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#### Solar axions



- **Primakoff conversion** of thermal photons, axion-photon coupling *g*<sub>*av*</sub>
- atomic recombination/deexcitation, Bremsstrahlung and Compton scattering (ABC), axion-electron coupling  $g_{ae}$  [Redondo, JCAP 12 (2013) 008]





[Dafni et al., PRD 99 (2019) 3, 035037]

• **nuclear transitions**, axion-nucleon coupling  $g_{aN}$ , e.g. 14.4 keV from <sup>57</sup>Fe M1 transition

## Helioscope concept

• **reconversion** of solar axions into **X-ray photons** in magnetic field [Sikivie, PRL 51 (1983) 1415-1417]



- conversion over macroscopic length, loss of coherence for axion masses above O(0.01) eV
- → buffer gas, refractive photon mass, restore coherence, access to larger axion masses [van Bibber et al., PRD 39 (1989) 2089]

### Helioscopes

#### **CAST** at CERN X-ray focusing optics, low-background techniques, buffer gas operation

[Anastassopoulos et al., Nature Phys. 13 (2017) 584-590]

#### Brookhaven

stationary magnet

[Lazarus et al., PRL 69 (1992) 2333-2336]



[Altenmüller et al., arXiv:2406.16840]





# International Axion Observatory (

X-ray detectors

- **next generation** helioscope, expand solar axion search by > 1 order of magnitude [Armengaud et al., JINST 9 (2014) T05002]
- > 10<sup>4</sup> signal-to-noise improvement (w.r.t. CAST)  $\rightarrow$
- **purpose-built magnet**, up to 5.4 T over more than 20 m, 8 conversion bores with 60 cm diameter
- high-efficiency X-ray focusing optics, focal spot smaller than 0.2 cm<sup>2</sup>
- high-efficiency ultra-low background • X-ray detectors, background less than 10<sup>-8</sup> cts/keV/cm<sup>2</sup>/s
- 50% sun tracking time





- **intermediate stage**, technological prototype
- → **inform IAXO** design, optimization, risk mitigation
- fully fledged helioscope, > 10<sup>2</sup> signalto-noise improvement (w.r.t. CAST)
- → relevant **physics results**
- **DESY** HERA south hall



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Baby **ACO** magnet

- common-coil **dipole**, racetrack configuration, counter-flowing currents
- → winding layout **similar to IAXO** toroid
- **two bores** with 70 cm diameter, **2-3 T** average magnetic field strength
- Al-stabilized **Nb-Ti-based** superconductor
- → cost-effective, but delays due to cable availability





# Baby **A** drive system

- CTA MST prototype parts, support frame holding magnet, optics, vacuum system and detectors, 90 t load
- 50% sun-tracking time, → < 0.01° pointing precision







Baby **X** X-ray optics

- profit from **space instruments**
- XMM Newton flight spare
  - $\circ$  7.5 m focal length
- custom hybrid telescope
  - **inner** optics based on NuSTAR/XRISM experience
  - **outer** optics from cold-slumped Willow-glass
  - 5 m focal length

70 cm

# Baby **XO** X-ray detectors

- sufficient energy resolution and threshold
- high X-ray **detection** efficiency
- $\rightarrow$ different detector technologies







GridPix

resolution

excellent **spatial** 

**Micromegas** 

- $O(10^{-7})$  cts/keV/cm<sup>2</sup>/s achieved, proven in CAST
- baseline technology  $\rightarrow$



cryogenic

- excellent energy resolution and threshold
- $\rightarrow$ axion spectroscopy

priority

metallic micro calorimeters (**MMC**) transition edge sensors (**TES**)

ultra-low **background**, background goal < 10<sup>-7</sup> cts/keV/cm<sup>2</sup>/s, **limited overburden** 



semiconductor

 $\Rightarrow$ 

# Silicon drift detectors

- tiny read-out electrode, low capacitance
- → good energy resolution, O(100) eV
- → low energy threshold, < 1 keV
- thin deadlayer, no entrance window needed [Mertens et al., J.Phys.G 48 (2020) 1, 015008]
- → high X-ray **detection efficiency**
- pure materials, little auxiliaries
- → great, yet unproven, potential for **low background** applications, O(10<sup>-6</sup>) cts/keV/cm<sup>2</sup>/s achieved
- → alternative technology to Micromegas





**passive-shield** SDD design

#### **active-shield** SDD-in-HPGe concept



### Physics case

- wide **unexplored ALP parameter space**,  $g_{ay}$  down to few 10<sup>-12</sup> GeV<sup>-1</sup>,  $g_{ae}$  down to few 10<sup>-13</sup>
- test **astrophysically hinted regions**, transparency, stellar cooling
- **QCD axions** in the meV to eV range, buffer gas phase
- → independent of dark-matter hypothesis



- **post-discovery** opportunities
  - **axion spectroscopy**, detection of Primakoff  $(g_{a\gamma})$  and ABC  $(g_{ae})$  flux to distinguish axion mass and models [Jaeckel, Thormaehlen, JCAP 03 (2019) 039]
  - **solar magnetometry**, sub-keV axions from longitudinal plasmons, information about magnetic field profile [O'Hare, PRD 102 (2020) 4, 043019]

#### Non-solar axions

• **supernova axions**, high energy photon detector, O(10) MeV, supernovascope [Ge et al., JCAP 11 (2020) 059]



**relic axions**, large magnetic field volume, haloscope setups, **RADES** [Alvarez Melcon et al., JCAP 05 (2018) 040]

resonant cavity insert

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- **exotic WISPs**, e.g. dark photons, chameleons
- high-frequency gravitational waves, inverse Gertsenshtein effect [Franciolini, PRD 106 (2022) 10, 103520]



### Conclusions

- solar axions searches able to **probe large ALP parameter space**
- **new result by CAST**,  $g_{a\gamma} < 5.7 \cdot 10^{-11}$  GeV<sup>-1</sup> (95% CL) for  $m_a \leq 0.02$  eV [Altenmüller et al., arXiv:2406.16840]
- **IAXO is next generation helioscope**, expand solar axion search by > 1 order of magnitude









