

Dark photon and scalar field for tabletop experiments

The axion's underdogs

Etienne SAVALLE

DPHP, CEA Saclay

30 Septembre 2024

1 Dark Matter

2 Dark photon

- Properties
- Resonant experiments
- Broadband experiments

3 Scalar field

- Properties
- Colocated
- Space-time separated

Types of particles Occupation number (cf slide 24 Pierre's talk) :

$$n \propto \frac{h^3 \rho_{DM}}{m_{DM}^4 v_{max}^3} \simeq 10^{29} \left(\frac{\mu eV}{m} \right)^4 \gg 1$$

Pseudo-scalar

Vector

Scalar



Go back in time 40 minutes ago



Underdog n°1



Underdog n°2

1 Dark Matter

2 Dark photon

- Properties
- Resonant experiments
- Broadband experiments

3 Scalar field

- Properties
- Colocated
- Space-time separated



What is it ?

Hypothetical gauge boson, similar to the regular photon but associated with a hidden/dark sector.

Interaction

Kinetic mixing

Mass

Not necessarily massless as the photon

Detections strategies

Direct and indirect

Axion Lagrangian:

$$\mathcal{L}_{\text{axion}} = \frac{1}{2}(\partial_\mu\phi)(\partial^\mu\phi) - \frac{1}{2}m_\phi^2\phi^2 - \frac{g_{\phi\gamma}}{4}\phi F_{\mu\nu}\tilde{F}^{\mu\nu} + \dots$$

Dark Photon Lagrangian:

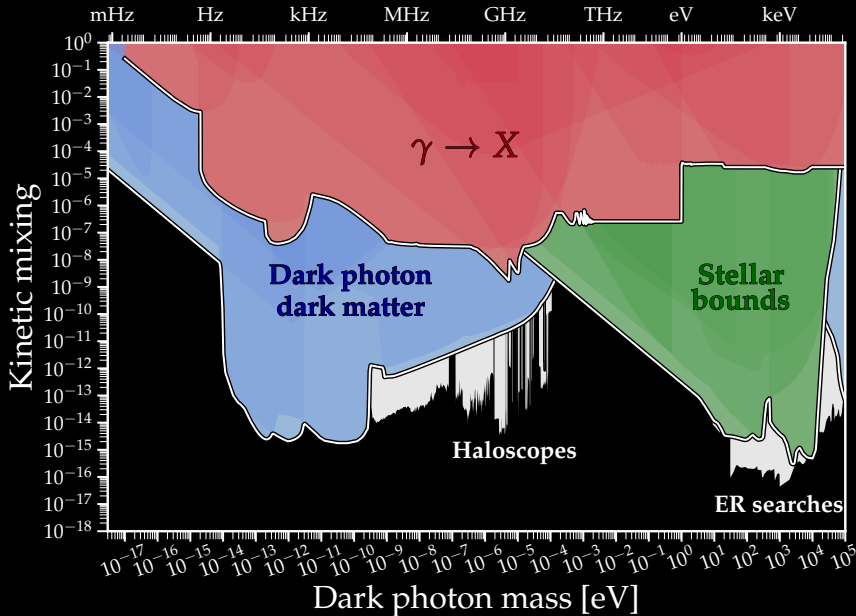
$$\mathcal{L}_{\text{dark photon}} = -\frac{1}{4}\phi_{\mu\nu}\phi^{\mu\nu} + \frac{1}{2}m_{\gamma'}^2\phi_\mu\phi^\mu + \frac{\chi}{2}F_{\mu\nu}\phi^{\mu\nu}$$

Massless

- Photon of the hidden sector.
- Very weak kinetic mixing.
- Not a DM candidate.

Massive

- Different behavior at low/high energies.
- Kinetic mixing with photons.
- DM candidate.



Caputo et al., PRD 2021

Dark Photon Lagrangian:

$$\mathcal{L}_{\text{dark photon}} = -\frac{1}{4}\phi_{\mu\nu}\phi^{\mu\nu} + \frac{1}{2}m_{\gamma'}^2\phi_\mu\phi^\mu + \frac{\chi}{2}F_{\mu\nu}\phi^{\mu\nu}$$

Kinetic shift

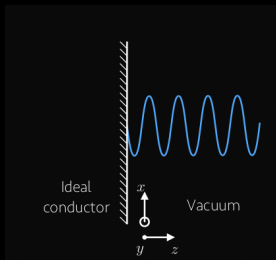
$$\phi^\mu \rightarrow \tilde{\phi}^\mu - \chi A^\mu$$

Modified Maxwell's equation

$$\vec{\nabla} \times \vec{H} - \dot{\vec{D}} = \chi\epsilon_0 \left(\dot{\vec{E}}_{DM} - c^2\vec{\nabla} \times \vec{B}_{DM} \right)$$

Additional electric field in vacuum

$$\vec{E}_{DM} = -\partial_t \vec{A} = \chi\omega_{\gamma'}\vec{\phi}_{DM} = i\chi\omega_{\gamma'}\phi_0 e^{i\omega_{\gamma'}t - ik_{\gamma'}\cdot\vec{x}}\vec{e}$$



Dark photon in a mirror: Conversion to an

"usual" electric field :

$$\vec{E}_{\text{out}} = -i\omega_{\gamma'}\phi_0\chi e^{-i\omega_{\gamma'}t - ik_{\gamma'}x}\vec{e}$$

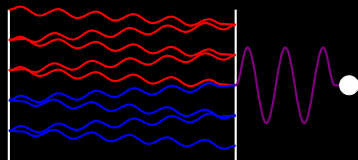
Thanks again Pierre !

Output Power: $\frac{\Pi}{\text{W/m}^2} = 1.44 \times 10^{-20} \left(\frac{\rho_{\text{DM}}}{\text{GeV/cm}^3} \right) \left(\frac{\chi}{10^{-12}} \right)^{-2}$

Emission: Normal to surface

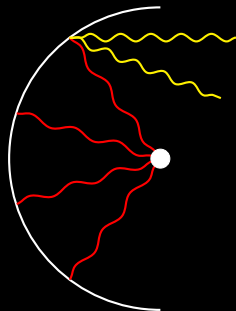
Resonant Search

- Narrow mass range
- High sensitivity
- Tuned detectors



Broadband Search

- Wide mass range
- Moderate sensitivity
- Broad scan



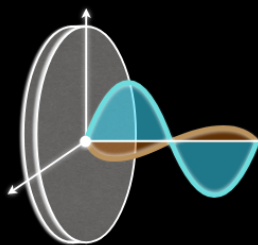
Horns et al., JCAP 2013

Polarization: Big difference with the axion

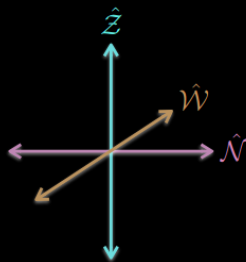
Axial experiment
(Zenith-pointing)



Planar experiment
(North-facing)



Possible DP Polarisations



Caputo et al., PRD 2021

Earth revolution creates a modulation

Sensitivity : Planar > Axial and measurement timing optimizable

$$\text{SNR}(f) \sim \frac{\Pi S Q \alpha}{k_B (T_{\text{sys}} + T_{\text{ext}})} \sqrt{\frac{\tau}{\Delta f}} \delta(m - m_\varphi)$$

m_φ Monochromatic signal at the dark photon mass*

- Increase sensitivity with

S High surface area for broadband

Q High quality factor for resonant

α Optimally paced measurement method for polarization

T_{ext} External noise is already attenuated

- Out of focus for broadband
- Not enhanced by resonance

T_{sys} Limiting noise due to the measurement electronics temperature.

$\tau, \Delta f$ Can be tuned with duration and resolution

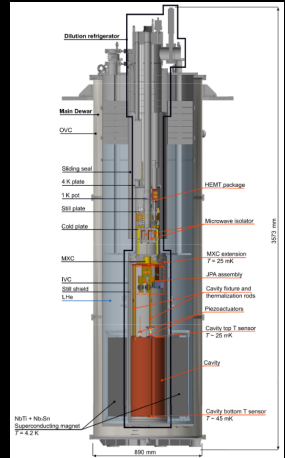
Center for Axion and Precision Physics :

- Resonant cavity : 1, 1.8, 2.5, 5 GHz
- Bandwidth : 150 MHz
- System noise : 0.4 K
- Sensitivity at the 10^{-14} level

Ahn et al., PRX 2024

Other resonant experiments :

- Orpheus paper by Cervantes et al., PRD 2022
- Haystack paper by Backes et al., Nature 2021
- QUAX-LNF paper by Rettaroli et al., PRD 2024
- ORGAN paper by Mc Allister et al., ADP 2023

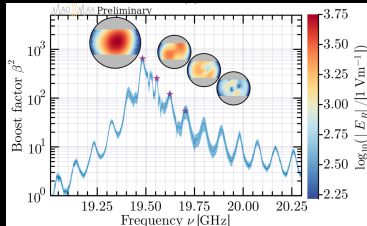
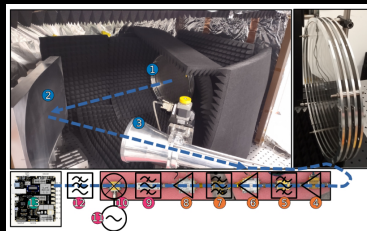


Magnetized Disc and Mirror Axion eXperiment:

- Broadband antenna
- Bandwidth : 19 – 20.4 GHz
- System noise : 240 K
- Sensitivity at the 10^{-12} level

Special features:

- Tunable resonance creating a sensitivity boost
- Experimental sensitivity measurement



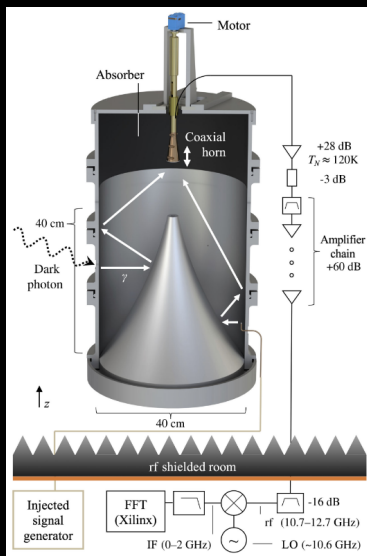
Egge et al., arXiv 2024

Broadband Reflector Experiment for Axion Detection

- Broadband antenna
- Bandwidth : 10.7 – 12.5 GHz
- System noise : 120 K
- Sensitivity at the 10^{-12} level

Special features:

- Hershey's kiss reflector



Knirck et al., PRL 2024

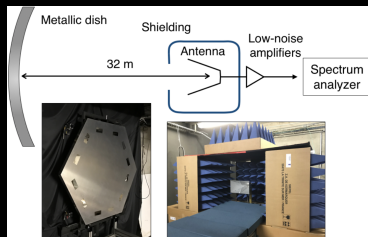
Search for U(1) dark matter with an Electromagnetic Telescope

- Broadband antenna
- Bandwidth : 5 – 7 GHz
- System noise : 554 K
- Sensitivity at the 10^{-10} level*

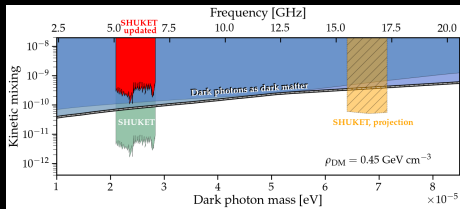
Special features:

- Lowest cost per PRL publication*

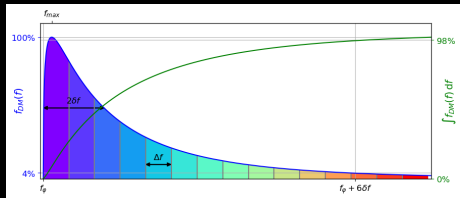
Brun et al., PRL 2019



Sensitivity: Decrease due to diffraction and mismatch in the mode antenna Gué et al., PRD 2024



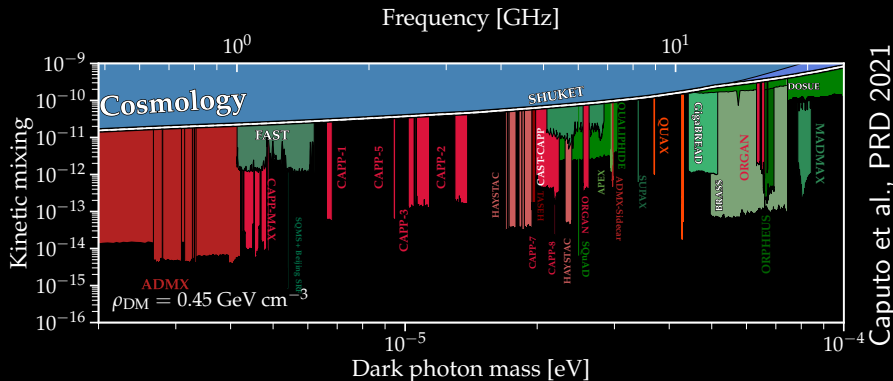
Data analysis: Improved software searching for non monochromatic shape



Upgrade: Improved electronic

$$T_{sys} \searrow \text{ and } \tau \nearrow$$

Stay tuned for revised constraints thanks to Jordan Gué's PHD work and Robin Signoret's future PHD work.



- Good training before an axion search
- Still some mass range to cover

1 Dark Matter

2 Dark photon

- Properties
- Resonant experiments
- Broadband experiments

3 Scalar field

- Properties
- Colocated
- Space-time separated

BASE **Scalar field** **PV 130**



Pokémon basique pré-dessiné par Gilles de la Motte - Fonds de la recherche en santé Québec

Vacuum fluctuations **60**
 Shuffle your hand into your deck and draw 5 new cards. Also, remove any status effect from Scalar Field.

Symmetry break **90**
 If the opponent's Pokémon has more HP than Scalar Field deal an additional 40 damage.

Faiblesse **Résistance**
 +20 -10

Retraite

Axion bends light and manipulates energy fields, making it nearly invisible.

©2012 Vevep.fr Illus. Sébastien

What is it ?

A field represented by a scalar value at every point in space and time.

Interaction

Couples to other fields via Yukawa interactions or portal interactions.

Mass

From ultra-light to very heavy.

Detection strategies

Cosmic background, gravitational effects, precision experiments, and colliders.

Scalar field theory action

The theory relies on an action where φ is the massive scalar field :

$$S = \int d^4x \frac{\sqrt{-g}}{c} \frac{c^4}{16\pi G} \underbrace{[R - 2g^{\mu\nu} \partial_\mu \varphi \partial_\nu \varphi - V(\varphi)]}_{\text{GENERAL RELATIVITY + SCALAR FIELD}}$$

$$+ \int d^4x \frac{\sqrt{-g}}{c} \underbrace{[\mathcal{L}_{SM}[g_{\mu\nu}, \Psi_i] + \mathcal{L}_{int}[g_{\mu\nu}, \varphi, \Psi_i]]}_{\text{STANDARD MODEL + SCALAR FIELD}}$$

General relativity part

- 1 Spatial dependence
- 2 Time dependence

Lagrangian part

- 1 Coupling SM-DM
- 2 Constants modulation

Lagrangien part

$$\mathcal{L}_{int}[g_{\mu\nu}, \varphi, \Psi] = \frac{\varphi^i}{i} \left[\frac{d_e^{(i)}}{4e^2} F_{\mu\nu} F^{\mu\nu} - \frac{d_g^{(i)} \beta_3}{2g_3} F_{\mu\nu}^A F_A^{\mu\nu} - \sum_{i=e,u,d} \left(d_{m_i}^{(i)} + \gamma_{m_i} d_g^{(i)} \right) m_i \bar{\psi}_i \bar{\psi}_i \right]$$

Damour and Donoghue, PRD 2010

i = 1 Linear coupling

i = 2 Quadratic coupling

Coupling parameter

d_e Fine structure constant

d_{m_e} Electron mass

$d_{m_{u,d}}$ Quark mass

d_g Lambda QCD scale

General relativity action part

$$S = \int d^4x \frac{\sqrt{-g}}{c} \frac{c^4}{16\pi G} \underbrace{[R - 2g^{\mu\nu} \partial_\mu \varphi \partial_\nu \varphi - V(\varphi)]}_{\text{GENERAL RELATIVITY + SCALAR FIELD}}$$

How φ varies with space-time ? Applying the least action principle we get the following scalar field equation :

Linear coupling

$$\square \varphi + \left(\frac{m_\varphi c^2}{\hbar} \right)^2 \varphi = - \frac{4\pi G}{c^2} \alpha_A \rho_A$$

Quadratic coupling

$$\square \varphi + \left(\frac{m_\varphi c^2}{\hbar} + \frac{4\pi G}{c^2} \alpha_A \rho_A \right)^2 \varphi = 0$$

Linear coupling

Classical phenomenology

$$\varphi(t, \vec{r}) = \varphi_0 \sin(\omega_\varphi t - \vec{k}_\varphi \cdot \vec{r}) - s_A \frac{GM_A}{c^2 r} e^{-r/\lambda_\varphi}$$

Phenomenology

Time oscillation

Spatial dependence

Quadratic coupling

Richer phenomenology

$$\varphi(t, \vec{r}) = \varphi(r) \varphi_0 \sin(\omega_\varphi t - \vec{k}_\varphi \cdot \vec{r})$$

Phenomenology

Screening

Enhancement

Energy-impulsion tensor and scalar field energy density

$$T^{\mu\nu}(\varphi) = \frac{2}{\sqrt{-g}} \frac{\delta \mathcal{L}}{\delta g_{\mu\nu}} \Rightarrow \langle \rho_\varphi \rangle = \frac{T_{00}}{c^2} = \left\langle \frac{1}{2\kappa c^2} \dot{\varphi}^2 + \frac{1}{2\kappa c^2} \omega_\varphi^2 \varphi^2 \right\rangle$$

Energy density

$$\langle \rho_\varphi \rangle = \rho_{DM} \Rightarrow \varphi_0 = \frac{\sqrt{8\pi G \rho_{DM}}}{\omega_\varphi c}$$

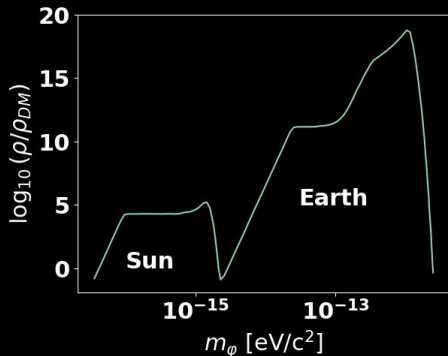
Density



Galactic halo



Earth halo



A. Banerjee et al., Nature Communications, (2020)

Fine structure constant variation

When considering only the electromagnetic effect, the effective lagrangien $\mathcal{L}_{int} + \mathcal{L}_{SM}$ leads to variation of the fine structure constant :

$$\mathcal{L}_{eff}^{EM} = \underbrace{-\frac{e^2 c}{16\pi\hbar\alpha} F^2}_{\text{ELECTROMAGNETISM FROM STANDARD MODEL}} + \underbrace{d_e \varphi \frac{e^2 c}{16\pi\hbar\alpha} F^2}_{\text{ELECTROMAGNETISM FROM SCALAR FIELD}} \simeq \frac{-e^2 c}{16\pi\hbar\alpha (1 + d_e \varphi)} F^2$$

Variation of the fine structure constant

$$\alpha(t) = \alpha \left(1 + d_e \sqrt{\frac{8\pi G \rho_{DM}}{\omega_\varphi c^2}} \cos(\omega_\varphi t) \right)$$

Other constants

$$m_j(\varphi) = m_j \left(1 + d_{m_j}^{(i)} \varphi^i \right)$$

for $j = e, u, d$

$$\Lambda_3(\varphi) = \Lambda_3 \left(1 + d_g^{(i)} \varphi^i \right)$$

Colocated clocks

Comparison of different clocks at the same space and time :

$$\frac{\delta(\nu_A/\nu_C)}{(\nu_A/\nu_C)_0} = \left(d_e + (d_{m_e} - d_g) \right) \varphi$$

Dimensionless

$$m_x/\Lambda_3, \Rightarrow d_{m_x} - d_g.$$

vs

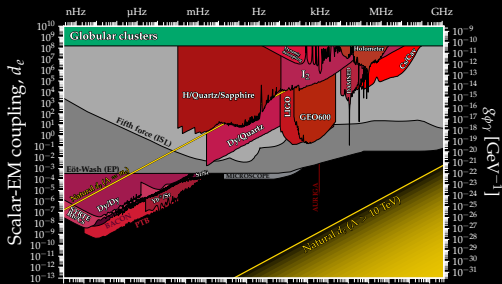
Space-time separated clocks

Comparison of the same clocks at different space and/or time :

$$\frac{\delta(\nu_{A_1}/\nu_{A_2})}{(\nu_{A_1}/\nu_{A_2})_0} = (2d_e + d_{m_e}) \varphi$$

Dimensional

$$m_x \Rightarrow d_{m_x}.$$



Universality of free fall

Eot-Wash : Wagner et al., CQG 2012
 MICROSCOPE : Bergé et al., PRL 2018

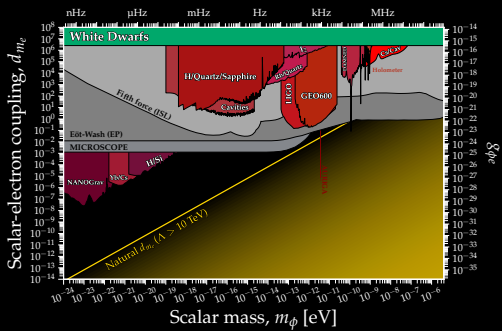
Colocated

SYRTE : Hees et al. PRL 2016
 JILA : Kennedy et al., PRL 2020
 BACON : Bely et al., Nature, 2021
 PTB : Filzinger et al., PRL 2023

Space-time separated

DAMNED : Savalle et al., PRL 2021
 GEO600 : Vermeulen et al., Nature 2021
 LIGO : Göttel et al., PRL 2024

Caputo et al., PRD 2021

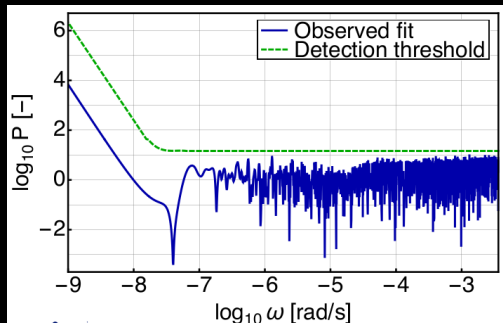


Clocks comparison:

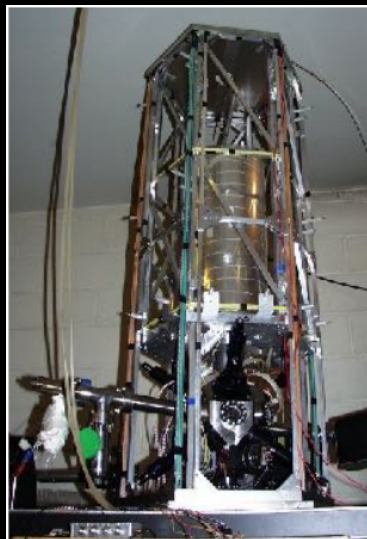
Searching for a periodic signal in two clocks desynchronization

Sensitivity:

$$\frac{\delta(\nu_A/\nu_C)}{(\nu_A/\nu_C)_0} = d_e \varphi + \dots$$



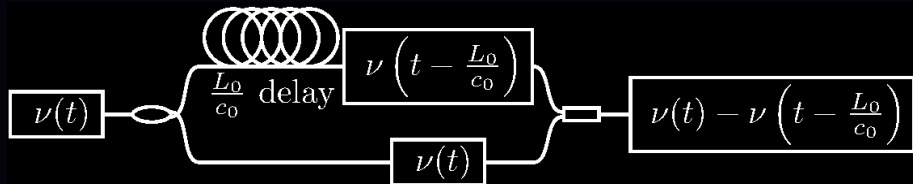
Hees et al., PRL 2016



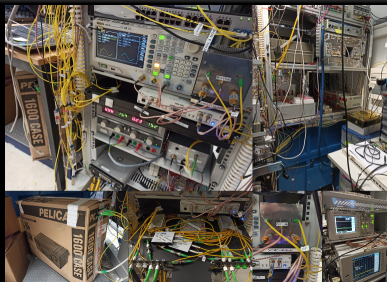
SYRTE atomic clock

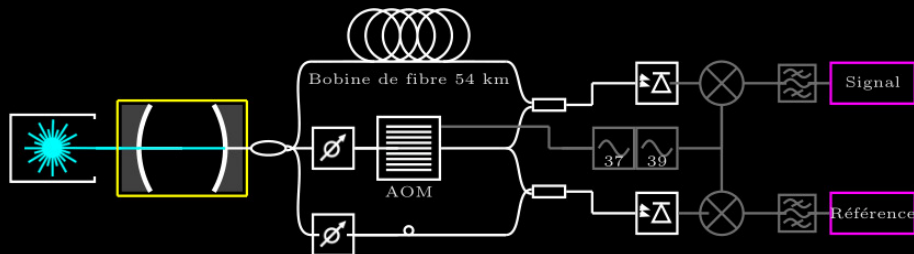
DARk Matter from Non Equal Delays

"DAMNED" allows to compare an ultrastable cavity to itself in the past.



Unequal-arm length Mach-Zender interferometer





Phase difference between the delayed and non delayed signals

$$\Delta\Phi(t) = \omega_0 T_0 + \omega_0 \int_{t-T_0}^t \left(\frac{\Delta T(t')}{T_0} + \frac{\Delta\omega(t')}{\omega_0} \right) dt'$$

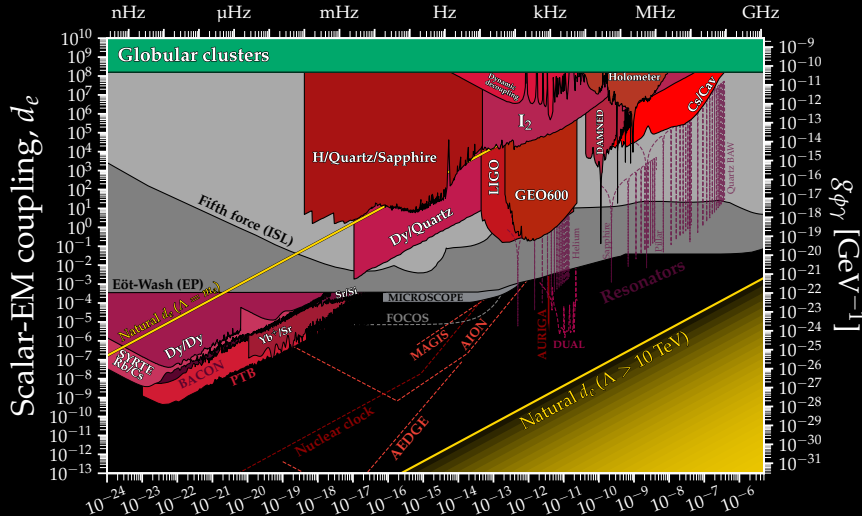
$$+ \omega_0 T_0 \left(\frac{\delta T}{T_0} + \frac{\delta\omega}{\omega_0} \right) \sin \left(\omega_\varphi t - \omega_\varphi \frac{T_0}{2} \right) \text{sinc} \left(\omega_\varphi \frac{T_0}{2} \right)$$

Color code

NOMINAL VALUE

NOISE

DARK MATTER EFFECT



$$[1 - \Lambda^2 \mathcal{G}] \kappa \phi^8$$

Caputo et al., PRD 2021

- Available with any stable/accurate metrology equipment
- Still some mass range to cover
- Same data analysis as an axion search