## **Polarization of spin-half particles with effective spacetime-dependent masses**

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Hydrodynamics and related observables in heavy-ion collisions October 28 - 31, 2024, Nantes, France

S. Bhadury, A. Das, W. Florkowski, Gowthama K. K., R. R., *Phys. Lett.* B 849 (2024) 138464

#### **Features of non-central collisions :**



Before collision

After collision

Figure 1: Heavy-ion collision experiments. [LHC Collaboration, JINST 17 (2022) 05, P05009]

- Special features of Non-Central Collisions :
	- Large Magnetic Field. [A. Bzdak and, V. Skokov, Phys. Lett. B 710 (2012) 171–174]
	- Large Angular Momentum. [F. Becattini et. al. Phys. Rev. C 77 (2008) 204906]
	- Barge Tingular Momentum, [p. becausinel. al. Filys. Rev. C // (2006) 204306]<br>
	Particle polarization at small  $\sqrt{S_{NN}}$ . [STAR Collaboration, Nature 548 62-65, 2017]

## **Particle polarization :**



Figure 2: Origin of particle polarization. [W. Florkowski *et al*, PPNP 108 (2019) 103709]

 $\circ$  Large orbital angular momentum  $\rightarrow$  local vorticity  $\rightarrow$  spin alignment [Z.-T. Liang and X.-N. Wang, Phys. Rev. Lett. 94, 102301 (2005); Phys. Lett. B 629, 20 (2005)]

### **Particle polarization :**



Experimental evidence, [STAR Collaboration, Nature 548, 62 (2017), Phys. Rev. Lett. 123, 132301 (2019), Phys. Rev. Lett. 126, 162301 (2021)]

◦ Theoretical models assuming equilibration of spin d.o.f. explain this data.

## **Particle polarization :**



Figure 3: Observation (L) and prediction (R) of longitudinal polarization. [Left: Phys. Rev. Lett. **123** 132301 (2019); Right: Phys. Rev. Lett. **120** 012302 (2018)]

 $\circ$  Theoretical models assuming equilibration of spin d.o.f. predict the opposite sign.

• Non-local collisions have been considered.

[N. Weickgenannt et. al., Phys. Rev. Lett. **<sup>127</sup>** 052301 (2021); Phys. Rev. D **<sup>106</sup>**, 116021 (2022)]

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• Dissipative spin-hydrodynamics has been formulated.

[S. Bhadury et. al., Phys.Lett.B **<sup>814</sup>** 136096 (2021); Phys. Rev. D **<sup>103</sup>** 014030 (2021)]

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• A theory with spin should be constructed from Quantum Field Theory.

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 $\mathrm{QFT} \overset{\mathrm{WF}}{\longrightarrow} \mathrm{Kinetic}\ \mathrm{Equation}\ \frac{\int_\mathrm{p}}{\longrightarrow} \mathrm{Macroscopic}\ \mathrm{theory}.$ 

#### **NJL model in the mean field approximation :**

• Let us consider the Lagrangian of the Nambu-Jona-Lasinio (NJL) type [W. Florkowski, J. Hufner , S.P. Klevansky, L. Neise, Annals Phys. **245** 445-463 (1996)]

$$
\mathcal{L} = \bar{\psi} \left( i \hat{\phi} - m_0 \right) \psi + G \left[ \left( \bar{\psi} \psi \right)^2 + \left( \bar{\psi} i \gamma_5 \psi \right)^2 \right].
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$$

• This leads to the equation of motion (we assume  $m_0 = 0$ )

$$
\[\dot{i}\,\hat{\phi}-\sigma(x)-\dot{i}\gamma_5\,\pi(x)\]\psi=0,
$$

where we restrict ourselves to the mean field approximation

$$
\sigma = \langle \hat{\sigma} \rangle = -2G \langle \bar{\psi} \psi \rangle, \qquad \qquad \pi = \langle \hat{\pi} \rangle = -2G \langle \bar{\psi} i \gamma_5 \psi \rangle.
$$

• The Wigner function is defined as

$$
\mathcal{W}_{\alpha\beta}(x,k) \equiv \int d^4y \, e^{ik \cdot y} \, G_{\alpha\beta} \left( x + \frac{y}{2}, x - \frac{y}{2} \right)
$$

where  $G_{\alpha\beta}(x,y) = \langle \bar{\psi}_{\beta}(y) \psi_{\alpha}(x) \rangle$ .

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where  $G_{\alpha\beta}(x,y) = \langle \bar{\psi}_{\beta}(y) \psi_{\alpha}(x) \rangle$ .

• The kinetic equation satisfied by the Wigner function is

$$
\left[\gamma_{\mu}K^{\mu}-\sigma+\frac{i\hbar}{2}\left(\partial_{\mu}\sigma\right)\partial_{k}^{\mu}-i\gamma_{5}\pi-\frac{\hbar}{2}\gamma_{5}\left(\partial_{\mu}\pi\right)\partial_{k}^{\mu}\right]\mathscr{W}(x,k)=0.
$$

where  $K^{\mu} = k^{\mu} + \frac{i\hbar}{2} \partial^{\mu}$ .

## **Clifford-algebra decomposition :**

• We can decompose the Wigner function in the Clifford-algebra basis as

$$
\mathcal{W}=\mathcal{F}+i\gamma_5\mathcal{P}+\gamma_\mu\mathcal{V}^\mu+\gamma^\mu\gamma_5\mathcal{A}_\mu+\frac{1}{2}\sigma^{\mu\nu}\mathcal{S}_{\mu\nu},
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• The components are real and obtained by respective traces

$$
\mathcal{F} = \mathrm{Tr} \Big[ \mathcal{W} \Big], \quad \mathcal{P} = -i \mathrm{Tr} \Big[ \gamma_5 \mathcal{W} \Big], \quad \mathcal{V}^{\mu} = \mathrm{Tr} \Big[ \gamma^{\mu} \mathcal{W} \Big], \quad \mathcal{A}^{\mu} = \mathrm{Tr} \Big[ \gamma_5 \gamma^{\mu} \mathcal{W} \Big], \quad (\ldots)
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$$

• The quantum kinetic equations for the components resulting from the kinetic equation for the Wigner function are

$$
K^{\mu}\mathcal{V}_{\mu} - \sigma\mathcal{F} + \pi\mathcal{P} = -\frac{i\hbar}{2} \Big[ \left( \partial_{\nu}\sigma \right) \left( \partial_{k}^{\nu}\mathcal{F} \right) - \left( \partial_{\nu}\pi \right) \left( \partial_{k}^{\nu}\mathcal{P} \right) \Big]
$$

$$
-iK^{\mu}\mathcal{A}_{\mu} - \sigma\mathcal{P} - \pi\mathcal{F} = -\frac{i\hbar}{2} \Big[ \left( \partial_{\nu}\sigma \right) \left( \partial_{k}^{\nu}\mathcal{P} \right) + \left( \partial_{\nu}\pi \right) \left( \partial_{k}^{\nu}\mathcal{F} \right) \Big]
$$

$$
K_{\mu}\mathcal{F} + iK^{\nu}\mathcal{S}_{\nu\mu} - \sigma\mathcal{V}_{\mu} + i\pi\mathcal{A}_{\mu} = -\frac{i\hbar}{2} \Big[ \left( \partial_{\nu}\sigma \right) \left( \partial_{k}^{\nu}\mathcal{V}_{\mu} \right) - \left( \partial_{\nu}\pi \right) \left( \partial_{k}^{\nu}\mathcal{A}_{\mu} \right) \Big]
$$

$$
iK^{\mu}\mathcal{P} - K_{\nu}\tilde{\mathcal{S}}^{\nu\mu} - \sigma\mathcal{A}^{\mu} + i\pi\mathcal{V}^{\mu} = -\frac{i\hbar}{2} \Big[ \left( \partial_{\nu}\sigma \right) \left( \partial_{k}^{\nu}\mathcal{A}^{\mu} \right) - \left( \partial_{\nu}\pi \right) \left( \partial_{k}^{\nu}\mathcal{V}^{\mu} \right) \Big]
$$

$$
2iK^{[\mu}\mathcal{V}^{\nu]} - \varepsilon^{\mu\nu\alpha\beta}K_{\alpha}\mathcal{A}_{\beta} - \pi\tilde{\mathcal{S}}^{\mu\nu} + \sigma\mathcal{S}^{\mu\nu} = \frac{i\hbar}{2} \Big[ \left( \partial_{\gamma}\sigma \right) \left( \partial_{k}^{\gamma}\mathcal{S}^{\mu\nu} \right) - \left( \partial_{\gamma}\pi \right) \left( \partial_{k}^{\gamma}\tilde{\mathcal{S}}^{\mu\nu} \right) \Big]
$$

where  $\tilde{\delta}^{\mu\nu} = \frac{1}{2} \varepsilon^{\mu\nu\alpha\beta} \delta_{\alpha\beta}$ .

• In order to obtain the classical transport equations one makes semiclassical expansion of the WF components as follows

$$
\mathfrak{X} = \mathfrak{X}_{(o)} + \hbar \mathfrak{X}_{(1)} + \hbar^2 \mathfrak{X}_{(2)} + \cdots
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$$

- the expressions at the appropriate powers of  $\hbar$  are compared
- In the classical limit the spin dynamics is described by the behavior of the axial current density  $\mathcal{A}_{(0)}^{\mu}$  whose evolution equation is determined by considering the transport equations up to the first order in  $\hbar$ .
- In the following we will set  $\pi = 0$  and  $\sigma_{(0)}(x) = M(x)$ .
- $M(x)$  is the in-medium mass of particles, which is treated as externally given and plays the role of a background scalar field. [M.I. Gorenstein, S.-N. Yang, Phys. Rev. D 52 (1995) 5206–5212 P. Romatschke, Phys. Rev. D 85 (2012) 065012 L. Tinti, A. Jaiswal, R. Ryblewski, Phys. Rev. D 95 (5) (2017) 054007 A. Czajka, S. Hauksson, C. Shen, S. Jeon, C. Gale, Phys. Rev. C 97 (4) (2018) 044914]

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• For the scalar component one can get the dynamic equation

 $p^{\mu}\partial_{\mu}\mathcal{F}(x,p) + M(x)\partial_{\mu}M(x)\partial_{p}^{\mu}\mathcal{F}(x,p) = 0,$ 

#### **Kinetic equation for axial current**

• The kinetic equation for axial current is

[W. Florkowski, J. Hufner , S.P. Klevansky, L. Neise, Annals Phys. **245** 445-463 (1996)]

$$
k^{\alpha}\left(\partial_{\alpha}\mathfrak{A}^{\mu}\right)+M\left(\partial_{\alpha}M\right)\left(\partial_{(k)}^{\alpha}\mathfrak{A}^{\mu}\right)+\left(\partial_{\alpha}\ln M\right)\left(k^{\mu}\mathfrak{A}^{\alpha}-k^{\alpha}\mathfrak{A}^{\mu}\right)=0.
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$$

• In the leading order of the semiclassical expansion, one can use the ansatz [S. Bhadury, W. Florkowski, A. Jaiswal, A. Kumar, R. R., Phys. Lett. B **814** (2021) 136096]

$$
\mathcal{A}^{\mu}(x,k) = 2M \int dP \, dS \, s^{\mu} \Big[ f^{+}(x,p,s) \delta^{(4)}(k-p) + f^{-}(x,p,s) \delta^{(4)}(k+p) \Big].
$$

where  $f^{\pm}(x, p, s)$  are the distribution functions for particles (+) and antiparticles  $(-)$  in the extended phase-space of position x, on-shell momentum  $p^{\mu} = (p^{\circ}, \mathbf{p})$   $(p^2 = M^2(x))$ , and spin  $s^{\mu} = (s^{\circ}, \mathbf{s})$ .

- Integration measures are:  $dP = \frac{d^3p}{E_p}$ ,  $dS = \left(\frac{M}{\pi s}\right)d^4s \,\delta(s\cdot s + \mathfrak{s}^2)\,\delta(p\cdot s)$ .
- s denotes the value of the Casimir operator  $s = (1/2)(1/2 + 1) = 3/4$
- This ansatz satisfies the condition:  $k \cdot \mathcal{A}(x, k) = 0$ .

• The hydrodynamic variable describing the dynamics of spin is the spin tensor  $S^{\lambda\mu\nu}(x).$ 

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- The **canonical** spin tensor is defined as

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S_{\text{can}}^{\lambda\mu\nu}(x) = \frac{1}{2} \varepsilon^{\lambda\mu\nu\alpha} \int d^4k \,\mathfrak{A}_{\alpha}(x,k).
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• The **GLW (de Groot, van Leeuwen, van Weert)** spin tensor is defined as

[S. Bhadury, W. Florkowski, A. Jaiswal, A. Kumar, R. R., Phys. Lett. B **814** (2021) 136096]

$$
S^{\lambda,\mu\nu}(x) = \int dP \, dS \, p^{\lambda} s^{\mu\nu} [f^+(x,p,s) + f^-(x,p,s)].
$$

where  $s^{\alpha\beta} = \frac{1}{M} \varepsilon^{\alpha\beta\mu\nu} p_{\mu}s_{\nu}$  is the internal angular momentum tensor originally introduced by Mathisson. [M. Mathisson, Acta Phys. Pol. 6 (1937) 163–200]

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•  $S^{\lambda\mu\nu}(x)$  and  $S^{\lambda\mu\nu}_{\text{can}}(x)$  are related by

[W. Florkowski, A. Kumar, R.R., Prog. Part. Nucl. Phys. **108** 103709 (2019)]

$$
S_{\text{can}}^{\lambda\mu\nu} = S^{\lambda,\mu\nu} + S^{\mu,\nu\lambda} + S^{\nu,\lambda\mu}.
$$
 (\*)

#### **Evolution of Spin Tensor :**

• Recall the kinetic equation for the axial current

$$
k^{\alpha} \left(\partial_{\alpha} \mathfrak{A}^{\mu}\right) + M\left(\partial_{\alpha} M\right) \left(\partial^{\alpha}_{\left(k\right)} \mathfrak{A}^{\mu}\right) + \left(\partial_{\alpha} \ln M\right) \left(k^{\mu} \mathfrak{A}^{\alpha} - k^{\alpha} \mathfrak{A}^{\mu}\right) = 0.
$$

• Multiplying this by  $k_\beta \varepsilon_\mu^{\beta\gamma\delta}$  and integrating over k we get evolution equation for the GLW spin tensor

[S. Bhadury, A. Das, W. Florkowski, Gowthama K. K., R. R., Phys. Lett. B 849 (2024) 138464]

$$
\partial_{\alpha} S^{\alpha,\gamma\delta} = (\partial_{\alpha} \ln M) \left( S^{\gamma,\delta \alpha} - S^{\delta,\gamma \alpha} \right) \Bigg| \neq 0
$$

- As expected, the spin tensor is conserved when  $M$  is constant.
- However, if  $M$  varies the spin tensor is sourced through its derivative.

## **Conservation of Total Angular Momentum :**

• Conservation of total angular momentum implies:

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\partial_\lambda J^{\lambda,\mu\nu}=0.
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$$
\partial_{\lambda} S^{\lambda, \mu\nu}_{\text{can}} = T^{\nu\mu}_{\text{can(a)}} - T^{\mu\nu}_{\text{can(a)}}
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$$
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$$

• Using the results for the semiclassical expansion of the Wigner function (for spin-1/2 particles) one finds [W. Florkowski, A. Kumar, R. Ryblewski, Phys. Rev. C 98 (4) (2018) 044906]

$$
T^{\mu\nu}_{(a) \text{ can}}(x) = \int d^4k k^{\nu} \mathcal{V}^{\mu}_{(1)}(x,k)
$$

where  $\mathcal{V}_{(1)}^{\mu}(x,k) = -(1/(2M)) \partial^{\alpha} \mathcal{S}_{\alpha}^{\mu}(x,k)$  and  $\mathcal{S}_{\alpha\mu}(x,k) = \frac{1}{M} \varepsilon_{\alpha\mu\rho\sigma} k^{\rho} \mathcal{A}^{\sigma}(x,k)$ 

• Hence we get

$$
M\partial_{\lambda}S^{\lambda,\mu\nu}_{\text{can}} = \partial_{\lambda}\left(MS^{\nu,\lambda\mu}\right) - \partial_{\lambda}\left(MS^{\mu,\lambda\nu}\right)
$$

which using  $\star$  can be shown to reproduce equation for GLW spin tensor.

## Analytic Solutions

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Figure 4: Transverse view of non-central collisions. [S. Bhadury et. al. PLB 849 (2024) 138464]

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Figure 4: Transverse view of non-central collisions. [S. Bhadury et. al. PLB 849 (2024) 138464]

 $\bullet$  Consider a system expanding longitudinally along the  $z$ -axis:

$$
f(x, p, s) = g(x, p, s)\delta(p_x)\delta(p_y)
$$

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Figure 4: Transverse view of non-central collisions. [S. Bhadury et. al. PLB 849 (2024) 138464]

• Consider a system expanding longitudinally along the  $z$ -axis:

$$
f(x, p, s) = g(x, p, s)\delta(p_x)\delta(p_y)
$$

• Hence:  $S^{\lambda,\mu\nu}(x) = \int dP \, dS \, p^{\lambda} s^{\mu\nu} g(x, p, s) \delta(p_x) \delta(p_y) \implies S^{1,\mu\nu} = S^{2,\mu\nu} = 0.$ 

• Transverse polarization implies:

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g(x, p, s) = h(x, p, s)\delta(s_x)\delta(s_z)
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Figure 5: Transverse polarization schematic diagram. [S. Bhadury et. al. PLB 849 (2024) 138464]

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• Furthermore, if  $M = M(t, z)$ , then the only non-zero components of spin-tensor are  $S^{0,01}$ ,  $S^{3,01}$ ,  $S^{0,31}$ ,  $S^{3,31}$ .

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$$

• 2 equations, 4 unknowns, we still need further contraints!

• Let us consider longitudinal boost invariance and use the following vector basis :

$$
u^{\mu} = \begin{pmatrix} \frac{t}{\tau} \\ 0 \\ 0 \\ \frac{z}{\tau} \end{pmatrix}, \quad S_x^{\mu} = \begin{pmatrix} 0 \\ 1 \\ 0 \\ 0 \end{pmatrix}, \quad S_y^{\mu} = \begin{pmatrix} 0 \\ 0 \\ 1 \\ 0 \end{pmatrix}, \quad S_z^{\mu} = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 1 \end{pmatrix}
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\frac{d\sigma}{d\tau} + \frac{\sigma}{\tau} = 0.
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• The solution is equivalent to conservation law in Bjorken model.

## **Analytic Solutions - II (Longitudinal Polarization):**

• Longitudinal polarization implies:

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g(x, p, s) = h(x, p, s)\delta(s_x)\delta(s_y)
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Figure 6: Longitudinal polarization schematic diagram. [S. Bhadury et. al. PLB 849 (2024) 138464]

• The spin tensor under longitudinal polarization becomes :

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• Similar to transverse case, the spin decouples from the gradient of  $M(x)$  and we have a similar solution.

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- We assume the simple case:

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M = M(t), \qquad S^{3,01}(t,z) = \nu S^{0,01}(t,z) = \nu \sigma(t,z).
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• Then this leads to:

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\left(\frac{\partial}{\partial t} + \nu \frac{\partial}{\partial z}\right) \sigma(t, z) = -\left(\frac{\partial \ln M(t)}{\partial t}\right) \sigma(t, z).
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• This solution implies the increase of spin density with decreasing mass, indicating a connection between chiral restoration and spin polarization.

- Gradients of effective mass can act like a source of spin polarization.
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- Spin evolution decouples from the source term in a highly symmetric system.
- By giving up boost-invariance, we find a connection between spin polarization and chiral restoration.
- A self-consistently determined  $M(x)$  should be used to study the evolution.

## Thank you for your attention!

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# Other Aspects