

Theory Session

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Theory Session



What is theoretical Physics?

Field of theoretical physics is incredibly broad:

- ▶ Methods and Phenomena strongly depend on **Energy scale**
- ▶ In general: try to find mathematical means to *model* phenomena in Nature
- ▶ Theory \Leftrightarrow **Testable** mathematical description of Nature
 - \Rightarrow Make new predictions, **test** predictions in experiments

Example: Einstein's **General theory of Relativity**

- \Rightarrow Explain Perihelion of **Mercury**, not possible with Newtonian Mechanics
- \Rightarrow Test GR in bending of light around the Sun
- \Rightarrow predict **gravitational red-shift** and **gravitational waves**

What is theoretical **high-energy** Physics?

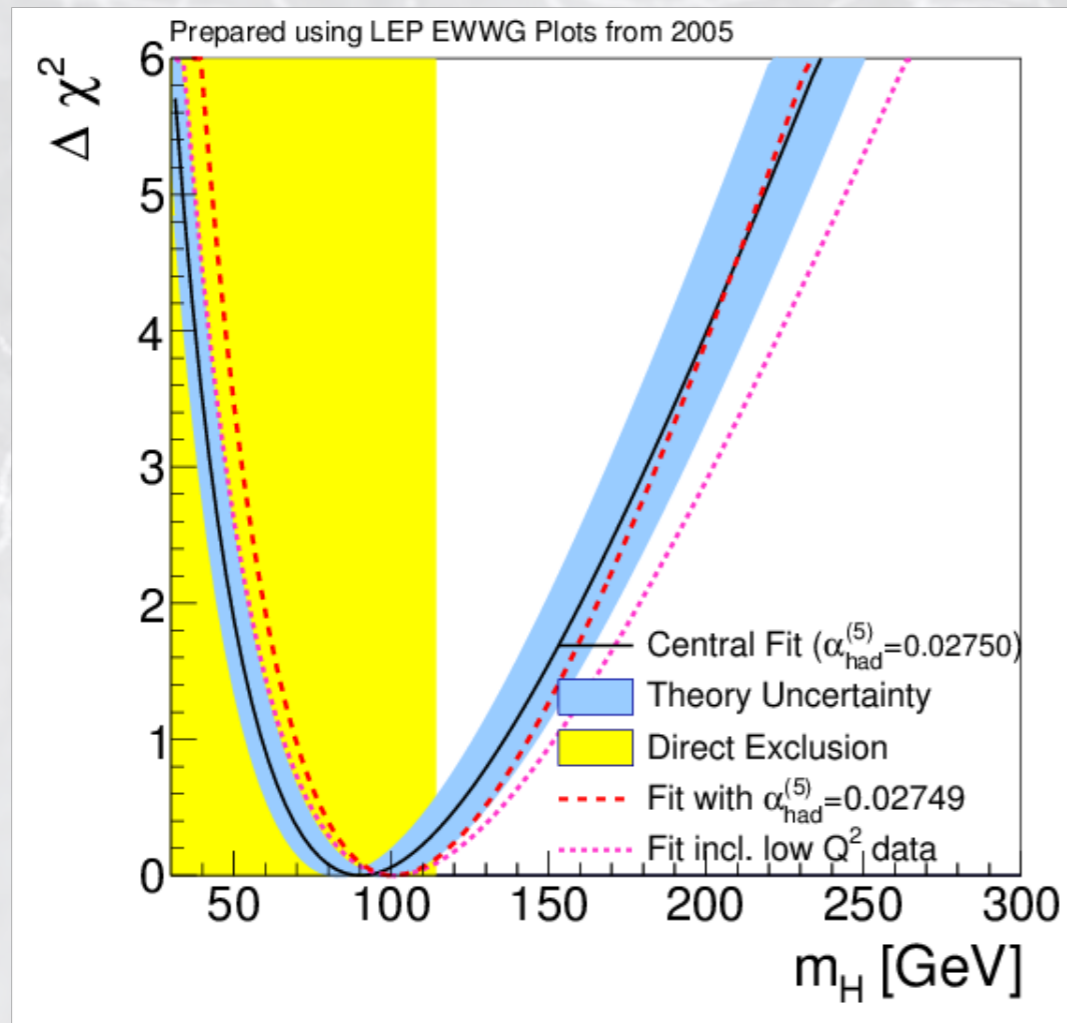
Let's divide in two categories:

“Pure” Theory

- ▶ Develop **Mathematical Foundations**:
Quantum Field Theory, String Theory,
Renormalisation, ...
- ▶ Develop **Computational Methods**:
Feynman Diagrams, Perturbation
Theory, Lattice Field Theory, ...
- ▶ Predict New Phenomena by demanding
mathematical consistency:
Symmetries, Gauge Theory, ...



What is theoretical **high-energy** Physics?



No lose “theorem”

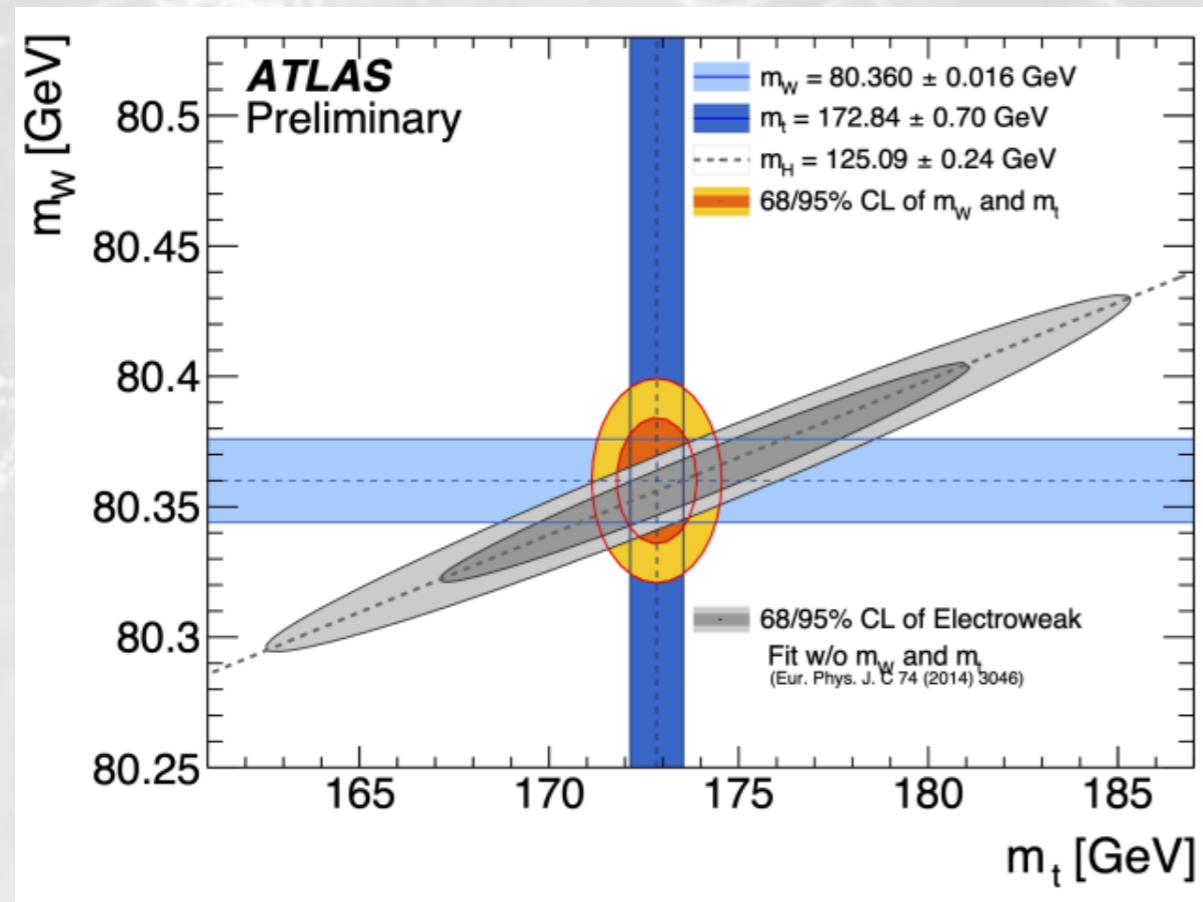
Phenomenology

- ▶ **Apply QFT** to develop Models of Nature: Shell Model, Fermi Theory, Standard Model, Effective Field Theories, ...
- ▶ Fit Models to **Data** and predict Observables to **test** models
- ▶ Construct new models to describe **“anomalous data”**

What is theoretical **high-energy** Physics?

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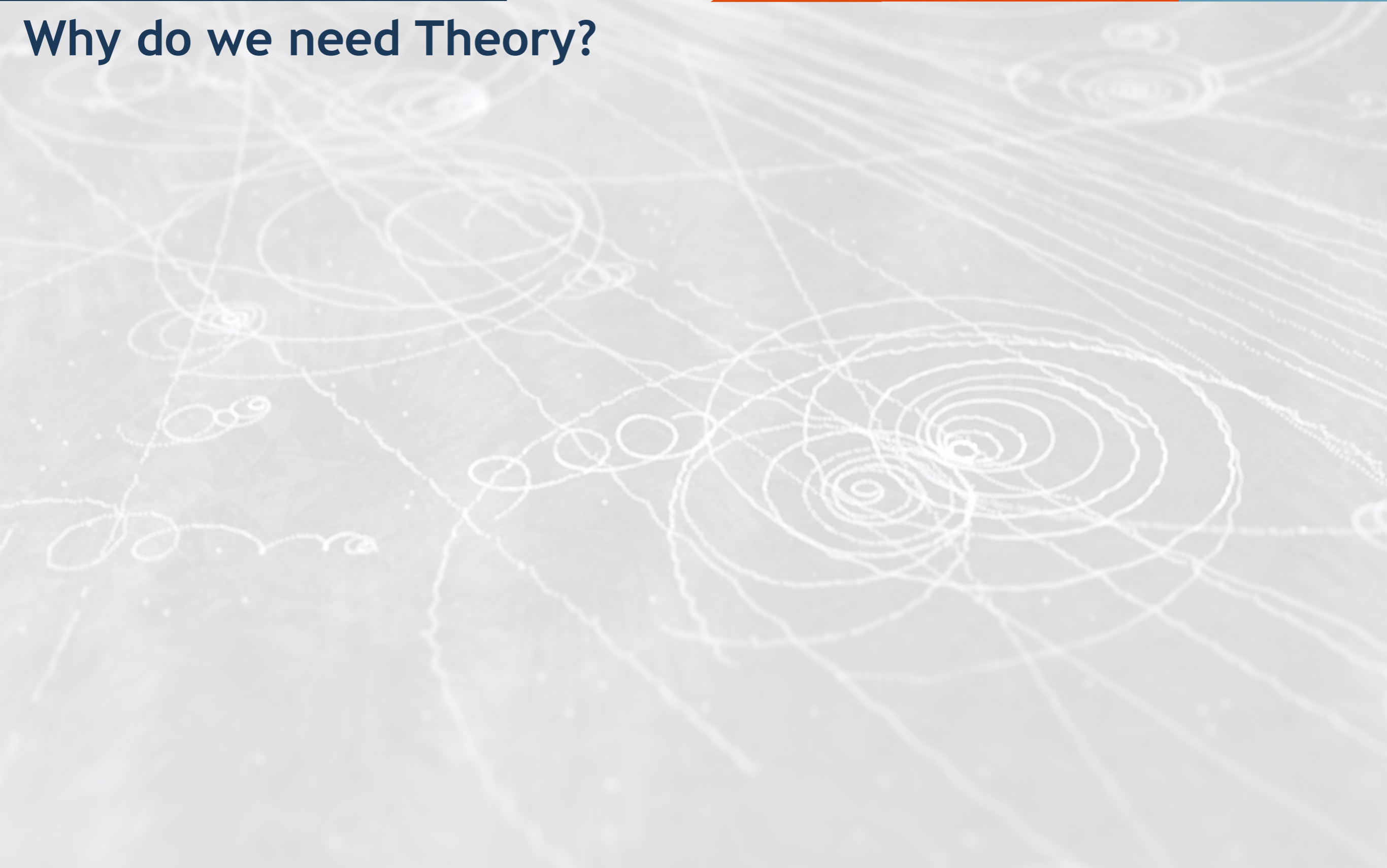
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No lose “theorem”

Why do we need Theory?



Why do we need Theory?

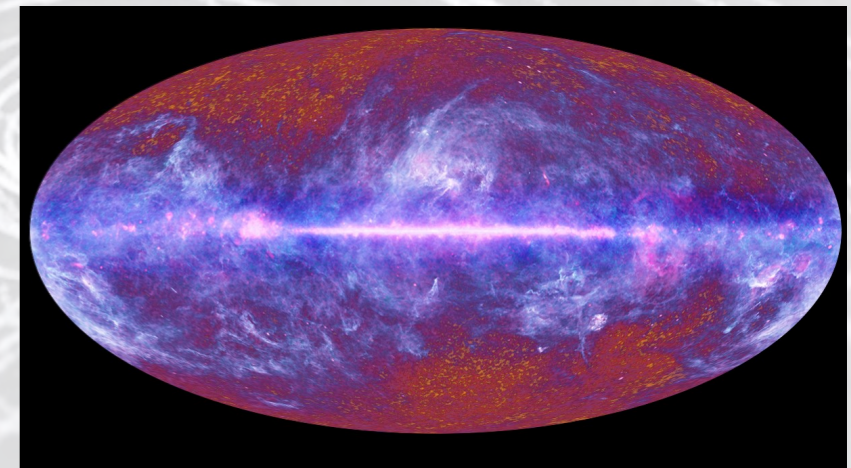
Ehhhh... Because it's fun 🤪??

The fun today:

- ▶ **Axion** emission from **Strange Matter** in Core Collapse **Supernovae** (Maël Cavan)

- ▶ Cosmological and astrophysical constraints on resonant s -wave **Dark Matter annihilation** (Margaux Jomain)

- ▶ **Giant Dipole Resonances** and pygmy Resonances within the **Large Scale Nuclear Shell Model** approach (Oscar Le Noan)



(picture unrelated)

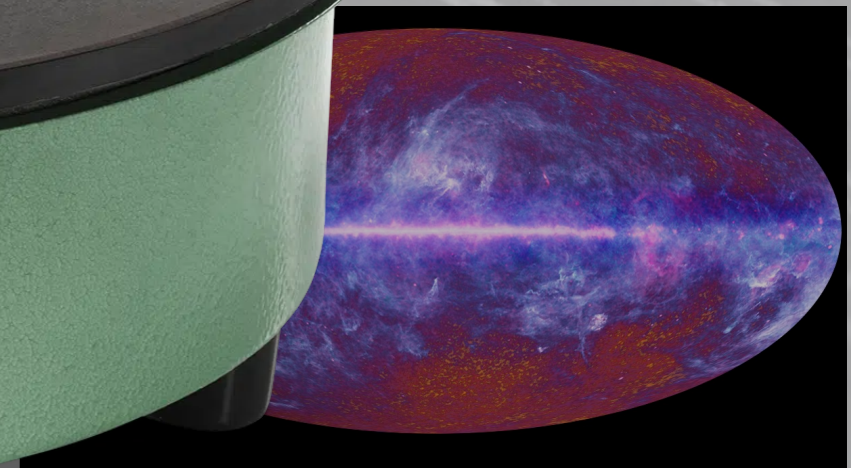
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Some Particle Physics
Phenomenology Crash
Course

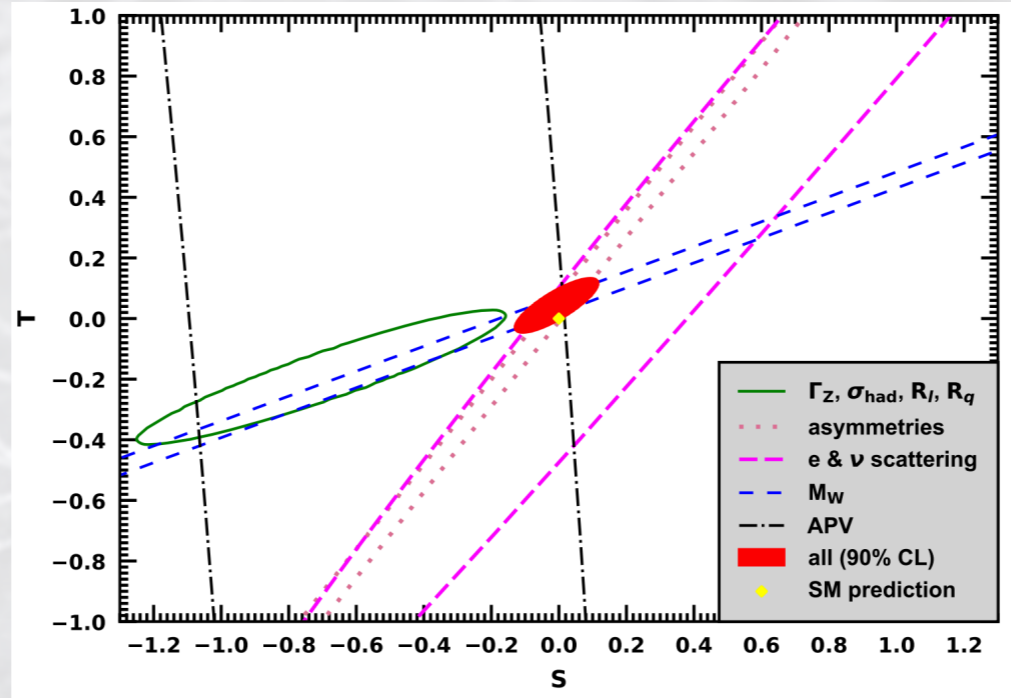


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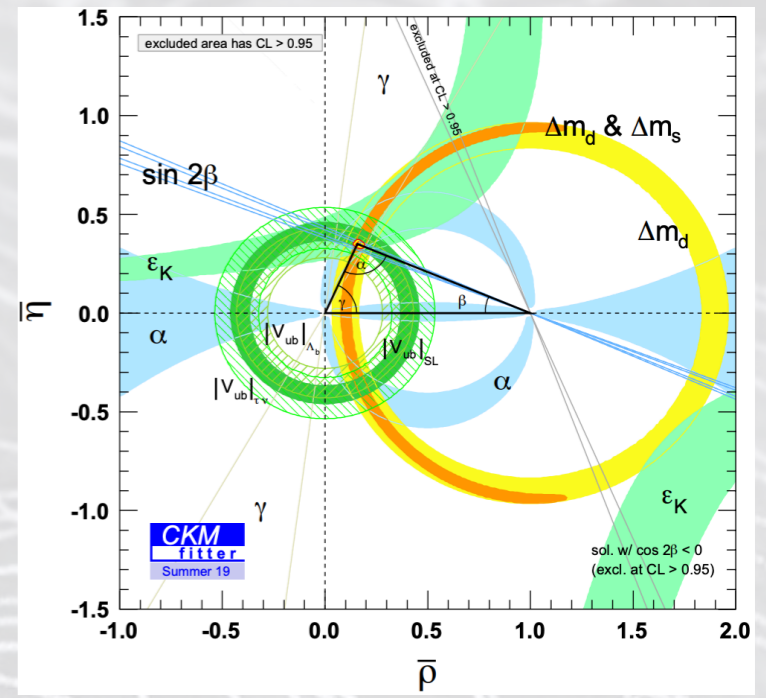
The Standard Model – A success story

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + i\bar{\psi}\not{D}\psi + h.c. + \chi_i y_{ij} \chi_j \phi + h.c. + |D_\mu \phi|^2 - V(\phi)$$

$\mathcal{G}_{SM} = SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$
+ some *accidental* symmetries



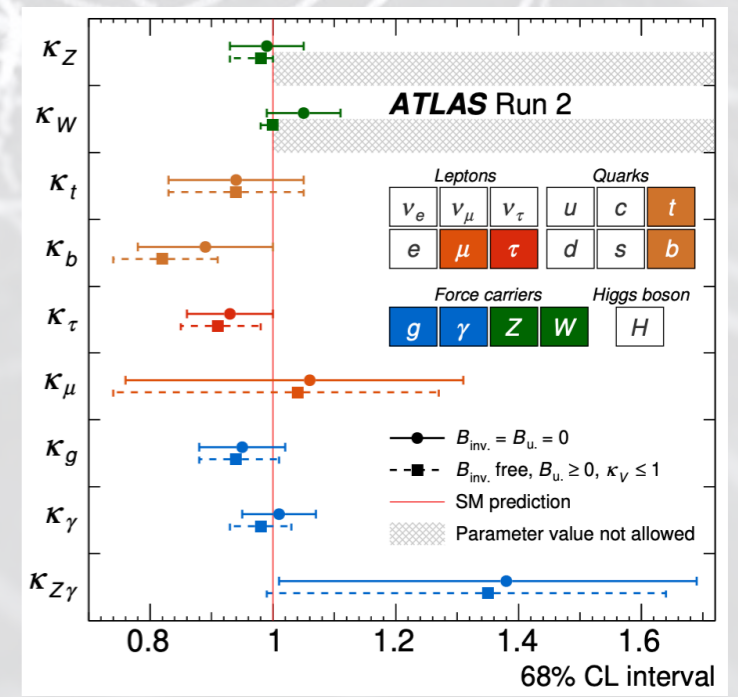
Electroweak fit



CKM paradigm of flavour mixing

Hundreds of experimental measurements overwhelmingly confirm the SM!

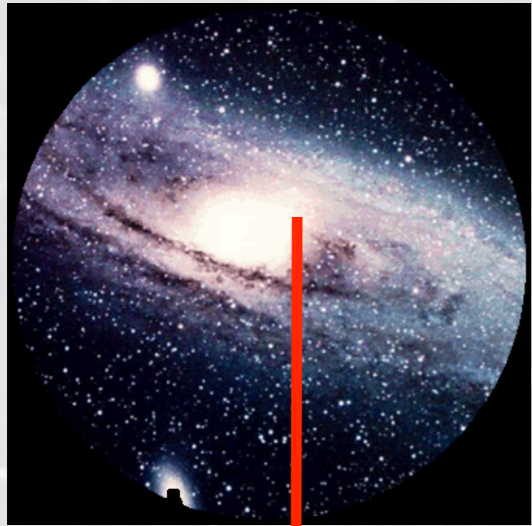
⇒ Just need to completely understand the Higgs sector now?



Higgs couplings

Strong arguments in **f(l)avour** of New Physics!

Some remaining puzzles of Nature:

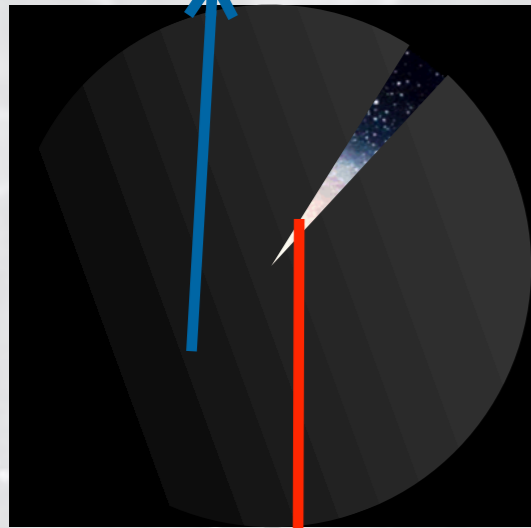


SM matter

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Dark matter

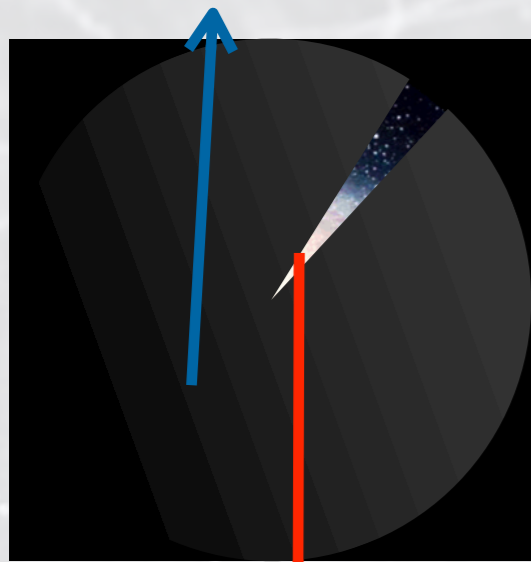


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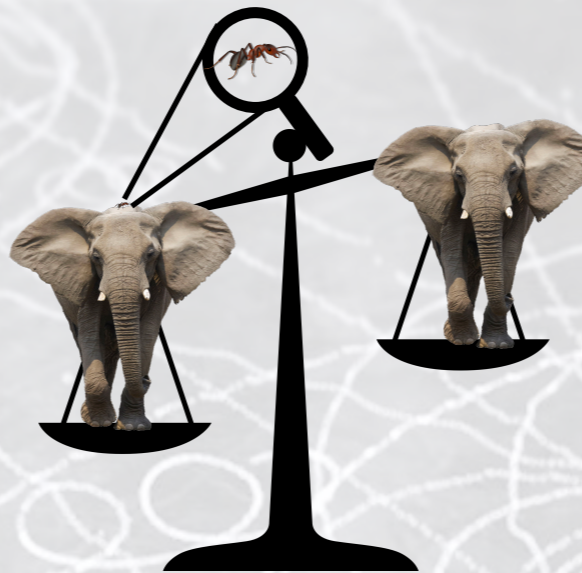
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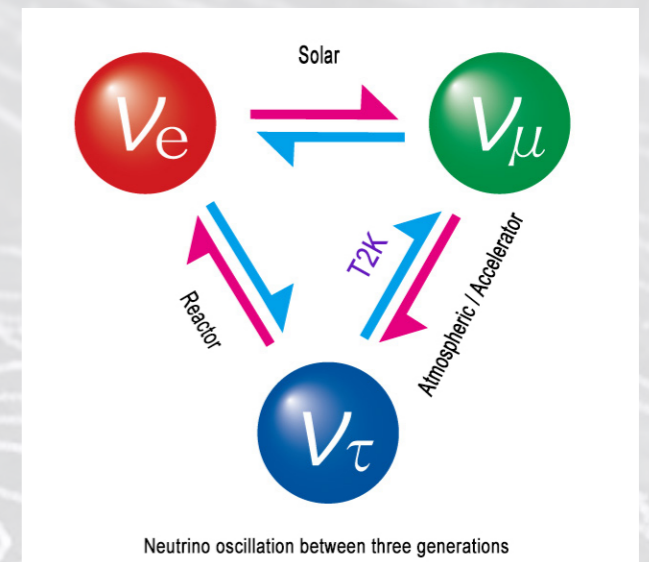
SM matter

Baryon asymmetry of the Universe



$$\frac{n_B - n_{\bar{B}}}{n_\gamma} \simeq \mathcal{O}(10^{-10})$$

ν -oscillations



Some “*theoretical*” issues:

► Quantum theory of Gravity

► Flavour puzzle

► “Hierarchy problem”: why is the Higgs so light and the Planck-scale so high?

► “Strong CP problem”: what is the mechanism behind the absence of $G\tilde{G}$?

Axions & **Supernovae**

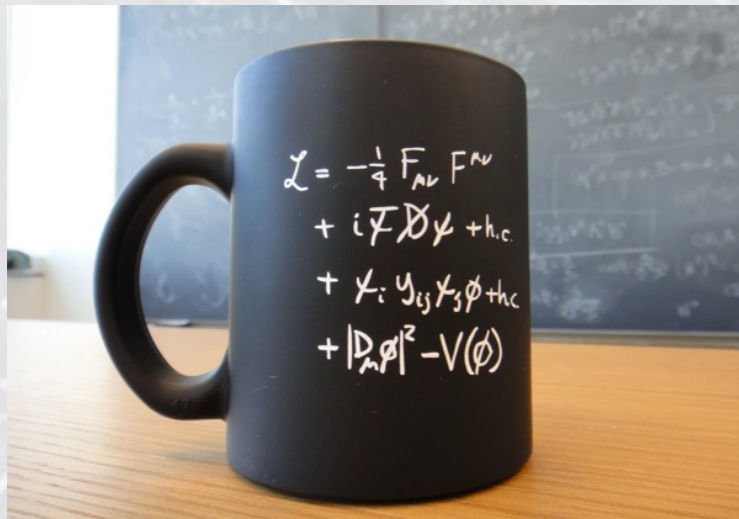


The strong CP problem

Quantum ChromoDynamics (QCD) symmetry is $SU(3)$

Non-abelian symmetries allow for “special” term in the Lagrangian:

(For $SU(2)$ the term can be re-absorbed in the W fields)



$$+ \theta \frac{g}{32\pi} G_{\mu\nu} \tilde{G}^{\mu\nu} \text{ with } \tilde{G}_{\mu\nu} = \epsilon_{\alpha\beta\mu\nu} G^{\alpha\beta}$$

$\Rightarrow G\tilde{G}$ term is CP violating

Measurements of neutron EDM:

$$\bar{\theta} = \theta - \arg \det Y_u Y_d \lesssim 10^{-10}$$

Coincidence?

Axions and the strong CP problem

Technical Naturalness à la 't Hooft: a given parameter is allowed to be small if it enhances the symmetry of the Theory

Give Symmetry origin to a small $\bar{\theta}$? e.g. Left-Right symmetry, many approaches ...

(QCD-) Axions:

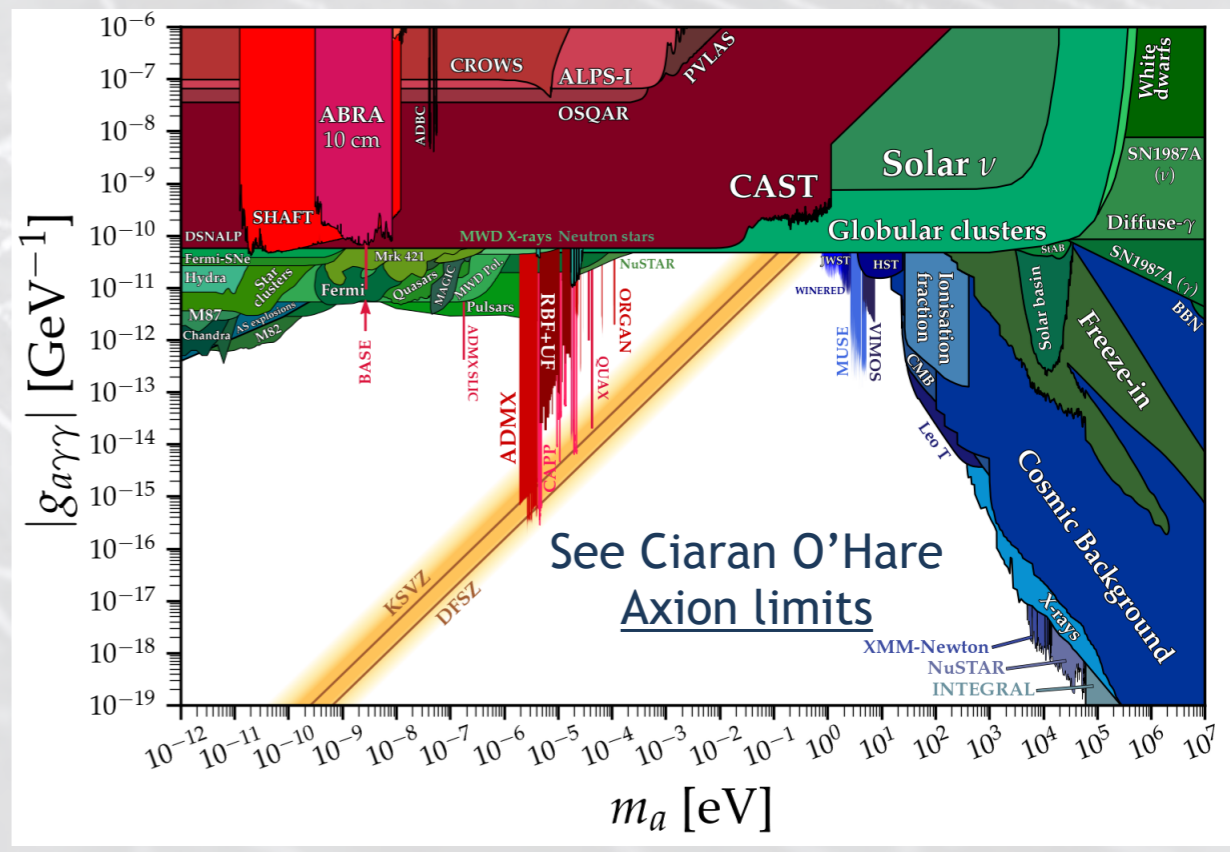
$$\mathcal{L} \supset \left(\frac{a}{f_a} + \theta \right) \frac{1}{32\pi^2} G\tilde{G}$$

Symmetry: $a \rightarrow a + \alpha f_a, \theta \rightarrow \theta - \alpha$

Generate other couplings via RGE:

$$\mathcal{L} \supset \frac{\partial_\mu a}{f_Q} Q^\dagger \sigma^\mu Q$$

Many technical details to UV origin, here just few comments on axion EFT...



See Ciaran O'Hare Axion limits

For a proper intro axions see e.g. the [TASI lecture notes](#) by Anson Hook

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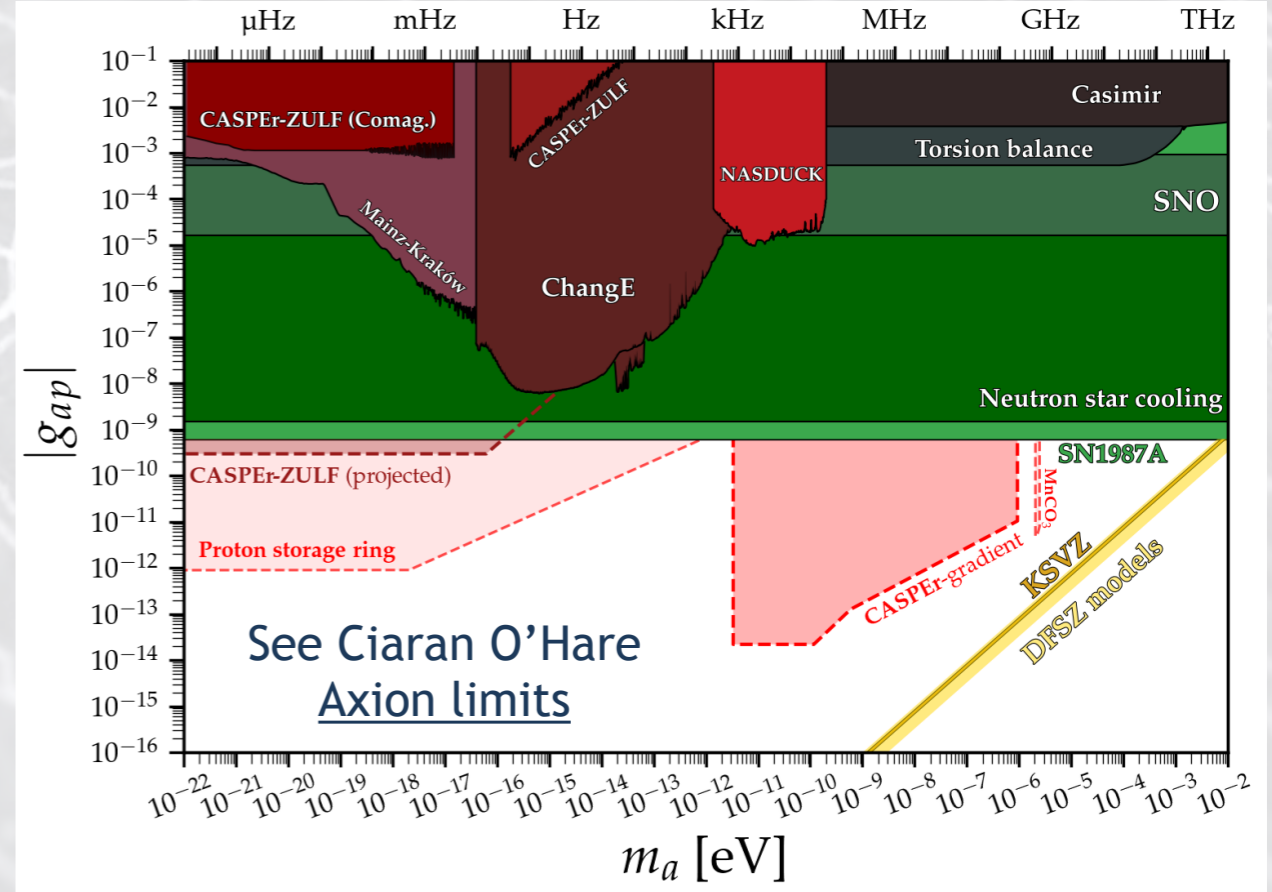
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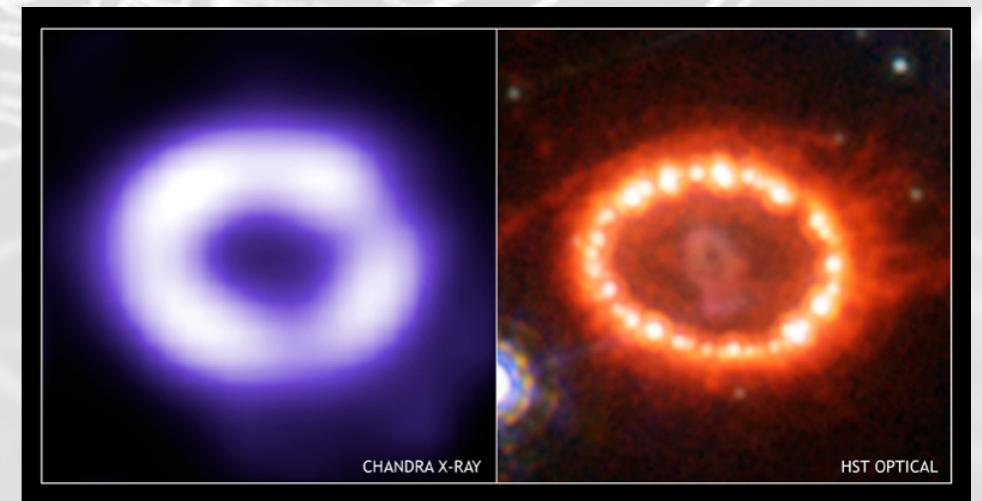
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Supernovae

Some fun facts:

(For everything else see Cosmo-session and [Mael's talk](#))

- ▶ Use Type Ia Supernovae e.g. as Standard Candles for distance measurements
Maaaaany observations
- ▶ Observed 1 (one) supernova with Neutrino burst: **SN 1987A**
- ▶ Supernova 1987A was a **Core collapse Supernova (Type II)**
- ▶ $\sim 10^{58}$ were neutrinos emitted, 25 observed
TWENTY-FIVE
- ▶ Also spectral measurements
- ▶ Supernova models rely on this data
- ▶ Use Supernovae as Ultra-high-luminosity particle physics laboratories

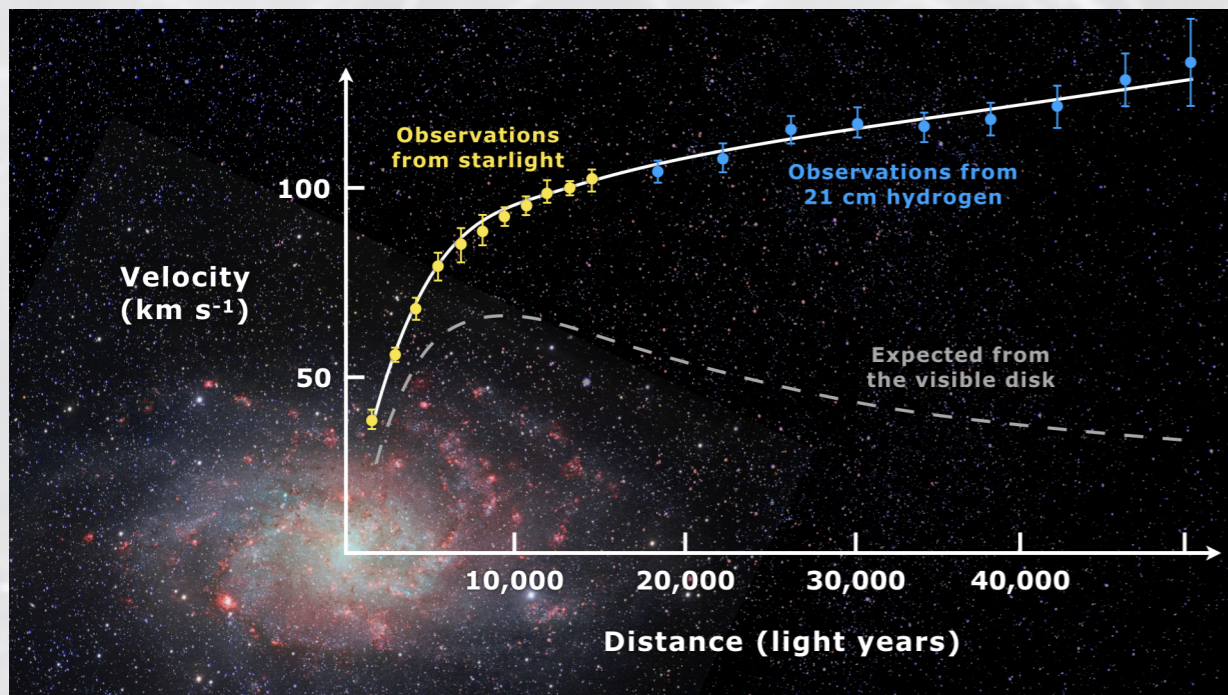


Dark Matter & Astrophysics

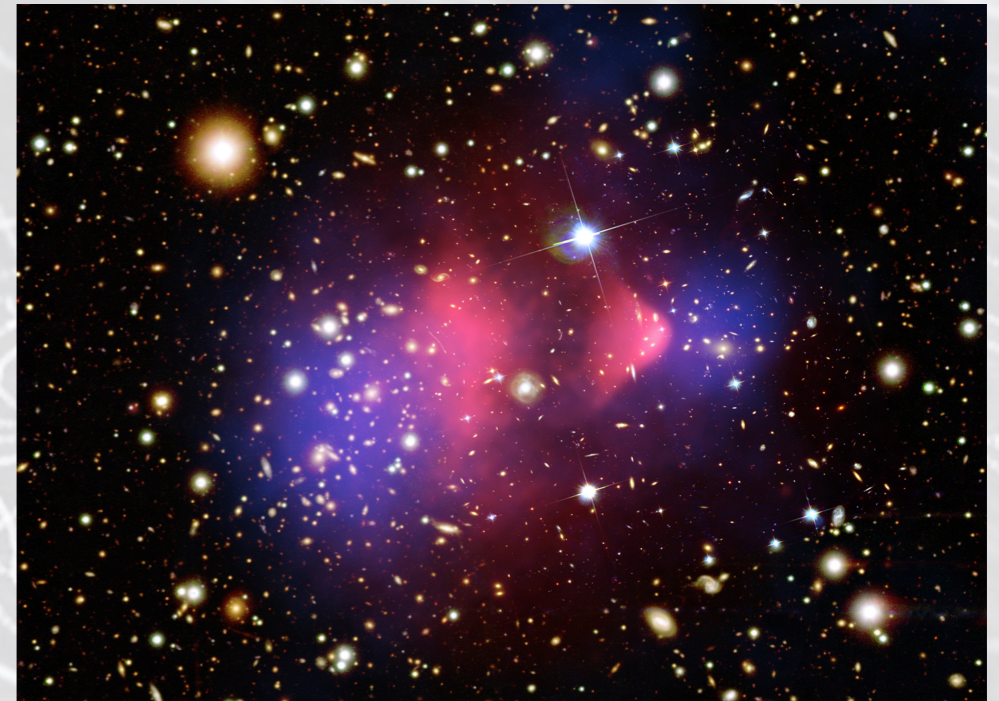


Evidence for Dark Matter

- ▶ Galaxy rotation curves do not follow expectations for visible matter



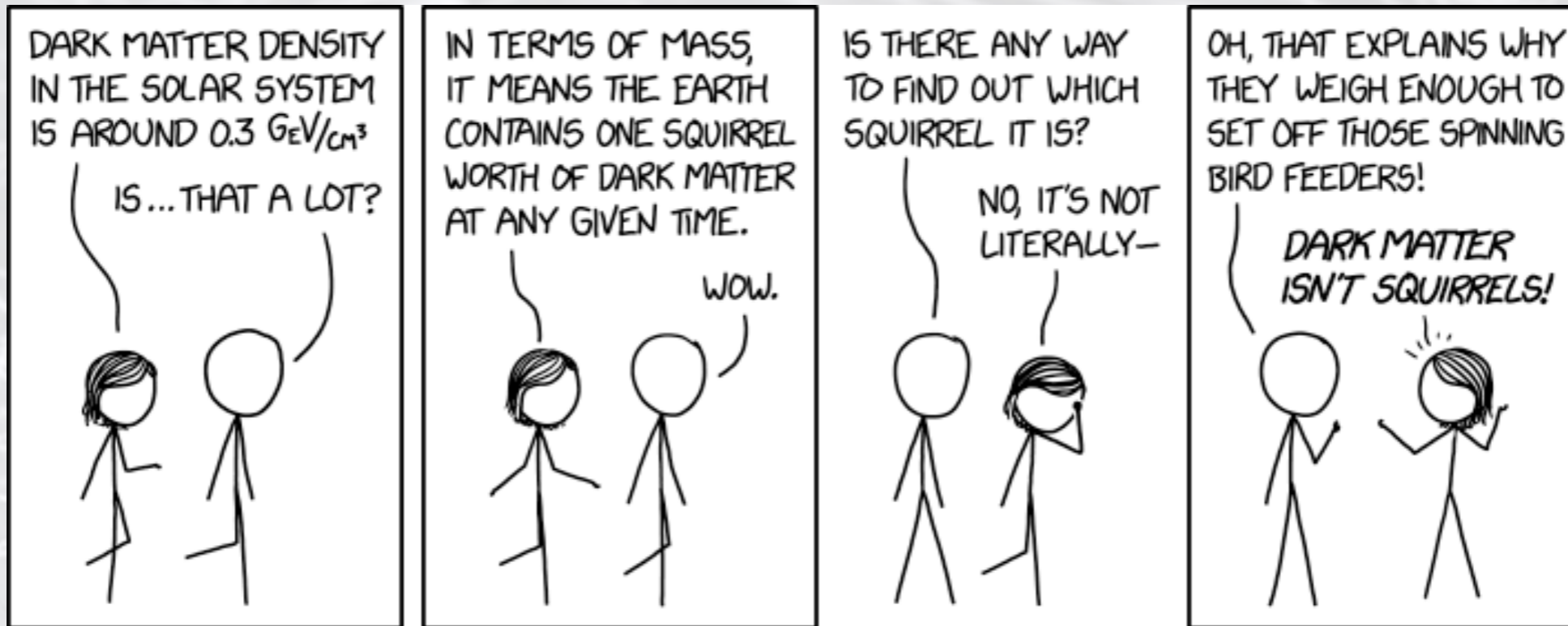
- ▶ **Bullet Cluster**: Colliding galaxies visible matter does not explain result



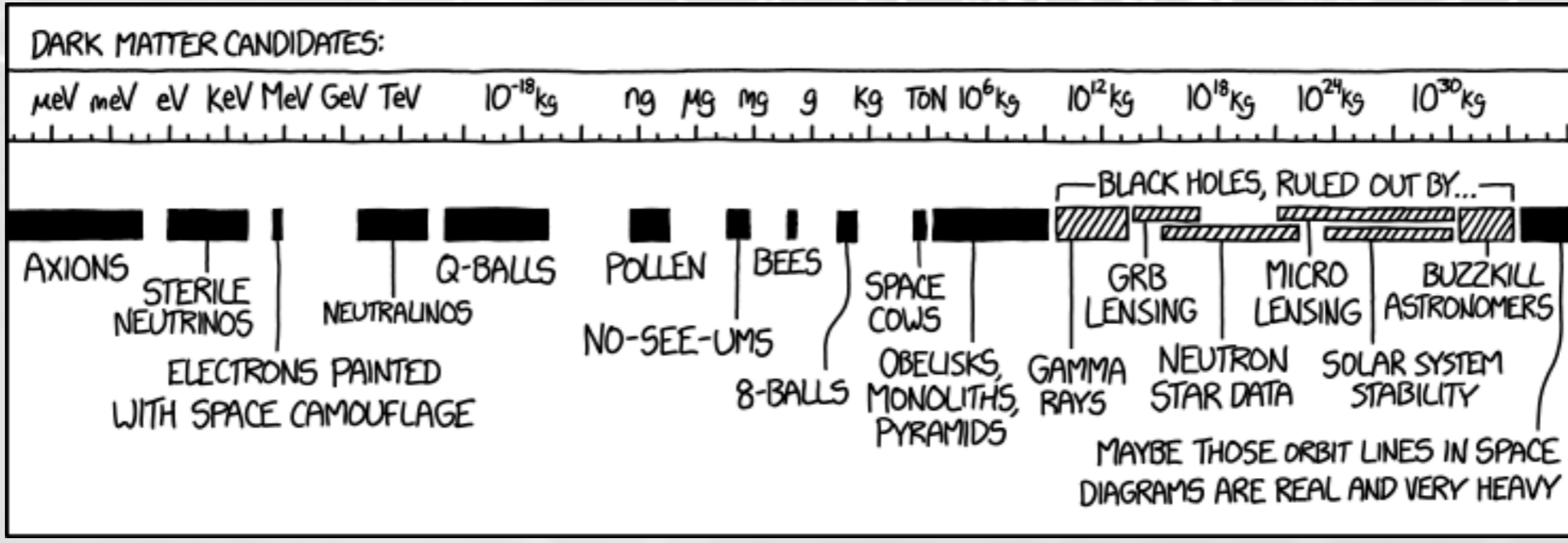
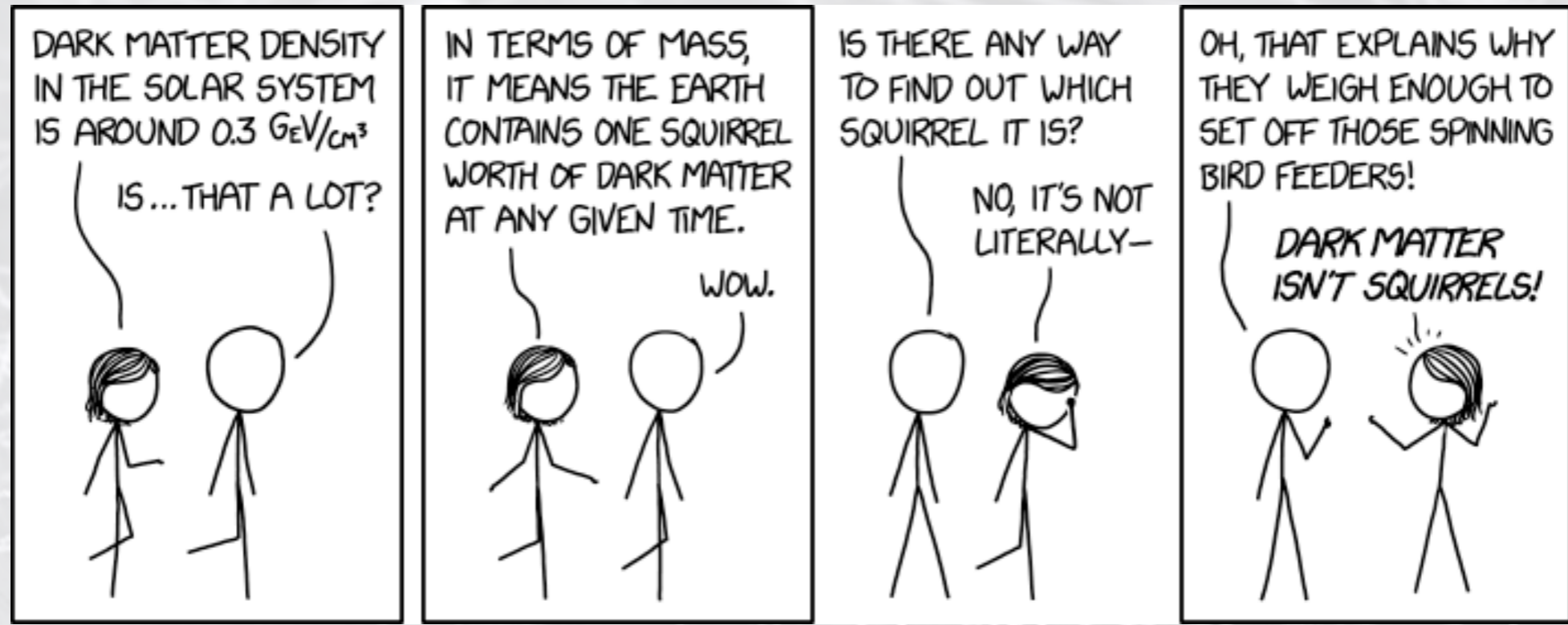
- ▶ (Large) Structure formation does not work with visible matter
- ▶ **CMB** multipole spectrum, **BAO's**, **Redshift** distortions, Lyman- α forrest, ...

Some of these (but not all) are explainable with modified gravity, or **MO**dified Newtonian Dynamics (**MOND**)

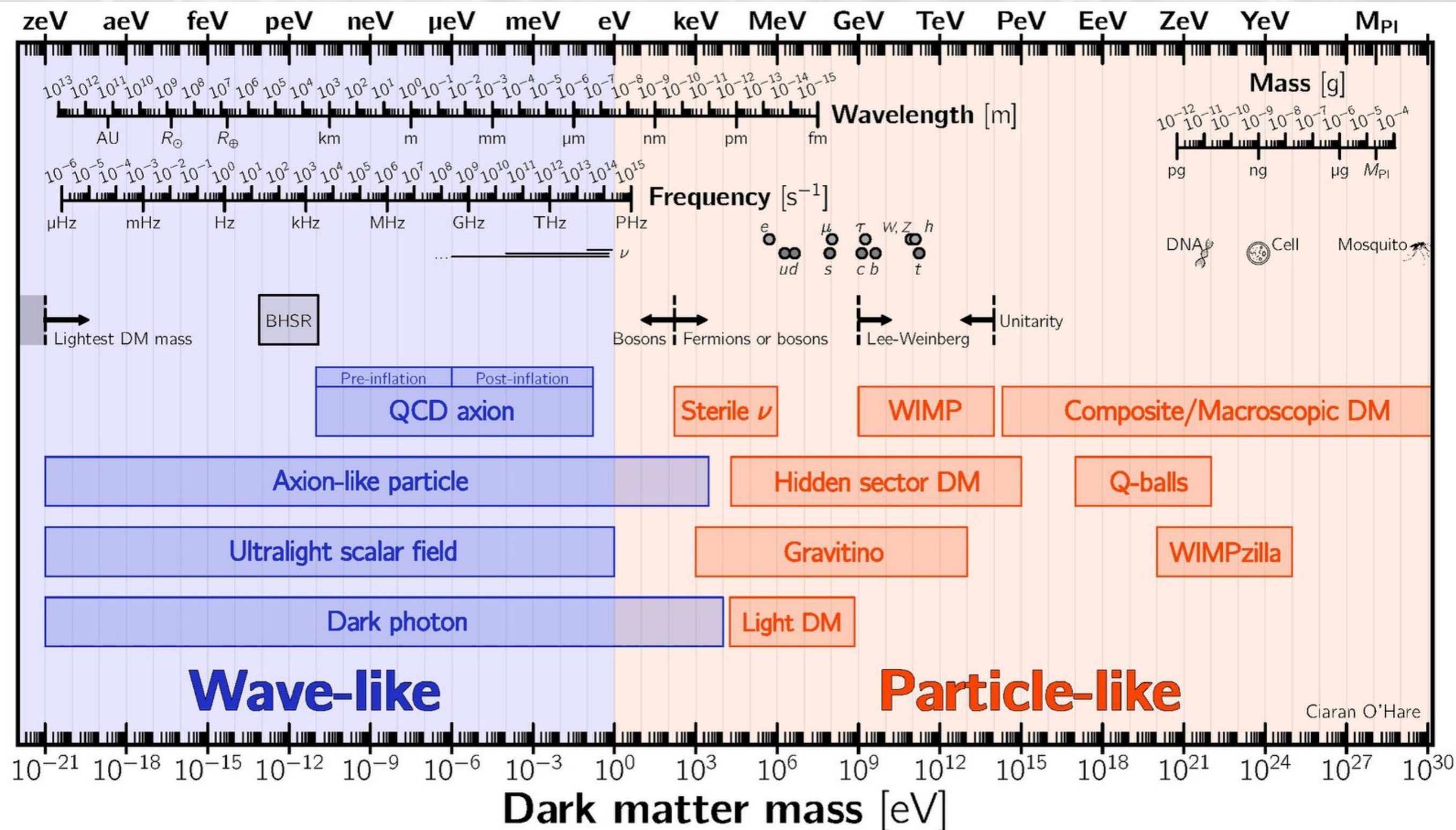
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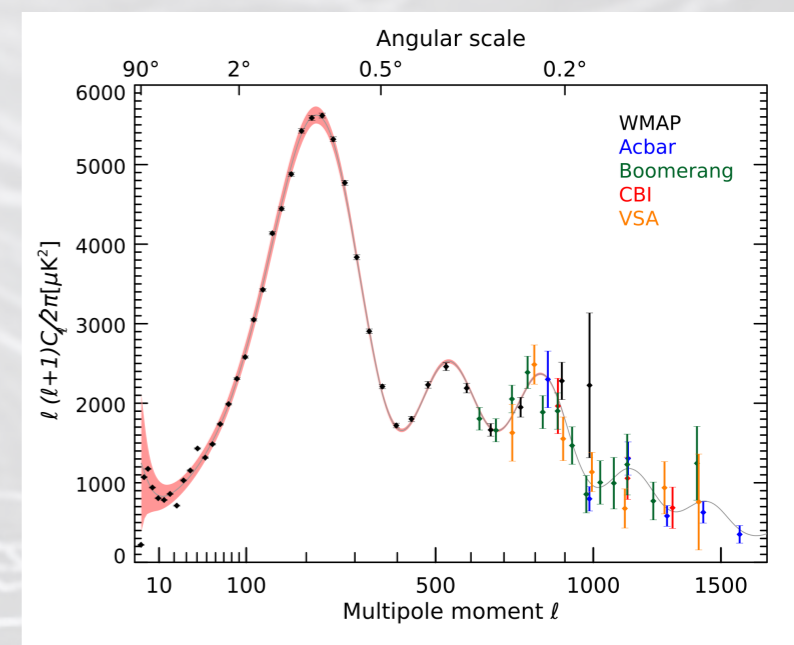
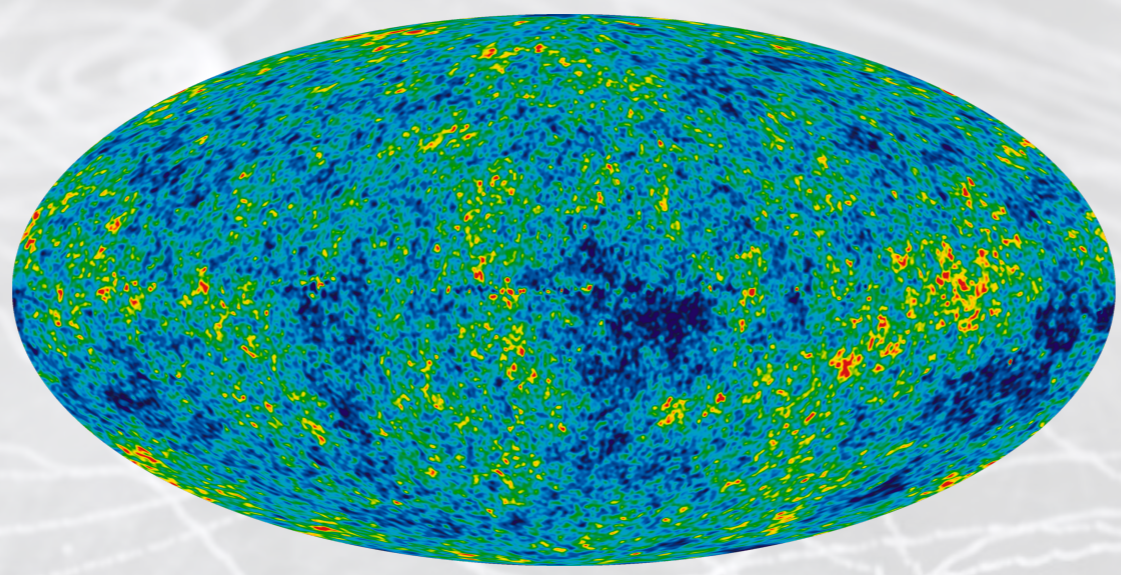
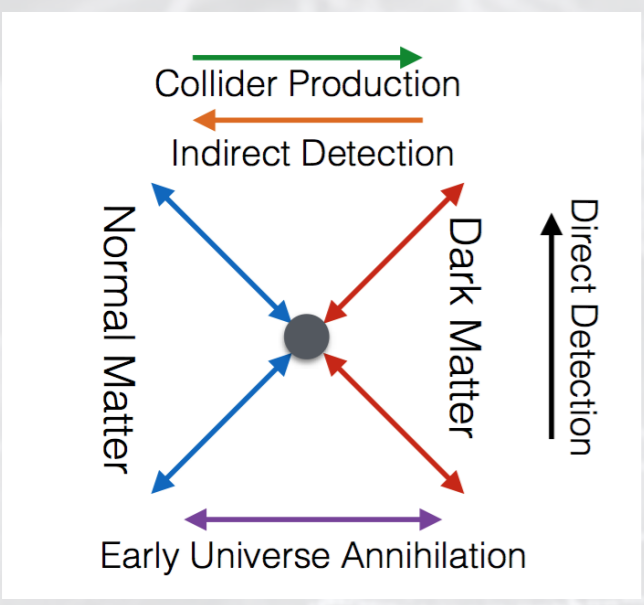
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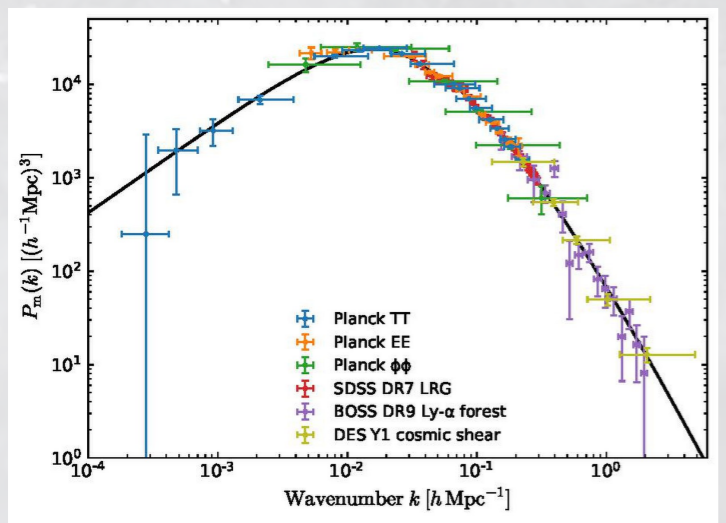


How do we find Dark Matter?

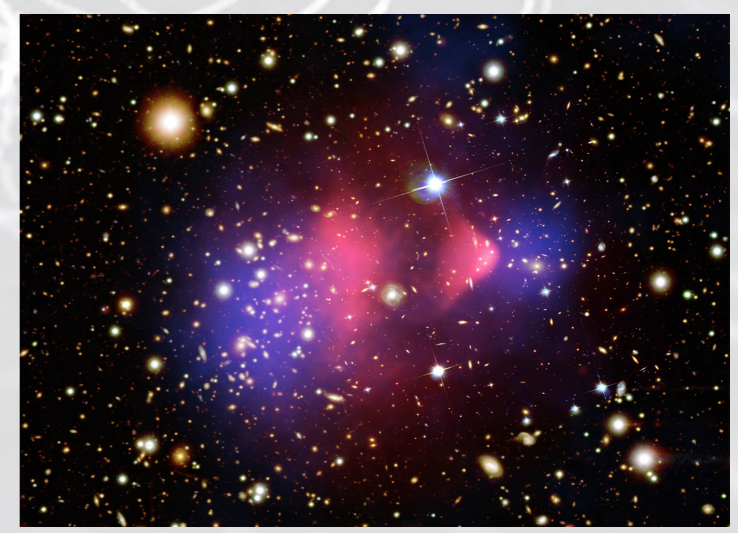
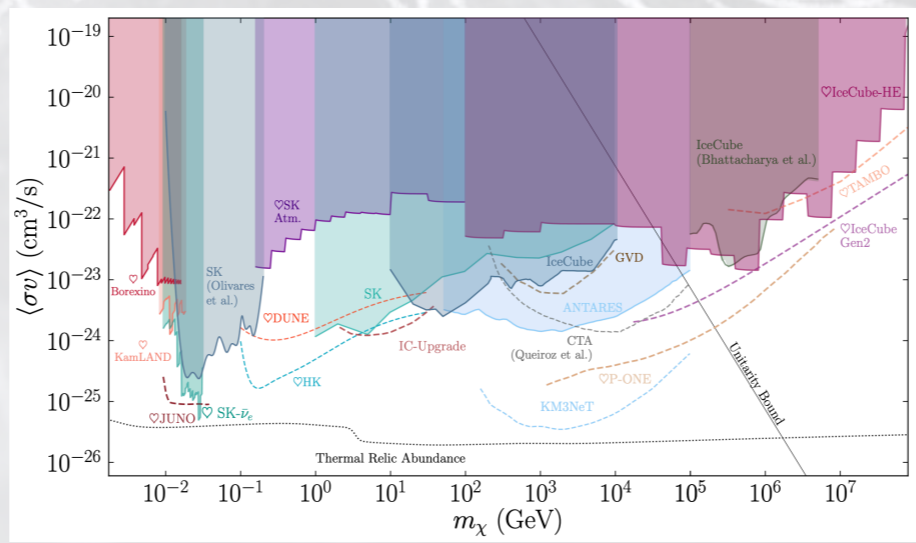


Structure formation: Cold or warm DM?

Bullet Cluster: Self-interactions?



Detection on Earth?



Today: Indirect detection of Dark Matter (and how to escape it): Margaux's talk

Nuclear Theory



Nuclear Theory in a nutshell

The problem: many body interactions

Few particles: $2 \rightarrow 2$, $2 \rightarrow N$ scatterings, $1 \rightarrow N$ decays

Many many particles: Describe *collective behaviour*

⇒ Statistical Mechanics: **Gases, Fluids, Plasmas**, ...

⇒ Low energy: **condensed matter, soft matter, crystals**, ...

Nuclear Physics somewhere in between ...

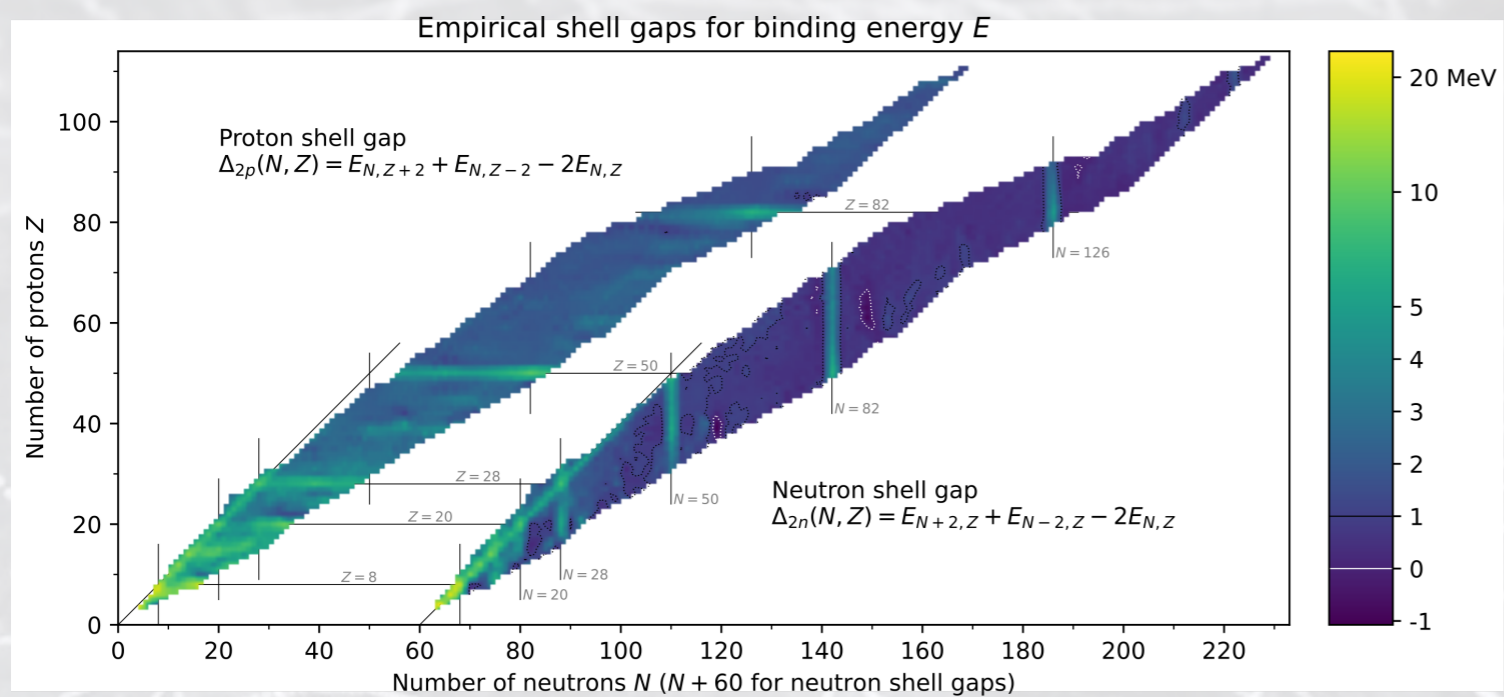
Need to model **strong interactions** between (many) nucleons

Require *large-scale* computations (expensive)



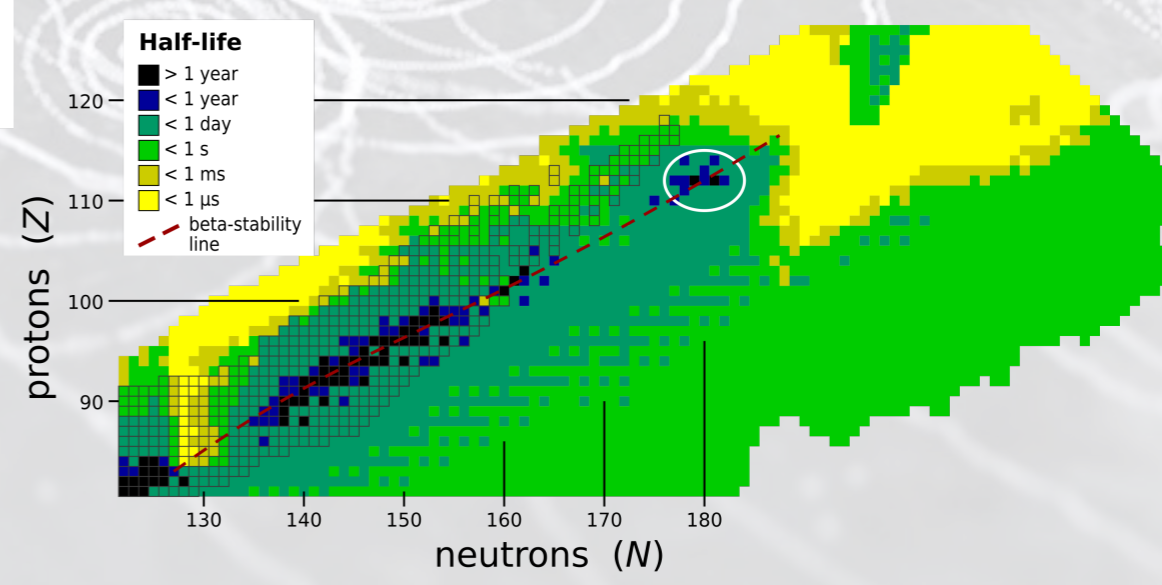
Shell model

Explain **Magic Numbers**: 2, 8, 20, 28, 52, 82, 126, ...



...184 ???

Describe basic Nuclear structure



Predict “Island of Stability”

Many more (and more advanced approaches):

Large Scale Shell Model (**LSSM**), (Relativistic) (Quasi) Random Phase Approximation, Open Quantum Systems, ...

(For everything else see Nuclear session and Oscar's talk)

Why Nuclear Theory?

Experimentalists and Theoreticians across all Energy scales rely on Nuclear Theory:

- ▶ Cosmology: **Big Bang Nucleosynthesis (BBN)**
- ▶ Astrophysics: Theory of **Stars, Supernovae, etc, ...**
- ▶ Particle Physics: **Neutrino Experiments @ Nuclear Reactors, test SM in Nuclear transitions, β -decay experiments, ...**
- ▶ Industry/everyday life: **Nuclear Reactors, Imaging technologies for Medicine, Radiotherapy, ...**

Why theoretical **high-energy** Physics ?

because it is *fun*, duh! 🎉

Why theoretical **high-energy** Physics ?

Ok it's also useful:

- ▶ Describe and make sense of phenomena across all settings and **energy scales**
- ▶ Even if abstract, it is **physics**, think e.g. about the **no-lose “theorem”**
- ▶ Necessary for experiments: think of all the theory in **GEANT4, Pythia, ...**

Enjoy the session :)



