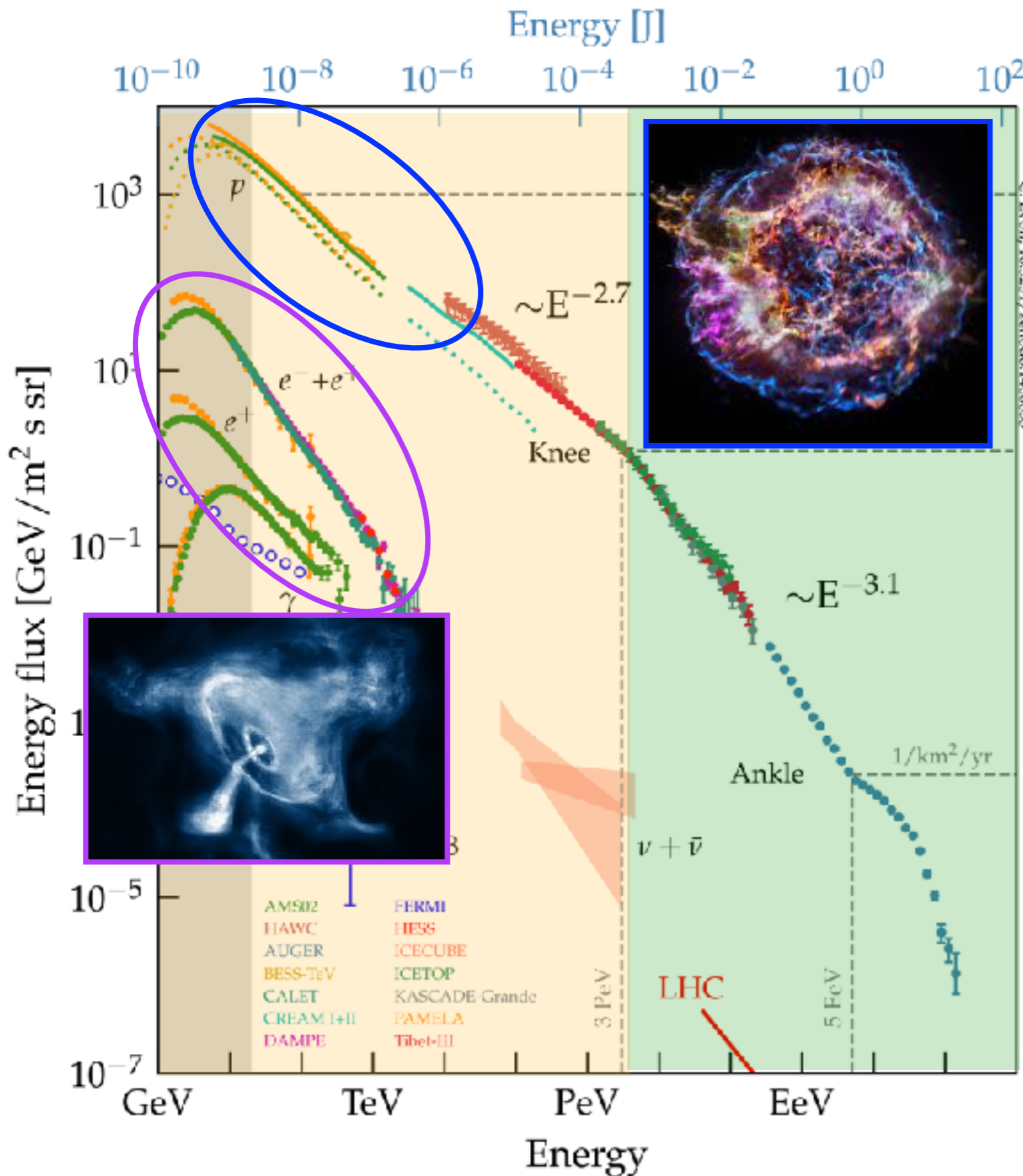
Emission processes often not clearly understood...

# And the Origin of Cosmic rays?



- Electrons can be accelerated in pulsar wind nebulae up to PeV energies
  - Accelerated protons seen in our favorite candidates (Supernova remnants) but  $E_{\text{max}} \ll \text{PeV}$
- ➔ Ancient PeVatrons?
- ➔ Other sources? Massive Stellar Clusters?

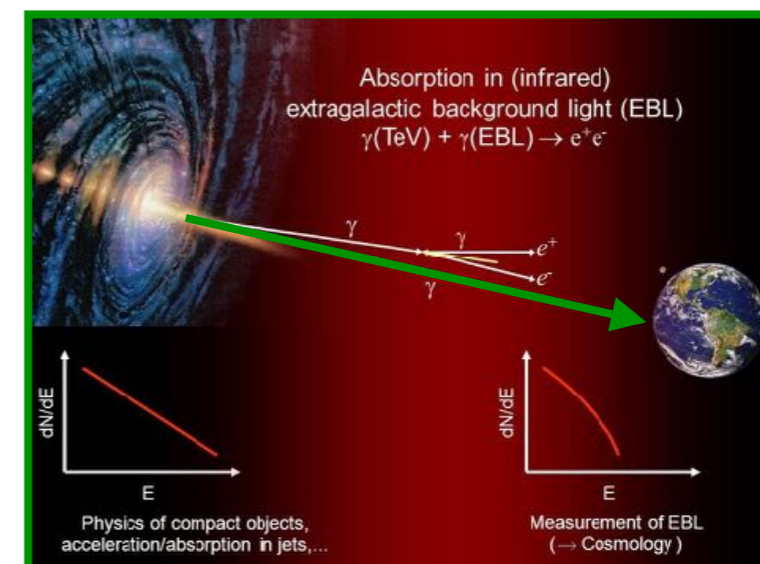
# And the Origin of Cosmic rays?



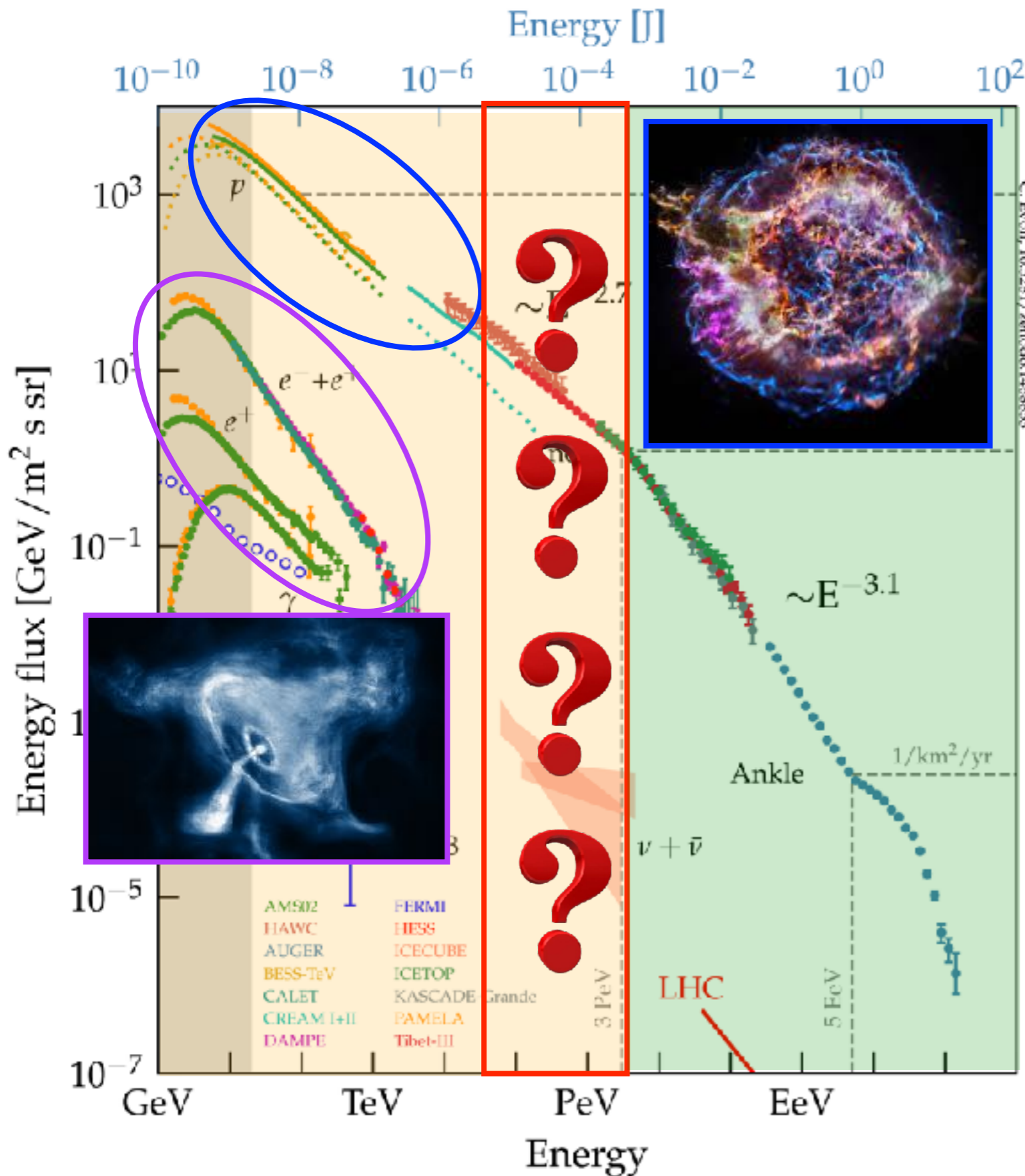
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Need some help from neutrinos!

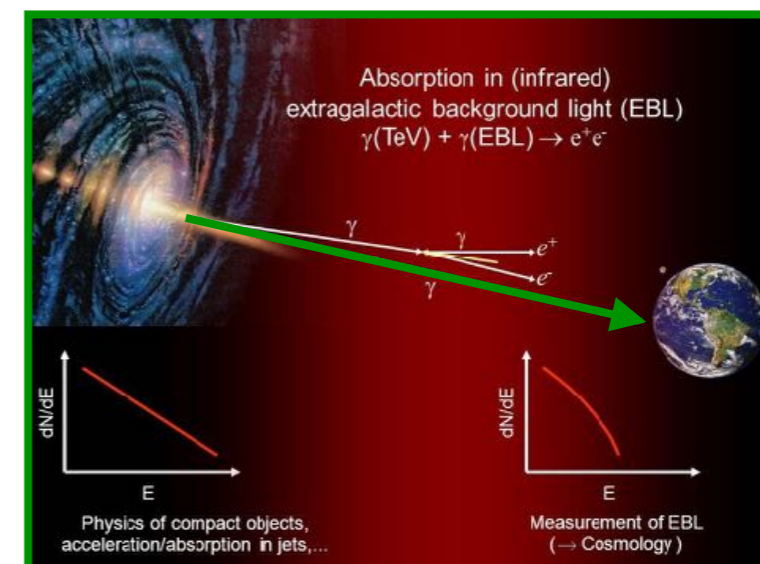


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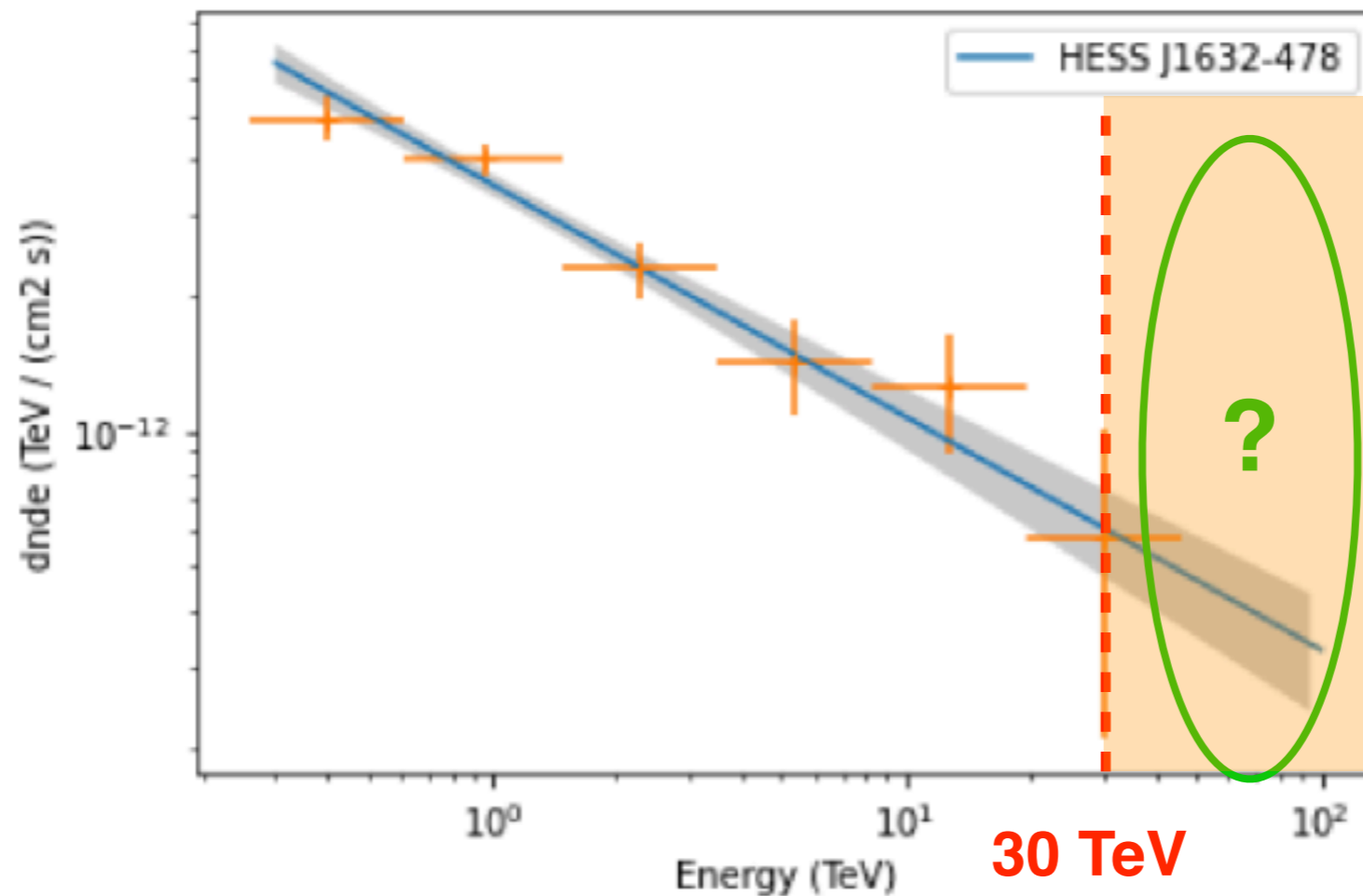
Need some help from neutrinos!



# Lack of Statistics!

If  $E_p = 3 \times 10^{15}$  eV needed,  $E_{\text{g-ray}} \sim 3 \times 10^{14}$  eV (= 300 TeV)

For many TeV sources, we do not know what happens at the highest energy:



Does the spectrum has a cutoff ( $E_{\text{max}}$ ) or continue up to 300 TeV?  
➔ At the sensitivity limit with H.E.S.S.!

# The Cherenkov Telescope Array (CTA)



**Canary Islands (Northern Hemisphere)  
13 telescopes**

**Chile (Southern Hemisphere)  
> 50 telescopes**

● **Leo's talk:**  
Correction for mispointing with LSTs

3 different sizes:

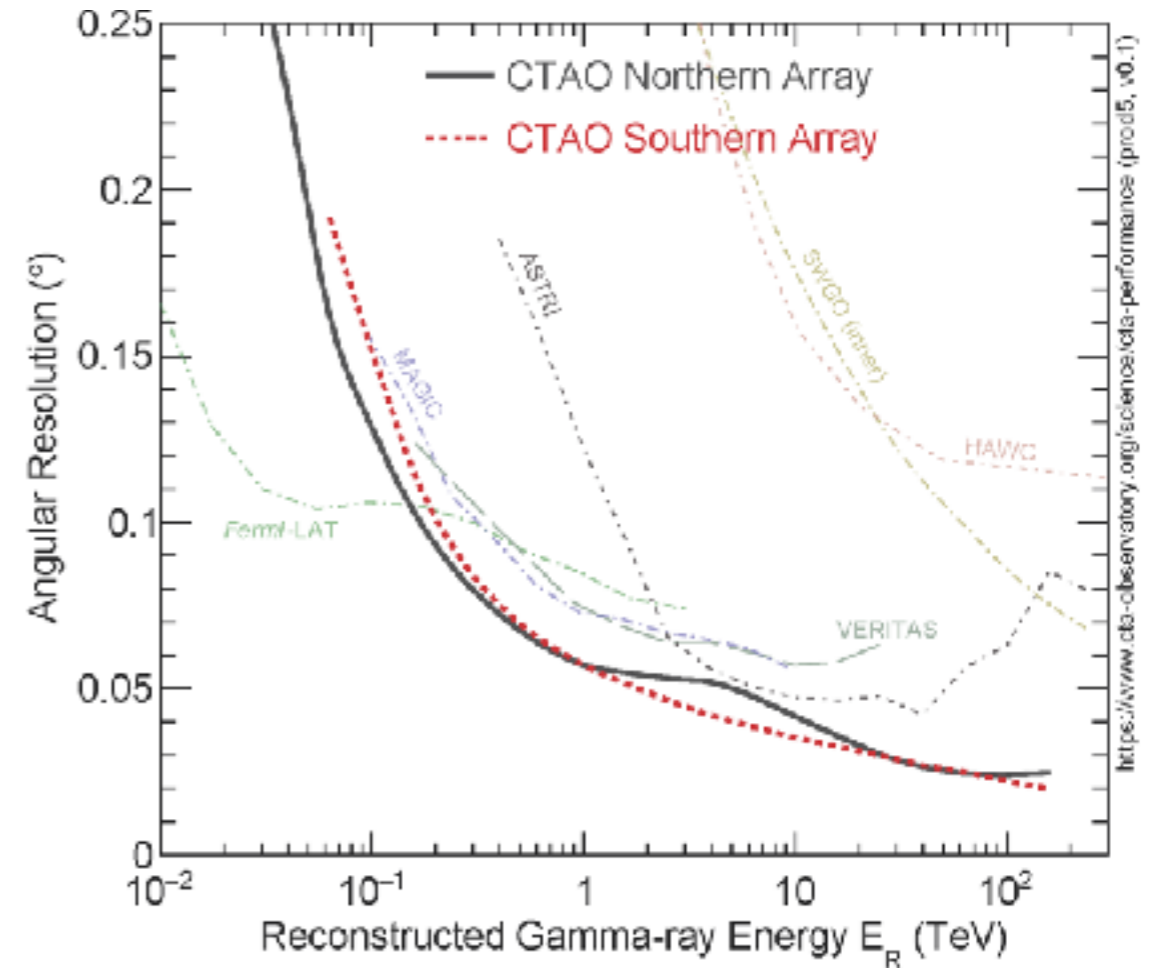
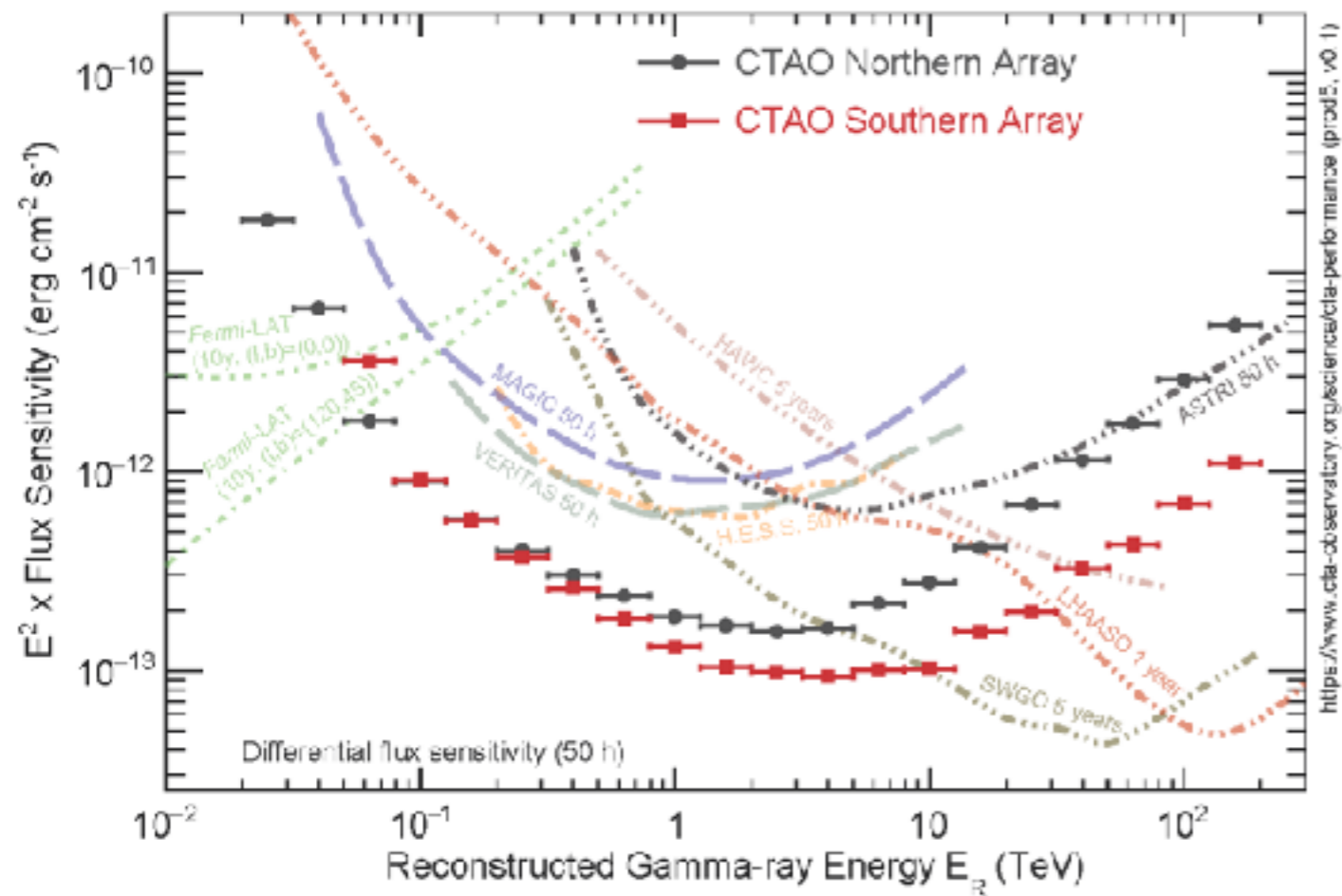
- ➔ Large Size Telescopes (LST)
- ➔ Medium Size Telescopes (MST)
- ➔ Small Size Telescopes (SST)

(currently under construction)





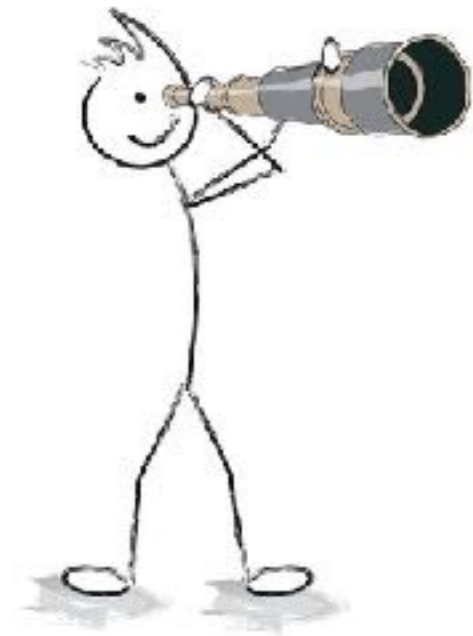
# Better sensitivity and angular resolution



→ Detection of fainter signals

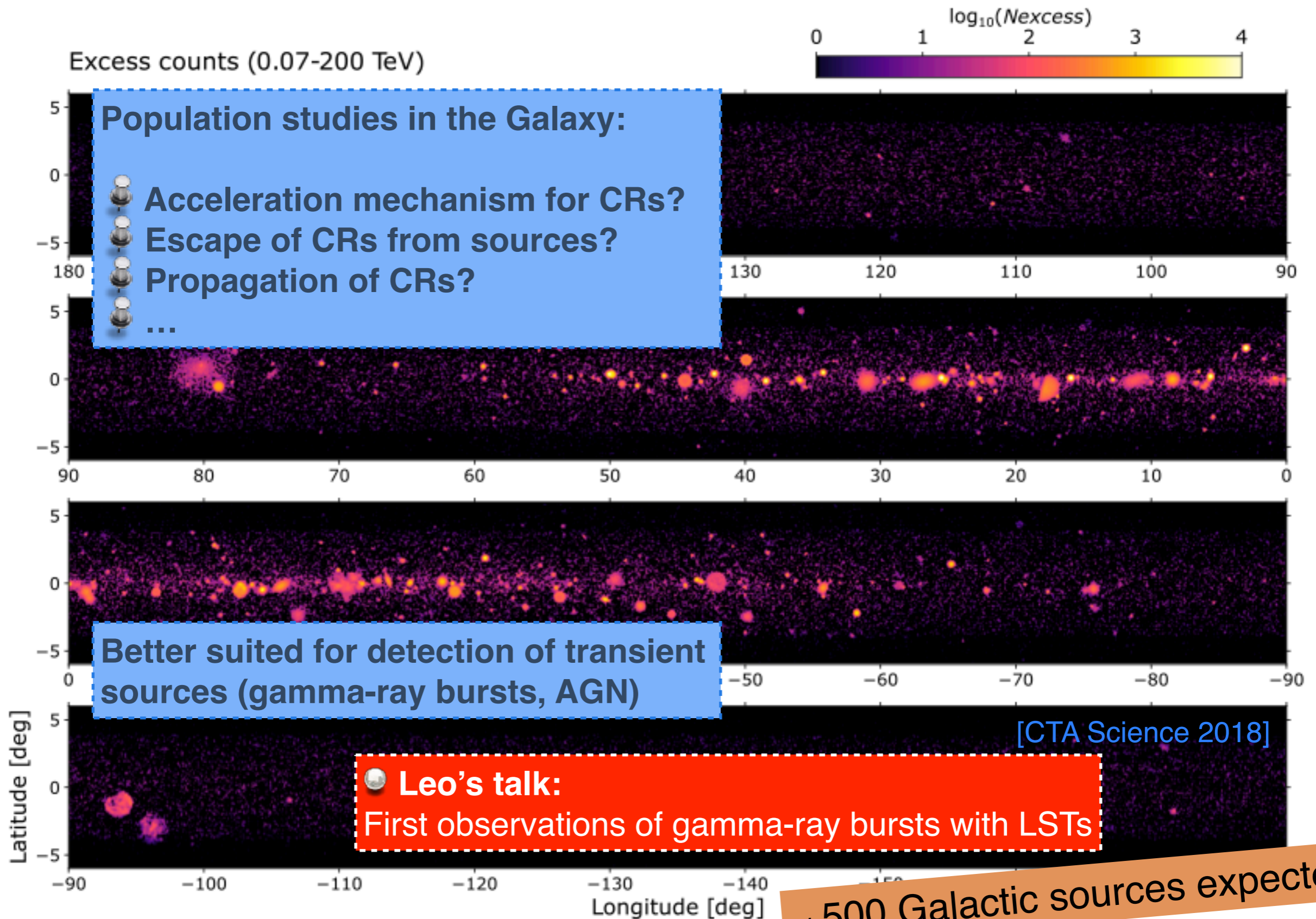


... and farther



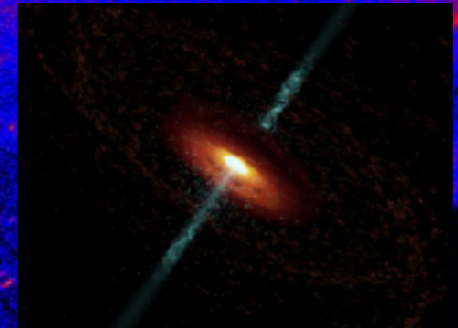
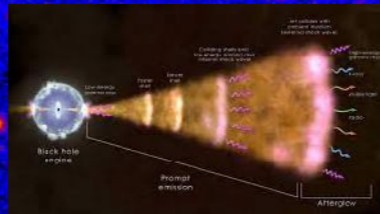
= Detection of more sources!

# Simulation of the CTA Galactic Plane Survey



# Conclusions

- Cosmic rays are relativistic charged particles detected at Earth  
→ **We want to address the century-old question of their origin**
- **We use gamma rays (or neutrinos) to pinpoint their acceleration site**
- Gamma-ray astronomy allows to probe:
  - **acceleration, escape and propagation** of cosmic rays
  - a better understanding of the most extreme phenomena



- CTA should be able to answer the question of the origin of cosmic rays and will likely discover unexpected things!

**FUTURE IS BRIGHT FOR GALACTIC AND EXTRAGALACTIC SCIENCE WITH GAMMA RAYS!**



# Backup Slides

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# Detection of gamma rays

HE: 0.3—100 GeV

Extensive Air Shower experiments (EAS)

UHE: 0.1—30 PeV

Fermi-LAT



Electro-Magnetic Cascade



WCDA,  
KM2A



Cherenkov Light



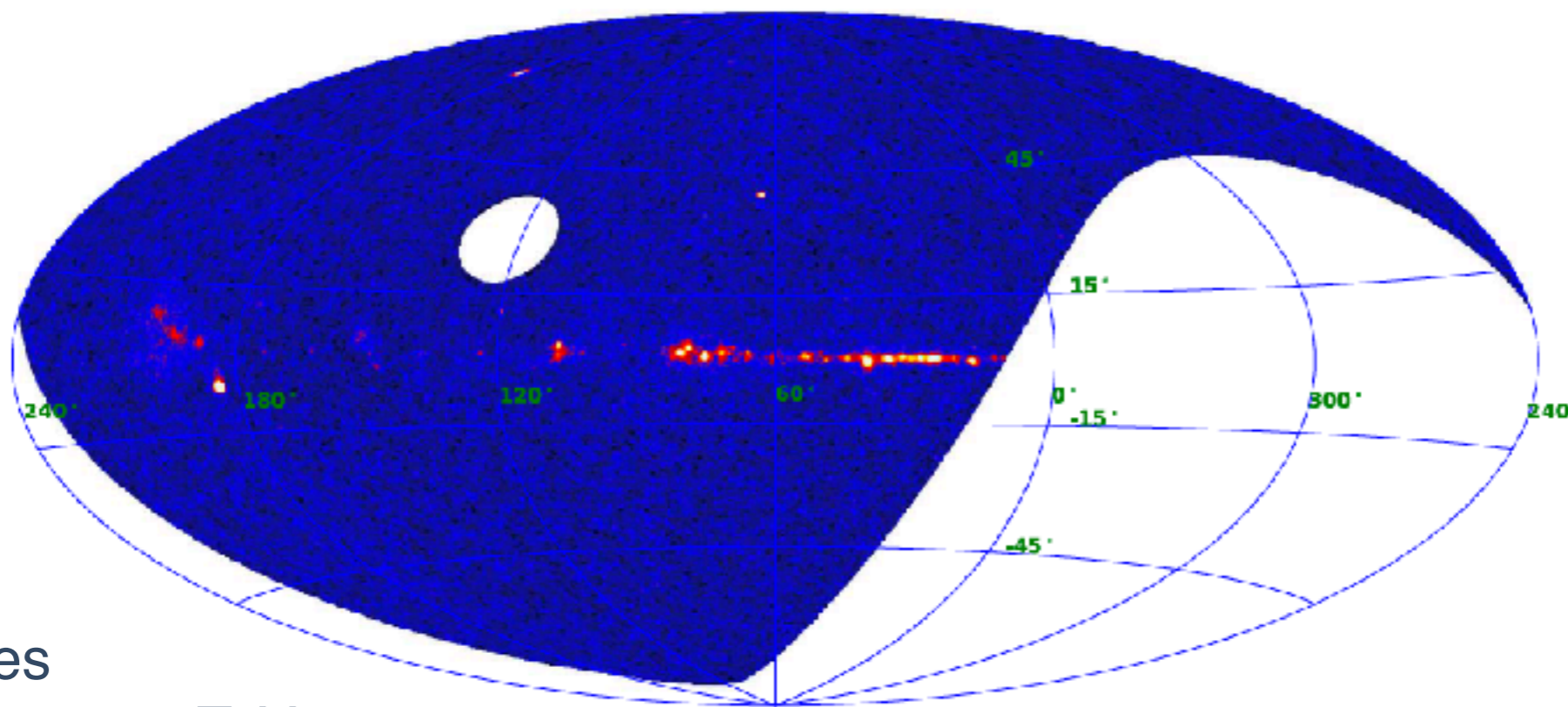
VHE: 0.1—100 TeV

Imaging Atmospheric Cherenkov Telescopes (IACTs)

# The 1st LHAASO catalog: a breakthrough

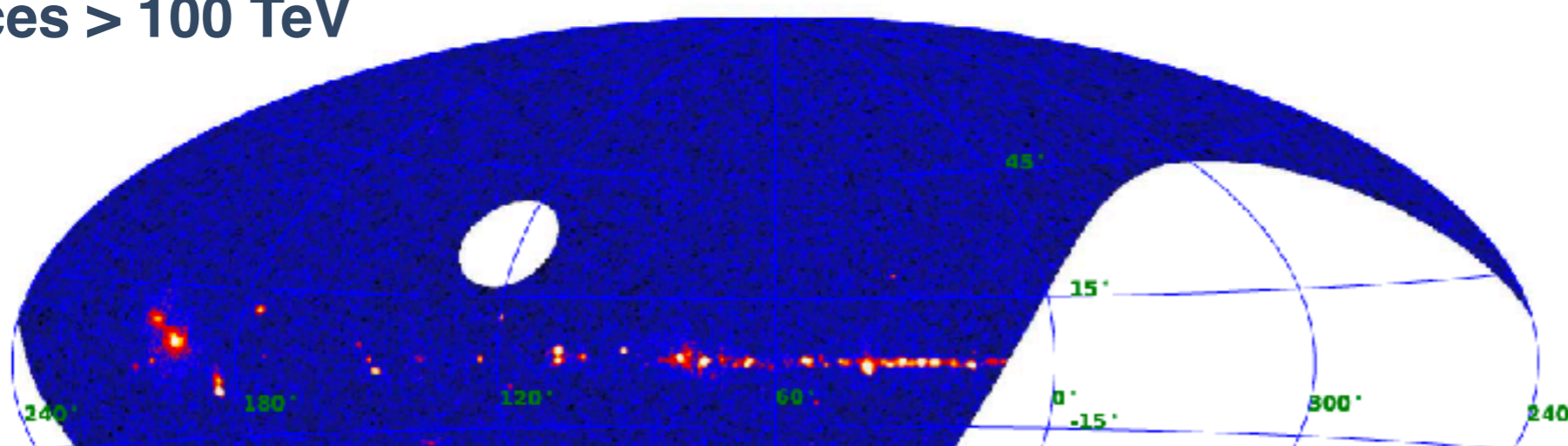
[LHAASO, Cao et al. 2024]

WCDA (1 TeV < E < 25 TeV) Significance Map



- 90 sources
- 75 sources > 45 TeV
- **43 sources > 100 TeV**

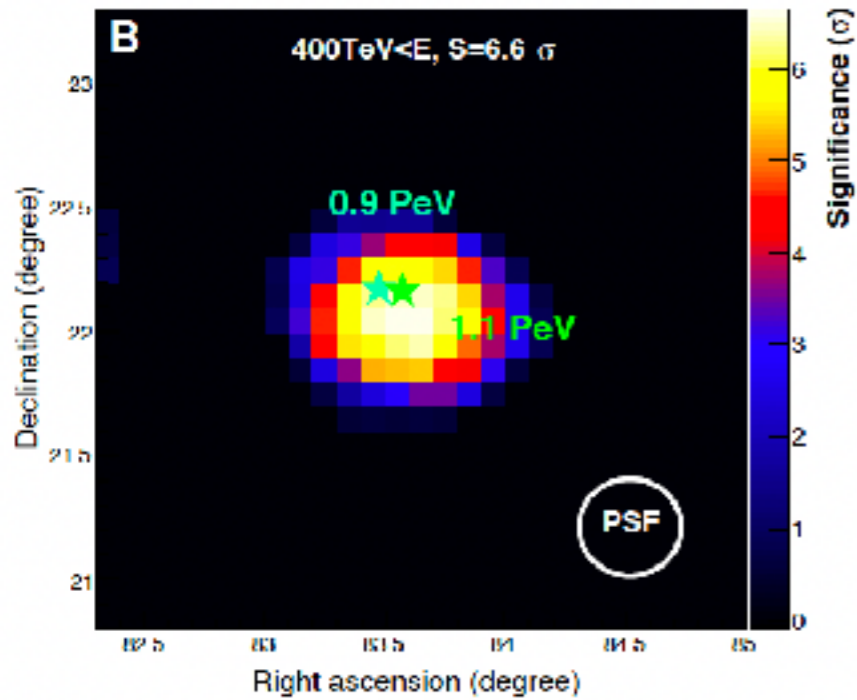
KM2A (E > 25 TeV) Significance Map



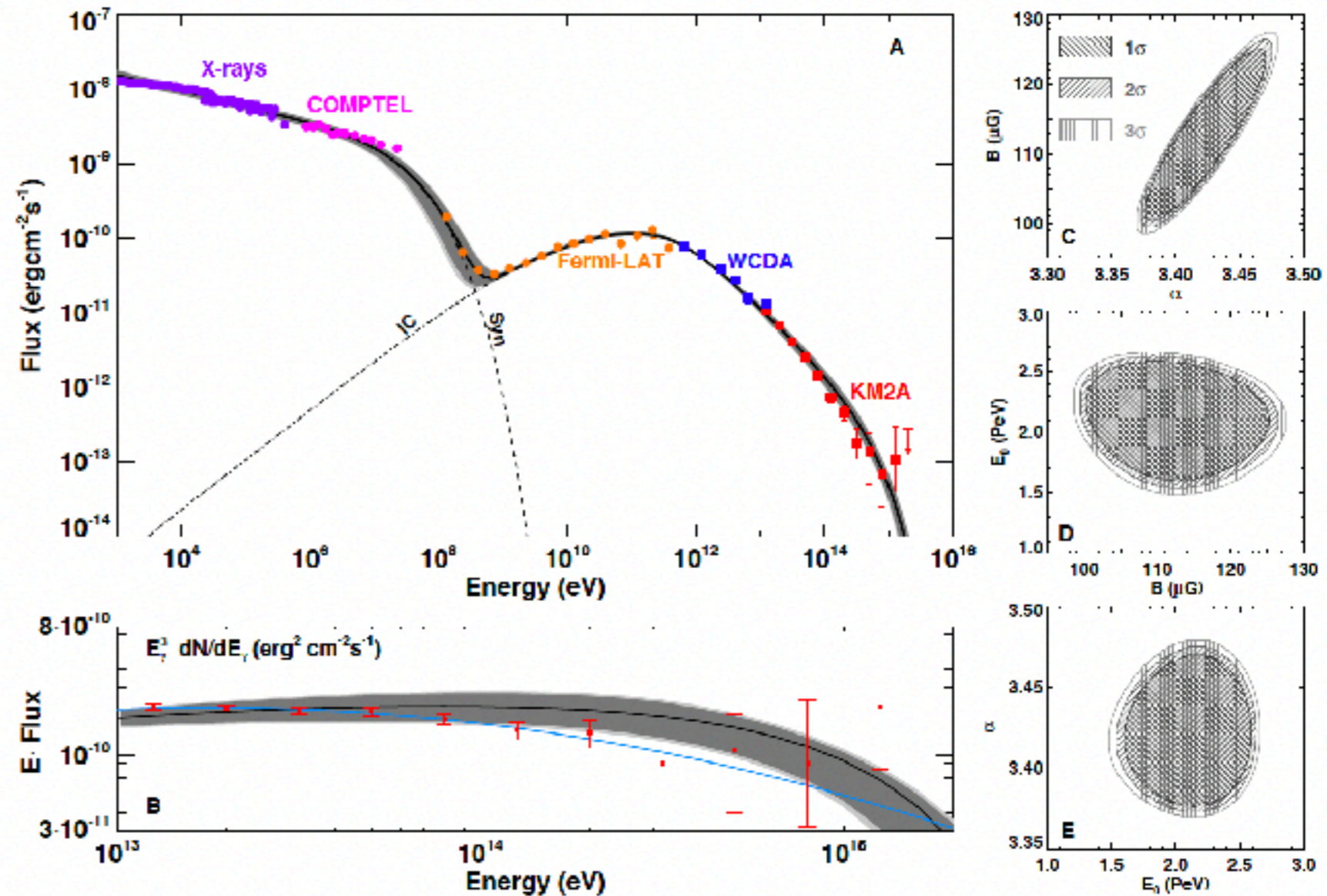
Surprisingly: 22/43 UHE sources spatially coincident with energetic pulsars  
→ If leptonic, where are the hadronic PeVatrons?

# The Crab PWN as a likely PeV electron accelerator

Photon of 1.1 PeV !



[LHAASO, Cao et al. 2021]



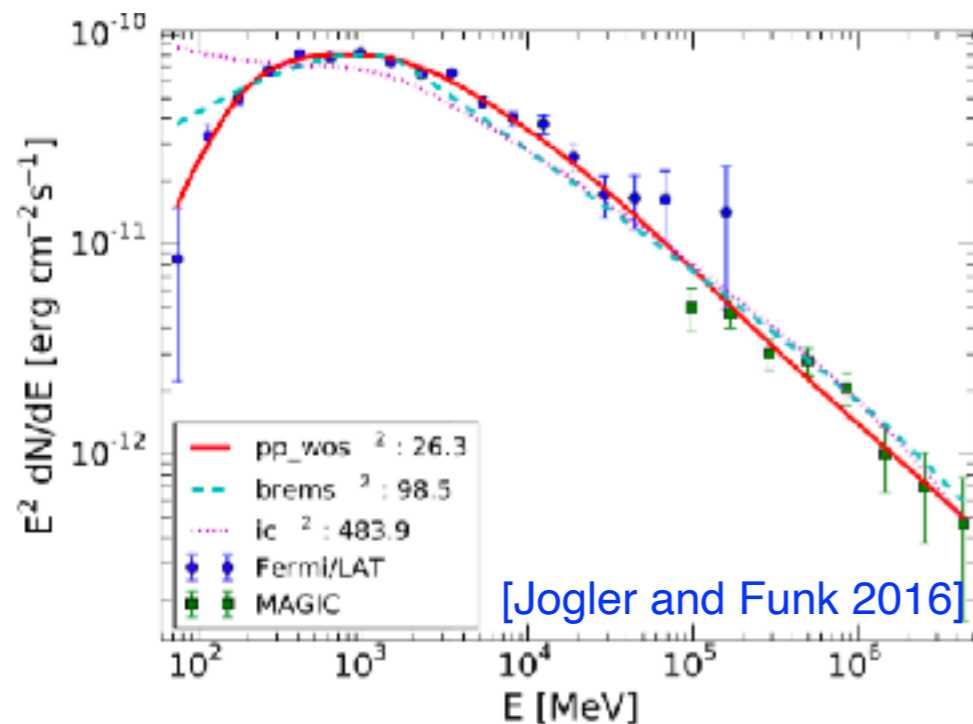
➔ Likely leptonic:  $E_{\text{max},e} = 2.15 \text{ PeV}$ ,  $\alpha = 3.42$ ,  $B = 112 \text{ muG}$

Leptonic PeVatron

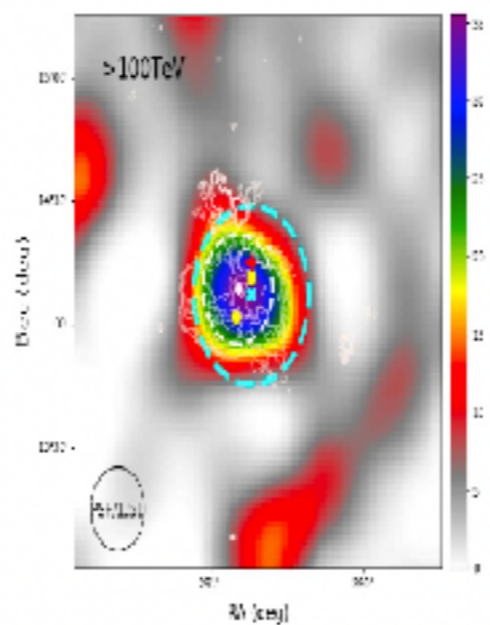
- Steepening of the spectrum between 60 TeV and 500 TeV?
- Hardening of the spectrum at 1 PeV (second population?) ? Clearly not significant at this stage

# Are SNRs still on the stage?

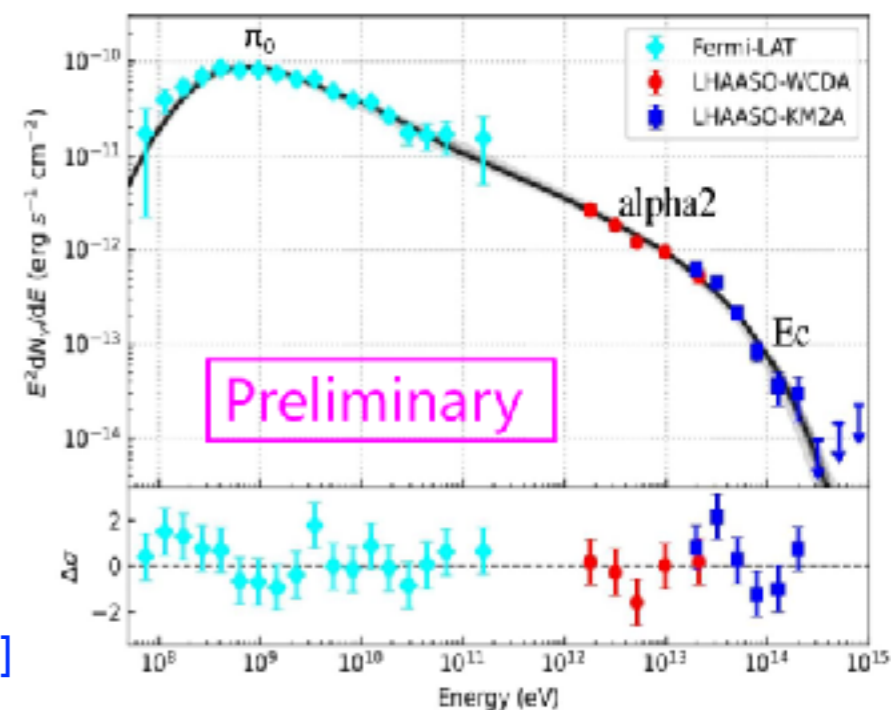
- W51C ( $t \sim 30$  kyr)



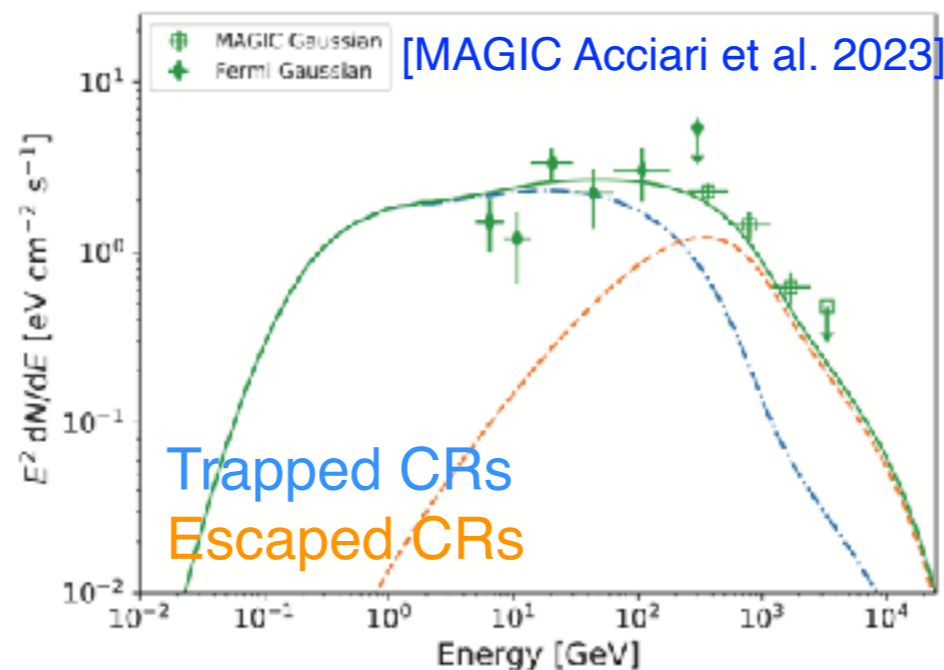
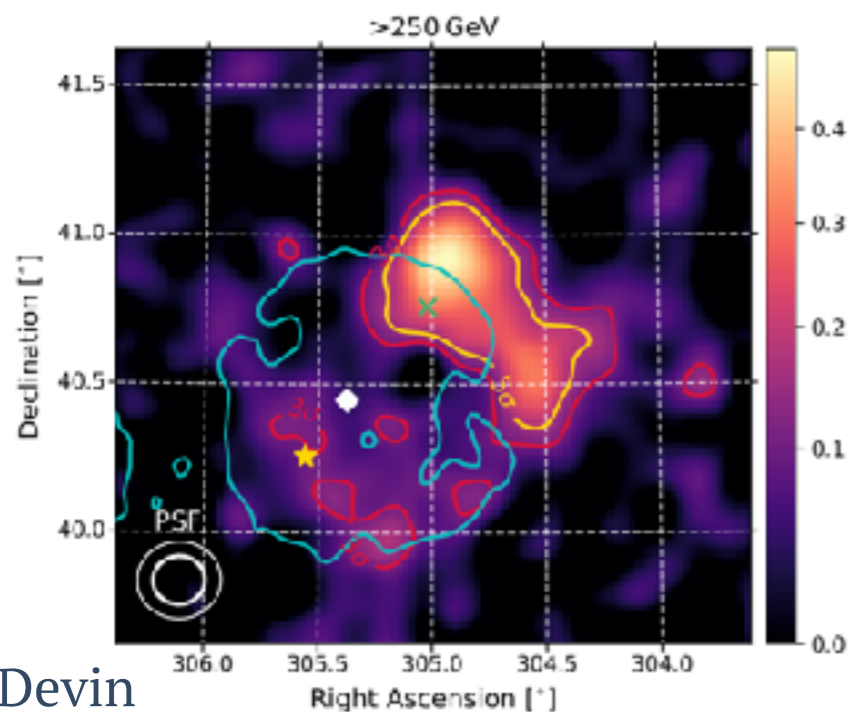
LHAASO detection but  $E_{\text{cut,p}} \sim 400$  TeV



[LHAASO, Fang et al. ICRC 2023]



- Gamma Cygni SNR (G78.2+2.1,  $t \sim 7$  kyr)



Ancient PeVatron?

+ observations with  
LHAASO,  
PoS ICRC2023 (2023) 602



# IceCube diffuse neutrinos flux

