



KILONOVA CATCHER

# An overview of 2023-2024 activities towards 2025

2024 July 3<sup>rd</sup>



Université  
de Paris





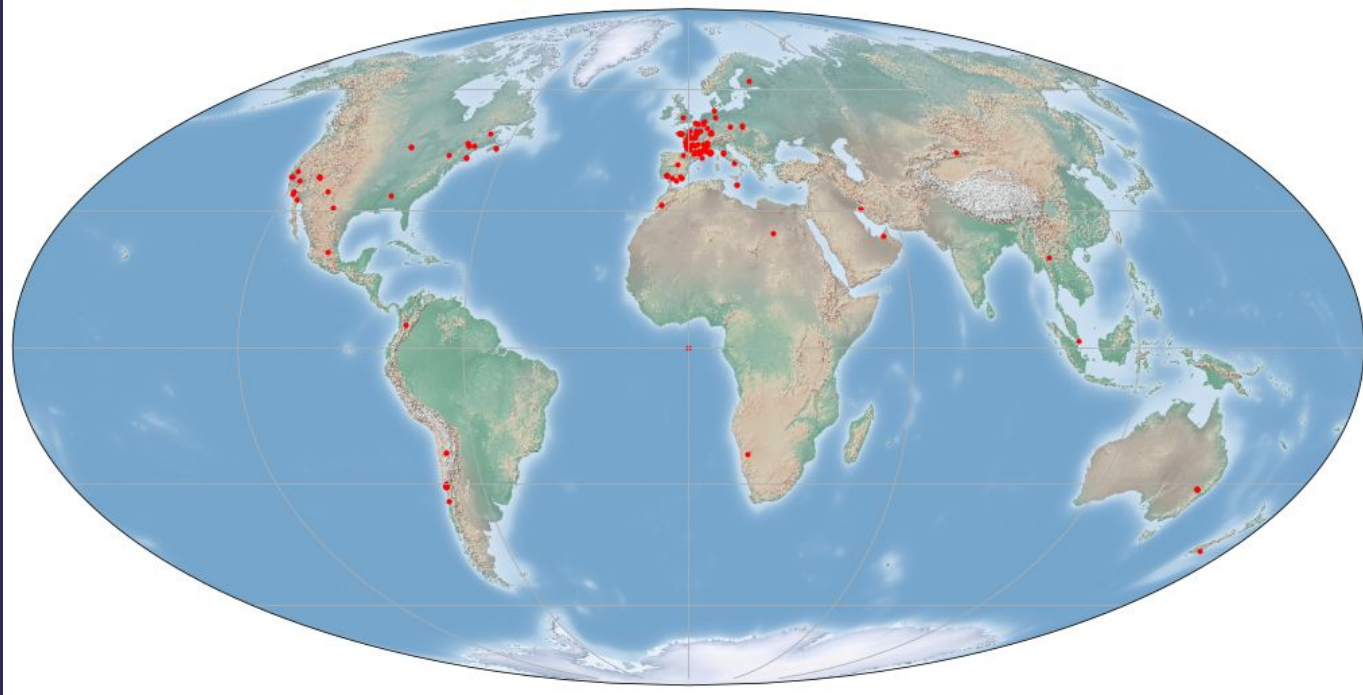
# OUTLINES

1. Overview of the 2023 year scientific activities
2. Overview of the 2024 year scientific activities
3. What are we missing to be (much) better

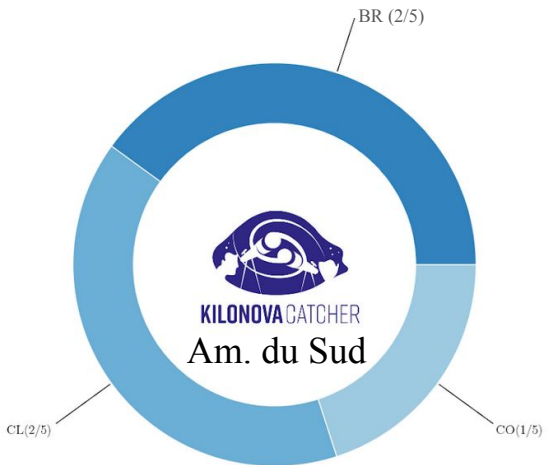
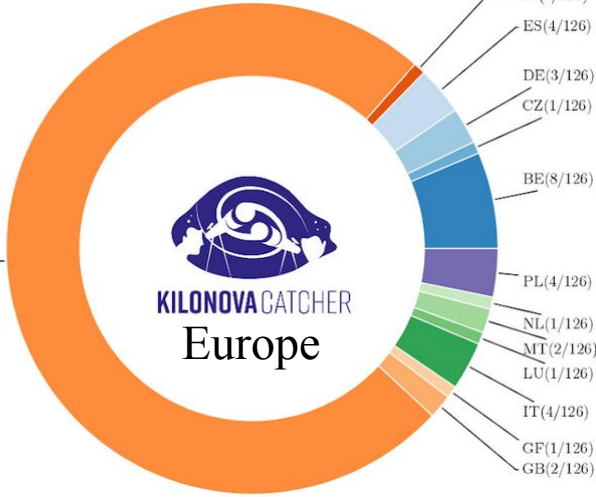
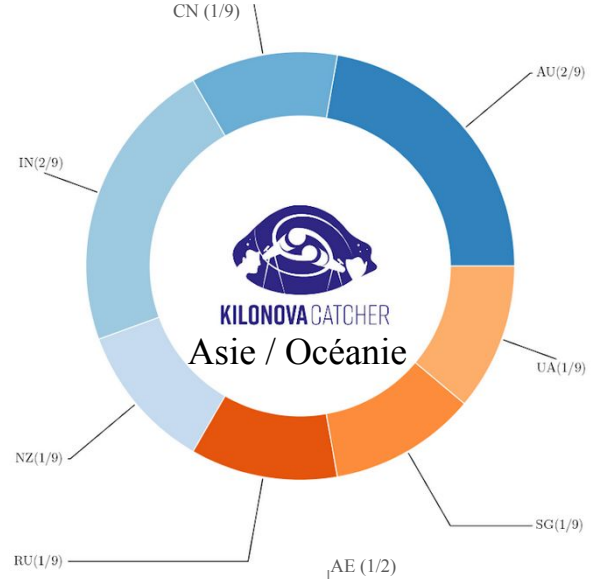


# The KNC network

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- 177 users
- 139 individual telescopes





# The KNC network connected to GRANDMA



MM alerts & Obs. plan

Follow-up images  
and  
photometric results



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**1**

Overview of the 2023 year scientific activities



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# The follow-up campaign that we engaged in 2023



Alert Name	Alert type	# images	filters	Publications
GRB 230328B	Gamma-ray burst	1	Rc	<a href="#">GCN 33550</a>
S230627c	Gravitational waves	60	L, B, Rc, TB, TG, TR	-
GRB 230812B	Gamma-ray burst	46	V, Rc, Ic, r, i	<a href="#">MNRAS paper</a>
ZTF23abnevza	Supernova	54	B, V, Rc, Ic	<a href="#">GCN 34928</a>
GRB 231115A	Gamma-ray burst	9	C, g, B, V, Rc, Ic	<a href="#">GCN 35051</a>
SN 2023wrk	Supernova	461	B,V,Rc,Ic, gri	paper to be submitted



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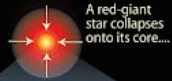
# The follow-up campaign that we engaged in 2023

## A collection of very diverse stellar cataclysmic events in the Universe

### Gamma-ray Bursts

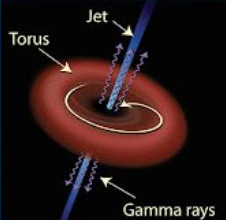
Gamma-Ray Bursts (GRBs): The Long and Short of It

**Long gamma-ray burst**  
( $>2$  seconds' duration)



A red-giant star collapses onto its core...

...becoming so dense that it expels its outer layers in a supernova explosion.



Jet  
Torus  
Gamma rays

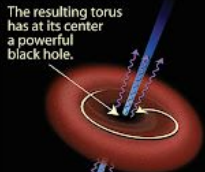
**Short gamma-ray burst**  
( $<2$  seconds' duration)



Stars\* in a compact binary system begin to spiral inward...



...eventually colliding.

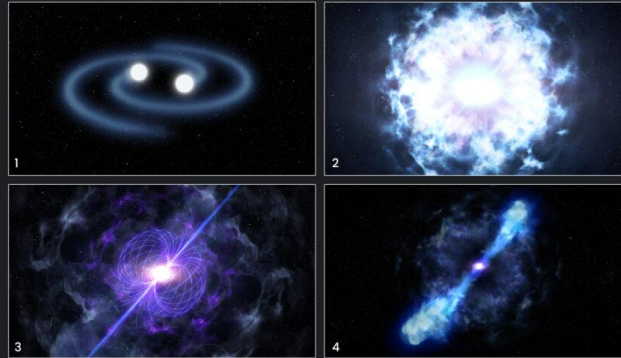


The resulting torus has at its center a powerful black hole.

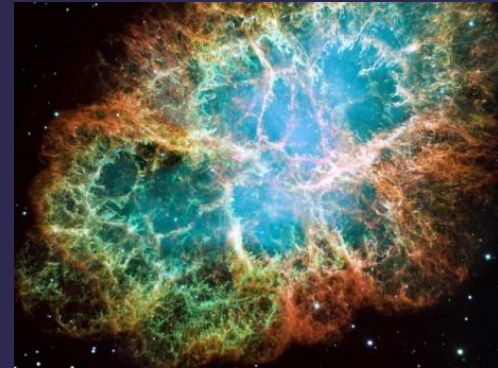
\*Possibly neutron stars.

### Optical counterparts from the mergers of compact objects (Gravitational waves)

Collision Sequence for a Magnetar-Powered Kilonova Blast



### Supernovae







# The follow-up campaign that we engaged in 2023

## GRB 230328B (a long GRB with an early detection of the optical afterglow):

- 1 image provided (F. Kugel)
- We detected the GRB optical afterglow!

- A GRANDMA GCN was emitted: <https://gcn.nasa.gov/circulars/33550>
- No more studies about this GRB because of a lack of multi wavelength data from the GRANDMA side

**Subject** GRB 230328B : GRANDMA/Kilonova-Catcher optical detection  
**Date** 2023-03-31T20:37:22Z (9 months ago)  
**From** Damien Turpin at NAOC (CAS) <dturpin-astro@hotmail.com>

F. Kugel (KNC), D. Turpin (CEA), S. Karpov (FZU), T. Hussenot-Desenonges (IJCLab), S. Antier (OCA/Artemis), P.A. Duverne (APC, U. Paris Cit ), P. Hello (IJCLab), A. Klotz (OMP/IRAP) report on behalf of the GRANDMA/Kilonova-Catcher collaboration:

The Kilonova-Catcher telescope network responded to the alert of GRB 230328B (Swift detection: Gropp et al., GCN [33527](#); Fermi GBM detection: Veres et al., GCN [33526](#)).

The KNC observations were taken by F. Kugel with an ARTEMIS CCD ATIK-460ex camera mounted in the 0.4-m f/2.8 reflector telescope in the Chante-Perdrix Observatory (France). The afterglow is detected in the 36x60s unfiltered coadded images at about 5.2 hours (midtime of the exposure) after the Swift/BAT trigger time.

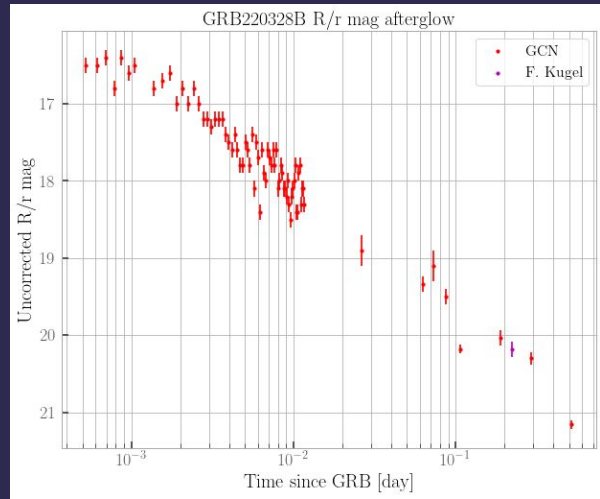
Below, we report our photometric measurement in Rc magnitude.

T-T0 (day) | Exposure | Filter | Mag +/- err | Mag.Lim. (AB)  
-----  
0.22172 | 36 x 60s | Rc | 20.19 +/- 0.10 | 21.4 (3 sigma)

Our detection is consistent with the detections and limits previously reported in Pankov et al., GCN [33528](#); Belkin, GCN [33530](#); Lu et al., GCN [33534](#); Catapano et al., GCN [33535](#); Suresh et al., GCN [33536](#); Adami et al., GCN [33537](#); Gompertz et al., GCN [33538](#); Komesh et al., GCN [33539](#); Lu et al., GCN [33540](#); Ror et al., GCN [33547](#); Agui Fernandez et al., GCN [33549](#)

The GRANDMA/Kilonova-Catcher images have been calibrated using field stars from the PanSTARRS-DR1 catalog using the STDpipe pipeline (Karpov 2022) and the ps1/r to Rc mag conversion from (Pancino et al. 2022).

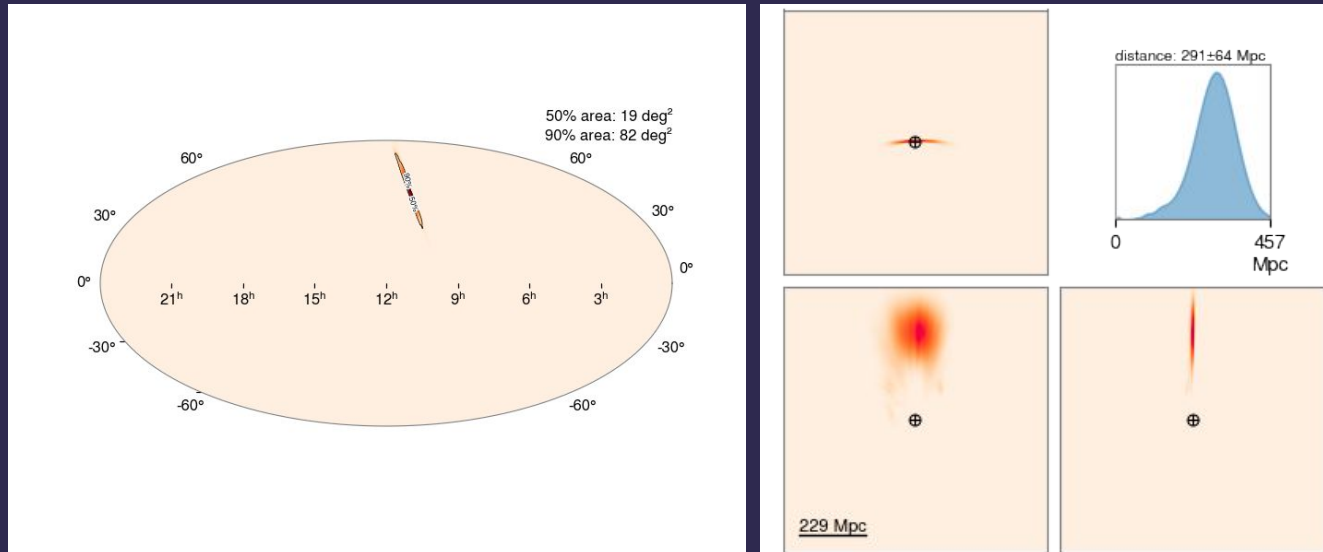
GRANDMA is a worldwide telescope network ([grandma.ijclab.in2p3.fr](http://grandma.ijclab.in2p3.fr)) devoted to the observation of transients in the context of multi-messenger astrophysics (Antier et al. 2020 MNRAS 497, 5518). Kilonova-Catcher (KNC) is the citizen science program of GRANDMA (<http://kilonovacatcher.in2p3.fr/>).



# The follow-up campaign that we engaged in 2023

**S230627c (GW NSBH merger 48% / loc. area = 82 deg<sup>2</sup>):**

- Follow-up of two ZTF transient candidate ZTF23aaptudb (O. Aguerre) & ZTF23aaptusa (A. Popowicz, M. Freeberg, D. ST-Gelais)
- Sadly, nothing interesting found in association with the GW NSBH event



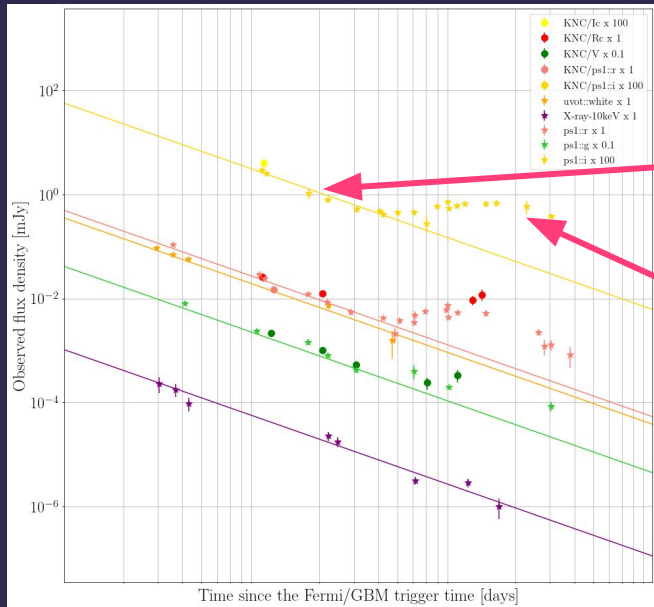


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# The follow-up campaign that we engaged in 2023

**GRB 230812B (Long GRB from the collapse of a massive star, the 3rd brightest GRB ever observed, having multi-wavelength detections:**

- **46 images provided (M. Serrau, S. Leonini, M. Freeberg, F. Romanov, M. Odeh, J. Nicolas, L. Rousselot)**
- **We detected the GRB optical afterglow in 10 images (from the stacking of single images)**



Emission from electrons accelerated in the shock front between the (decelerating) GRB relativistic jet and the interstellar medium

emergence of the Supernova type Ic emission at late times



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# The follow-up campaign that we engaged in 2023

**GRB 230812B (Long GRB from the collapse of a massive star, the 3rd brightest GRB ever observed with multi-wavelength detections:**



**Best fit light curve with a GRB afterglow and SN Ic model**

**Hussenot-Desenonges, T. et al 2024 (MNRAS)**

**<https://ui.adsabs.harvard.edu/abs/2023arXiv231014310H/abstract>**

**co-authors: M. Serrau, S. Leonini, M. Freeberg, L. Rousselot, M. Odeh, J. Nicolas**

- The spectral analysis of the x-ray to optical emission shows the explosion site of the GRB in its host galaxy



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# The follow-up campaign that we engaged in 2023

**ZTF23abnevza / SN2023vuc (was possibly associated with the Fermi/GBM long GRB GRB 231024A but the association was discarded then):**

- **54 images provided (M. Freeberg, M. Odeh, A. Popowicz) under poor conditions (bright moon)**
- **We detected the GRB optical afterglow in the M. Freeberg images**

**Subject** GRB 231024A: GRANDMA observations of [A2023vuc](#)/GOTO23baj candidate  
**Date** 2023-11-02T13:35:29Z (7 months ago)  
**From** Dilya Akl at American Uni. SHJ <dalyaakl.d@gmail.com>  
**Via** Web form

D. Akl (AUS), D. Turpin (CEA-Saclay/Irfu), P. Gokuldass (ERAU), R. Strausbaugh (EUI), Z. Vidadi (SHAO), S. Antier, A. de Ugarte Postigo (OCA/Artemis), M. Coughlin (UMN), Q. Andre (OCA), J. Peloton, P. Hello (JCLAB), I. Tosta e Melo (UniCT-DFA), T. Pradier (Unistra/IPHC), S. Karpov, M. Prouza, M. Mašek, M. Blazek (FZJ), A. Klotz (IRAP), A. Takey, M. Ismail, M. Abdelkareem, M. Molham (NRIAG), L. Almeida, L. Fraga, W. Corradi, N. Sasaki (LNA), F. Navarete (NOIRLab/SOAR), M. Freeberg (KNC) on behalf of the GRANDMA and Kilonova-Catcher collaborations:

The GRANDMA and Kilonova-Catcher telescope networks observed the source GOTO23baj/SN2023vuc (Gompertz et al., GCN [34878](#), Iglesias-López et al, GCN [34881](#)), finally classified as a young supernova (Saccardi et al, GCN [34882](#)). We used Skyportal (Coughlin et al., 2023) for the coordination of our observations.

Observations were conducted from 2023-10-25 17:15:51 to 2023-11-01 03:00:12, ~0.91-7.15 days after the GOTO detection time (midtime ref = 2023-10-24T23:23:27.50 from Gompertz et al., GCN [34878](#)).

In the following table, we report a subset of the preliminary photometry of our observations. Magnitudes are reported in the AB and Vega system depending on the filter set.

T-T0 day MJD	Observed	Exposure	Filter	Mag + -err
0.91	60242.886	Les Makes	50x120s	r 18.56+/-0.19 (AB)
1.04	60243.019	FRAM-CTA-N	76x120s	V 18.51 U.L (Vega)
5.32	60247.290	KNC-T11	17x180s	Rc 18.03+/-0.09 (Vega)
5.92	60247.889	KAO	11x180s	r 18.27+/-0.01 (AB)
5.96	60247.932	KAO	11x180s	i 18.51+/-0.02 (AB)
6.23	60248.206	KNC-T11	17x180s	Rc 18.17+/-0.08 (Vega)
6.81	60248.793	KAO	15x180s	i 18.46+/-0.01 (AB)
6.85	60248.829	KAO	14x180s	r 18.42+/-0.01 (AB)
7.13	60249.105	SOAR	5x75s	r 18.33+/-0.01 (AB)
7.14	60249.111	SOAR	3x150s	i 18.37+/-0.01 (AB)
7.14	60249.117	SOAR	3x200s	g 18.03+/-0.01 (AB)
7.15	60249.125	SOAR	2x120s	z 18.04+/-0.05 (AB)

Our images were taken under poor conditions and a bright moon. All the data in Sloan filters have been calibrated with respect to the PS1 catalog. Johnson filters have been measured using the Gaia catalog. All the data have been reduced by a single data processing pipeline STPIPE (Karpov et al., 2022). Please contact us for more information and collaboration on this source.

- **A GRANDMA GCN was emitted: <https://gcn.nasa.gov/circulars/34928>**
- **No more studies about this SN finally classified as a young type II (Core Collapse Supernova)**



# The follow-up campaign that we engaged in 2023

## GRB 231115A (a short GRB - so possibly associated with a GW signal detected in M82!!!):

- 9 images provided (M. Freeberg, M. Serrau, E. Broens)
- We **did not** detect the GRB optical afterglow

**Subject** GRB231115A: GRANDMA Observations  
**Date** 2023-11-15T23:22:01Z (2 months ago)  
**From** Cristina Andrade at UMN <andra104@umn.edu>  
**Via** Web form

A. Iskandar (XAO), F. Wang (THU/BJP), J. Zhu (BJP), L. Wang, X. Zeng, C. Andrade (UMN), A. de Ugarte Postigo (CNRS/OCA), D. Akl (AUS), E. Broens (KNC), S. Antier (OCA-Artemis), I. Tosta e Melo (UniCT-DFA), P. Hello (IJCLAB), D. Turpin (CEA-Saclay/Irfu), T. Pradier (Unistra/IPHC), M. Coughlin (UMN), S. Karpov (FZU), J. Peloton (IJCLab) report on behalf of the GRANDMA collaboration:

We observed the field of GRB 231115A (Fermi GBM team, GCN [35035](#)) covering the complete INTEGRAL error box (D'Avanzo et al. GCN [35036](#); Mereghetti et al. GCN [35037](#)) within the GRANDMA collaboration. Imaging with the 0.4m SNOVA telescope did not find any candidate in r-band around 2023-11-15 17:37:53 (e.g. 2h after the trigger time) down to an upper limit of 18.9 (5-sigma threshold) or 19.3 (3-sigma threshold) using PS1 catalog as photometric comparison. We also looked carefully at the location of [AT\\_2023xvj](#) (Kumar et al. GCN [35041](#)).

The amateur contribution to GRANDMA, Kilonova Catcher (KNC), made no detection with a 30x180s image using a clear filter on 2023-11-15T20:00 UTC (TGRB + 3.28h). We determine a detection limit of 20 mag in r-band, using PS1 for calibration and color term correction. At 2023-11-15T18:53:25.219, we obtained R>18 from 5x180s exposure. The upper limit is given at 5-sigma averaged over all the images.

These upper limits are consistent with previous reports by MASTER (Lipunov et al. GCN [35046](#)).

GRANDMA is a worldwide telescope network ([grandma.ijclab.in2p3.fr](#)) devoted to the observation of transients in the context of multi-messenger astrophysics (Antier et al. 2020 MNRAS 497, 5518). Kilonova-Catcher (KNC) is the citizen science program of GRANDMA ([http://kilonovacatcher.in2p3.fr/@](#)).

- A GRANDMA GCN was emitted: <https://gcn.nasa.gov/circulars/35051>
- Actually this was not a GRB but the first Giant flare from a Magnetar observed in M 82, see Mereghetti et al. 2023 <https://arxiv.org/abs/2312.14645>



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# The follow-up campaign that we engaged in 2023

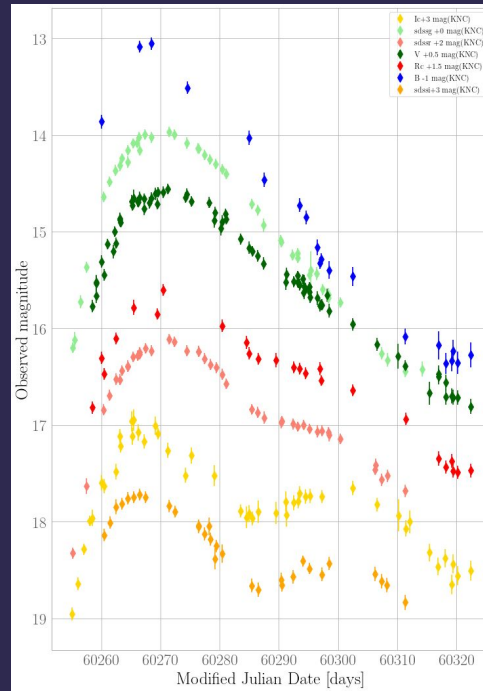
## SN 2023wrk (a SN type Ia classified as a subclass SN 99aa-like):

- 461 images provided, it's just amazing (M. Freeberg, M. Serrau, E. Broens, D. Marchais, R. Ménard, F. Dubois, C. Galdies, D. ST-Gelais, A. Popowicz, M. Odeh, G. Parent)
- We detect the SN in all the images.

SN type  
**Ia**

Origin

Thermonuclear runaway in  
a C-O white dwarf exceeding  
the Chandrasekhar mass  
limit after matter accretion  
or through a DD merger  
channel







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# The follow-up campaign that we engaged in 2023

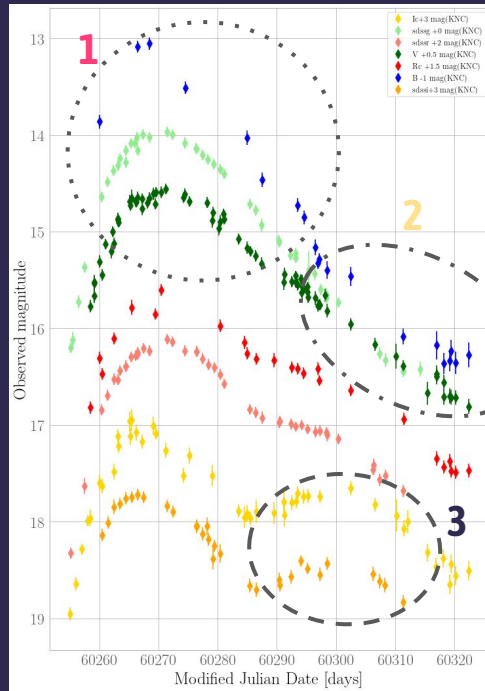
## SN 2023wrk (a SN type Ia classified as a subclass SN 99aa-like):

- 472 images provided, it's just amazing (M. Freeberg, M. Serrau, E. Broens, D. Marchais, R. Ménard, F. Dubois, C. Galdies, D. ST-Gelais, A. Popowicz, M. Odeh, G. Parent)
- We detect the SN in all the images.

SN type  
**Ia**

Origin

Thermonuclear runaway in a C-O white dwarf exceeding the Chandrasekhar mass limit after matter accretion



## A typical SNIa light curve powered

1. by the radiating external envelope of the star and heated by the radioactive decay of freshly formed  $^{56}\text{Ni}$
2. by the radiating inner part of the ejecta heated by the radioactive decay of  $^{56}\text{Co}$
3. by the recombination process of newly formed and ionized  $^{56}\text{Fe}$  into stable  $^{56}\text{Fe}$  resulting in NIR and IR radiation





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The GW 04 run activity in 2023



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**2**

**Overview of the 2024 year scientific activities**



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# The follow-up campaign that we engaged in 2024



Alert Name	Alert type	# images	filters	Publications
EP240331A	??	2	Rc	<a href="#">GCN 36010</a>
IC240327B	HEN	21	L,B,V,R,g,r	<a href="#">GCN 36050</a>
S240422ed	Gravitational waves	46	L,,Rc,V,g,r	<a href="#">GCN 36299</a>
EP240618a	??	8	G,V,Rc,Ic,g,r	<a href="#">GCN 36712</a>
EP240625a	??	5	G,B,V,Rc	<a href="#">GCN 36771</a>
EP240702A	??	1	V	Not planned



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**3**

What are we missing to be (much) better

Alert transmission to the KNC  
community  
for non GW alerts

Quick data (image) availability  
(through STDweb?)



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Working environment  
KNC website and server

Communication



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# Software development plans

1. Renovate the coding architecture during the Summer to include much more **automatization** everywhere
2. We should propose a more diverse set of MM transient alerts (EP, SVOM, KM3NeT, etc.) and we could even think about using KNC for your research besides GRANDMA BUT KNC data are GRANDMA (keep this in mind). Need
3. **Automatize** even more the data analysis of their images (all the building blocks are there...., we may think about connecting STDweb to KNC also) to produce quick light curve (I would need an intern in python code for doing this)
4. Refreshing the KNC web interface (bug fixes at least): but a big patch of the web interface is likely not feasible in the second semester of 2024

**Any other idea?**



# Some people who have expressed interests in KNC project

**I did not forget you, I need to find time to set up plans and actions**

- **PA Duverne:** image analysis and tuning of the KNC image analysis pipeline (STDPipe)
- **JG Ducoin:** automatization of the KNC image analysis pipeline (Quicklook) + new follow-up proposal with KM3NeT alerts
- **N Guessoum:** Communication and networking
- **C Andrade:** Communication ?