

# Exhaustive Symbolic Regression

Learning Astrophysics directly from Data

Harry Desmond



w/ Deaglan Bartlett & Pedro Ferreira

[arXiv:2211.11461](https://arxiv.org/abs/2211.11461)

[arXiv:2301.04368](https://arxiv.org/abs/2301.04368)

[arXiv:2304.06333](https://arxiv.org/abs/2304.06333)

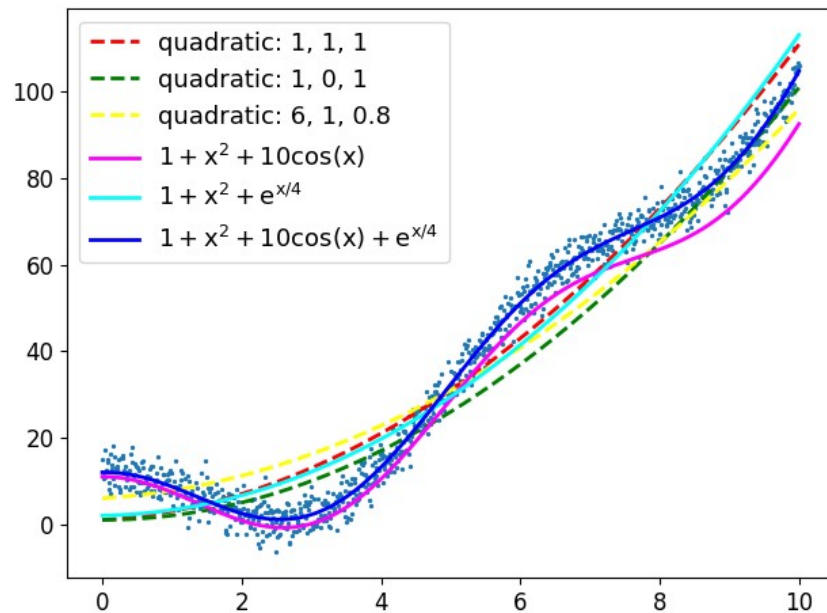
[arXiv:2310.16786](https://arxiv.org/abs/2310.16786)

# Symbolic Regression overview

- Discover *functions* describing a dataset rather than parameters of predefined function

Numerical regression:  $y = 6 + 1x + 0.8x^2$

Symbolic regression:  $y = 1 + x^2 + 10\cos(x) + e^{x/4}$

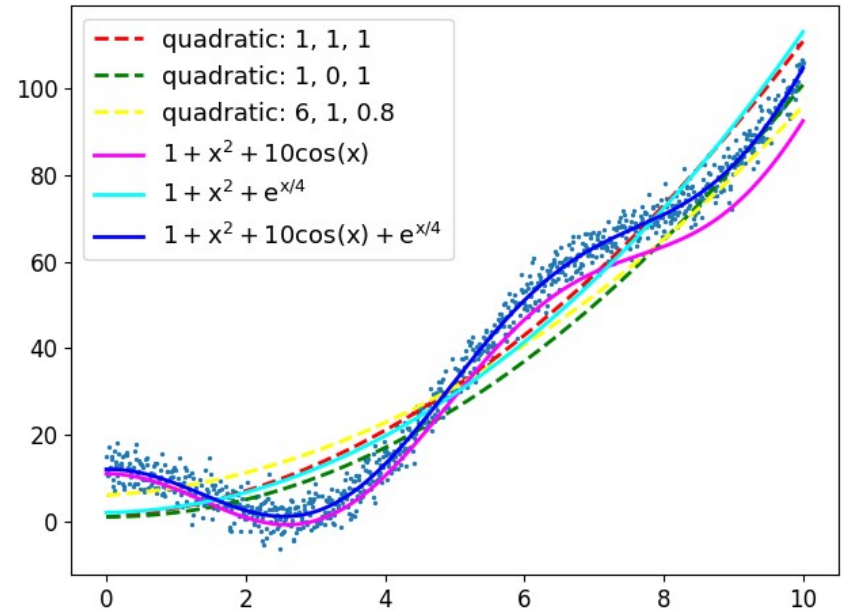


# Symbolic Regression overview

- Discover *functions* describing a dataset rather than parameters of predefined function
- Advantages:
  - Much more general (reduces confirmation bias)
  - Highly interpretable
  - Easy to prevent overfitting
- Difficulties:
  - Larger search space makes convergence harder
  - Optimisation methods of numerical regression not applicable

Numerical regression:  $y = 6 + 1x + 0.8x^2$

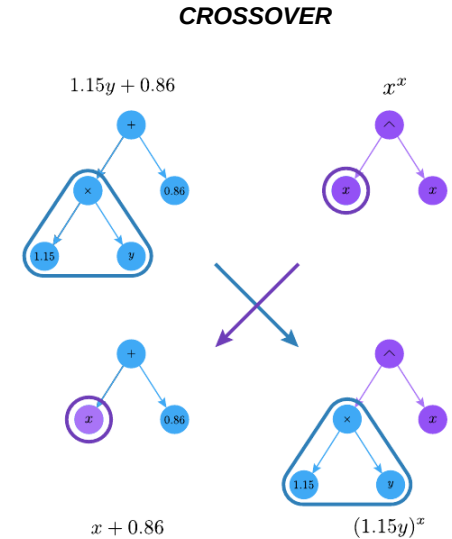
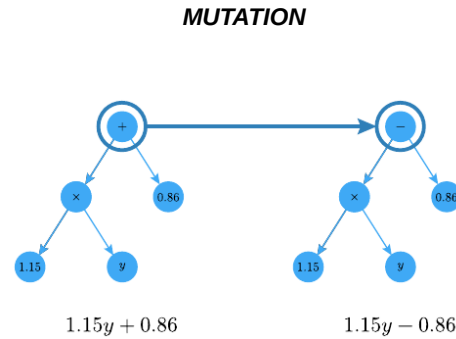
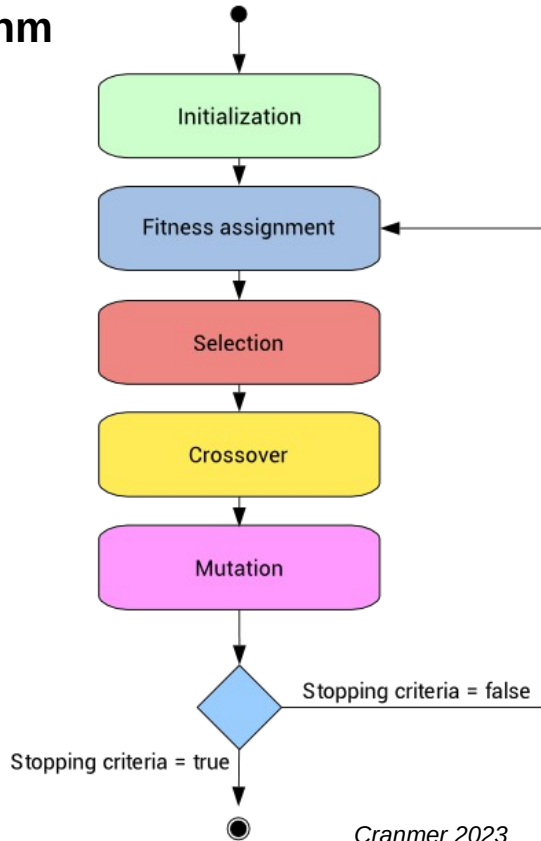
Symbolic regression:  $y = 1 + x^2 + 10\cos(x) + e^{x/4}$



# Traditional Symbolic Regression

## I. Generating functions

**Genetic Algorithm**  
(e.g. *PySR*,  
*DataModeler*)



# Traditional Symbolic Regression

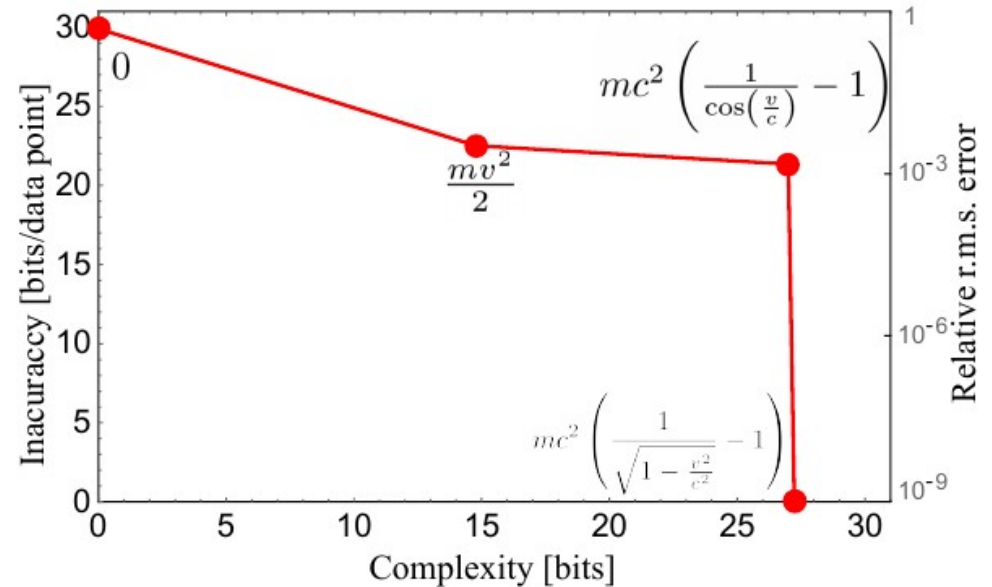
## II. Assessing functions

- Problem: Can typically get 0 error with some (very complex) overfitted function
- Solution: *two* objectives, accuracy and simplicity
- The best equations are the ones that cannot be made more accurate without also being made more complex

# Traditional Symbolic Regression

## II. Assessing functions

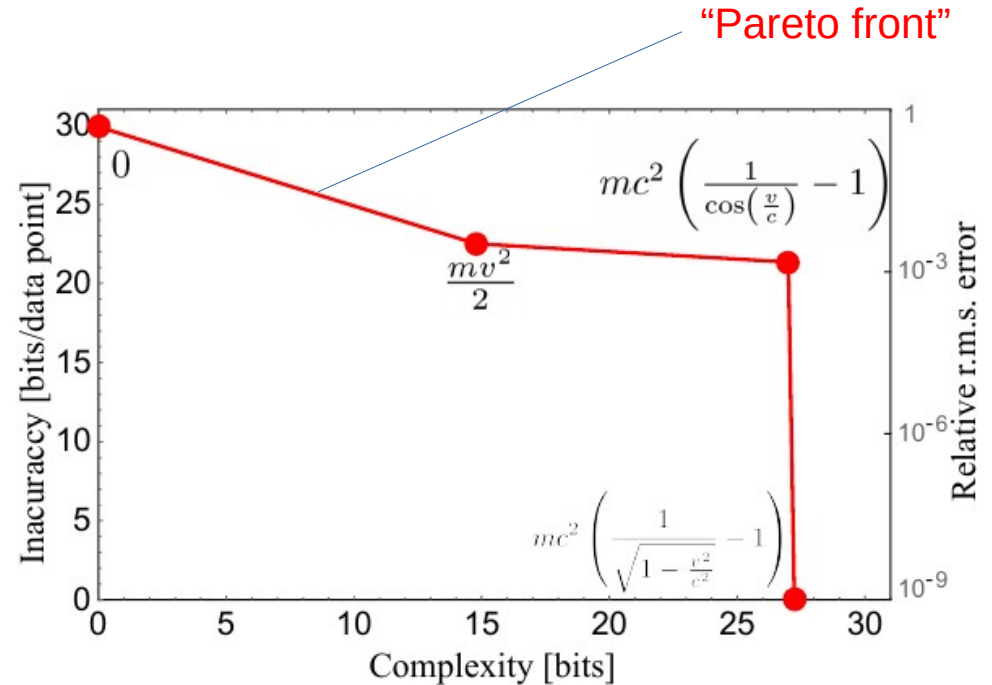
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# Traditional Symbolic Regression

## II. Assessing functions

- Problem: Can typically get 0 error with some (very complex) overfitted function
- Solution: *two* objectives, accuracy and simplicity
- The best equations are the ones that cannot be made more accurate without also being made more complex (“Pareto-optimal”)



# Exhaustive Symbolic Regression

Designed to overcome two problems:

- 1 Stochastic method may fail to find any given function
- 2 Typical accuracy definitions fail to account for data uncertainties, and complexity definition is largely arbitrary. The two are incommensurable.



# Exhaustive Symbolic Regression

Designed to overcome two problems:

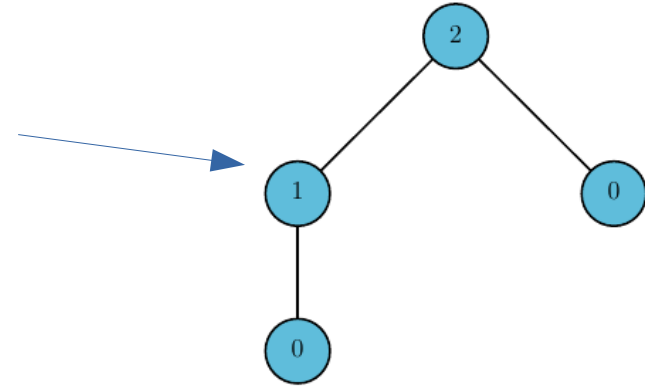
- 1 Stochastic method may fail to find any given function  
⇒ Search exhaustively, complexity by complexity
- 2 Typical accuracy definitions fail to account for data uncertainties, and complexity definition is largely arbitrary. The two are incommensurable.  
⇒ Use information-theoretic *Minimum Description Length (MDL) principle*

# Exhaustive Symbolic Regression

## I. Function generation & optimisation

1) Generate all possible trees with given complexity = #nodes, with placeholder operators labelled by arity (number of arguments to operator)

2) Decorate trees with all operator permutations



# Exhaustive Symbolic Regression

## I. Function generation & optimisation

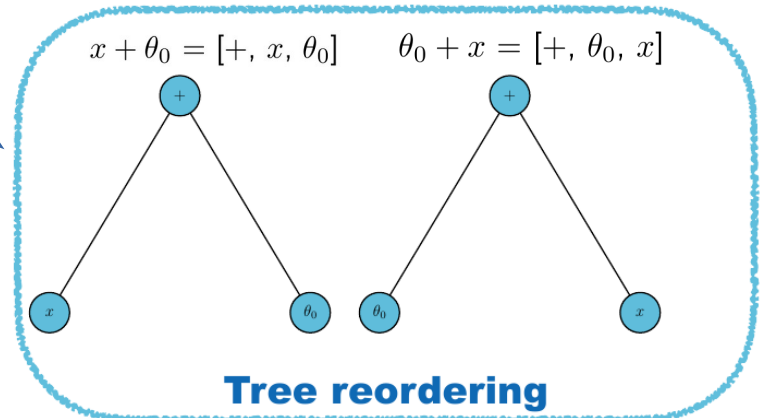
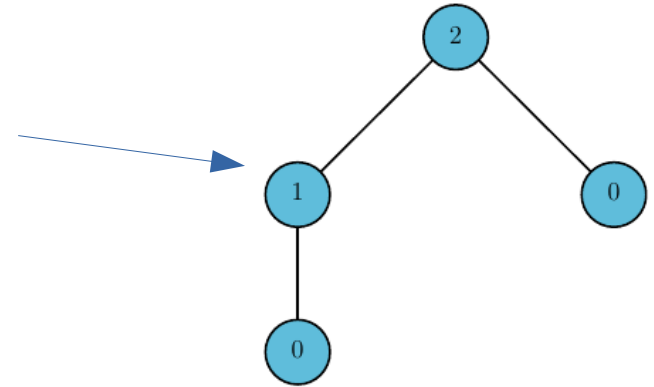
1) Generate all possible trees with given complexity = #nodes, with placeholder operators labelled by arity (number of arguments to operator)

2) Decorate trees with all operator permutations

3) Simplify and remove duplicates (*tree reordering, parameter permutations, simplifications, reparametrisation invariance, parameter combinations*)

4) Calculate maximum-likelihood parameter values

5) Repeat for all desired complexities



# Exhaustive Symbolic Regression

## II. Model selection principle: *minimum description length*

$$L(D) = L(H) + L(D | H)$$

Description length      Hypothesis      Residuals

The diagram shows the equation  $L(D) = L(H) + L(D | H)$ . Each term in the equation is enclosed in a light blue dashed-line box. A blue line points from the label 'Description length' to the box around  $L(D)$ . Another blue line points from the label 'Hypothesis' to the box around  $L(H)$ . A third blue line points from the label 'Residuals' to the box around  $L(D | H)$ .

- Purpose of functional fit is *data compression*
- Most information-efficient function has minimum  $L(D)$

# Exhaustive Symbolic Regression

## II. Model selection principle: *minimum description length*

$$L(D) = L(H) + L(D | H)$$

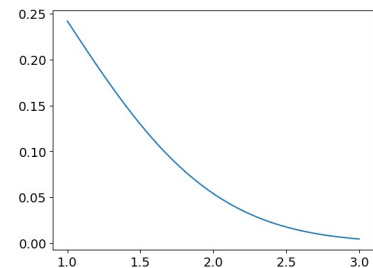
Description length      Hypothesis      Residuals

$$L(D) = -\log(\mathcal{L}(\hat{\theta})) + k \log(n) + \sum_i^p \log(|\hat{\theta}_i|/\Delta_i) + p \log(2) + \sum_j \log(c_j)$$

Shannon-Fano coding       $k$  nodes with  $n$  basis functions needs  $\log(n^k)$  nats      Send  $p$  parameters, each with precision  $\Delta_i$       Integer constants

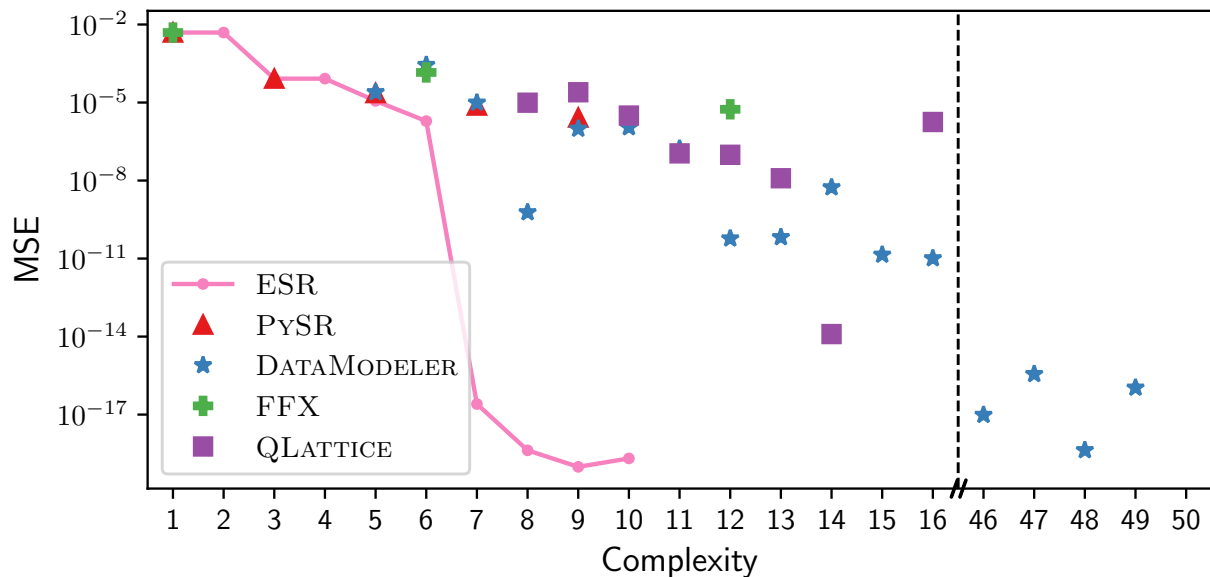
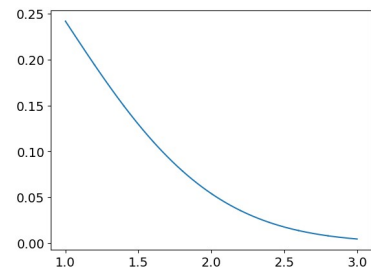
- Purpose of functional fit is *data compression*
- Most information-efficient function has minimum  $L(D)$
- Both accuracy and complexity expressed in nats  $\Rightarrow$  can be combined
- Accounts for both functional and parametric complexity. Accuracy is likelihood.

# *Test case 0: Benchmarking*



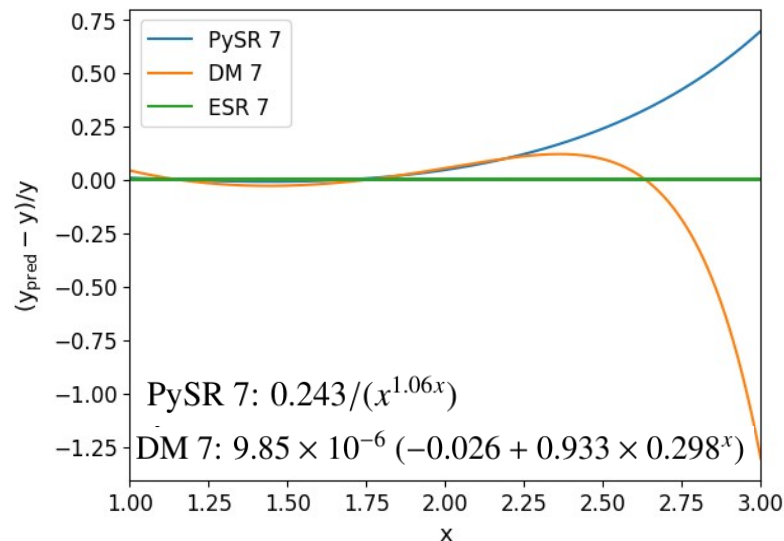
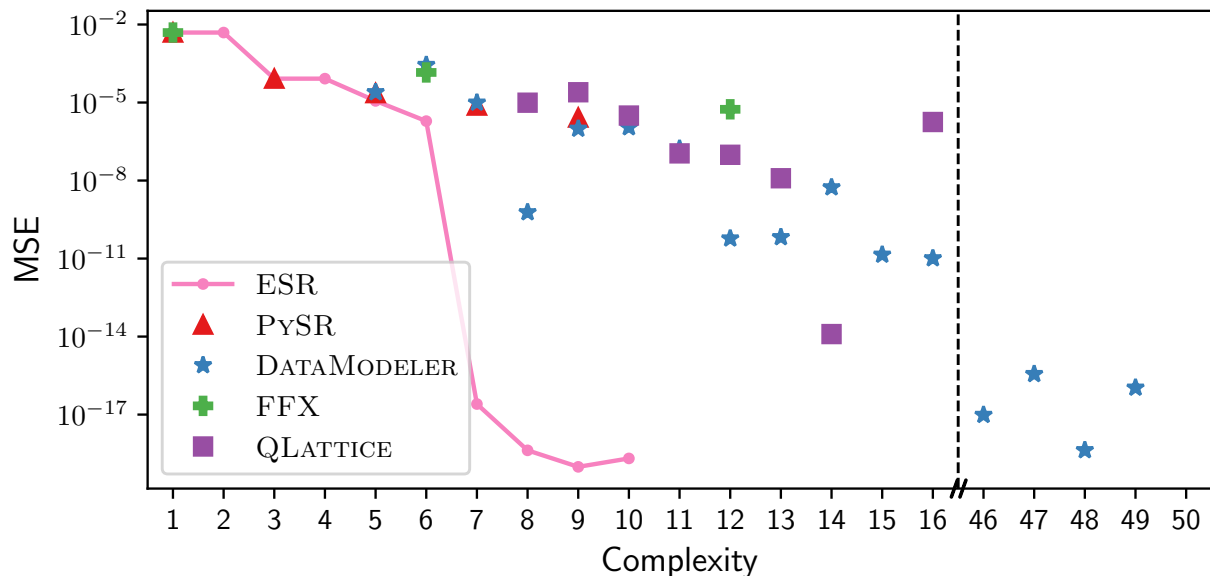
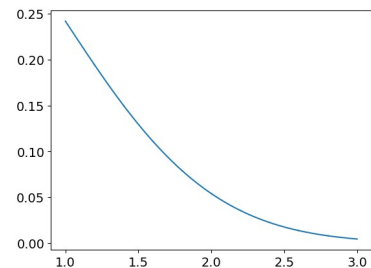
- *feynman\_1\_6\_2a* dataset from the *SRBench 2022 Competition*

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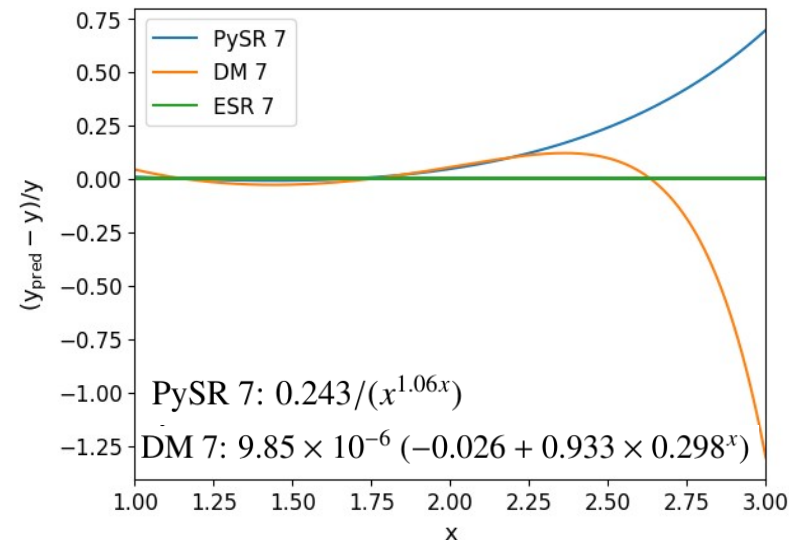
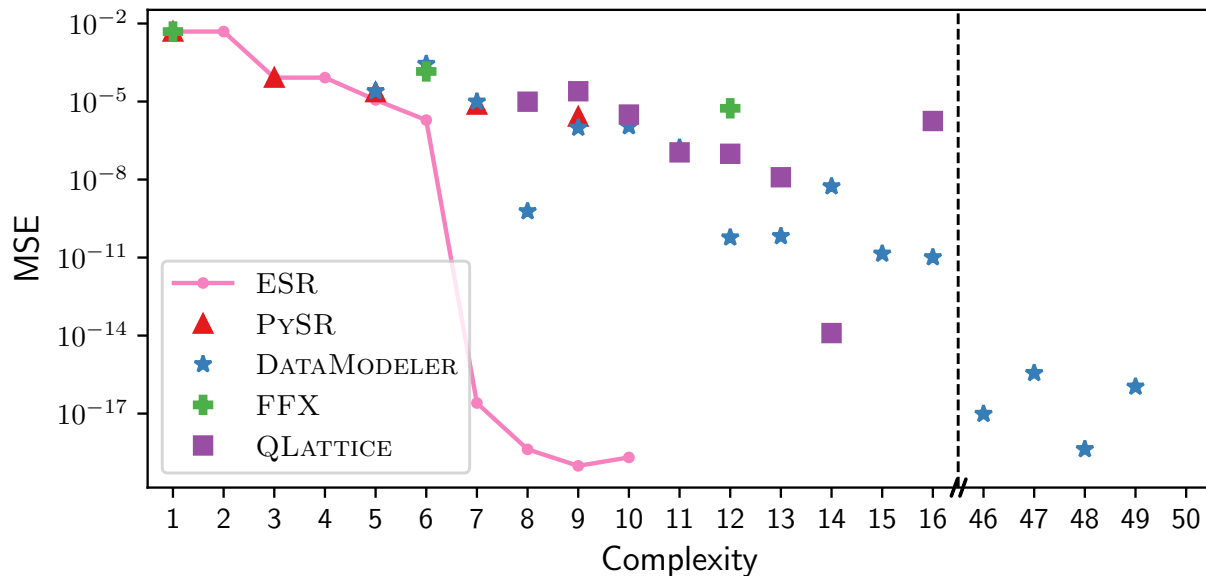
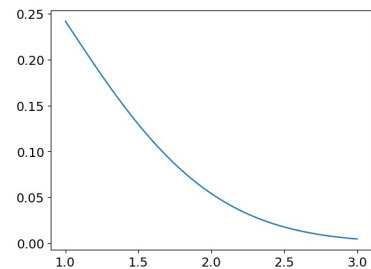


- *feynman\_1\_6\_2a* dataset from the *SRBench 2022 Competition*
- Not only does ESR get by far the lowest error...

$$y = \theta_1 \theta_0^{x^2}$$
$$\theta_0 = 0.6065$$
$$\theta_1 = 0.3989$$



# Test case 0: Benchmarking



- *feynman\_1\_6\_2a* dataset from the *SRBench 2022 Competition*
- Not only does ESR get by far the lowest error... it discovers the standard normal!

$$y = \theta_1 \theta_0^2$$
$$\theta_0 = 1/\sqrt{e}$$
$$\theta_1 = 1/\sqrt{2\pi}$$

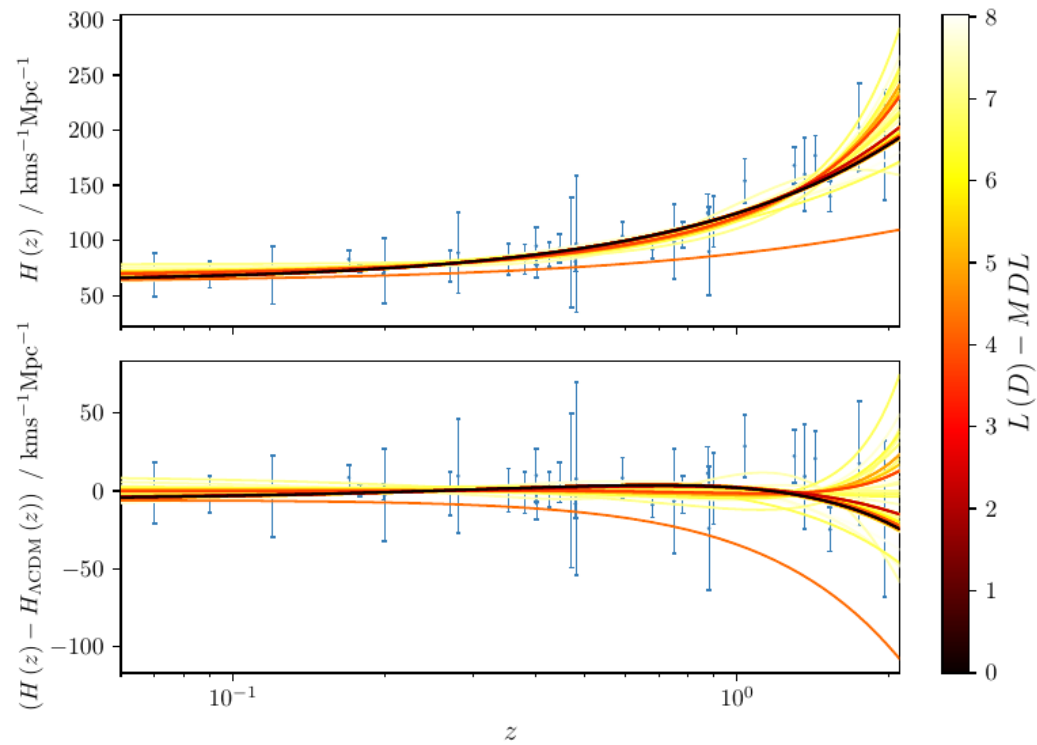
# *Test case 1: The law of cosmic expansion*

- Can we determine the functional form of cosmic expansion without assuming GR?
- How good is the Friedmann equation relative to other simple functions?

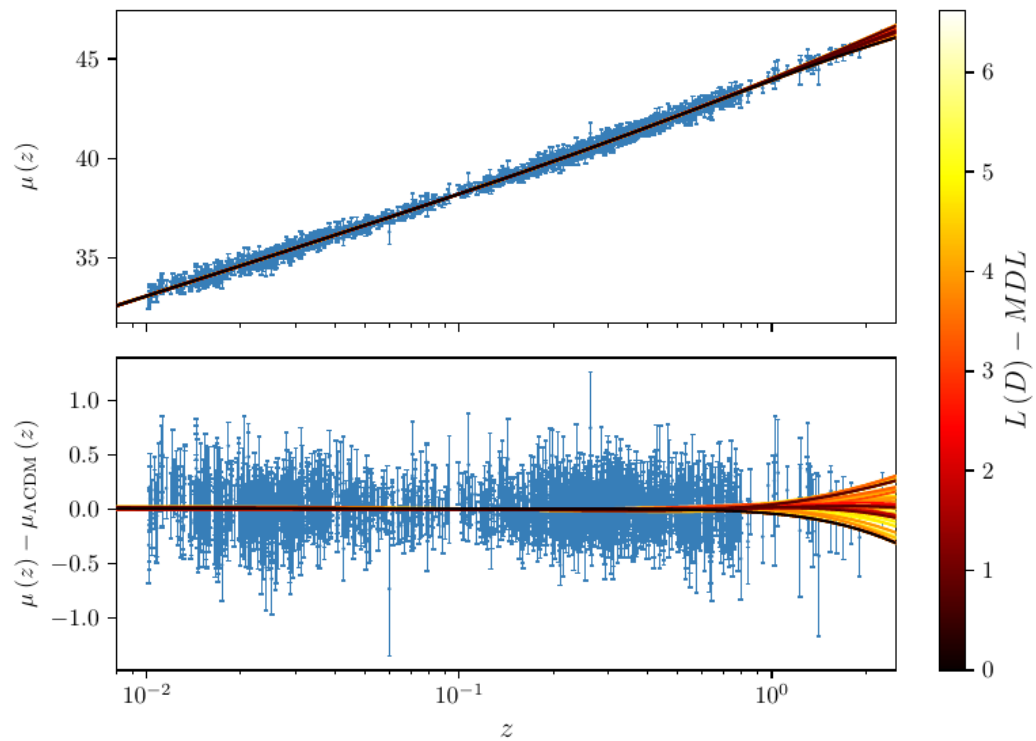
$$H(z)_{\Lambda\text{CDM}}^2 = \theta_0 + \theta_1 (1 + z)^3 \quad H(z)_{\Lambda\text{fluid}}^2 = \theta_0 + \theta_1 (1 + z)^{\theta_2}$$

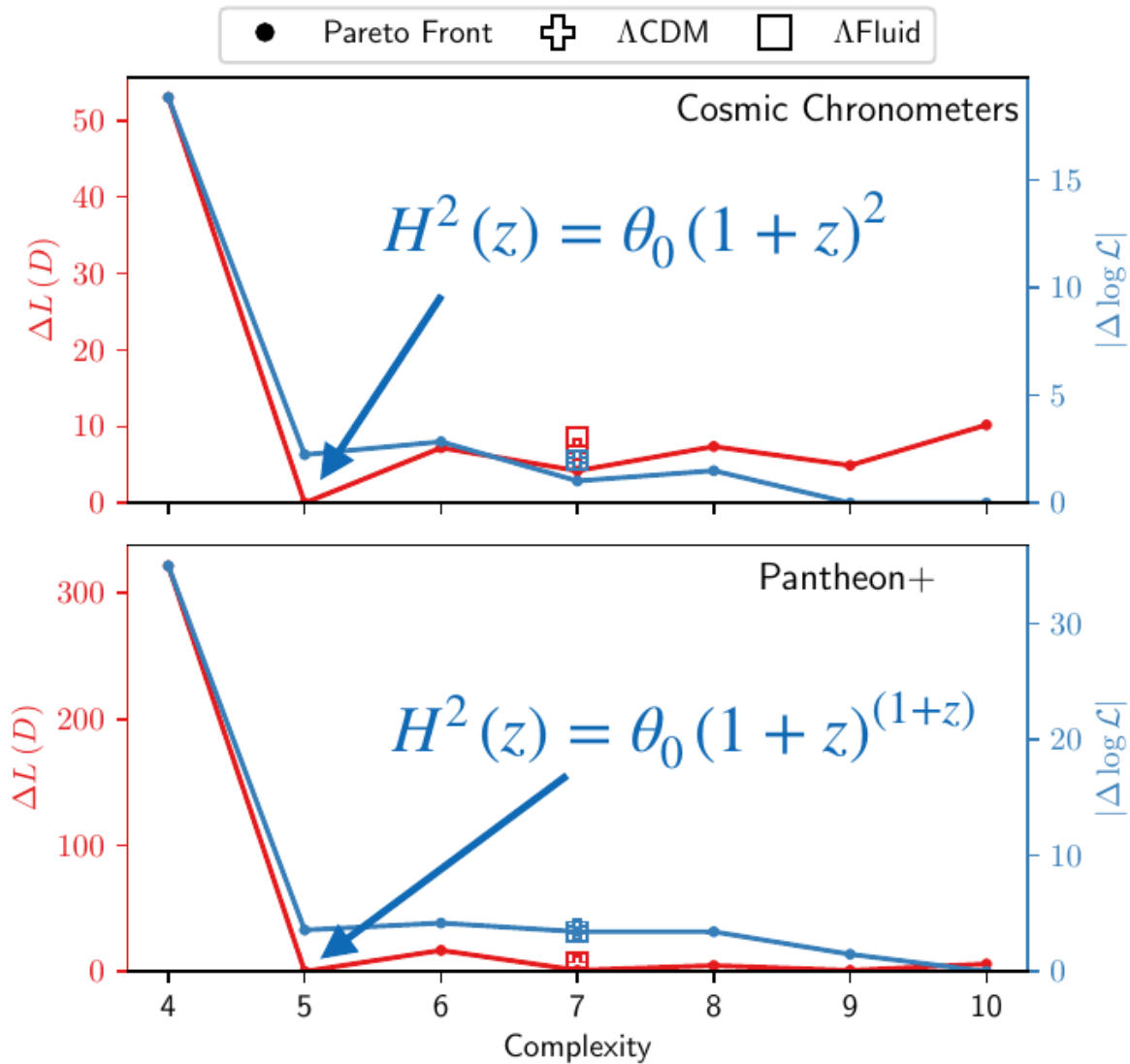
- Data:
  - Cosmic chronometers (32 data points) (Moresco et al 2022)
  - Type Ia Supernovae (1590 data points) (Pantheon+, Scolnic et al 2021)
- Basis operators:  $\{x \equiv 1 + z, \theta, \text{inv}, +, -, \times, \div, \text{pow}\}$

## Cosmic Chronometers



## Pantheon+ Supernovae





- $\Lambda$ CDM ranked 39<sup>th</sup> for cosmic chronometers and 37<sup>th</sup> for SNe

- Best functions approximate  $\Lambda$ CDM at low  $z$ , but are simpler

- ~200 functions (up to complexity 10) more accurate than  $\Lambda$ CDM for Pantheon+

# Test case 2: Potential of the inflaton

## What we know

$$A_S = (0.027 \pm 0.0027) M_{\text{Pl}},$$

$$n_S = 0.9649 \pm 0.0042,$$

$$r < 0.028 \text{ (95\% CL)}.$$

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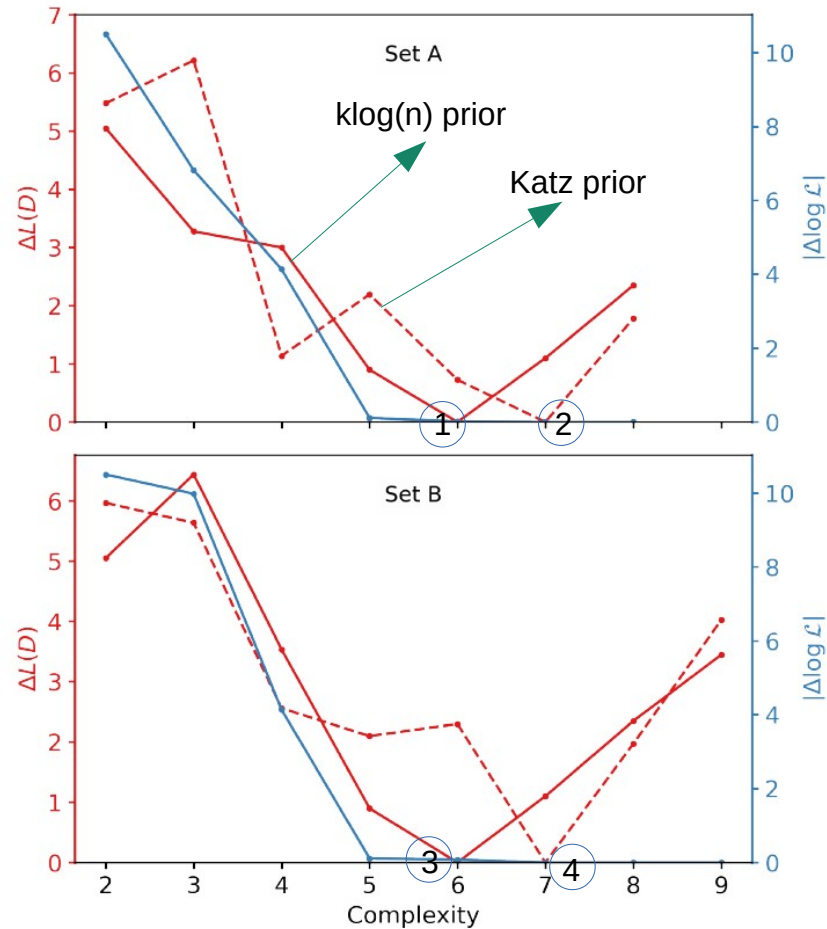
## Operator basis sets

A:  $\{x, a, \text{inv}, \text{exp}, \text{log}, \text{sin}, \sqrt{|\cdot|}, \text{cube}, \text{square}, +, *, -, /\}$

B:  $\{x, a, \text{inv}, \text{exp}, \text{log}, +, *, -, /, \text{power}\}$

## Best functions

- |                                         |  |                                            |
|-----------------------------------------|--|--------------------------------------------|
| 1. $e^{-e^{e^{e\phi}}}$                 |  | 3. $e^{-e^{e^{e\phi}}}$                    |
| 2. $\theta_0 (\theta_1 + \log(\phi)^2)$ |  | 4. $\theta_0 \phi^{\frac{\theta_1}{\phi}}$ |



Operator set B,  
klog(n) prior

Rank	$V(\phi)$	Complexity	Prediction		Description length			
			$n_s$	$r$	Residuals	Function	Parameter	Total
1	$e^{-e^{e^{\phi}}}$	6	0.9678	0.002	-12.65	6.59	0.00	-6.06
2	$\theta_0 e^{-e^{\phi}}$	5	0.9668	0.002	-12.79	6.93	0.69	-5.16
3	$ \theta_0 ^{e^{e^{\phi}}}$	5	0.9674	0.002	-12.72	6.93	0.69	-5.09
4	$e^{e^{\phi}} e^{-e^{e^{\phi}}}$	8	0.9685	0.002	-12.50	8.79	0.00	-3.71
5	$\theta_0 e^{-e^{e^{\phi}}}$	6	0.9680	0.002	-12.62	8.32	0.69	-3.61
6	$e^{-e^{\frac{1}{\phi}}}$	5	0.9768	0.008	-8.65	5.49	0.00	-3.16
7	$e^{\theta_0 e^{e^{\phi}}}$	6	0.9674	0.002	-12.72	8.32	1.36	-3.04
8	$ \theta_0 ^{e^{e^{e^{\phi}}}}$	6	0.9678	0.002	-12.65	8.32	1.39	-2.94
9	$e^{e^{\phi}} e^{-e^{e^{e^{\phi}}}}$	9	0.9685	0.002	-12.50	9.89	0.00	-2.62
10	$ \theta_0 ^{\frac{1}{\phi}}$	4	0.9756	0.019	-8.77	5.55	0.69	-2.53
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
1272	$\theta_0 \left(1 - e^{-\sqrt{2/3}\phi}\right)^2$	9	0.9678	0.003	-12.63	17.51	0.69	5.57
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
8697	$\theta_0 \phi^2$	4	0.9669	0.132	31.82	5.55	0.69	38.01
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
10839	$\theta_0 \phi^4$	5	0.9508	0.262	168.23	6.93	0.69	178.81

# Conclusions

- *Exhaustive Symbolic Regression*: Guaranteed to find best simple function for any data
- *Minimum description length* affords principled combination of accuracy and simplicity
- *A Katz language model* can assign function priors based on a training set

<https://github.com/DeaglanBartlett/ESR>

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So far we've learnt...

- 1 Cosmic chronometers and SNe don't uniquely favour Friedmann
- 2 Planck implies optimal inflationary potentials
- 3 The radial acceleration relation doesn't clearly support modified gravity

Extra Slides

# MDL as a Bayesian statistic

$$P(f_i|D) = \frac{1}{P(D)} \int P(D|f_i, \theta_i) P(\theta_i|f_i) P(f_i) d\theta_i \quad \log P(f_i|D) = -\log P(f_i) - \log \mathcal{Z}(D|f_i)$$

$$\log \mathcal{Z}(D|f_i) \simeq \log H(D, f_i, \hat{\theta}_i) + \frac{p}{2} \log 2\pi - \frac{1}{2} \log |\det \hat{I}^H| \quad \left( \hat{I}_{\alpha\beta}^H = -\partial_\alpha \partial_\beta \log H(D, f_i, \theta_i) |_{\hat{\theta}_i} \right)$$

## Functional Priors

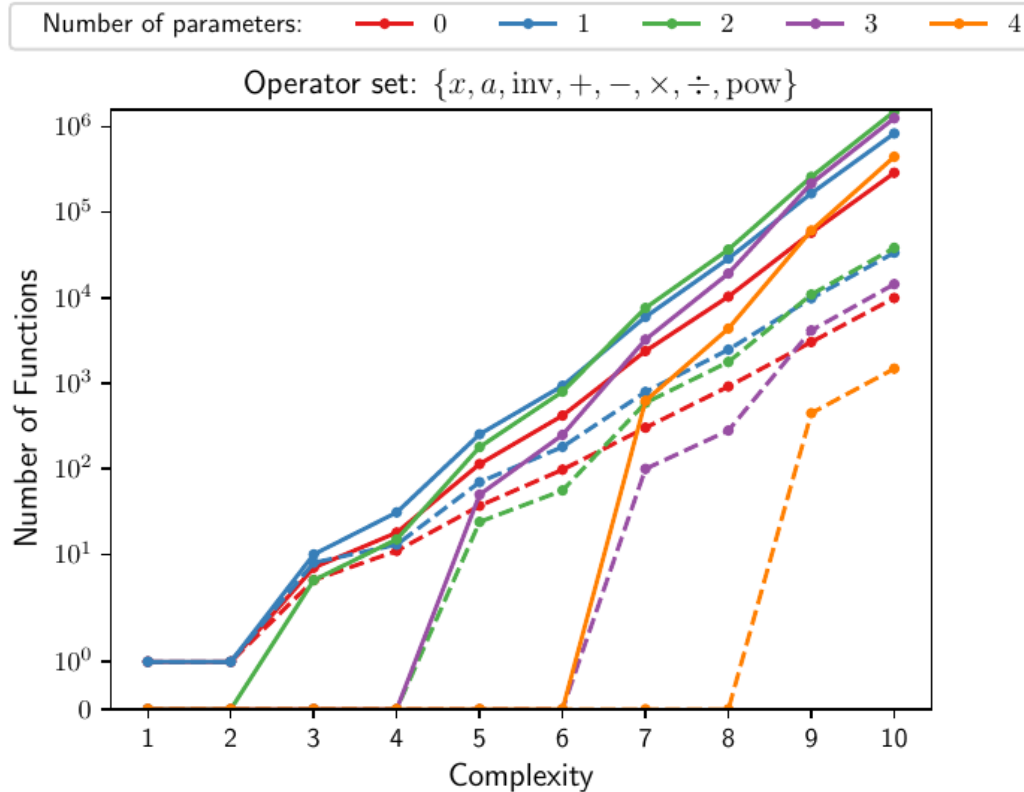
MDL implies  $-\log P(f_i) = k \log n + \sum_{\alpha} \log c_{\alpha}$

OR

Quantify our lower prior on e.g.  $\sin(\sin(x_0+x_1))$  compared to  $\sin(x_0)+\sin(x_1)$

“Katz back-off model” determines probability of next operator given  $n$  preceding operators based on a training set of equations

# Simplifications make an exhaustive search feasible



**168 million**

Naïve estimate:  $\sum_{j=1}^n j^k = H_n^{(-k)}$

**5.2 million**

Total number of decorated trees

**134,234**

Number of unique equations

**1400**

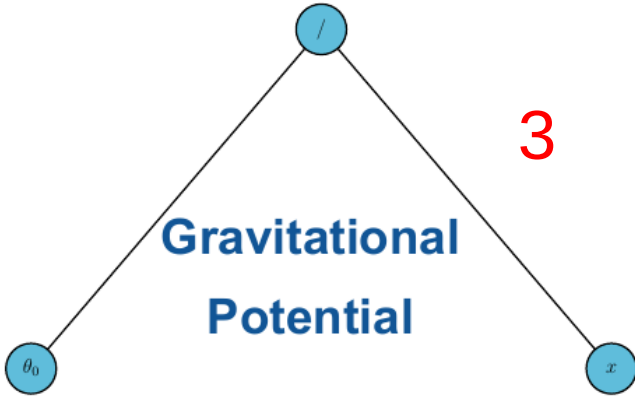
Times fewer equations to consider than our naïve guess

**119,861**

Number of equations containing at least one parameter

# Many physics functions have complexity < 10

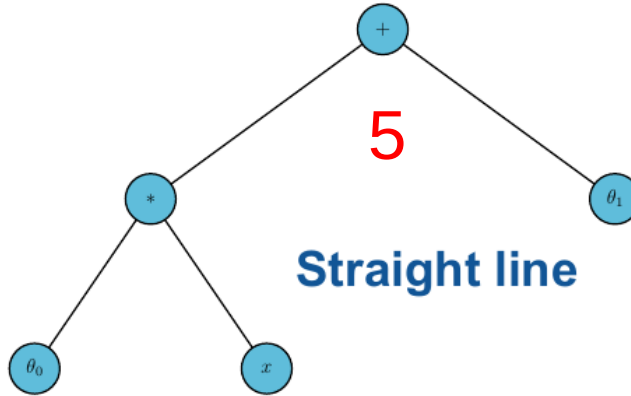
$$\frac{\theta_0}{x} = [/, \theta_0, x]$$



3

**Gravitational  
Potential**

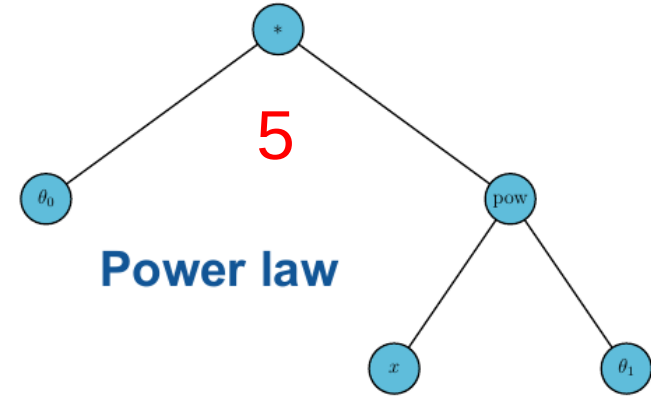
$$\theta_0 x + \theta_1 = [+ , * , \theta_0 , x , \theta_1]$$



5

**Straight line**

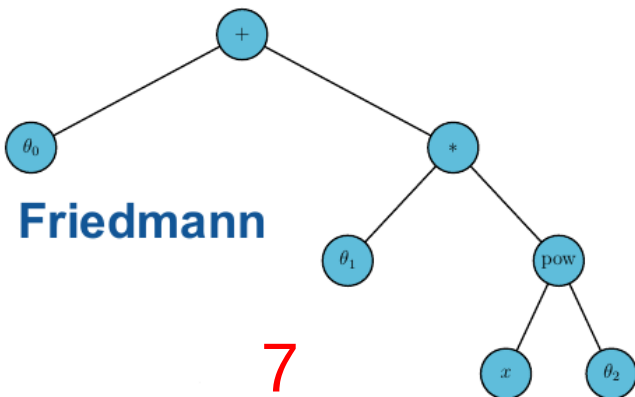
$$\theta_0 x^{\theta_1} = [* , \theta_0 , \text{pow} , x , \theta_1]$$



5

**Power law**

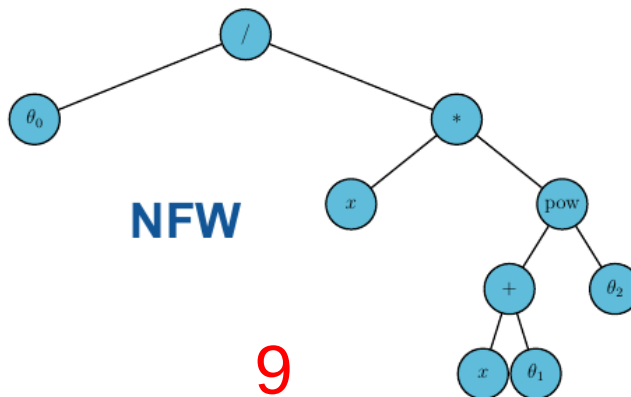
$$\theta_0 + \theta_1 x^{\theta_2} = [+ , \theta_0 , * , \theta_1 , \text{pow} , x , \theta_2]$$



7

**Friedmann**

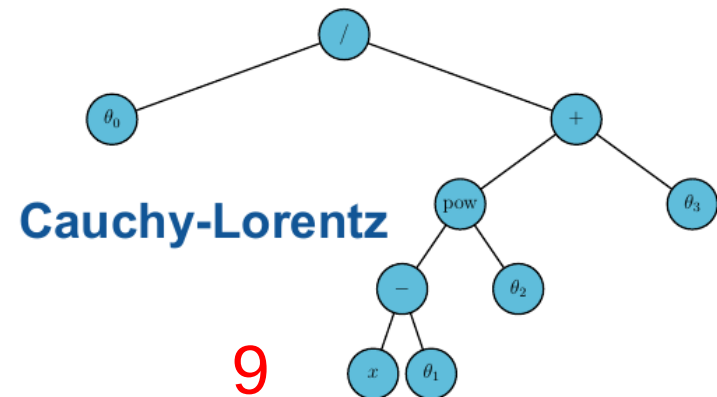
$$\frac{\theta_0}{x(x+\theta_1)^{\theta_2}} = [/, \theta_0 , * , x , \text{pow} , + , x , \theta_1 , \theta_2]$$



9

**NFW**

$$\frac{\theta_0}{(x-\theta_1)^{\theta_2+\theta_3}} = [/, \theta_0 , + , \text{pow} , - , x , \theta_1 , \theta_2 , \theta_3]$$



9

**Cauchy-Lorentz**

## Precision of constants - tradeoff between accuracy and information needed

True  $\hat{\theta}_i$  uniformly distributed within  $\pm\Delta_i/2$  of transmitted value - may not transmit true MLP

Taylor expand log-likelihood about true value:

$$-\log(\mathcal{L}(\hat{\theta} + \mathbf{d})) \approx -\log(\mathcal{L}(\hat{\theta})) + \frac{1}{2} \mathbf{d}^T \mathbf{I} \mathbf{d} \quad \mathbf{I}_{ij} = \left. \frac{d^2(-\log \mathcal{L})}{d\theta_i d\theta_j} \right|_{\hat{\theta}}$$

Giving expected contribution to description length

$$L(\Delta) = \frac{1}{2} \sum_{ij} \langle \mathbf{I}_{ij} d_i d_j \rangle - \sum_i \log(\Delta_i)$$

Minimise this:

$$L(\Delta_i) = \frac{1}{24} \mathbf{I}_{ii} \Delta_i^2 - \log(\Delta_i) \quad \Rightarrow \quad \Delta_i = \left( \frac{12}{\mathbf{I}_{ii}} \right)^{1/2}$$

## ■ The Description Length of a Function

$$L(D) = -\log(\mathcal{L}(\hat{\theta})) + k \log(n) - \frac{p}{2} \log(3) + \sum_i^p \left( \frac{1}{2} \log(\mathbf{I}_{ii}) + \log(|\hat{\theta}_i|) \right)$$

$$\frac{1}{2} \left( p \log(N) - 2 \log(\mathcal{L}) \right) = \frac{1}{2} \text{BIC} \quad (\text{for large number of data points, } N)$$

- Description length looks like BIC plus corrections due to structural complexity (prior on model)
- For large  $N$ , equivalent to minimising the BIC (an approximation to the evidence)

$$P(H) = \exp(-L(D)) / \sum(\exp(-L(D)))$$

## Cosmic Chronometers - Standard Clocks

$$H(z) = -\frac{1}{1+z} \frac{dz}{dt} \approx -\frac{1}{1+z} \frac{\delta z}{\delta t}$$



Image credit: AAS NOVA

- Passively evolving stellar population are standard clocks - measure  $\delta t$
- Directly measure  $\delta z$
- (Cosmological) model-independent measurement of  $H(z)$
- We use sample of 32 CC  $H(z)$  measurements

Jimenez & Loeb 2001 (arXiv:astro-ph/0106145)  
Moresco 2015 (arXiv:1503.01116)  
Moresco et al. 2016 (arXiv:1601.01701)  
Ratsimbazafy et al. 2017 (arXiv:1702.00418)  
Stern et al. 2010 (arXiv:0907.3149)  
Simon et al. 2004 (arXiv:astro-ph/0412269)  
Borghi et al. 2021 (arXiv:2110.04304)  
Zhang et al. 2012 (arXiv:1207.4541)  
Moresco et al. 2012 (arXiv:1201.3609)  
Moresco et al. 2022 (arXiv:2201.07241)



## Type Ia Supernovae - Standard Candles

$$d_L(z) = (1+z) \int_0^z \frac{dz'}{H(z')}$$

$$\mu(z) = 5 \log_{10} \left( \frac{d_L(z)}{10 \text{ pc}} \right)$$

$$\mu = m_B + \alpha x_1 - \beta c - M_0$$

**Amplitude** **Stretch** **Colour** **Rest-frame magnitude**

Pantheon+ sample with SH0ES Cepheid calibration

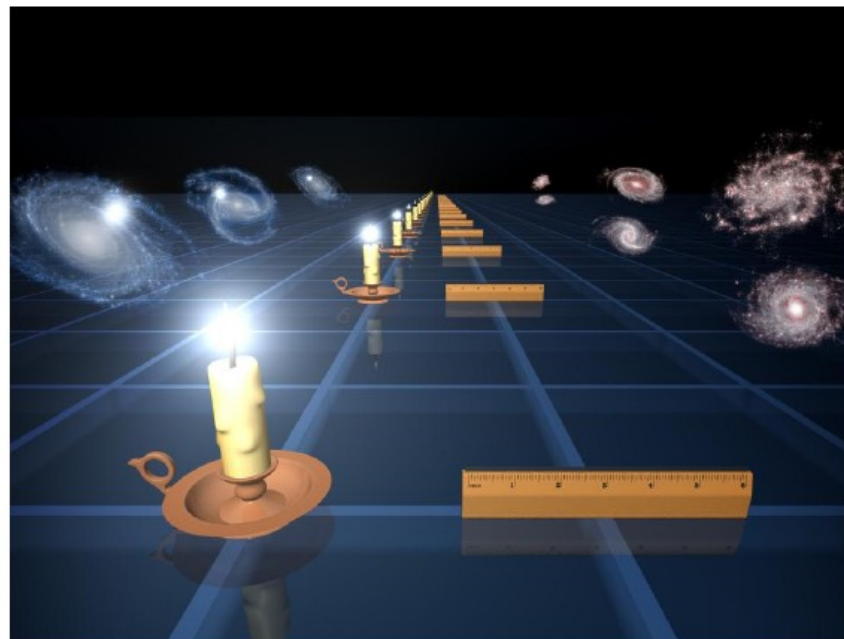


Image credit: NASA

Scolnic et al. 2021 (arXiv:2112.03863)  
Riess et al. 2022 (arXiv:2112.04510)

CCs

Rank	$y(x) / \text{km}^2\text{s}^{-2}\text{Mpc}^{-2}$	Complexity	Parameters			Codelength			
			$\theta_0$	$\theta_1$	$\theta_2$	Residuals <sup>1</sup>	Function <sup>2</sup>	Parameter <sup>3</sup>	Total
1	$\theta_0 x^2$	5	3883.44	-	-	8.36	5.49	2.53	16.39
2	$ \theta_0  x^{\theta_1}$	5	3982.43	0.22	-	7.97	5.49	5.24	18.70
3	$\theta_0  \theta_1 ^{-x}$	5	1414.43	0.31	-	7.57	6.93	5.58	20.08
4	$\theta_0 x^{\theta_1}$	5	3834.51	2.03	-	8.35	6.93	5.08	20.36
5	$x^2 (\theta_0 + x)$	7	3881.85	-	-	8.36	9.70	2.53	20.60
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
39	$\theta_0 + \theta_1 x^3$	9	3164.02	1481.71	-	7.28	12.48	3.76	23.51
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
84	$\theta_0 + \theta_1 x^{\theta_2}$	7	3322.96	1374.97	3.08	7.27	11.27	6.52	25.06

<sup>1</sup>  $-\log \mathcal{L}(\hat{\theta})$      
<sup>2</sup>  $k \log(n) + \sum_j \log(c_j)$      
<sup>3</sup>  $-\frac{p}{2} \log(3) + \sum_i^p (\frac{1}{2} \log(I_{ii}) + \log(|\hat{\theta}_i|))$

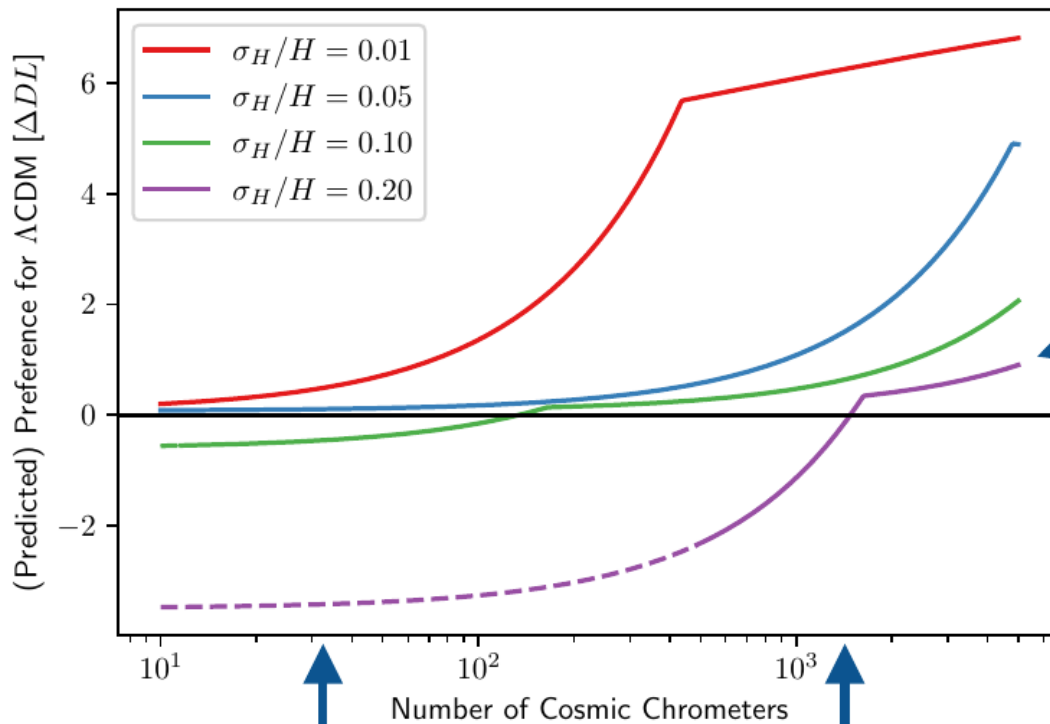
SNe

Rank	$y(x) / \text{km}^2\text{s}^{-2}\text{Mpc}^{-2}$	Complexity	Parameters			Codelength			
			$\theta_0$	$\theta_1$	$\theta_2$	Residuals <sup>1</sup>	Function <sup>2</sup>	Parameter <sup>3</sup>	Total
1	$\theta_0 x^x$	5	5345.02	-	-	706.18	6.93	5.11	718.22
2	$ \theta_0  x^{\theta_1}$	9	5280.11	0.16	-	705.11	5.49	8.41	719.01
3	$\theta_0  \theta_1 ^{-x}$	5	1694.95	0.32	-	701.79	6.93	10.33	719.05
4	$\theta_0 x^{\theta_1}$	7	5378.69	0.78	-	702.45	9.70	6.98	719.13
5	$ \theta_0   \theta_1 ^x$	5	1898.47	1.14	-	701.88	5.49	12.64	720.02
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
37	$\theta_0 + \theta_1 x^3$	9	3591.09	1773.63	-	701.85	12.48	8.81	723.13
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
96	$\theta_0 + \theta_1 x^{\theta_2}$	7	3280.83	2069.32	2.73	701.64	11.27	12.19	725.10

<sup>1</sup>  $-\log \mathcal{L}(\hat{\theta})$      
<sup>2</sup>  $k \log(n) + \sum_j \log(c_j)$      
<sup>3</sup>  $-\frac{p}{2} \log(3) + \sum_i^p (\frac{1}{2} \log(I_{ii}) + \log(|\hat{\theta}_i|))$

# Should we have seen $\Lambda$ CDM? No.

Mock cosmic  
chronometer  
data assuming  
 $\Lambda$ CDM  
(Planck18)



**Current  
observational  
uncertainties**

**We used  
this many**

**Would need  
this many**

# Test case 2: The radial acceleration relation

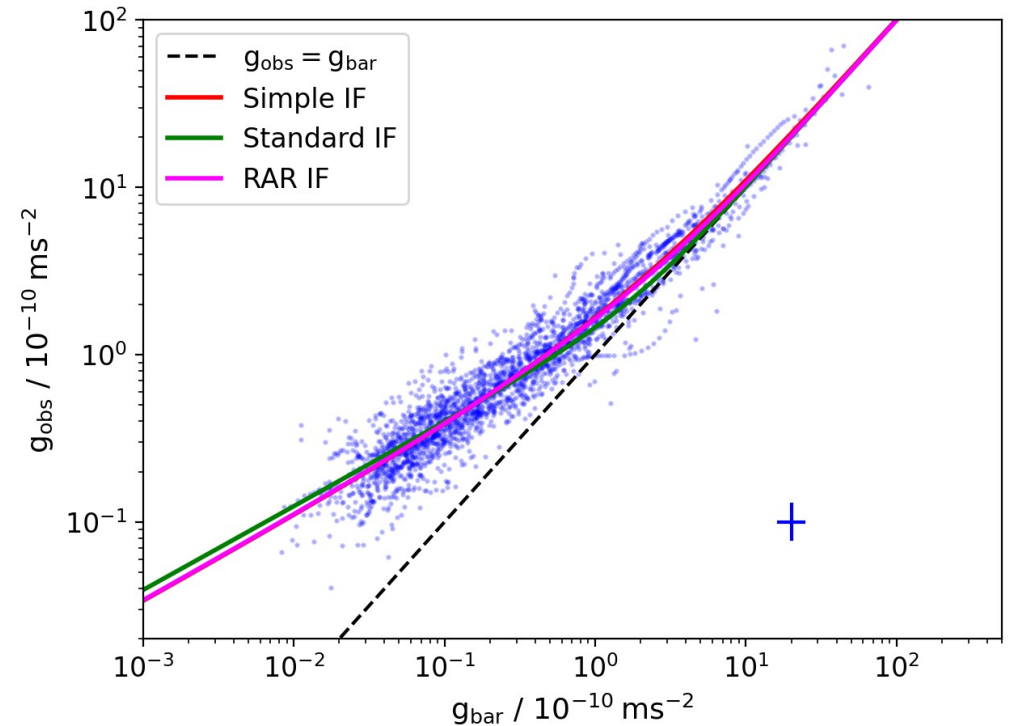
- Relates acceleration sourced by baryons ( $g_{\text{bar}}$ ) to total acceleration as measured by rotation velocity ( $g_{\text{obs}}$ )
- 2,696 points from 153 late-type galaxies (SPARC sample)
- Regularity and low scatter hard to understand in  $\Lambda$ CDM

## MOND Interpolating Functions (IFs)

**Simple** —  $g_{\text{obs}} = g_{\text{bar}}/2 + \sqrt{g_{\text{bar}}^2/4 + g_{\text{bar}} a_0}$

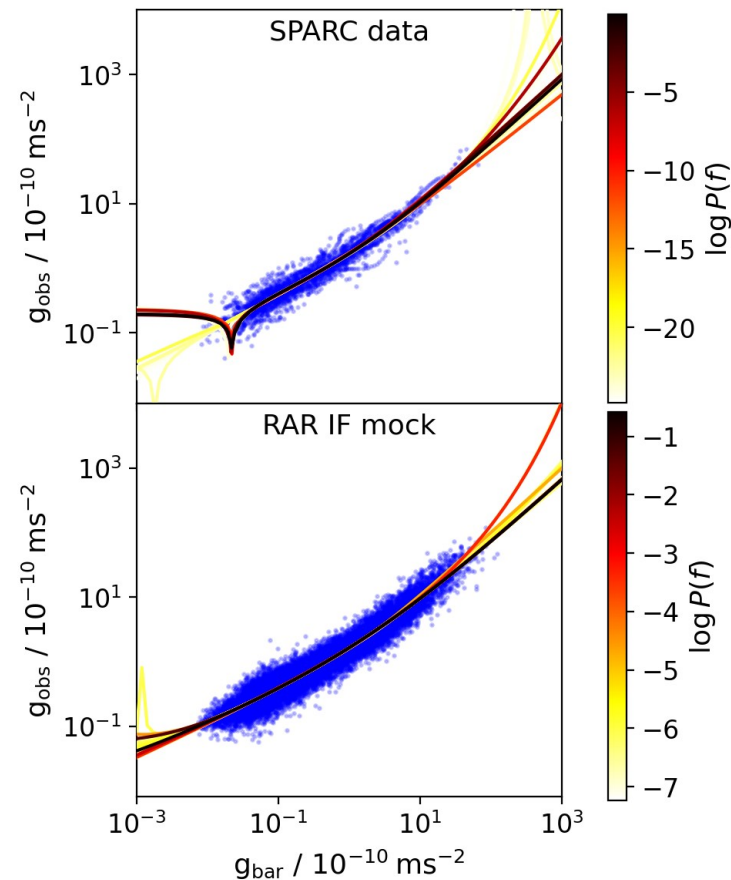
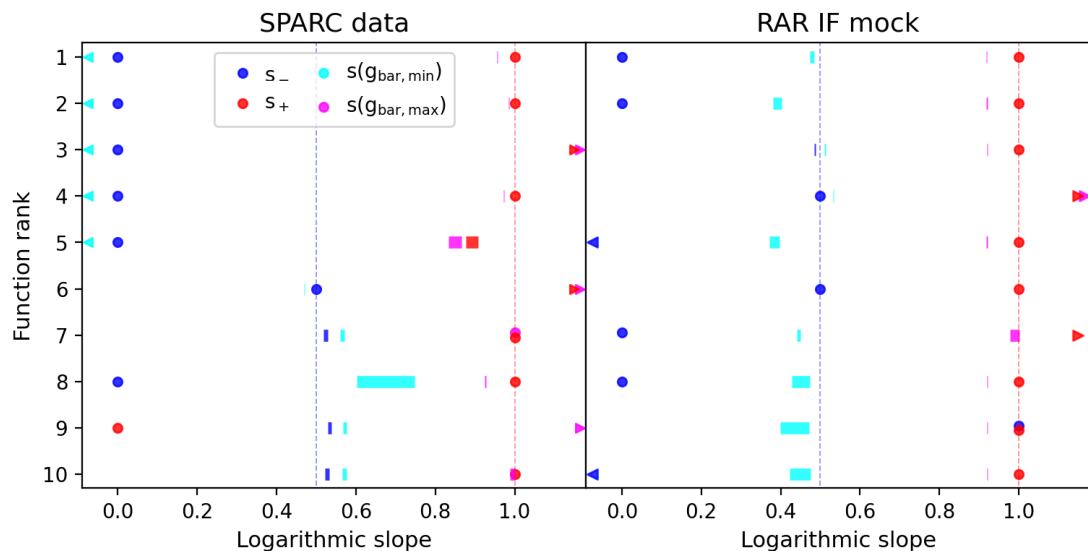
**Standard** —  $g_{\text{obs}} = \frac{1}{\sqrt{2}} \sqrt{g_{\text{bar}}^2 + \sqrt{g_{\text{bar}}^2 (g_{\text{bar}}^2 + 4g_0^2)}}$

**RAR** —  $g_{\text{obs}} = g_{\text{bar}} / (1 - \exp(-\sqrt{g_{\text{bar}}/g_0}))$

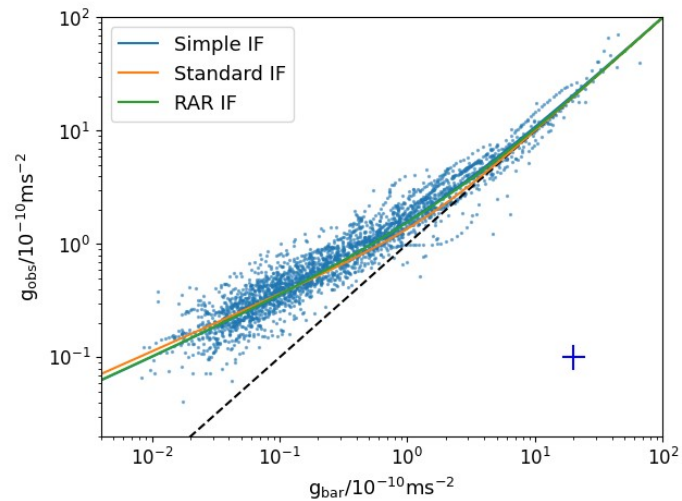


1) Are the MOND IFs optimal descriptions of the RAR?

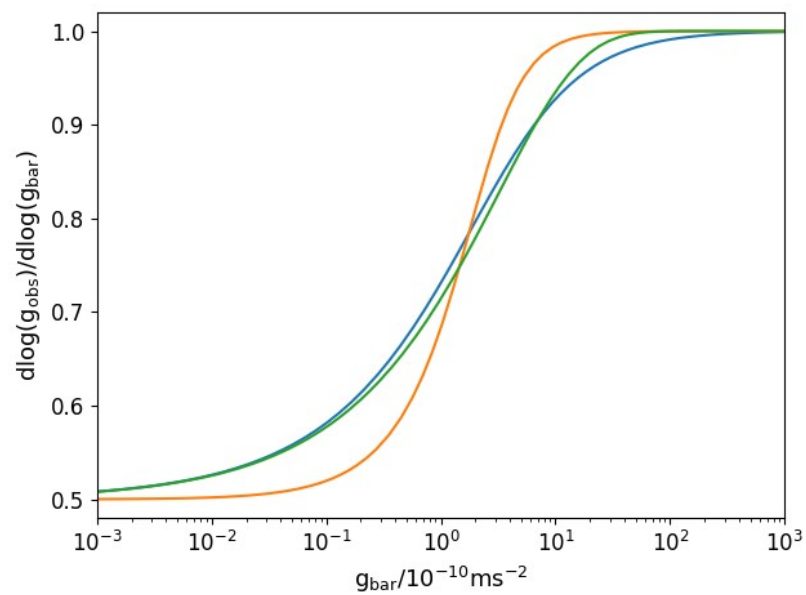
2) Do optimal solutions satisfy the MOND limits (and hence may be considered new IFs)?



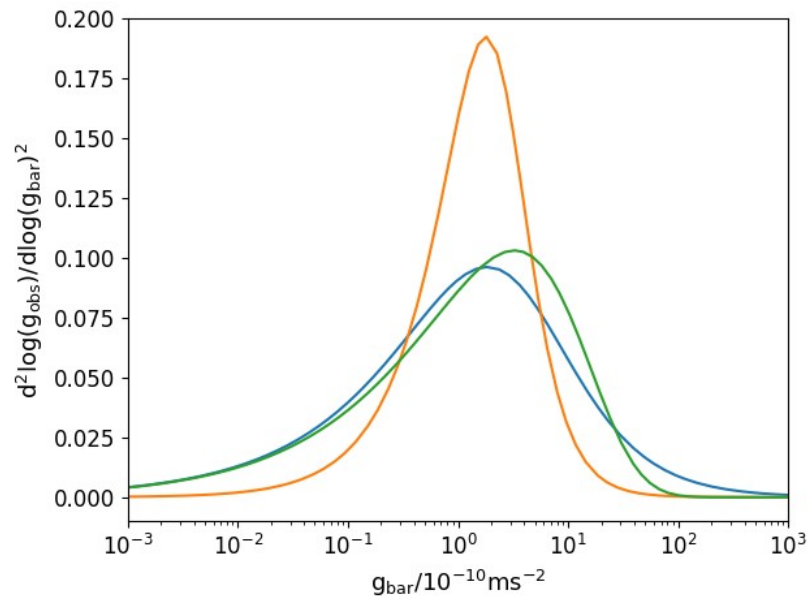
- Newtonian limit often found; deep-MOND limit rarely
- Can't recover MOND behaviour even from MOND mocks!  
⇒ Precision and dynamic range of data insufficient

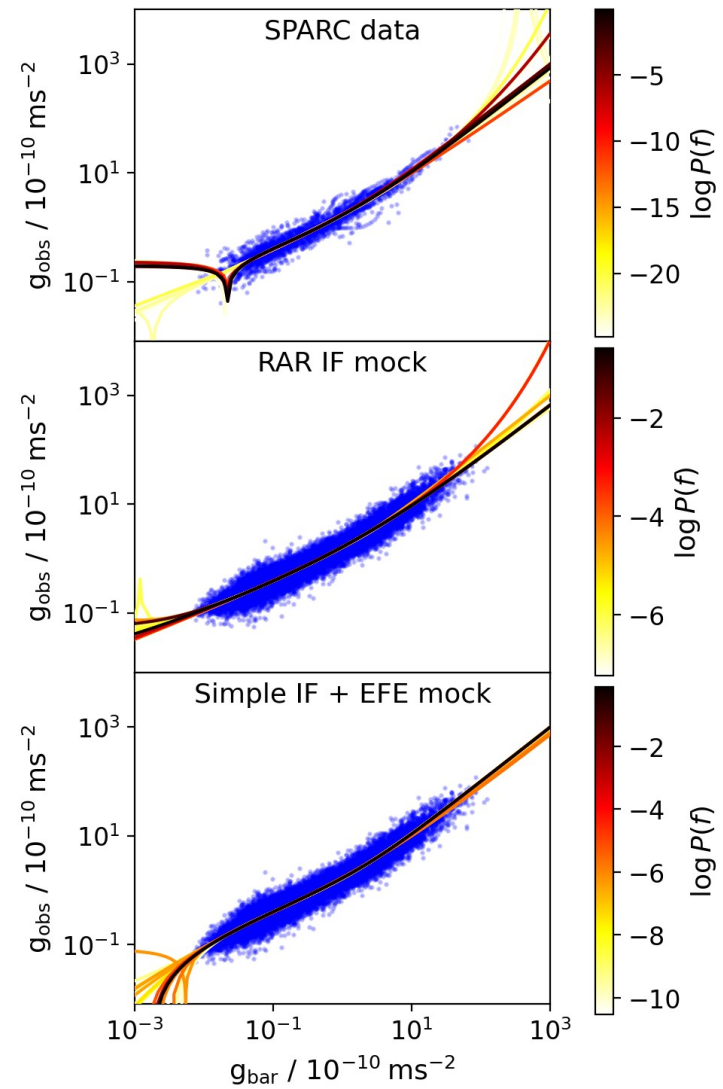
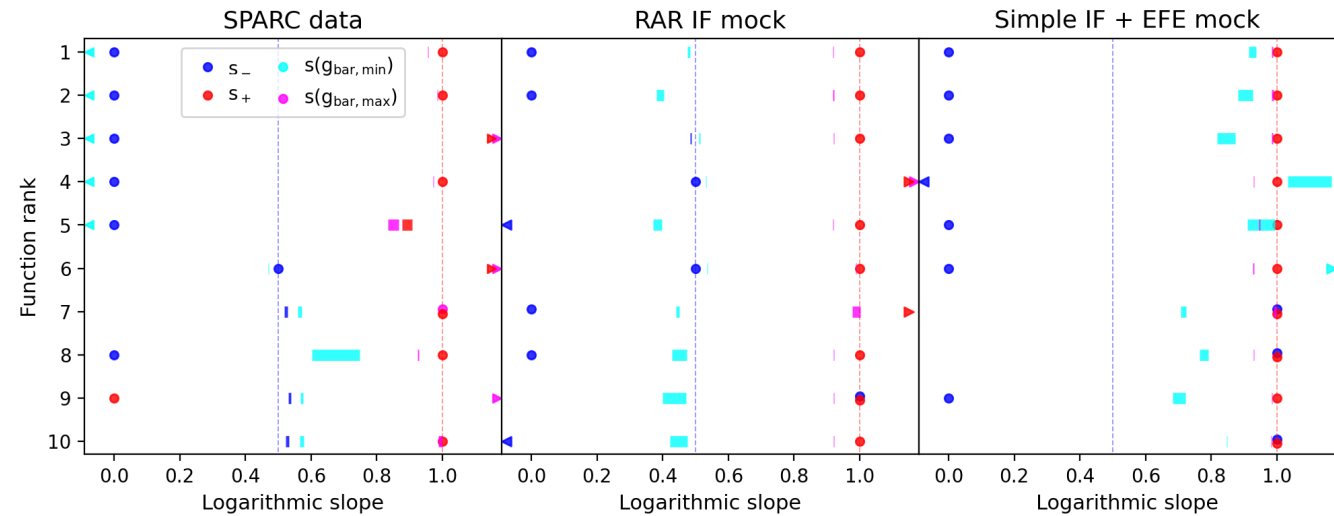


First  
logarithmic  
derivative



Second  
logarithmic  
derivative





**SPARC**  
**data**

Rank	Function	Comp.	$P(f)$	Parameters					Description length		
				$\theta_0$	$\theta_1$	$\theta_2$	$\theta_3$	Resid. <sup>1</sup>	Func. <sup>2</sup>	Param. <sup>3</sup>	Total
1	$\theta_0 ( \theta_1 + x ^{\theta_2} + x)$	9	$9.3 \times 10^{-1}$	0.84	-0.02	0.38	—	-1279.1	14.5	14.0	-1250.6
2	$  \theta_1 ^x + \theta_0 ^{\theta_2} + x$	9	$6.4 \times 10^{-2}$	-0.99	0.64	0.36	—	-1279.9	12.5	19.6	-1247.9
3	$ \theta_0 ^{ \theta_1 - x ^{\theta_2} - \theta_3}$	9	$2.0 \times 10^{-3}$	$-1.4 \times 10^2$	0.02	0.14	0.89	-1276.4	12.5	19.5	-1244.4
4	$ \theta_0(\theta_1 + x) ^{\theta_2} + x$	9	$1.4 \times 10^{-4}$	0.35	-0.02	0.34	—	-1268.9	14.5	12.7	-1241.7
5	$ \theta_0 -  \theta_1 - x ^{\theta_2} ^{\theta_3}$	9	$1.0 \times 10^{-5}$	-0.30	0.02	0.42	2.14	-1271.1	12.5	19.5	-1239.1
6	$\sqrt{x} \exp\left(\frac{ \theta_0 + x ^{\theta_1}}{2}\right)$	9	$1.5 \times 10^{-9}$	-0.02	0.36	—	—	-1257.9	17.5	10.0	-1230.3
7	$\left(\frac{ \theta_0 ^x}{x}\right)^{\theta_1} + x$	9	$2.4 \times 10^{-10}$	1.87	-0.52	—	—	-1250.6	14.5	7.6	-1228.5
8	$\sqrt{ \theta_0 + x } + \theta_1 x$	8	$1.8 \times 10^{-10}$	$-1.8 \times 10^{-3}$	0.72	—	—	-1245.6	12.9	4.5	-1228.2
9	$\left \theta_0 + \frac{1}{\sqrt{x}}\right ^{\theta_1}$	8	$9.6 \times 10^{-11}$	-0.22	-2.14	—	—	-1251.1	14.3	9.2	-1227.6
10	$(\sqrt{x} + \frac{1}{x})^{\theta_0} + x$	9	$8.2 \times 10^{-11}$	-0.53	—	—	—	-1248.3	16.1	4.8	-1227.4
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
17	$x/(\exp(\theta_0) -  \theta_1 ^{\sqrt{x}})$	9	$2.2 \times 10^{-11}$	0.03	0.44	—	—	-1250.9	17.5	7.3	-1226.1
—	Double power law	11	$9.7 \times 10^{-16}$	4.65	3.96	1.03	0.60	-1252.3	17.7	18.5	-1216.1
—	Simple IF	10	$5.5 \times 10^{-25}$	1.11	—	—	—	-1217.3	18.6	3.9	-1194.8
—	RAR IF	9	$6.7 \times 10^{-26}$	1.13	—	—	—	-1212.8	16.1	3.9	-1192.7
—	Simple IF + EFE	59	$5.0 \times 10^{-69}$	1.16	$6.8 \times 10^{-3}$	—	—	-1238.9	139.9	5.6	-1093.4
—	Standard IF	14	$9 \times 10^{-150}$	1.54	—	—	—	-939.5	27.9	4.1	-907.5

$$^1 - \log \mathcal{L}(\hat{\theta})$$

$$^2 k \log(n) + \sum_j \log(c_j)$$

$$^3 - \frac{p}{2} \log(3) + \sum_i^p (\log(I_{ii})^{1/2} + \log(|\hat{\theta}_i|))$$



RAR IF  
mock

Rank	Function	Comp.	$P(f)$	Parameters				Description length			
				$\theta_0$	$\theta_1$	$\theta_2$	$\theta_3$	Resid. <sup>1</sup>	Func. <sup>2</sup>	Param. <sup>3</sup>	Total
1	$\theta_0 + \theta_1 x + \sqrt{x}$	8	$5.6 \times 10^{-1}$	$9.1 \times 10^{-3}$	0.63	—	—	-2045.2	12.9	4.9	-2027.4
2	$\sqrt{ \theta_0 + x } + \theta_1 x$	8	$2.8 \times 10^{-1}$	$3.0 \times 10^{-3}$	0.64	—	—	-2044.4	12.9	4.8	-2026.7
3	$\theta_0 x + x^{\theta_1}$	7	$8.2 \times 10^{-2}$	0.64	0.49	—	—	-2045.2	11.3	8.5	-2025.5
4	$\sqrt{x} \exp\left(\frac{x^{\theta_0}}{2}\right)$	7	$3.5 \times 10^{-2}$	0.36	—	—	—	-2040.7	12.5	3.5	-2024.7
5	$(\theta_0 + x)\left(\theta_1 + \frac{1}{\sqrt{x}}\right)$	9	$1.1 \times 10^{-2}$	$1.3 \times 10^{-3}$	0.64	—	—	-2044.5	16.1	4.8	-2023.5
6	$\frac{1}{\sqrt{ \theta_0 + \frac{1}{x} }} + x$	8	$8.8 \times 10^{-3}$	1.74	—	—	—	-2038.5	12.9	2.3	-2023.3
7	$(x \theta_0 )^{(x \theta_1 )^{\theta_2}}$	9	$3.1 \times 10^{-3}$	-2.09	$-1.4 \times 10^{-4}$	0.04	—	-2045.3	12.5	10.6	-2022.2
8	$\theta_0 x +  \theta_1 + x ^{\theta_2}$	9	$2.4 \times 10^{-3}$	0.64	$1.4 \times 10^{-3}$	0.49	—	-2045.4	14.5	8.9	-2022.0
9	$x( \theta_0 - x ^{\theta_1} - \theta_2)$	9	$2.3 \times 10^{-3}$	$1.2 \times 10^{-3}$	-0.51	-0.64	—	-2045.3	14.5	8.9	-2021.9
10	$(\theta_0 - x)(\theta_1 - x^{\theta_2})$	9	$2.2 \times 10^{-3}$	$-6.5 \times 10^{-4}$	-0.64	-0.51	—	-2045.4	14.5	9.0	-2021.9
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
27	$x/(\exp(\theta_0) - \exp(-\sqrt{x}))$	9	$3.2 \times 10^{-4}$	-0.01	—	—	—	-2039.3	17.5	1.9	-2020.0
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
41	$x/(\exp(\theta_0) -  \theta_1 \sqrt{x})$	9	$1.1 \times 10^{-4}$	$-5.0 \times 10^{-3}$	0.38	—	—	-2042.1	17.5	5.7	-2018.9
—	RAR IF	9	$1.0 \times 10^{-3}$	1.14	—	—	—	-2041.1	16.1	3.9	-2021.1
—	Double power law	11	$3.4 \times 10^{-8}$	1.25	1.47	0.90	0.54	-2047.2	17.7	18.7	-2010.8
—	Simple IF	10	$2.8 \times 10^{-11}$	1.12	—	—	—	-2026.2	18.6	3.9	-2003.7
—	Standard IF	14	$2.9 \times 10^{-55}$	1.54	—	—	—	-1934.4	27.9	4.1	-1902.4
—	Simple IF + EFE	59	$5.9 \times 10^{-64}$	1.12	0	—	—	-2026.2	139.9	3.9	-1882.4

$$^1 - \log \mathcal{L}(\hat{\theta})$$

$$^2 k \log(n) + \sum_j \log(c_j)$$

$$^3 - \frac{p}{2} \log(3) + \sum_i^p (\log(I_{ii})^{1/2} + \log(|\hat{\theta}_i|))$$

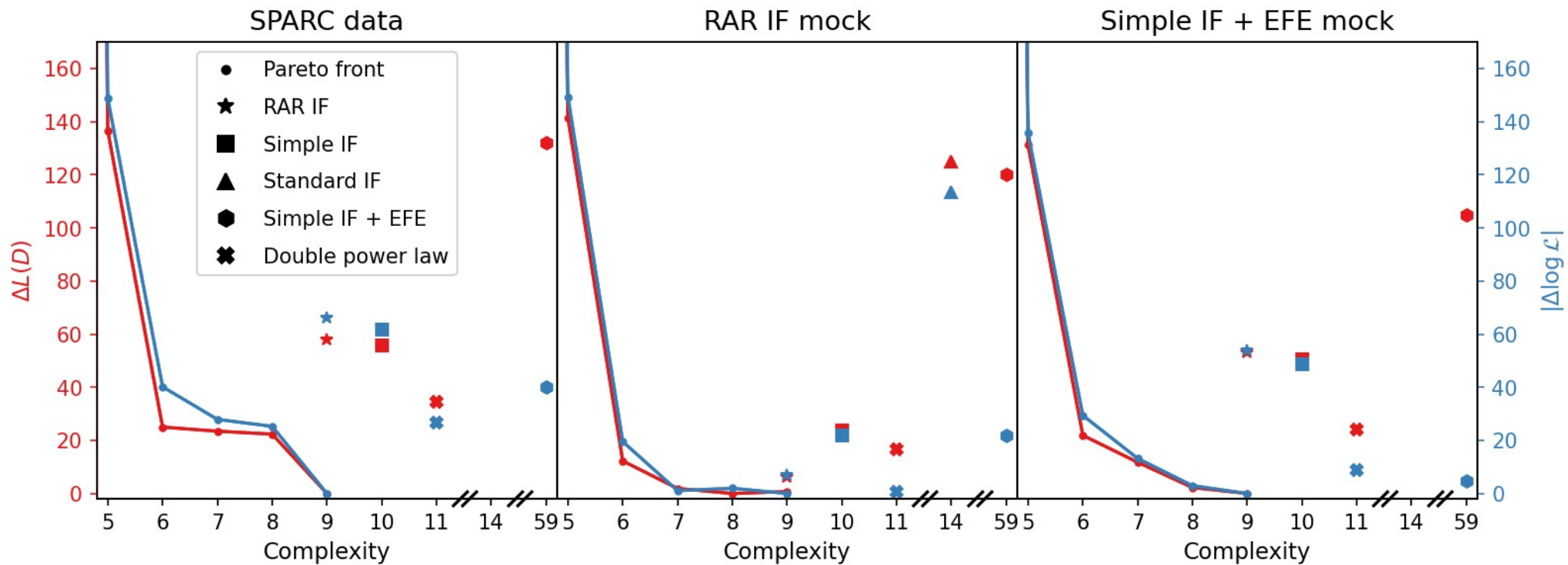
Simple  
IF +  
EFE  
mock

Rank	Function	Comp.	$P(f)$	Parameters					Description length		
				$\theta_0$	$\theta_1$	$\theta_2$	$\theta_3$	Resid. <sup>1</sup>	Func. <sup>2</sup>	Param. <sup>3</sup>	Total
1	$\theta_0 + \sqrt{x^2 + 2x}$	9	$8.9 \times 10^{-1}$	-0.06	—	—	—	-2017.7	14.5	3.1	-2000.0
2	$\theta_0 + \sqrt{x \theta_1 + x }$	8	$9.3 \times 10^{-2}$	-0.06	1.97	—	—	-2017.9	12.9	7.3	-1997.8
3	$- \theta_0 ^{\sqrt{x}} + \theta_1 + x$	8	$5.6 \times 10^{-3}$	0.26	0.95	—	—	-2017.9	12.9	10.1	-1995.0
4	$(\theta_0 - x)(\theta_1 - x^{\theta_2})$	9	$3.3 \times 10^{-3}$	$3.1 \times 10^{-3}$	-0.71	-0.53	—	-2019.7	14.5	10.7	-1994.4
5	$x^{\theta_0} - \theta_1(\theta_2 - x)$	9	$2.4 \times 10^{-3}$	0.39	0.79	0.12	—	-2020.9	14.5	12.3	-1994.1
6	$ \theta_0 - x ^{\theta_1} - \theta_2 x$	9	$2.0 \times 10^{-3}$	$5.5 \times 10^{-3}$	0.48	-0.71	—	-2019.1	14.5	10.6	-1994.0
7	$x \theta_0 ^{- \theta_1 ^{x^{\theta_2}}}$	9	$1.7 \times 10^{-3}$	0.04	-0.16	0.33	—	-2018.1	12.5	11.9	-1993.8
8	$x(\theta_0 +  \theta_1 + x ^{\theta_2})$	9	$1.5 \times 10^{-3}$	0.71	0.01	-0.53	—	-2018.7	14.5	10.6	-1993.7
9	$ \theta_0 ^{ \theta_1 ^{x^{\theta_2}}} + x$	9	$6.5 \times 10^{-4}$	$7.0 \times 10^{-6}$	0.03	0.17	—	-2016.7	12.5	11.4	-1992.8
10	$\exp\left(\theta_0 - \frac{1}{\sqrt[3]{x}}\right) + x$	9	$5.5 \times 10^{-4}$	0.57	—	—	—	-2014.0	17.5	3.9	-1992.6
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
21	$x/(\exp(\theta_0) -  \theta_1 ^{\sqrt{x}})$	9	$1.8 \times 10^{-5}$	0.03	0.44	—	—	-2014.2	17.5	7.4	-1989.3
—	Double power law	11	$3.4 \times 10^{-11}$	3.53	3.31	0.98	0.60	-2012.3	17.7	18.6	-1976.0
—	Simple IF	10	$1.2 \times 10^{-22}$	1.11	—	—	—	-1972.1	18.6	3.9	-1949.6
—	RAR IF	9	$7.0 \times 10^{-24}$	1.13	—	—	—	-1966.9	16.1	3.9	-1946.8
—	Simple IF + EFE	59	$3.8 \times 10^{-57}$	1.19	$8.6 \times 10^{-3}$	—	—	-2016.0	139.9	5.9	-1870.2
—	Standard IF	14	$2 \times 10^{-141}$	1.54	—	—	—	-1708.3	27.9	4.1	-1676.3

$$^1 - \log \mathcal{L}(\hat{\theta})$$

$$^2 k \log(n) + \sum_j \log(c_j)$$

$$^3 - \frac{p}{2} \log(3) + \sum_i^p (\log(I_{ii})^{1/2} + \log(|\hat{\theta}_i|))$$

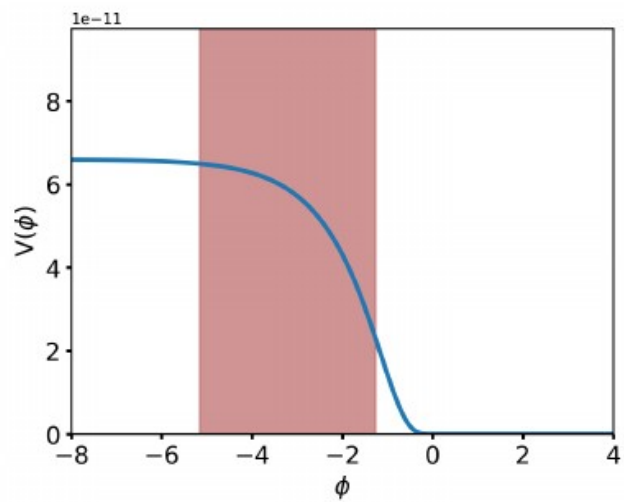


$$\theta_0 \left( |\theta_1 + x|^{\theta_2} + x \right)$$

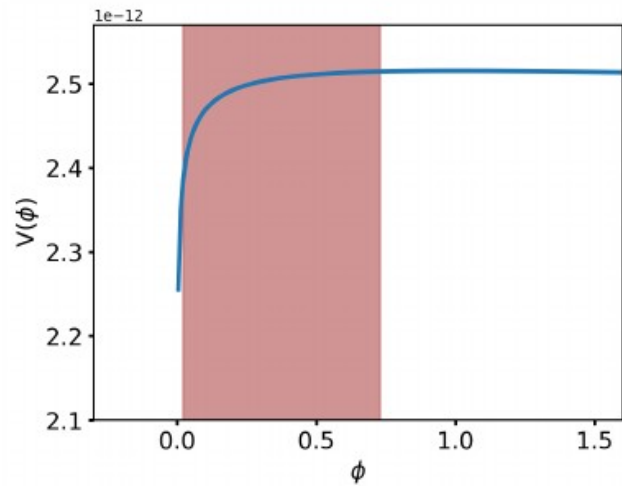
$$\theta_0 + \theta_1 x + \sqrt{x}$$

$$\theta_0 + \sqrt{x^2 + 2x}$$

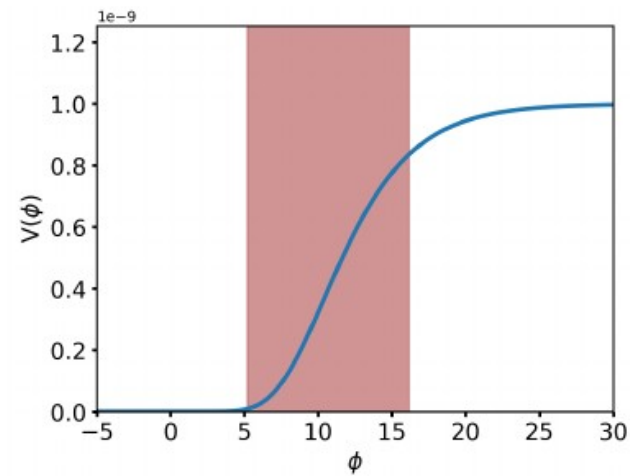
ESR readily Pareto-dominates all literature fits



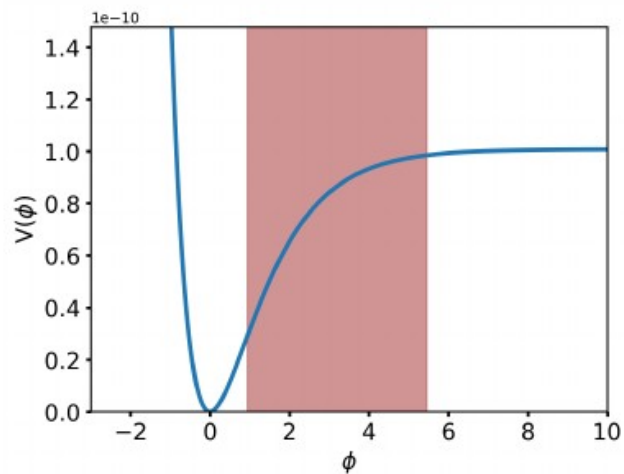
(a)  $V(\phi) = e^{-e^{e\phi}}$



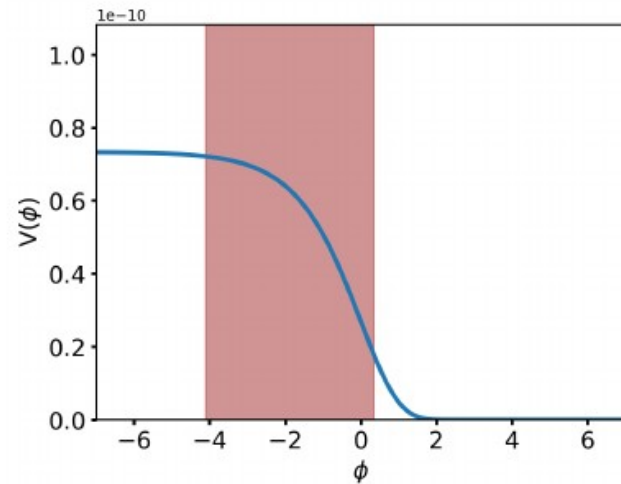
(b)  $V(\phi) = \theta_0 (\theta_1 + \log(\phi)^2)$



(c)  $V(\phi) = |\theta_0|^{\theta_1} \phi$

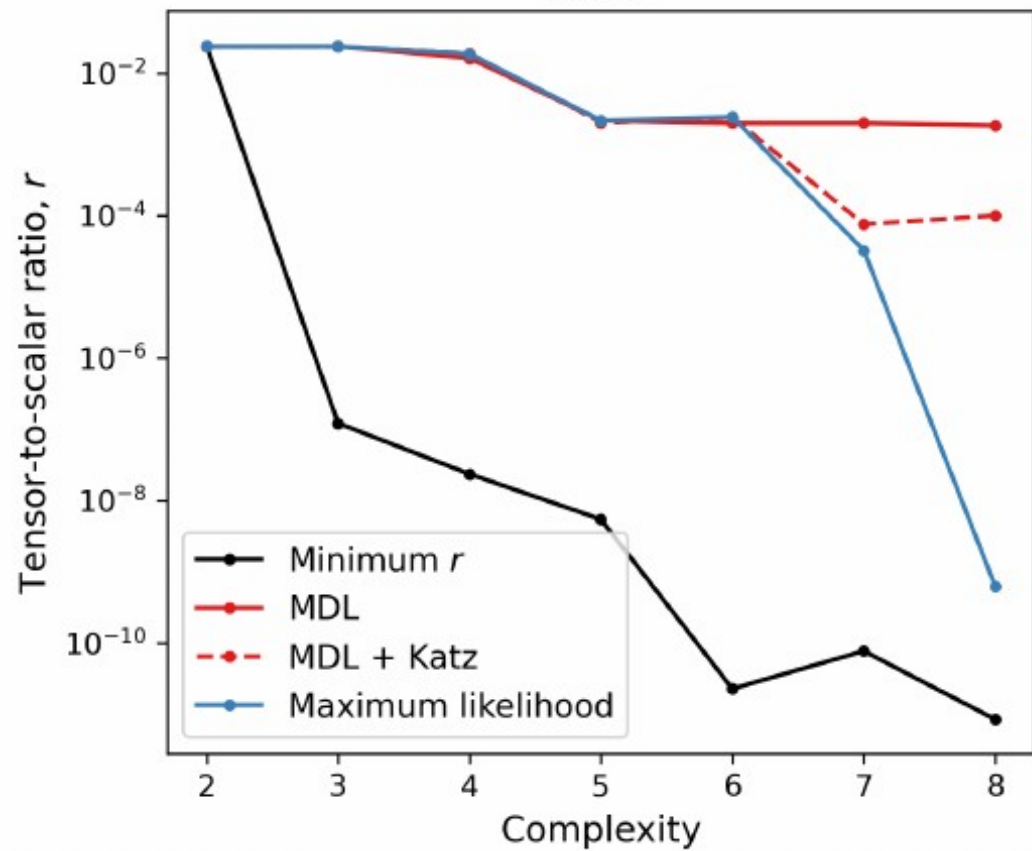


(d) Starobinsky

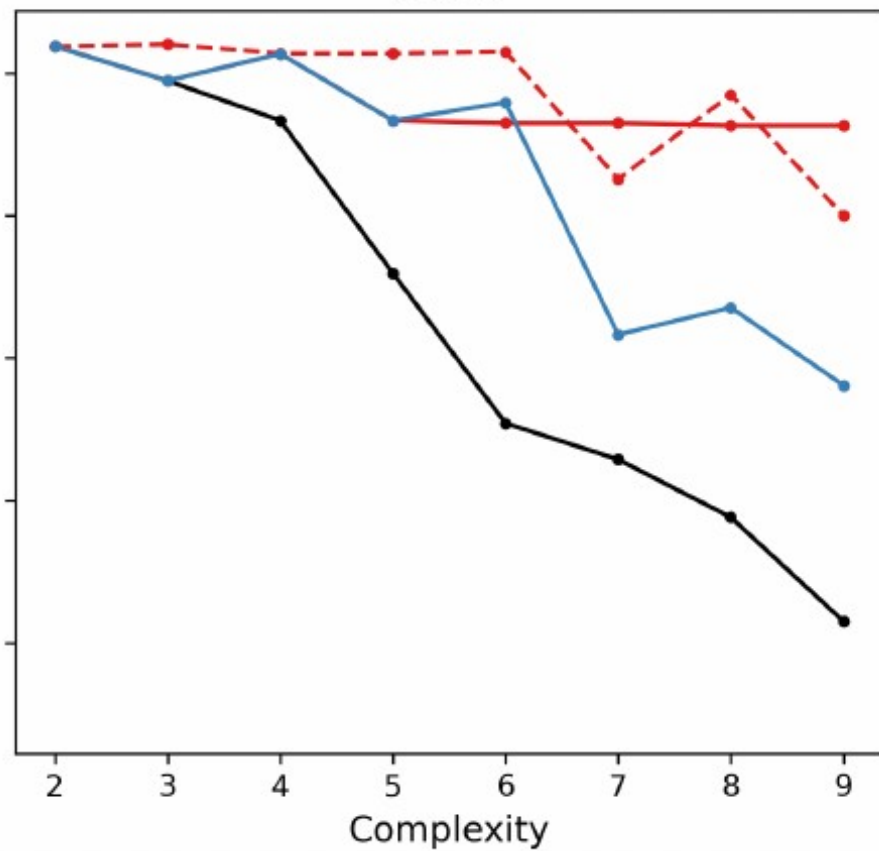


(e)  $V(\phi) = \theta_0 e^{-e^\phi}$

Set A



Set B



Set A

klog(n)

Rank	$V(\phi)$	Complexity	Prediction		Codelength			
			$n_s$	$r$	Residuals <sup>1</sup>	Function <sup>2</sup>	Parameter <sup>3</sup>	Total
1	$e^{-e^e e^\phi}$	6	0.9678	0.002	-12.65	6.59	0.00	-6.06
2	$\theta_0 e^{-e^\phi}$	5	0.9668	0.002	-12.79	6.93	0.69	-5.16
3	$e^{\frac{1}{\sin(\sqrt{\phi})}}$	5	0.9634	0.012	-12.39	8.05	0.00	-4.35
4	$e^{-3e^e e^\phi}$	6	0.9673	0.002	-12.38	8.32	0.00	-4.07
5	$\theta_0 - e^{3\phi}$	5	0.9670	$2 \times 10^{-4}$	-12.77	8.05	0.69	-4.03
6	$\theta_0 - e^{2\phi}$	5	0.9673	$5 \times 10^{-4}$	-12.74	8.05	0.69	-4.00
7	$\sin^2(\sin(\sin(\sqrt{\phi})))$	6	0.9619	0.013	-12.06	8.32	0.00	-3.74
8	$e^{e^\phi - e^e e^\phi}$	8	0.9685	0.002	-12.50	8.79	0.00	-3.71
9	$\theta_0 e^{-e^e e^\phi}$	6	0.9680	0.002	-12.62	8.32	0.69	-3.61
10	$(\theta_0 + e^\phi)^2$	5	0.9676	0.002	-12.68	8.05	1.06	-3.58
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
—	$\theta_0 \left(1 - e^{-\sqrt{2/3}\phi}\right)^2$	9	0.9678	0.003	-12.63	17.51	0.69	5.57
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
24273	$\theta_0 \phi^2$	4	0.9669	0.132	31.82	5.55	0.69	38.06
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
38717	$\theta_0 \phi^4$	5	0.9508	0.262	168.23	6.93	0.69	175.85

$$^1 - \log \mathcal{L}(\hat{\theta}) \quad ^2 k \log(n) + \sum_j \log(c_j) \quad ^3 - \frac{p}{2} \log(3) + \sum_i^p (\log(I_{ii})^{1/2} + \log(|\hat{\theta}_i|))$$

Set B

klog(n)

Rank	$V(\phi)$	Complexity	Prediction		Codelength			
			$n_s$	$r$	Residuals	Function	Parameter	Total
1	$e^{-e^{e^e\phi}}$	6	0.9678	0.002	-12.65	6.59	0.00	-6.06
2	$\theta_0 e^{-e^\phi}$	5	0.9668	0.002	-12.79	6.93	0.69	-5.16
3	$ \theta_0  e^{e^\phi}$	5	0.9674	0.002	-12.72	6.93	0.69	-5.09
4	$e^{e^\phi} e^{-e^{e^e\phi}}$	8	0.9685	0.002	-12.50	8.79	0.00	-3.71
5	$\theta_0 e^{-e^{e^\phi}}$	6	0.9680	0.002	-12.62	8.32	0.69	-3.61
6	$e^{-e^{\frac{1}{\phi}}}$	5	0.9768	0.008	-8.65	5.49	0.00	-3.16
7	$e^{\theta_0 e^{e^\phi}}$	6	0.9674	0.002	-12.72	8.32	1.36	-3.04
8	$ \theta_0  e^{e^{e^e\phi}}$	6	0.9678	0.002	-12.65	8.32	1.39	-2.94
9	$e^{e^\phi} e^{-e^{e^e\phi}}$	9	0.9685	0.002	-12.50	9.89	0.00	-2.62
10	$ \theta_0 ^{\frac{1}{\phi}}$	4	0.9756	0.019	-8.77	5.55	0.69	-2.53
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
1272	$\theta_0 \left(1 - e^{-\sqrt{2/3}\phi}\right)^2$	9	0.9678	0.003	-12.63	17.51	0.69	5.57
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
8697	$\theta_0 \phi^2$	4	0.9669	0.132	31.82	5.55	0.69	38.01
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
10839	$\theta_0 \phi^4$	5	0.9508	0.262	168.23	6.93	0.69	178.81

Set A

Katz

Rank	$V(\phi)$	Complexity	Prediction		Codelength			
			$n_s$	$r$	Residuals	Function	Parameter	Total
1	$\theta_0 (\theta_1 + \log(\phi))^2$	7	0.9649	$8 \times 10^{-5}$	-12.90	9.28	1.39	-2.23
2	$\frac{1}{(\theta_0 + e^\phi)^2}$	6	0.9654	0.002	-12.88	10.31	1.06	-1.51
3	$e^{\frac{\theta_0}{\phi}}$	4	0.9756	0.019	-8.77	6.98	0.69	-1.10
4	$\theta_0 (\theta_1 - e^\phi)^2$	7	0.9676	0.002	-12.68	10.77	1.39	-0.53
5	$\frac{\theta_0}{\phi + \log\left(\frac{ \theta_1 }{\phi}\right)}$	8	0.9649	$1 \times 10^{-4}$	-12.90	11.06	1.39	-0.45
6	$(\theta_0 - e^\phi)^2$	5	0.9676	0.002	-12.68	11.59	1.06	-0.04
7	$\theta_0 \log(\phi)$	4	0.9783	0.024	-6.39	6.10	0.69	0.40
8	$\frac{1}{\theta_0 + e^{\frac{\theta_1}{\phi}}}$	7	0.9649	0.001	-12.90	11.23	2.08	0.42
9	$\theta_0 - \frac{\theta_1}{\phi}$	5	0.9768	0.010	-8.62	8.12	1.65	1.16
10	$\theta_0 e^{\theta_1 e^\phi}$	7	0.9668	0.002	-12.79	12.74	1.39	1.34
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
16653	$\theta_0 \phi^2$	4	0.9669	0.132	31.82	5.49	0.69	38.01
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
37635	$\theta_0 \phi^4$	5	0.9508	0.262	168.23	7.15	0.69	176.08
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
—	$\theta_0 \left(1 - e^{-\sqrt{2/3}\phi}\right)^2$	9	0.9678	0.003	-12.63	11.42	0.69	-0.52



Set B

Katz

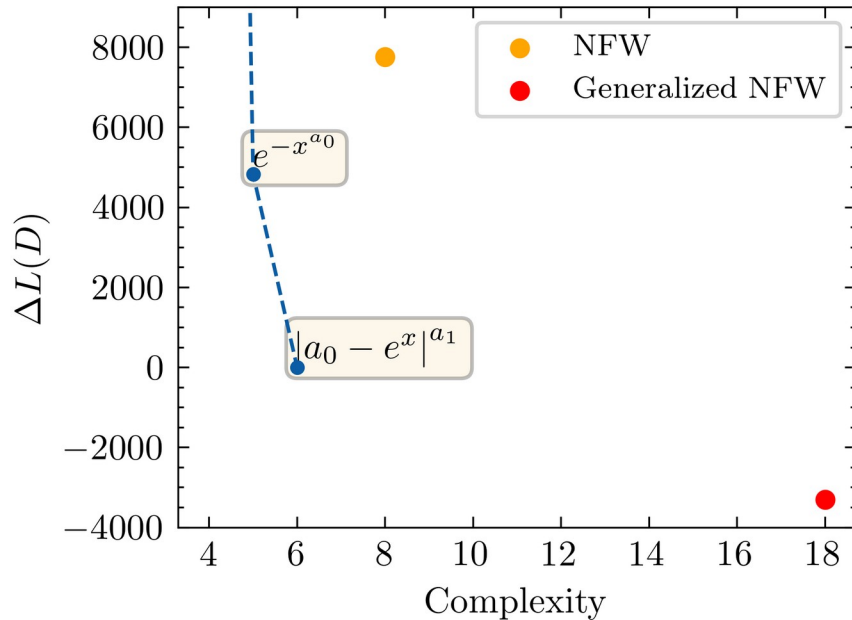
Rank	$V(\phi)$	Complexity	Prediction		Codelength			
			$n_s$	$r$	Residuals	Function	Parameter	Total
1	$\theta_0 \phi^{\frac{\theta_1}{\phi}}$	7	0.9649	$3 \times 10^{-4}$	-12.90	8.92	1.39	-2.59
2	$\theta_0 (\theta_1 + \phi^\phi)$	7	0.9649	$6 \times 10^{-4}$	-12.90	10.12	1.39	-1.39
3	$\theta_0 \phi^{\phi^{\theta_1}}$	7	0.9649	0.002	-12.89	9.55	2.71	-0.63
4	$\phi^{\theta_0} e^{\frac{\theta_1}{\phi}}$	8	0.9647	0.005	-12.83	9.37	2.83	-0.62
5	$\left(\phi + e^{\frac{\theta_0}{\phi}}\right)^{\theta_1}$	8	0.9636	0.012	-12.48	10.05	1.83	-0.60
6	$ \theta_0 ^{\frac{\theta_1}{\phi}}$	5	0.9756	0.019	-8.77	6.88	1.39	-0.50
7	$e^{\theta_0  \theta_1 ^\phi}$	8	0.9653	0.020	-11.66	8.11	3.26	-0.30
8	$e^{\frac{\theta_0}{\phi}}$	4	0.9756	0.019	-8.77	8.04	0.69	-0.03
9	$\theta_0 e^{\frac{\theta_1}{\phi}}$	6	0.9653	0.020	-8.82	7.84	1.39	0.40
10	$\theta_0 \log(\phi)$	4	0.9783	0.024	-6.39	6.26	0.69	0.57
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
12	$\theta_0 \left(1 - e^{-\sqrt{2/3}\phi}\right)^2$	9	0.9678	0.003	-12.63	12.64	0.69	0.70
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
5401	$\theta_0 \phi^2$	4	0.9669	0.132	31.82	4.67	0.69	38.03
⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮	⋮
8938	$\theta_0 \phi^4$	5	0.9508	0.262	168.23	4.67	0.69	180.84

# Inferring halo density profiles



## *From simulations*

Richard Stiskalek



## *From observations*

Alicia Martin

