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### **High-Temperature Superconducting Cavities**

### for Dark Matter Axion Search at CAPP

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## Introduction:

### **Cavity Haloscope Experiment for Dark Matter Axion Search**



## Introduction:

### High Quality Factor Cavity is Needed for Faster Axion Search



## Introduction:

### High Quality Factor Should be Maintained in a High Magnetic Field



### **Material Evaluation Criteria**

**Two Preferred Conditions for Cavity Haloscopes** 

## **High-Temperature Superconductor**

 $\frac{R_s}{R_{Cu}} \ll 1$  in a magnetic field

## **Biaxially-Textured Film**

Minimizing Surface Defect & Magnetic Field Degradation



### **Vortex Dynamics & Energy Dissipation Mechanism**

> Type II superconductor forms vortices (quantized magnetic flux,  $\overline{\Phi_0}$ ) in a magnetic field.



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- Above a depinning frequency, pinning force become negligible.  $\succ$



**High Depinning** 

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- Vortices respond to the incident electromagnetic wave.
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- Above a melting field, vortices are mixed each other. >





### **Vortex Dynamics & Energy Dissipation Mechanism**

- > Type II superconductor forms vortices (quantized magnetic flux,  $\overline{\Phi_0}$ ) in a magnetic field.
- Vortices respond to the incident electromagnetic wave.
- > Every vortex is trapped in each pinning potential well.
- > In the pinning regime, the surface resistance is much smaller than that of copper.



### **Considering Cavity Haloscope Experimental Condition**

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100 mK 8 GHz	R <sub>s</sub> (B = 0 T) (Ohm)	R <sub>s</sub> (B = 8 T, ॥c) (Ohm)	Critical Field (H <sub>c2</sub> )	Depinning Frequency
OFHC Cu (Metal) Low Temperature Superconduct	~ 7E-3 ors (LTS)	~ 7E-3	None	None
<b>NbTi (LTS)</b> Gatti <i>et al</i> . PRD(2019)	~ 1E-6	~ 4e-3	Small ~ 13 T	~ 45 GHz
Nb <sub>3</sub> Sn (LTS) Alimenti <i>et al.</i> SUST(2020) High Temperature Superconduct	~ 1E-6	?	~ 25 T	small ~ 6 GHz
Bi-2212 (HTS) Bi-2223 (HTS)	~ 1E-5	?	> 100 T (IIab) Larbalestier <i>et al.</i> Nature(2001)	?
TI-1223 (HTS)	~ 1E-5	<b>~ 1e-4</b> Calatroni <i>et al</i> . SUST(2017)	> 100 T (IIab) Larbalestier <i>et al.</i> Nature(2001)	<b>12 — 480 MHz</b> Calatroni <i>et al.</i> SUST(2017)
ReBCO (HTS) Re = Y, Gd, Eu,	<b>~ 1E-5</b> Ormeno <i>et al.</i> PRB(2001)	<b>~ 1e-4</b> Romanov <i>et al.</i> Scientific Reports(2020)	> 100 T (IIab) Larbalestier <i>et al.</i> Nature(2001)	Strong Pinning 10 — 100 GHz Romanov <i>et al.</i> Scientific Reports(2020)

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#### **Biaxially-Textured ReBCO Coated Conductor Architecture**

> Larger energy dissipation is originated due to flowing current at high angle grain boundaries.



M. J. Lancaster, "Passive microwave device applications of HTS", Cambridge University Press (2006).

#### **Biaxially-Textured ReBCO Coated Conductor Architecture**

- > Larger energy dissipation is originated due to flowing current at high angle grain boundaries.
- Second generation ReBCO coated conductor technology realizes low angle grain boundaries which have misorientation angles less than 4 degree. (biaxially-textured)



#### **Biaxially-Textured ReBCO Coated Conductor Architecture**

- > Larger energy dissipation is originated due to flowing current at high angle grain boundaries.
- Second generation ReBCO coated conductor technology realizes low angle grain boundaries which have misorientation angles less than 4 degree. (biaxially-textured)
- > 2 dimensional high-quality ReBCO coated conductor is flexible to manipulate.

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#### **Biaxially-Textured ReBCO Coated Conductors Shows Low Surface Resistance**

Cavity Haloscope: 
$$\nu = O(1 \sim 10^2)$$
 GHz,  $T_{phy} = O(10^2)$  mK,  $B = O(10)$  T

 $R_{s}(\nu, T_{phy}, B) = R_{BCS}(\nu, T_{phy}, 0) + R_{defect}(\nu, T_{phy}, 0) + R_{vortex}(\nu, T_{phy}, B) + R_{magnetism}(\nu, T_{phy}, B)$ 



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For Y & Eu atoms





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 $\sim \mathbf{v}^2$  For Y & Eu atoms



### Thin Film Does Not Degrade External Tesla-Scale Magnetic Field



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### Thin Film Does Not Degrade External Tesla-Scale Magnetic Field



### **Development Methods**

#### Strategy 1



Pros: Clean Surface Cons: Slow Fabrication, Electrically Disconnected

#### Strategy 2



Pros: Easy to Fabricate, Electrically Connected Cons: Surface Defect (~ 10%)



### **Strategy 1**

- $\succ$  Simulation study can estimate energy loss from gaps.
- $\succ$  Misalignments and defects are considered based on fabrication error.
- $\succ$  Only evanescent field enter into a gap.

Ahn et al. PRApplied(2022), 17, L061005





### **Strategy 1**





#### **History of Development**



Date	Таре	f (GHz)	n <sub>gap</sub>	Q (0 T)	Q (8 T)	$\mathbf{Q}_{gap}$	Experiment
1	YBCO	6.9	12	0.22 M	0.33 M		Prototype
2	GdBCO	2.3	32	0.60 M	0.50 M		Cavity Haloscope
3	EuBCO+APC	2.3	34	5.0 M	3.5 M	> 10 M	Axion Quark Nugget Search
4	EuBCO+APC	5.4	14	20 M	13 M		Cavity Haloscope
5	EuBCO+APC	1.2 ~ 1.5	?	?	?		Axion Haloscope (CAPP-MAX)

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### Superconducting Cavity for CAPP-MAX

- > CAPP's flagship experiment to search for axion above 1GHz
- Dine-Fischler-Srednicki-Zhitnitsky (DFSZ) sensitivity



B-field map

Ahn's talk on Mon 13:30

Ivanov's talk on Fri 13:55



### Superconducting Cavity for CAPP-MAX

#### arXiv:2402.12892, Submitted to PRX, Accepted



 $\nu_a$  [GHz]

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Below DFSZ sensitivity or Axion as 20% of Dark Matter

Yannis K. Semertzidis's Opening Talk on Mon 10:00

 $\nu_a$  [GHz]



### Superconducting Cavity for CAPP-MAX

arXiv:2402.12892, Submitted to PRX, Accepted



### Strategy 2

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- Sensitive to Contact Problem
- ➢ Gaps should be closed electrically to prevent a radiation



### **Development Methods**

#### Strategy 1



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### Strategy 2



### Strategy 2



### Superconducting Cavity for CAPP-MAX





### Superconducting Cavity for CAPP-MAX



### Superconducting Cavity for High-frequency Axion Search



### Superconducting Cavity for High-frequency Axion Search



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### Superconducting Cavity for CAPP-MAX





## Summary:

### High-temperature Superconducting Cavities for Axion Search at CAPP

- Superconducting RF technology can enhance the scan rate of axion haloscope. >
- High-temperature superconductors (HTS) are one of the most promising superconductors  $\succ$ for realizing a high Q factor cavity in a high magnetic field.
- The Center for Axion and Precision Physics Research (CAPP) have successfully fabricated  $\succ$ 10 million Q factor cavity.
- Recently, CAPP developed large-scale HTS cavities for the CAPP-MAX experiment, aiming  $\succ$ to achieve sensitivity below DFSZ levels or even to detect 20% of axions as dark matter.
- Stay tuned! The physics run with the large-scale HTS cavity for CAPP-MAX will start soon.
- HTS cavities for high-mass axion search in R&D process.  $\succ$





### **Strategy 1**





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