



The ORGAN Experiment: Phase 1 results

The Axion Quest, Rencontres du Vietnam



Australian Government Australian Research Council





Australian National University





Aaron Quiskamp











How to detect axions?

Detect photons with resonant cavity \rightarrow **Axion Haloscope** (P. Sikivie 1983) \bullet







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 $\langle \rangle$

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Axion mass m_a determines ν_a







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Cavities are narrow-band so must be tunable!



High-mass (frequency) axion haloscope at UWA

[1] Ballesteros, G., et al. PRL (2017)

[2] Buschmann, M., et al. Nat Commun (2022)







- High-mass (frequency) axion haloscope at UWA
- Why "high mass" (>40 μ eV or 10 GHz)?
- Largely unexplored but many predictions..

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•
$$\frac{df}{dt} \propto f^{-14/3} \rightarrow$$
 Falling off a sensitivity cliff..

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- Scan between 15 16 GHz
- Tuning rod cavity



4





Rotation stage



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 $\overrightarrow{E}_{cav} \bullet \overrightarrow{B}_{ext} \neq 0$





Rotation stage

 TM_{010} mode



4



- Scan between 15 16 GHz
- Tuning rod cavity
- Moving the rod radially "tunes" the mode frequency









 $\vec{E}_{cav} \bullet \vec{B}_{ext} \neq 0$





Rotation stage

TM₀₁₀ mode



4

Dilution fridge

Phase 1a cavity





5

- Zero-dead-time FFT on FPGA •
- Python and LabVIEW based DAQ ullet
- Operated at 5K with LNF HEMT amp lacksquare



Phase 1a cavity









ORGAN DAQ

Paul Altin (ANU)

Dilution fridge





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Phase 1a cavity









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"Mode Map"







- ~24 days of data-taking
- Limits set between 15.28 16.23 \bullet GHz at $g_{a\gamma\gamma} \sim 3 \times 10^{-12}$ (ALP cogenesis)

-1) $(GeV^{-}$ 10^{-11} $|g_{a\gamma\gamma}|$

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- Gaps to be filled in phase 2 of ulletORGAN

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- 26 27 GHz target range
- Remember resonator is necessarily small, $\nu_c \propto R_c^{-1}$



Rotation stage





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- Extremely sensitive to alignment and rod tilt
- Challenging to set antenna coupling
- Novel resonator needed..

















• Rectangular cavity with a tunable wall





- Rectangular cavity with a tunable wall
- New tunable rectangular cavity solves many problems! •





- Rectangular cavity with a tunable wall
- New tunable rectangular cavity solves many problems!
- Rectangle cavity VS Tuning rod cavity



Tunable rectangular resonant cavities for axion haloscopes

Ben T. McAllister,^{1,2,*} Aaron P. Quiskamp,² and Michael E. Tobar² ¹ARC Centre of Excellence for Dark Matter Particle Physics, Swinburne University of Technology, John St, Hawthorn Victoria 3122, Australia ²QDM Lab, Department of Physics, University of Western Australia, 35 Stirling Highway, Crawley, Western Australia 6009, Australia

(Received 2 October 2023; accepted 28 November 2023; published 16 January 2024)

Axions are a compelling dark matter candidate, and one of the primary techniques employed to search for them is the axion haloscope, in which a resonant cavity is deployed inside a strong magnetic field so that some of the surrounding axions may convert into photons via the inverse Primakoff effect and become trapped inside the resonator. Resonant cavity design is critical to the sensitivity of a haloscope, and several geometries have been utilized and proposed. Here we consider a relatively simple concept—a rectangular resonant cavity with a tunable wall—and compare it to the standard tuning rod-type resonators employed in the field. We find that the rectangular cavities support similar modes to cylindrical tuning rod cavities, and have some advantages in terms of axion sensitivity and practicality, particularly when moving to higher frequencies which are of great and growing interest in the international axion dark matter community.

DOI: 10.1103/PhysRevD.109.015013





Using materials on-hand







Using materials on-hand









RF choke lid $Q_L : \sim 1000 \rightarrow 4000$



- Using materials on-hand
- <u>NO</u> mode crossings between 26 27 GHz



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- Using materials on-hand
- <u>NO</u> mode crossings between 26 27 GHz
- Very easy to tune minimal motor steps required → reduced cryogenic heat load

RF choke lid $Q_L :\sim 1000 \rightarrow 4000$

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Parameter	Rectangle	Tuning-rod
С	~ 0.4*	~ 0.43











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Antenna Coupling	Relatively easy	X Challenging











Limits on $g_{a\gamma\gamma}$





Limits on $g_{a\gamma\gamma}$

Can exclude ALP cogenesis with 95% confidence!





Limits on $g_{a\gamma\gamma}$

- Can exclude ALP cogenesis with 95% confidence! \bullet
- sensitivity!





(A quick detour..)





(A quick detour..)

→ Quantum • Q





(A quick detour..)

- → Quantum Q ullet
- Search around 6 GHz





(A quick detour..)

- $Q \rightarrow Quantum$
- Search around 6 GHz
- **Uses Josephson Parametric** \bullet **Amplifier** (JPA): $\downarrow T_s$





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- $Q \rightarrow Quantum$
- Search around 6 GHz
- **Uses Josephson Parametric** lacksquare**Amplifier** (JPA): $\downarrow T_s$
- Operates at mK: $\downarrow T_s$

Variable coupling: † $Q_{\rm L}$

Useful R&D for future ORGAN phases







(A quick detour..)

- $Q \rightarrow Quantum$
- Search around 6 GHz
- Uses Josephson Parametric Amplifier (JPA): $\downarrow T_s$
- Operates at mK: $\downarrow T_s$

• Variable coupling: 1 $Q_L \frac{\beta^2}{(1+\beta)^2}$

Useful R&D for future ORGAN phases









Raytheon JPA

• JPA very sensitive to environment





Raytheon JPA

- JPA very sensitive to environment
- Tune the JPA resonance using flux bias





Normalised reflection phase





-1

ORGAN-Q Setup





- Scanned between 6.15 6.35 GHz
- ~17 days of data



Mode Map @ 4K









- Scanned between 6.15 6.35 GHz ullet
- ~17 days of data
- Initial 4K measurements $Q_0 \sim 14,000$
- Degraded to $Q_0 \sim 4,000$ over many cools













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- Tradeoff between tuning, Q and T_{rod}
- Adjusted antenna coupling $\beta \simeq 1.45 \pm 0.32$
- $T_{sys} \simeq 928 \pm 108 \,\mathrm{mK}$ using SNRI method













ORGAN-Q Limits

Most sensitive limits in the region





arXiv:2407.18586





ORGAN-Q Limits

- Most sensitive limits in the region ullet
- Comparable sensitivity to neighbouring \bullet experiments





arXiv:2407.18586





ORGAN-Q Limits

- Most sensitive limits in the region \bullet
- Comparable sensitivity to neighbouring \bullet experiments
- ORGAN-Q was a successful testbed for \bullet future phases of ORGAN
 - Clamshell cavity
 - Millikelvin operation
 - JPA operation and characterisation \bullet
 - Cryogenic switch + hot load \bullet calibration
 - Adjustable antenna coupling \bullet

 $|g_{a\gamma\gamma}|$ [GeV⁻





arXiv:2407.18586







Future (optimistic) Plans



How to increase V?



"Pizza cavity"

Jeong, Junu, et al. 'Search for Invisible Axion Dark Matter with a Multiple-Cell Haloscope'. Physical Review Letters, vol. 125, no. 22, APS, Nov. 2020





How to increase V?

- The rectangle geometry is a natural candidate for a 2-• cavity array with iris coupling
- Only 1 readout required ullet
- ~3x increase in scan rate for 2-cavity array
- Easy to tune compared to tuning-rod iris cavity \bullet



"Pizza cavity"

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Top view

Side view



Phase 1

 Targeted searches between 15-16 GHz and 26-27 GHz
 ~ month scale









Phase 2a

 Wider search (15-20 GHz) building on current expertise
 ~ year scale









 10^{-10}

 10^{-11}

 10^{-12}

 10^{-13}

 $\overline{}$

(GeV

 $g_{a\gamma\gamma}$

- Wider search (15-20 GHz) building on current expertise ~ year scale
- mK temperatures SQL ampifiers: $T_{sys} \lesssim 1\,{
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JeV

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- Superconducting cavities: $Q \gtrsim 50,000$









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- Larger $\overrightarrow{B_0}$: 11.5T \rightarrow 16T








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- Larger V: Iris cavity









Phase 2a

 Develop efficent single photon counting (SPC) devices





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- Reach QCD axion model bands





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- Develop efficent single photon counting (SPC) devices
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Phase 2a

- Develop efficent single photon counting (SPC) devices
- Reach QCD axion model bands
- $T_{amb} \lesssim 50 \mathrm{mk}$
- + All previous upgrades





Phase 1 of ORGAN complete







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- Most sensitive limits above 15 GHz







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 - Multiple cavity array: 1V







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- ORGAN-Q most sensitive search between 6.15-6.35 GHz







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- ORGAN-Q most sensitive search between 6.15-6.35 GHz
- Utilise ORGAN-Q expertise in Phase 2











Questions? Thank you for your attention!



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