



Dark matter phenomenology in 2HDM + complex singlet, probe at lepton colliders

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• 2HDM is one of the most popular extensions of Standard Model scalar sector.

- However, it is not easy to accommodate a dark matter candidate in the 2HDM, inert doublet scenario is quite strongly constrained by the data.
- The singlet extension of 2HDM, can give rise to a viable dark matter candidate.
- In addition, recently observed 95 GeV excess, can be explained if we consider complex singlet extension of 2HDM, simultaneously having the possibility of dark matter.
- We have considered all the constraints (experimental and theoretical) on the model, studied the dark matter phenomenology and the prospect of detecting such a scenario at the future colliders.

Scalar potential

2HDM scalar potential with imposed softly broken $\mathcal{Z}_2\text{-symmetry}$ and complex singlet with \mathcal{Z}_2' symmetry :

$$\begin{split} V_{2HDMS} &= -m_{11}^{2}\Phi_{1}^{\dagger}\Phi_{1} - m_{22}^{2}\Phi_{2}^{\dagger}\Phi_{2} - [m_{12}^{2}\Phi_{1}^{\dagger}\Phi_{2} + h.c.] + \frac{\lambda_{1}}{2}(\Phi_{1}^{\dagger}\Phi_{1})^{2} \\ &+ \frac{\lambda_{2}}{2}(\Phi_{2}^{\dagger}\Phi_{2})^{2} + \lambda_{3}(\Phi_{1}^{\dagger}\Phi_{1})(\Phi_{2}^{\dagger}\Phi_{2}) + \lambda_{4}(\Phi_{1}^{\dagger}\Phi_{2})(\Phi_{2}^{\dagger}\Phi_{1}) + \frac{\lambda_{5}}{2}(\Phi_{1}^{\dagger}\Phi_{2})^{2} \\ &+ m_{5}^{2}S^{\dagger}S + \left[\frac{m_{5}^{\prime 2}}{2}S^{2} + h.c.\right] + \left[\frac{\lambda_{1}^{\prime \prime}}{24}S^{4} + h.c.\right] + \left[\frac{\lambda_{2}^{\prime \prime}}{6}(S^{2}S^{\dagger}S) + h.c.\right] \\ &+ \frac{\lambda_{3}^{\prime \prime \prime}}{4}(S^{\dagger}S)^{2} + S^{\dagger}S[\lambda_{1}^{\prime}\Phi_{1}^{\dagger}\Phi_{1} + \lambda_{2}^{\prime}\Phi_{2}^{\dagger}\Phi_{2}] + [S^{2}(\lambda_{4}^{\prime}\Phi_{1}^{\dagger}\Phi_{1} + \lambda_{5}^{\prime}\Phi_{2}^{\dagger}\Phi_{2}) + h.c.] \\ &\Phi_{i} = \left(\begin{array}{c} \phi_{i}^{+} \\ v_{i} + \rho_{i} + i\eta_{i} \end{array}\right), \quad \langle\Phi_{i}\rangle = \left(\begin{array}{c} 0 \\ v_{i} \end{array}\right) \\ &S = \frac{1}{\sqrt{2}}(v_{S} + \rho_{S} + iA_{S}), \quad \langle S\rangle = v_{S} \end{split}$$

Three scalars h_1, h_2, h_3 , charged scalars H^{\pm} , pseudoscalar A and dark matter A_5 . In our analysis h_1 is the 95 GeV scalar and h_2 is the observed Higgs at 125 GeV₃₀₀

$$\left(\begin{array}{c}h_1\\h_2\\h_3\end{array}\right)=R(\alpha_1,\alpha_2,\alpha_3)\left(\begin{array}{c}\rho_1\\\rho_2\\\rho_3\end{array}\right),$$

Interaction basis parameters :

$$\begin{split} \lambda_1,\lambda_2,\lambda_3,\lambda_4,\lambda_5,m_{12}^2,\tan\beta, \mathsf{v}_5,m_5'^2,\\ \lambda_1',\lambda_2',\lambda_4',\lambda_5',\lambda_1''=\lambda_2'',\lambda_3'' \end{split}$$

Mass basis parameters :

 $\begin{array}{l} m_{h_1}, m_{h_2}, m_{h_3}, m_{H^{\pm}}, m_A, m_{A_5}, \tan\beta, v_S, \\ \lambda'_1 - 2\lambda'_4, \lambda'_2 - 2\lambda'_5, \lambda''_1 - \lambda''_3, c_{h_1bb}, c_{h_1tt}, \\ alignm, m_{12}^2 \end{array}$

$$\mathbf{v} = \sqrt{\mathbf{v}_1^2 + \mathbf{v}_2^2}, \quad \tan \beta = \frac{\mathbf{v}_2}{\mathbf{v}_1}$$

$$c_{h_1 b b} = \frac{R_{11}}{\cos \beta}, \quad c_{h_1 t t} = \frac{R_{12}}{\sin \beta}, \quad \text{alignm} = \sin(\beta - \alpha_1 - \alpha_3 \text{sgn}(\alpha_2)) \approx 1$$

Trilinear coupling:

$$\frac{\lambda_{h_jA_SA_S}}{v} = -[(\lambda_1'-2\lambda_4')c_\beta R_{j1} + (\lambda_2'-2\lambda_5')s_\beta R_{j2} - \frac{v_S}{2v}(\lambda_1''-\lambda_3'')R_{j3}]$$

Quartic coupling:

$$\lambda_{h_{j}h_{k}A_{5}A_{5}} = -[(\lambda_{1}' - 2\lambda_{4}')R_{j1}R_{k1} + (\lambda_{2}' - 2\lambda_{5}')R_{j2}R_{k2} - \frac{1}{2}(\lambda_{1}'' - \lambda_{3}'')R_{j3}R_{k3}]$$

Testing against the Constraints

The model is written in SARAH and spectrum is generated via SPheno.

Theoretical constraints

• Boundedness from below, unitarity, vacuum stability

Oblique parameters

• Constraints from S,T,U parameters are taken into account.

Collider constraints

• Constraints from collider checked via HiggsBounds and HiggsSignals

Dark matter constraints

Dark matter observables such as relic density, direct and indirect detection cross-section are generated via micrOMEGAs.

- Observed relic density from Planck Experiments.
- Upper bound on direct detection cross-section from LUX-ZEPPLIN (LZ) experiments.
- Upper limit on indirect detection cross-section from FERMI-LAT experiments.

BP1

BP1									
	<pre>mh1 95 tan β 10</pre>	m _{h2} 125.09 ^c h ₁ bb 0.2096	^m h ₃ 900 ^c h ₁ tt 0.4192	m _A 900 Ilignm λ 0.98	${}^{m}_{H\pm}_{900}_{1'}_{1'} - 2\lambda'_{4}_{4'}_{12.3327}$	m_{12}^{2} 8.0456 × 10 ⁴ $\lambda'_{2} - 2\lambda'_{5}$ -0.3109	$^{mA_{S}}_{325.86}$ $\lambda_{1}^{\prime\prime} - \lambda_{3}^{\prime\prime}_{1.3645}$	v_{S} 239.86 Ωh^{2} 8.71 × 10 ⁻³	
BP2									
	m _{h1} 95 tan β 6.6	m _{h2} 125.09 ^c h ₁ bb 0.258	^m h ₃ 700 ^c h ₁ tt a 0.372	^m A 700 lignm λ 0.98	$m_{H\pm}{700}_{1 - 2\lambda'_{4}}{12.75}$	m_{12}^2 7.2576 × 10 ⁴ $\lambda'_2 - 2\lambda'_5$ -0.3135	m_{A_S} 325.86 $\lambda_1'' - \lambda_3''$ -1.0112	v_{S} 239.86 Ωh^{2} 3.16 × 10 ⁻⁴	
BP55									
	m _h 95 tan 2	$^{h_{1}}$ $^{m_{h_{2}}}$ $^{125.09}$ $^{\beta}$ $^{c_{h_{1}bb}}$ $^{c_{h_{1}bb}}$ $^{0.2323}$	m_{h_3} 650 $h_1 tt$ 3 0.3105	<i>m</i> _A 800 alignm 0.97	$m_{H\pm} 800$ $\lambda'_{1} - 2.$ 0.00209	$\begin{array}{c} m_{12}^2 \\ 1.69 \times 10^1 \\ \lambda_4' & \lambda_2' - 2\lambda_2' \\ 9 & 0.000746 \end{array}$	$ \begin{array}{c} & m_{A_{S}} \\ 5 & 55.596 \\ 5 & \lambda_{1}^{\prime\prime} - \lambda_{3}^{\prime\prime} \\ -0.025735 \end{array} $	νς 300 Ωh ² 0.11	
BP2900									
	m _{h1} 95 tan β 5	m _{h2} 125.09 ^c h ₁ bb 0.3669	^m h ₃ 2900 ^c h ₁ tt 0.3393	<i>m_A</i> 2900 alignm 0.99995	$\begin{array}{c} {}^{m}_{H^{\pm}}\\ {}^{2900}\\ \lambda_{1}^{\prime}-2\lambda\\ {}^{7.616}\end{array}$	$ \begin{array}{c} m_{12}^{2} \\ 1.6173 \times 1 \\ \lambda_{2}^{\prime} - 2\lambda_{3} \\ 0.0 \end{array} $	${}^{m_{A_{S}}}_{5}$	V_{5} 1000 M^{2} Ωh^{2} 0.111	



- The DM relic density has a dip at $2m_{A_S} \sim m_{h_3}$, where the resonant annihilation channel via h_3 opens up.
- The direct detection bound gets relaxed in certain regions of the parameter space where cancellation between various contribution takes place.
- The DM direct detection limits from LZ relaxes for underabundant DM due to the rescaling factor $\zeta = \Omega h^2 / (\Omega h_{\rm PLANCK}^2)$ rescaled by DM relic density.

Search for BP1 at HL-LHC

BP1 corresponds to $m_{h_3} = 900$ GeV and $m_{A_S} = 326$ GeV



Process	Signal Significance
Gluon fusion	1.36 σ
Vector boson fusion	0.007 σ

Dark matter search at lepton colliders

 $\mathsf{Mono-}\gamma$



Mono-Z



The cross-sections of mono-photon and mono-Z processes



• The muon-collider has larger cross-section of the mono-photon and mono-Z processes than the e^+e^- collider, due to larger muon Yukawa coupling compared to electron Yukawa.

We have used WHIZARD for the analysis.

Kinematic distributions



	m_{h_3} (GeV)	$m_\chi~({ m GeV})$	Ωh^2	$BR(h_3 \rightarrow \chi \chi)$	$BR(h_2 \rightarrow \chi \chi)$
BP1	900	325.86	8.71×10^{-3}	0.25	-
BP2	700	325.86	3.16×10^{-4}	0.48	-
BP3	700	156.0	1.61×10^{-4}	0.69	-
DM55	650	55.6	0.11	3.81×10^{-9}	0.0199
BP2900	2900	1000	0.111	0.0359	-

Process	Production cross-section (pb) at $\sqrt{s} =$		
	1 TeV	3 TeV	
$\gamma \nu \bar{\nu}$	2.447	2.964	

Table: Whizard cross sections for SM background at $\sqrt{s} = 1$ and 3 TeV.

Benchmark	$S(1 \ ab^{-1})$	$S(10 ab^{-1})$
BP1	0.76	2.4
BP2	0.59	2.2
BP3	1.7	5.3

 Table: 1 TeV mono- γ +missing energy

 signal

Analysis of low mass dark matter-DM55

- In this benchmark the the dark matter comes from the decay of 125 GeV Higgs boson h_2 .
- We will look into mono-Z (in the dilepton final state), Higgsstrahlung is the major production.



At low \sqrt{s} , required for the low mass DM scenario, the major background comes from $\nu\nu Z$ where $\nu\nu$ comes from Z-boson.

To reduce this background, we construct the variable M missing-mass which is the invariant mass of the missing particles.

$$\begin{split} M &= E_{inv}^2 - |p_{inv}|^2 \\ &= (\sqrt{s} - E_Z)^2 - |p_Z^2|^2 \\ &= (\sqrt{s} - E_Z)^2 - (E_Z^2 - m_Z^2) \\ &= s - 2\sqrt{s}E_Z + m_Z^2 \end{split}$$



After applying a cut M > 100 GeV we achieve :

Benchmark	$S(\sqrt{s}=250 \text{ GeV})$	$S(\sqrt{s}=500 \text{ GeV})$		
BP55	4.3 $(1ab^{-1})$, 7.4 $(3ab^{-1})$	1.2 $(1ab^{-1})$, 2.0 $(3ab^{-1})$		

Table: At ILC

Benchmark	$S(\sqrt{s}=1 \text{ TeV})$	$S(\sqrt{s}=3 \text{ TeV})$
BP55	5.4 (10 <i>ab</i> ⁻¹),	$0.38~(10ab^{-1})$

Table: At muon collider

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Summary

- We perform a scan of the 2HDMS parameter space, choose benchmarks with varied masses and portal couplings.
- The benchmarks that are under-relic, easier to probe at the collider, due to larger portal coupling and invisible branching.
- The DM mass above a few hundred GeV will be difficult to probe at the LHC, ILC or muon collider will be more sensitive.
- Low mass DM $\lesssim \frac{m_{h_2}}{2}$ can be best probed at the mono-Z final state, at ILC or muon collider.

Further things to do

- The high mass DM BP2900, can be probed only at muon collider, analysis still to be done, VBF production process looks most promising.
- We would like to establish a complementarity between different machines as well as final states.
- The impact of polarization is also under study.